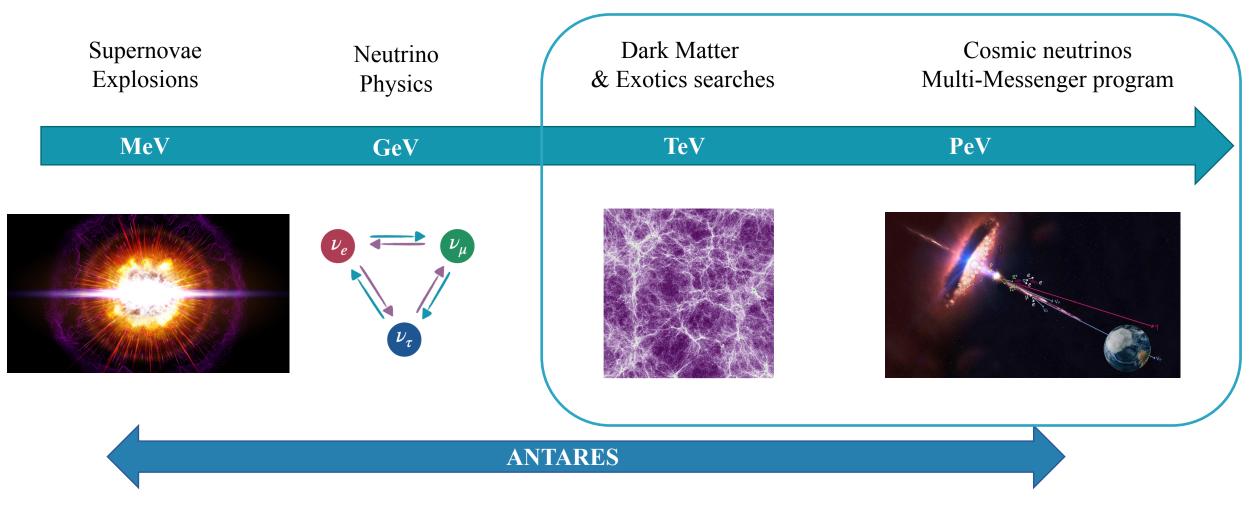
#### XX International Workshop on Neutrino Telescopes Venice, October 23<sup>rd</sup>-27<sup>th</sup> 2023

#### **Overview of ANTARES, the first Mediterranean Sea telescope**

#### Dr. Chiara Poirè, on behalf of the ANTARES collaboration



#### **Physics motivations**

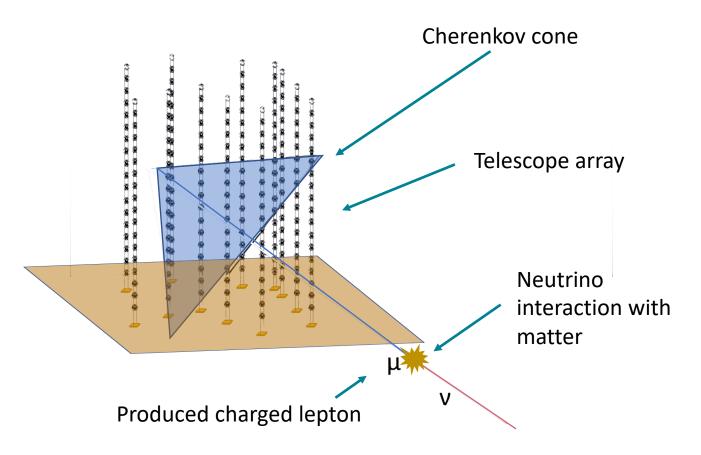


+ Not only physics: Environmental research (Earth & Sea Science).

#### **Characteristics of neutrino telescopes**

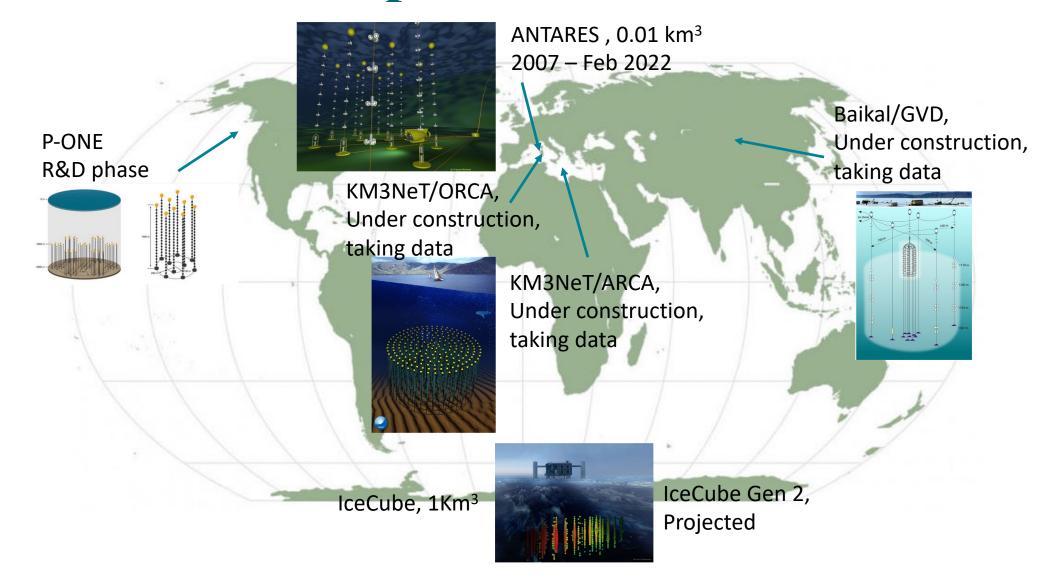
- <u>Large volume</u>: offers a large number of free target nucleons for neutrino interactions.
- <u>Great depth</u>: provides shielding against secondary particle produced by CRs.
- <u>Transparent medium</u>: allows propagation of Cherenkov photons emitted by relativistic charged particles produced by neutrino interactions

Sensitive from few GeV to few PeV



Time and position of photons to reconstruct the Cherenkov cone  $\rightarrow$  neutrino incoming direction

#### Neutrino telescopes around the world

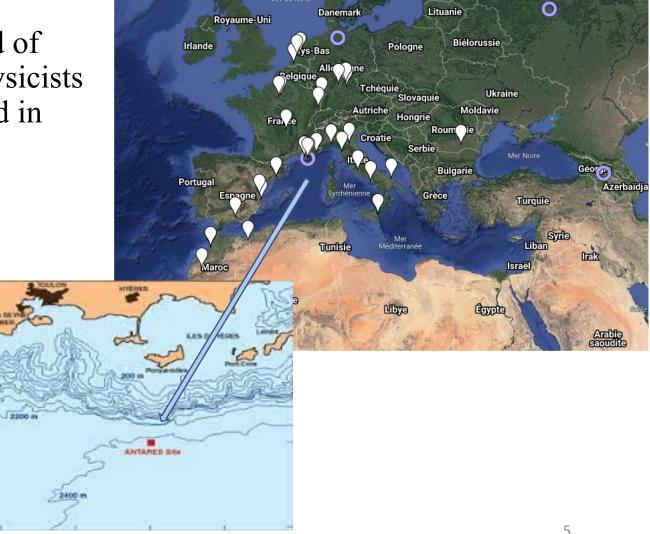


#### **The ANTARES Collaboration**

The ANTARES collaboration is composed of around 150 engineers, technicians and physicists from different institutes principally located in Europe.

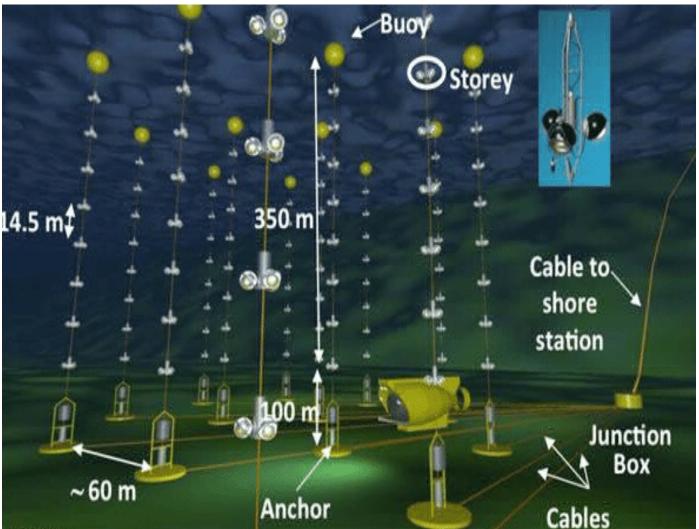
The collaboration was born in 1997.

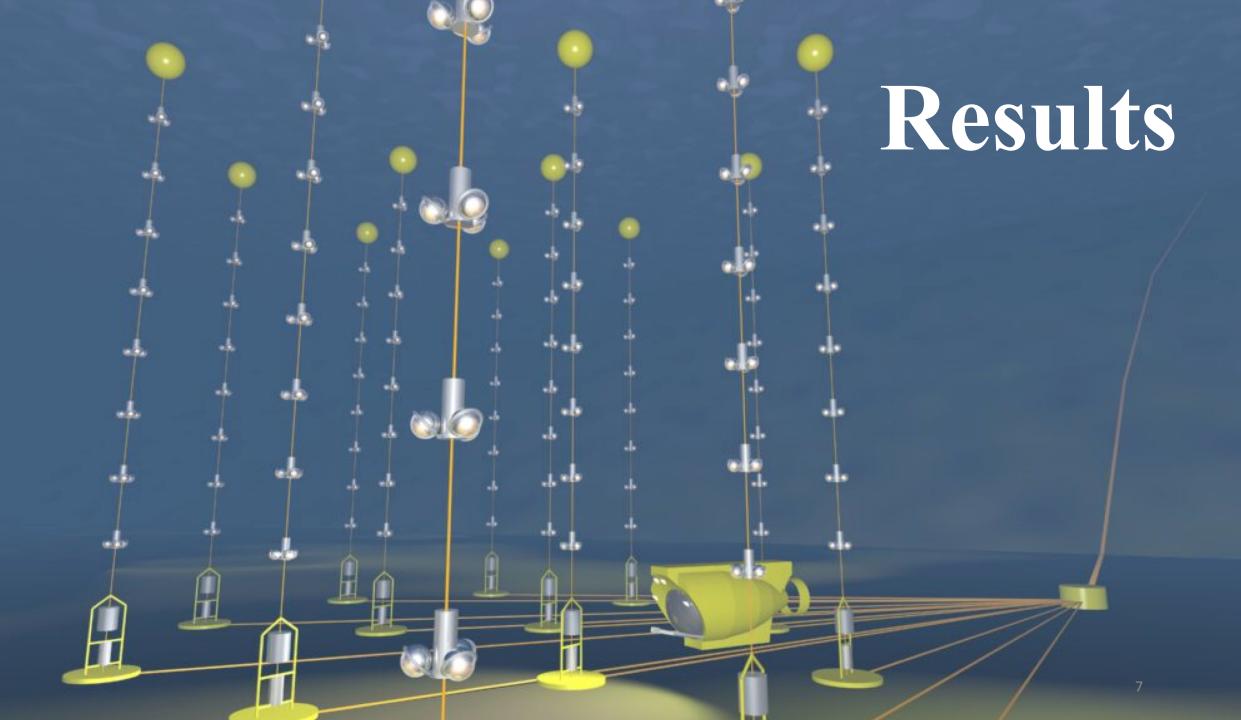




### **The ANTARES neutrino telescope**

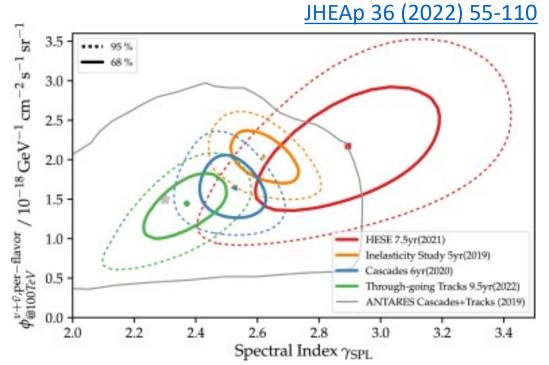
- Toulon, France
- ~2500 m depth
- Data taking: 2007-> February 2022 🤍 Array infrastructure:
- 12 lines
- 25 storeys (3 OM x storey)
- ~900 PMTs
- Volume = 0.01 km<sup>3</sup>
- ~60 m  $\rightarrow$  > 10000 neutrinos recorded
- $\rightarrow$  ~100 publications in the field of neutrino physics and astrophysics

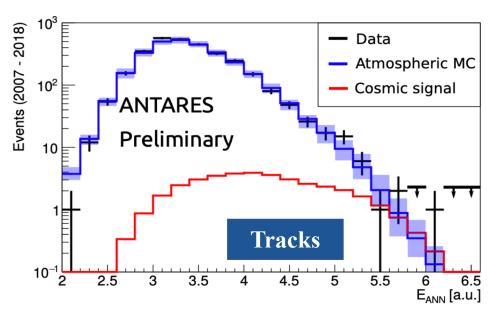




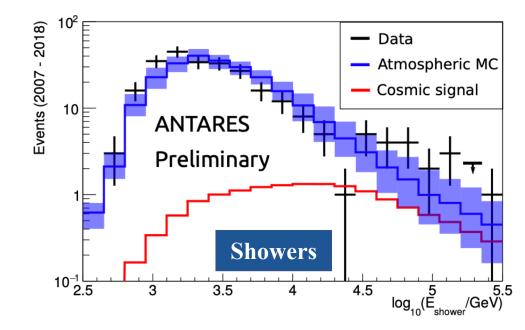
#### **Cosmic diffuse emission**

- IceCube has confirmed the existence of high-energy cosmic neutrinos in 2013
- ANTARES using data from <u>2007-2018</u> (3330 days) has observed 1.8 σ excess → 50 events observed while 36.1 ± 8.7 expected from background (tracks + showers).
- <u>New results are in preparation</u>.



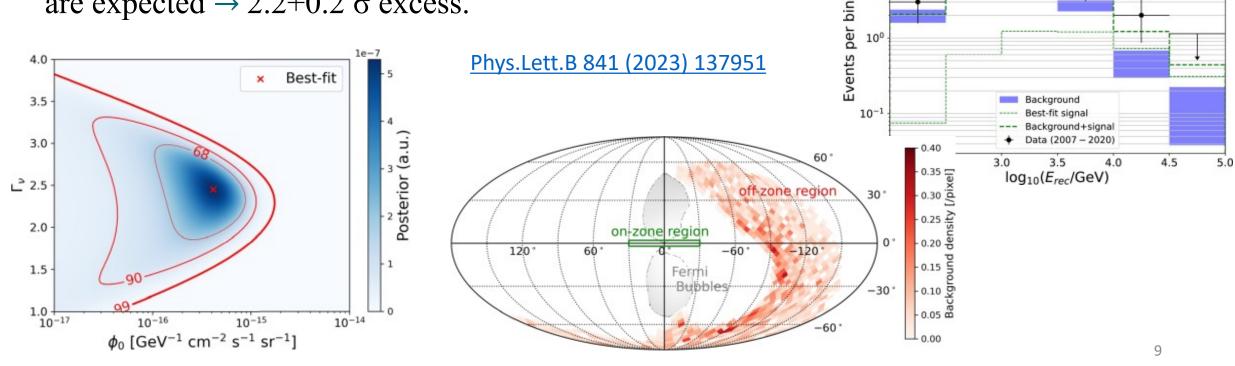


PoS ICRC2021 (891)



#### Galactic diffuse emission

- This year two announcements for Galactic component contributing to the cosmic neutrino diffuse flux: ANTARES (~2  $\sigma$ ) and IceCube (~4.5  $\sigma$ ).
- ANTARES ON/OFF analysis at E>1 TeV detects 21+13 track+shower events while  $11.7 \pm 0.6 + 11.2 \pm 0.9$  events are expected  $\rightarrow 2.2+0.2 \sigma$  excess.



**Tracks** 

4.5

Showers

5.0

Background Best-fit signal

Background+signal

Data (2007 - 2020)

4.0

3.5

log10(Erec/GeV)

101

 $10^{-1}$ 

101

2.0

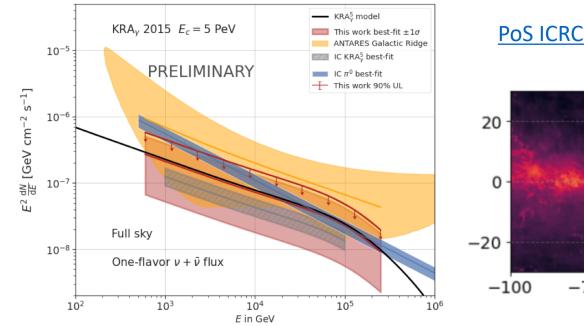
2.5

3.0

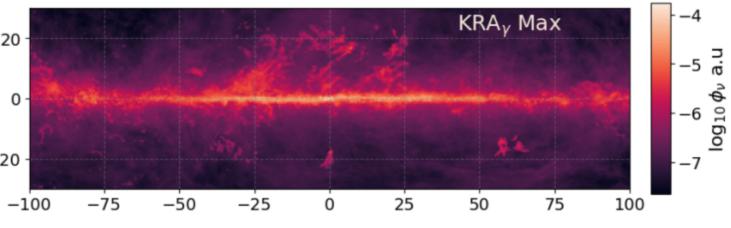
Events per bin 00

#### Galactic diffuse emission

- ANTARES template analysis using the most recent KRA $\gamma$  models shows a 1.5-1.8  $\sigma$  excess.
- This is the same method used by IceCube (~4.5  $\sigma$ ).



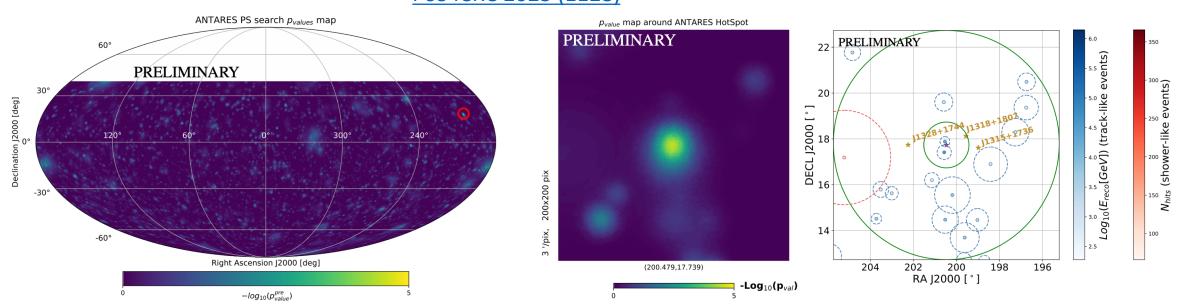
#### PoS ICRC 2023 (1084)



#### **Search for neutrino sources**

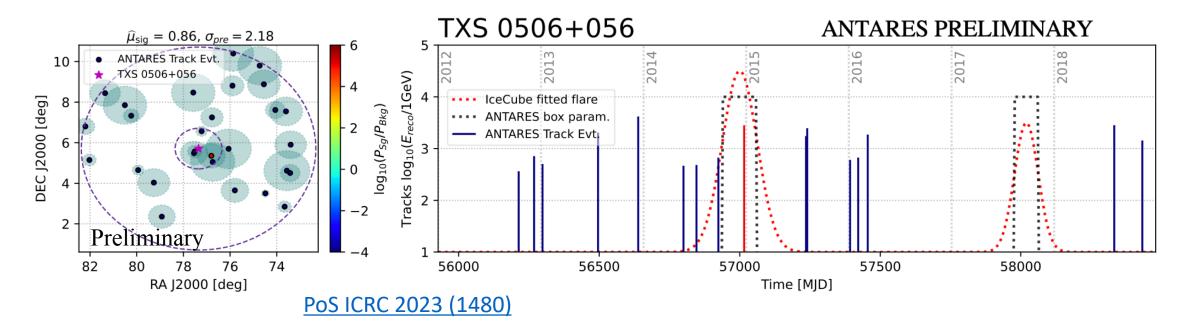
- Several types of searches performed: time-integrated, all-sky search and candidate list search
- All-sky time-integrated search: the most significant spot,  $(RA, dec) = (200.5, 17.7)^\circ$ , has a pre(post)-trial significance of 4.0(1.2)  $\sigma$ , with <u>no evident association</u> (closest source is 1° away).
- Candidate list search: 163 sources were tested to reduce the trial factor. <u>No significant excess</u> observed but some sources show interesting upper fluctuations.
  → Best candidates are MG3 J225517+2409 and 3C403 with 3.4(1.7) σ pre(post)-trial

significance. PoS ICRC 2023 (1128)



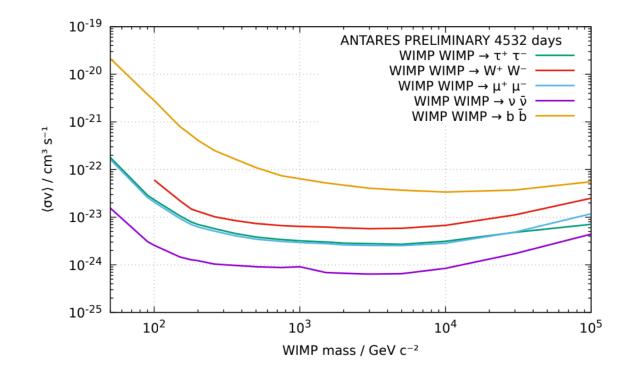
#### **Time-dependent Source Searches**

- Time-dependent search: using the temporal information from external observatories
- Tested potential neutrino flares by IceCube [ApJL 920 L45 (2021)] with ANTARES data.
- Out of 36 sources, 4 have fitted signal (~2  $\sigma$  upper-fluctuations).
- TXS 0506+056 has 1 event compatible with the "orphan" neutrino flare (2014-2015). Overall significance comparable to time-integrated search, with 3 times less signal.
- Searches based on Electro-Magnetic observatories ongoing.



### Dark matter searches from Galactic centre

- Limits on  $< \sigma \nu >$  from WIMP annihilation in galactic center
- ANTARES 2007-2022, 4532 days
- 11850 tracks, 235 showers
- Assuming NFW halo profile with 100% branching ratio for each annihilation channel



PoS ICRC 2023 (1375)

## **Other analysis**

Full list of publications: https://antares.in2p3.fr/publications/

- Exotic physics: search for magnetic monopoles [JHEAp 34 (2022)], Limits on the nuclearite flux [JCAP01(2023)012]
- Neutrino physics: Neutrino oscillations [JHEP 06 (2019) 113], Non-standard Interactions [JHEP 07 (2022) 048]
- Multi-messenger searches: with HAWC through AMON [<u>ApJ 944 166 (2023)</u>], GRBs [JCAP03(2021)092], GWs [JCAP04(2023)004]
- Catalog analysis [<u>ApJ 911 48 (2021)</u>]
- Combined analyses:
  - Dark matter with IceCube [Phys. Rev. D 102 (2020)]
  - Sources with IceCube [<u>ApJ 892 92 (2020)</u>]
  - Correlations of neutrinos (with IceCube) with UHE Cosmic rays with AUGER [<u>ApJ 934 164 (2022)</u>]
- Multi-disciplinary facility: Sperm whale [Sci Rep . 2017 Apr 12;7:45517]

#### What is next?

- The next generation neutrino telescope KM3NeT is under construction and is already taking data
- KM3NeT has an improved design, with two detectors (ARCA & ORCA) sensitive from MeV to PeV energies.
- Expected to become complete by 2028







#### Conclusions

- ANTARES has taken data from 2007 to early 2022
- More than 100 papers have been published in peer-review journals in the field of neutrino searches
- ANTARES was part of the multi-messenger world collaborating with other experiments
- Final legacy analysis are ongoing and will be finalized in 2024



# Thanks for the attention!

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