



Vertex and Energy Reconstruction in JUNO

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On behalf of the JUNO collaboration

October 26, 2023

The XX International Workshop on Neutrino Telescopes
(NeuTel 2023)

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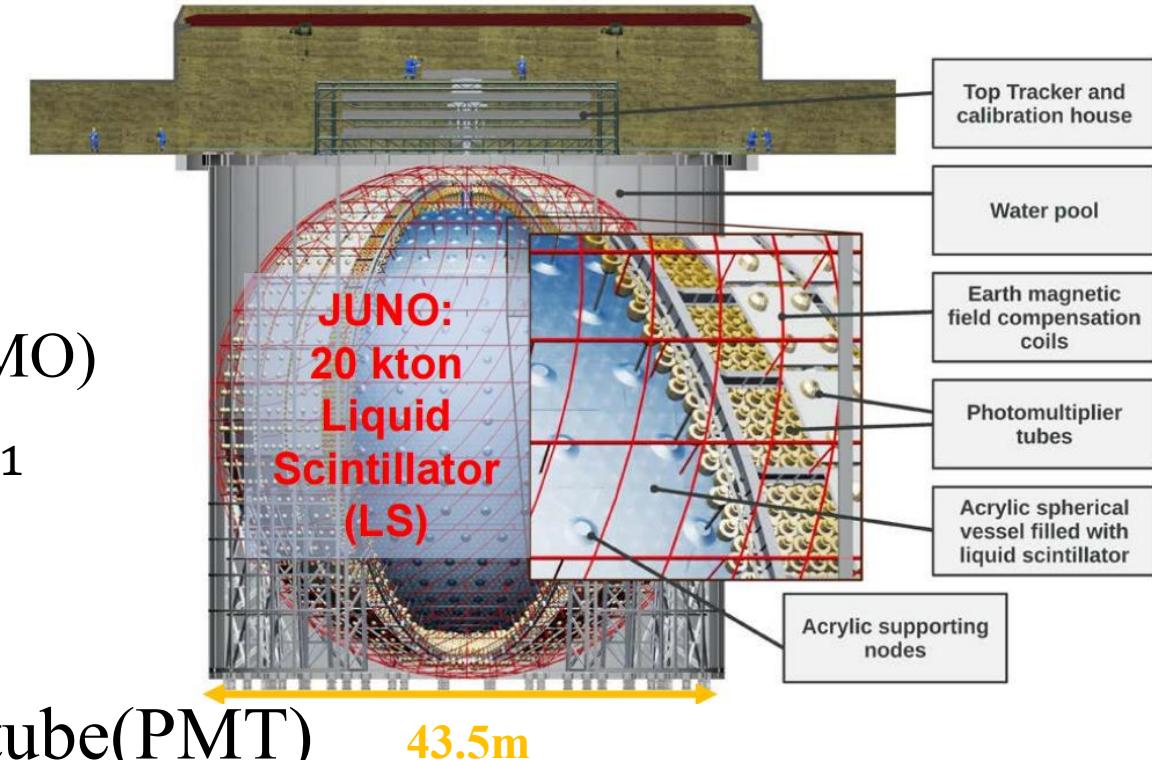
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The JUNO experiment

- A multi-purpose observatory
 - Determine the neutrino mass ordering(NMO)
 - Precisely measuring $\sin^2 \theta_{12}$, Δm_{21}^2 , Δm_{31}^2
 - Supernova, Atmosphere, Solar, etc.
- 20k tons liquid scintillator(LS)
- ~17612 large 20-inch photomultiplier tube(PMT)
- ~25600 small 3-inch PMT
- Unprecedented energy resolution
 - 3% @ 1 MeV



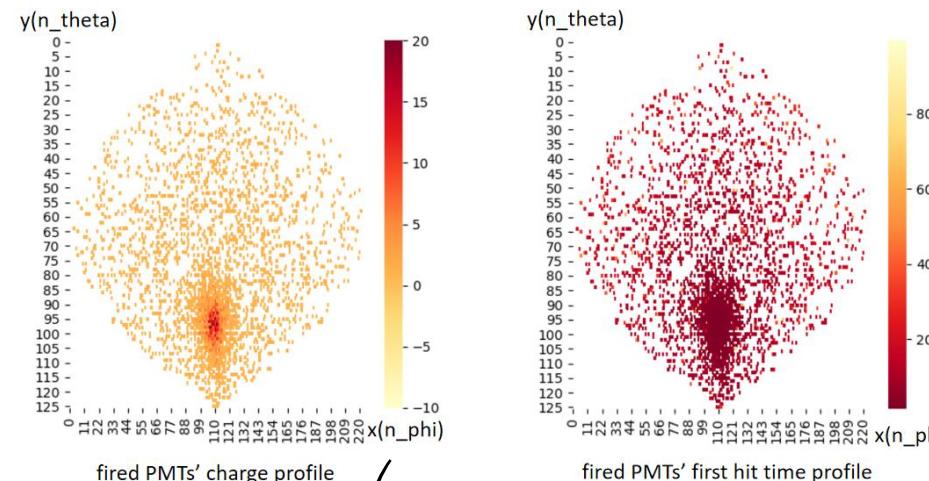
	LS detectors			
	KamLAND	Borexino	Daya Bay	JUNO
mass	1 kton	0.3 kton	20 ton	20 kton
resolution	6.5%	5%	8%	3%

Data driven reconstruction

Now the study is based on **simulation**



- The observables are large PMTs' hit **charge** $\{q_i\}$, residual hit **time** $\{t_{i,r}\}$ and small PMTs' hit **states**
- The nPE map $\hat{\mu}$, time PDF P_T and charge PDF P_Q are constructed from ^{68}Ge calibration data



r : vertex radial distance

d : distance to PMT

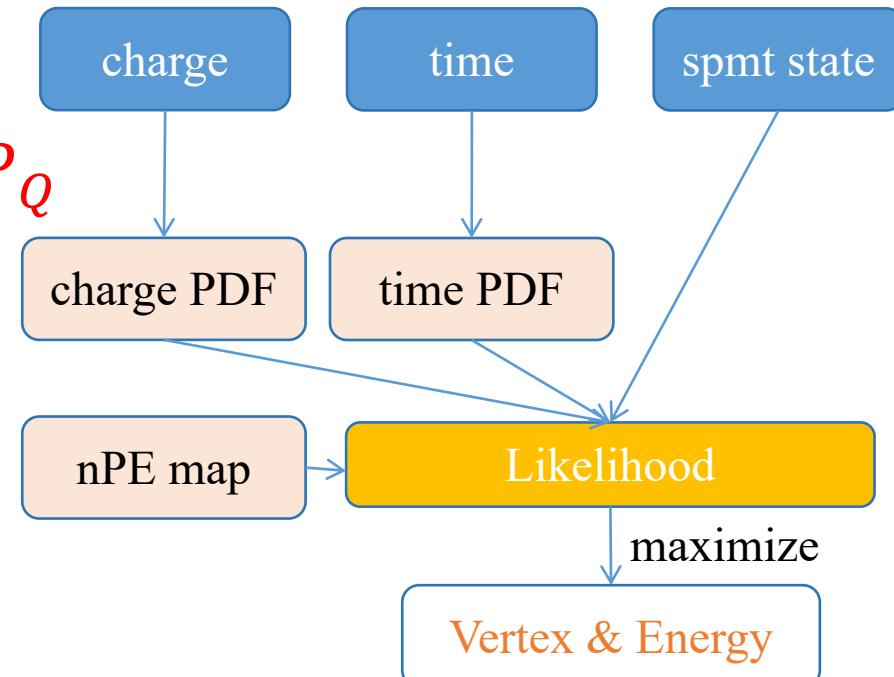
μ_l : LS PE

μ_d : dark noise PE

t_0 : event time

$$\mu_j = \hat{\mu}_j E + \mu_j^d$$

$$L(\{q_i\}; \{t_{i,r}\} | \vec{r}, E, t_0) = \prod_{unhit} e^{-\mu_j} \prod_{LPMT\ hit} \left(\left(\sum_{k=1}^{+\infty} \frac{e^{-\mu_j} \mu_j^k}{k!} P_Q(q_i|k) \right) P'_T(t_{i,r}|r, d_i, \mu_j^l, \mu_j^d) \right) \prod_{SPMT\ hit} (1 - e^{-\mu_j})$$

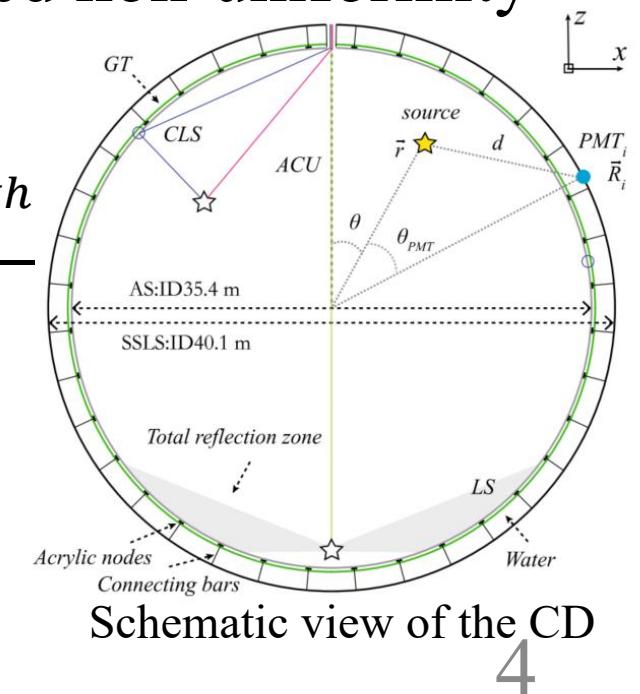




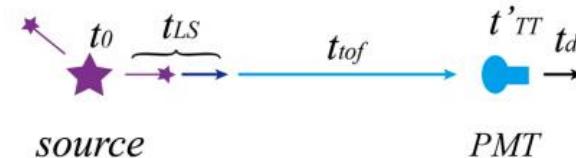
Construction of the nPE map

- nPE map $\hat{\mu}$: The expected number of LS photos per unit visible energy
- Method: Use ^{68}Ge to construct a 3D nPE map: $\hat{\mu}(r, \theta, \theta_{pmt})$, assume the nPE map is ϕ independent
- Advantage: More precise nPE map than 2D, optimized non-uniformity

$$\hat{\mu}(r, \theta, \theta_{pmt}) = \frac{1}{E_{vis}} \frac{1}{N_{evt}} \frac{1}{N_{pmt}} \sum_{i=1}^{N_{evt}} \sum_{j=1}^{N_{pmt}} \frac{\frac{q_{i,j}}{\hat{q}_j} - D_{ark} N_{oise} R_{ate} \cdot L_{ength}}{D_{etect} E_{fficiency}_j}$$



Construction of the time pdfs

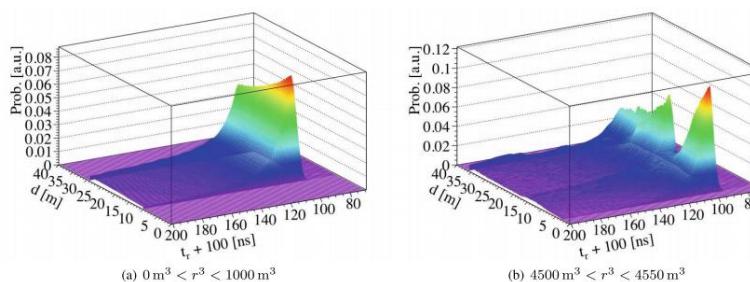


- The pdfs of the probability of the residual time of first hit

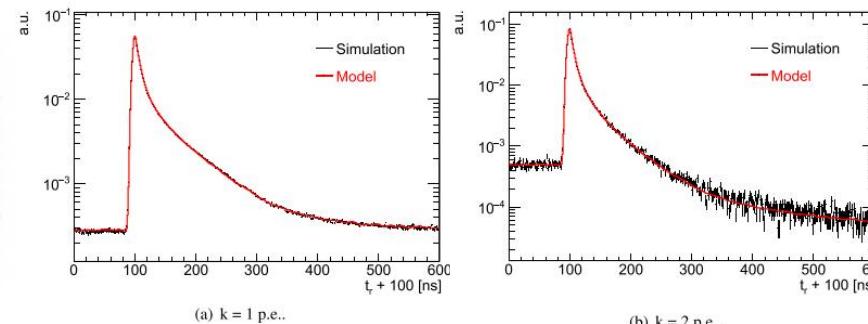
$$t_r = t_{hit} - t_{tof} - t_{delay} - t_0$$

r : vertex radial distance
 d : distance to PMT
 μ_l : LS PE
 μ_d : DN PE

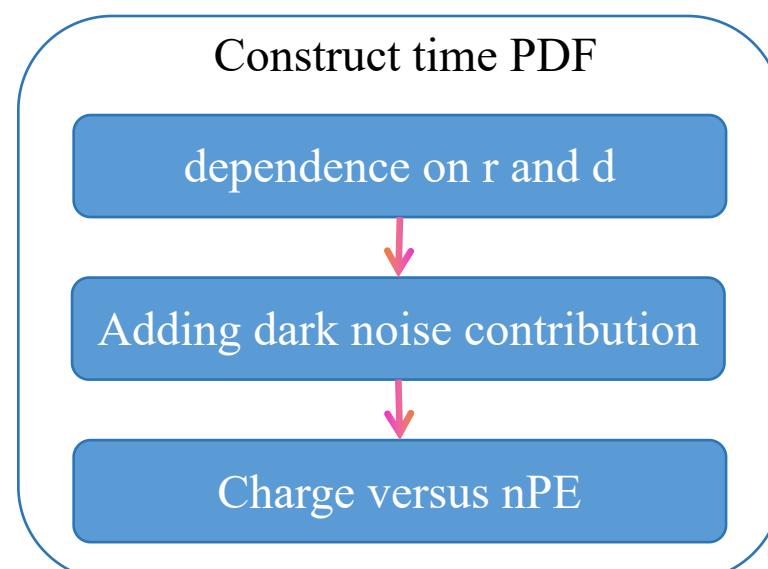
- Challenges: Vertex dependent; Dark noise(DN) contamination
- Method: Use ^{68}Ge calibration data to extract $P'_T(t_r|r, d, \mu_l, \mu_d)$
- Advantage: more precise time pdf, make resolution better near the centre detector edge



Examples of the time PDFs at different positions



Time PDF of Dynode-PMT of central ^{68}Ge events considering dark noise

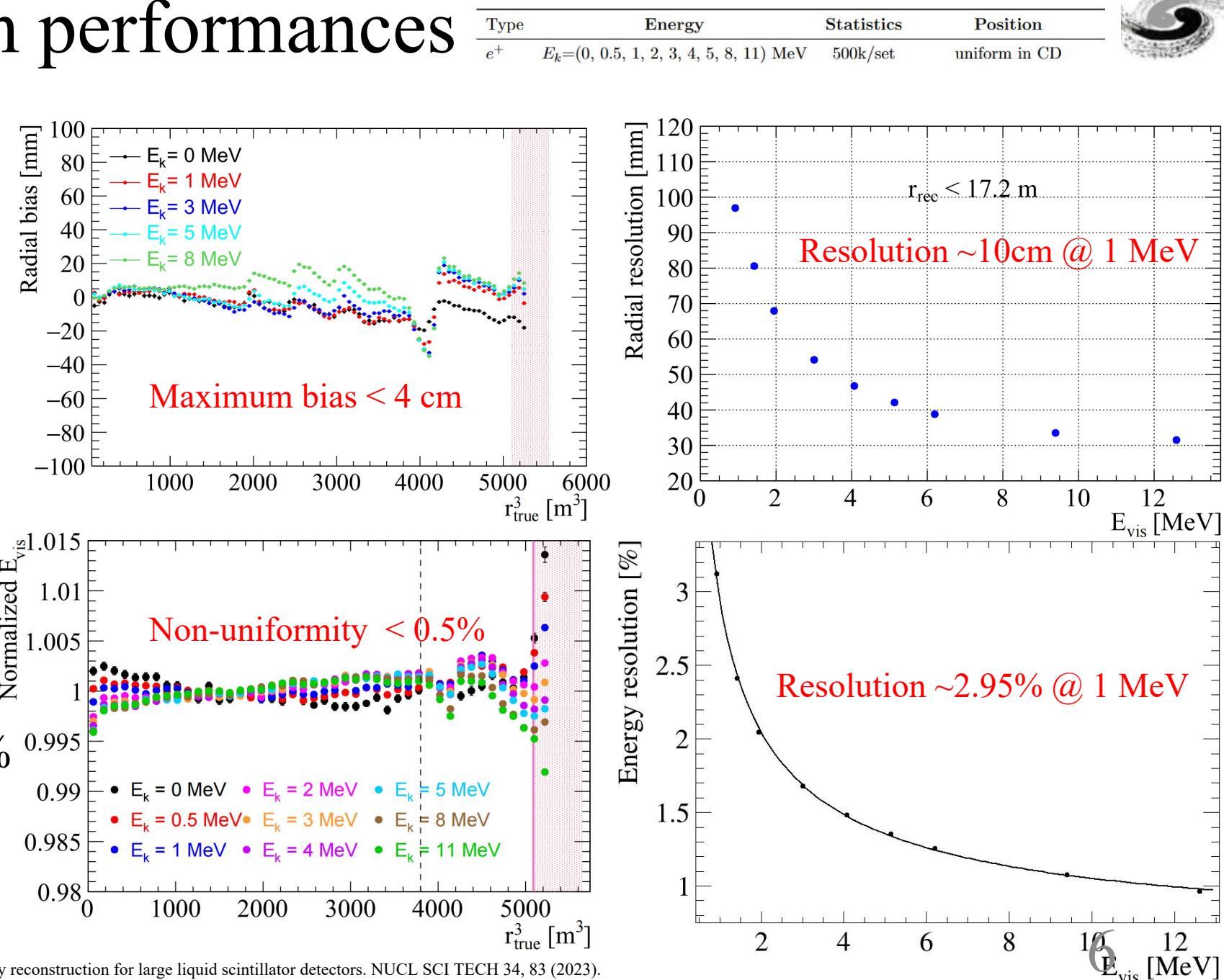


Reconstruction performances



The performances are evaluated with JUNO MC simulation data of uniformly generated positron samples.

- Radial performance
 - Maximum bias < 4 cm
 - Resolution ~ 10 cm
- Energy performance
 - Non-uniformity $< 0.5\%$
 - Resolution $\sim 2.95\%$

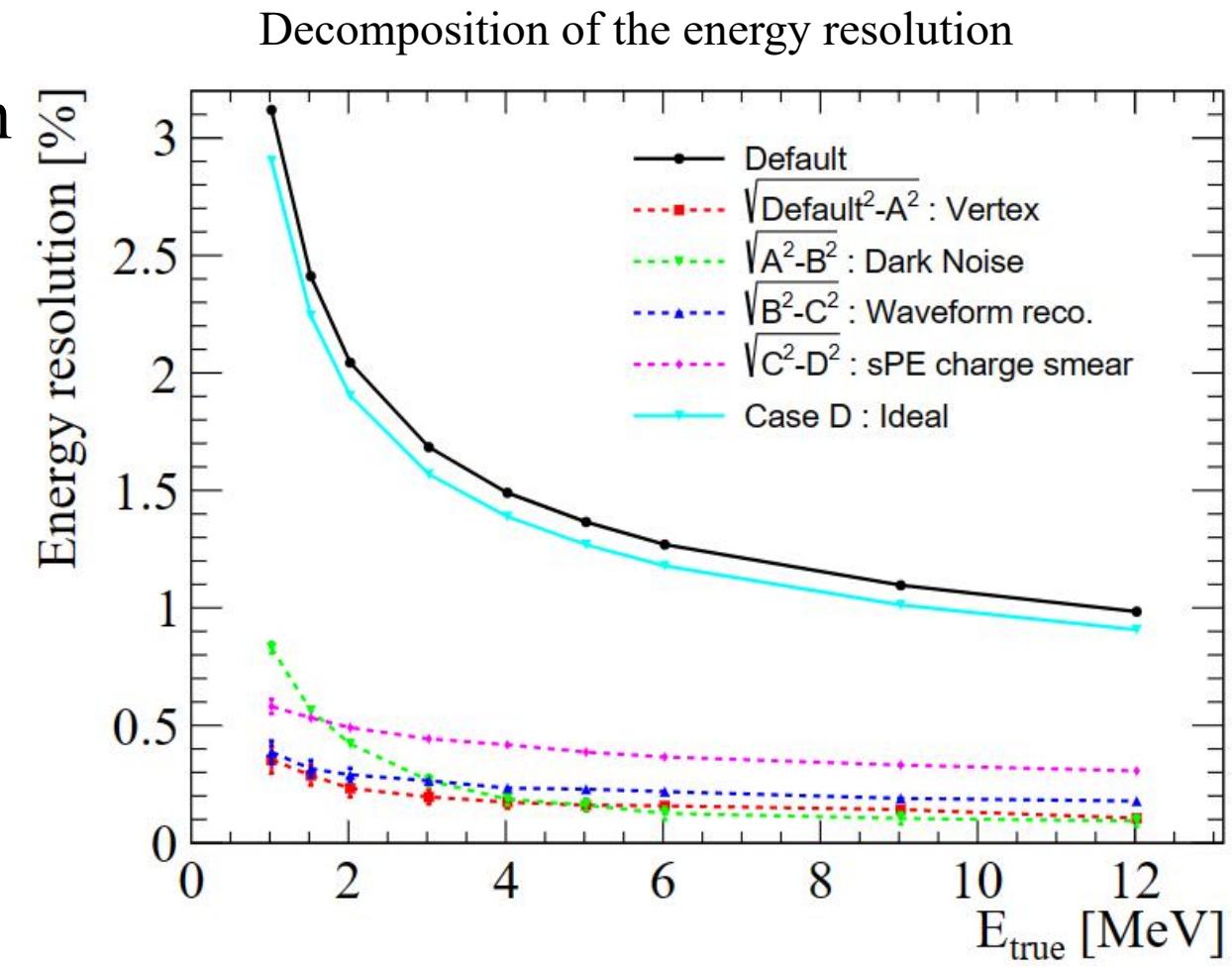


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Energy resolution decomposition

- Five components of energy resolution
- Two dominant factors:
 - Below 1.5MeV - Dark Noise
 - Above 1.5MeV - sPE charge smear
PMT convert PE to charge
- Works are ongoing to mitigate the impact of these factors





Summary

- Data driven method makes nPE map, time PDF and charge PDF realistic
- Simultaneous reconstruction of vertex and energy with both charge and time information of PMTs
- Latest performance
 - Vertex : bias < 4 cm, resolution ~ 10 cm @ 1 MeV
 - Energy: non-uniformity < 0.5%, resolution 2.95% @ 1 MeV
- Energy resolution decomposition shows two dominant factors: dark noise and sPE charge smear
- More works are ongoing



Thanks for your attention!