

Next-generation CE ν NS experiments at the ESS and beyond

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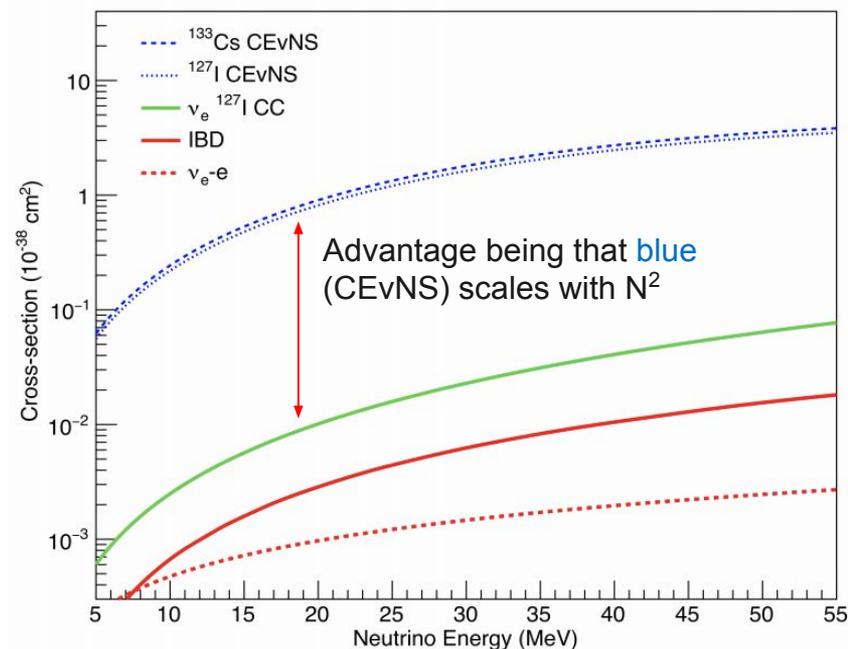
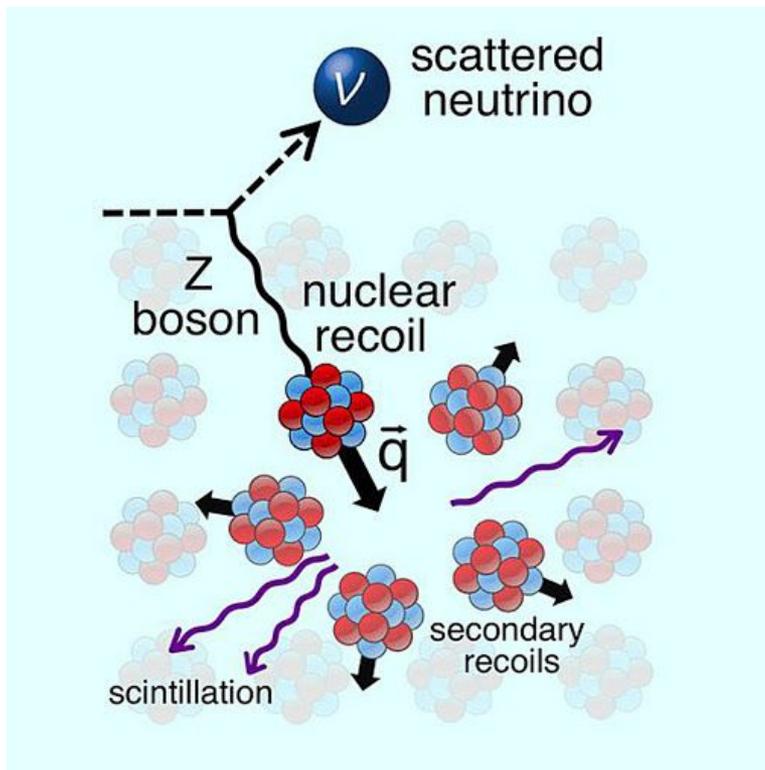
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Coherent Elastic Neutrino-Nucleus Scattering (CEvNS)

D.Z. Freedman, Phys. Rev. D 9 (1974) 1389



What you can do with it

Probe for non-standard interactions through cross section

(J.R. Wilson, PRL 32 (74) 849)
(J. Barranco et al., hep-ph/0508299, hep-ph-0512029)

Neutrino magnetic moment (enhancement of low-energy recoils) (A.C.Dodd et al, PLB 266 (91) 434)

Neutrino charge radius
(L.M.Krauss, PLB 269, 407)

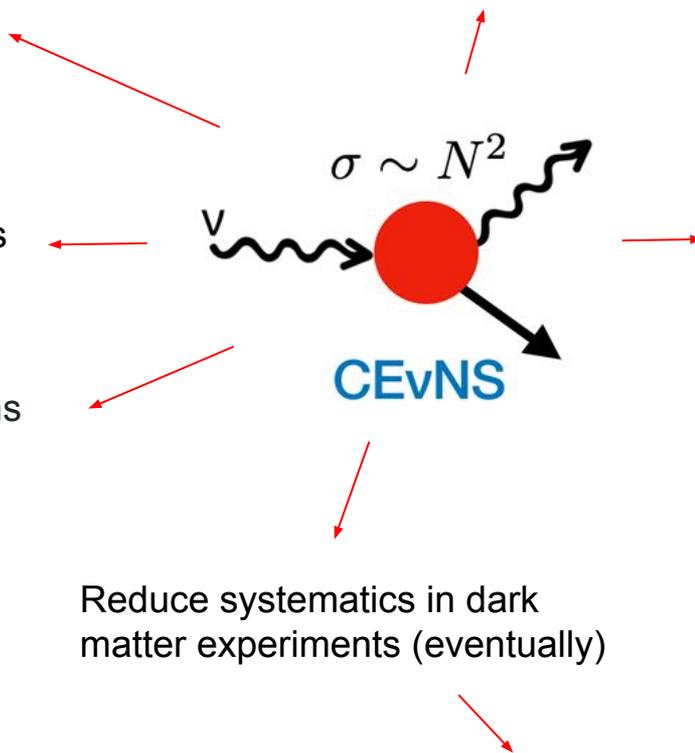
Sterile searches through oscillations
(CEvNS is a flavor-blind process)

(A.Drukier & L.Stodolsky, PRD 30 (84) 2295)

Reduce systematics in dark matter experiments (eventually)

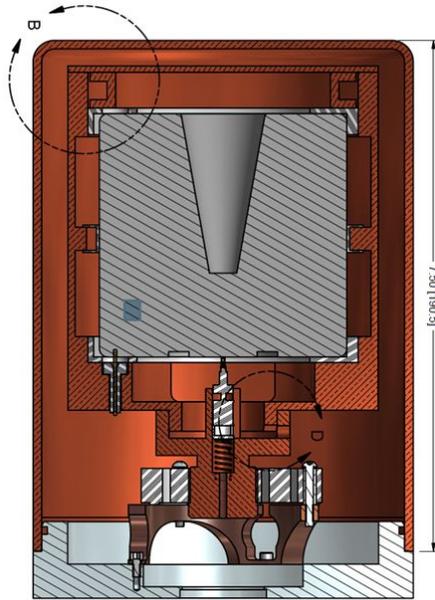
Probe for new neutral candidates

technology-scale detectors

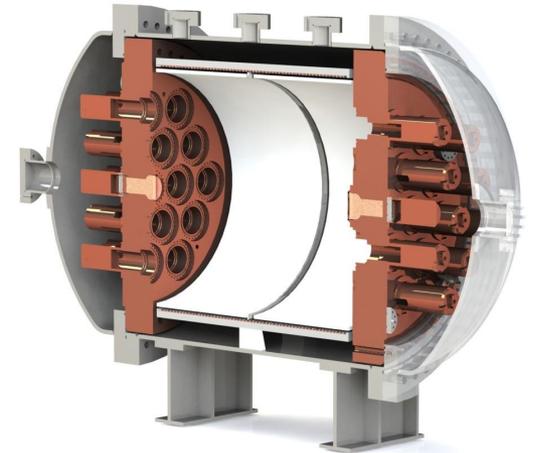




Cryogenic undoped CsI



p-type point contact Ge



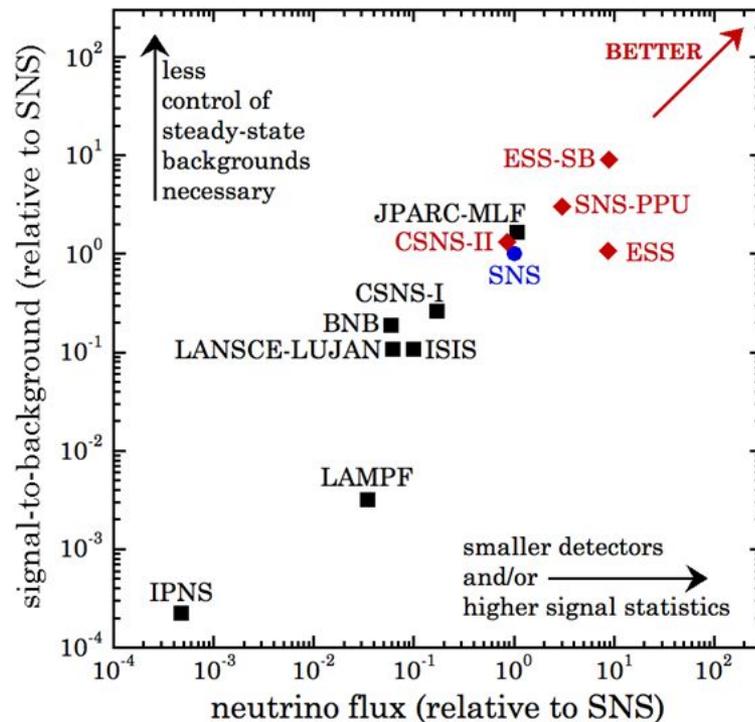
high pressure gas TPC

The best CE ν NS detectors deserve the best CE ν NS source

Enter... the European Spallation Source (ESS)



advantage: characteristic in both energy and time

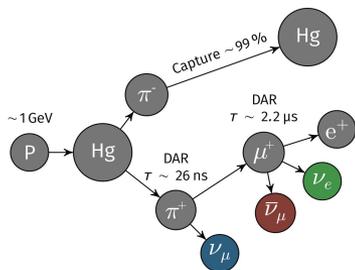


X10 the DAR ν production of the SNS

(=> signal statistics => sensitivity to new physics)

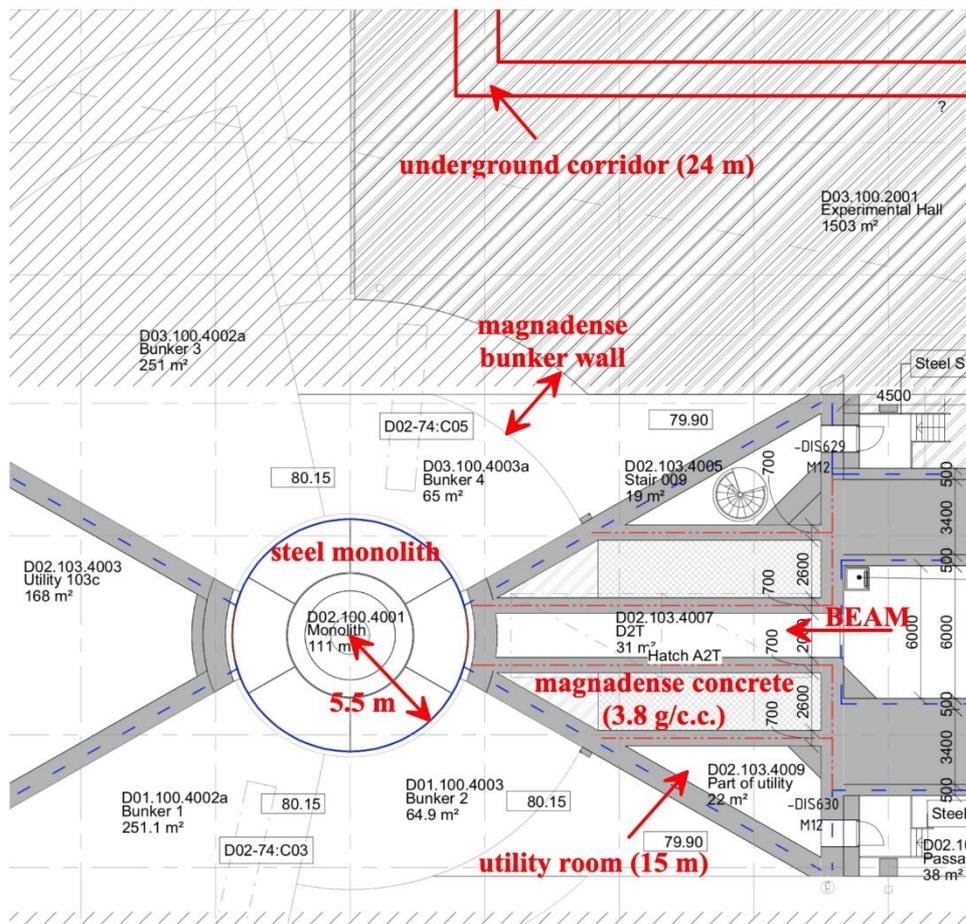
x2.5 SNS current (=> x2.5 SNS ν /p)

x2 SNS energy (=> ~x4 SNS ν /p)

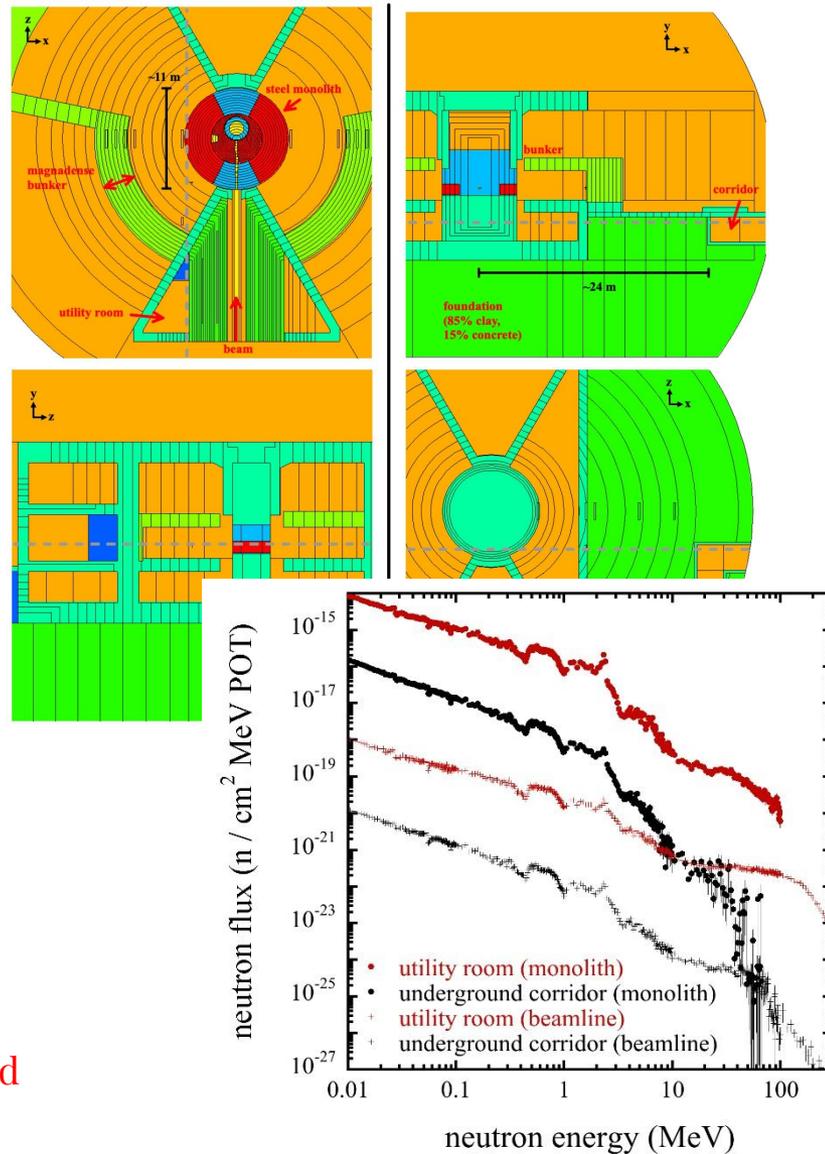


Detector Technology	Target nucleus	Mass (kg)	Steady-state background	E_{th} (keV $_{ee}$)	QF (%)	E_{th} (keV $_{nr}$)	$\frac{\Delta E}{E}$ (%) at E_{th}	E_{max} (keV $_{nr}$)	CE ν NS $\frac{NR}{yr}$ @20m, $>E_{th}$
Cryogenic scintillator	CsI	22.5	10 ckkd	0.1	~10 [71]	1	30	46.1	8,405
Charge-coupled device	Si	1	1 ckkd	0.007	4-30 [97]	0.16	60	212.9	80
High-pressure gaseous TPC	Xe	20	10 ckkd	0.18	20 [104]	0.9	40	45.6	7,770
p-type point contact HPGe	Ge	7	15 ckkd	0.12	20 [118]	0.6	15	78.9	1,610
Scintillating bubble chamber	Ar	10	0.1 c/kg-day	-	-	0.1	~40	150.0	1,380
Standard bubble chamber	C ₃ F ₈	10	0.1 c/kg-day	-	-	2	40	329.6	515

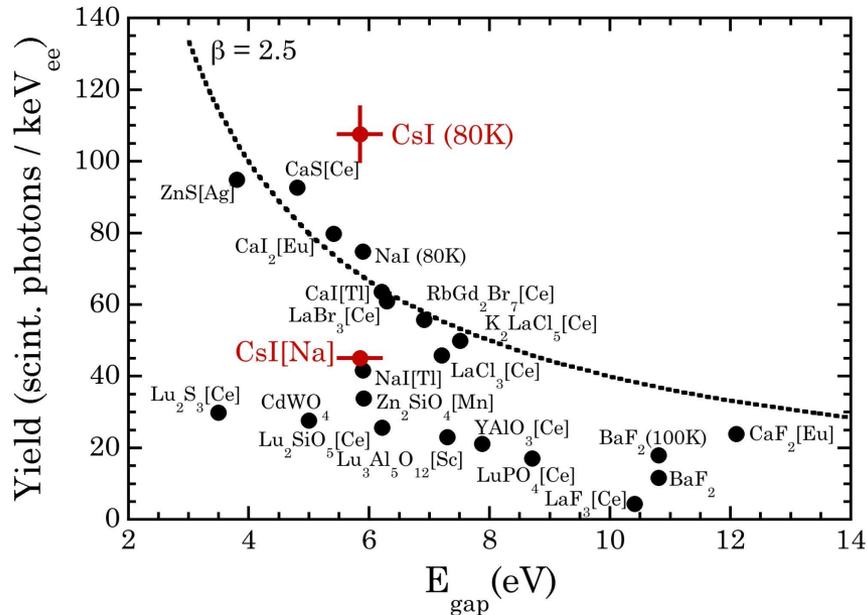
ESS (a detector home)



- Steady-state backgrounds subtractable
- beam-induced prompt neutrons are main background (candidate locations being evaluated)
- neutron camera for on-site measurements

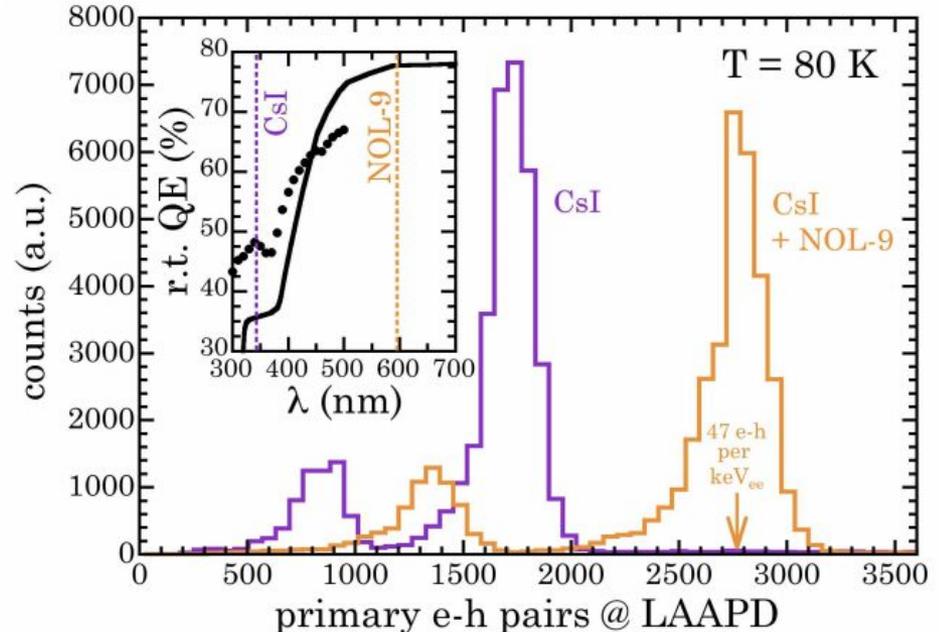


Signs of a deeper frontier (pure CsI)



- natural evolution from CsI[Na] measurement at SNS
- combine much higher light yield (x2.5-3) and more efficient photosensors
- large mass increase to ~60 kg (seven 7x7x40 cm crystals)

- LAAPDs with waveshifters (NOL-9) to increase synergy
 - Thresholds < 55 eV
 - limited by LAAPD-induced low-energy population

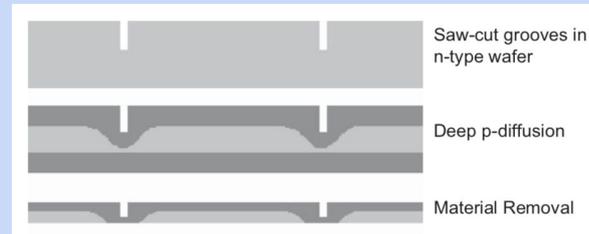


Joining of two houses

46 cm² LAAPD w/ 4 photon threshold from RMD



doping profiles boosting QE to visible spectrum



- can remove the complication of waveshifters if LAAPD QE can be brought to par
- R&D goals:
 - LAAPD sensitive to CsI light emission
 - radiopure cryostat stable at 80 K

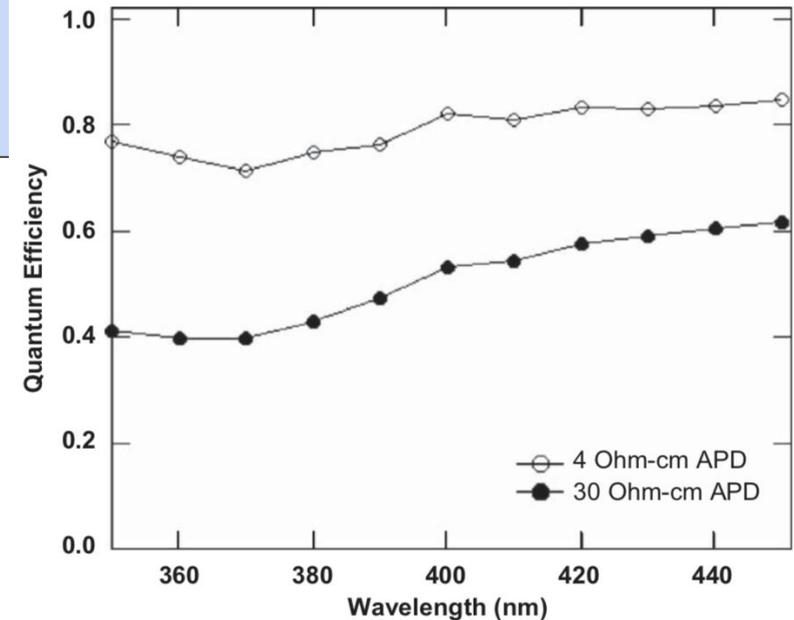
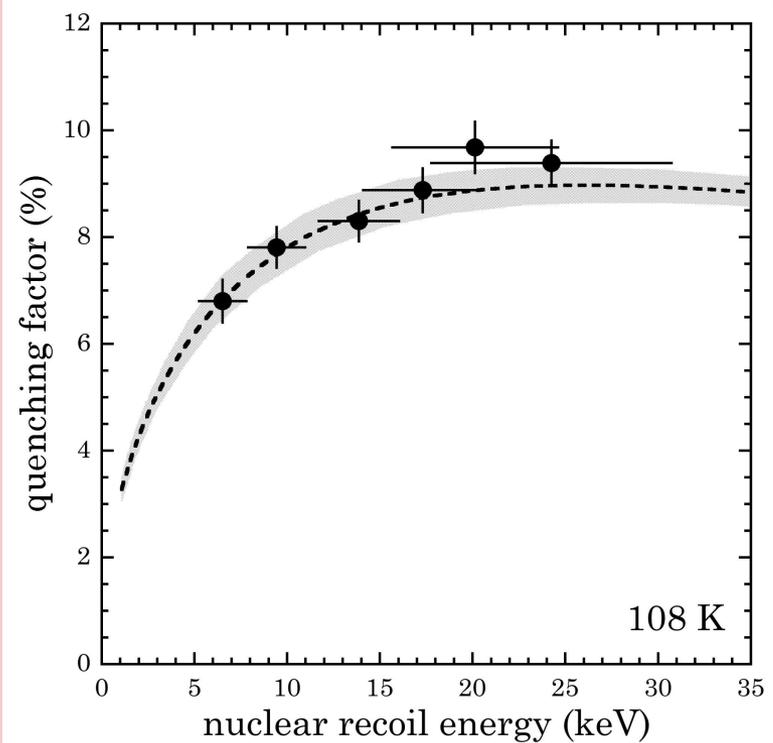
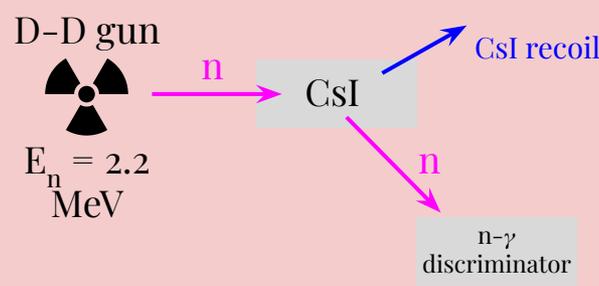
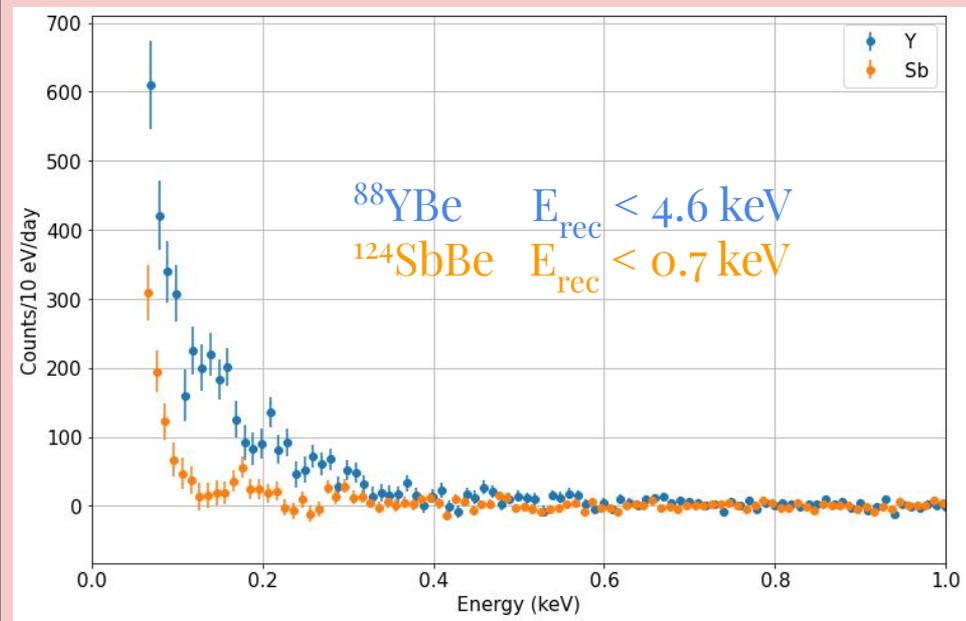
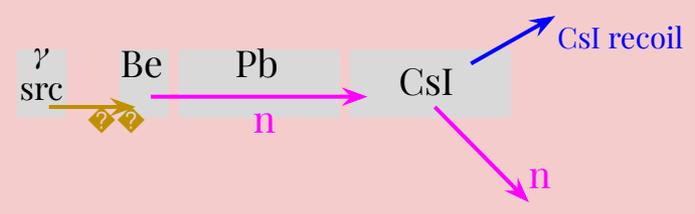


Fig. 3. Quantum efficiency vs. wavelength for a 4 and 30 Ω cm APD.

Cryogenic pure CsI quenching factor



arXiv:2101.03264



- switch Be with Al for background
- excess via signal - background
- **sub-keV recoils detected**

preliminary

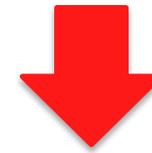
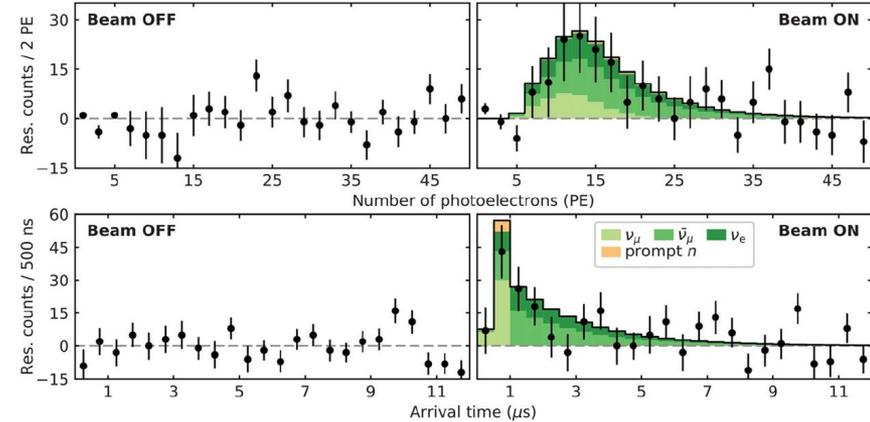
What does all that detector development get you? C^0SI



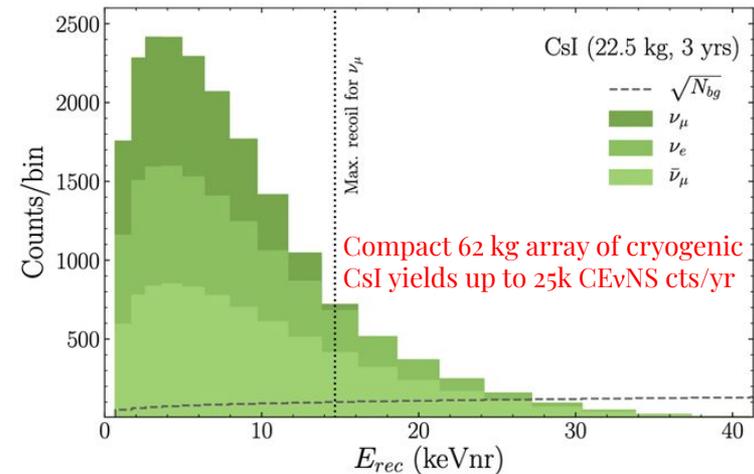
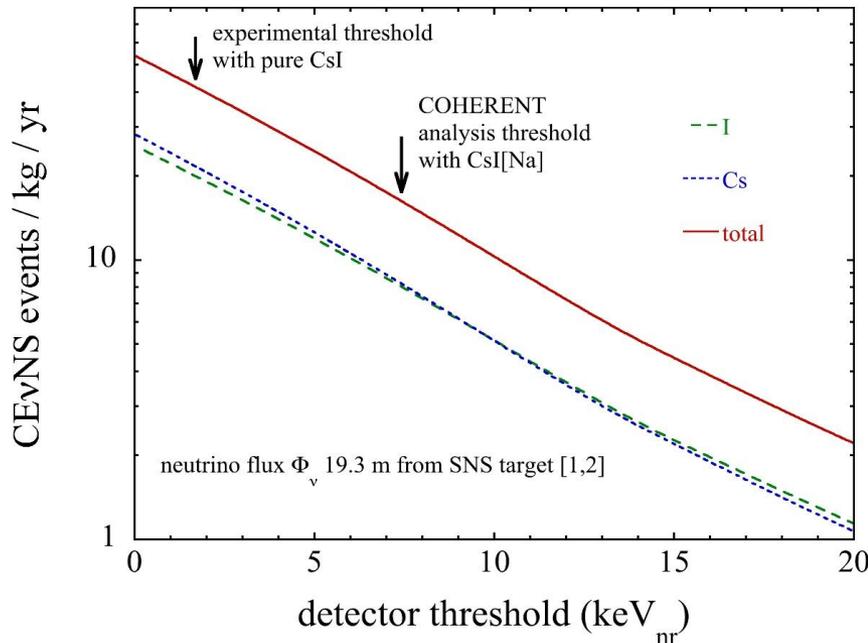
C^0SI

- QF that isn't a mystery
- High light yield
- Optimal QE
- Few-PE threshold

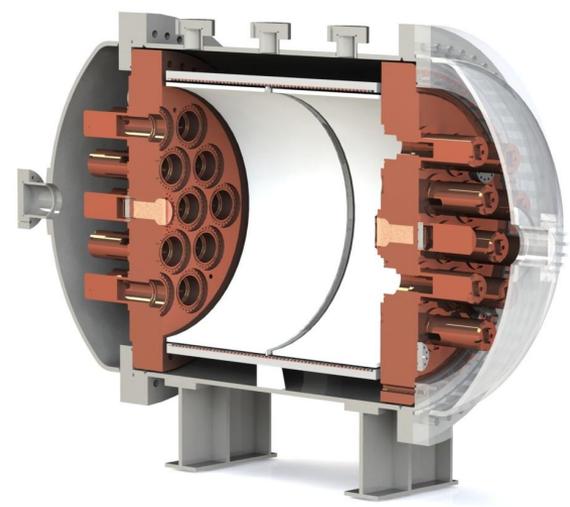
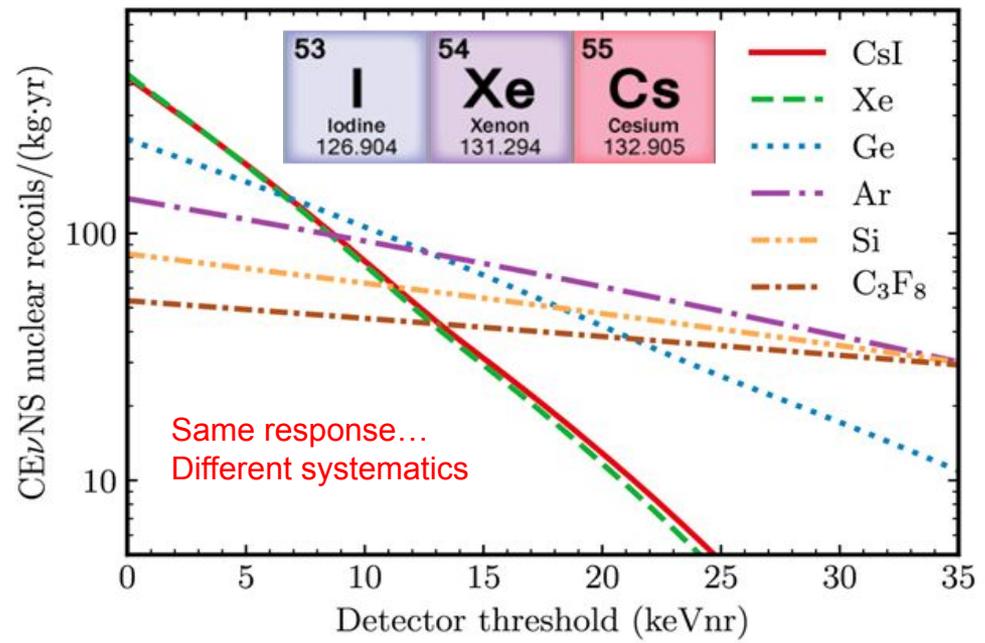
Moving from ~ 100 cts/yr



To $\sim 25k$ cts/yr



Sister detectors (G ν ESS)



- lack of cryogenics
- 1-2 e⁻ thresholds (~50 eV_{ee}) via EL
- Development already pushed by collaborators in NEXT
 - low bckgr
 - high pressures
- variety of nuclei \gg Ar, Kr, Xe

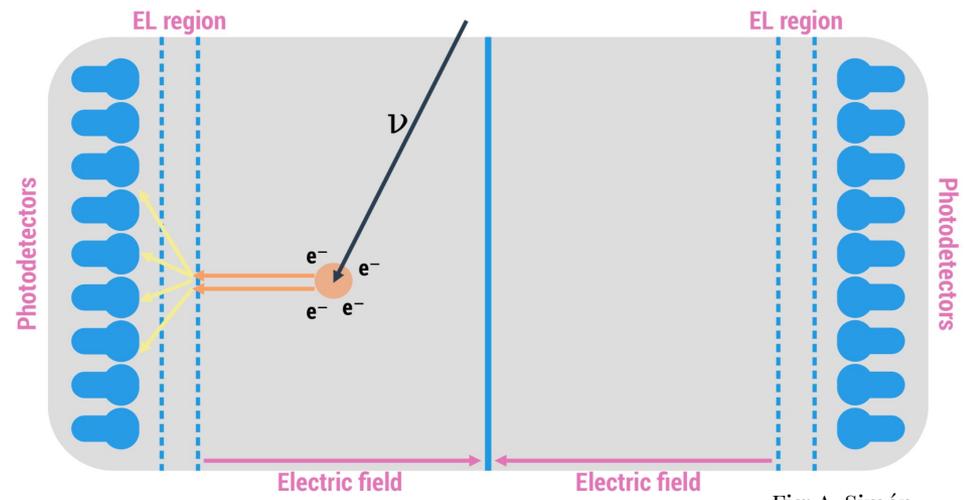


Fig: A. Simón

GavESS's gaseous prototype (GaP)

- opportunity to evaluate the technique in different conditions
 - multiple targets (Ar, Kr, Xe)
 - pressure up to 50 bar
- characterization of the low-energy response to nuclear recoils
 - **quenching factor measurements**
 - detection threshold

Currently characterizing ER signals with gaseous Ar at ~9.5 bar

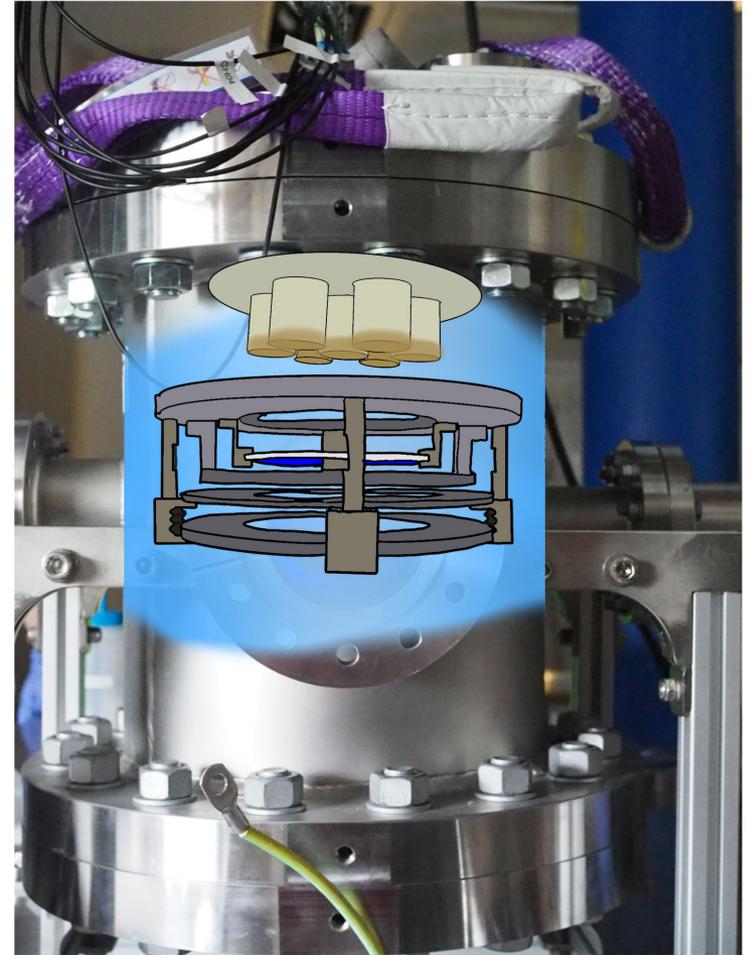
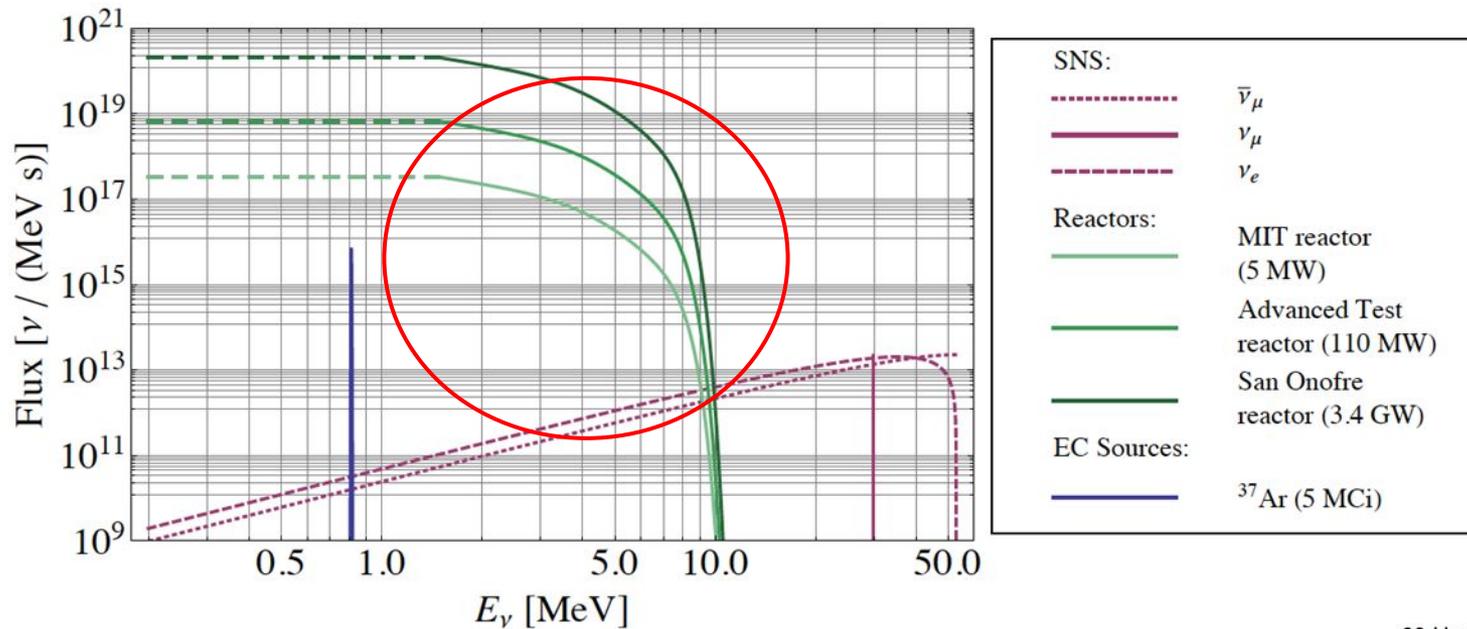


Fig: A. Simón

CE ν NS sources (reactors)

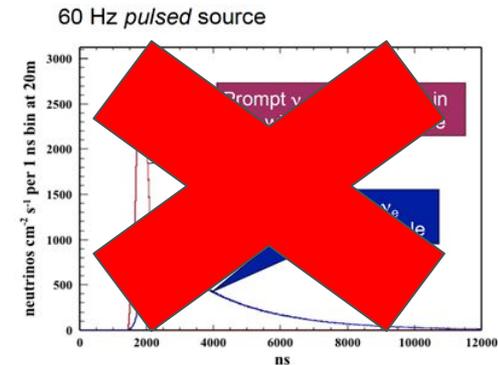


Enectali Figueroa-Feliciano / vMass 2013 / Milano

Low recoil energies... but high ν flux

No background subtraction (steady-state source)... but some locations have excellent background reduction

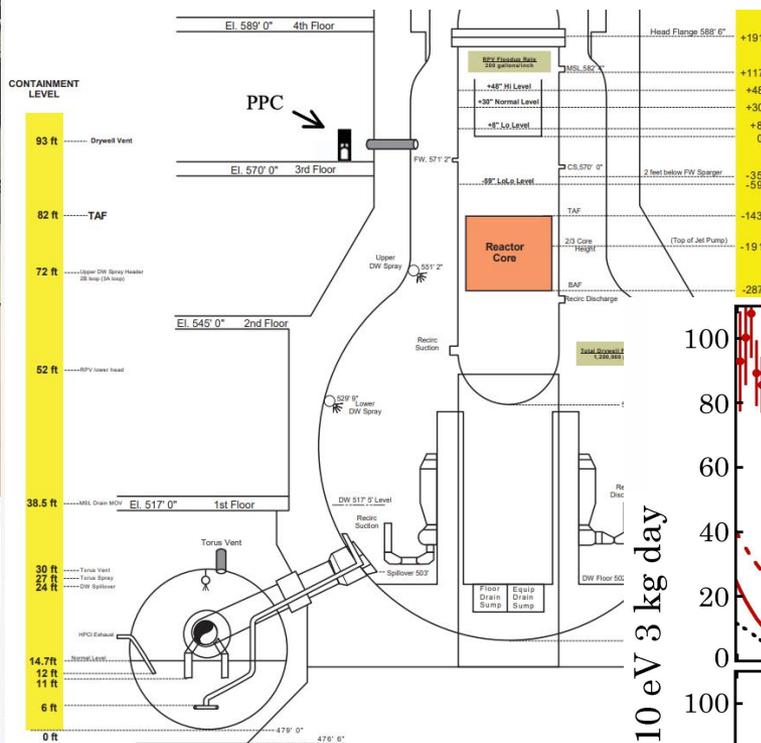
Spallation produces x200 the neutrons per ν



Cutting teeth at Dresden-II



- Mass
- Radiopurity
- Energy Resolution
- Low Threshold



- Only a 60x60 cm footprint
- Took 3 people one day to assemble
- $4.8E13 \nu/cm^2s$
- 10.4 m from core

Bayesian evidence ratio

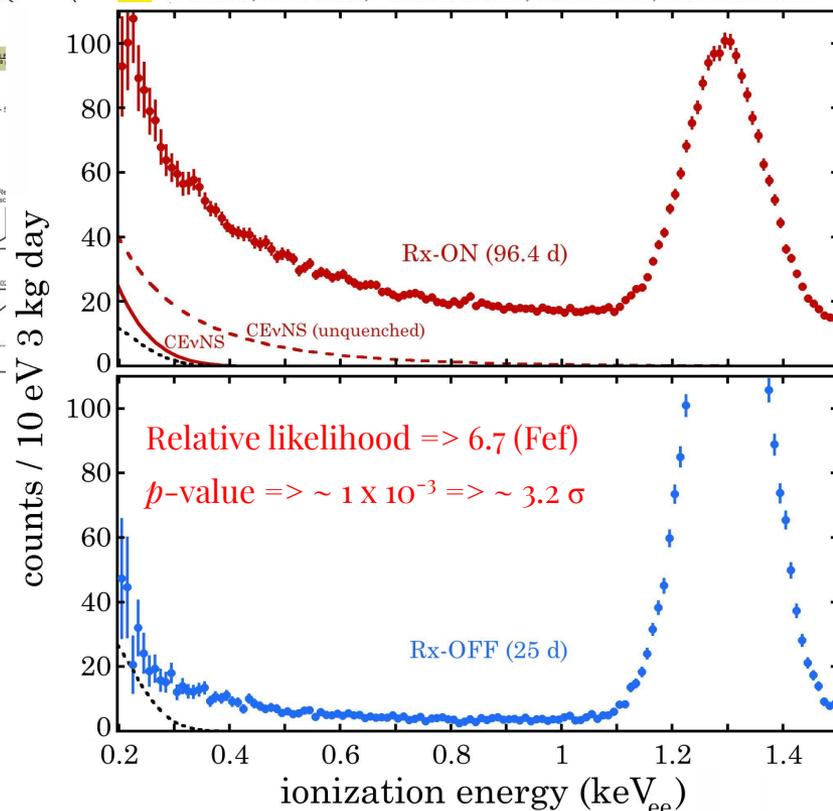
$\log_{10}(B_{10})$	B_{10}	Interpretation
0 to $\frac{1}{2}$	1 to 3.2	Weak
$\frac{1}{2}$ to 1	3.2 to 10	Moderate
1 to 2	10 to 100	Strong
>2	>100	Decisive

QF Model	B_{10} (MHVE)	B_{10} (KOP)
Fef	34.0	34.8
YBe	13.2	11.2
Lindhard	4.0	3.1

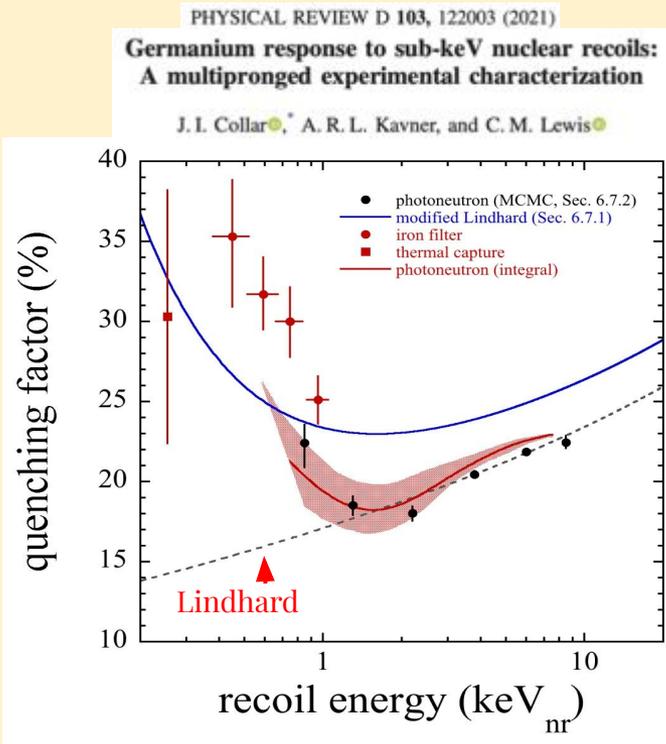
PHYSICAL REVIEW LETTERS **129**, 211802 (2022)

Measurement of Coherent Elastic Neutrino-Nucleus Scattering from Reactor Antineutrinos

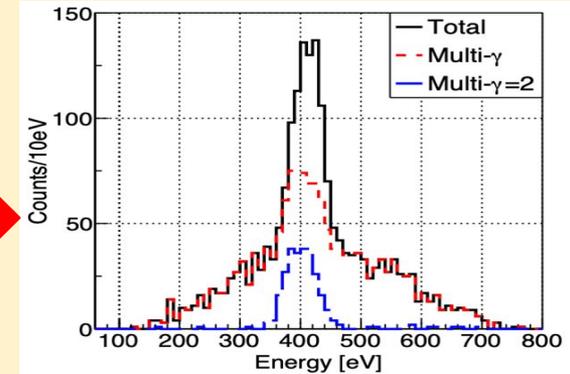
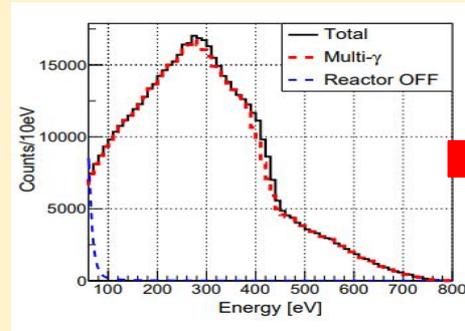
J. Colaresi,¹ J. I. Collar,^{2,*} T. W. Hossbach,³ C. M. Lewis,² and K. M. Yocum¹



A project of passion: Ge NR response



*comments on CONUS QF paper:
 arXiv:2203.00750



Calibration of nuclear recoils at the 100 eV scale using
 neutron capture

L. Thulliez,^a D. Lhuillier,^{a,*} F. Cappella,^b N. Casali,^b R. Cerulli,^{c,d} A. Challi,^a A. Chebboubi,^e
 E. Dumontell,^a A. Erhart,^f A. Giuliani,^g F. Gunsing,^h E. Jericha,^h M. Kaznacheeva,^f
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 L. Pattavina,^{d,i} D.V. Poda,^g R. Rogly,^g N. Schermer,^g O. Serot,^g G. Soum,^g L. Stodolsky,^g
 R. Strauss,^g M. Vignati,^{h,k} M. Vivier,^g V. Wagner,^g and A. Wex^g

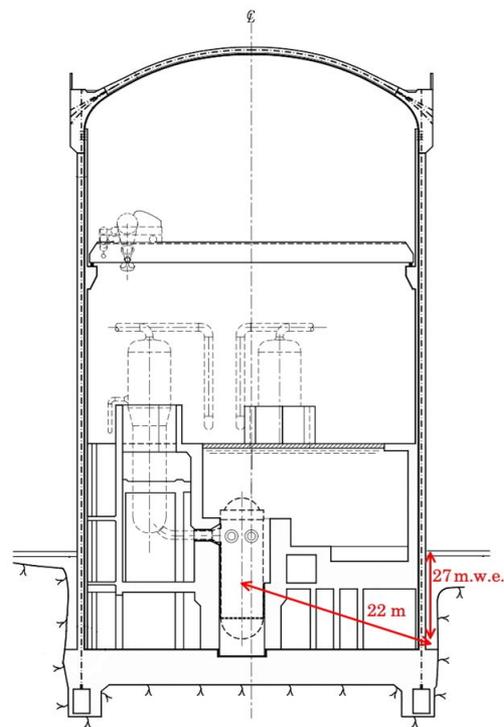
2021 JINST 16 P07032

- $0.4 \text{ keV}_{\text{nr}} \gg$ if Lindhard, it won't be visible
- A binary test for the form of the QF below 1 keV

CsI @ OSURR



The next pre-ESS step: Ringhals



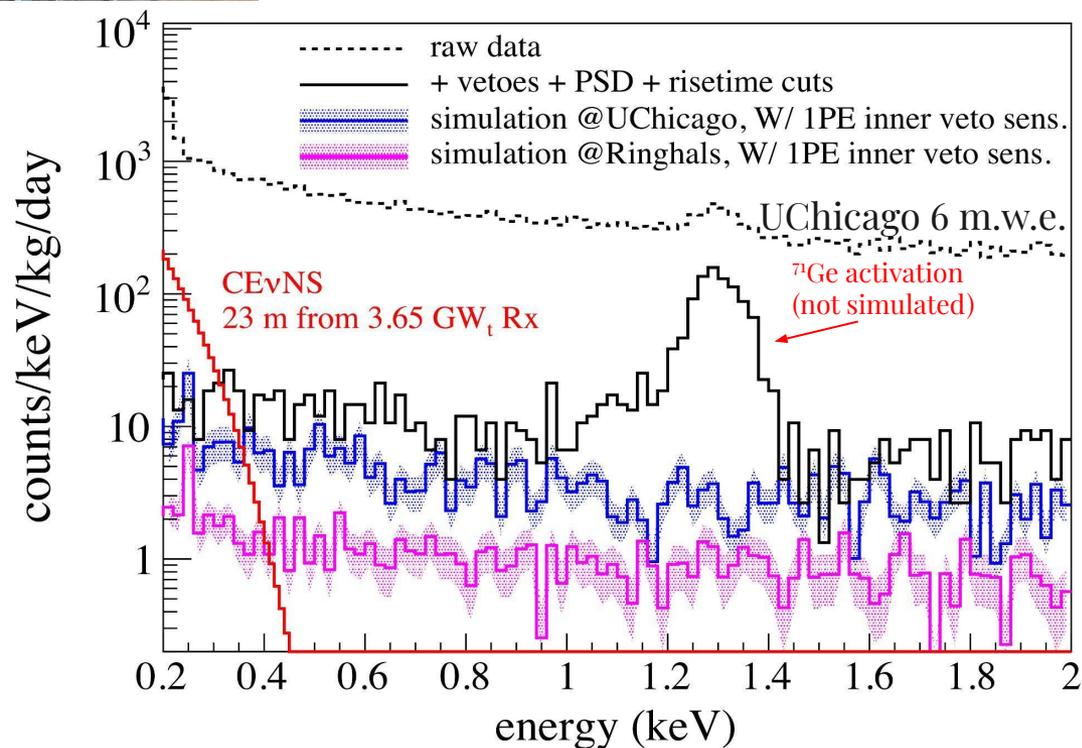
still with just a 6x60 cm footprint

- ^3He counter(s)
- NaI
- plastic scint
- overburden
- radon

30 m.w.e. in tendon gallery

Signal-to-background of ~ 40 expected
(present was $\frac{1}{4}$)

refurbished specs \gg significantly reduced
threshold



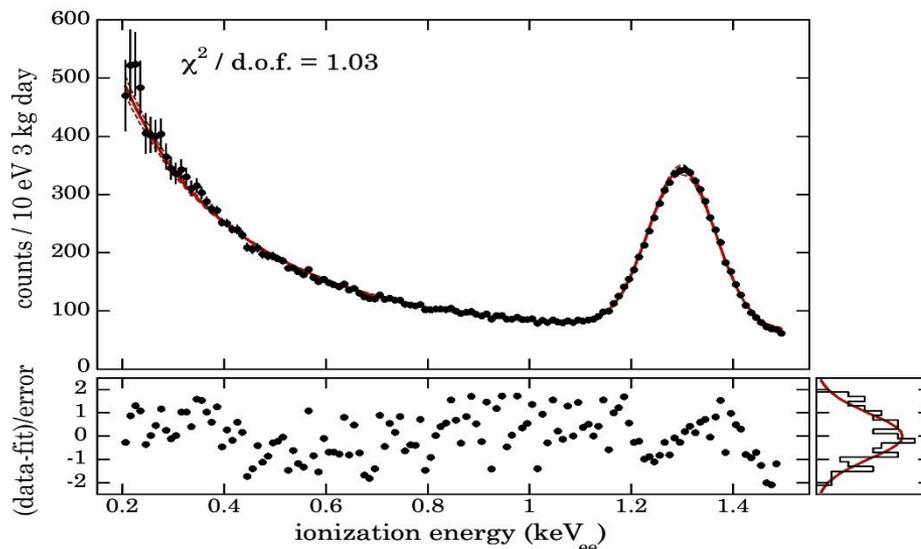
Thanks

Questions?



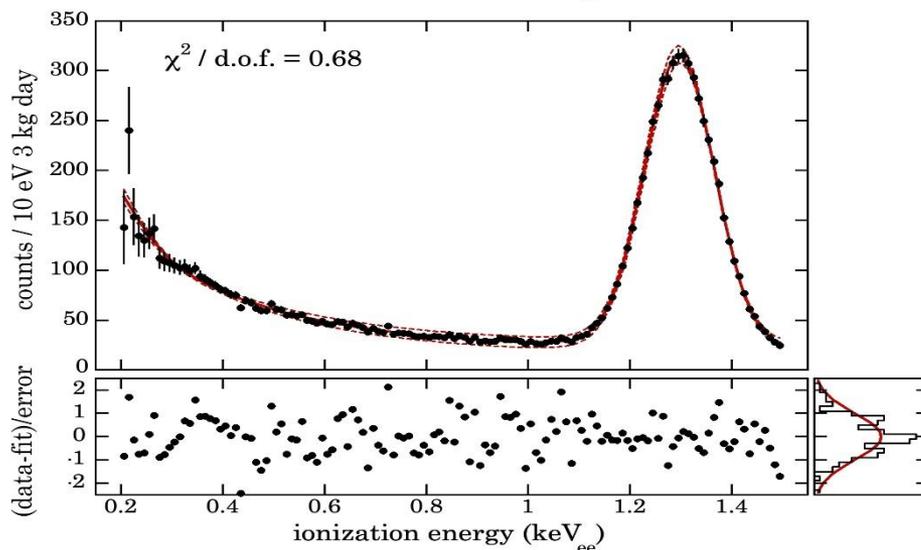
Extra: Background Model

No moderator



~1 keV wide ROI

1 inch moderator

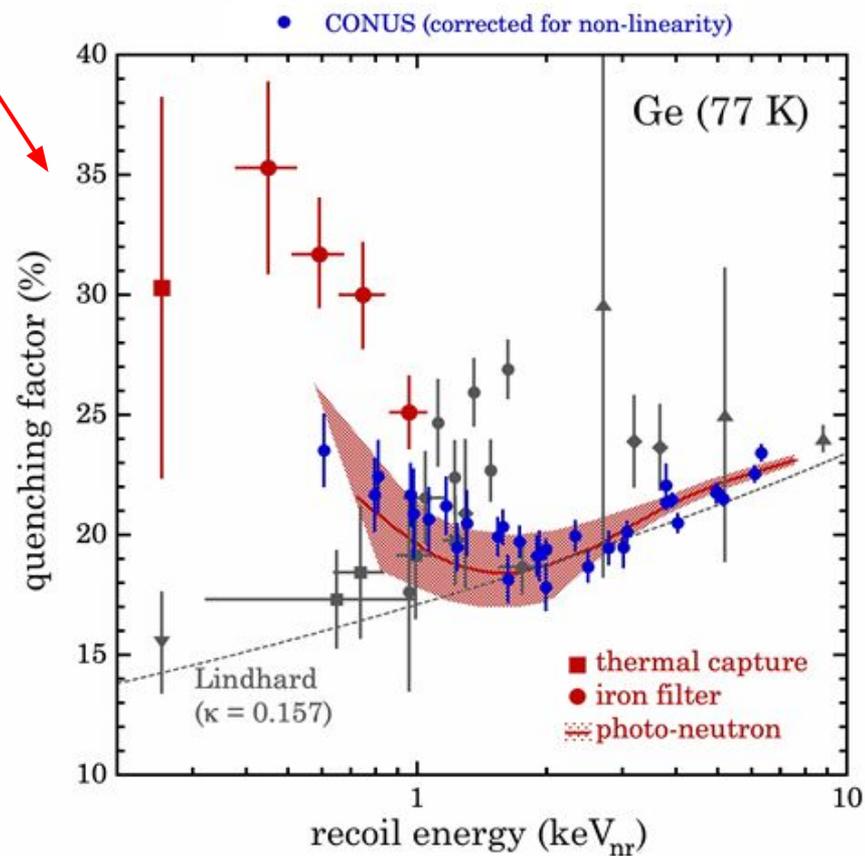
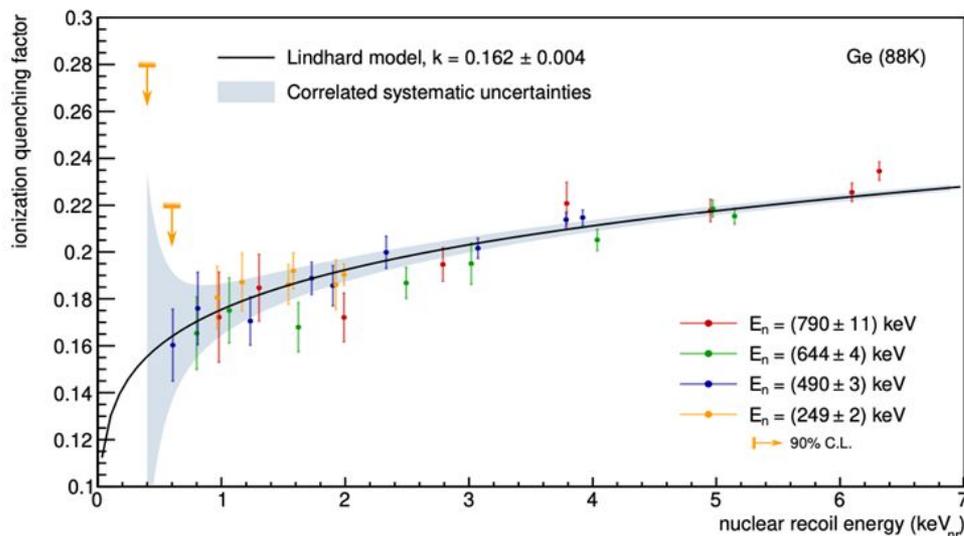


epithermal $\Rightarrow E_n^{-(1+\alpha)}$
where $\alpha \sim 0.2$
(best-match agrees)

no “shark tooth” peaks (fast)

simple falling exponential

Extra: QFs in Ge



two main points:

- improvised 2-point energy scale uncertainty \gg (Monte Carlo used to demonstrate)
- underestimated treatment (flat 10 eV) of ballistic deficit from DAQ \gg quoted numbers used to infer the correction (see right)
 - paper revision in response appears to soften ballistic deficit correction

*comments on CONUS sub-keV QF paper: arXiv:2203.00750