



Dark Matter search with the BDX-MINI experiment

Marco Spreafico

On behalf of BDX Collaboration

06 - 06 - 2026



- ① Introduction
 - Light Dark Matter
 - Dark Photon model
 - Beam Dump Experiments
- ② BDX
- ③ BDX-MINI
 - Experimental setup
 - Detector
 - Background
 - Detector performance
 - Background
 - Data analysis
 - Results
- ④ Outlook



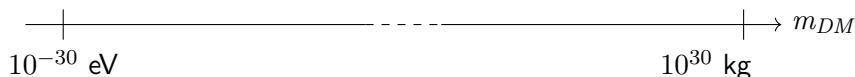
Dark Matter Problem

Astrophysical observations suggest existence of DM

- Information only from gravitational interaction
- ⇒ No clue on DM nature

Common assumption: **thermal origin of DM**

- DM we see comes from an epoch of thermodynamical equilibrium with SM
- constrain on available mass range
- strong constraint on viable DM → SM interaction



Dark Matter Problem

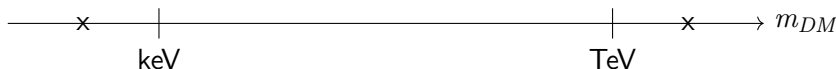
Astrophysical observations suggest existence of DM

- Information only from gravitational interaction
- ⇒ No clue on DM nature

Common assumption: **thermal origin of DM**

- DM we see comes from an epoch of thermodynamical equilibrium with SM
- constrain on available mass range
- strong constraint on viable DM → SM interaction

Thermal DM



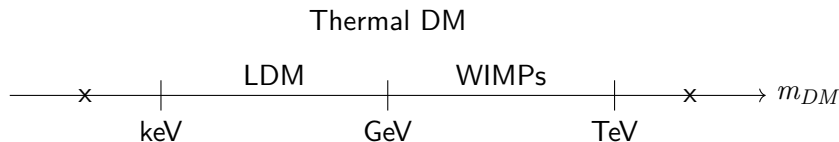
Dark Matter Problem

Astrophysical observations suggest existence of DM

- Information only from gravitational interaction
- ⇒ No clue on DM nature

Common assumption: **thermal origin of DM**

- DM we see comes from an epoch of thermodynamical equilibrium with SM
- constrain on available mass range
- strong constraint on viable DM → SM interaction



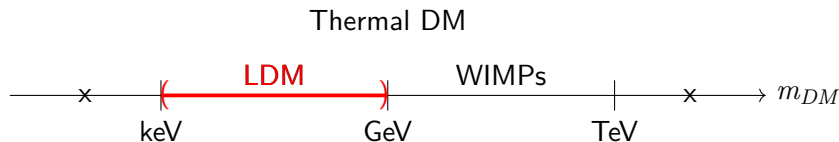
Dark Matter Problem

Astrophysical observations suggest existence of DM

- Information only from gravitational interaction
- ⇒ No clue on DM nature

Common assumption: **thermal origin of DM**

- DM we see comes from an epoch of thermodynamical equilibrium with SM
- constrain on available mass range
- strong constraint on viable DM → SM interaction



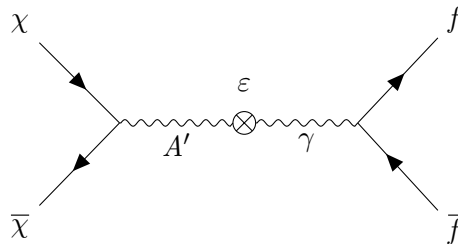
Light Dark Matter - Dark Photon model

Simplest possibility: "vector portal"¹

→ $U(1)$ gauge boson (**dark photon**) coupling to electric charge

$$\mathcal{L}_{LDM} \sim g_D A'_\mu J^\mu_\chi + \varepsilon e A'_\mu J^\mu_{EM} + [\dots]$$

Annihilation in SM:



¹ For a comprehensive review: 1707.04591, 2005.01515, 2011.02157

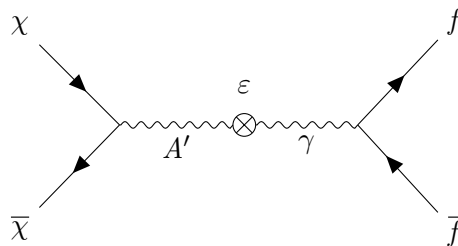
Light Dark Matter - Dark Photon model

Simplest possibility: "vector portal"¹

→ $U(1)$ gauge boson (**dark photon**) coupling to electric charge

$$\mathcal{L}_{LDM} \sim g_D A'_\mu J^\mu_\chi + \varepsilon e A'_\mu J^\mu_{EM} + [\dots]$$

Annihilation in SM:



Model parameters:

- Dark Photon mass $m_{A'}$, coupling to SM ε
- Dark Matter mass m_χ , coupling to DM g_D
($\alpha_D \equiv g_D^2/4\pi$)

$$y \equiv \frac{g_D^2 \varepsilon^2 e^2}{4\pi} \left(\frac{m_\chi}{m_{A'}} \right)^4 \sim \langle \sigma v \rangle_{\text{relic}} m_\chi^2$$

¹ For a comprehensive review: 1707.04591, 2005.01515, 2011.02157

Light Dark Matter

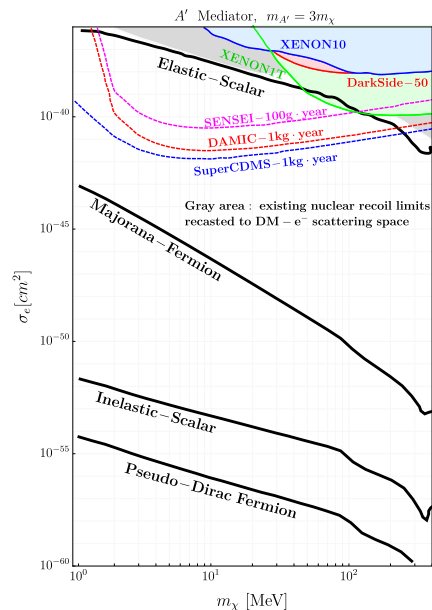
Direct detection not well suited for sub-GeV DM searches:

- DD experiments optimized for $m_\chi > \text{GeV}$
 - $E_R \propto m_\chi^2 / m_N$
 - ⇒ very low recoil energy
- LDM-SM interaction cross section depends on impinging particle velocity
 - DD sensitivity strongly model-dependent

LDM at accelerators

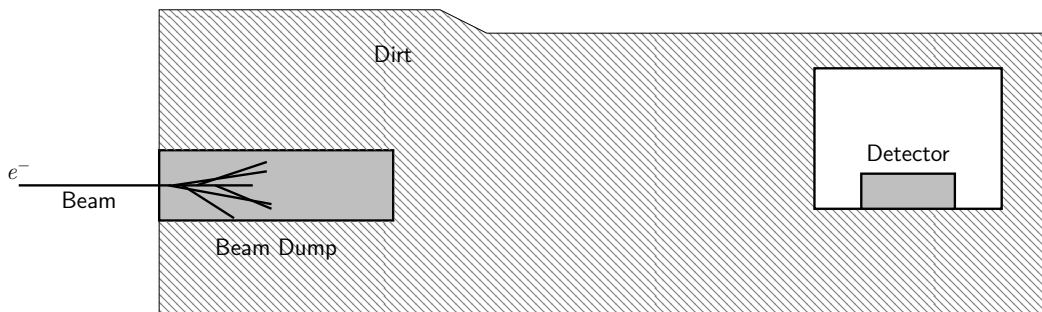
Accelerator based experiments at the *intensity frontier* uniquely suited to search for LDM:

- High intensity ⇒ increased possibility of DM production
- Production of relativistic DM ⇒ testing different models



Beam Dump experiments

Beam dump experiments: direct detection of LDM produced by beam impinging on fixed target (beam dump)²



χ production

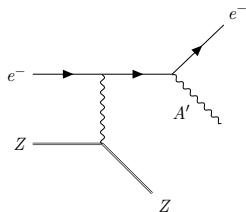
- e^- beam impinging on target
- χ from decay of A' produced in the dump

² Izaguirre et al., Phys. Rev. D 88, 114015

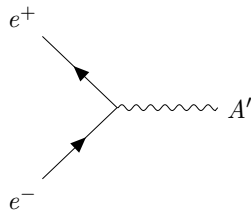
Dark Photon model

Dark Matter production³

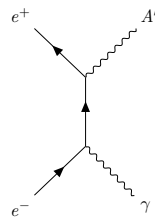
A' -strahlung



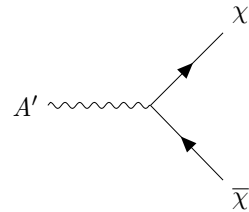
Resonant e^+e^-
annihilation



Non-resonant
 e^+e^- annihilation



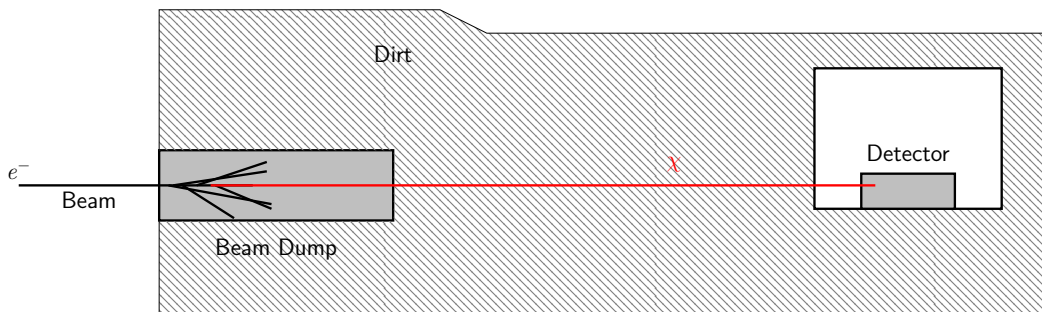
A' decay



³ L. Marsicano et al., Phys. Rev. Lett. 121, 041802

Beam Dump experiments

Beam dump experiments: direct detection of LDM produced by beam impinging on fixed target (beam dump)²



χ production

- e^- beam impinging on target
- χ from decay of A' produced in the dump

χ interaction

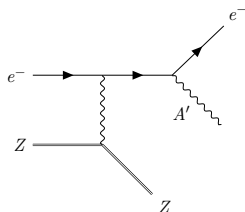
- Detector placed behind the dump (~ 10 m)
- χ scattering through A' exchange

² Izaguirre et al., Phys. Rev. D 88, 114015

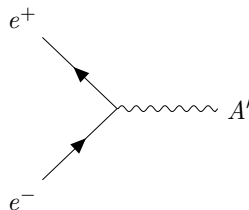
Dark Photon model

Dark Matter production³

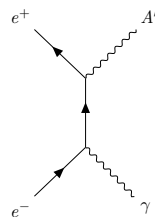
A' -strahlung



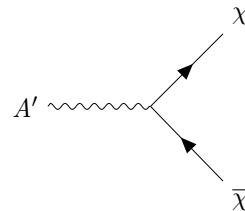
Resonant e^+e^-
annihilation



Non-resonant
 e^+e^- annihilation

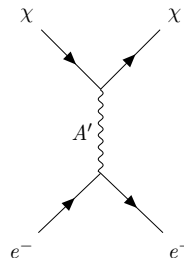


A' decay

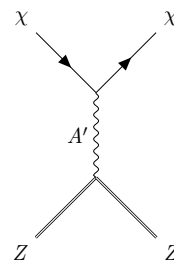


Dark Matter interaction

Scattering on e^-



Scattering on
nuclei



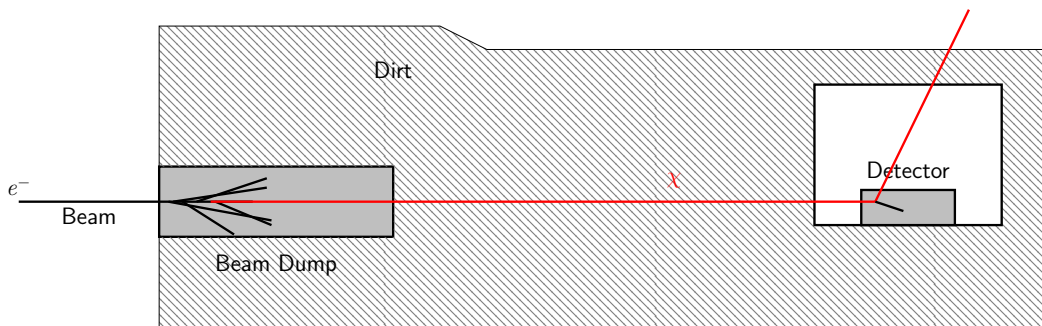
With \sim GeV DM recoil energy:

- $E_{rec}^e \sim 100$ MeV
- $E_{rec}^N \sim$ MeV

³ L. Marsicano et al., Phys. Rev. Lett. 121, 041802

Beam Dump experiments

Beam dump experiments: direct detection of LDM produced by beam impinging on fixed target (beam dump)²



χ production

- e^- beam impinging on target
- χ from decay of A' produced in the dump

χ interaction

- Detector placed behind the dump (~ 10 m)
- χ scattering through A' exchange

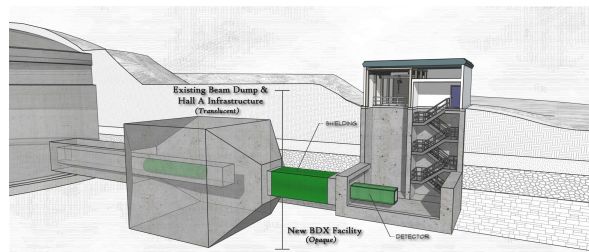
Number of signal events: $S \propto \frac{\alpha_D \varepsilon^4}{m_{A'}^4}$

² Izaguirre et al., Phys. Rev. D 88, 114015

BDX

BDX is a **JLab experiment** approved by PAC46

- unique experiment able to produce and detect LDM
- beam dump experiment specifically aimed at LDM searches



Experiment designed with two goals:

LDM production and detection

- Exploit CEBAF high-intensity beam
- Medium-high energy beam
- EM shower detection capability
- Fully parasitic

Minimize background

- Shielding to filter beam-related background
- Multi layer veto for cosmogenic background
- Segmented detector
- Time resolution for detector-veto coincidence

BDX - Experimental Setup

JLAB offers the best condition for BDX:

- Medium high energy beam (11 GeV)
- High electron beam current ($65 \mu\text{A}$)
- Fully parasitic wrt Hall-A physic program (Moeller)

New facility to be built in front of Hall-A beam dump:

- new underground ($\sim 8 \text{ m}$) hall
- 25 m downstream of Hall-A beam dump
- passive shielding ($\sim 6.6 \text{ m iron}$) to reduce beam related background
- $\sim 10 \text{ m}$ water equivalent overburden to reduce cosmogenic background



BDX - Detector

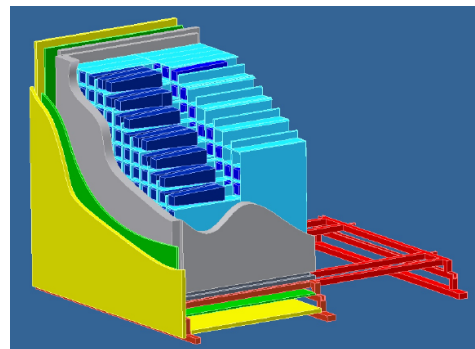
Detector design

Electromagnetic calorimeter:

- homogeneous calorimeter made with CsI(Tl) crystals read by SiPM

Veto system:

- hermetic multi layer veto
- 2 layer of plastic scintillator counters read by WLS fibers and SiPM
- 5 cm lead vault between veto and calorimeter



Modular detector arrangement:

- 1 module: 10×10 crystals
 - $50 \times 50 \text{ cm}^2$ front face, 30 cm long
 - Module surrounded by veto
- total: 8 modules ($\sim 2.6 \text{ m}$ length)

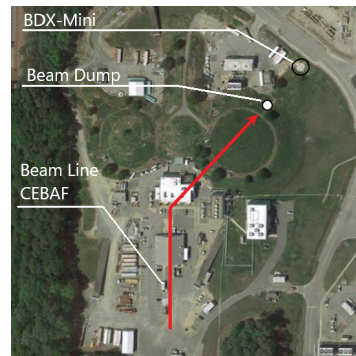
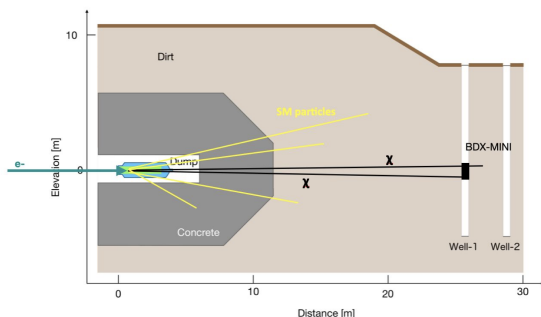
Signal detection:

- EM shower ($\gtrsim 100 \text{ MeV}$) and no corresponding activity in the active veto
- Signal efficiency $\sim 20\%$

BDX-MINI experiment

Pilot version of BDX:

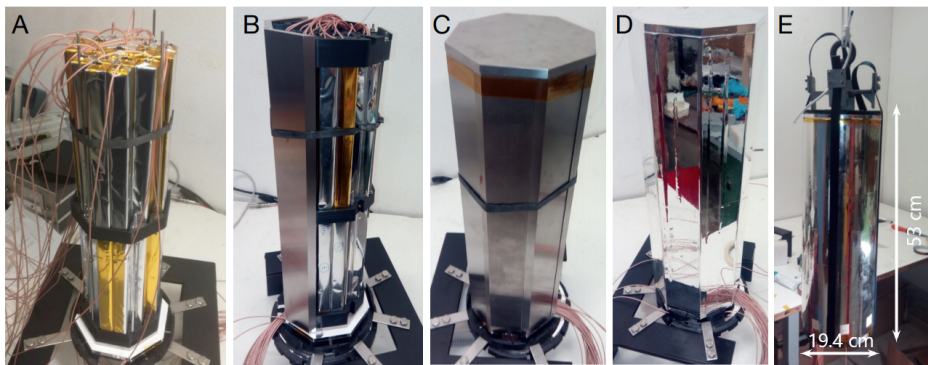
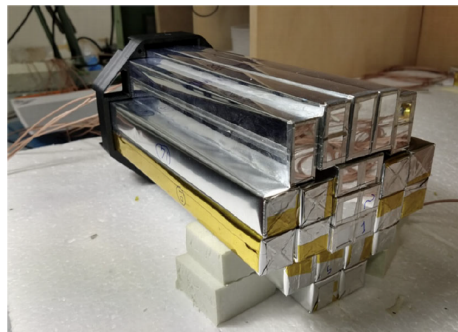
- 2.56 GeV e^- beam (10 GeV beam used for calibration)
- current up to 150 μA
- measurement alternating beam on and beam off data (beam on time $\sim 50\%$)
- accumulated 2.54×10^{21} EOT
- beam off measurements for cosmic background characterization



BDX-MINI detector

Electromagnetic calorimeter (ECal):

- 44 PbWO_4 crystals ($4 \times 10^{-3} \text{ m}^3$ active volume)
- SiPM readout



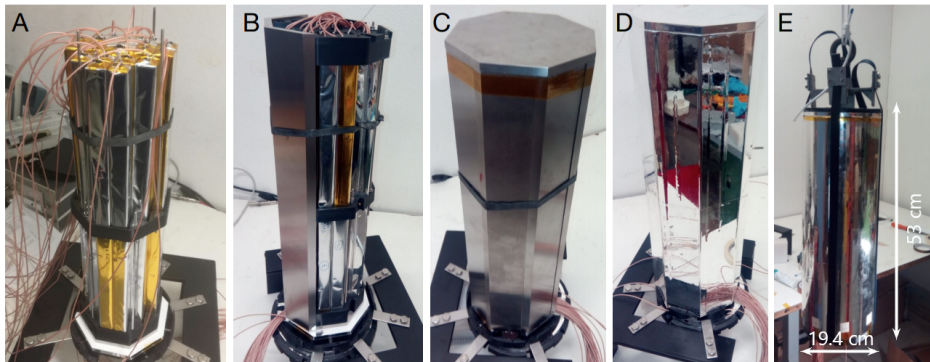
BDX-MINI detector

Electromagnetic calorimeter (ECal):

- 44 PbWO_4 crystals ($4 \times 10^{-3} \text{ m}^3$ active volume)
- SiPM readout

Veto system

- Active veto:
 - Octagonal (IV) and cylindrical (OV) plastic scintillator
 - Optically continuous
 - SiPM readout + WLS fibers light collection
- Passive tungsten shielding
 - 0.8 cm thick



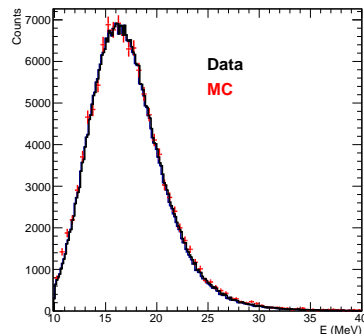
Detector performance

Detector performance:

- Calibration with special 10 GeV run
- Stability monitored with cosmic μ

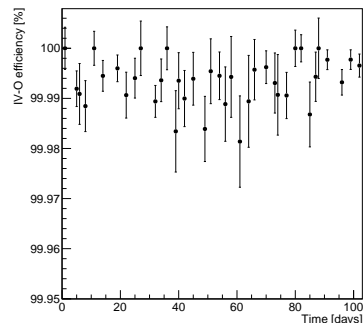
ECal energy response

- Energy calibration determined from 10 GeV data
 - compared secondary μ between data and MC
- Stability monitored with beam off data
 - Selected penetrating μ with Landau peak



VETO stability

- Veto efficiency monitored with cosmic muons
 - *tag-and-probe* method
 - trajectories selected with ECal energy deposition



Backgrounds

Two main sources of background:

Beam related background

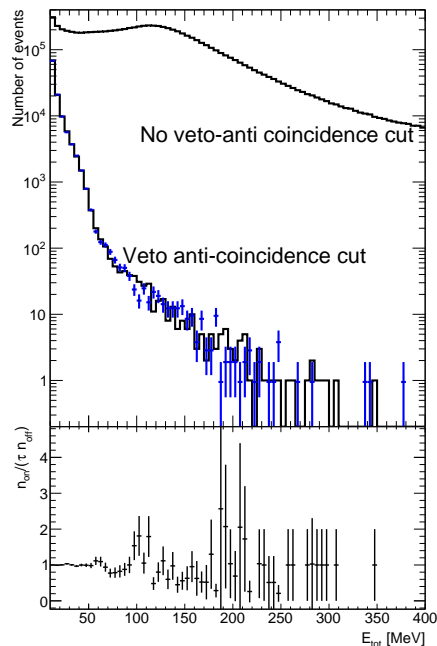
Yield estimate through MC simulations
(FLUKA+GENIE+Geant4)

- MC simulation validated with in-situ measurement
- ν only background \rightarrow negligible: $5.8 \times 10^{-23} \nu/\text{EOT}$

Cosmogenic background

- Continuous measurement \Rightarrow no rejection
- Charged particles rejected requiring veto anti-coincidence
- Further suppression can be achieved using energy cut

Measured cosmic background



Data analysis

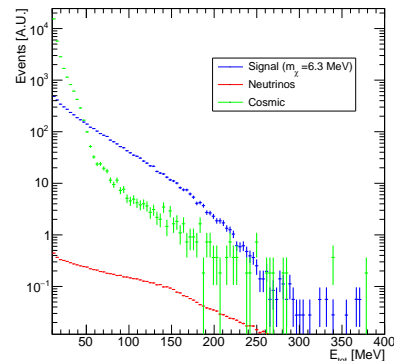
Approach

Blind analysis: experiment sensitivity optimized with MC simulations and beam-off data

Model: ON-OFF problem

$$\mathcal{L} = \prod_j \left[P(n_{\text{on}}^j; \mu_c^j + \mu_\nu^j + \alpha^j \cdot S) \cdot P(n_{\text{off}}^j; \mu_c^j \cdot \tau) \right]$$

- $n_{\text{on}}^j, n_{\text{off}}^j$: measured number of events during beam-on/beam-off intervals ($\tau = T_{\text{off}}/T_{\text{on}}$)
- μ_c^j/μ_ν^j : expected number of cosmogenic/beam-related backgrounds events
- μ_ν^j evaluated via MC, μ_c^j treated as nuisance parameter
- Data binned according to total energy deposition in ECal



Systematic uncertainties: described via ancillary pseudo-measurement factors in \mathcal{L} with Gaussian constraint

→ one-sided profile-likelihood test statistics to evaluate upper limit on S

Sensitivity optimization

Idea: improve sensitivity considering effect of data analysis cuts on background minimization and signal maximization

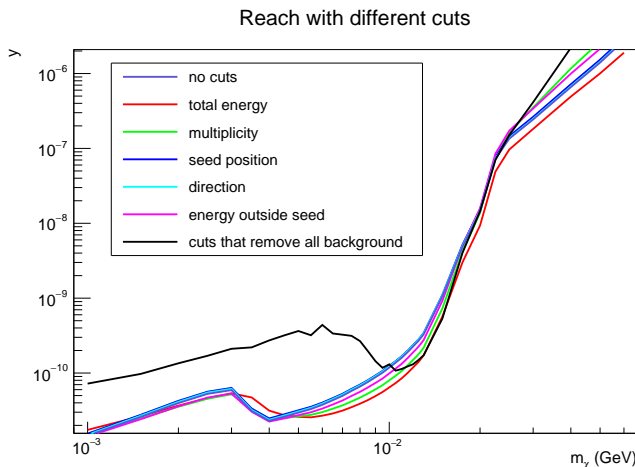
- beam-on data extrapolated from beam-off data
- used only events in anti-coincidence with veto
- evaluated upper limit on signal yield with different cuts and converted to exclusion limit

Tested different cuts

Best condition

- Anti-coincidence with veto
- $E_{tot} > 40$ MeV
- Data split into 8 energy bins

⇒ Sensitivity highly enhanced



Unblinding

Last step: unblinding and analysis of beam on data

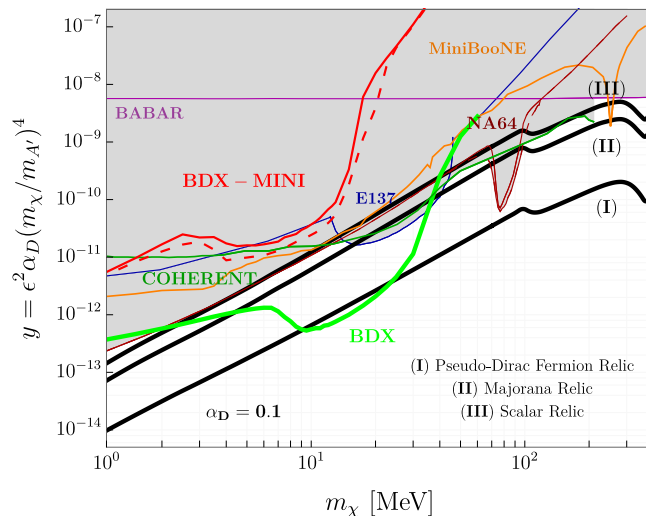
Experimental results

Yields (for $N_{EOT} = 2.54 \cdot 10^{21}$)

- $N_{on} = 3623$
- $N_{off} = 3822$ ($\tau = 1.054$)

No excess is observed

→ evaluated 90% exclusion limit in the LDM parameter space



Conclusions

- Dark matter in the MeV-to-GeV range is largely unexplored
- Beam dump experiments at the intensity frontier uniquely suited for DM searches
 - **Beam Dump eXperiment** at JLab: search for Dark Sector particles in the MeV-GeV mass range
- **BDX-MINI**: pilot version of BDX
 - First modern beam dump experiment searching for Light Dark Matter
 - Detector optimized for LDM searches
 - Analysis aimed to LDM detection
 - Evaluated exclusion limit → competitive to flagship experiments
- Beam dump experiment at e beam dump highly sensitive to Light Dark Matter in the MeV-GeV range
 - BDX-MINI remarkable results demonstrate that BDX is a mature, ready-to-run experiment (after the construction of a new underground experimental hall)