



# Super-Kamiokande physics & astrophysics results

Magdalena Posiadala-Zezula  
University of Warsaw  
for the Super-Kamiokande Collaboration

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# The Super-Kamiokande Collaboration

2023 Toyama meeting



Kamioka Observatory, ICRR, Univ. of Tokyo, Japan  
 RCCN, ICRR, Univ. of Tokyo, Japan  
 University Autonoma Madrid, Spain  
 BC Institute of Technology, Canada  
 Boston University, USA  
 University of California, Irvine, USA  
 California State University, USA  
 Chonnam National University, Korea  
 Duke University, USA  
 Gifu University, Japan  
 GIST, Korea  
 University of Glasgow, UK  
 University of Hawaii, USA  
 IBS, Korea  
 IFIRSE, Vietnam  
 Imperial College London, UK  
 ILANCE, France/Japan

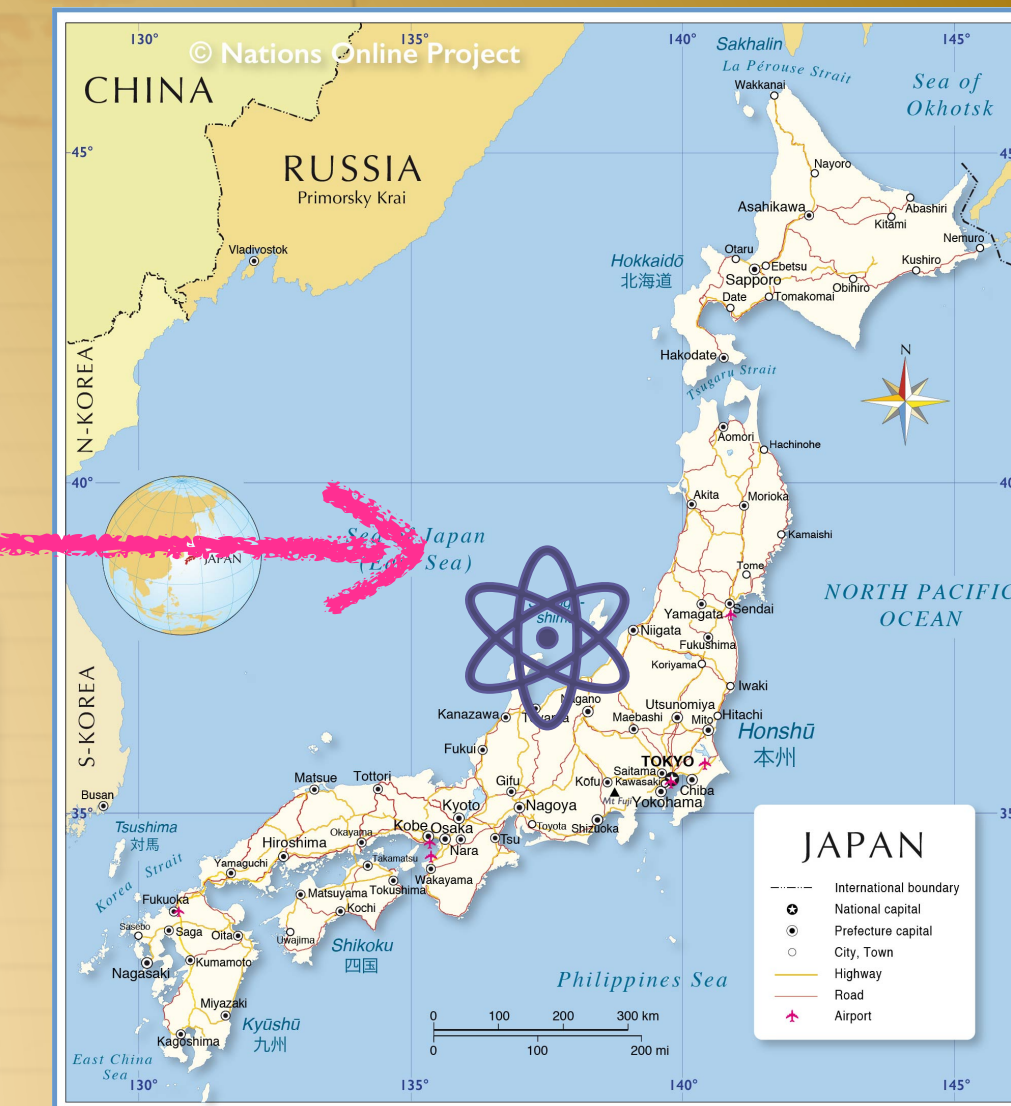
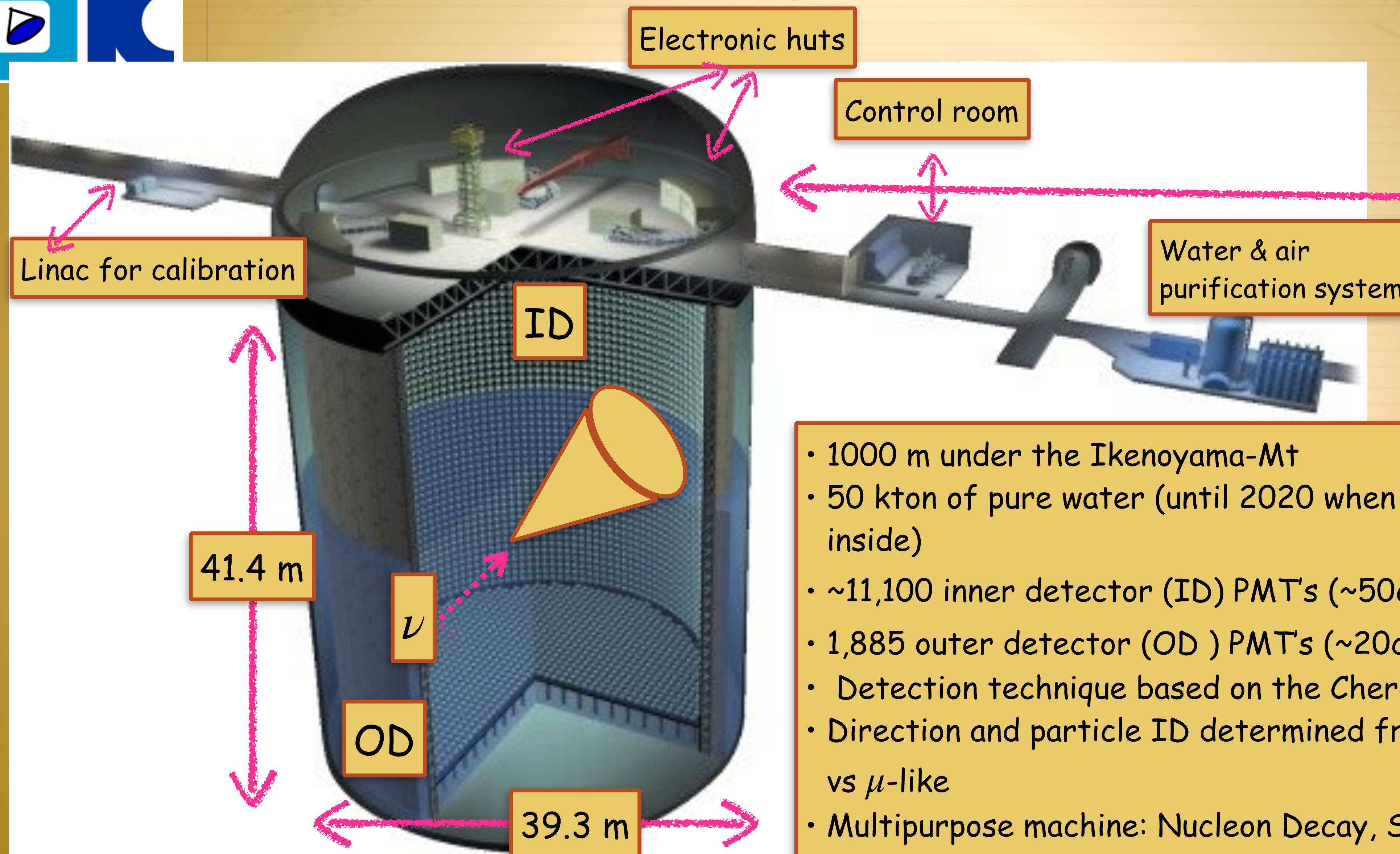
INFN Bari, Italy  
 INFN Napoli, Italy  
 INFN Padova, Italy  
 INFN Roma, Italy  
 Kavli IPMU, The Univ. of Tokyo, Japan  
 Keio University, Japan  
 KEK, Japan  
 King's College London, UK  
 Kobe University, Japan  
 Kyoto University, Japan  
 University of Liverpool, UK  
 LLR, Ecole polytechnique, France  
 Miyagi University of Education, Japan  
 ISEE, Nagoya University, Japan  
 NCBJ, Poland  
 Nihon University, Japan  
 Okayama University, Japan

Osaka Electro-Communication Univ., Japan  
 University of Oxford, UK  
 Rutherford Appleton Laboratory, UK  
 Seoul National University, Korea  
 University of Sheffield, UK  
 Shizuoka University of Welfare, Japan  
 Sungkyunkwan University, Korea  
 Stony Brook University, USA  
 Tohoku University, Japan  
 Tokai University, Japan  
 The University of Tokyo, Japan  
 Tokyo Institute of Technology, Japan  
 Tokyo University of Science, Japan  
 TRIUMF, Canada  
 Tsinghua University, China  
 University of Warsaw, Poland  
 Warwick University, UK  
 The University of Winnipeg, Canada  
 Yokohama National University, Japan

~230 collaborators from 53 institutes in 11 countries



# Super-Kamiokande detector



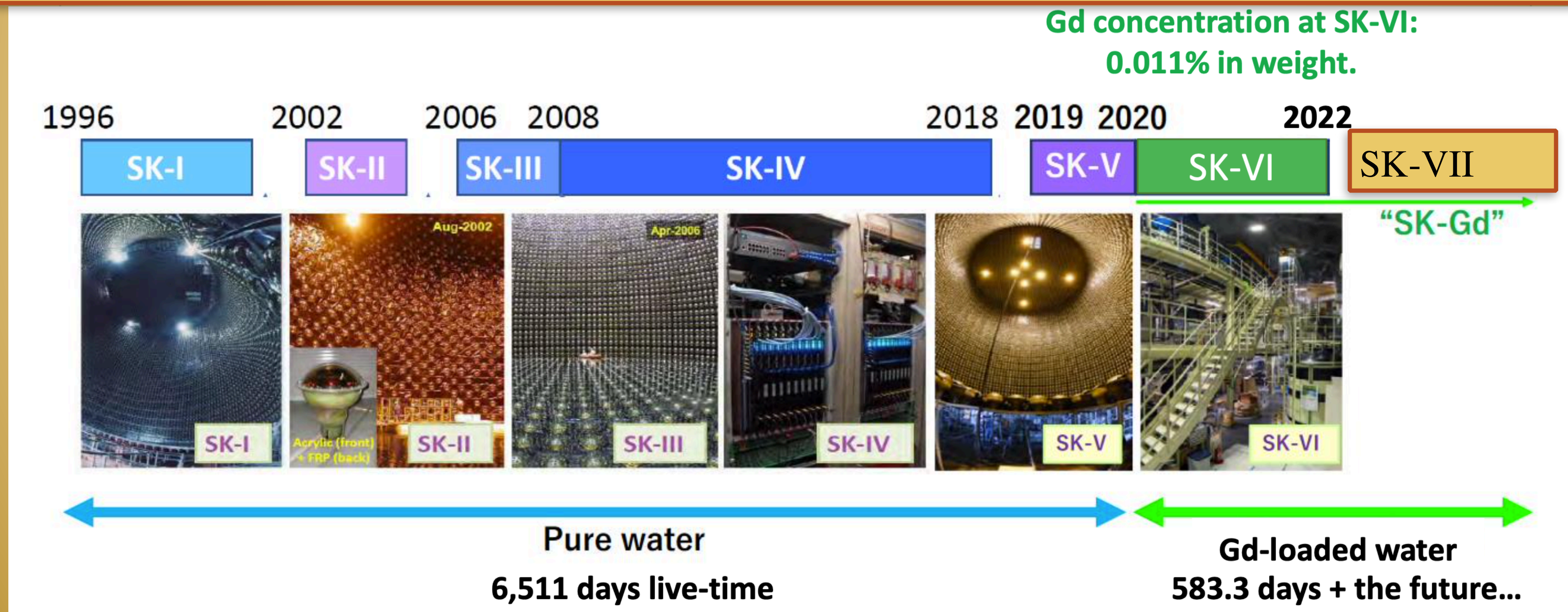
- 1000 m under the Ikenoyama-Mt
- 50 kton of pure water (until 2020 when Gd sulfate was added inside)
- ~11,100 inner detector (ID) PMT's (~50cm  $\phi$ )
- 1,885 outer detector (OD ) PMT's (~20cm  $\phi$ )
- Detection technique based on the Cherenkov radiation
- Direction and particle ID determined from the ring pattern:  $e$ -like vs  $\mu$ -like
- Multipurpose machine: Nucleon Decay, Solar and Supernova Neutrinos, Atmospheric Neutrinos, Far detector for T2K





# The Super-Kamiokande experiment

- Super-Kamiokande has been taking data since 1996 and has come through seven run periods
- Densely packed PMTs (40% / 20% for SK-II) and good water quality provide excellent sensitivity for various physics targets.
- In 2020 we have added Gd sulfate to the water in order to increase the sensitivity for neutron capture.





# Atmospheric neutrinos

- Neutrinos are produced when cosmic protons interact with the nuclei in the atmosphere:
  - with wide range of energy MeV- PeV produced isotropically about the Earth atmosphere
  - travel length varies 10km~1300km

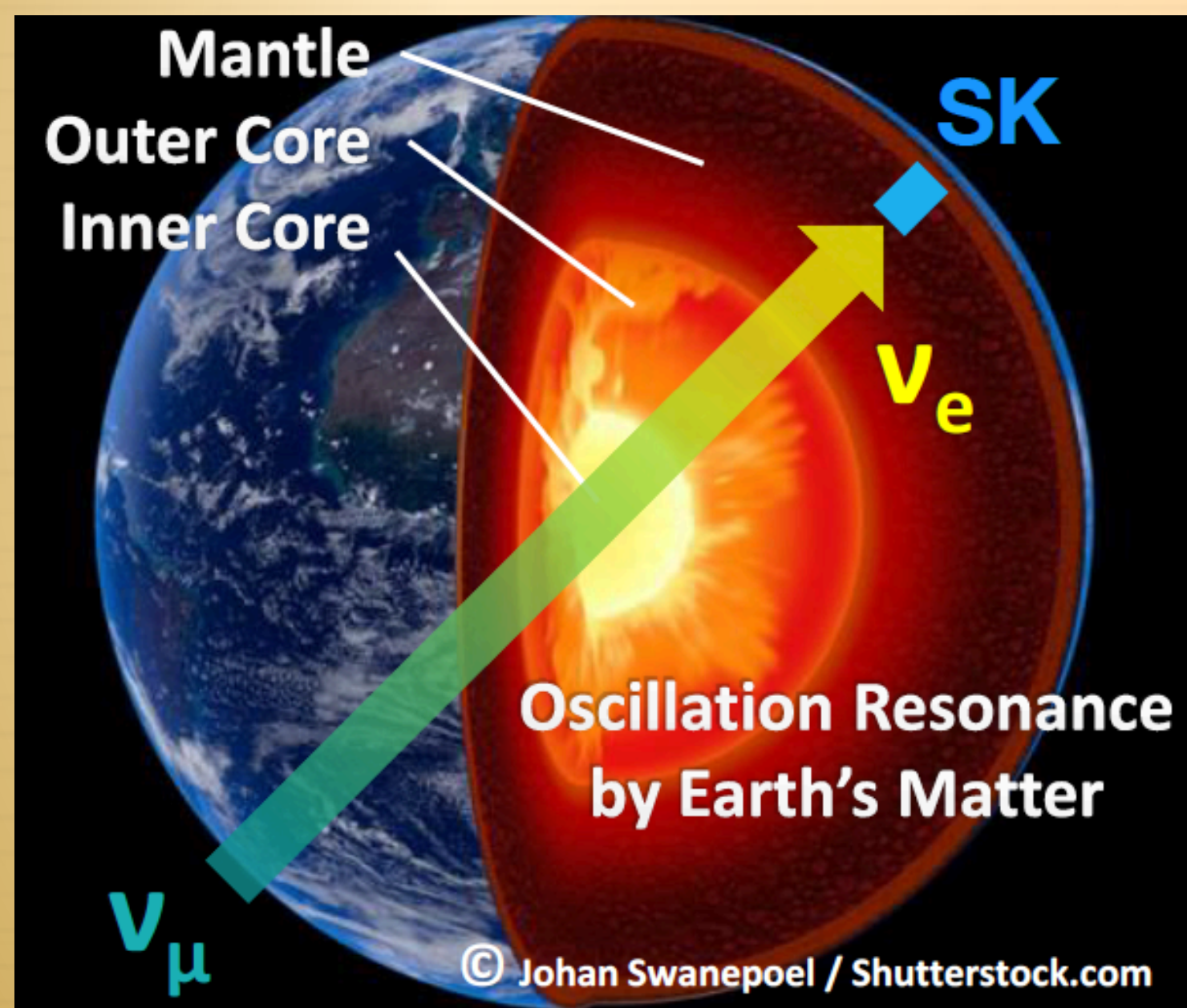
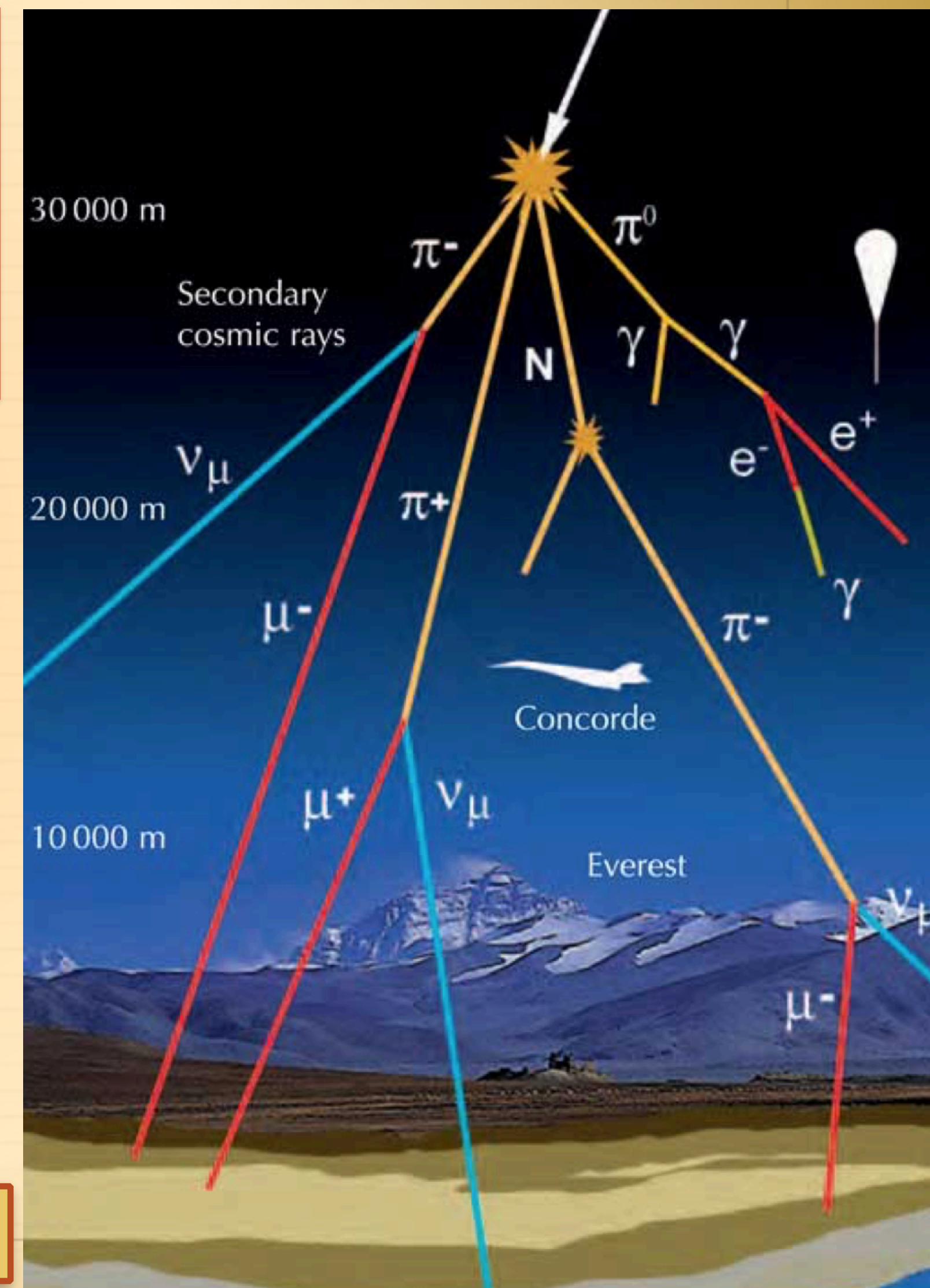


Photo: <https://physicsopenlab.org/2016/01/10/cosmic-muons-decay/>





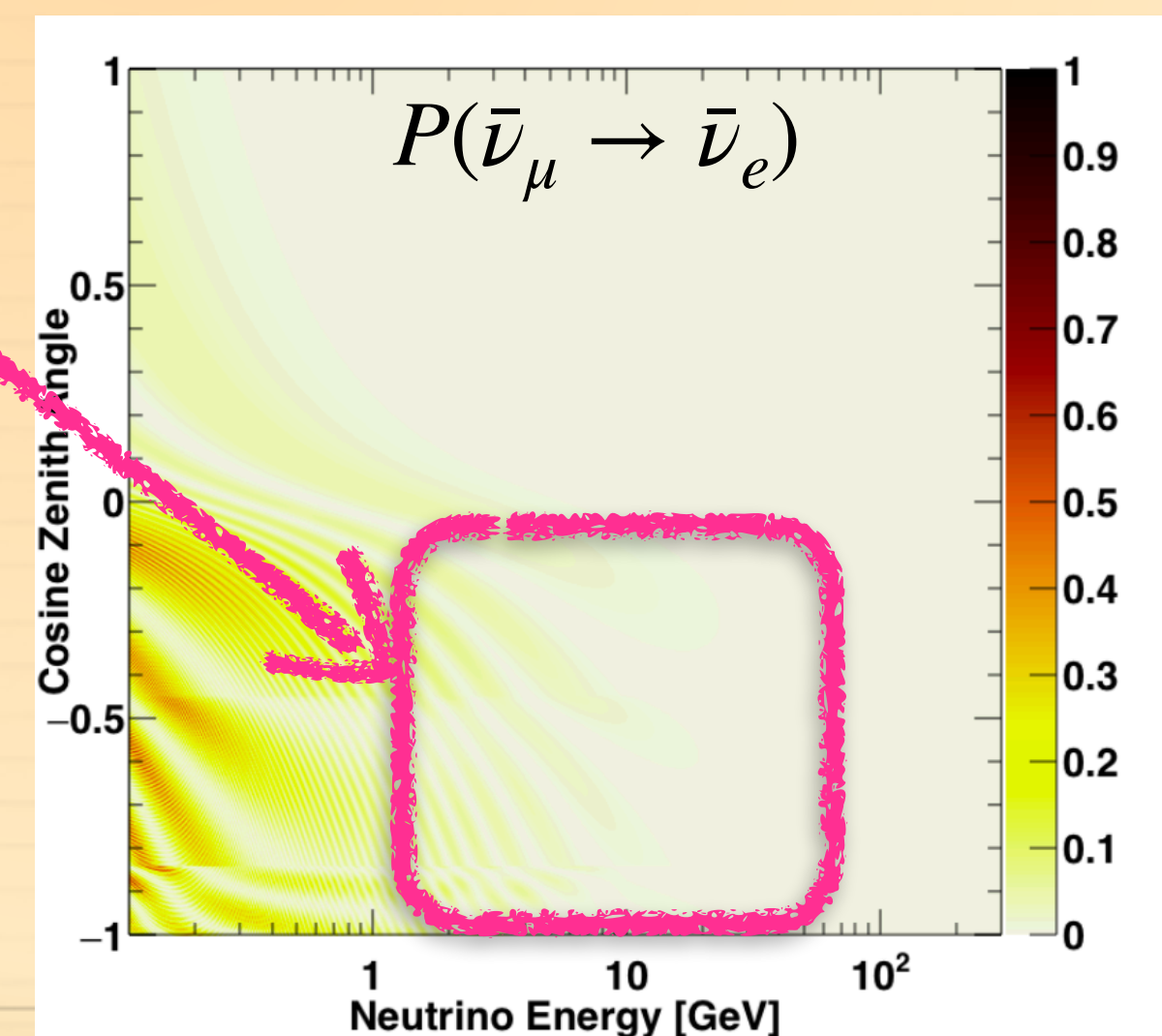
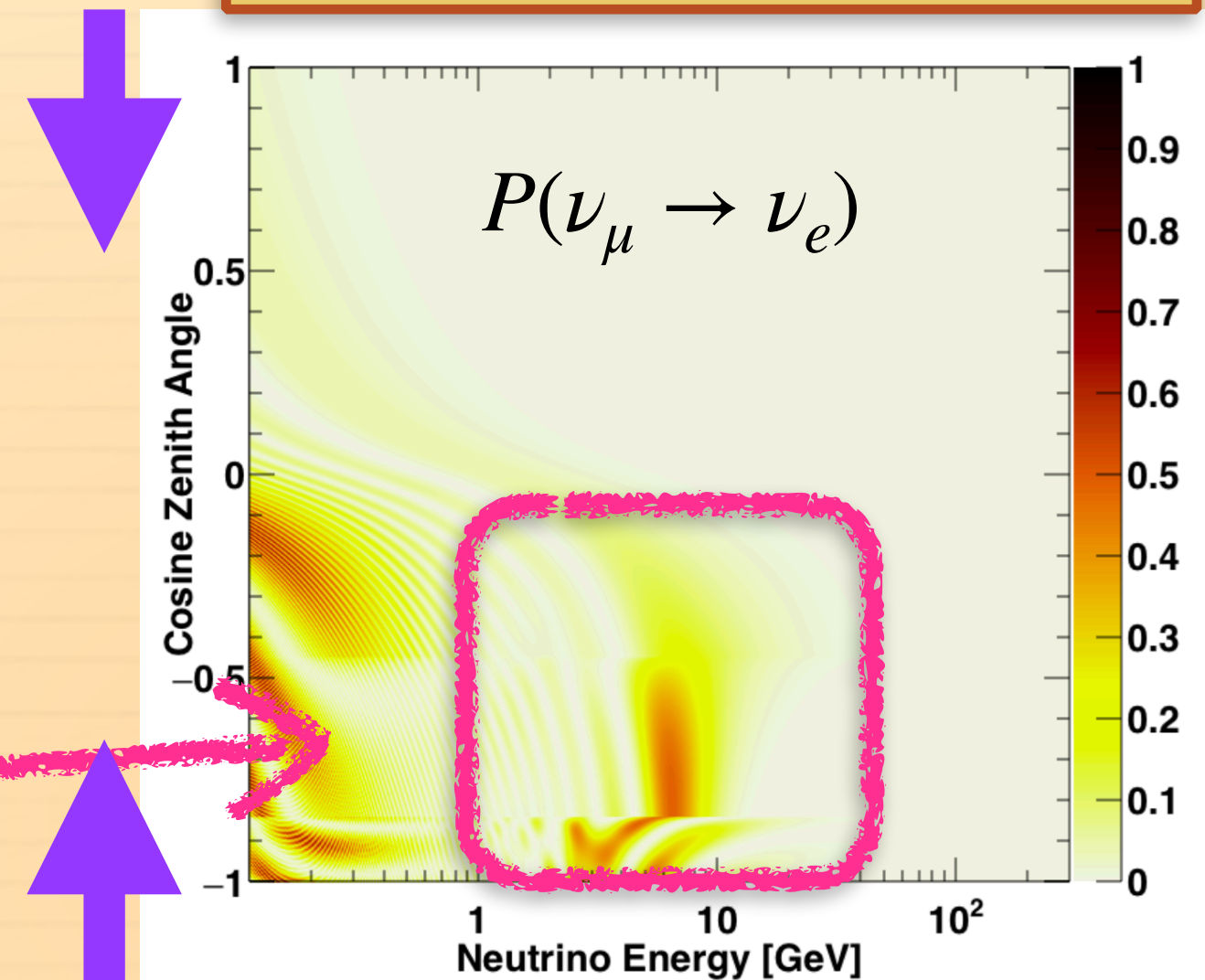
# Atmospheric neutrino oscillations

★ Oscillograms plotted with:  $\Delta m_{21}^2 = 7.7 \times 10^{-5}$ ,  
 $\sin^2 \theta_{23} = 0.50$ ,  $\sin^2 \theta_{12} = 0.30$ ,  
 $\sin^2 \theta_{13} = 0.0219$  and  $\delta_{CP} = 0$   
 ★ Phys. Rev. D. 97 072001

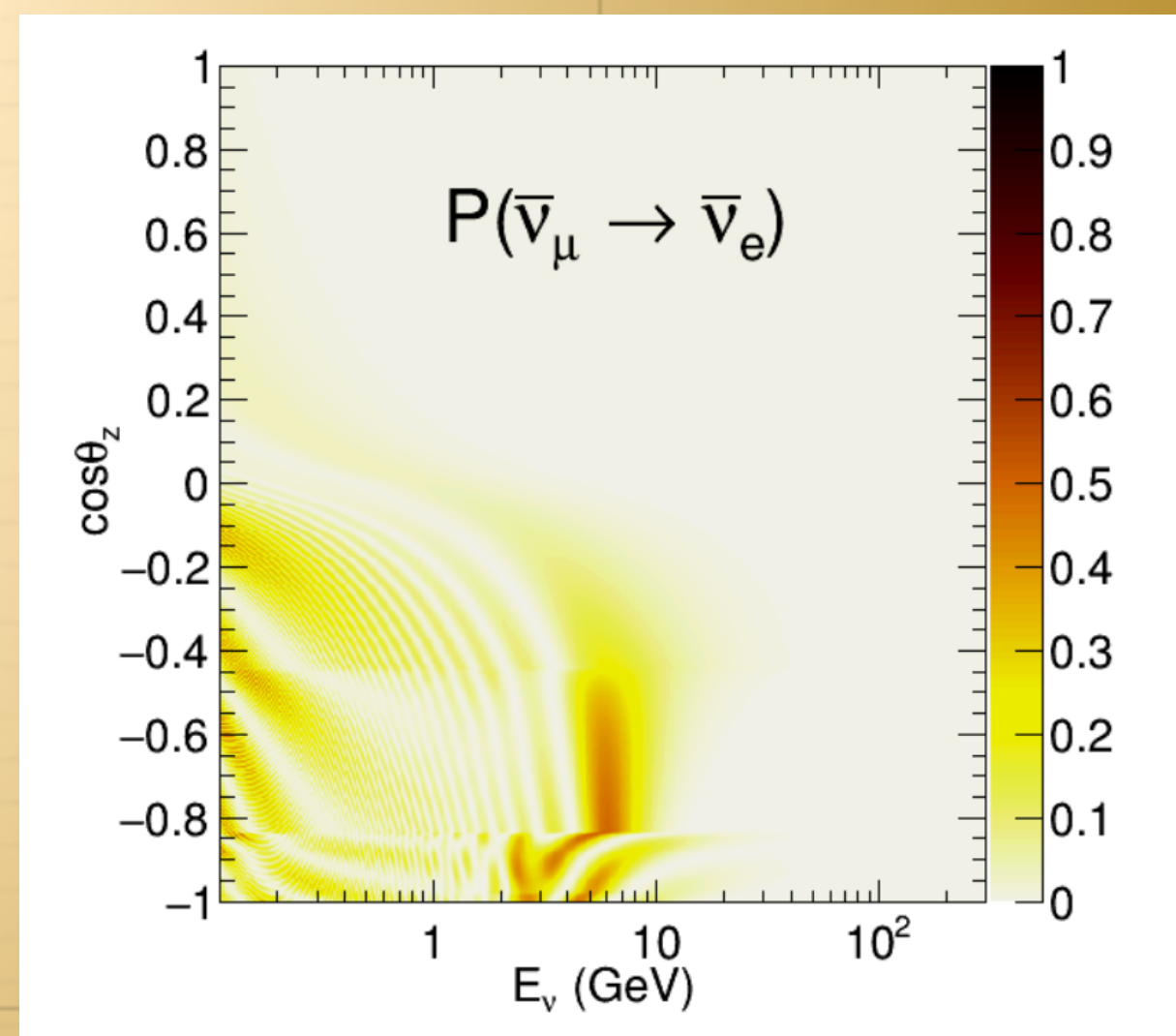
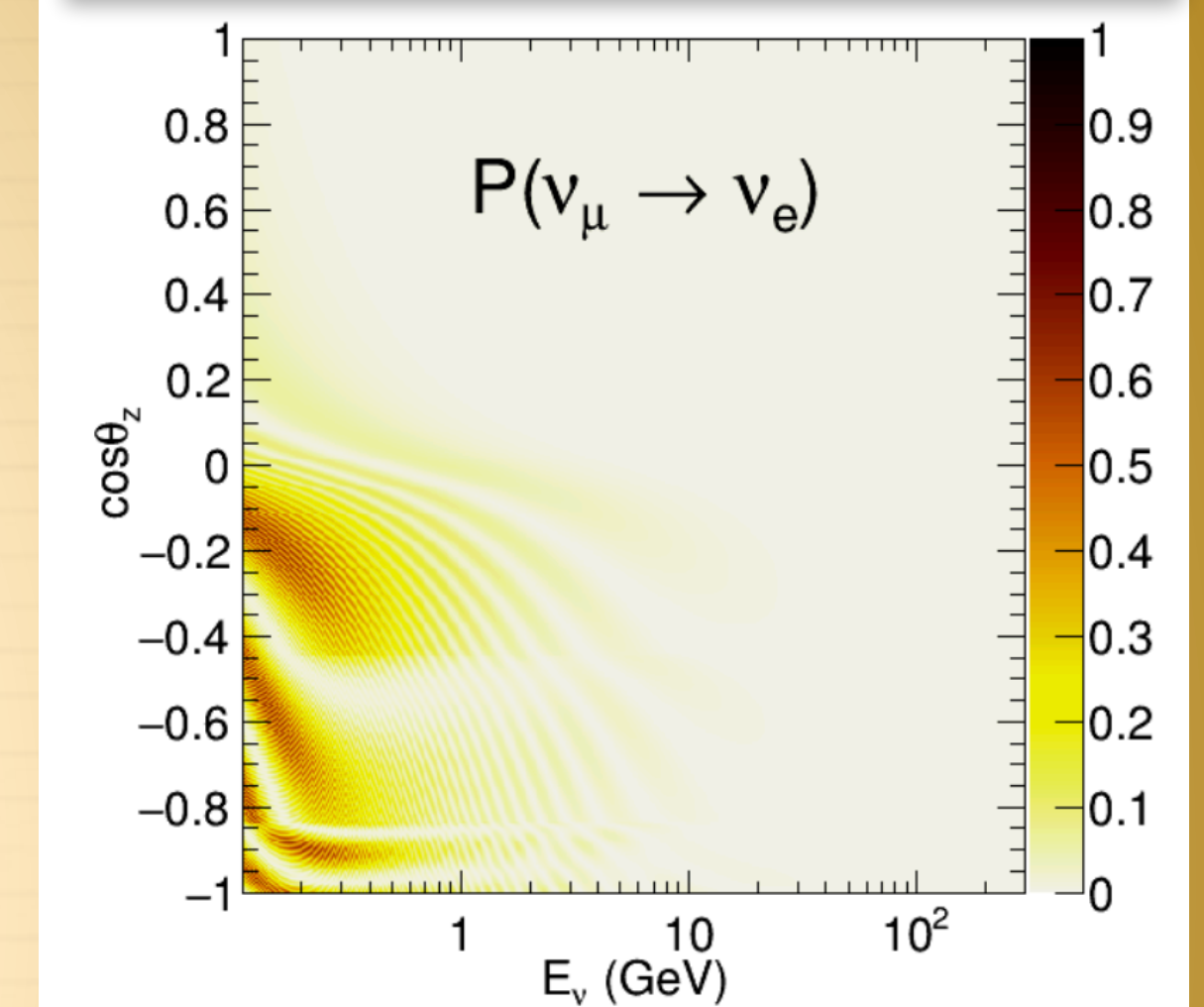
- Impact of matter effects:
  - NO: enhancement of  $\nu_e$  appearance
  - NO: effect is not present for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
  - IO; situation is reversed

• Thanks to presence of matter effects we are sensitive to neutrino mass ordering

Normal Ordering (NO)



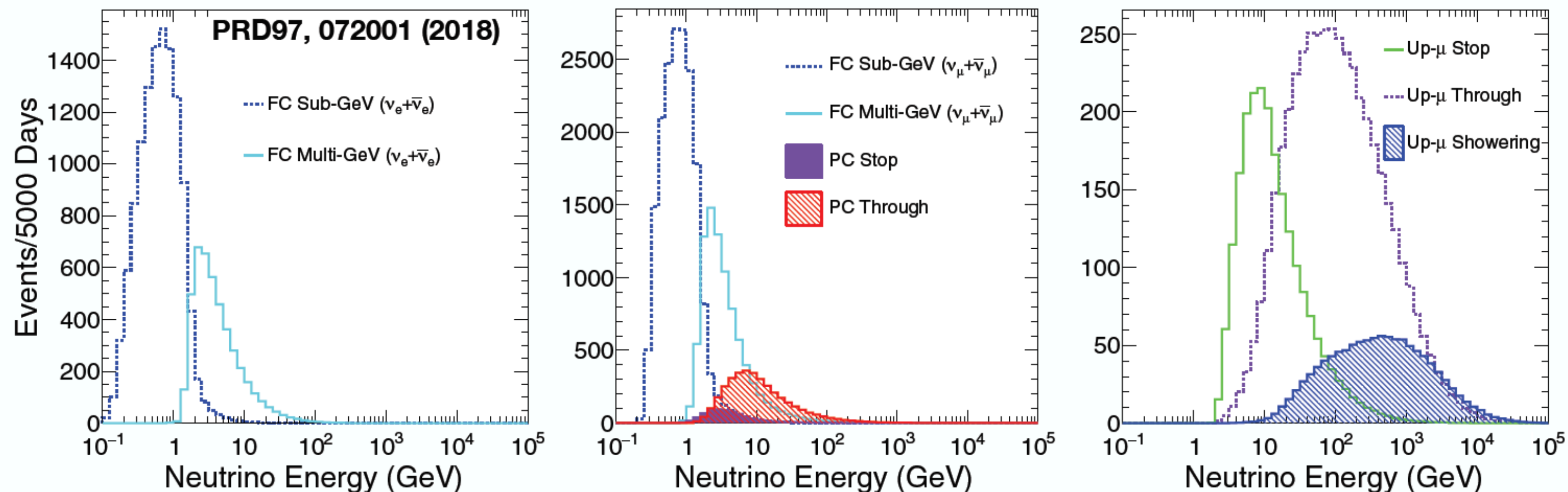
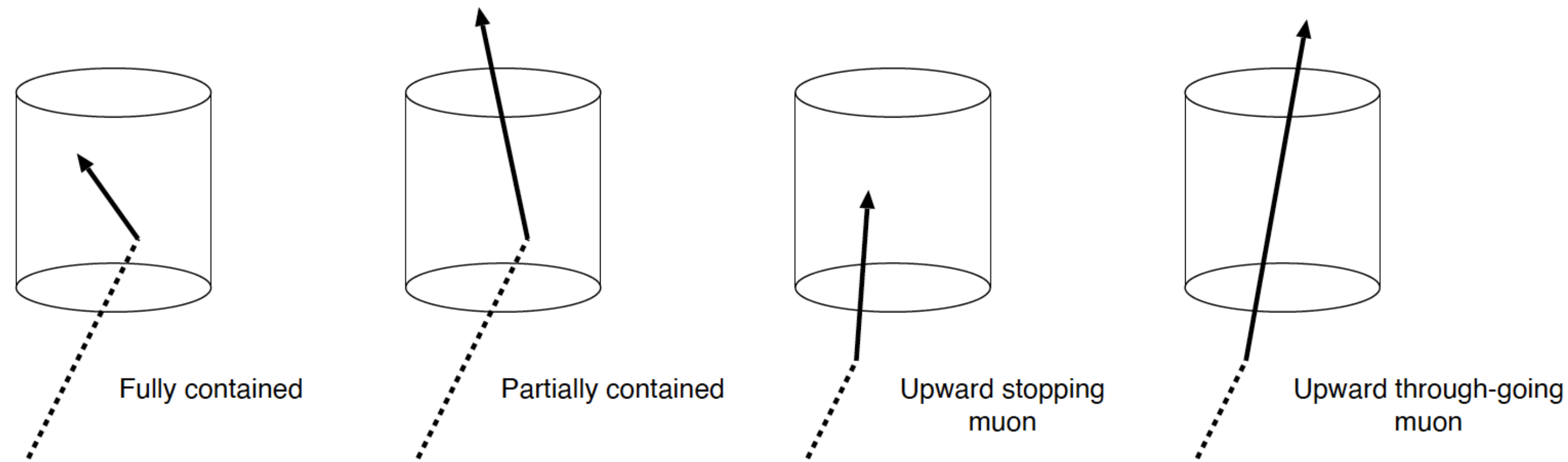
Inverted Ordering (IO)





# Super-K atmospheric neutrino event classification

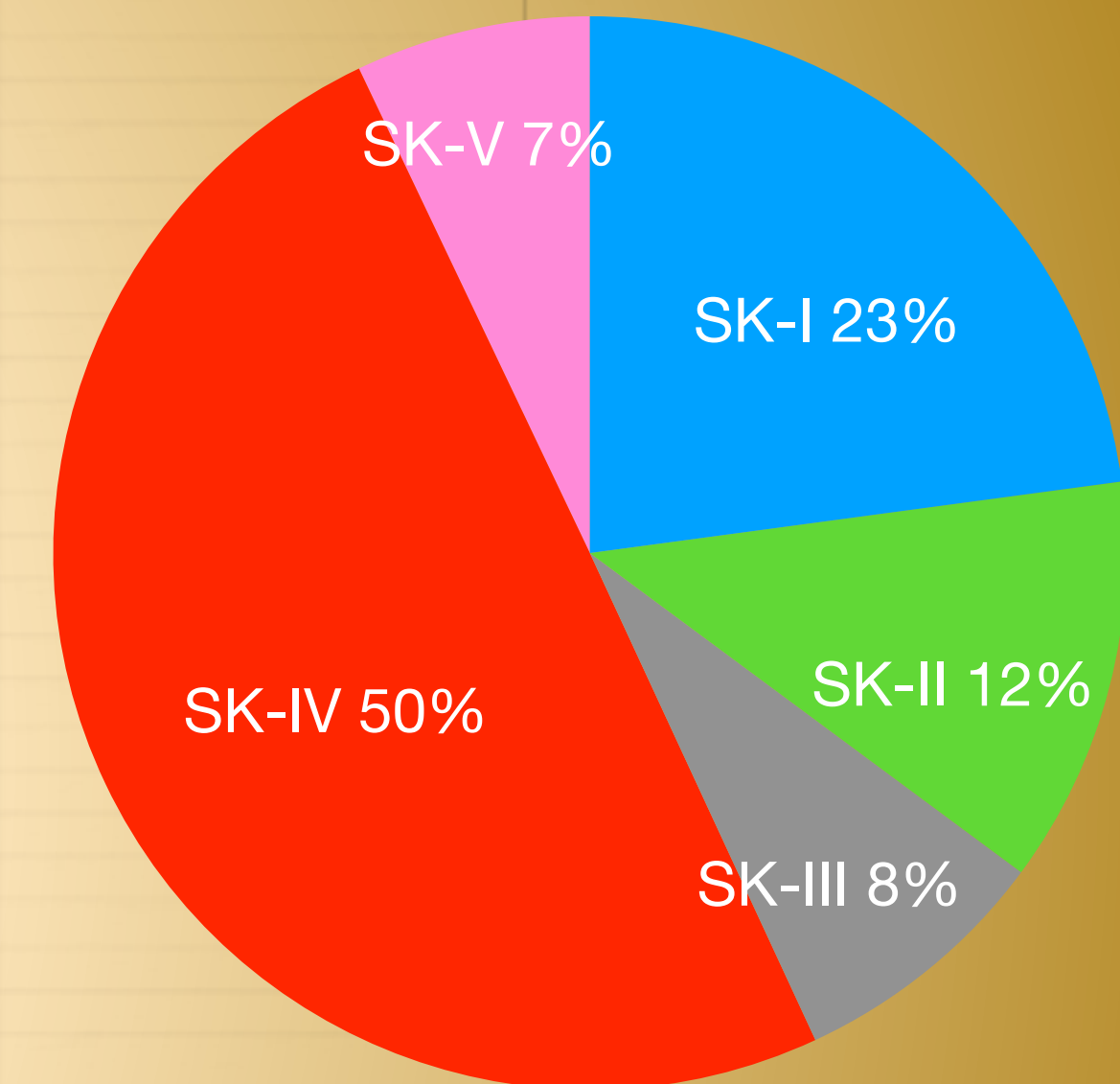
★ Atmospheric neutrino events at Super-K are classified into several categories:





# SK atm- $\nu$ oscillation analysis improvements

- New results with full SK pure water phase (SK-I ~V) with several updates
  - New publication is being written
    - previously published results: PRD97, 072001 (2018): SK 5326 days, 328 kt·yr
- Updates since the previous analysis:
  - Expansion of fiducial volume and more lifetime: 6511 days, 484 kt·yr in total +50% of statistics
  - Event selection with neutron tagging on hydrogen (SK-IV~V)
    - Enhancement of statistical separation between  $\nu$  and  $\bar{\nu}$  events
  - New multi-ring event classification using a Boosted Decision Tree (BDT)
    - Improved charged current/neutral current separation
- Atmospheric  $\nu$  oscillation fit with external constrains
  - $\theta_{13}$  from reactors

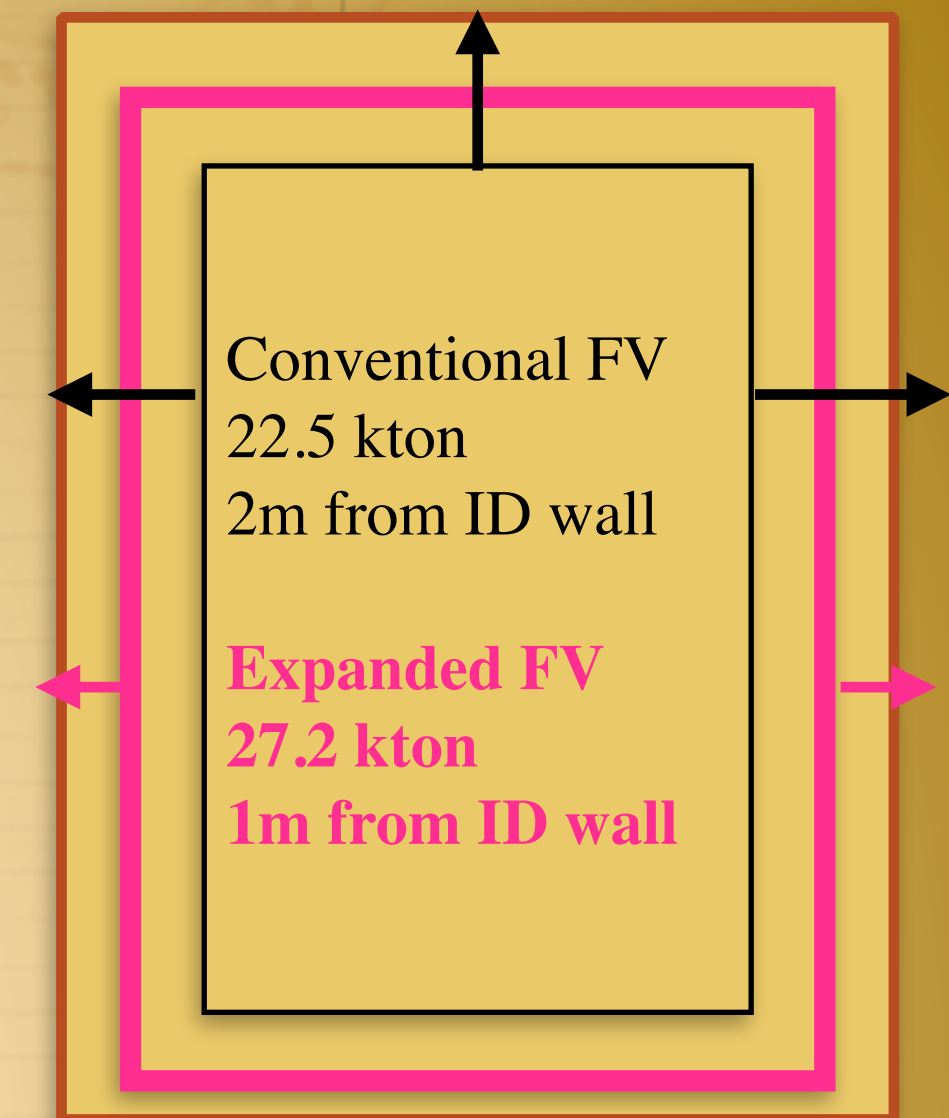




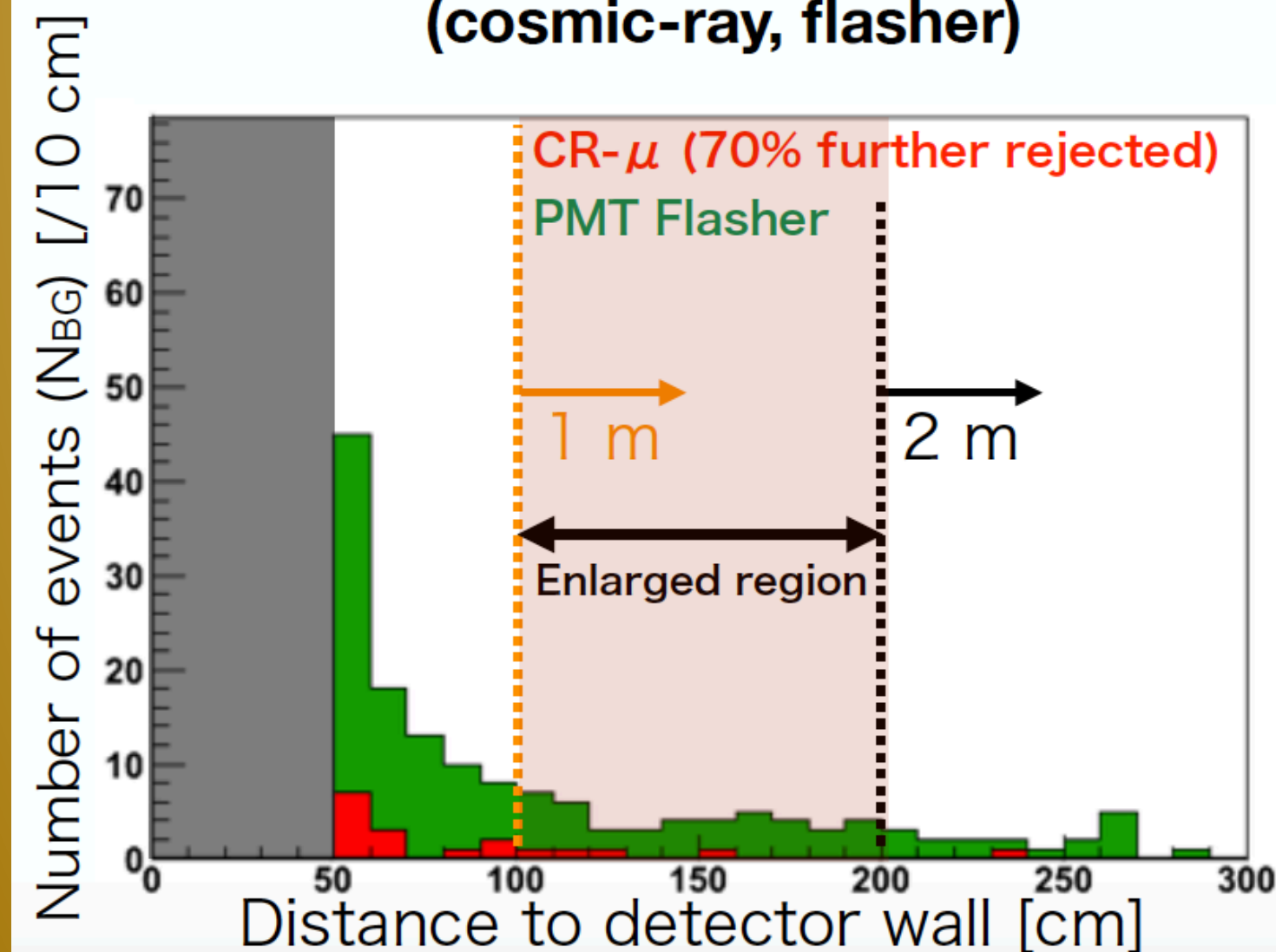
# Enlarging the Fiducial Volume

Distance btw vertex and nearest ID wall surface = "wall"

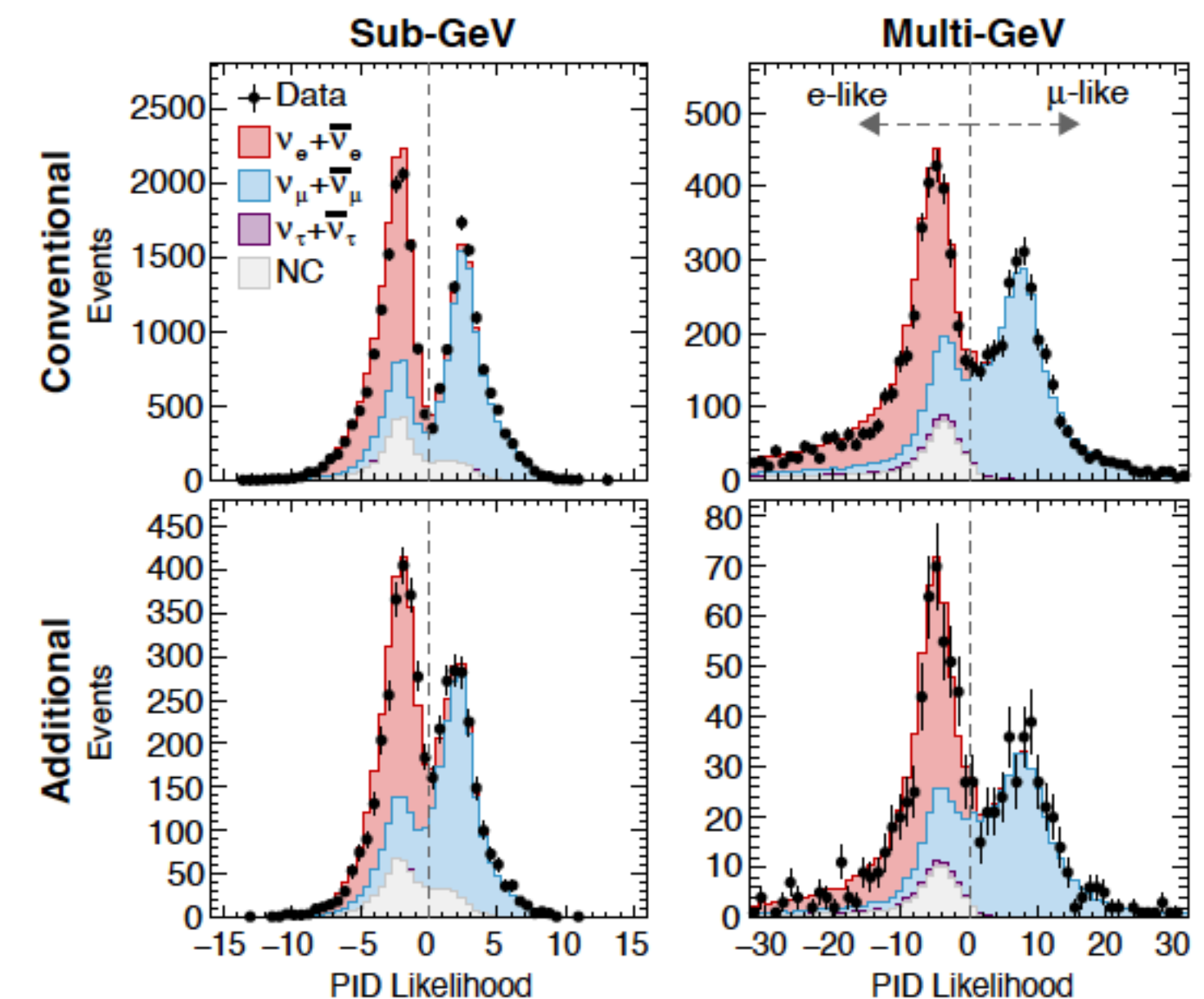
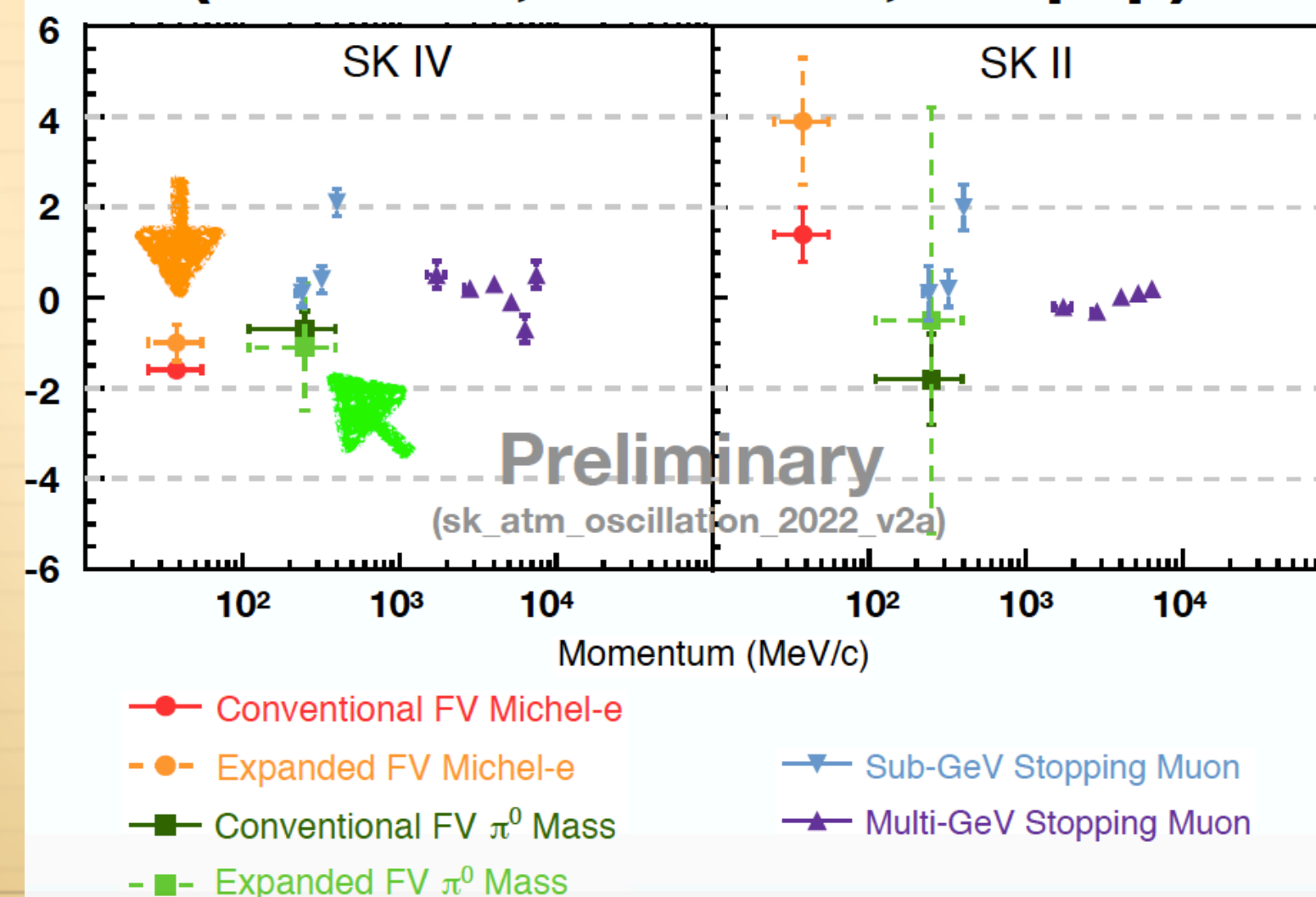
- Conventional fiducial volume defined as wall > 2m
- Expanded fiducial volume to wall > 1m (for all SK periods)
- ★ Increased fiducial volume by 20% (22.5kt → 27.2kt)
- Confirmed no significant increase of non-ν background and no significant bias in reconstruction (ex. energy scale)
- Systematics in the expanded region recalculated and under control



## External background (cosmic-ray, flasher)

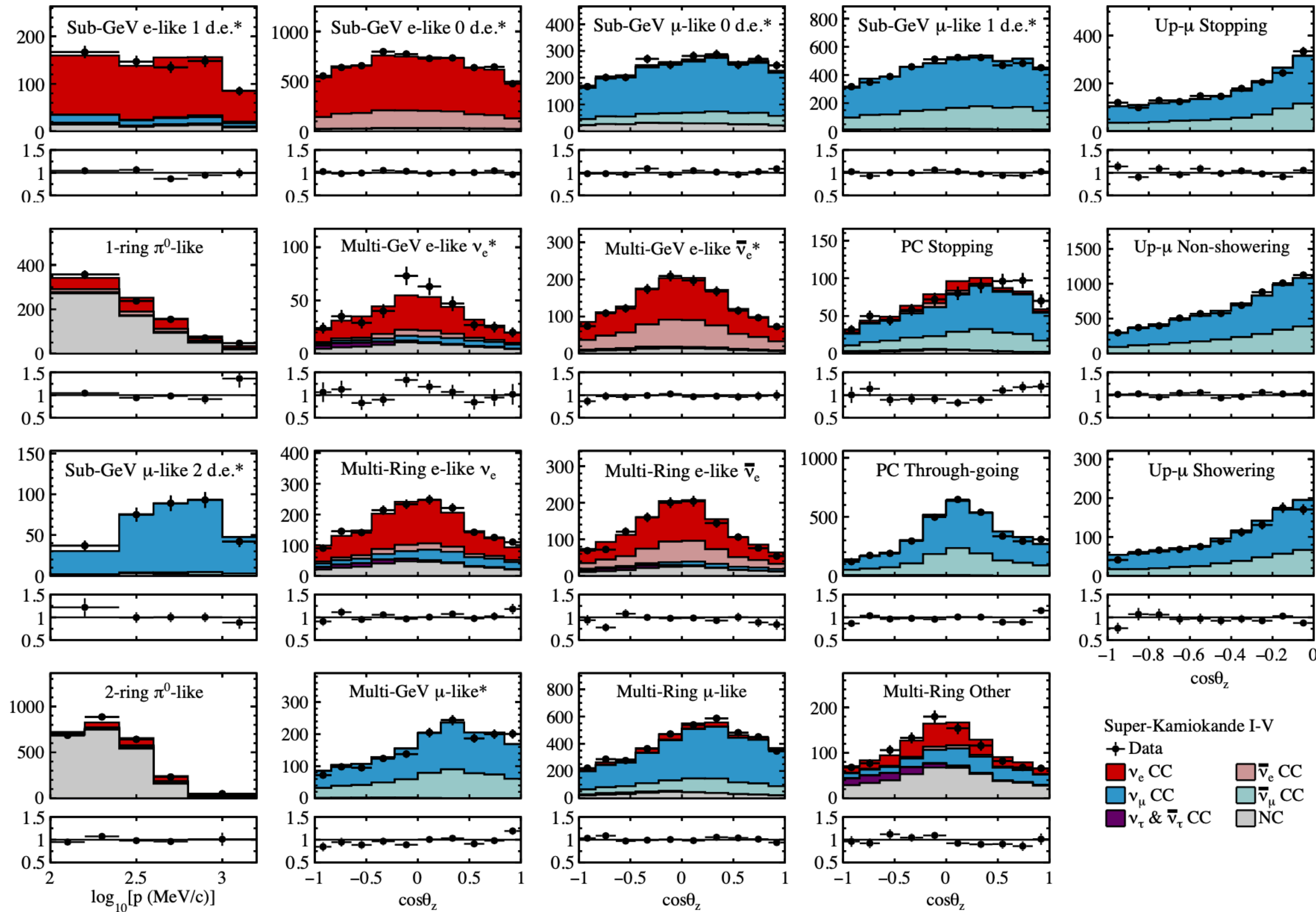


## Energy scale ( $\pi^0$ mass, Michel-e, stop- $\mu$ )





# Super-K samples

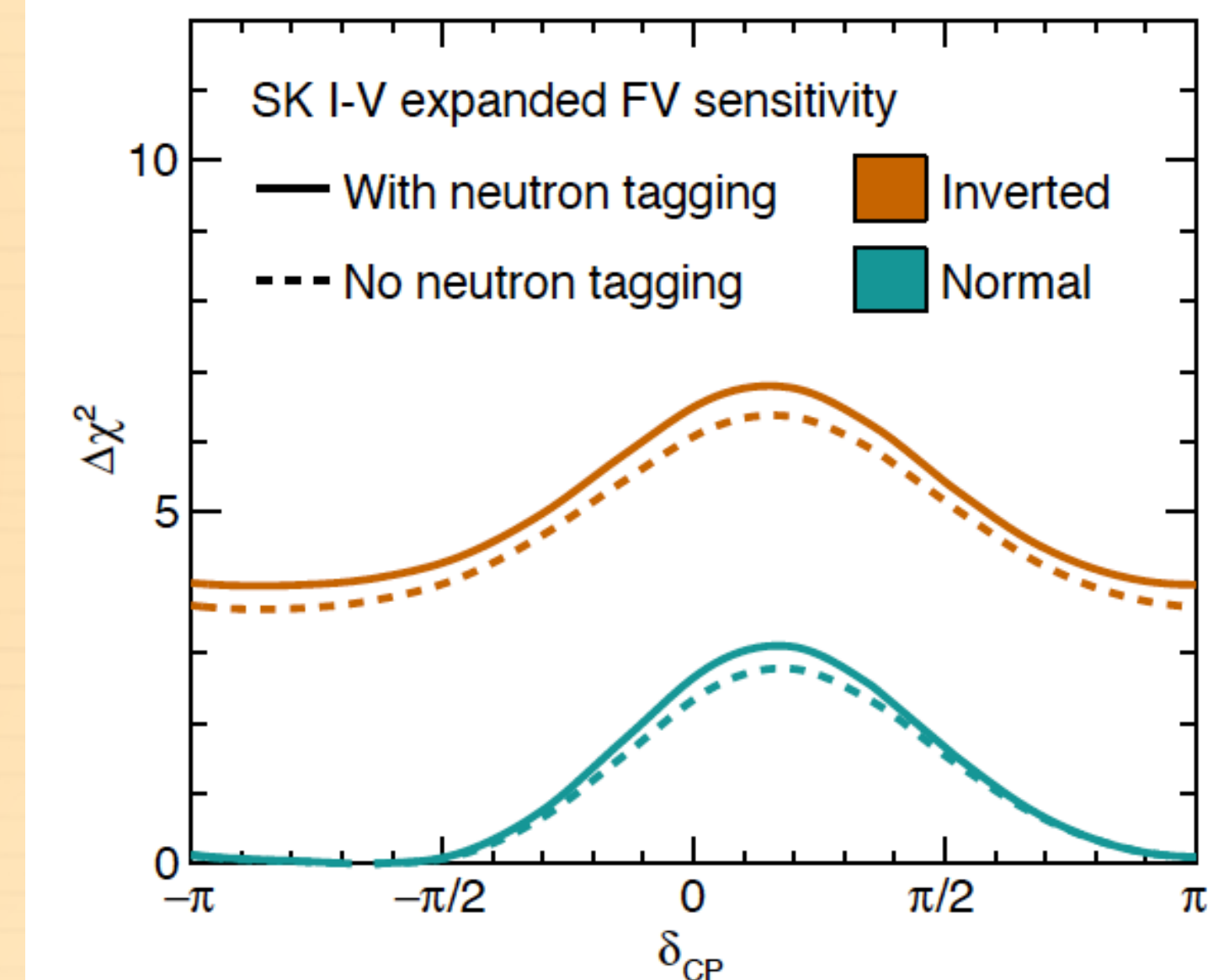
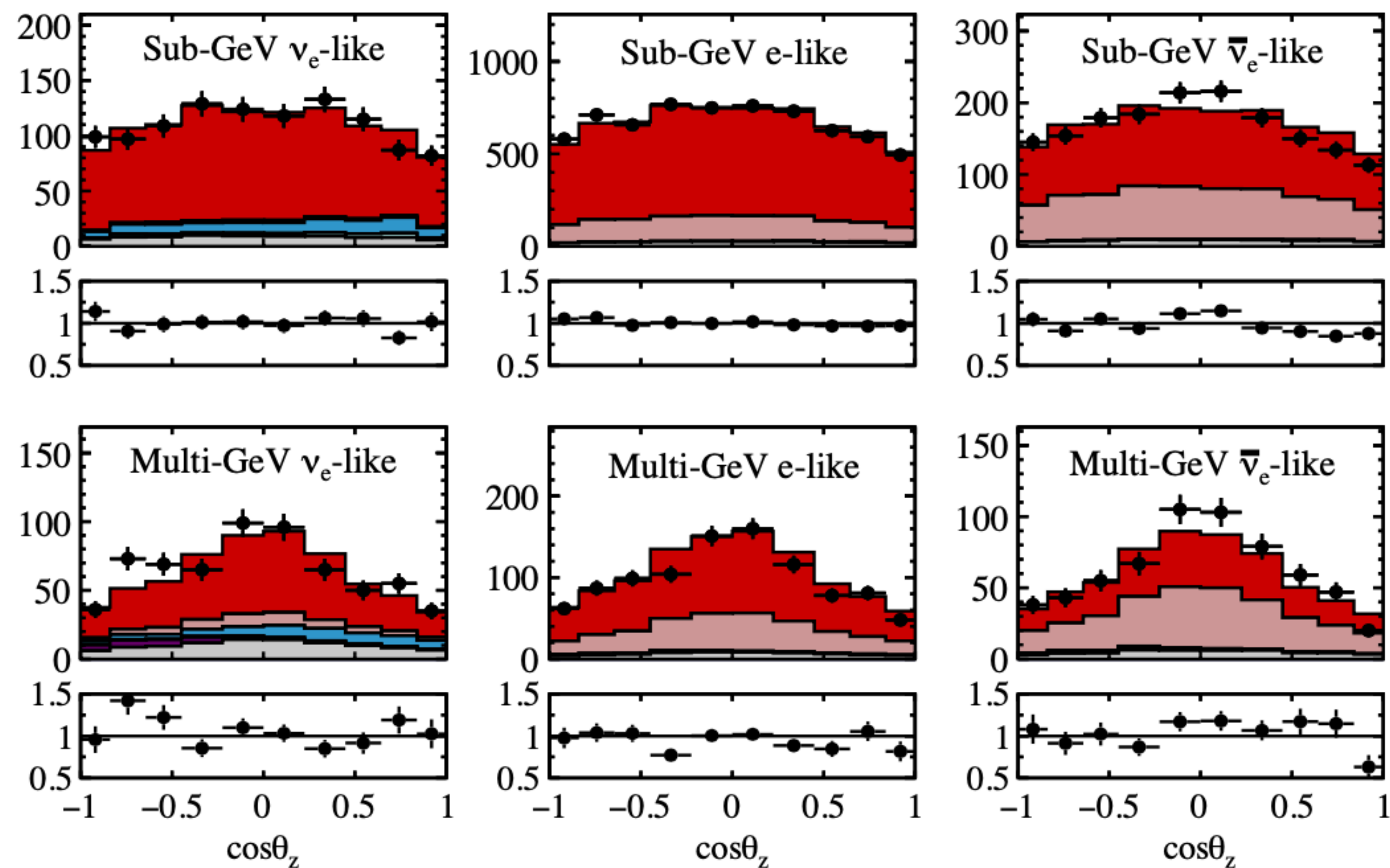


- FC: Sub-GeV and Multi-GeV samples with SK-I~III data, no neutron tagging included\*
- PC, UPMU, FC  $\pi^0$ , FC Multi-Ring samples uses SK-I~V data,



# Super-K samples with neutron tagging

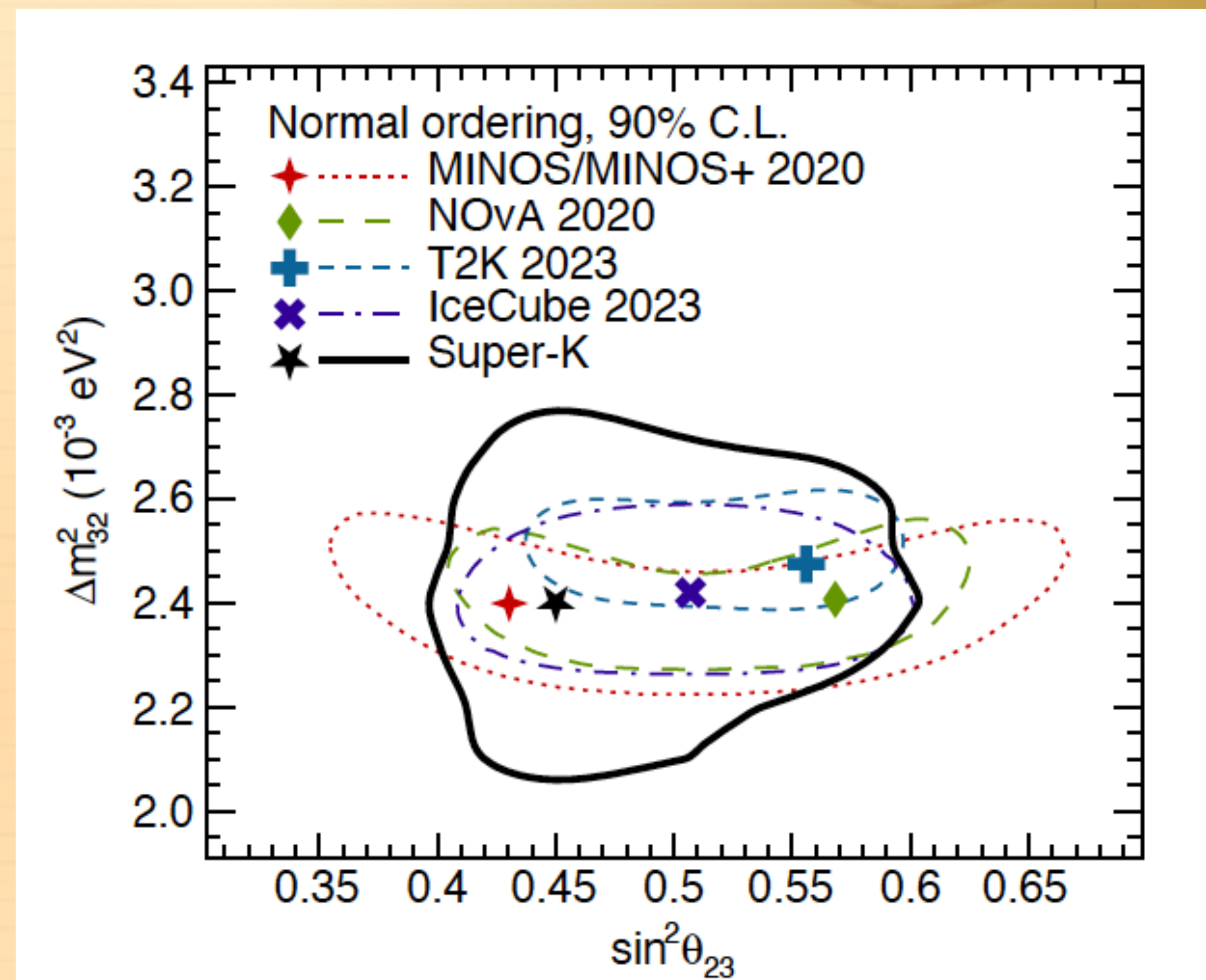
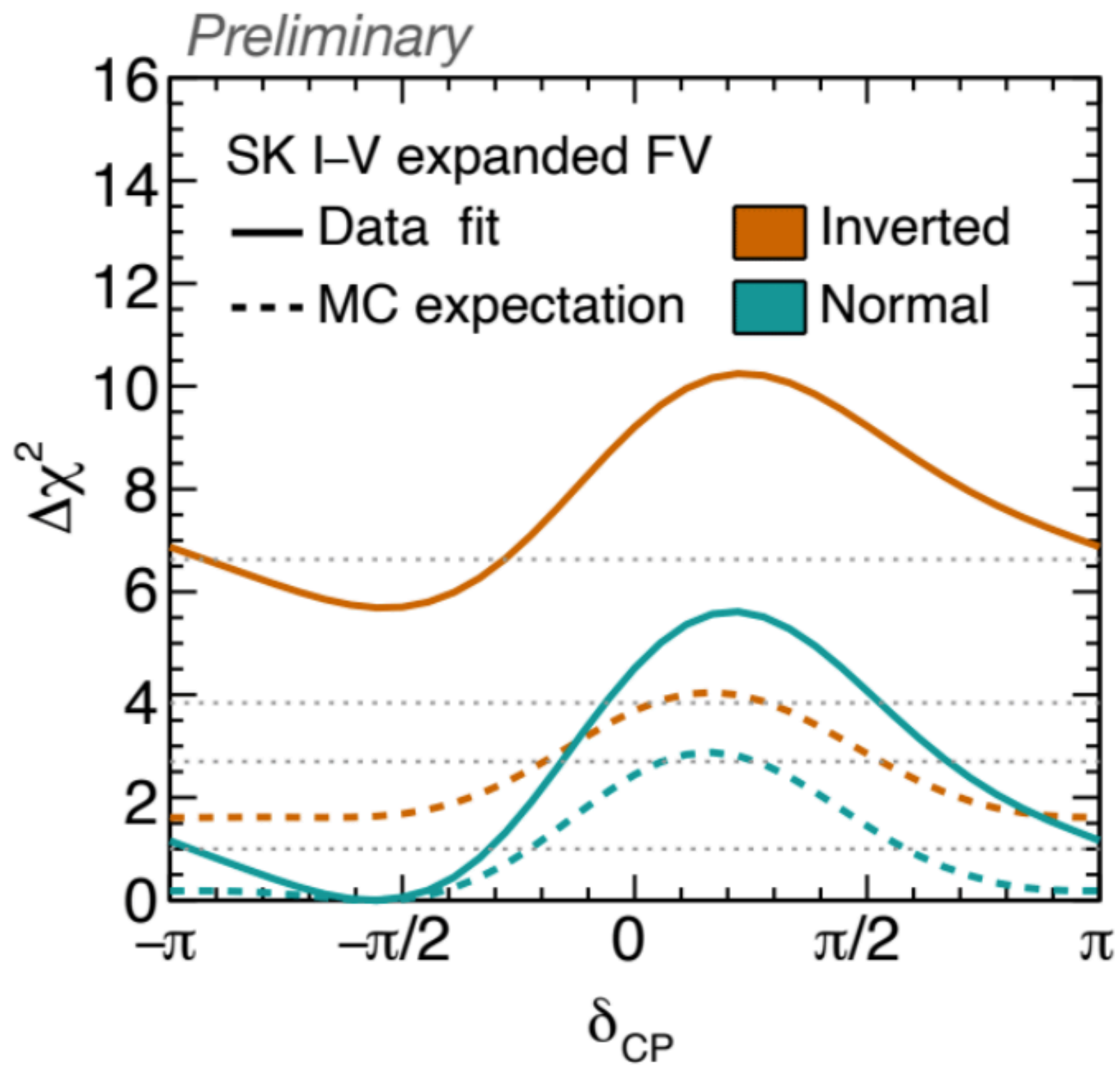
- Additional selections done for SK -IV and SK-V data period, with neutron tagging on Hydrogen.
- Improves separation between  $\nu$  and  $\bar{\nu}$  events





# SK atmospheric $\nu$ results

With  $\sin^2 \theta_{13}$   
constrained  
 $\sin^2 \theta_{13} = 0.0220 \pm 0.0007$   
[PTEP 2022,  
083C01 (2022)]



**SK 2023 best fit results:** Normal ordering,  $\delta_{CP} \simeq -\pi/2$ ,  $\Delta m_{32}^2 \simeq 2.4 \cdot 10^{-3} \text{ eV}^2$ ,  $\sin^2 \theta_{23} \simeq 0.45$

Mass ordering:  $\Delta\chi_{I.O-N.O}^2 \simeq 5.7$

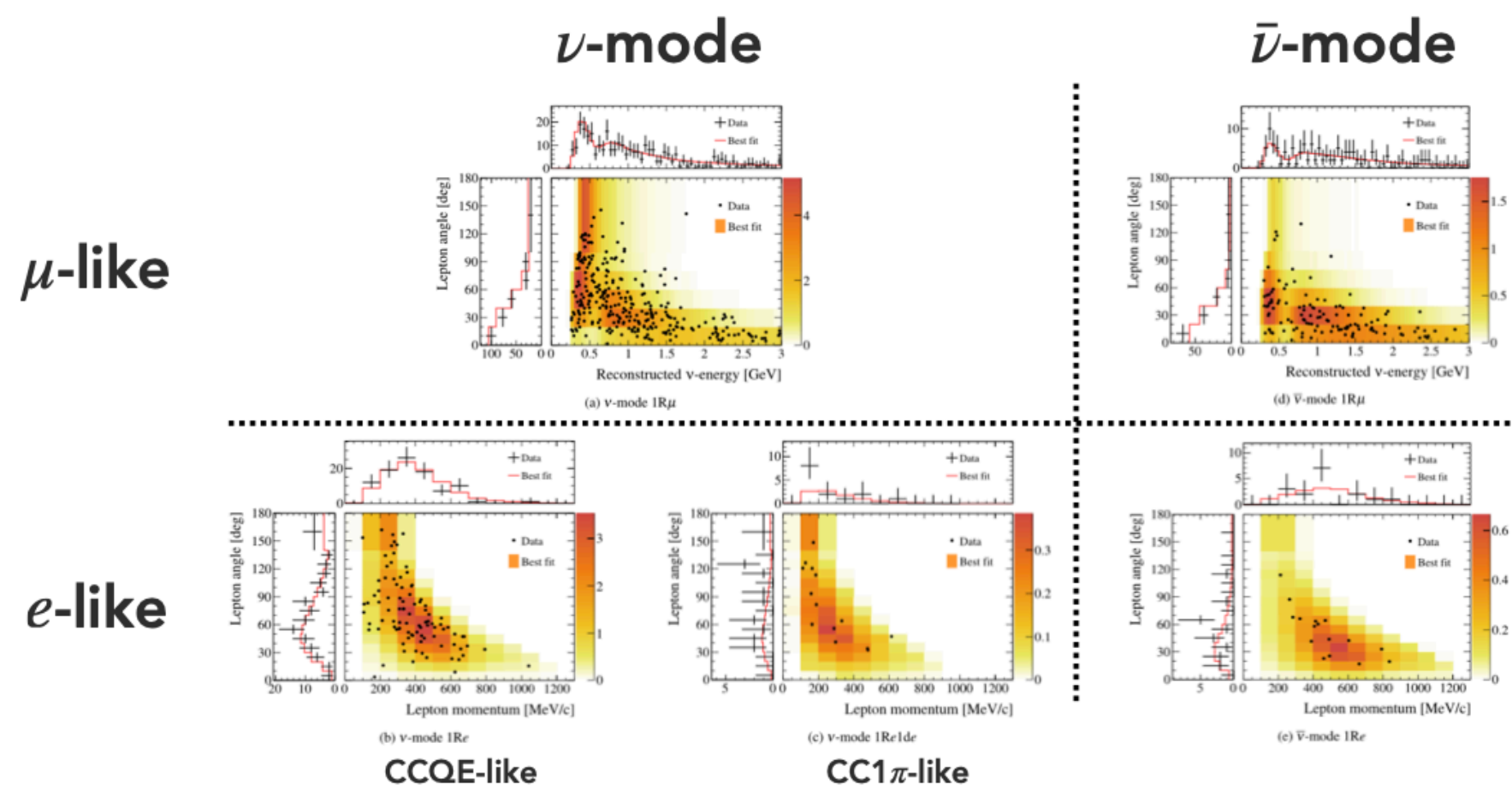
This analysis prefers NO over the IO at the 92.3% confidence level.



- Motivation of the Joint Analysis:
  - T2K and SK use the same detector and have samples with similar energy ranges and similar selections.
  - We can take into account **the correlations of the systematic uncertainties**
  - T2K near detector can be used to constrain **the cross-section uncertainties** for the low -energy atmospheric samples as well
  - T2K has better sensitivity to  $\delta_{CP}$  from  $\nu_\mu \rightarrow \nu_e$  oscillation and to  $\theta_{23}, \Delta m_{32}^2$  from  $\nu_\mu \rightarrow \nu_\mu$  channel (this is true for antineutrinos as well)
  - SK has **stronger discrimination of the mass ordering** thanks to the **matter effect** at the few GeV

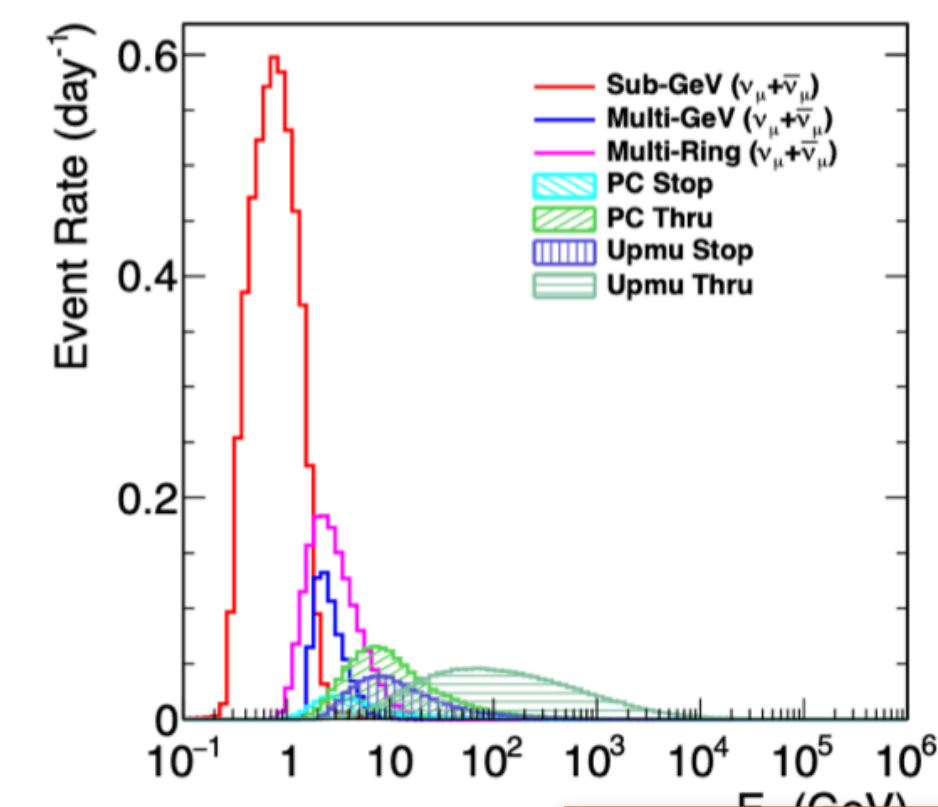
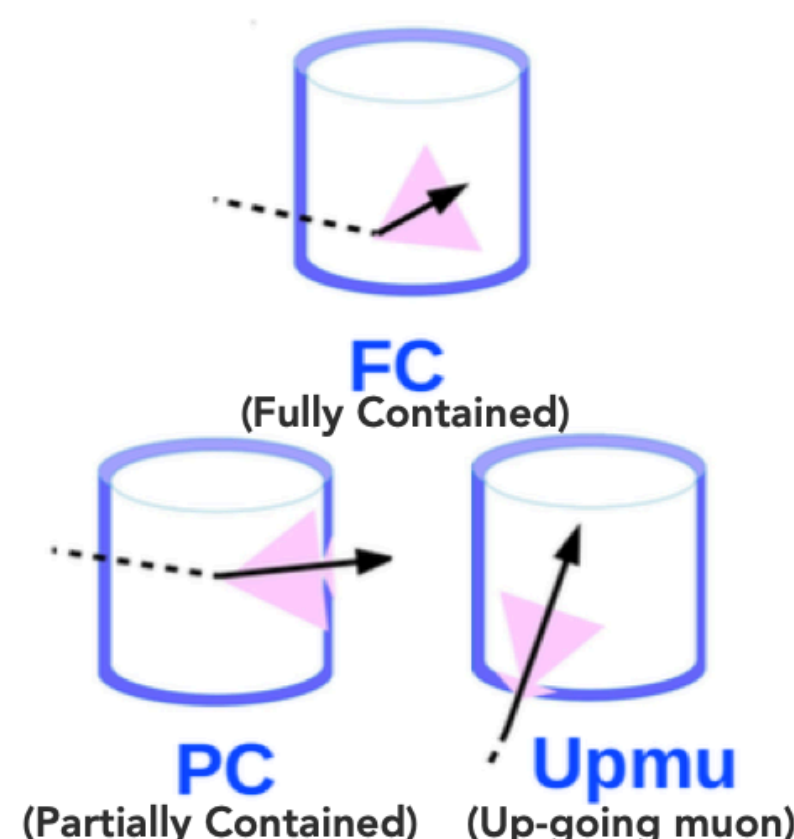
## T2K Run 1-10 (not the latest analysis) [[Eur.Phys.J.C 83 \(2023\) 9, 782](#)]

- Neutrino mode:  $19.7 \times 10^{20}$  POT
- Antineutrino mode:  $16.3 \times 10^{20}$  POT
- Mean neutrino energy  $\sim 0.6$  GeV



## SK IV atmospheric neutrinos [[PTEP 2019 \(2019\) 5, 053F01](#)]

- 3244.4 days of data taking
- 18 samples depending on the event topologies and neutrino energies
- **Wider energy ranges** than T2K

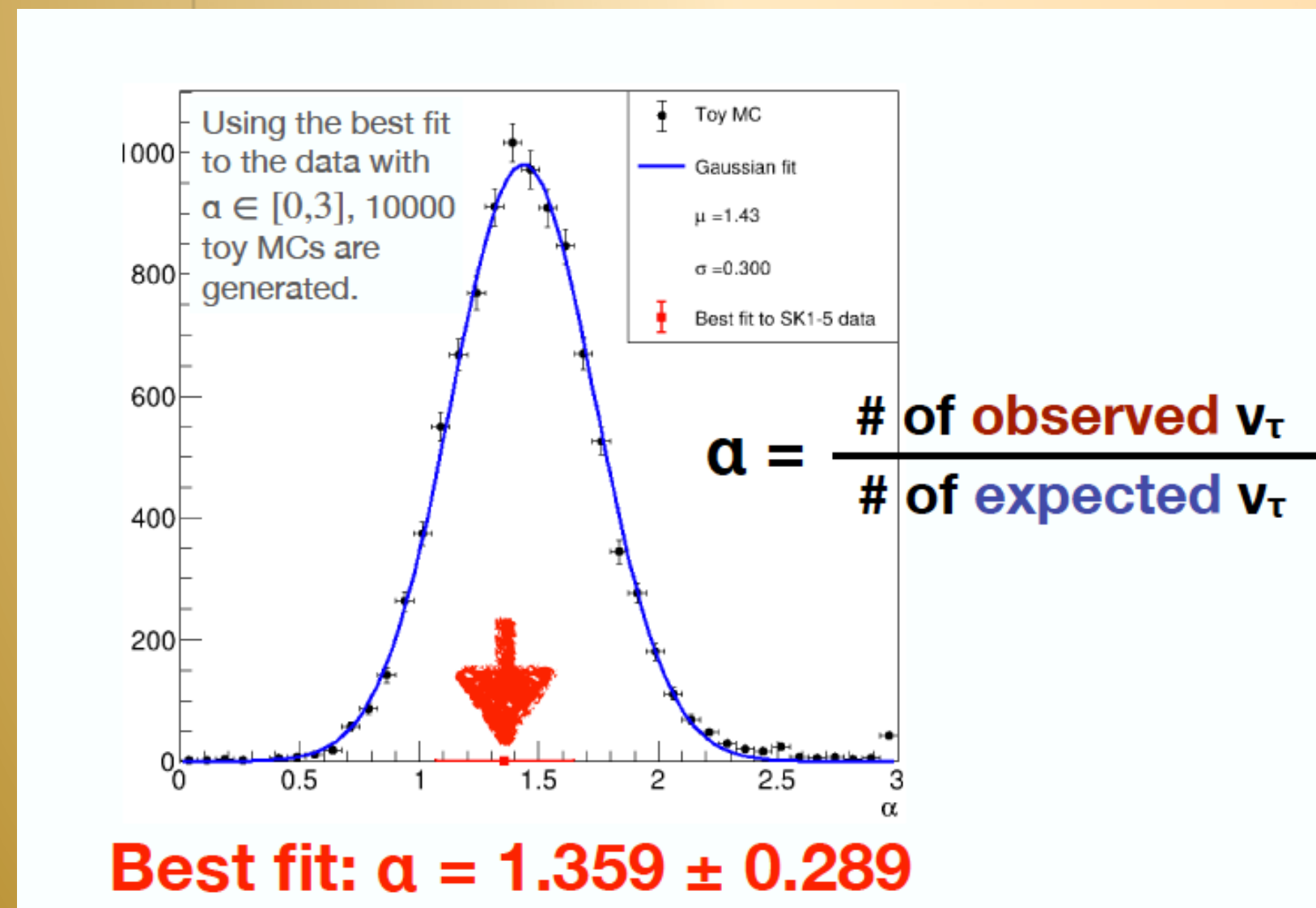


See the talks on this conference from Adrien Blanchet and Thomas Holvey

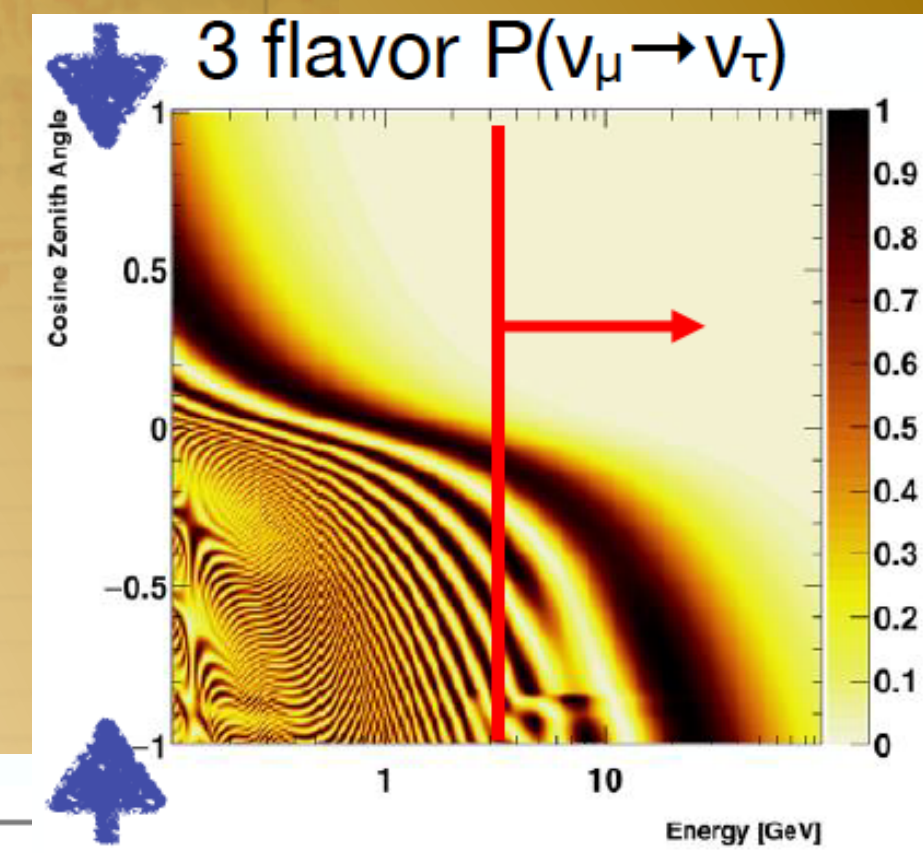
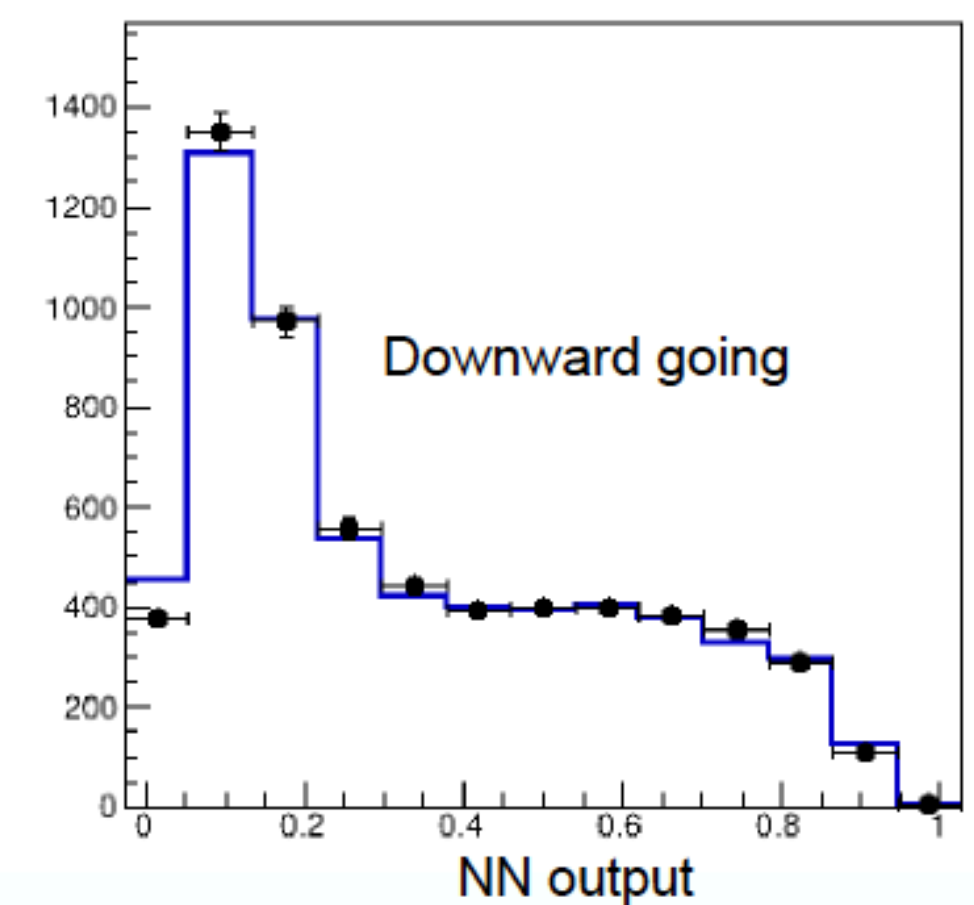
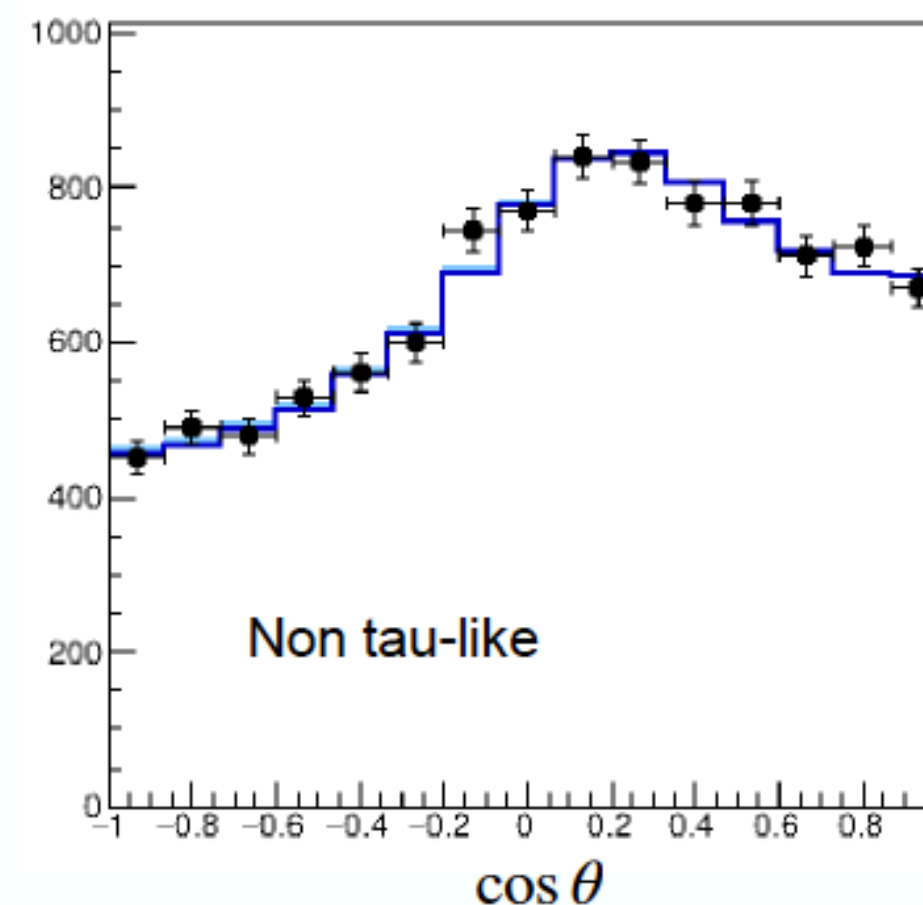
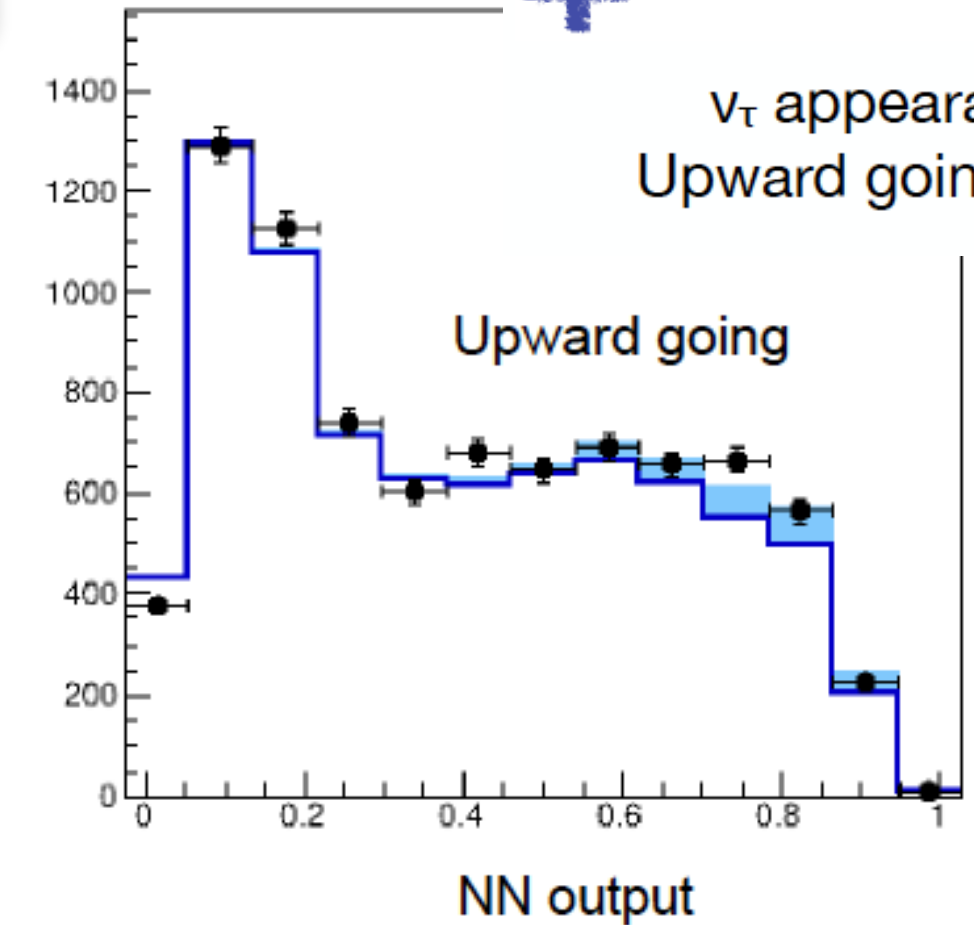
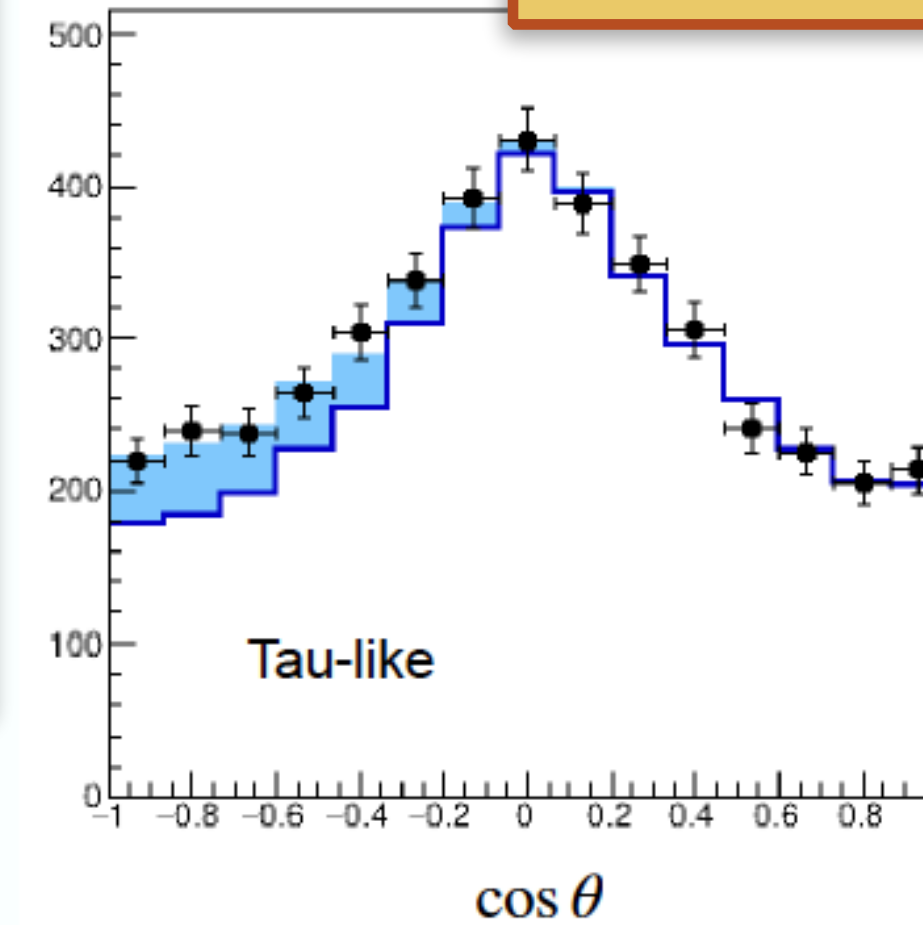


# $\nu_\tau$ appearance searches

- Updates since the last publication in 2018 PRD98 052006 (2018)
  - Full SK pure water phases (SK-I~V data)
    - Additional 2 years of SK-IV and SK-V data added
  - Expanded fiducial volume - overall 50% more data added
- Best fit of  $\nu_\tau$  normalisation parameter:  $\alpha = 1.359 \pm 0.289$
- Excluding no  $\nu_\tau$  appearance ( $\alpha = 0$ ) at  $4.8\sigma$  significance, p-value:  $7.5 \cdot 10^{-7}$
- Observed # of  $\nu_\tau$  CC events:  $428 \pm 92$  (normal MO)



Preliminary

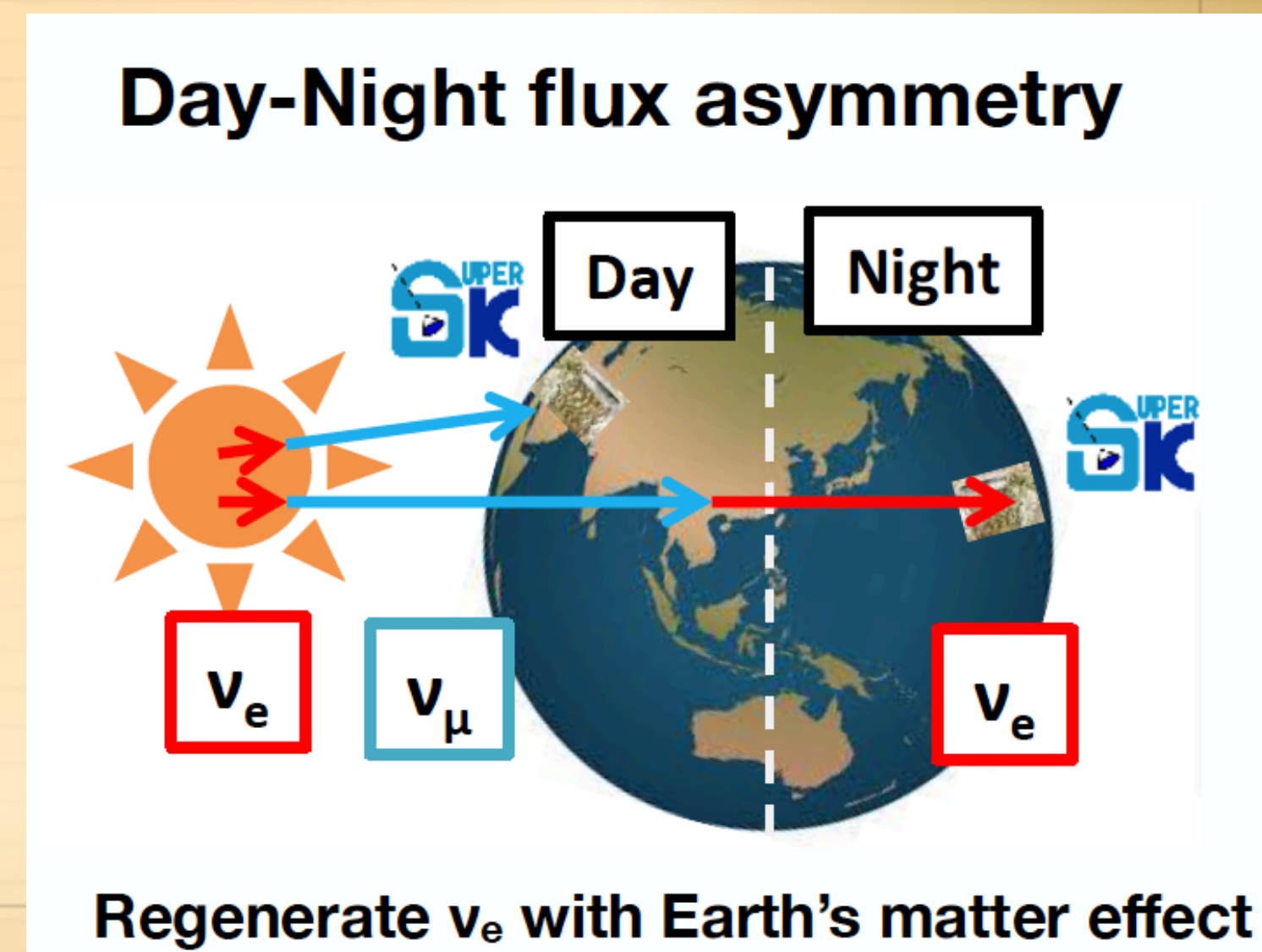
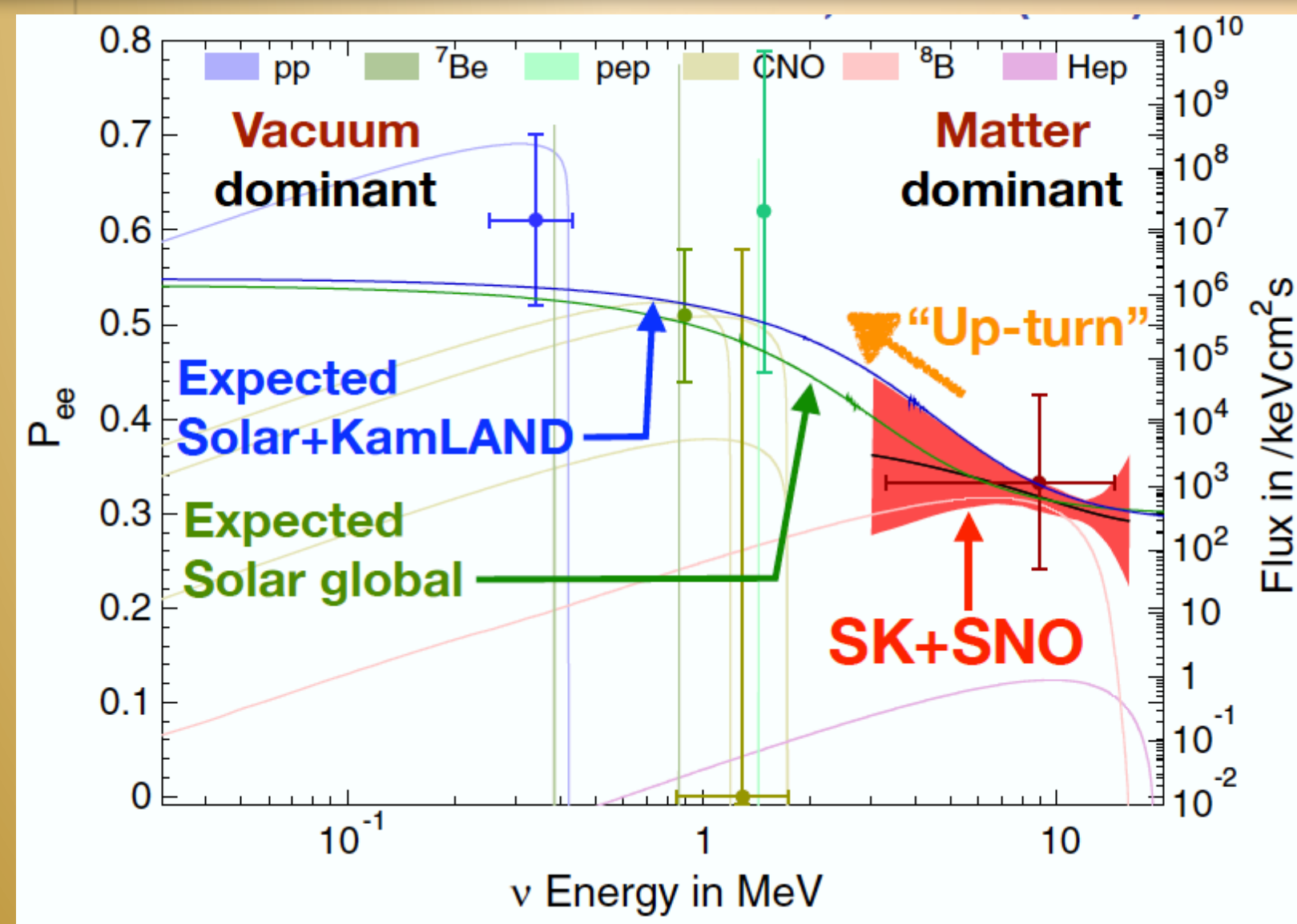


$\nu_\tau$  appearance signal:  
Upward going,  $E_\nu \geq 3.5\text{GeV}$



# Solar $\nu$ searches at SK

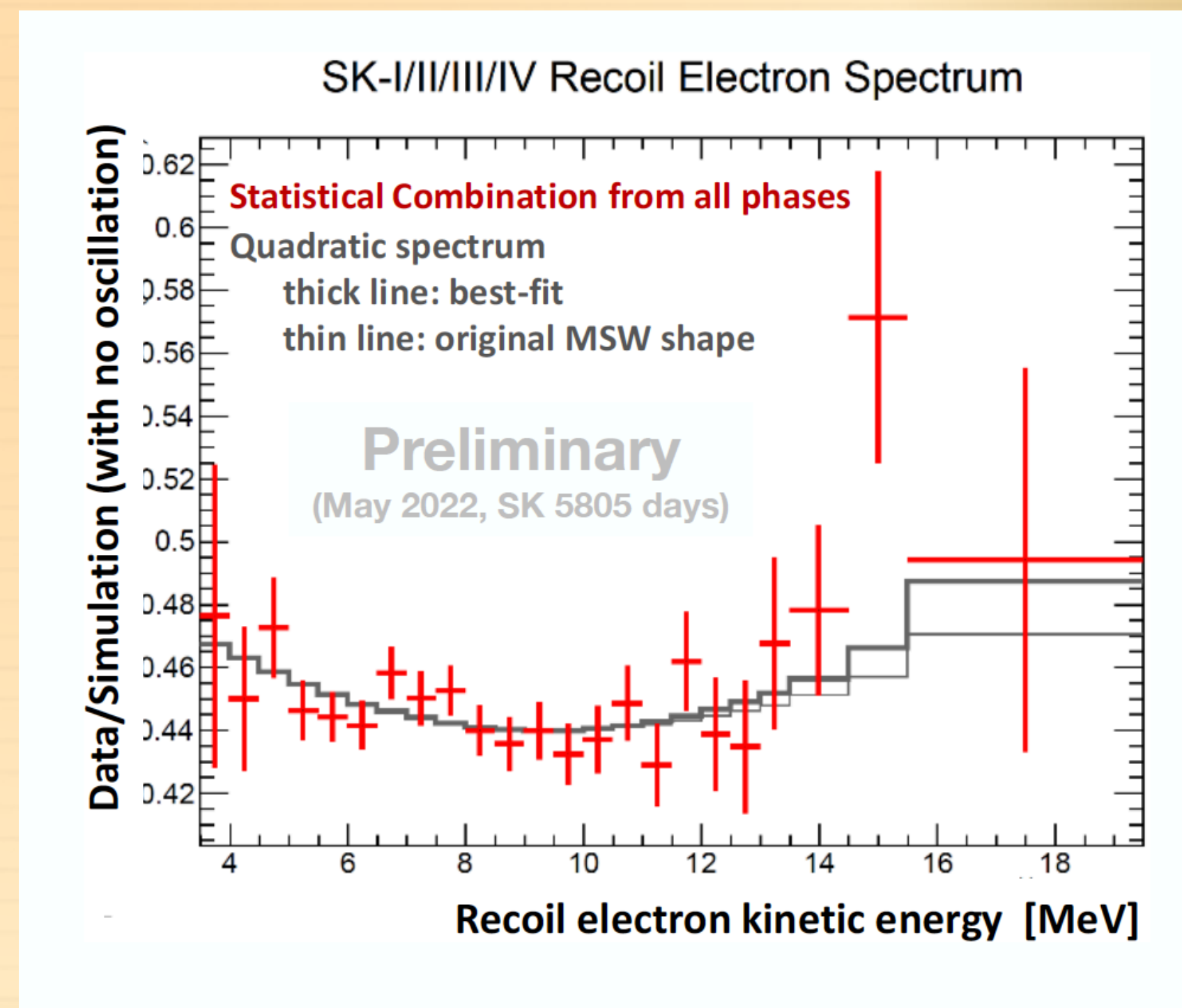
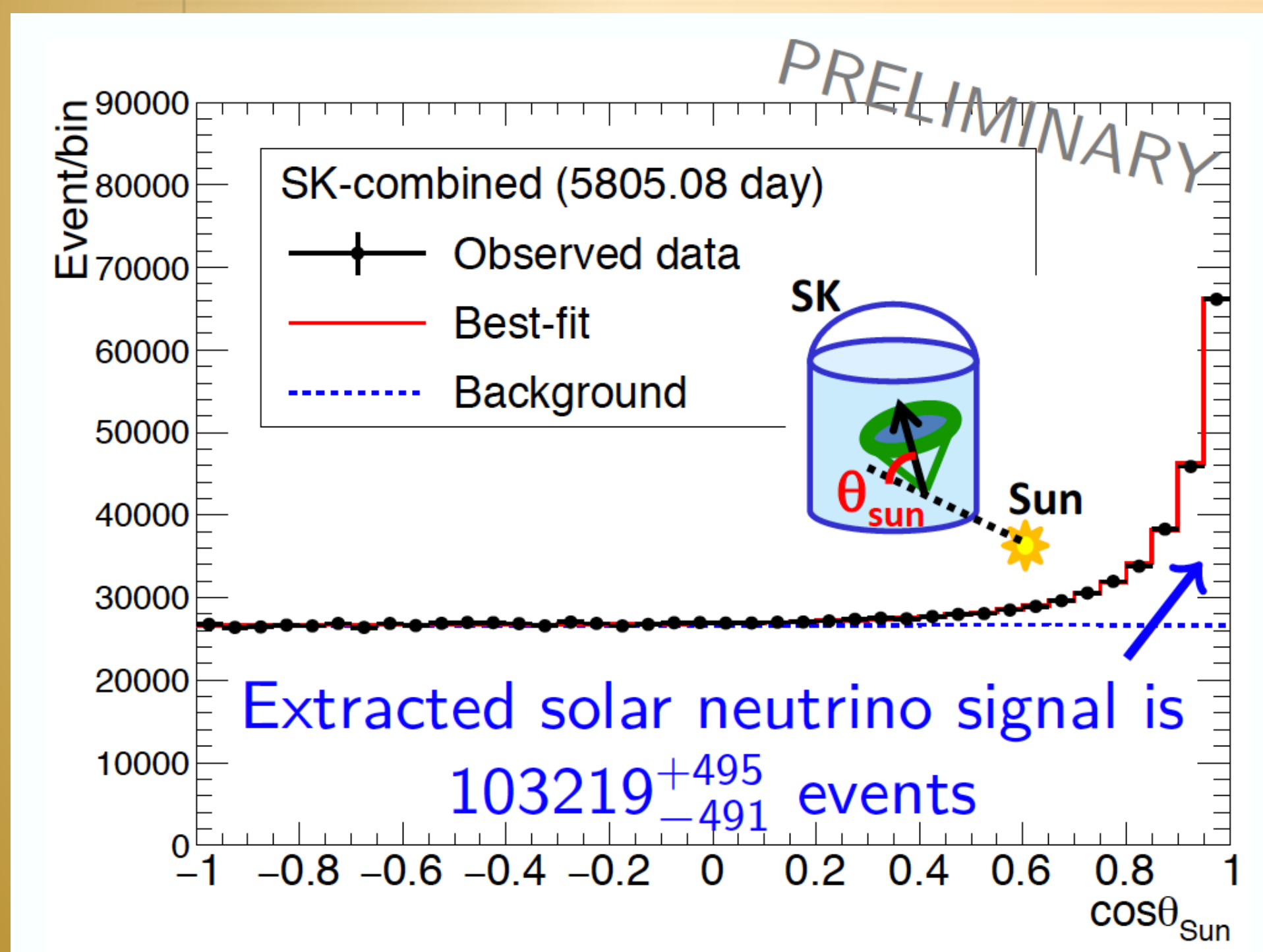
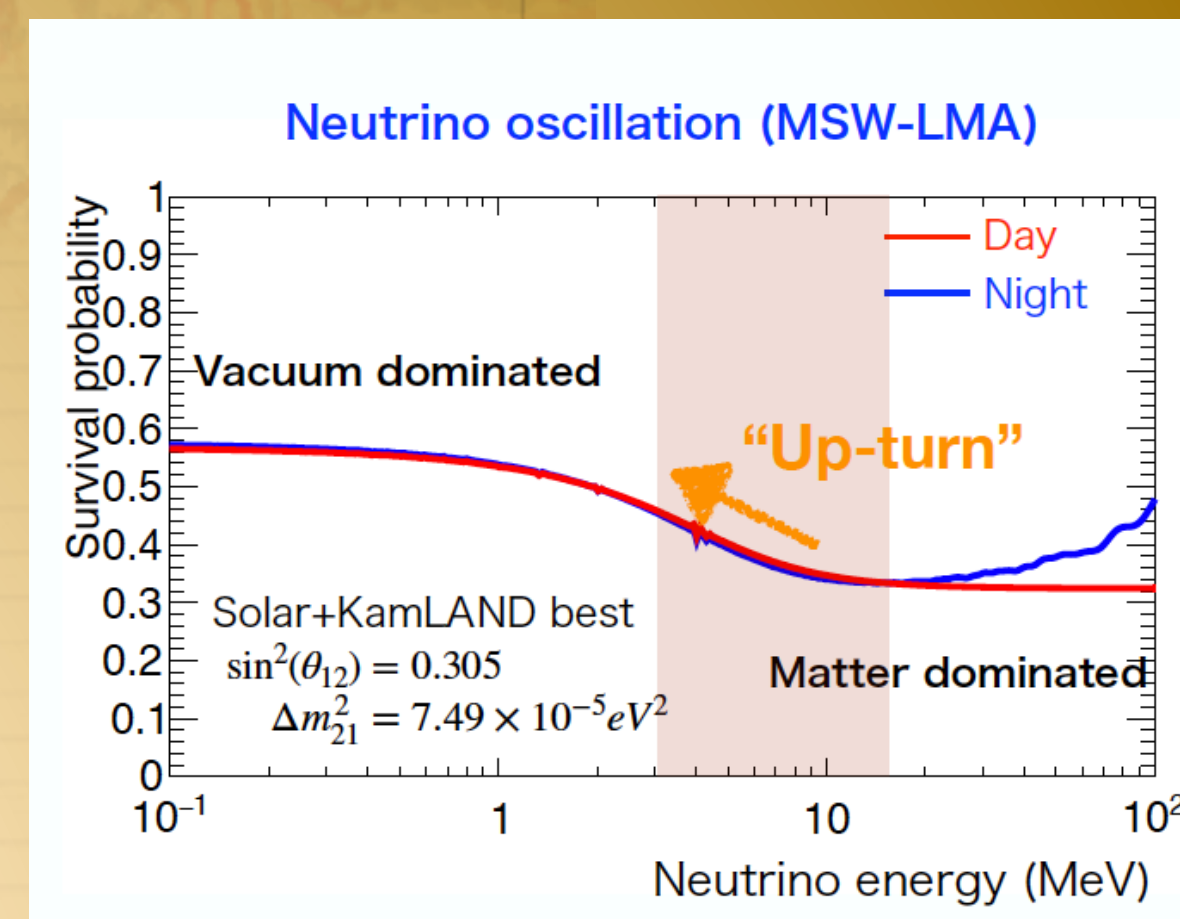
- Intense neutrinos coming from nuclear fusion in the Sun:
  - $^8\text{B}$  neutrinos recorded by SK with direction and energy information
  - Sensitive to  $\theta_{12}$  and  $\Delta m_{21}^2$
- Precision test of the MSW oscillation model
  - Precise measurement of spectrum at the vacuum-to-matter transition “up-turn”
    - Sensitive to matter effects in the Sun
  - Measurement of Day/Night flux asymmetry
    - Sensitive to matter effects in the Earth





# Solar $\nu$ searches at SK

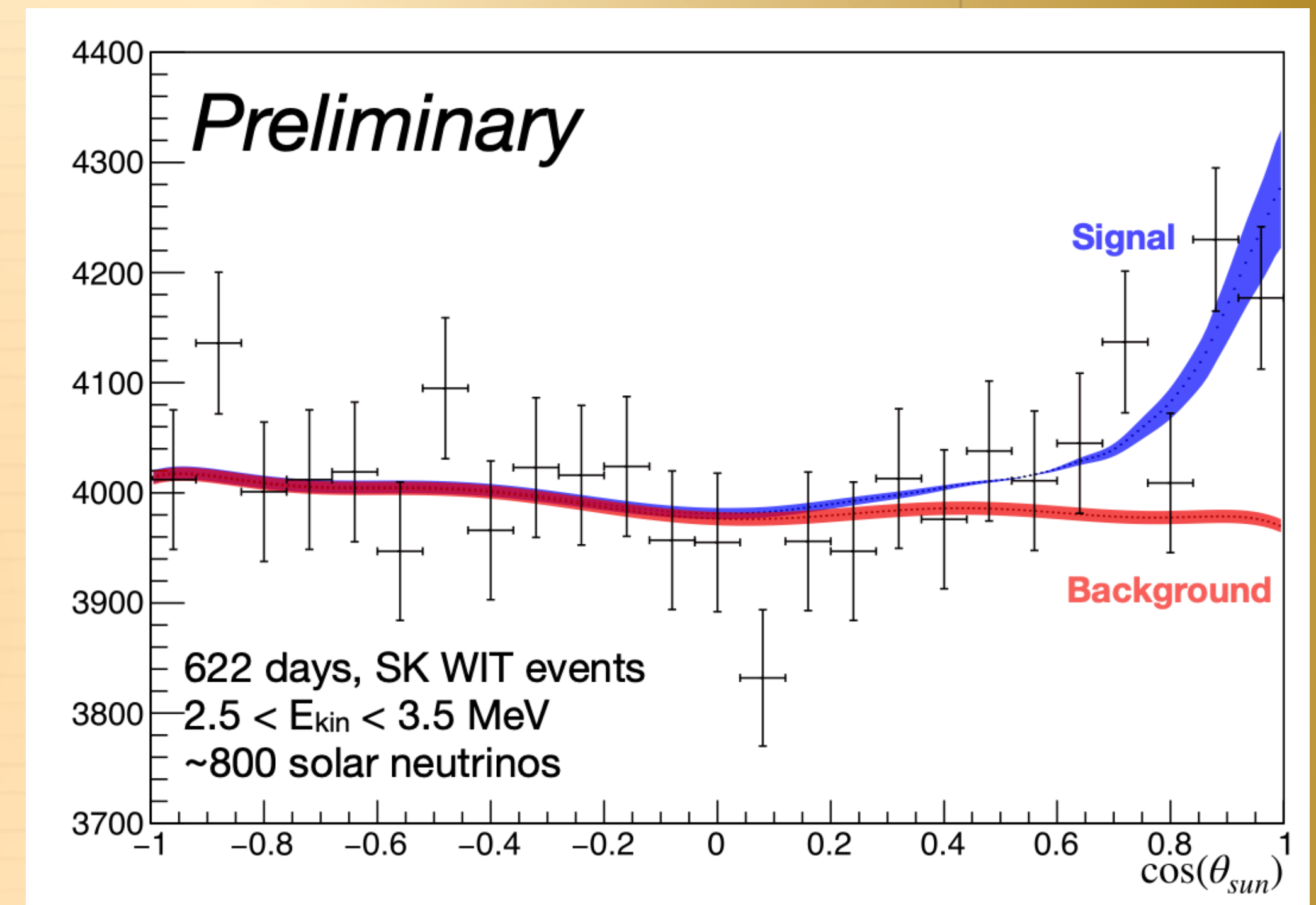
- SK measurement of solar neutrinos
  - Detecting recoil electron from elastic scattering
  - Robust signal extraction using angular correlation with the Sun
- Total # of observed solar  $\nu$  events  $>10^5$  in phases SK-I~IV (5805 days)
  - Energy spectrum slightly favours "up-turn", more data needed





# Solar $\nu$ searches at SK - lowering neutrino energy thresholds

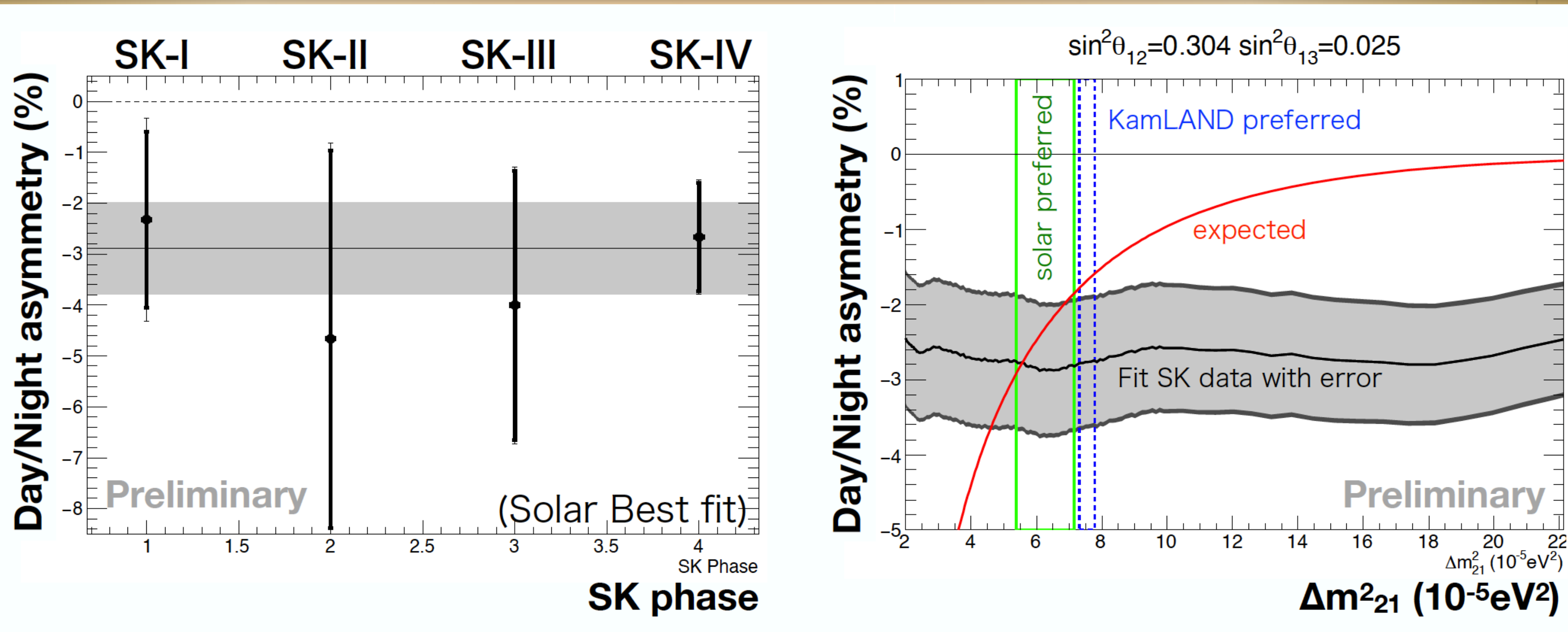
- Need to **lower current SK solar neutrino energy threshold** to observe transition between matter and vacuum oscillations in solar neutrinos. Background-limited
  - **Recoil electron kinetic energy 2.49 MeV**
- Dedicated hardware (WIT) searches un-triggered data for low-energy events & fits for fiducial vertex in real-time
- Solar neutrino identification down to 2.5 MeV recoil energy using BDT & WIT data





# Day/Night asymmetry at SK

- Day/Night flux asymmetry: 
$$A_{D/N} = \frac{\Phi_{day} - \Phi_{night}}{0.5(\Phi_{day} + \Phi_{night})}$$
  - Expected D/N asymmetry  $\sim 3\%$  in the SK energy range for solar neutrinos
  - Sk confirmed a higher solar  $\nu$  flux at the night than that at daytime with **significance  $3.2\sigma$  for the Solar best fit case,  $3.1\sigma$  for the Global best fit case**



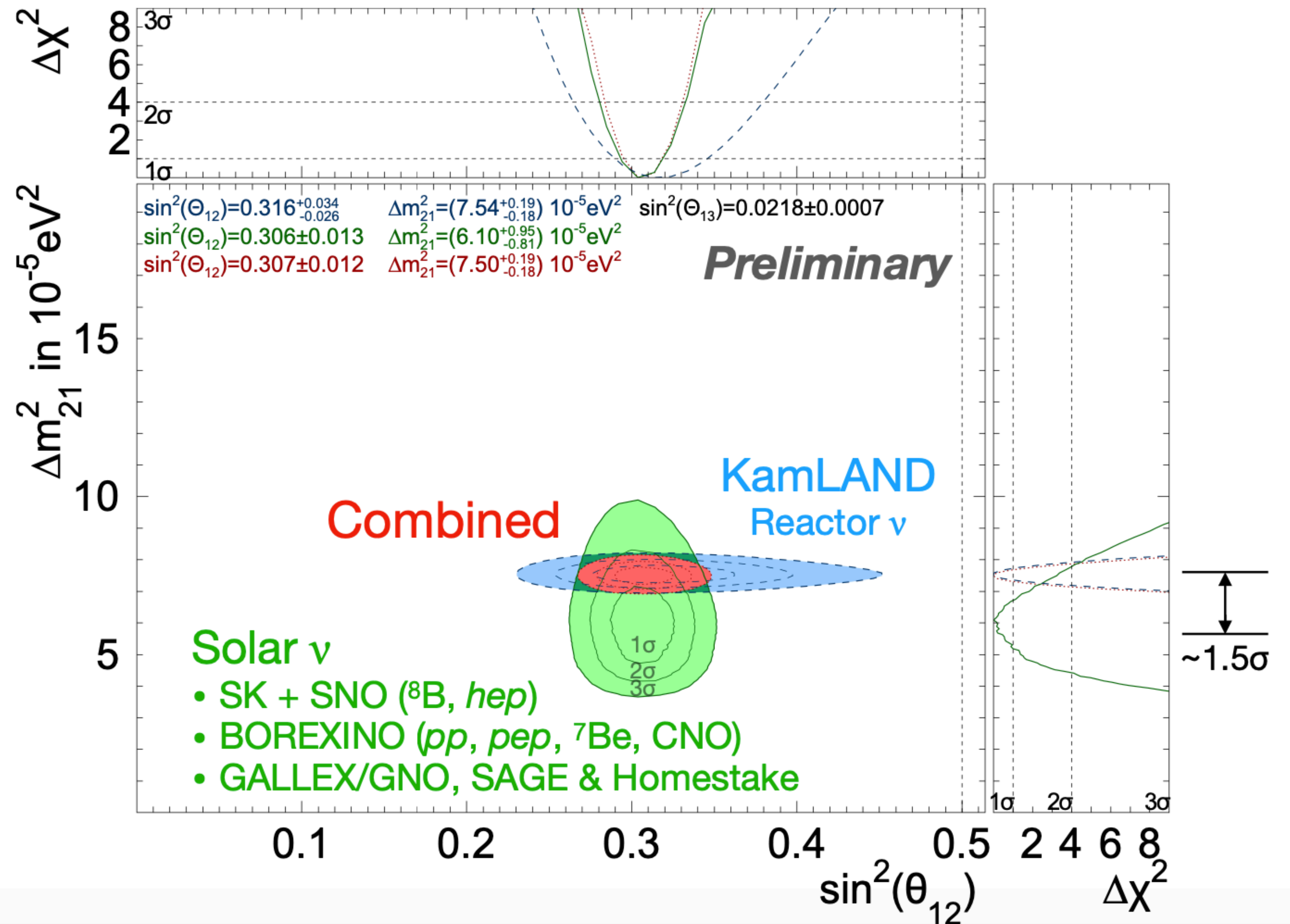


# Solar $\nu$ oscillation parameters

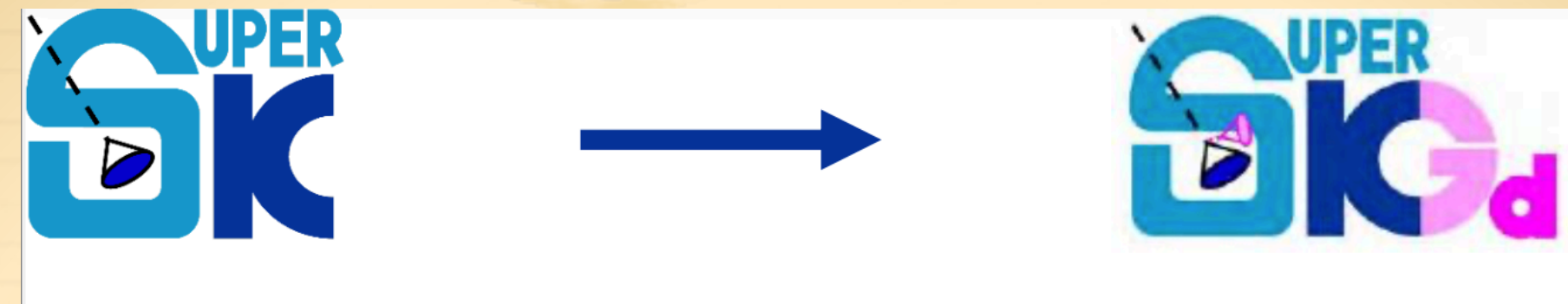
- There is  $1.5\sigma$  tension between SK+SNO and KamLAND in  $\Delta m_{21}^2$

Best fit oscillation parameters

Experiment	$\sin^2 \theta_{12}$	$\Delta m_{21}^2$
KamLAND	$0.316^{+0.034}_{-0.026}$	$7.54^{+0.19}_{-0.18} \times 10^{-5} \text{ eV}^2$
SK + SNO	$0.305 \pm 0.014$	$6.10^{+1.04}_{-0.75} \times 10^{-5} \text{ eV}^2$
Combined	$0.305^{+0.013}_{-0.012}$	$7.49^{+0.19}_{-0.17} \times 10^{-5} \text{ eV}^2$

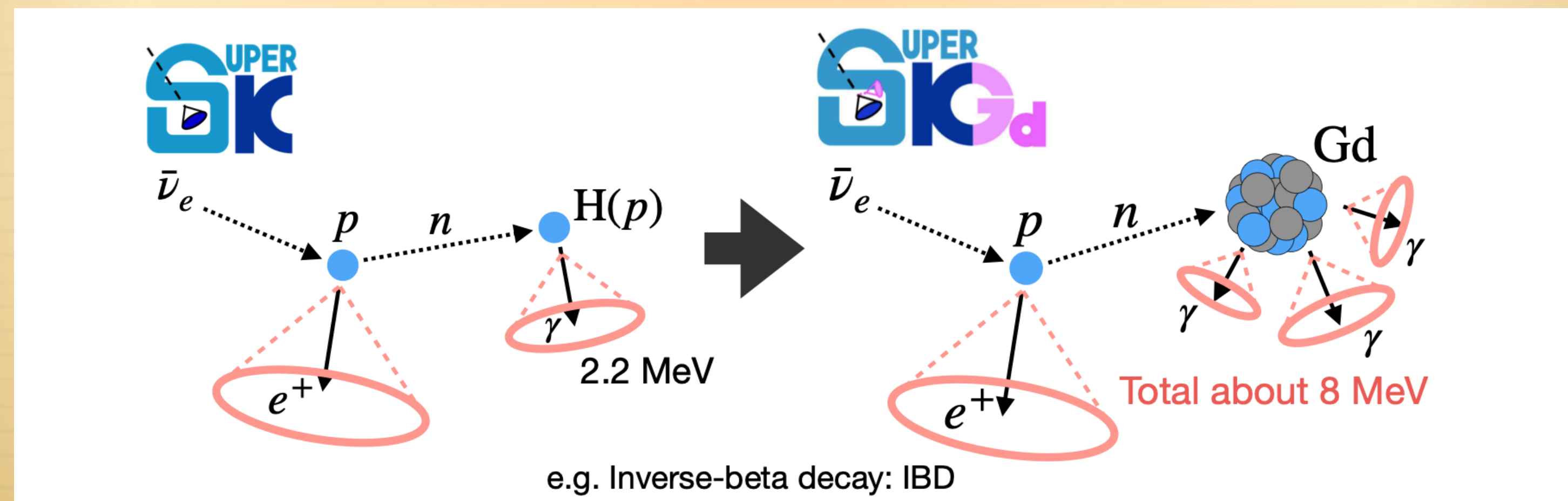






# SK-Gd era

Gadolinium project at Super-K: SK-Gd







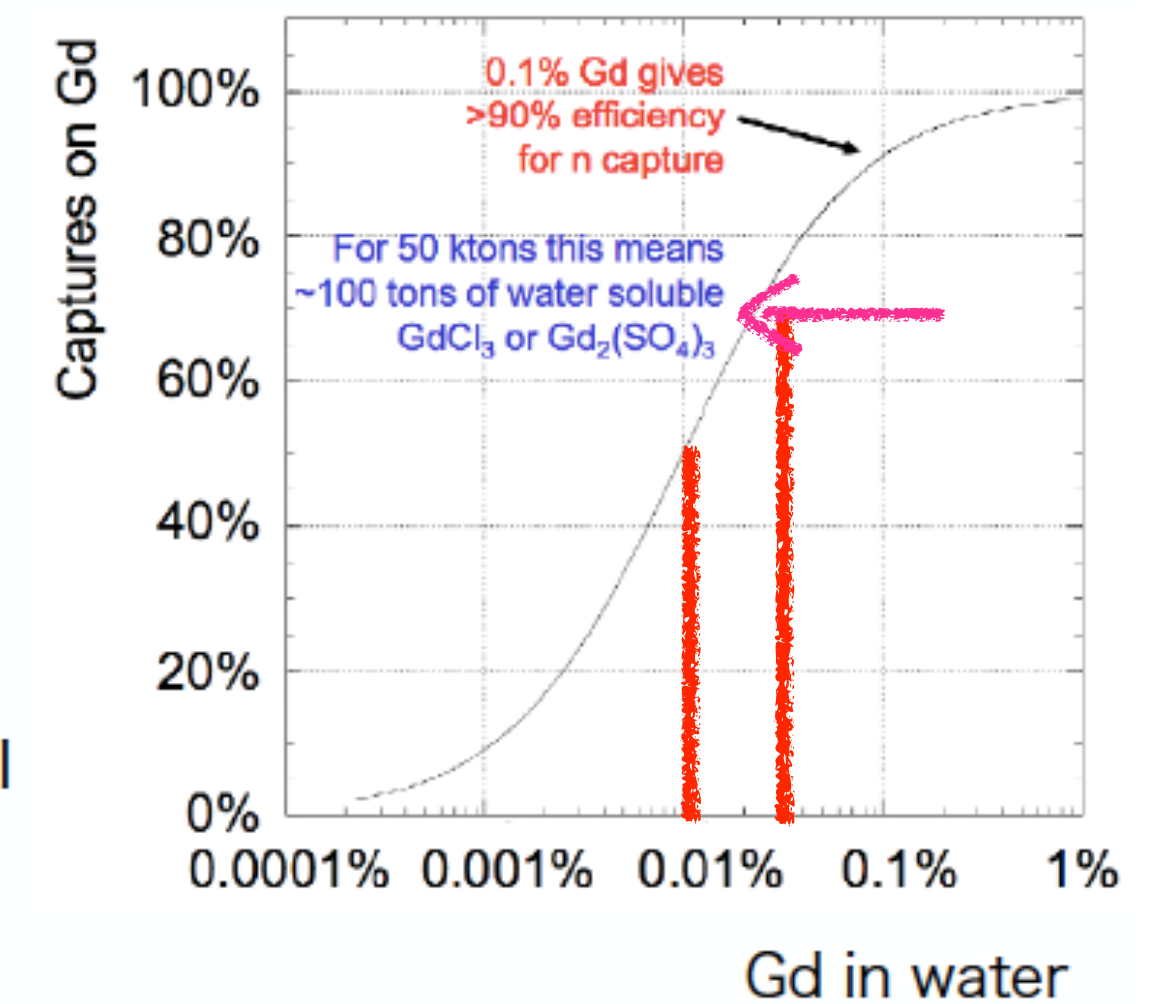
# Why Gd salt was added ?

- SK-Gd: add Gd sulphate to ultra pure water to enhance neutron tagging efficiency
- Physics targets:
  - Detect the world's first Diffuse Supernova Neutrino Background (DSNB)
  - Improvement of supernova direction pointing accuracy and allowing pre-supernova neutrino detection (early warning for SN).
  - Enhance  $\nu$  and  $\bar{\nu}$  identification in atmospheric  $\nu$  and T2K analyses
  - Reduce background in nucleon decay search
- Observed clear increase of neutron candidates and shorter capture time
- Confirmed uniformity of Gd concentration over SK detector and its stability over time

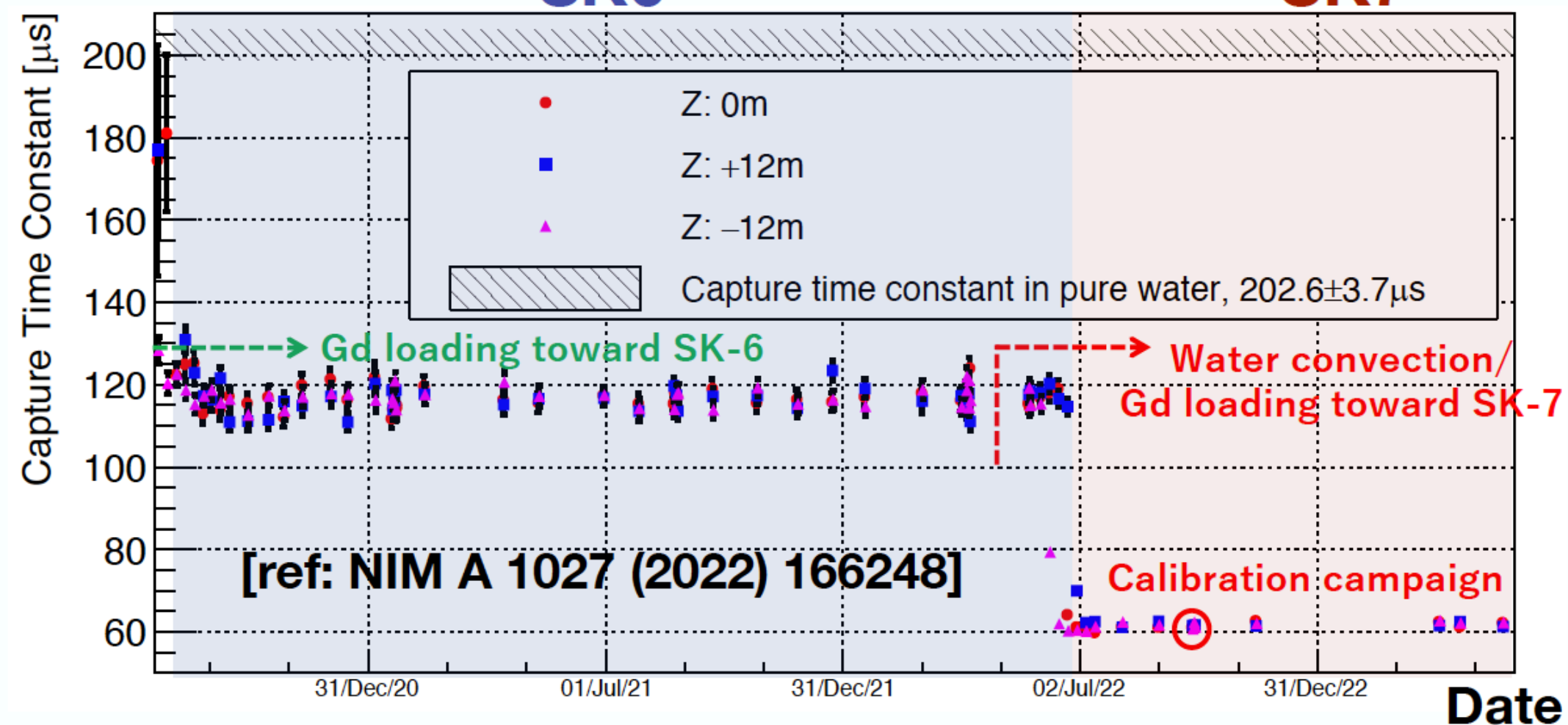
Dissolve Gadolinium into Super-K

J.Beacom and M.Vagins,

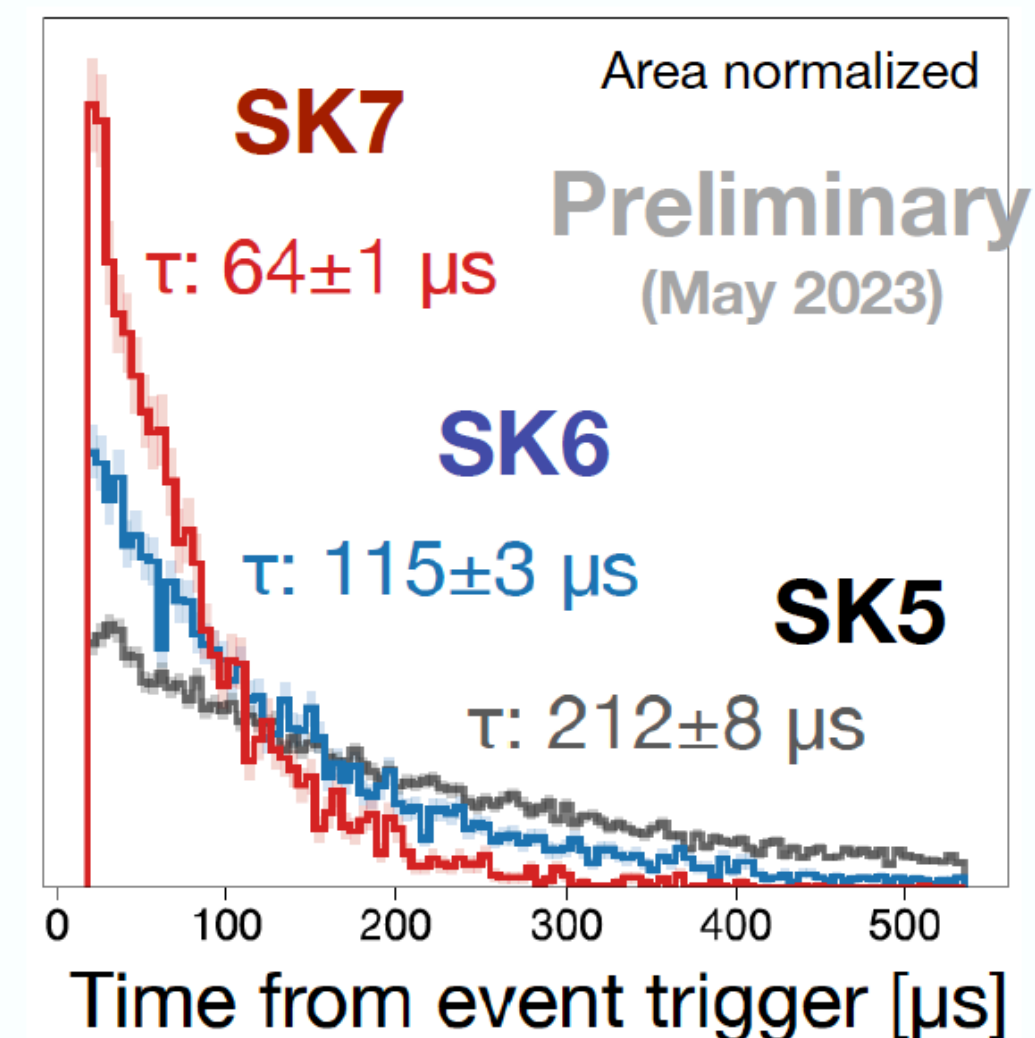
Phys.Rev.Lett.93(2004)171101



n-capture time with **AmBe** calibration source



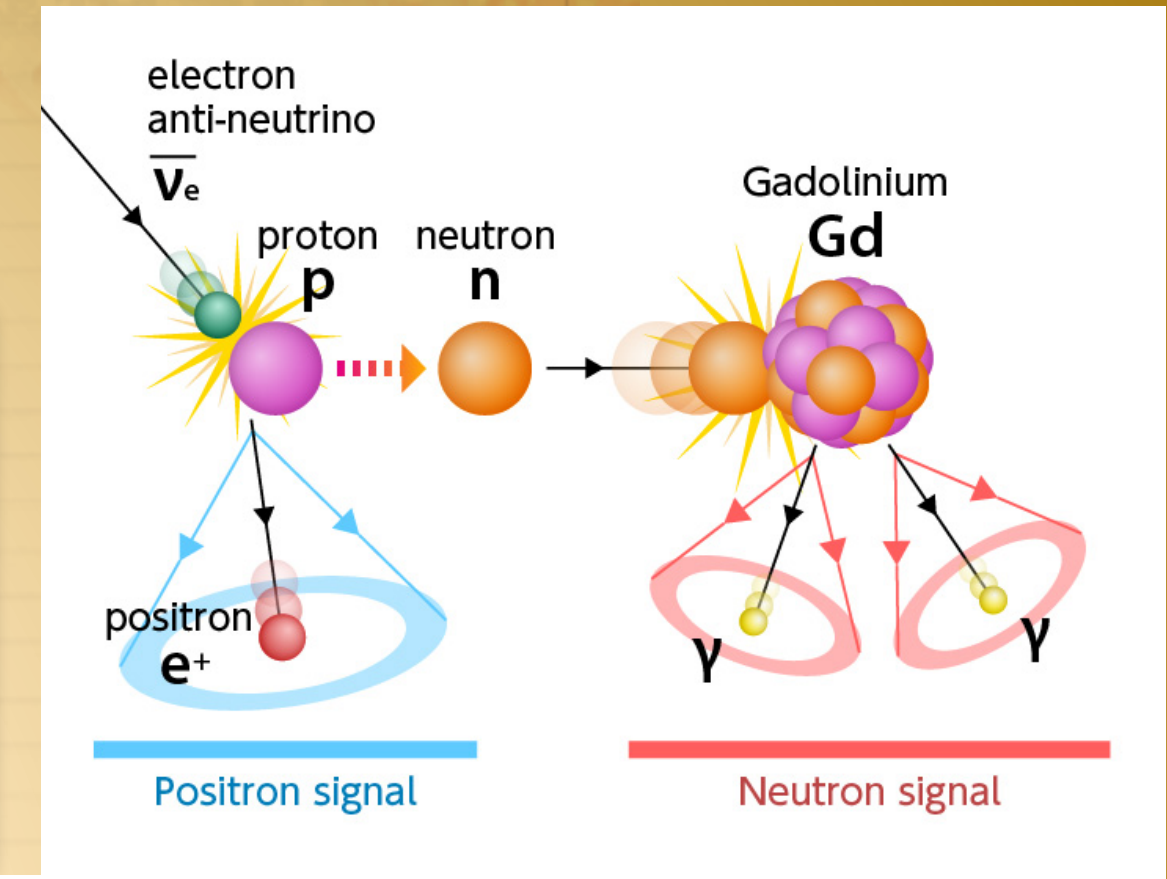
n-capture time in atm- $\nu$  events



We have finished Gd loading on July 25th!! 2022  
SK-VII era is now on

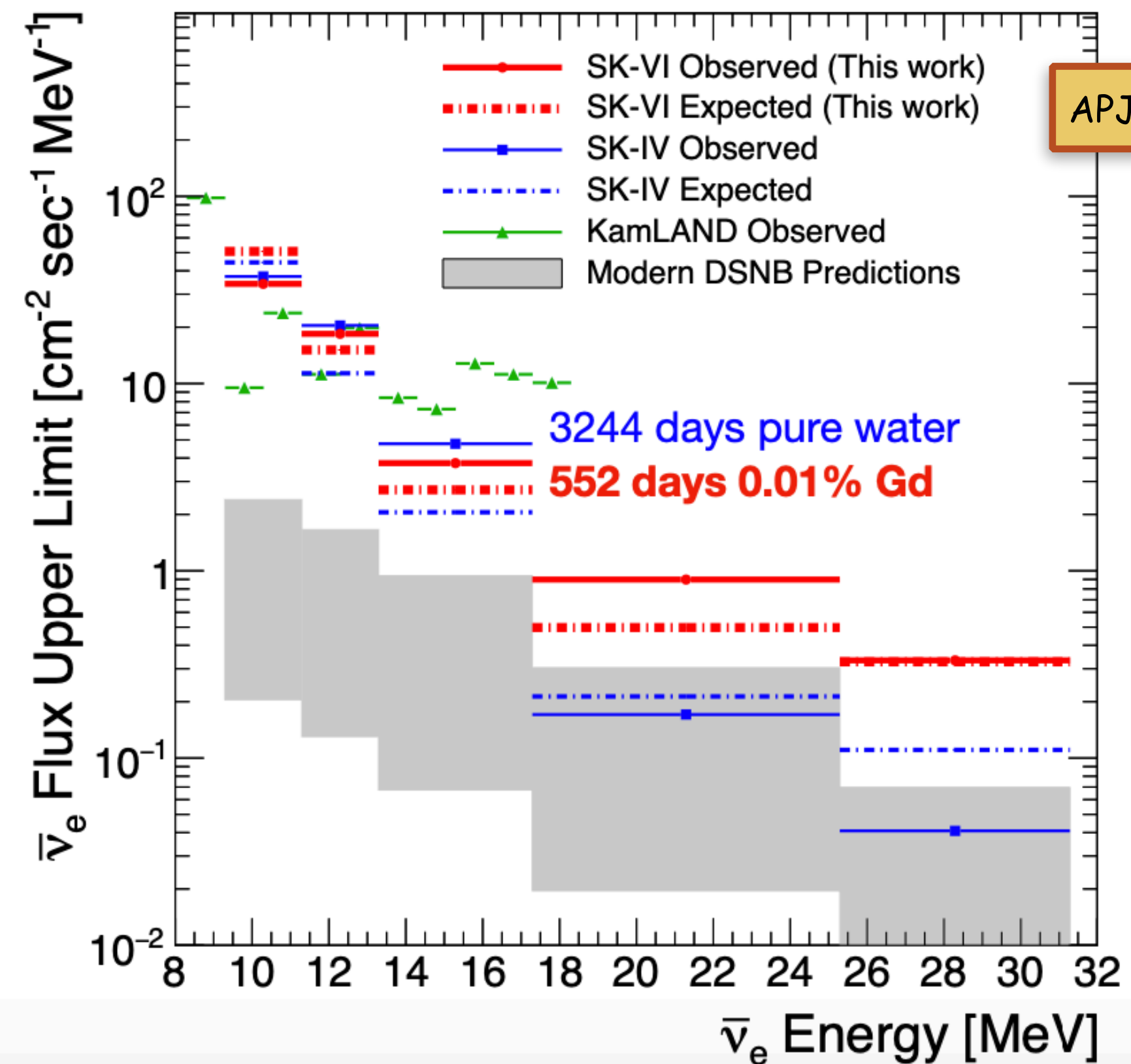
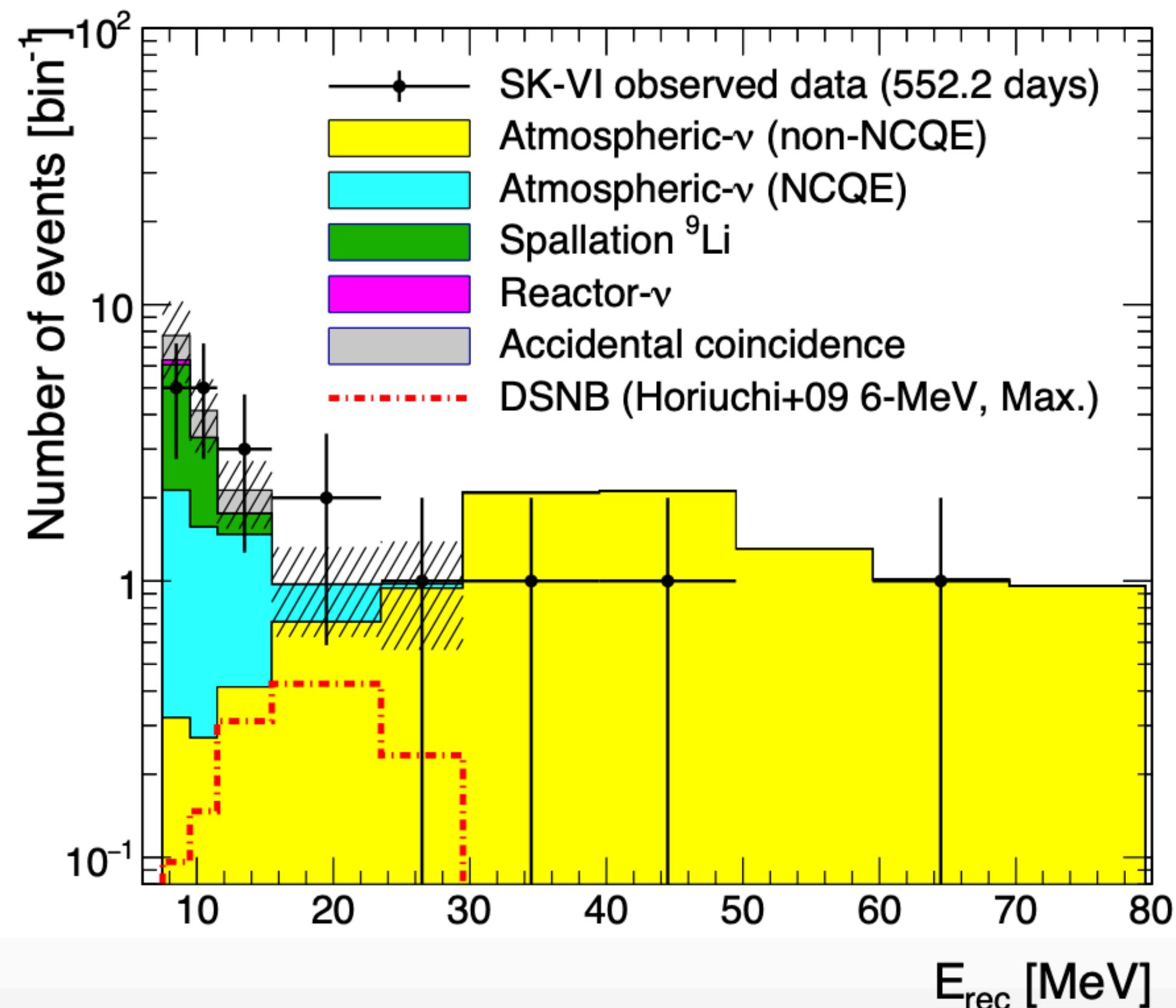


# SK-GD DSNB analysis



**Diffuse Supernova Neutrino Background (DSNB)** Integrated flux of the neutrinos emitted from past all Supernovae:

- Not-yet observed neutrino source expected from all past supernova
- ~7.5-30 MeV IBD signal window avoids reactor neutrinos & atmospheric background
- 0.01% SK-Gd data analysed: Sensitivity is close to theoretical predictions, competitive with pure water phases



APJ LETT 951:L27

See the dedicated talk on this conference from Masayuki Harada



# Summary

## Atmospheric neutrinos:

- Analysed all pure water data sets (SK-I ~V) and with expanded fiducial volume (total 27.2kton) and using information on neutron tagging on hydrogen - new paper in preparation
- New results of the SK+T2K joint fit
- New results on  $\nu_\tau$  appearance searches

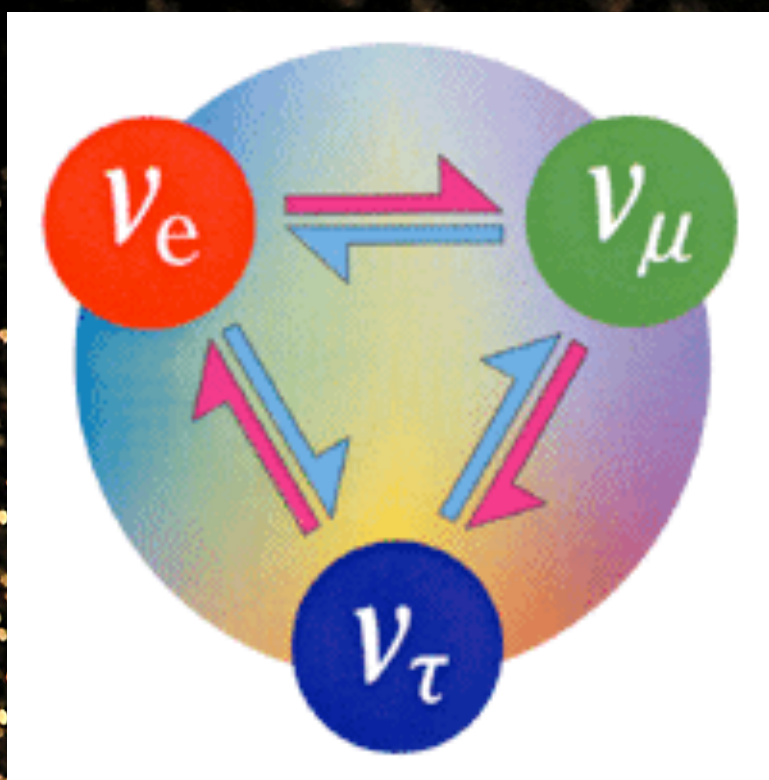
## Solar neutrinos:

- Global analysis including SK I-IV is compatible with KamLAND
- Lowering the energy threshold to 2.49MeV (kinetic energy) aiming to see the "up-turn"

## SK- Gd:

- Neutron tagging is working, observing many more captures
- First DSNB search in SK-Gd was performed using the data from Aug. 2020 to Jun. 2022. No significant excess over expected background was seen

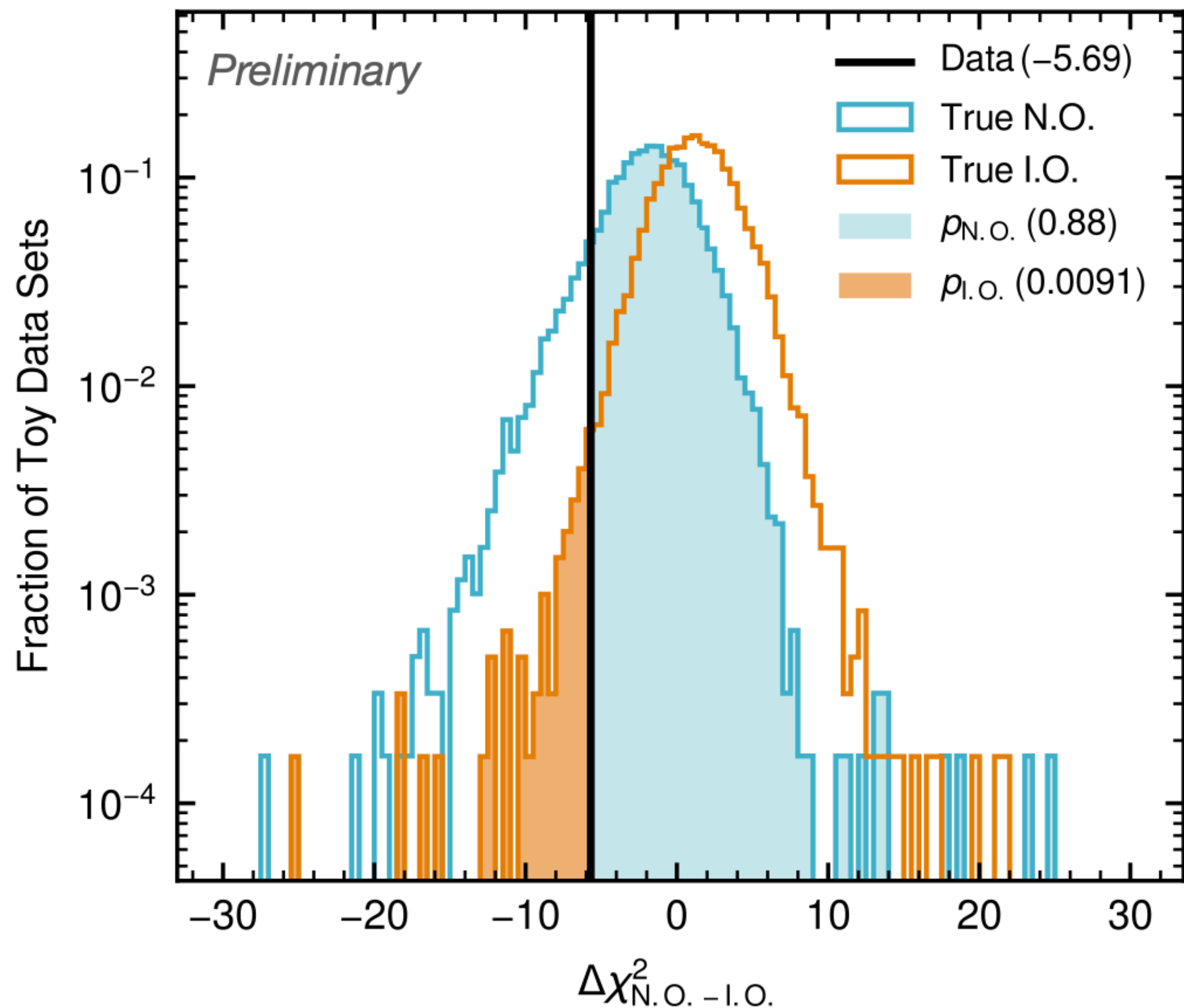




Thank you!



# Mass ordering significance



- Generate toy data sets to obtain distribution of  $\Delta\chi^2_{\text{NO-IO}}$
- SK data fit result has lower probability of occurring for a true inverted ordering than normal, but is also more extreme than normal ordering median (exceeds sensitivity)
- Compute CLS statistic to correct probability of rejecting the inverted ordering given simultaneous agreement with normal ordering:

$$\text{CL}_S = \frac{p_{\text{IO}}}{1 - p_{\text{NO}}} \approx 0.077$$

*Reject inverted ordering at the ~92% confidence level.*