



Latest Results from the CUORE Experiment

Alberto Ressa on behalf of the CUORE Collaboration



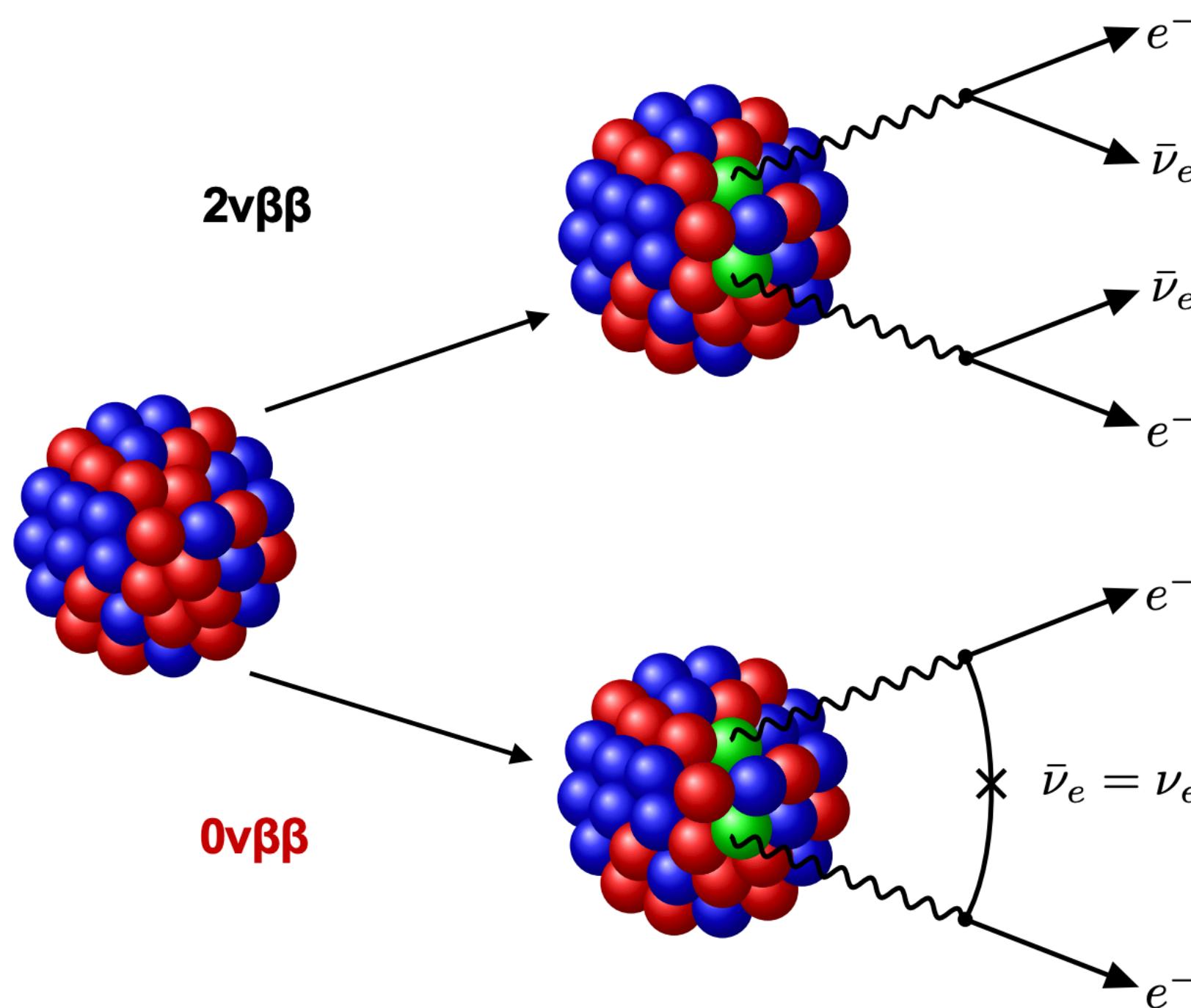
Istituto Nazionale di Fisica Nucleare

NeuTel 2023, Venezia



SAPIENZA
UNIVERSITÀ DI ROMA

Double β Decay



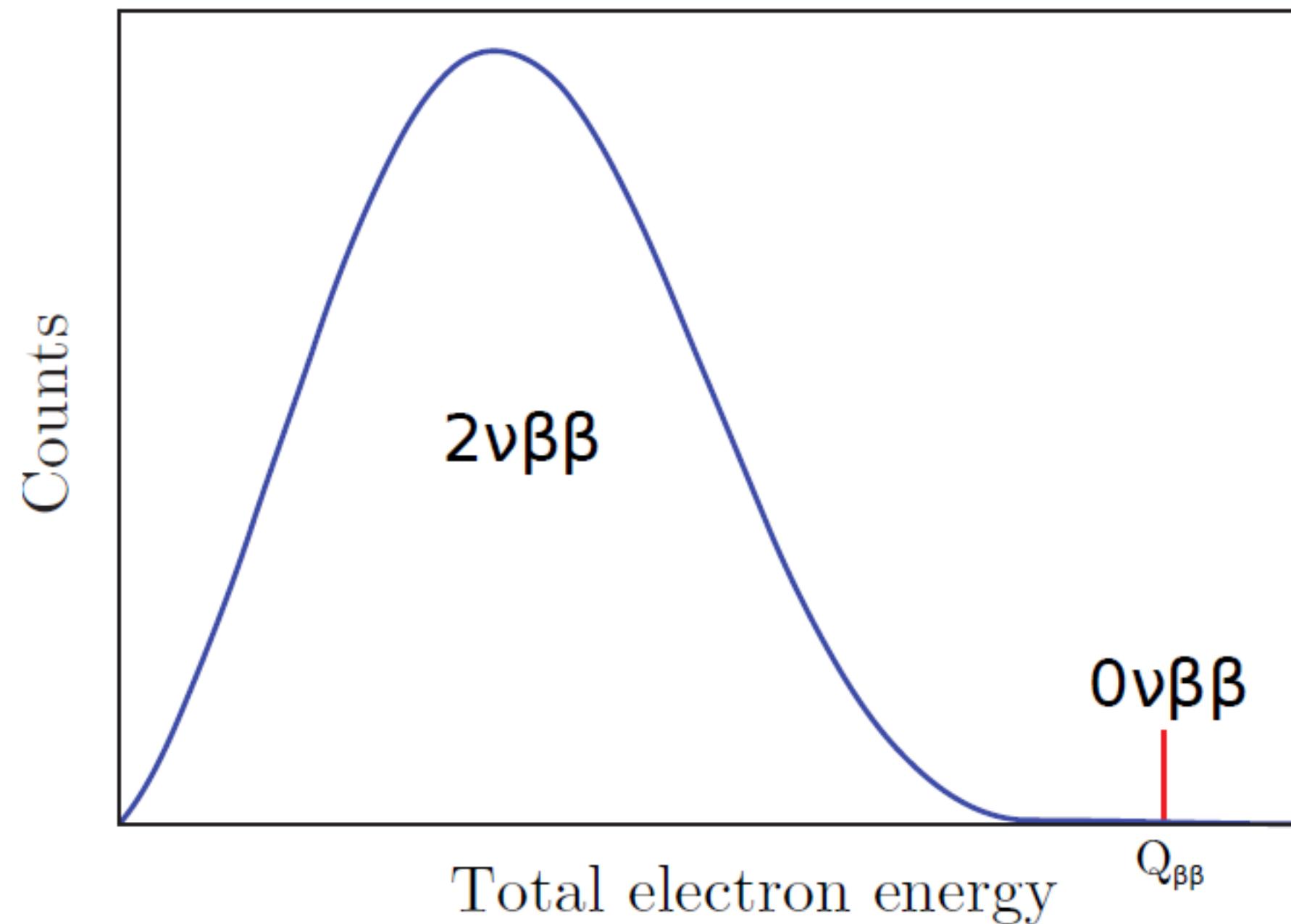
Standard Model allowed double β decay: $2\nu\beta\beta$

- **Observed** in 11 even-even nuclei in which single β decay is energetically forbidden
- $T_{1/2} \sim 10^{18} - 10^{24}$ years

Neutrinoless double β decay: $0\nu\beta\beta$

- **Neutrino Nature:** possible only if neutrino is a Majorana particle (coincides with its own antiparticle)
- **Total Lepton number violated of 2 units:** an ingredient to solve matter-antimatter asymmetry

Double β Decay



Standard Model allowed double β decay: $2\nu\beta\beta$

→ Continuous spectrum ending at the isotope Q-value

Neutrinoless double β decay: $0\nu\beta\beta$

→ Mono-energetic peak at the isotope Q-value
(a simple and clear experimental signature)

Experimental search for $0\nu\beta\beta$

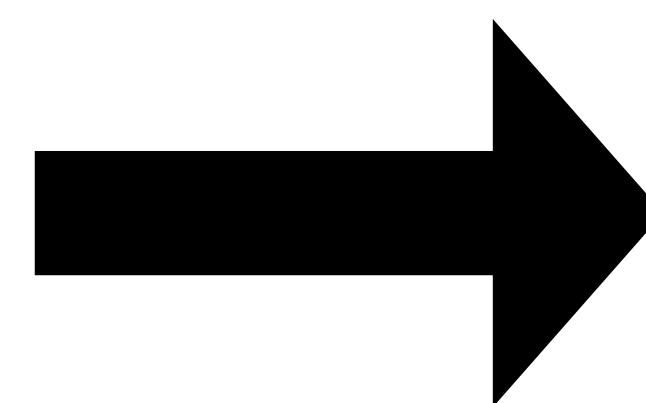
$$S_{0\nu} \propto \epsilon \sqrt{\frac{MT}{B\Delta}}$$

High Exposure

High Efficiency

Low Background

Good Energy Resolution

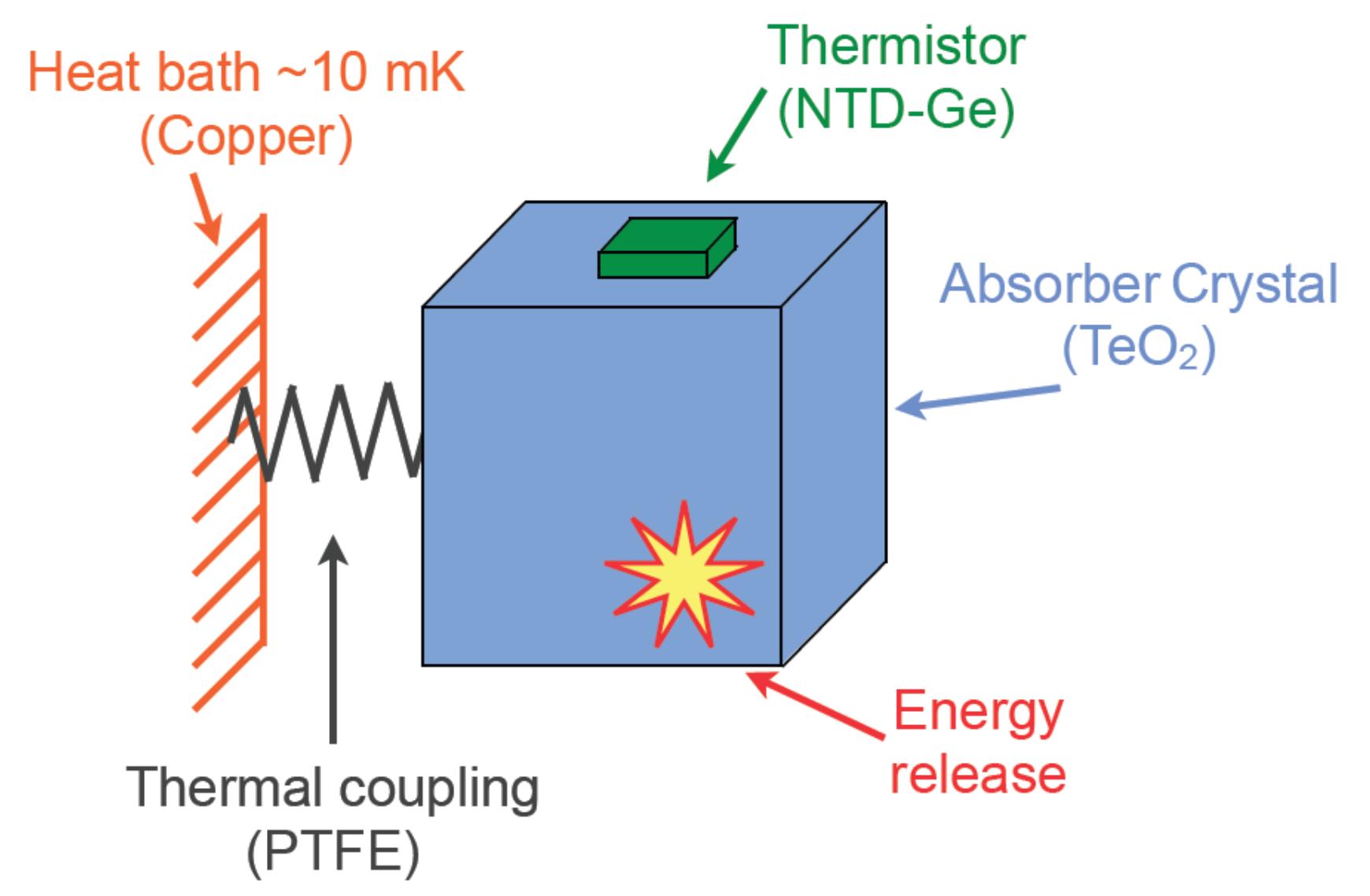
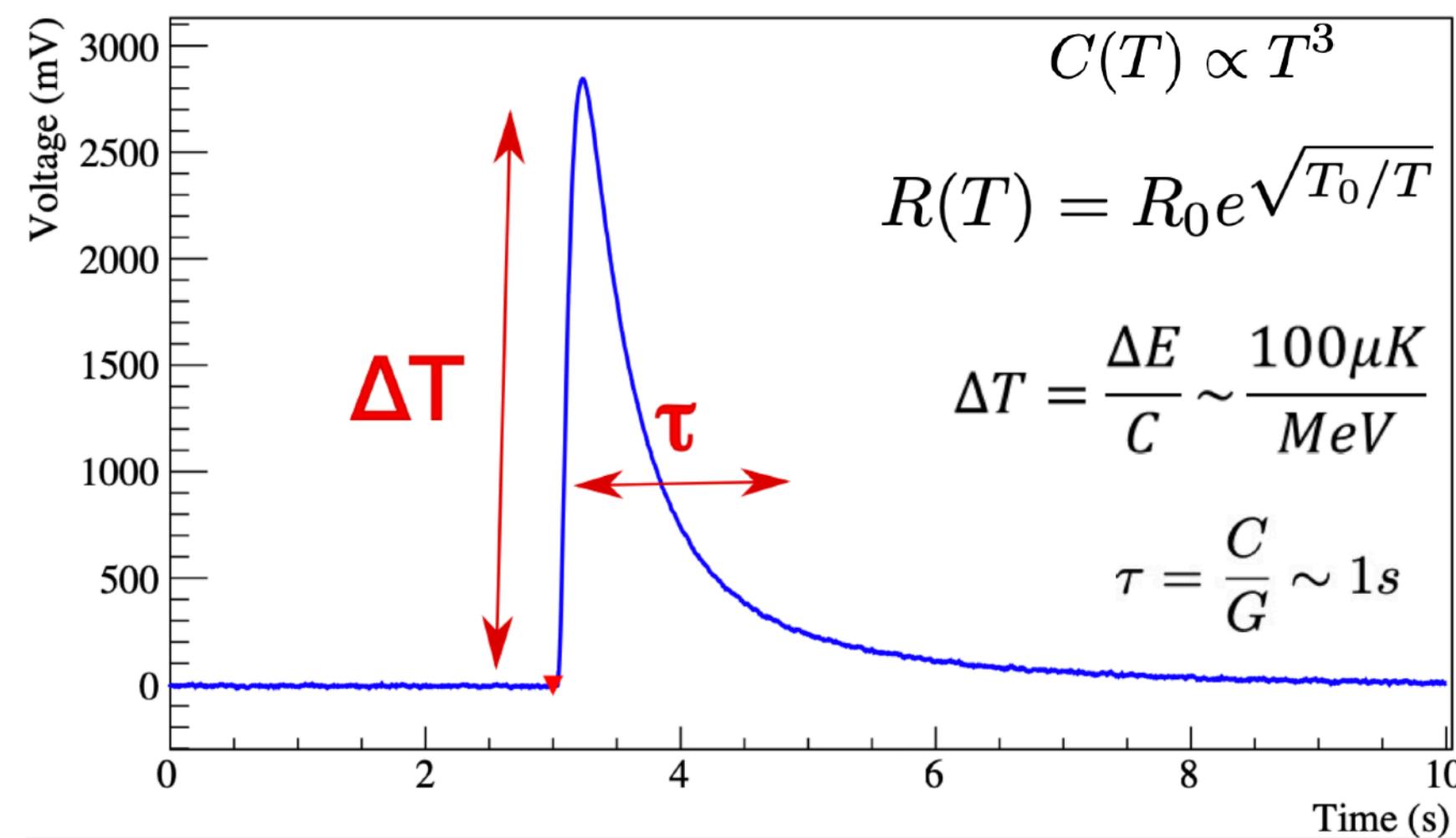


Cryogenic Calorimeters:

- A mature technology as demonstrated by several experiments in the $0\nu\beta\beta$ field (CUORE, CUPID-0, CUPID-Mo, AMoRE)
- A mature technology able to explore competitive regions for the $0\nu\beta\beta$ parameter space

Cryogenic Calorimeters

1. Interacting particles deposit energy in the crystal
2. The energy release heats up the crystal via thermal phonons
3. The temperature increase is converted into an electric signal by a cryogenic sensor
 - e.g. thermistors: resistance changes exponentially with temperature

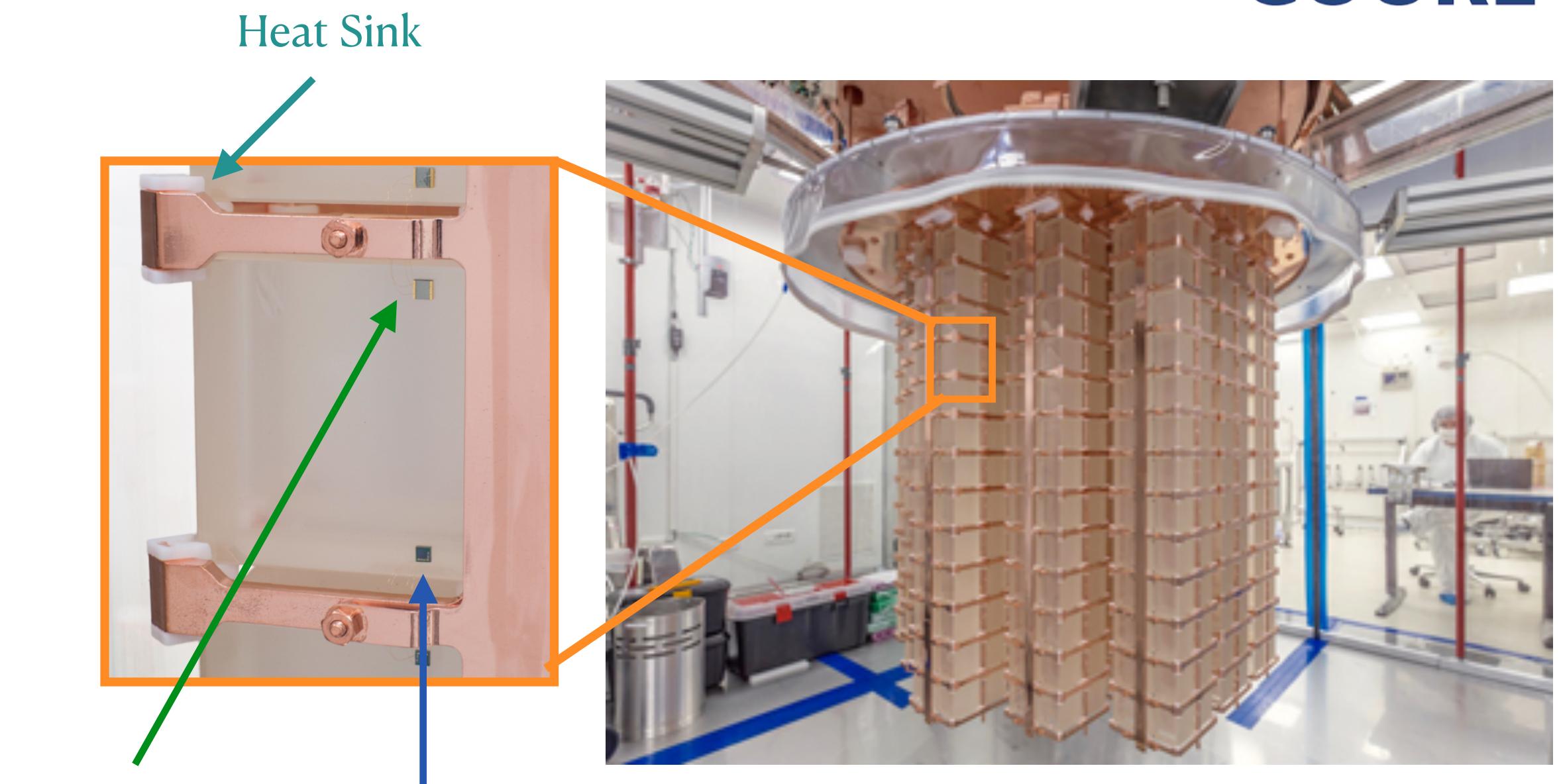


- ➔ Cryogenic temperatures (about 10 mK) make possible to turn the energy deposit into a readable temperature increase
- ➔ Slow pulses:
 - Able to reconstruct pulse shape
 - Higher pile-up (limited because of low event rate)

CUORE Experiment

Cryogenic Underground Observatory for Rare Events:

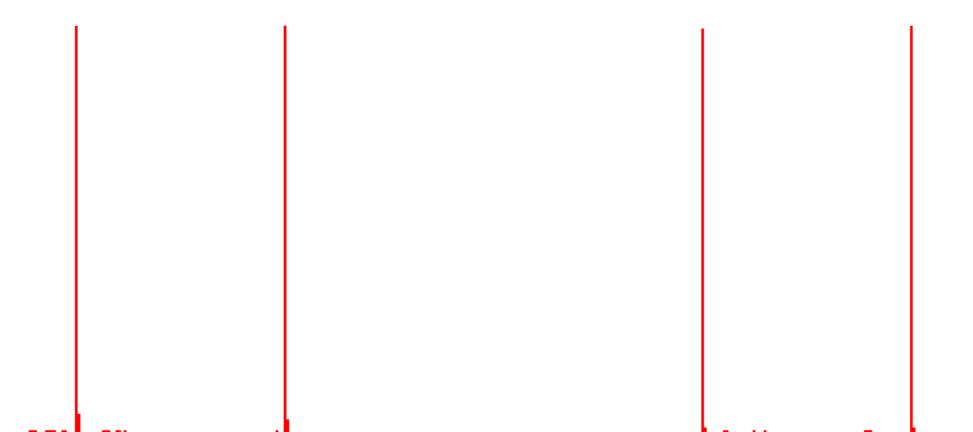
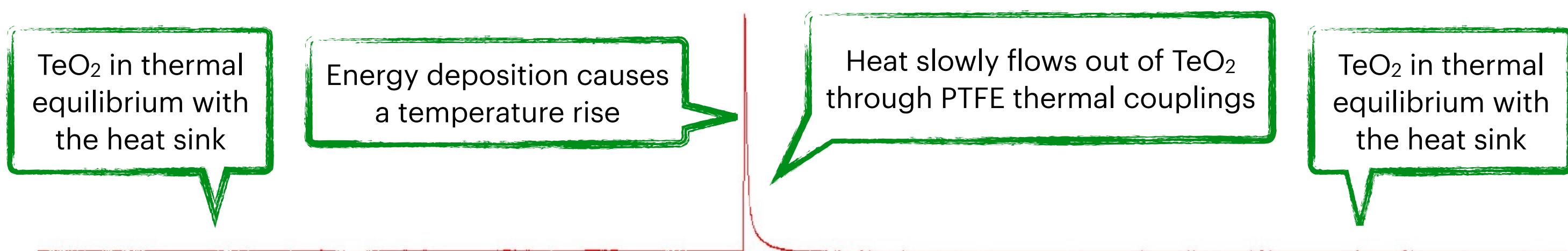
- Searching for $0\nu\beta\beta$ of ^{130}Te at 2528 keV
- Located at the Laboratori Nazionali del Gran Sasso in Italy
- 988 natural TeO_2 $5 \times 5 \times 5 \text{ cm}^3$ cubic crystals arranged in 19 towers of 13 floors
- 742 kg of TeO_2 (i.e. 206 kg of ^{130}Te)
- Taking data for over 4 years with >90% duty cycle



Heat Sink
Thermistors (NTD-Ge)
Silicon Heater

Silicon heaters provide standardized heat pulses for thermal gain correction

Example of CUORE data stream

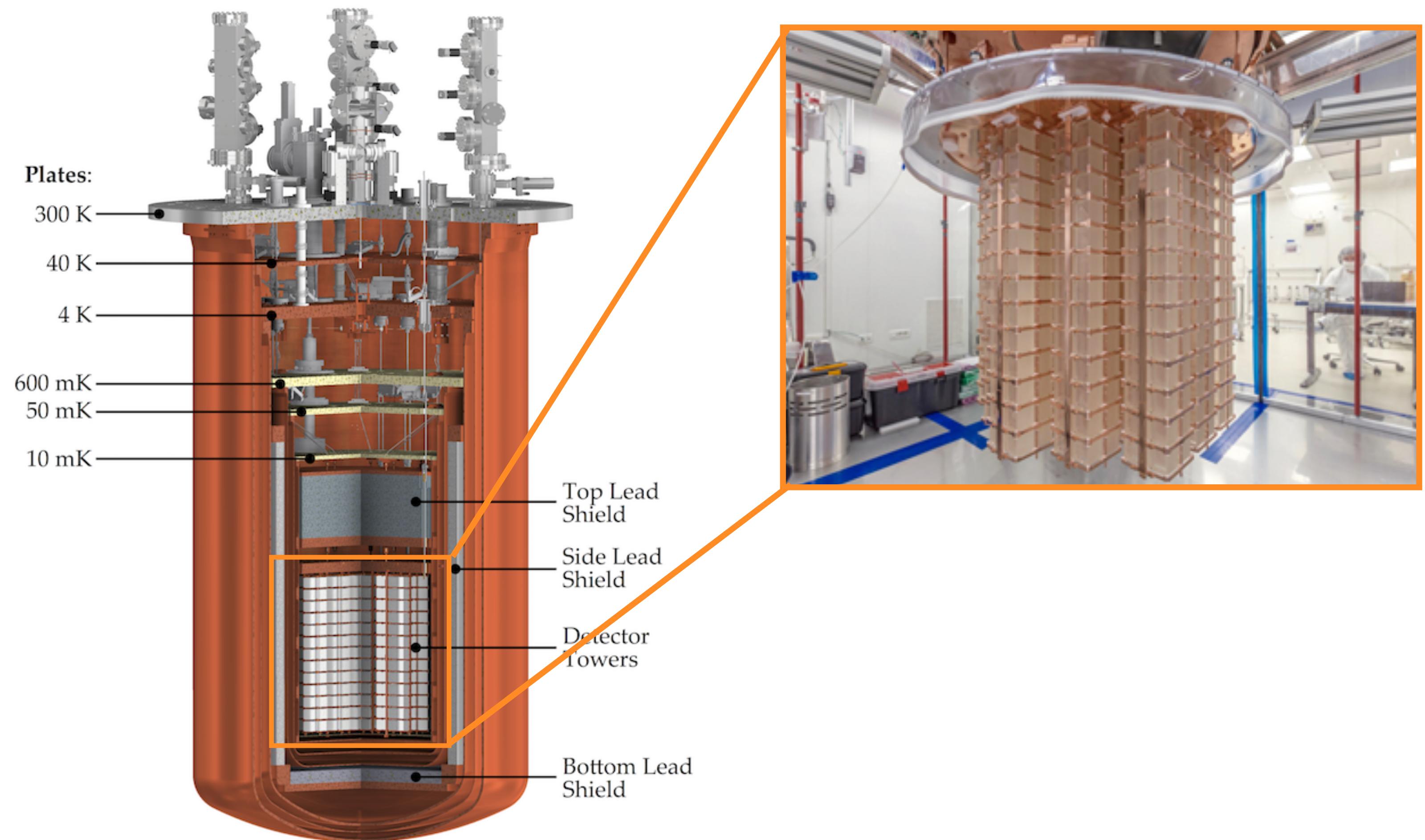


CUORE Experiment



Cryogenic Underground Observatory
for Rare Events:

- Operated in a world leading dilution refrigerator in terms of power and size
- Equipped with 4 Pulse Tubes for cooling to 4K
- Nested co-axial copper vessels at decreasing temperatures
- 15 tons cooled below 4 K and 3 tons below 50 mK
- TeO_2 crystals kept at 11-15 mK

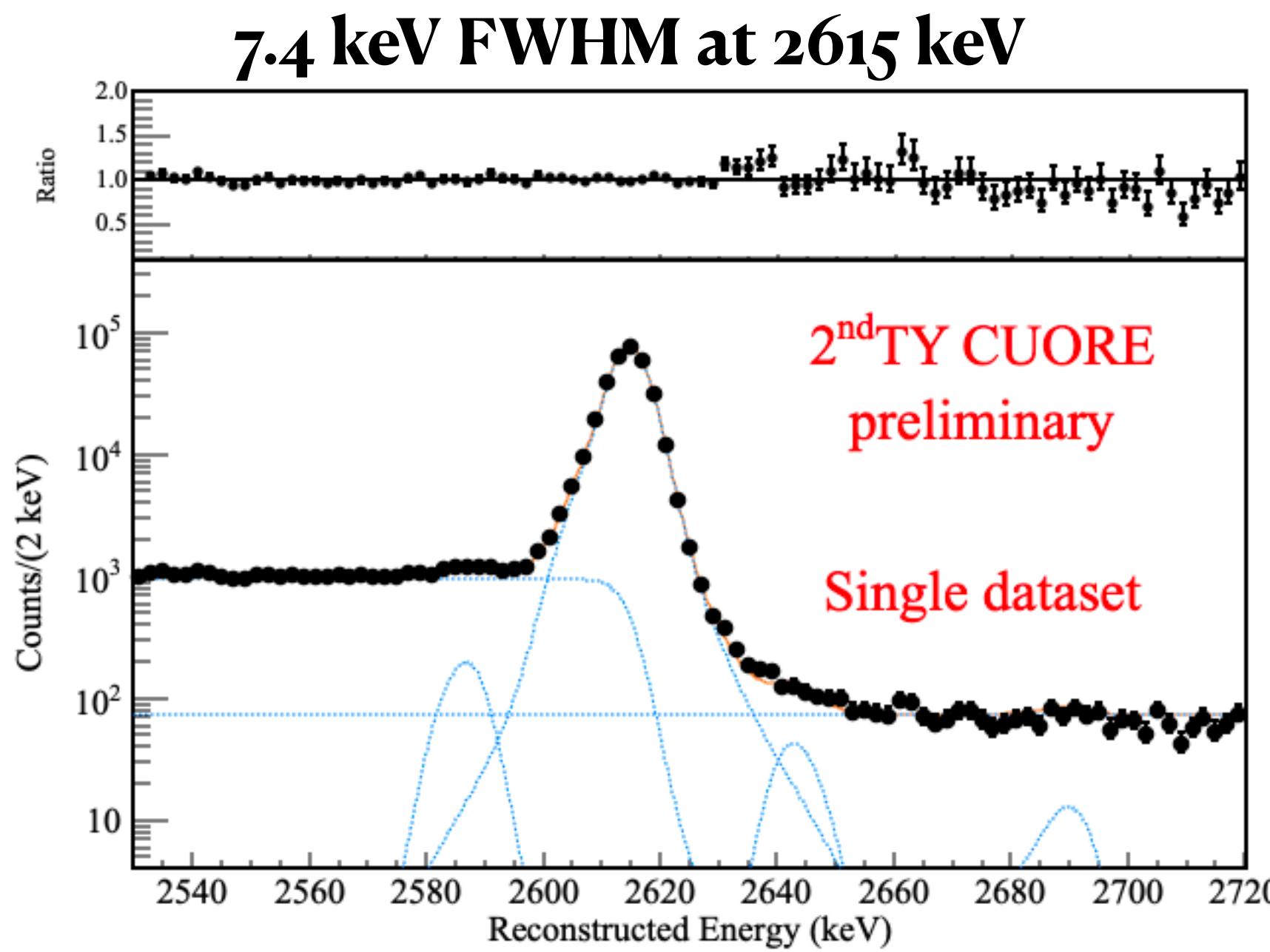


$0\nu\beta\beta$ with CUORE

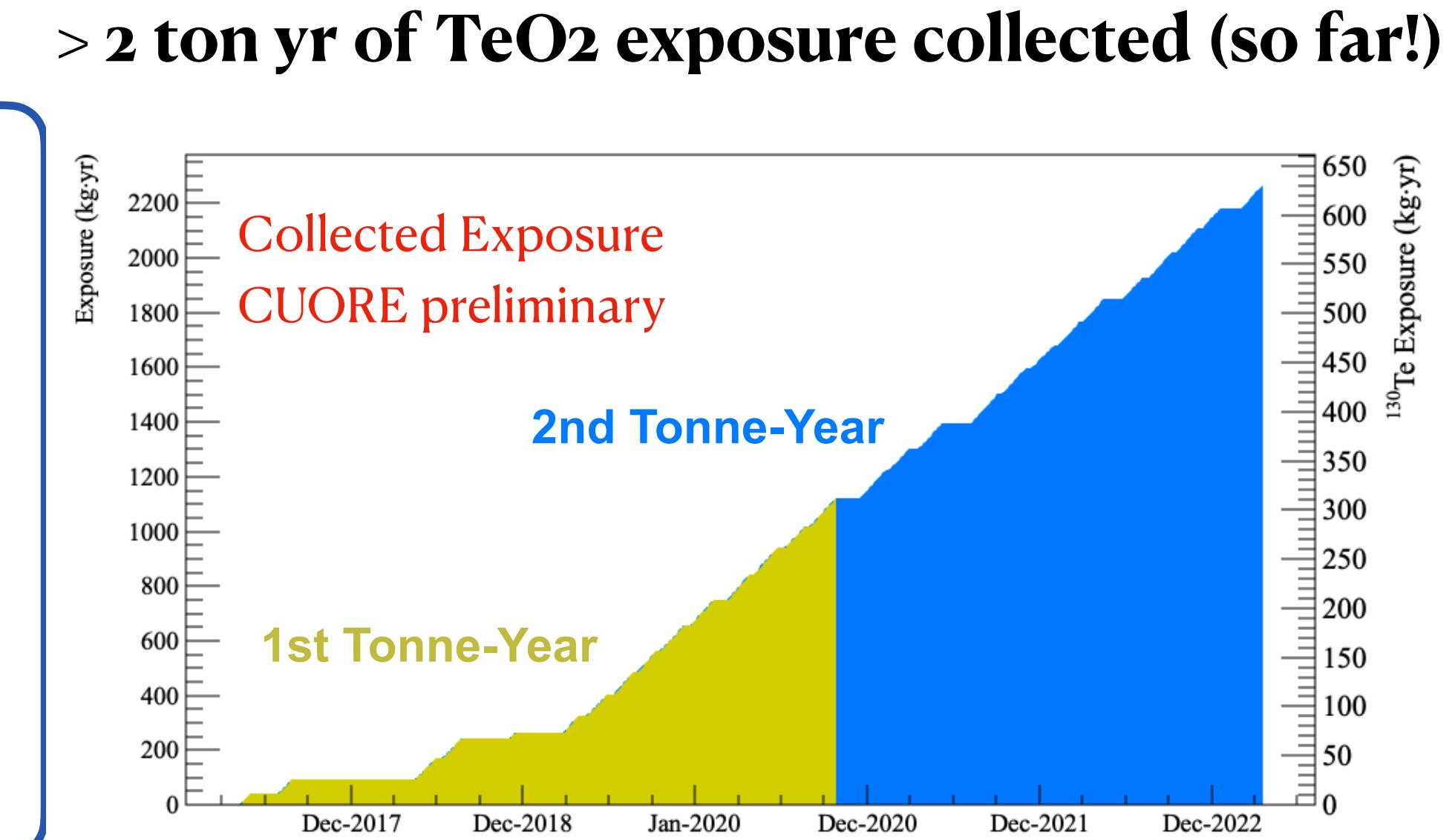
The source of the decay coincide with the absorber, implying an very high detection efficiency (88.4%)

$$S_{0\nu} \propto \epsilon \sqrt{\frac{MT}{B\Delta}}$$

Large scalability for ton-scale experiments (only limited by the cryogenics)



Thermal phonon detection allows for an excellent energy resolution (**0.3 %**)
 Negligible intrinsic resolution (~ 10 eV)
 Limited by Vibrational noise (\sim keV)

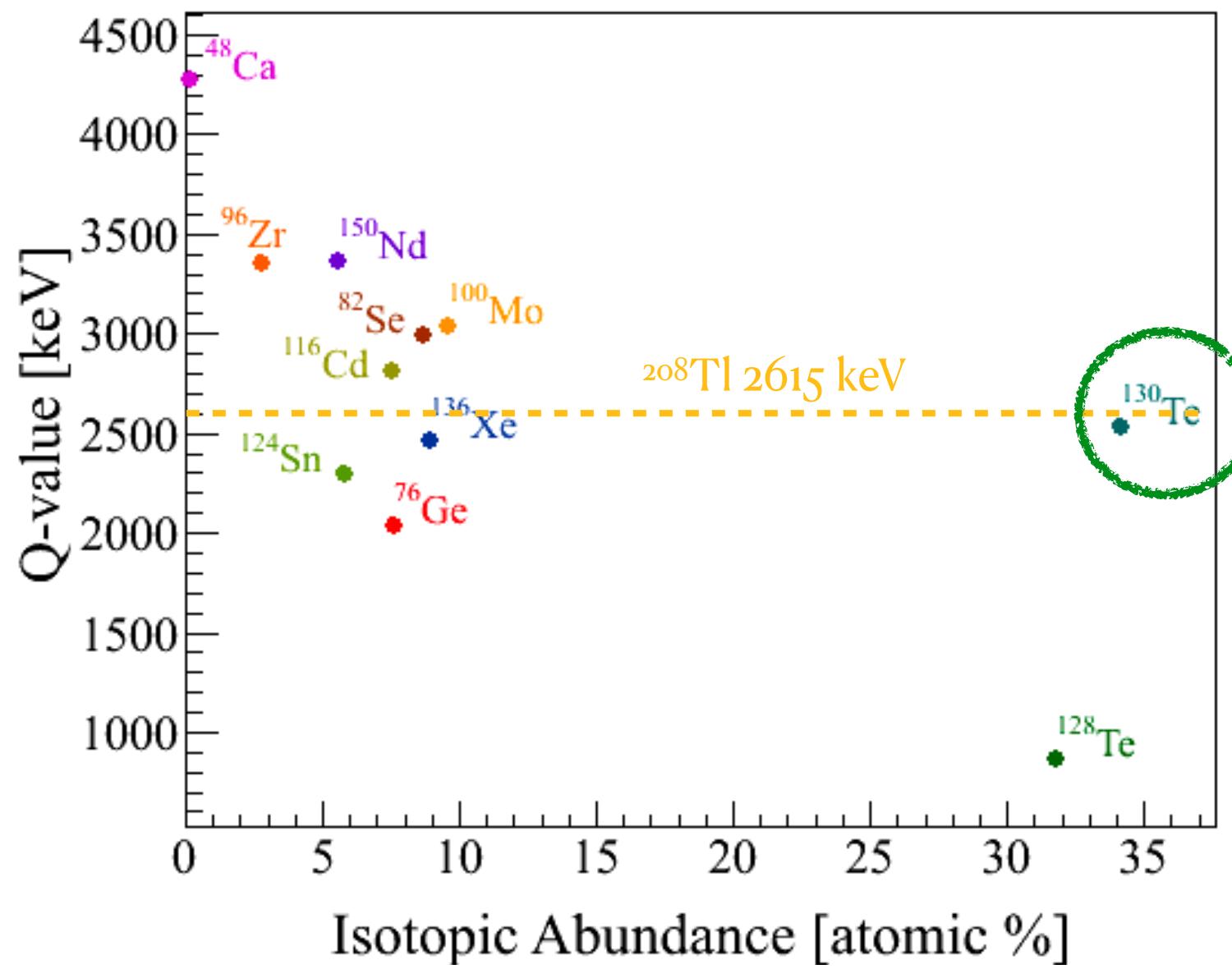


$0\nu\beta\beta$ with CUORE

^{130}Te :

Large natural abundance (34%), no need for enrichment

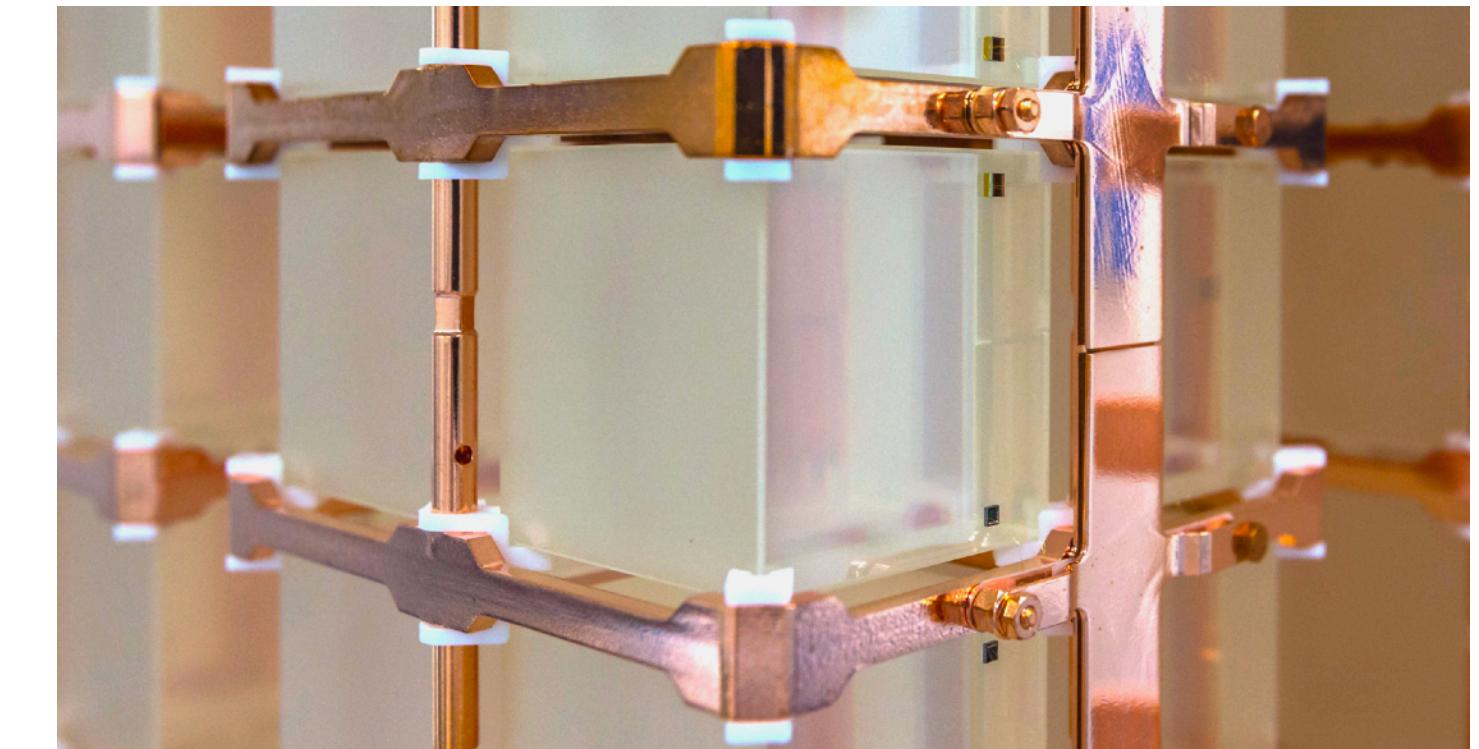
Q-value is close to the endpoint of gamma natural radioactivity



$$S_{0\nu} \propto \epsilon \sqrt{\frac{MT}{B\Delta}}$$

Cryogenic calorimeters technology presents a large flexibility in the materials choice

- Detector structure and components
- Crystal compounds



Select radiopure materials (copper and PTFE) and perform strict cleaning procedures

$0\nu\beta\beta$ with CUORE



Muons: $\sim 3 \times 10^{-8} \text{ s}^{-1} \text{ cm}^{-2}$

Neutrons: $< 4 \times 10^{-6} \text{ s}^{-1} \text{ cm}^{-2}$

Gammas: $\sim 0.73 \text{ s}^{-1} \text{ cm}^{-2}$

$$S_{0\nu} \propto \epsilon \sqrt{\frac{MT}{B\Delta}}$$

Operated Underground at LNGS with
3600 meters of water equivalent to
shield against cosmic rays

External layers against radioactivity (e.g.
Ancient Roman Lead)

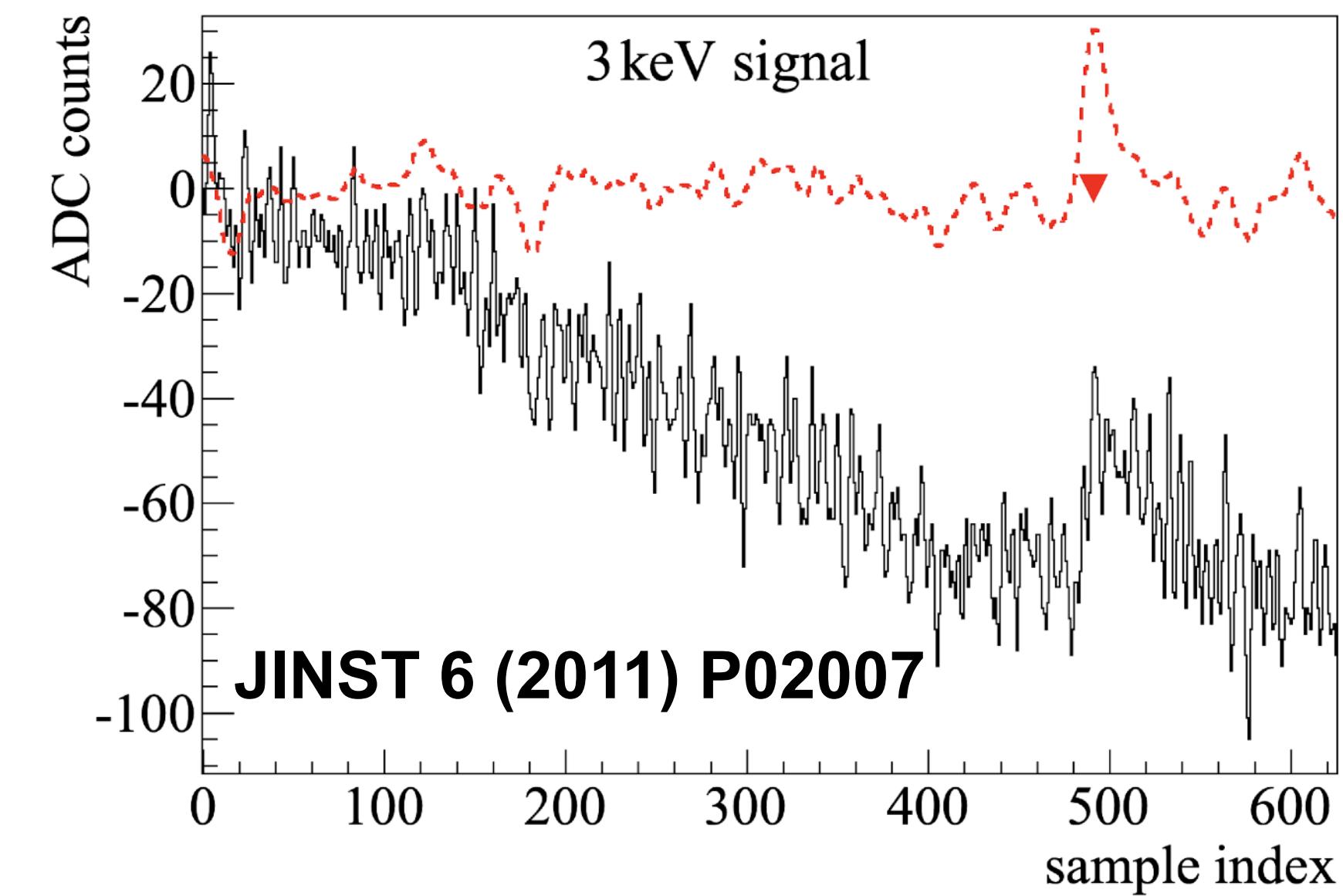
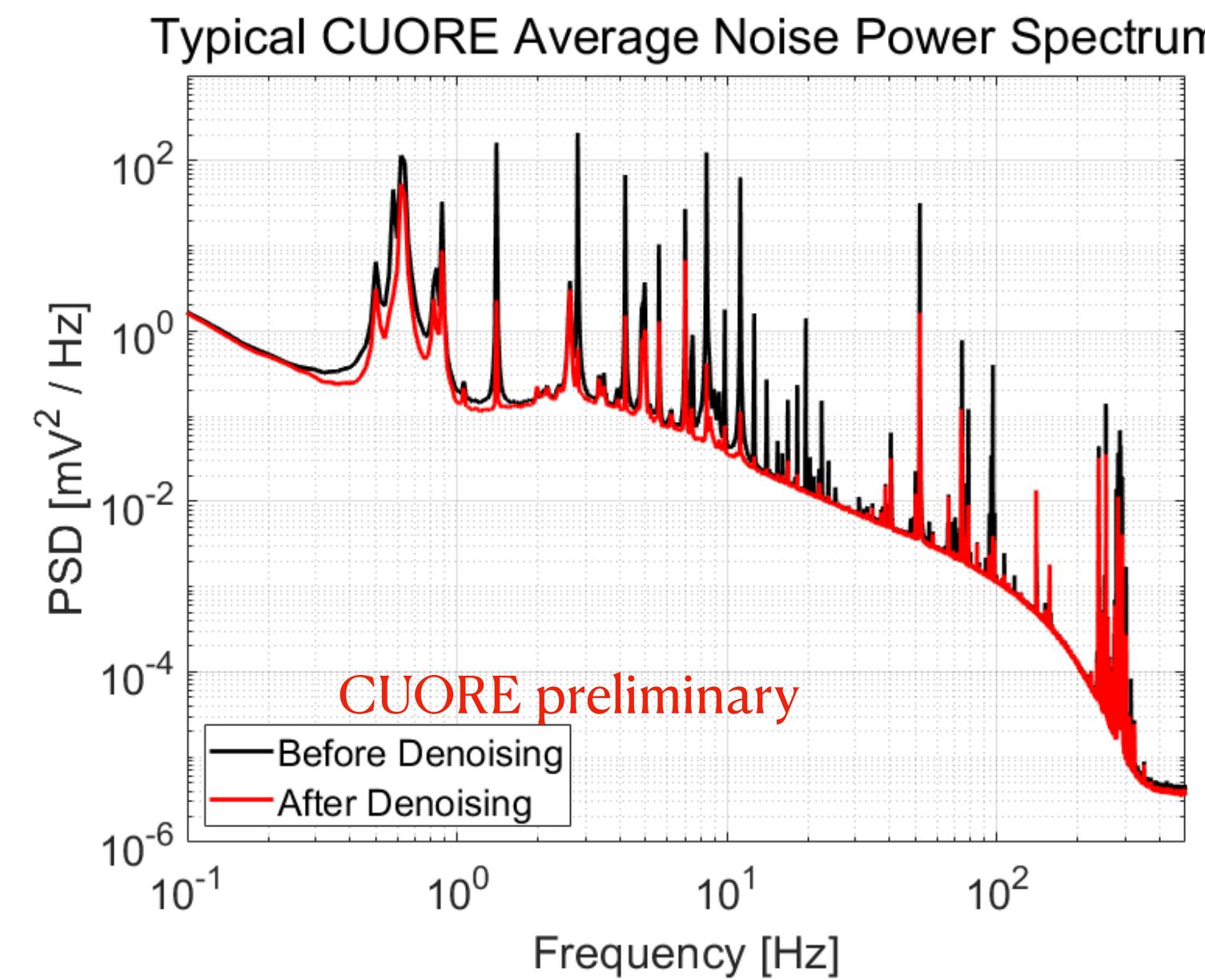
$1.3 \times 10^{-2} \text{ counts}/(\text{keV kg yr})$ at 2528 keV



Analysis Methods

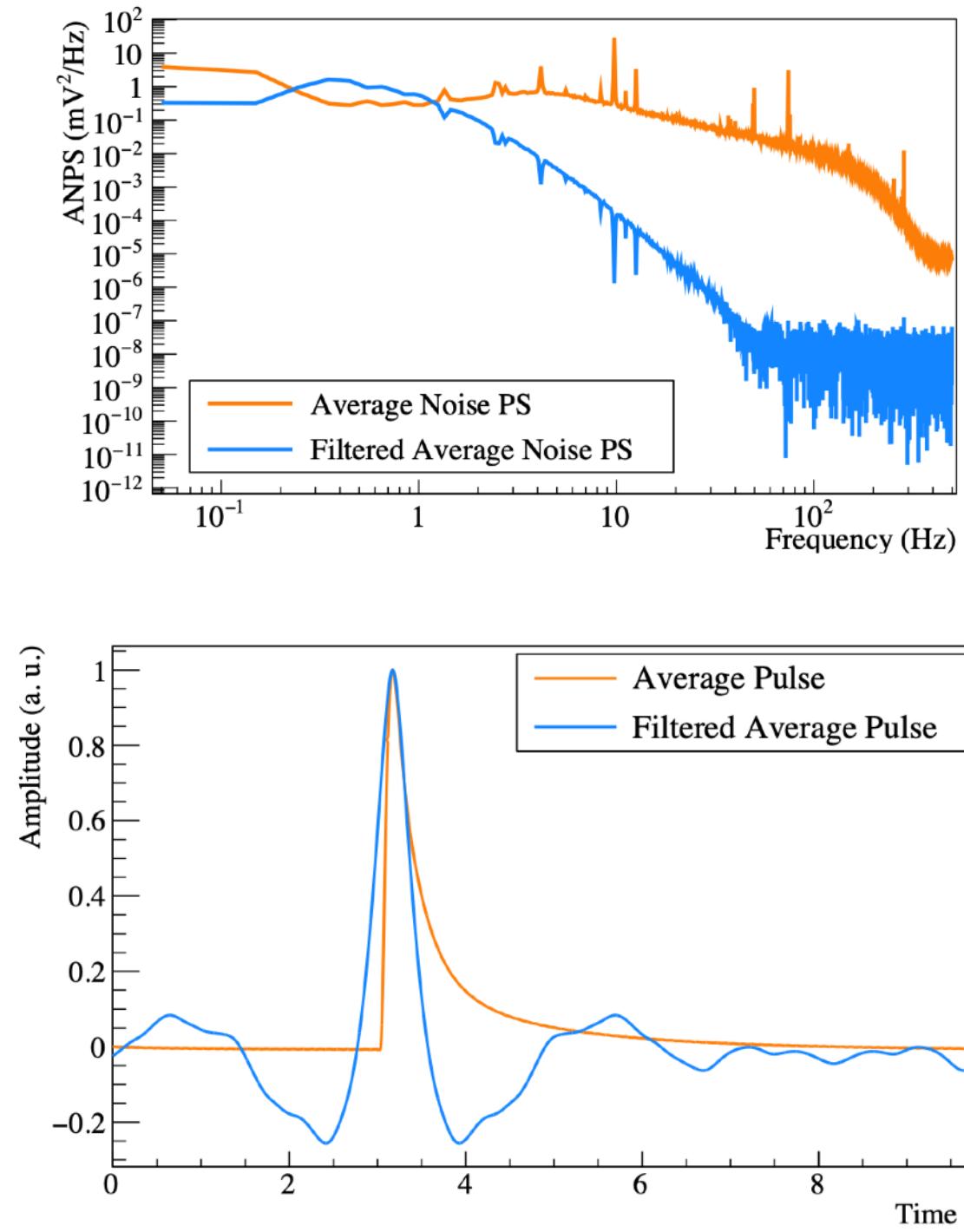
Signal Processing:

- **Denoising:** mitigate the noise by correlating it with auxiliary devices (microphones, accelerometers, seismometers)
- **Optimum Trigger:** apply an offline trigger on filtered waveforms to lower the energy threshold

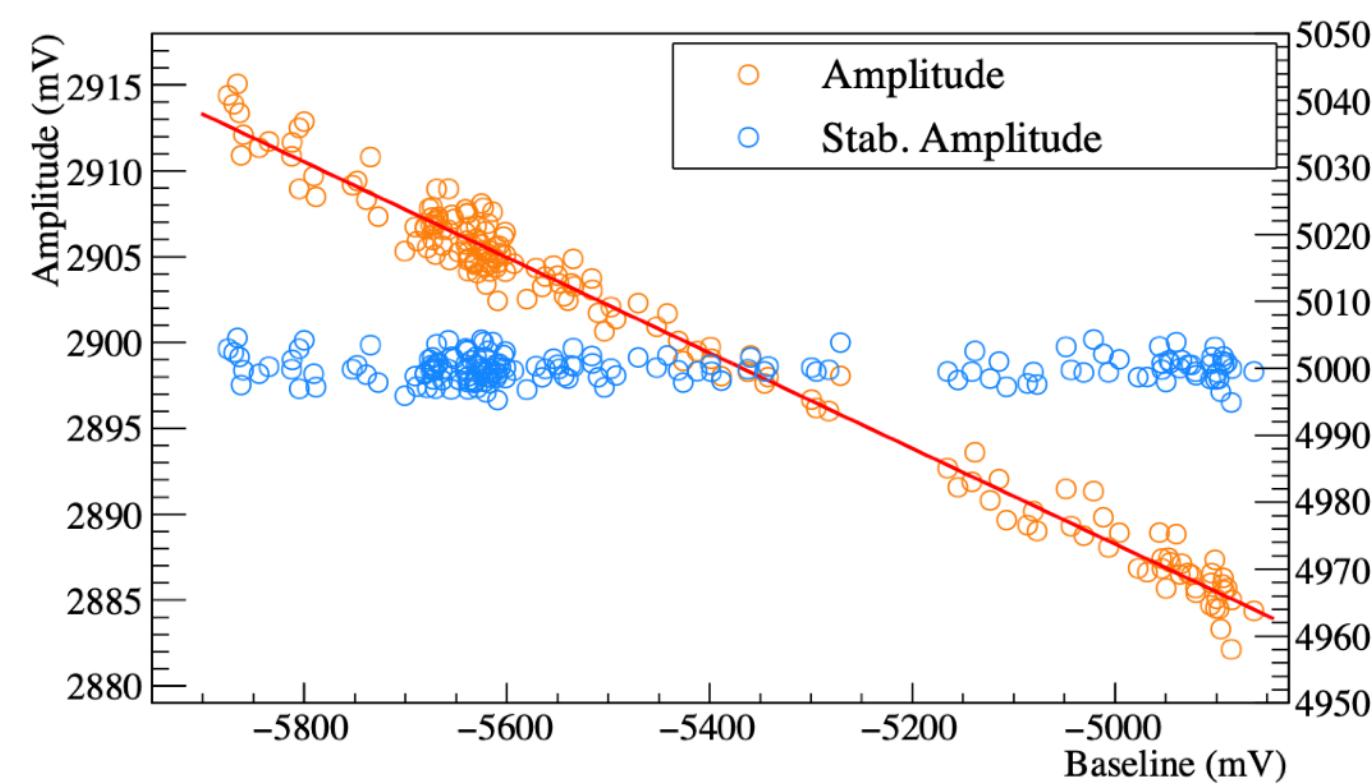
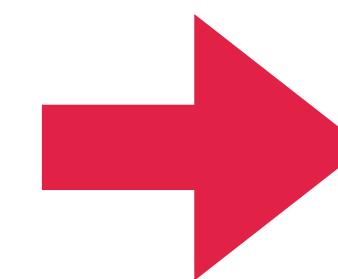


<https://indico.cern.ch/event/1199289/contributions/5447391/>

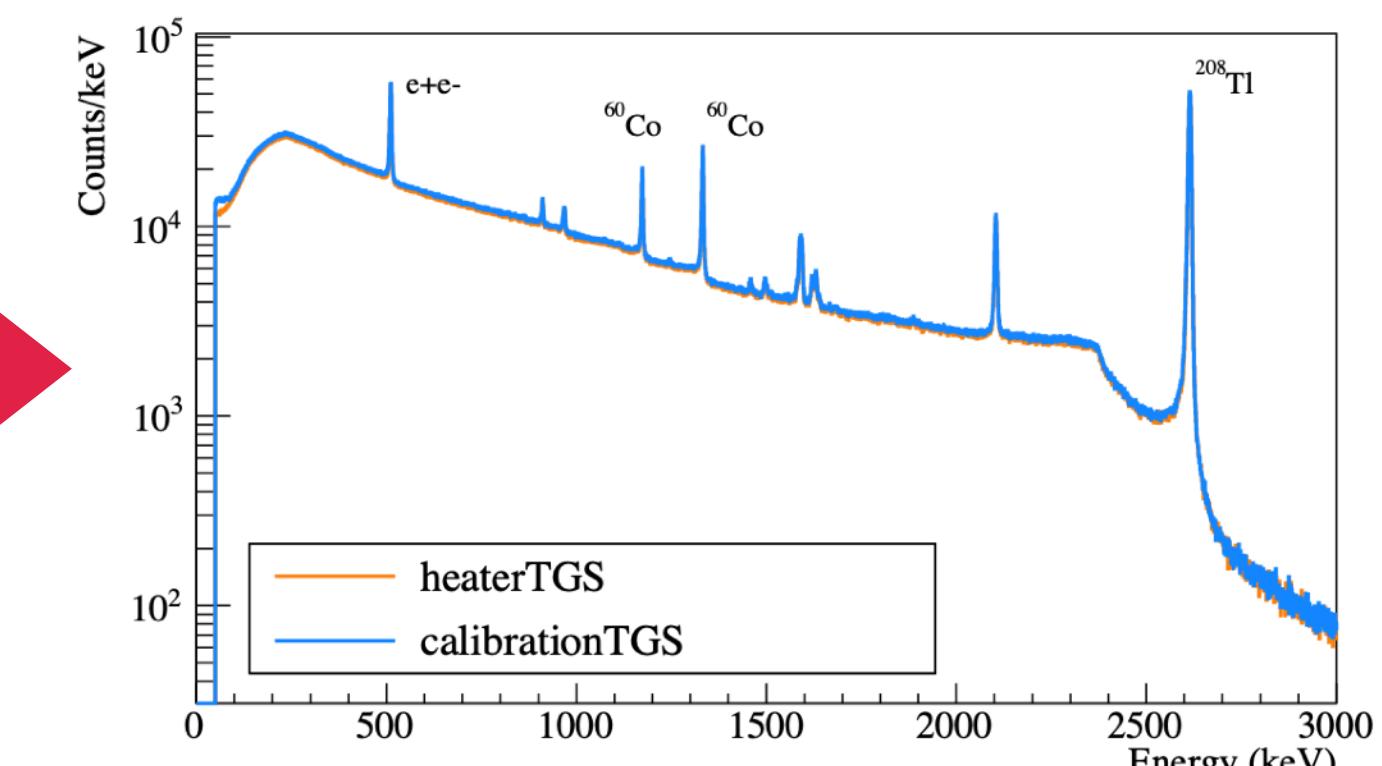
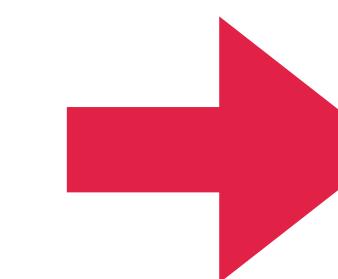
Analysis Methods



Optimum Filter: suppress the frequencies most affected by the noise relying with ideal pulse and noise spectrum



Thermal Gain Correction: correct amplitude dependence on the operating temperature (\sim baseline) drift by using the injected thermal pulses

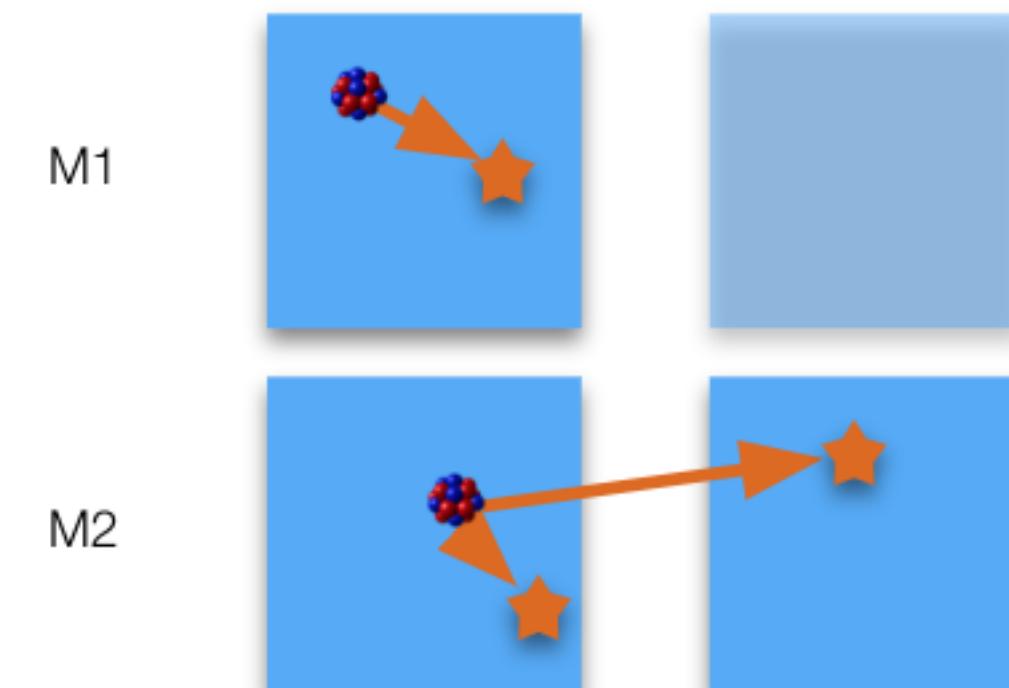


Energy Calibration: based on measurements with external ^{232}Th - ^{60}Co source deployment

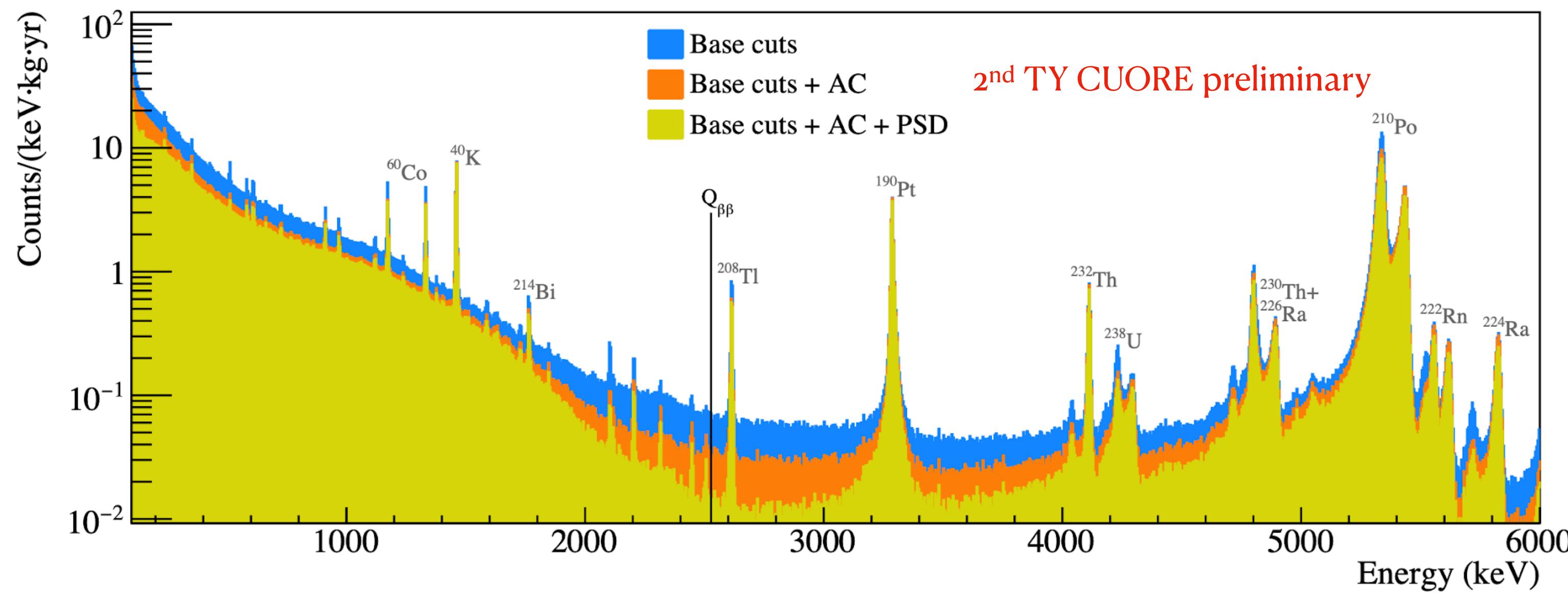
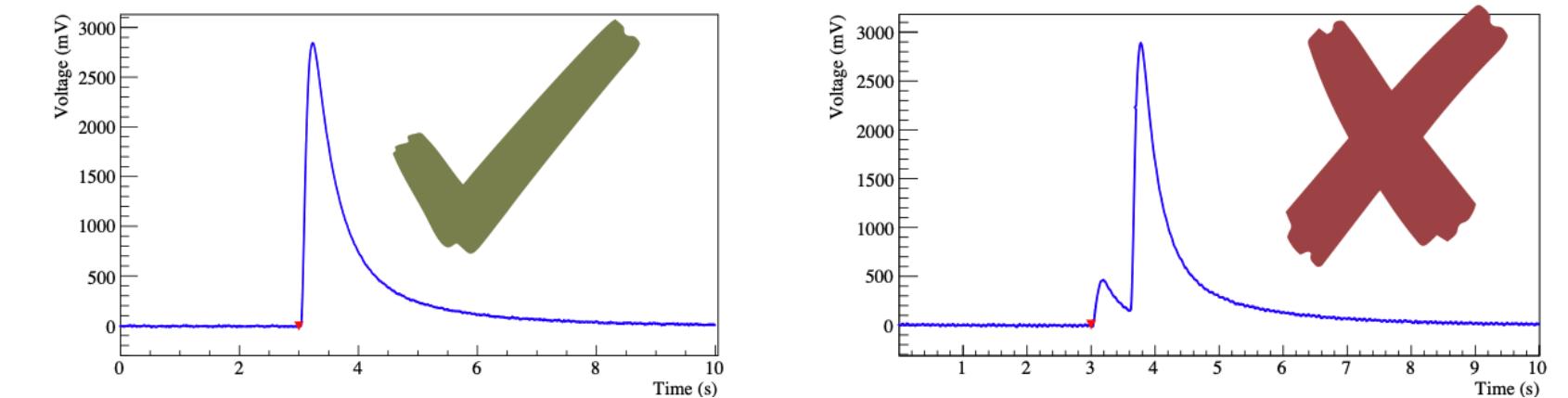
<https://indico.cern.ch/event/1199289/contributions/544739/>

Analysis Methods

- **Anti-Coincidence:** reject events depositing energy in multiple crystals



- **Pulse Shape Discrimination:** implemented using Principal Component Analysis (PCA)



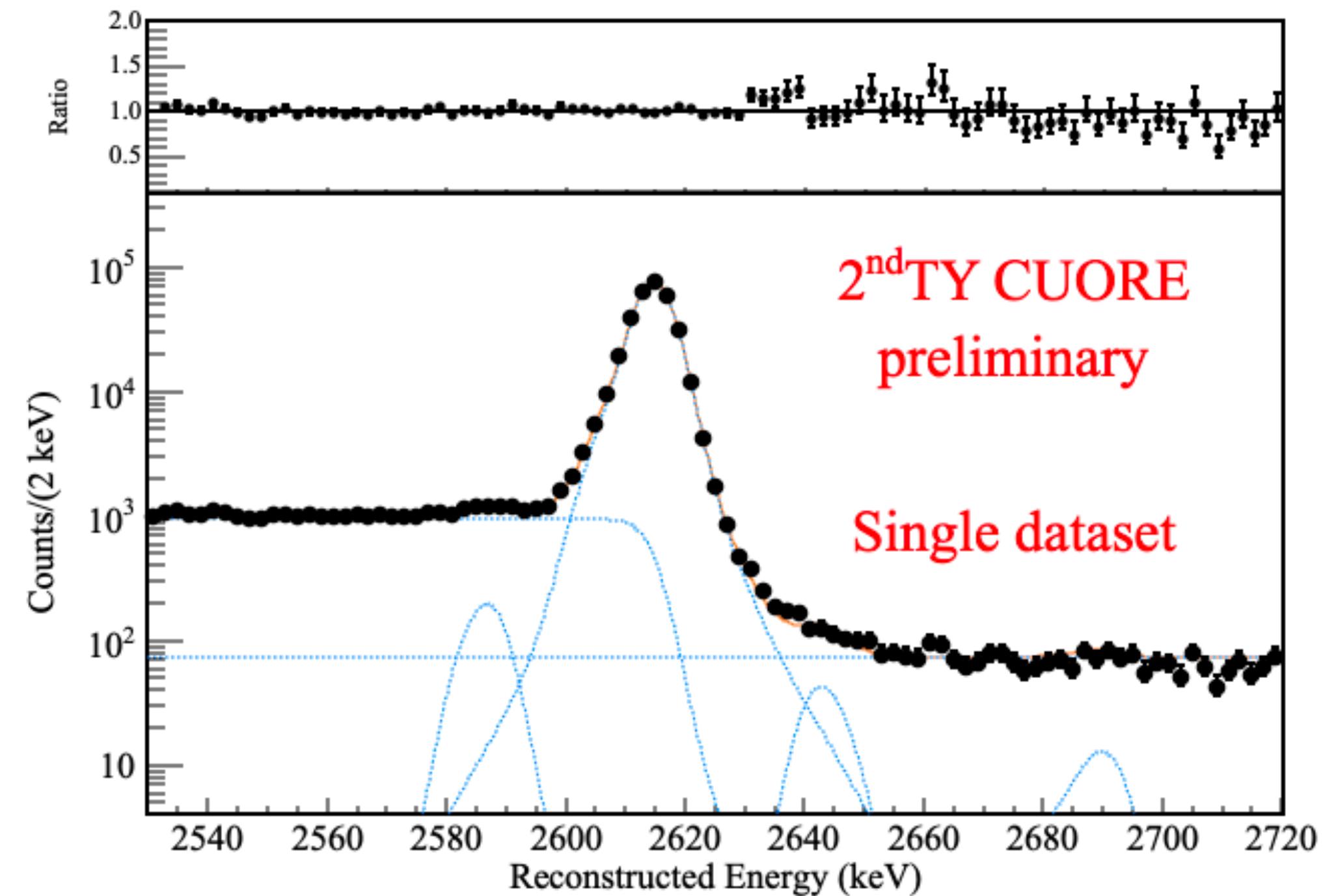
- **Blinding:** exchange events from ^{208}Tl 2615 keV line to the ^{130}Te Q-value

Analysis Methods

Physics peaks modeled on ^{208}Tl 2615 keV line (calibration data):

- 3 Gaussians
- Te x-rays (escape + coincident)
- 583 keV gamma line coincident with annihilation escape peak
- multi-Compton and flat background

$$\Delta E_{2615 \text{ keV}, 2^{\text{nd}}\text{TY}} = 7.43 \pm 0.37 \text{ keV}$$



Analysis Methods

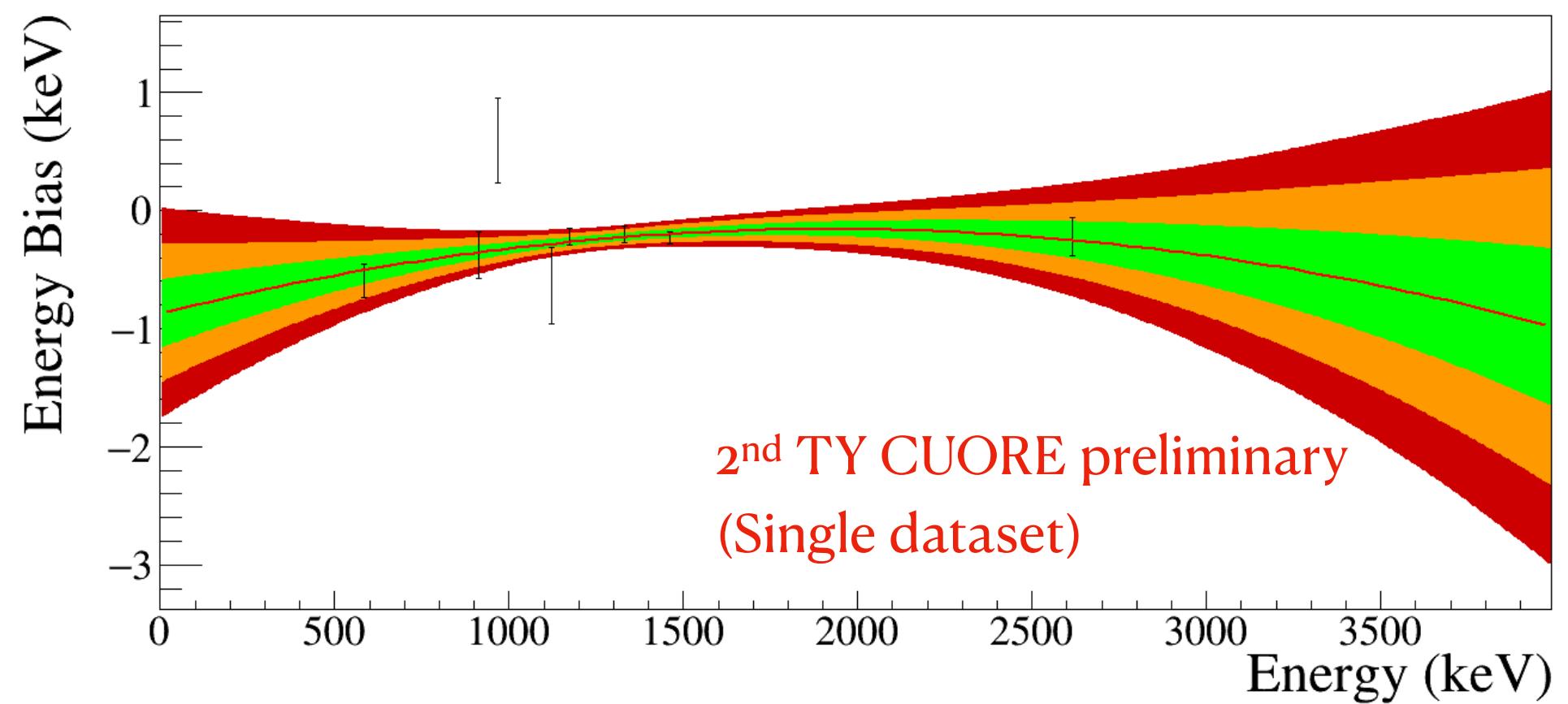
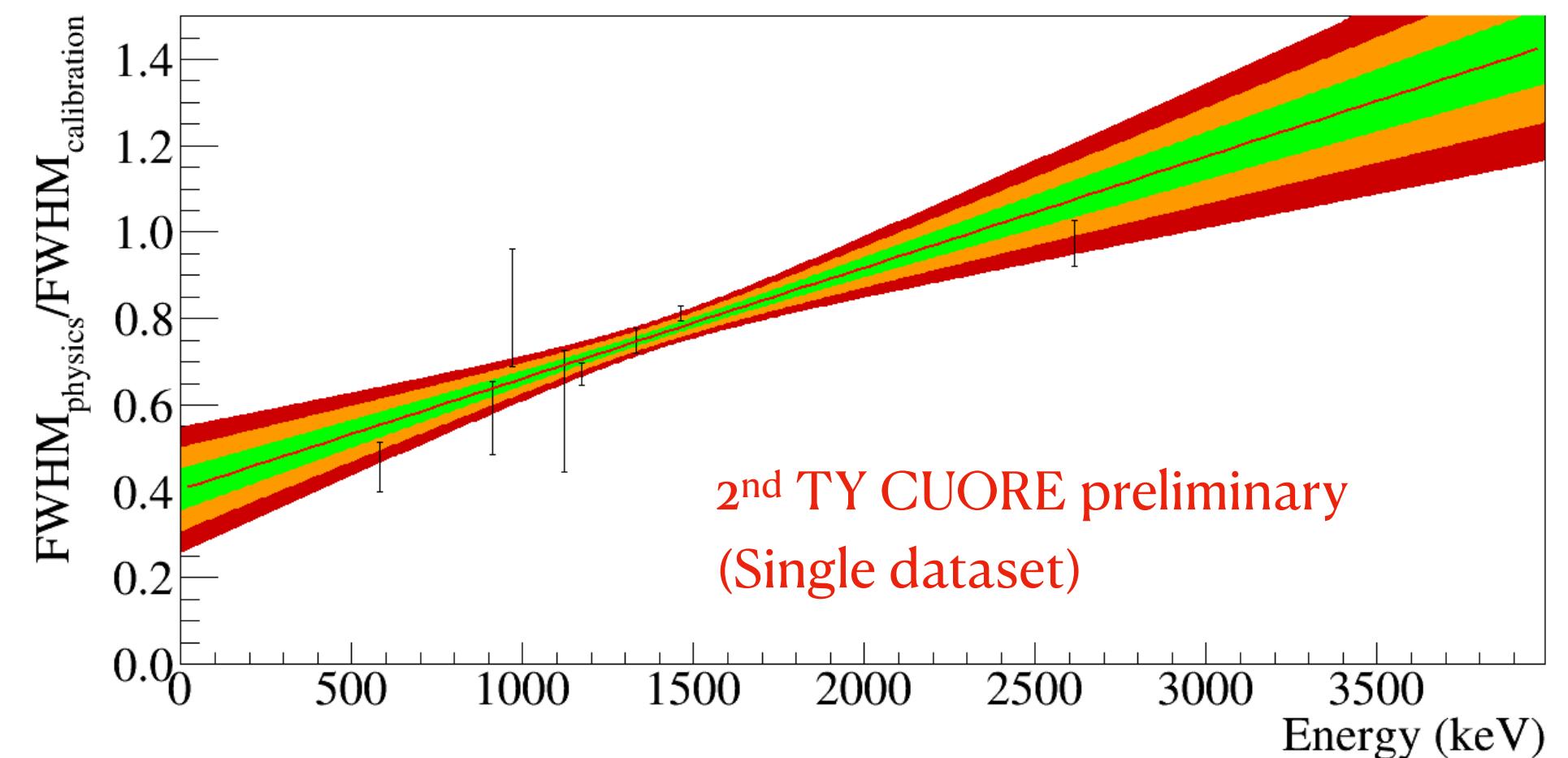
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$$\Delta E_{2615 \text{ keV}, 2^{\text{nd}} \text{ TY}} = 7.43 \pm 0.37 \text{ keV}$$

Scale the results at the Q-value of ^{130}Te fitting peaks in physics data
(noise and pile-up improved)

$$\Delta E_{Q_{\beta\beta}, 2^{\text{nd}} \text{ TY}} = 7.26^{+0.43}_{-0.47} \text{ keV}, E_{bias, 2^{\text{nd}} \text{ TY}} = -0.11^{+0.19}_{-0.25} \text{ keV}$$



Search for $0\nu\beta\beta$ with 2nd ton yr only

Region of Interest [2465, 2575] keV

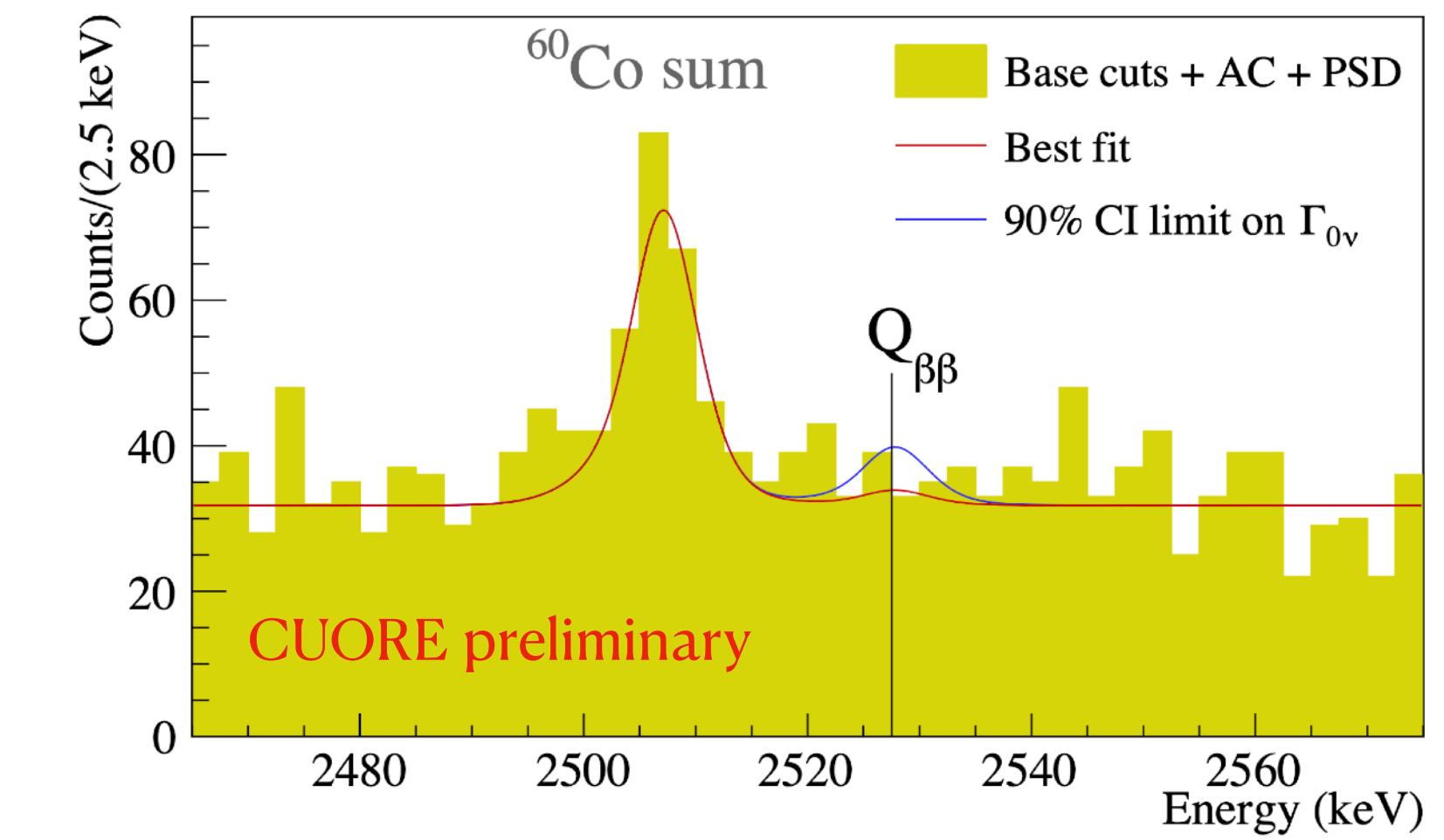
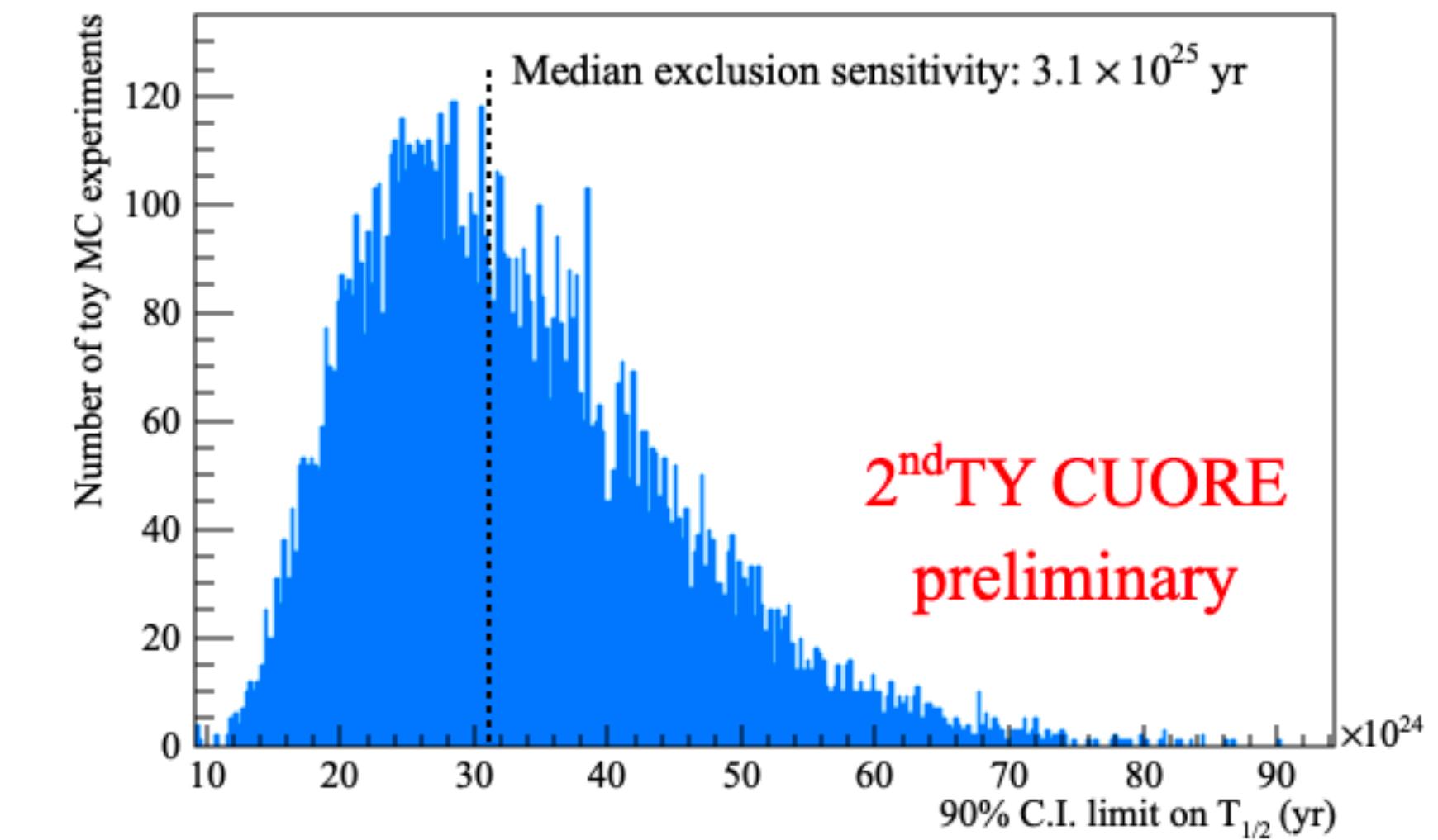
- Peak at 2528 keV ($0\nu\beta\beta$ signal)
- ^{60}Co peak
- Linear background
- Rates and background index and slope as free parameters
- Systematics treated as nuisance parameters
- Sensitivity evaluated from toy experiments in background only hypothesis

Average Background Index = 1.3×10^{-2} counts/(keV kg yr)

No evidence for $0\nu\beta\beta$

Decay Rate limit at 90% C.I. $\Gamma_{0\nu} < 2.5 \times 10^{-26}$ /yr

Half-life limit at 90% C.I. $T_{1/2} > 2.7 \times 10^{25}$ yr



Combine 1st and 2nd ton yr of data

We combined the posteriors on $0\nu\beta\beta$ half-life resulting from the analysis of the 1st ($T_{1/2} > 2.2 \times 10^{25}$ yr, [Nature 604, 53-58 \(2022\)](#)) and 2nd ton yr

→ Overall exposure 2023 kg yr

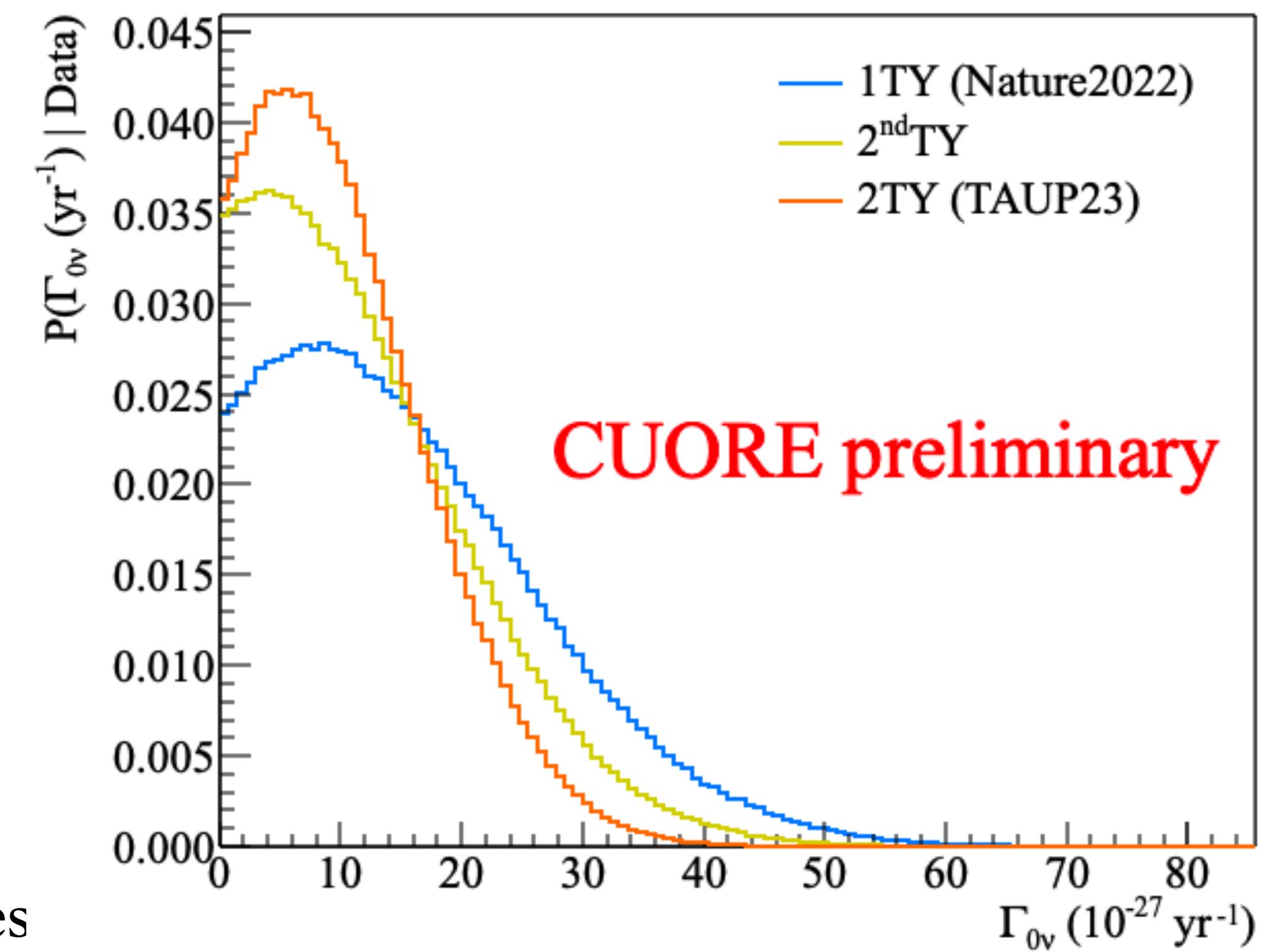
(Still) No evidence for $0\nu\beta\beta$

Decay Rate limit at 90% C.I. $\Gamma_{0\nu} < 2.1 \times 10^{-26}$ /yr

Half-life limit at 90% C.I. $T_{1/2} > 3.3 \times 10^{25}$ yr

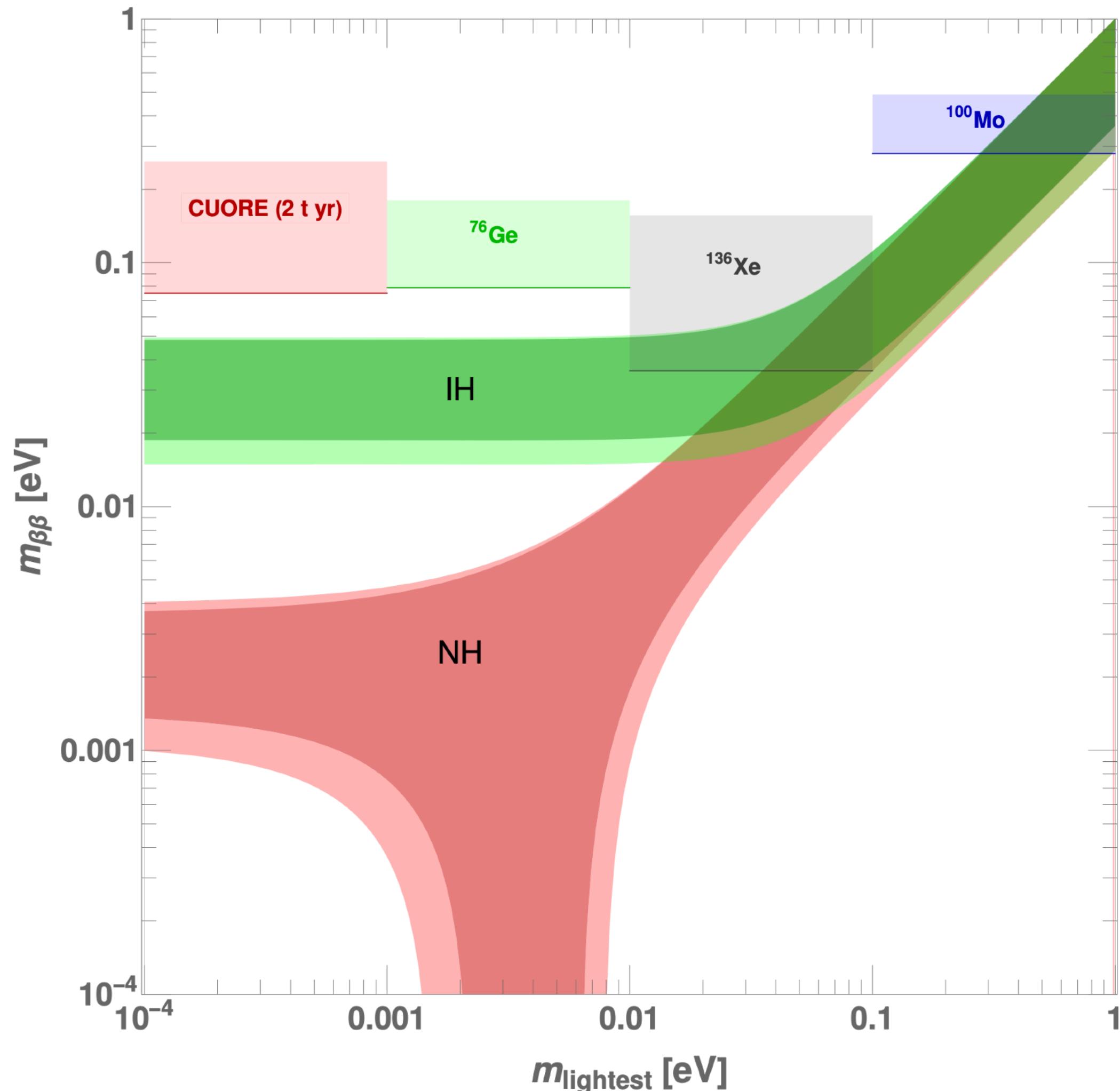
Stay tuned for the full analysis!

- Reprocess the 1st ton yr of data with the updated analysis techniques
- Repeat the $0\nu\beta\beta$ fit
- Finalize systematics



<https://indico.cern.ch/event/1199289/contributions/5447112/>

Limit on Majorana Mass



We converted half-life limit into an upper limit on the Majorana mass

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q, Z) |M^{0\nu}|^2 \frac{|\langle m_{\beta\beta} \rangle|^2}{m_e^2}$$

Phase Space Factor

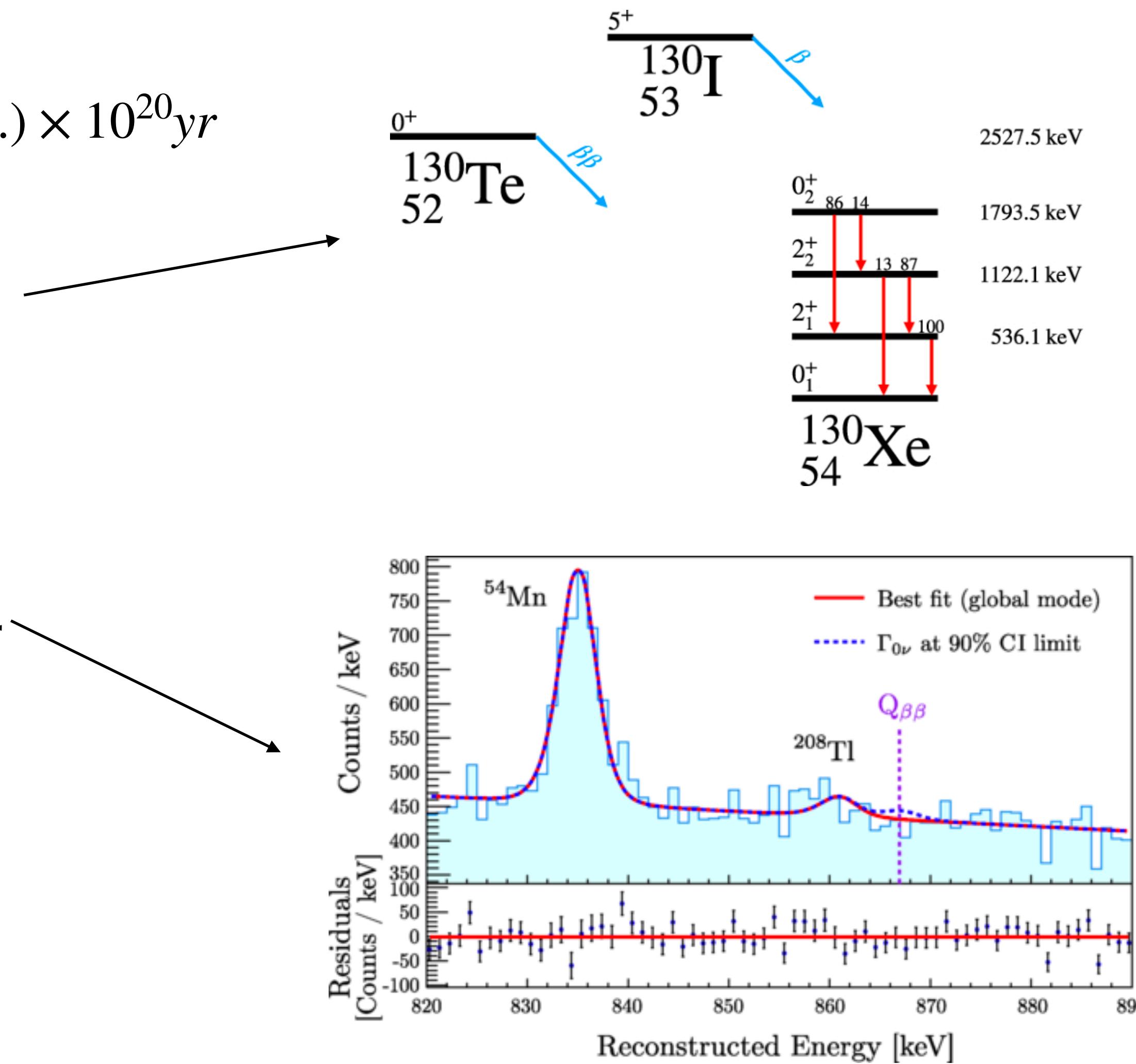
Nuclear Matrix Elements from several possible models (ISM, IBM, QRPA, ...) here assuming $g_A = 1.27$

$$m_{\beta\beta} < 75 - 255 \text{ meV}$$

More Double β decay results CUORE

Published results

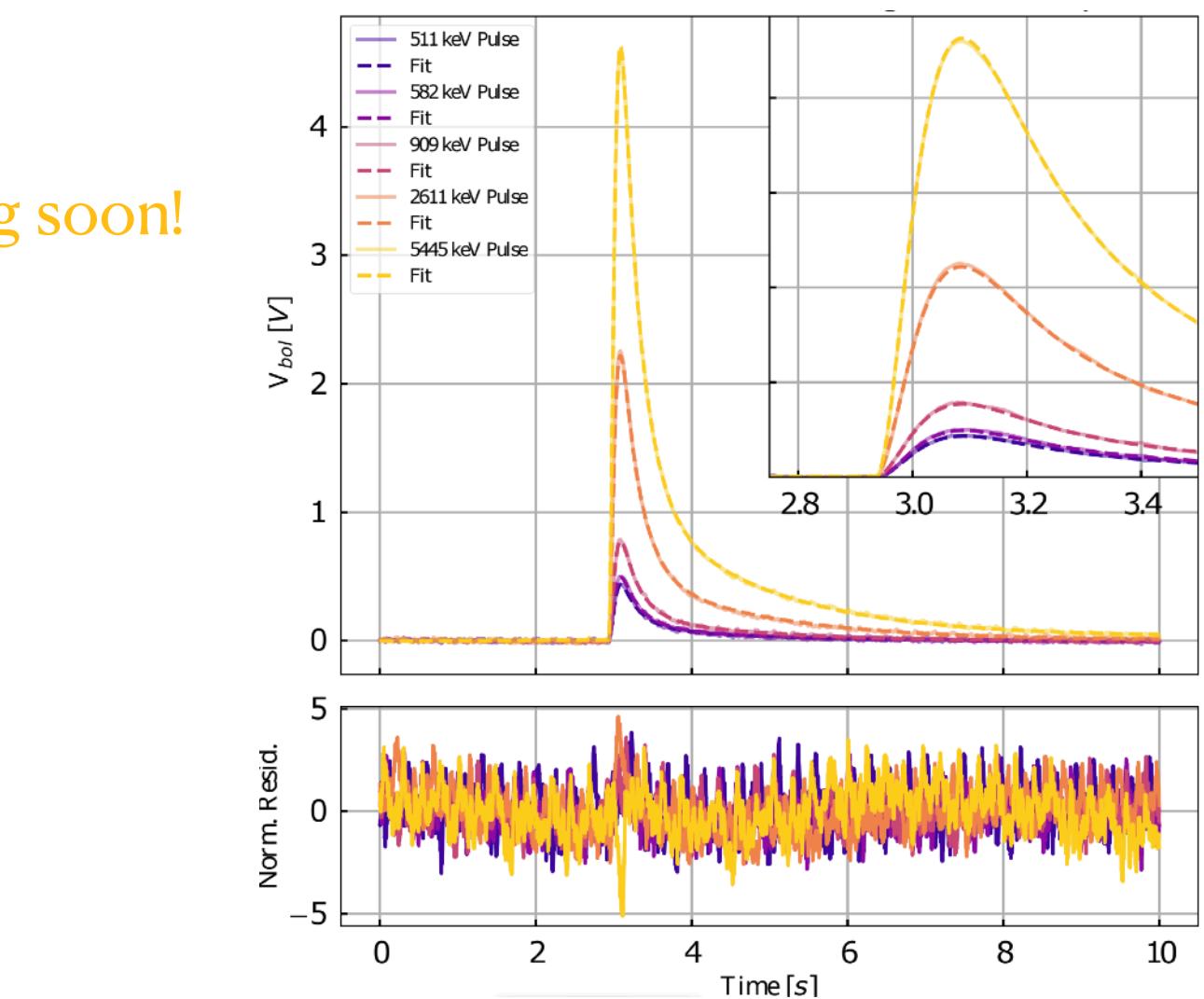
- $2\nu\beta\beta$ measurement: $T_{1/2} = 7.71^{+0.08}_{-0.06}(\text{stat.})^{+0.12}_{-0.15}(\text{syst.}) \times 10^{20} \text{ yr}$
<https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.126.171801>
- $0\nu\beta\beta$ and $2\nu\beta\beta$ to the first 0^+ excited state of ^{130}Te
<https://link.springer.com/article/10.1140/epjc/s10052-021-09317-z>
 - $T_{1/2}^{0\nu} > 5.9 \times 10^{24} \text{ yr}$
 - $T_{1/2}^{2\nu} > 1.3 \times 10^{24} \text{ yr}$
- $0\nu\beta\beta$ of ^{128}Te (867 keV - 188 kg): $T_{1/2} > 3.6 \times 10^{24} \text{ yr}$
<https://journals.aps.org/prl/abstract/10.1103/PhysRevLett.129.222501>
- $0\nu\beta^+EC$ of ^{120}Te : $T_{1/2} > 2.9 \times 10^{22} \text{ yr}$
<https://journals.aps.org/prc/abstract/10.1103/PhysRevC.105.065504>



Beyond Double β decay in CUORE

- Detailed thermal model of the detector response
<https://iopscience.iop.org/article/10.1088/1748-0221/17/11/P11023/meta>
- Denoising Techniques (already in the analysis!)
<https://indico.cern.ch/event/1199289/contributions/5447124/>
- Impact of marine microseism on detector response
<https://indico.cern.ch/event/1199289/contributions/5445886/>
- Background Model
<https://indico.cern.ch/event/1199289/contributions/5447161/>
- Dark matter search at low energies (Solar Axions, WIMPs...)
<https://indico.cern.ch/event/1199289/contributions/5445882/>

Published Results

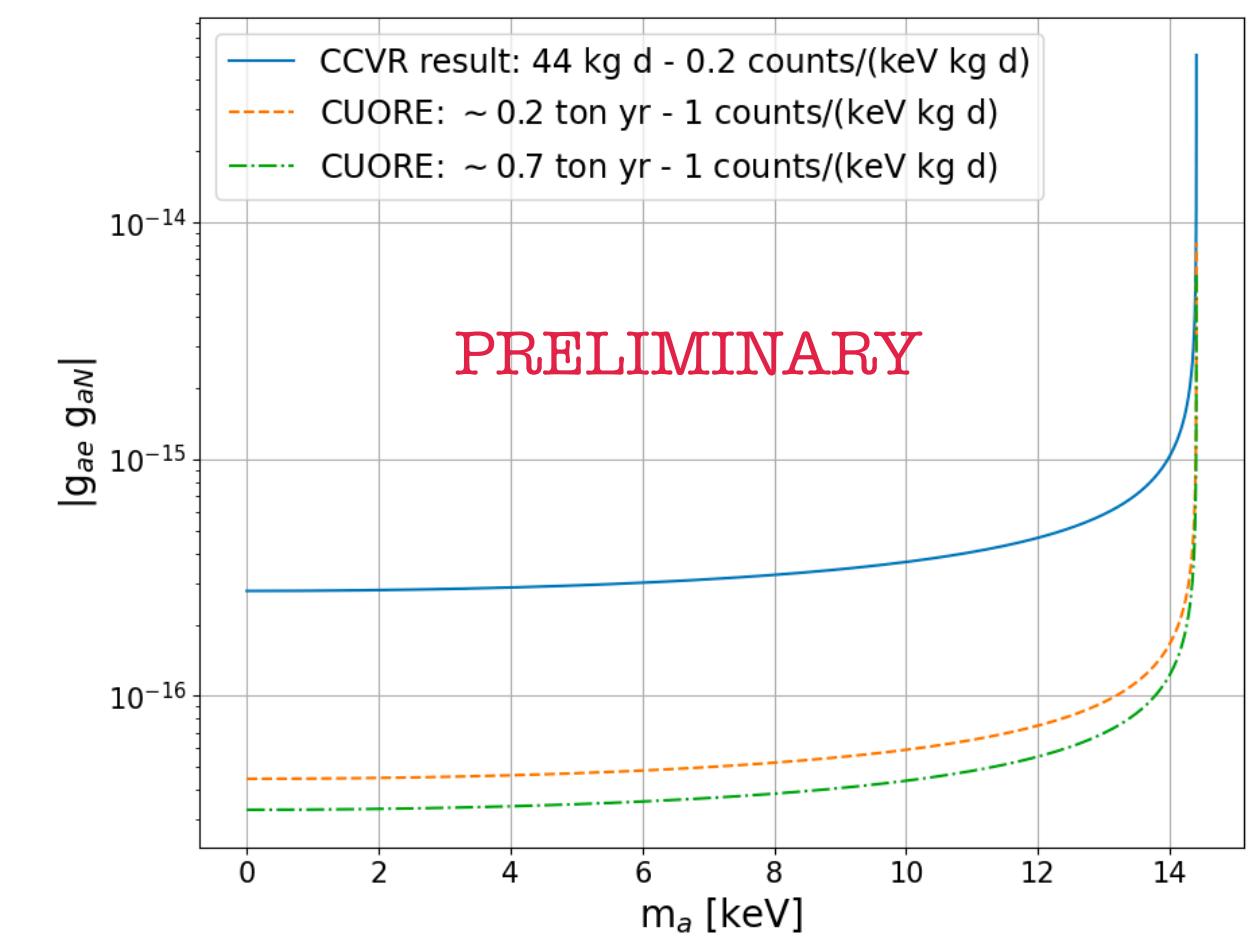
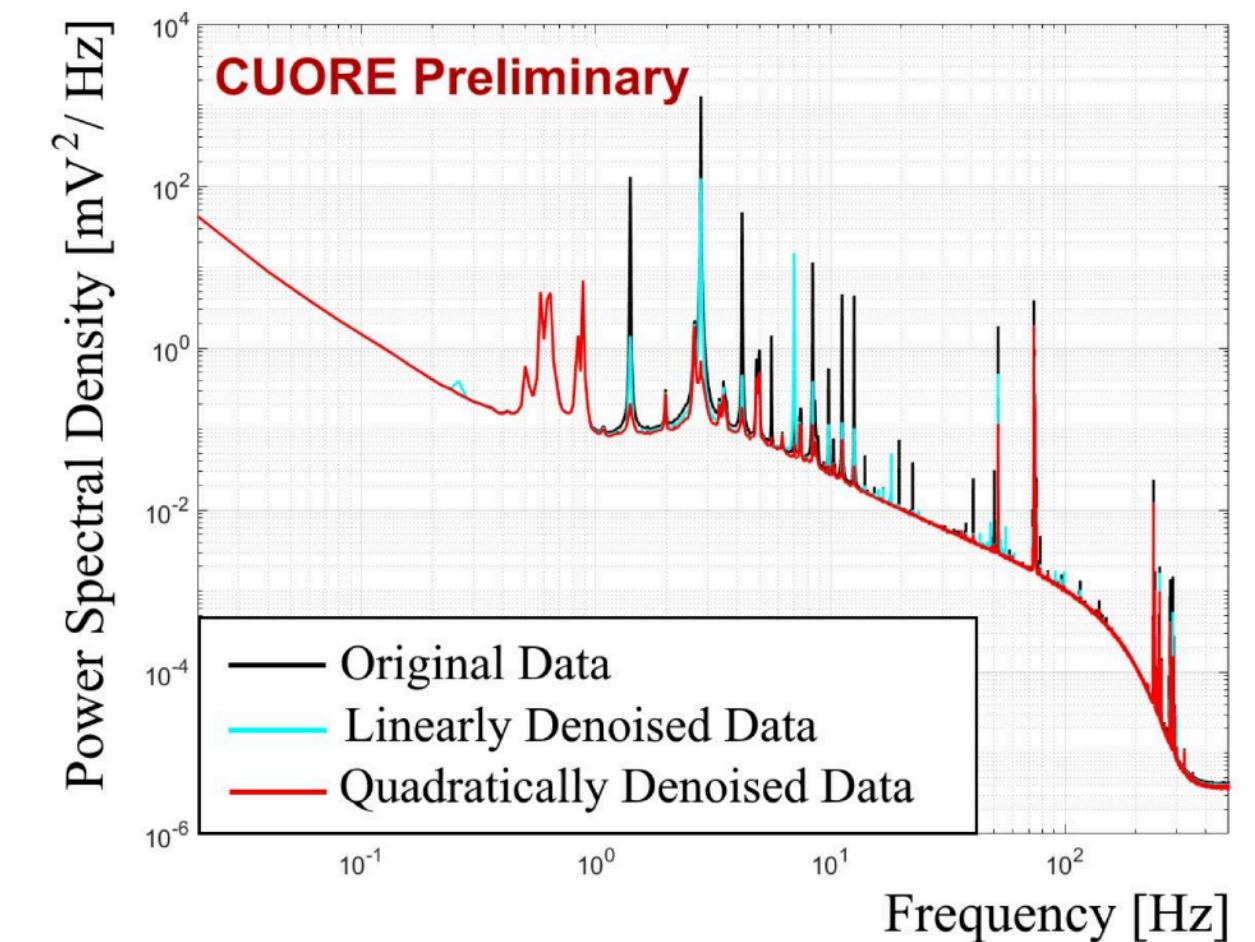


Coming soon!

Coming soon!

Coming soon!

Coming soon!





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CAL POLY
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Technology

Lawrence Livermore
National Laboratory

Thank You!



Yale

G S
S I

JOHNS HOPKINS
UNIVERSITY

Berkeley
UNIVERSITY OF CALIFORNIA



1st ton yr results

- $T_{1/2} > 2.2 \times 10^{25} \text{ yr}$
- Median Expected Sensitivity = $2.8 \times 10^{25} \text{ yr}$
- Background Index = $1.5 \times 10^{-2} \text{ counts/(keV kg yr)}$
- $\Delta E = 7.8(5) \text{ keV}$

