



nationale  
wetenschaps  
agenda

# Relic Neutrino detection with PTOLEMY

Auke-Pieter Colijn

Rome 30-01-2023

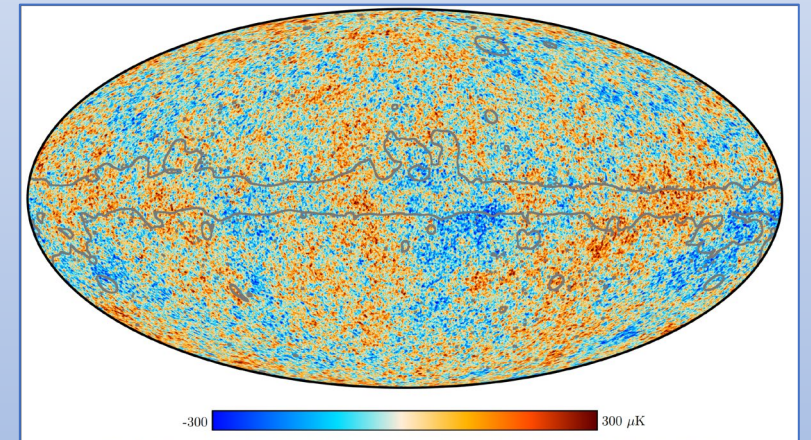


UNIVERSITEIT VAN AMSTERDAM



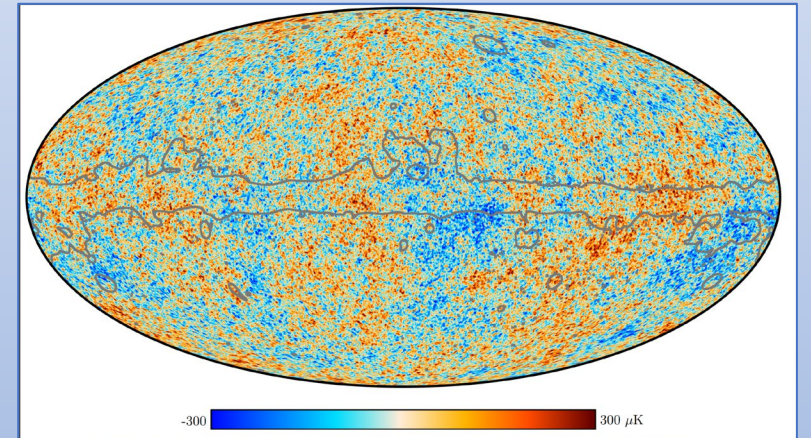
# Why believe Big Bang?

1. Expansion of Universe
2. Light element abundances
3. Cosmic Microwave Background



# Why believe Big Bang?

1. Expansion of Universe
2. Light element abundances
3. Cosmic Microwave Background
4. **Cosmic Neutrino Background**



# Big Bang



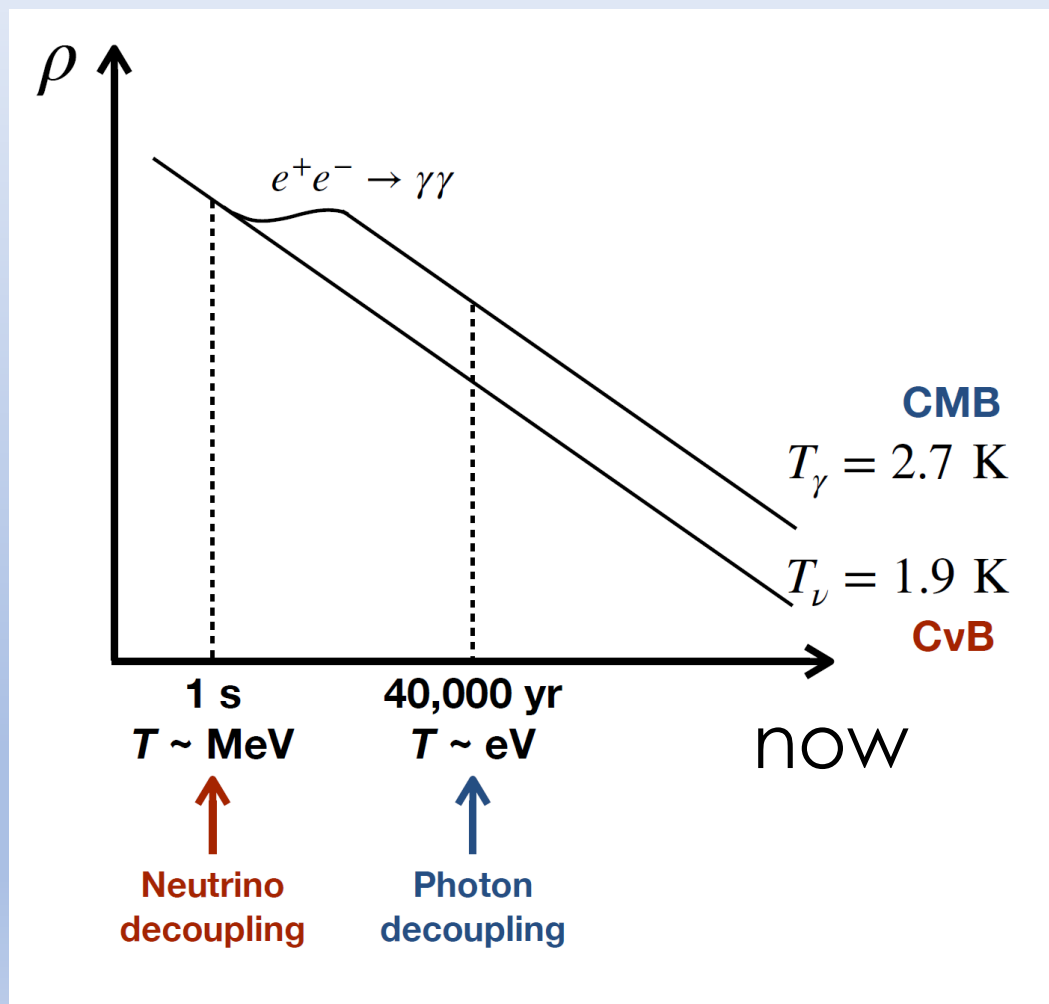
$10^{-36}$  sec

1 sec

380.000 year

30 January 2023

# Evolution of the relic neutrinos



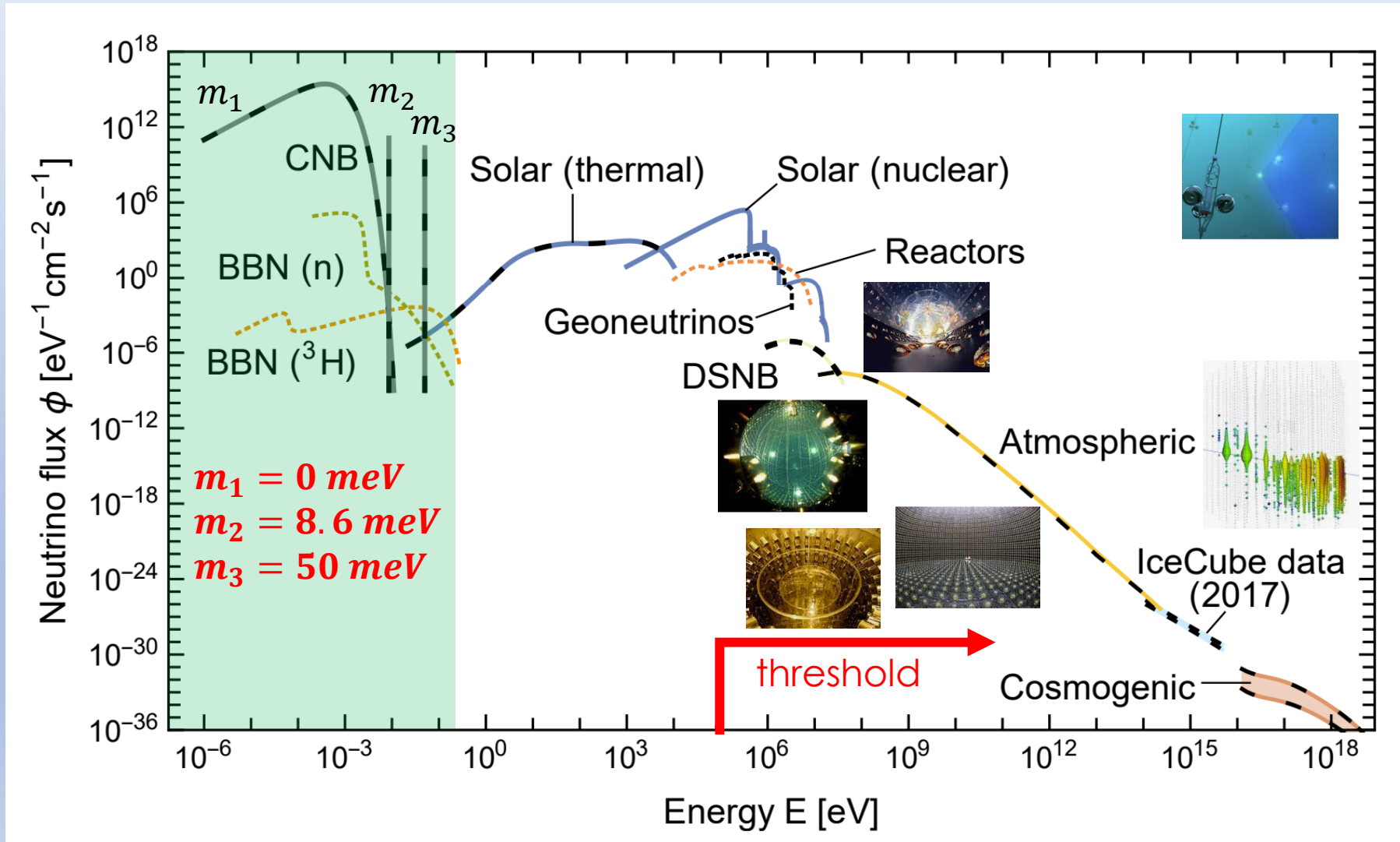
*CνB* and *CMB* temperature related:

$$\frac{T_\nu}{T_\gamma} = \left(\frac{4}{11}\right)^{\frac{1}{3}}$$

$$T_\nu \approx 1.9\text{ K} \Rightarrow p_\nu \approx 0.001\text{ eV}$$

$$n_\nu + n_{\bar{\nu}} \approx 56\text{ cm}^{-3} \times 6$$

# Grand Unified Neutrino Spectrum



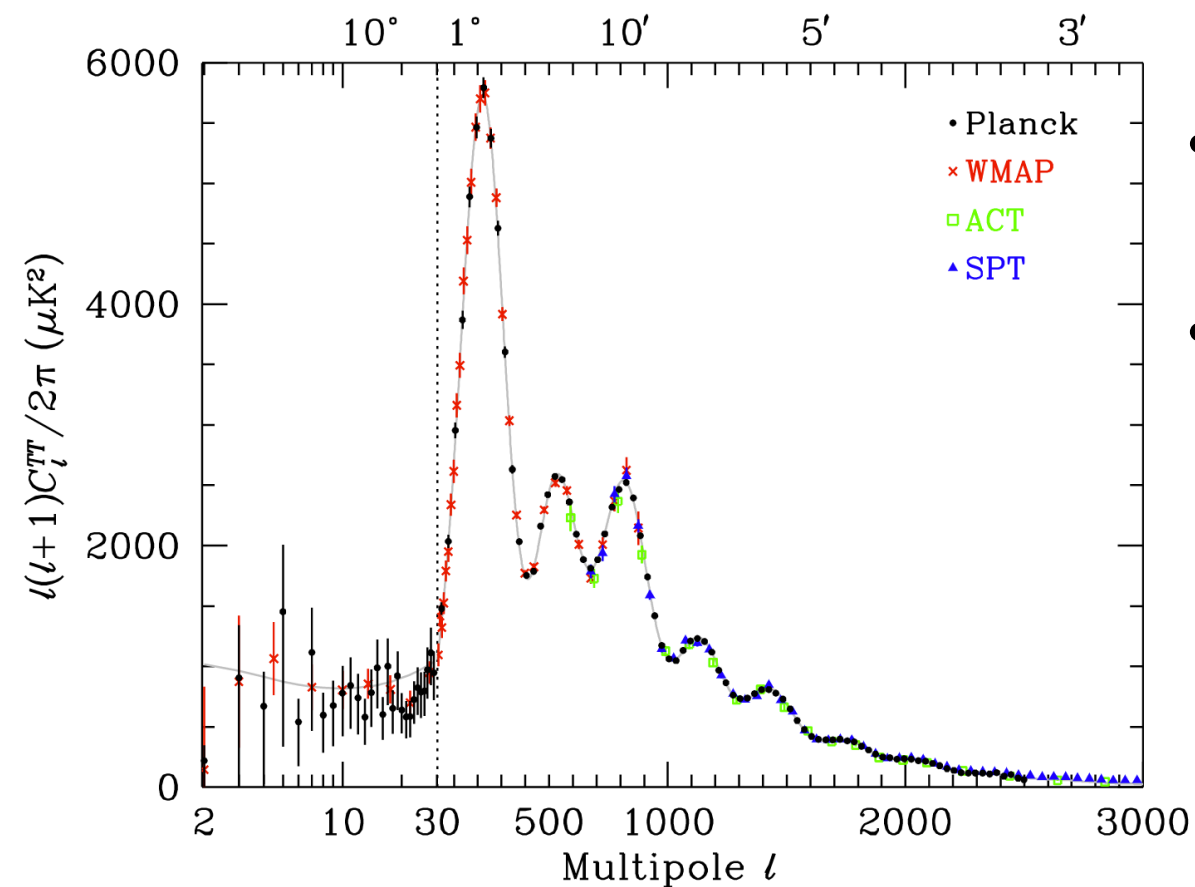
<https://arxiv.org/pdf/1910.11878.pdf>

Where to look for  $C\nu B$ ?

- or -

Is it possible to detect 0.001 eV neutrinos?

# 1. Early Universe & beyond



- $N_{eff} \approx 3$ . They exist
- $\Sigma m_\nu \leq 0.12 eV$ . They are light.

nature  
physics

ARTICLES

<https://doi.org/10.1038/s41567-019-0435-6>

**First constraint on the neutrino-induced phase shift in the spectrum of baryon acoustic oscillations**

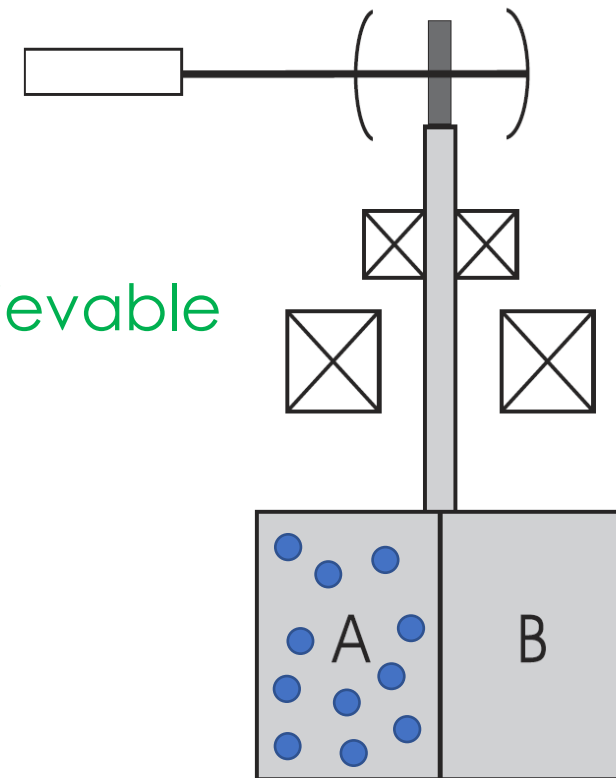
Daniel Baumann<sup>1</sup>, Florian Beutler<sup>2,3</sup>, Raphael Flauger<sup>4</sup>, Daniel Green<sup>4\*</sup>, Anže Slosar<sup>5</sup>, Mariana Vargas-Magaña<sup>6</sup>, Benjamin Wallisch<sup>1,7</sup> and Christophe Yèche<sup>3,8</sup>



## 2. Neutrino wind – coherent scatter

- Velocity of solar system wrt CMB frame  $\beta \approx 10^{-3}$
- **Coherent Acceleration** -  $\sigma_{\nu N} \propto G_F^2$ 
  - ✓ De Broglie  $\lambda \approx 2 - 3\text{mm}$
  - ✓  $a_{NR-D} = O\left(10^{-27} \frac{\text{cm}}{\text{s}^2}\right)$  for non-relativistic Dirac neutrinos
  - ✓  $a_{NR-M} \approx a_{NR-D} \cdot \beta^2 \approx a_{NR-D} \cdot 10^{-6}$

$a \approx 10^{-13} \text{cm/s}^2$  achievable



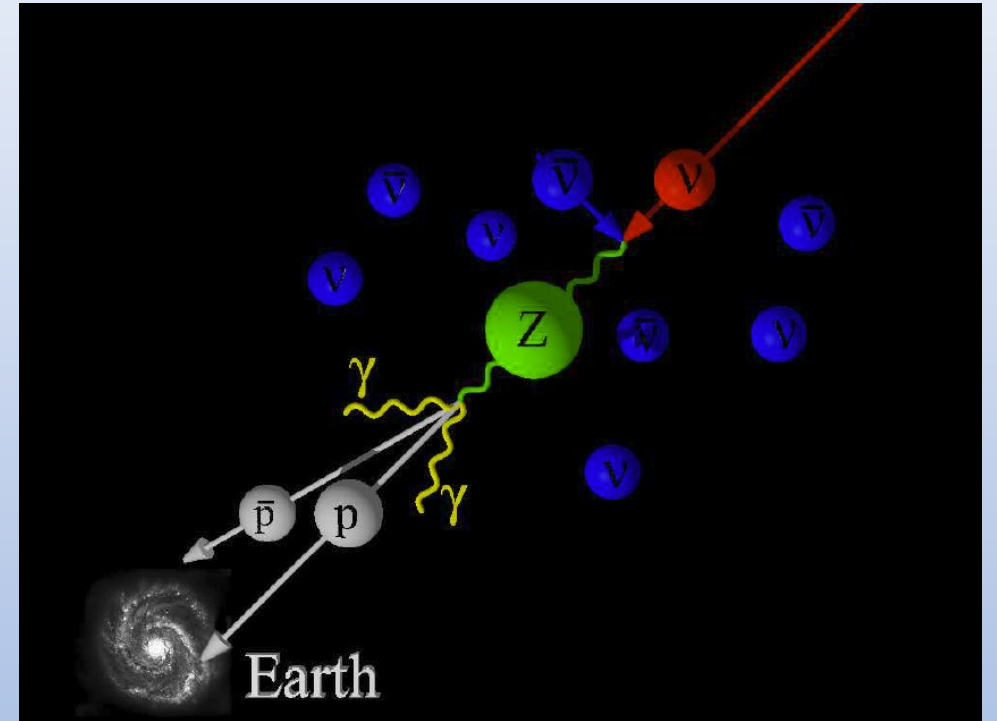
... and worry about solar  $\nu$  and WIMP backgrounds

# 3. Cosmic neutrinos

- Interact with high energy  $\nu$ :

$$E_{\nu_i} \approx 4 \cdot 10^{21} \left( \frac{eV}{m_{\nu_i}} \right) eV$$

- Result:
  1. Dip of high energy  $\nu$  flux
  2. Excess of high energy  $\gamma$ , proton flux



# 5. Induced beta decay

PHYSICAL REVIEW

VOLUME 128, NUMBER 3

NOVEMBER 1, 1962

## Universal Neutrino Degeneracy

STEVEN WEINBERG\*

*Imperial College of Science and Technology, London, England*

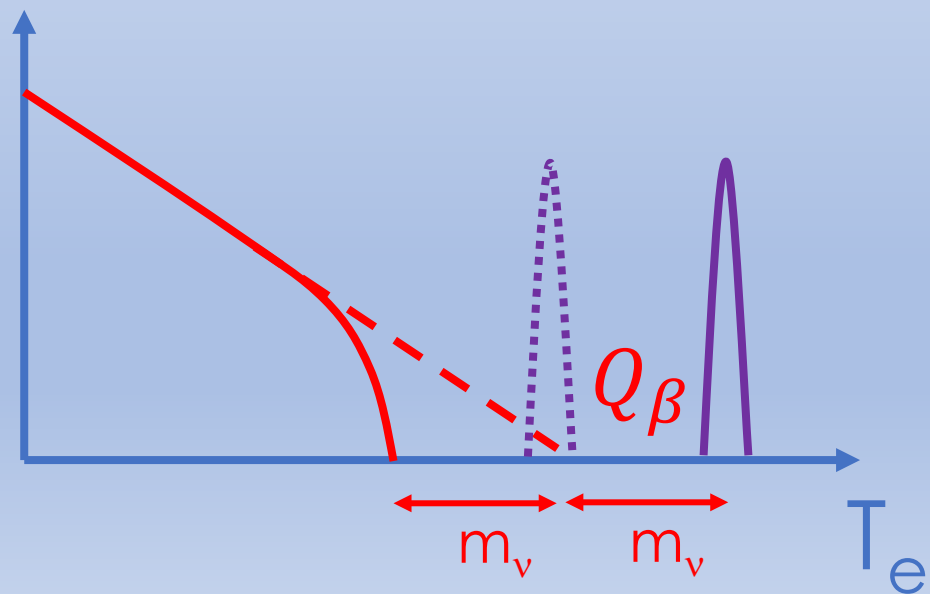
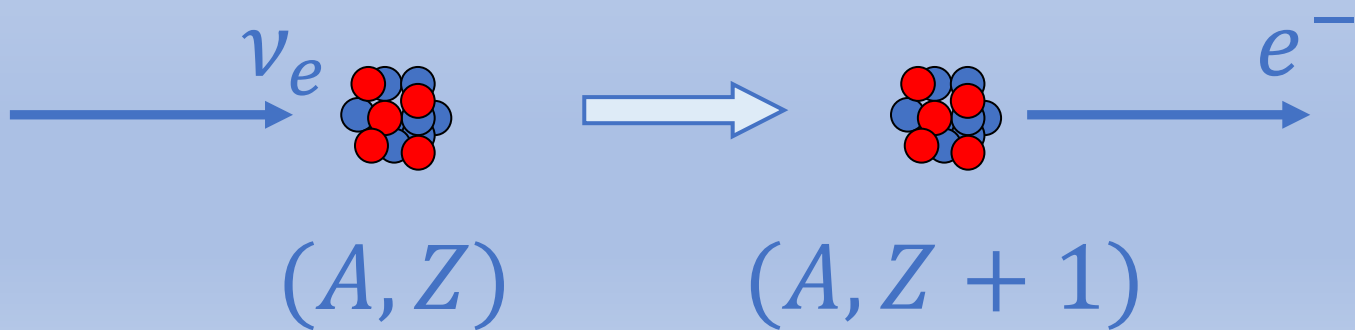
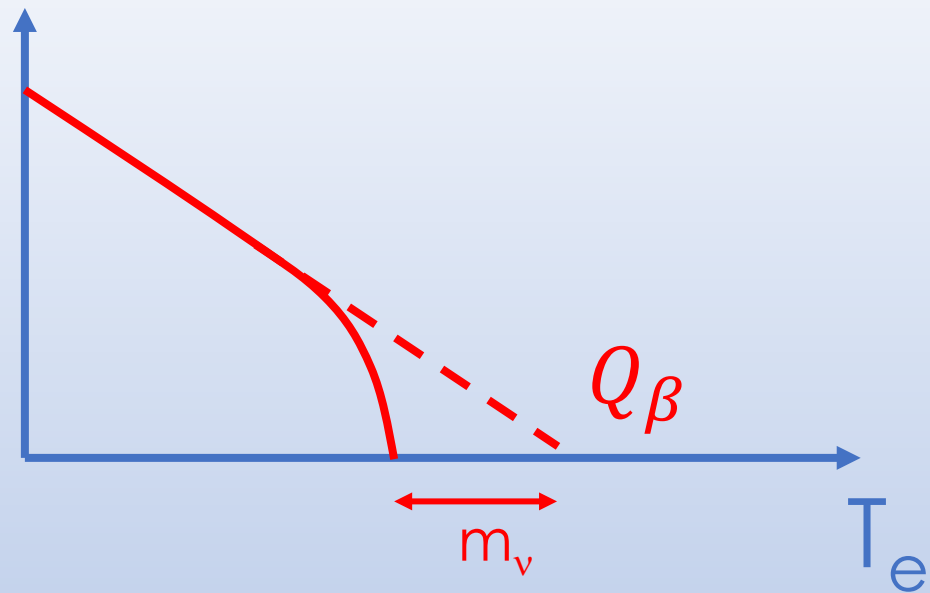
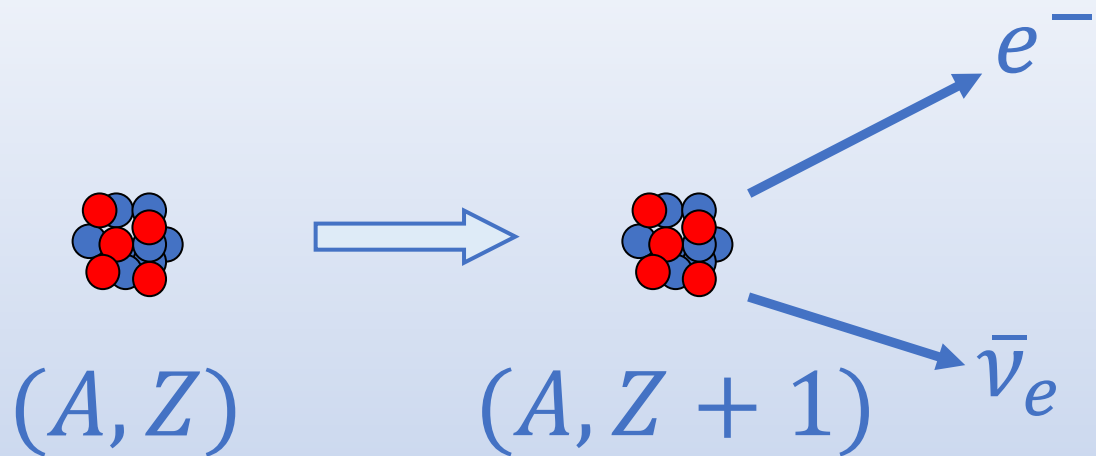
(Received March 22, 1962)

$$m_\nu = 0$$

Modern cosmological theories imply that the universe is filled with a shallow degenerate Fermi sea of neutrinos. In the steady state and oscillating models (and perhaps also the “big bang” theories) it can be shown rigorously that the proportion of filled neutrino levels (plus the proportion of filled antineutrino levels) is precisely one up to a finite Fermi energy  $E_F$ . The proof takes into account both absorption and the repulsive effects of already filled levels on neutrino emission. Experiment shows that  $E_F \leq 200$  eV for antineutrinos and  $E_F \leq 1000$  eV for neutrinos. The degenerate neutrinos could be observed (if  $E_F > 10$  eV) by looking for apparent violations of energy conservation in  $\beta^-$  decay. In the steady state and evolutionary cosmologies  $E_F$  is much too low to ever be observed, but in the oscillating cosmologies  $E_F \simeq 5R_c$  MeV, where  $R_c$  is the minimum radius of the universe in units of its present radius; thus experiment already shows that the universe will contract by a factor over  $10^3$ , if at all. Astronomical evidence plus Einstein’s field equation (without cosmological constant) require in an oscillating cosmology that  $E_F < 2 \times 10^{-3}$  eV (so  $R_c < 10^{-9}$ ) and suggest that higher energy neutrinos may represent the bulk of the energy of the universe. A model universe incorporating this idea is constructed.

Cocco, Mangano, Messina calculated  $m_\nu \neq 0$  case in 2007

<https://arxiv.org/abs/hep-ph/0703075>



# Selection of target

“Highish” cross-section

“Longish” lifetime

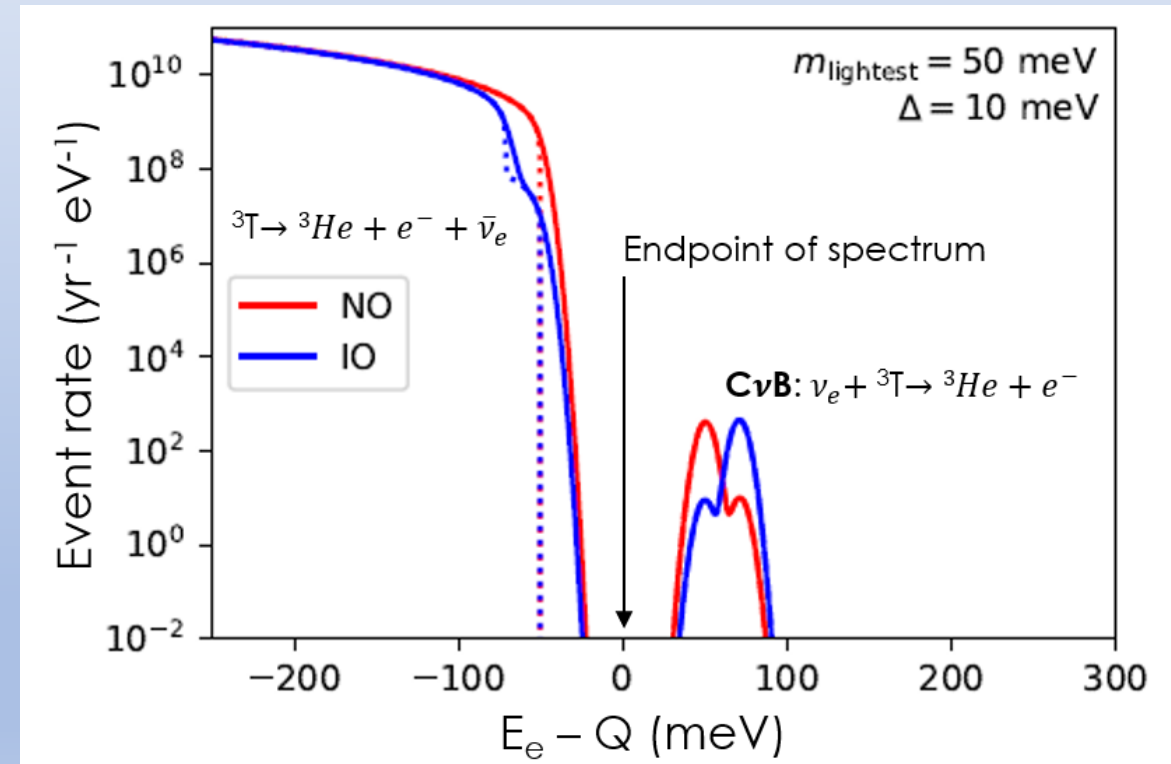
Isotope	Decay	$Q_\beta$ (keV)	Half-life (sec)	$\sigma_{\text{NCB}}(v_\nu/c)$ ( $10^{-41}$ cm <sup>2</sup> )
<sup>3</sup> H	$\beta^-$	18.591	$3.8878 \times 10^8$	$7.84 \times 10^{-4}$
<sup>63</sup> Ni	$\beta^-$	66.945	$3.1588 \times 10^9$	$1.38 \times 10^{-6}$
<sup>93</sup> Zr	$\beta^-$	60.63	$4.952 \times 10^{13}$	$2.39 \times 10^{-10}$
<sup>106</sup> Ru	$\beta^-$	39.4	$3.2278 \times 10^7$	$5.88 \times 10^{-4}$
<sup>107</sup> Pd	$\beta^-$	33	$2.0512 \times 10^{14}$	$2.58 \times 10^{-10}$
<sup>187</sup> Re	$\beta^-$	2.64	$1.3727 \times 10^{18}$	$4.32 \times 10^{-11}$
<sup>11</sup> C	$\beta^+$	960.2	$1.226 \times 10^3$	$4.66 \times 10^{-3}$
<sup>13</sup> N	$\beta^+$	1198.5	$5.99 \times 10^2$	$5.3 \times 10^{-3}$
<sup>15</sup> O	$\beta^+$	1732	$1.224 \times 10^2$	$9.75 \times 10^{-3}$
<sup>18</sup> F	$\beta^+$	633.5	$6.809 \times 10^3$	$2.63 \times 10^{-3}$
<sup>22</sup> Na	$\beta^+$	545.6	$9.07 \times 10^7$	$3.04 \times 10^{-7}$
<sup>45</sup> Ti	$\beta^+$	1040.4	$1.307 \times 10^4$	$3.87 \times 10^{-4}$



Tritium



- High cross-section for neutrino capture
- **No energy threshold**
- Sizeable lifetime
- Low Q-value of 18.6 keV
- **Tritium beta decay  $\sim 10^{15}$  Bq/gram**



# Expected rate: 100gram-year exposure

$m_\nu$ (eV)	FD (events yr <sup>-1</sup> )	NFW (events yr <sup>-1</sup> )	MW (events yrs <sup>-1</sup> )
0.6	7.5	90	150
0.3	7.5	23	33
0.15	7.5	10	12

Dirac

$m_\nu$ (eV)	FD (events yr <sup>-1</sup> )	NFW (events yr <sup>-1</sup> )	MW (events yrs <sup>-1</sup> )
0.6	7.5	90	150
0.3	7.5	23	33
0.15	7.5	10	12

Majorana

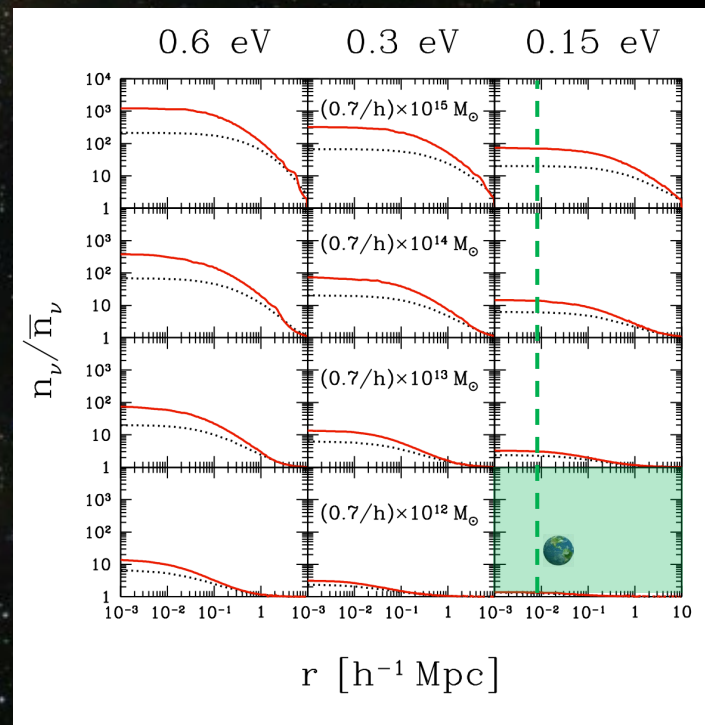
<https://arxiv.org/abs/hep-ph/0703075>

<https://arxiv.org/abs/1405.7654>



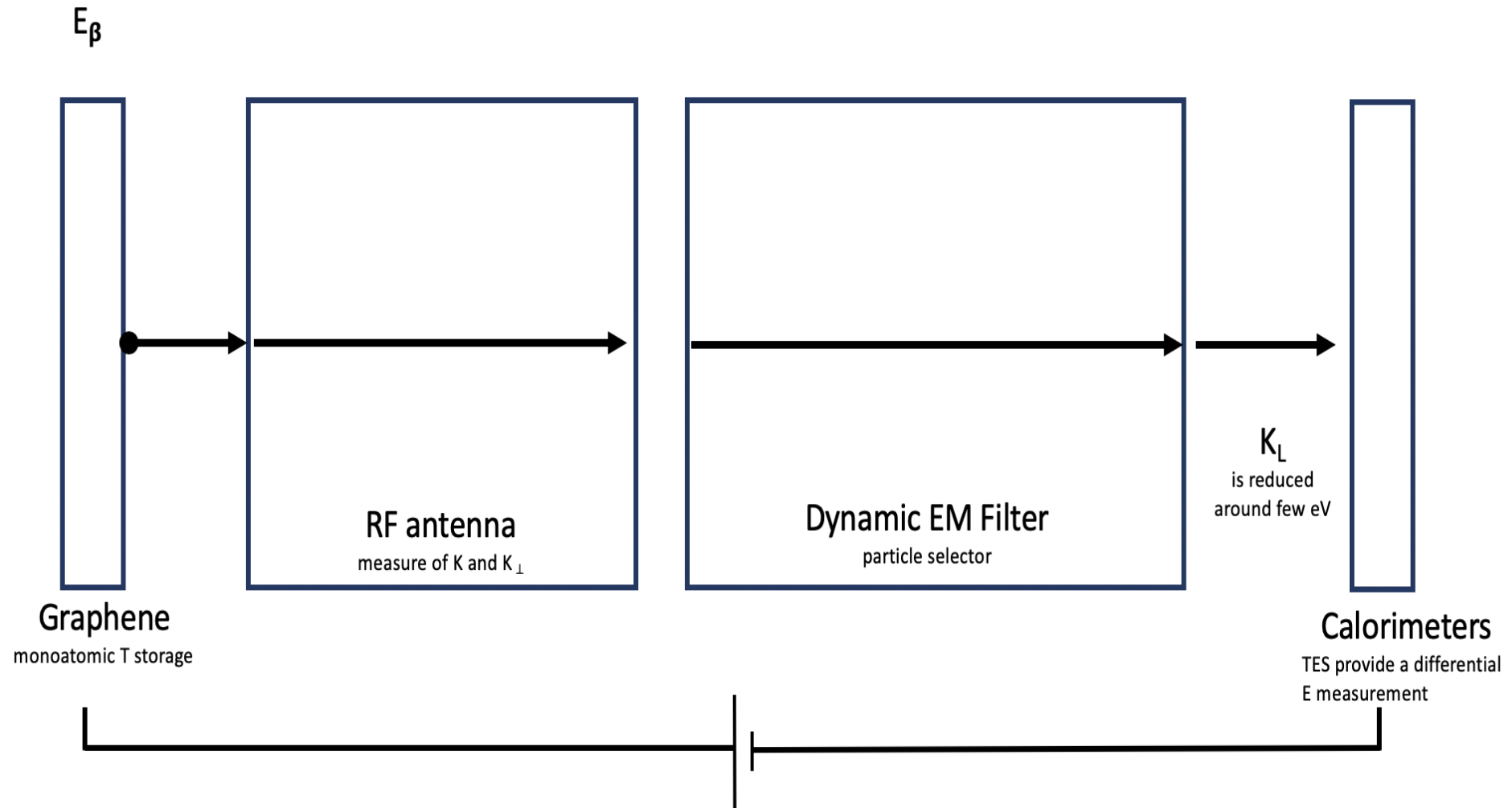
d rate

### Gravitational clustering



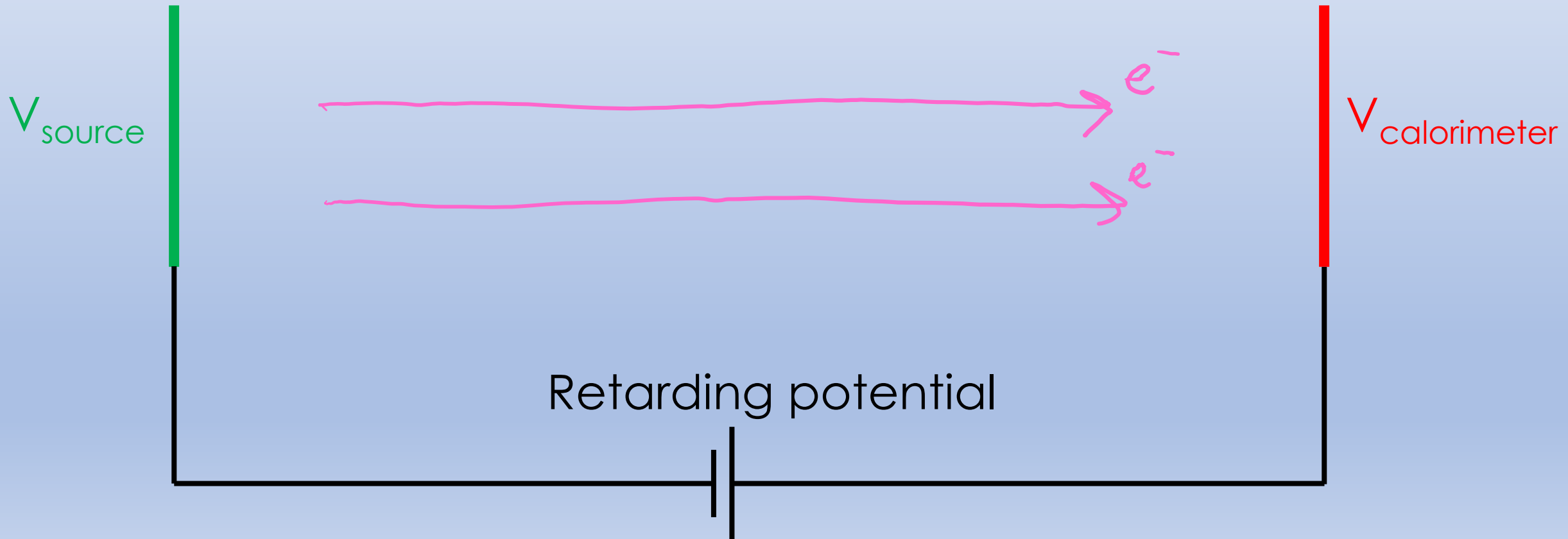


# PTOLEMY experiment - concept



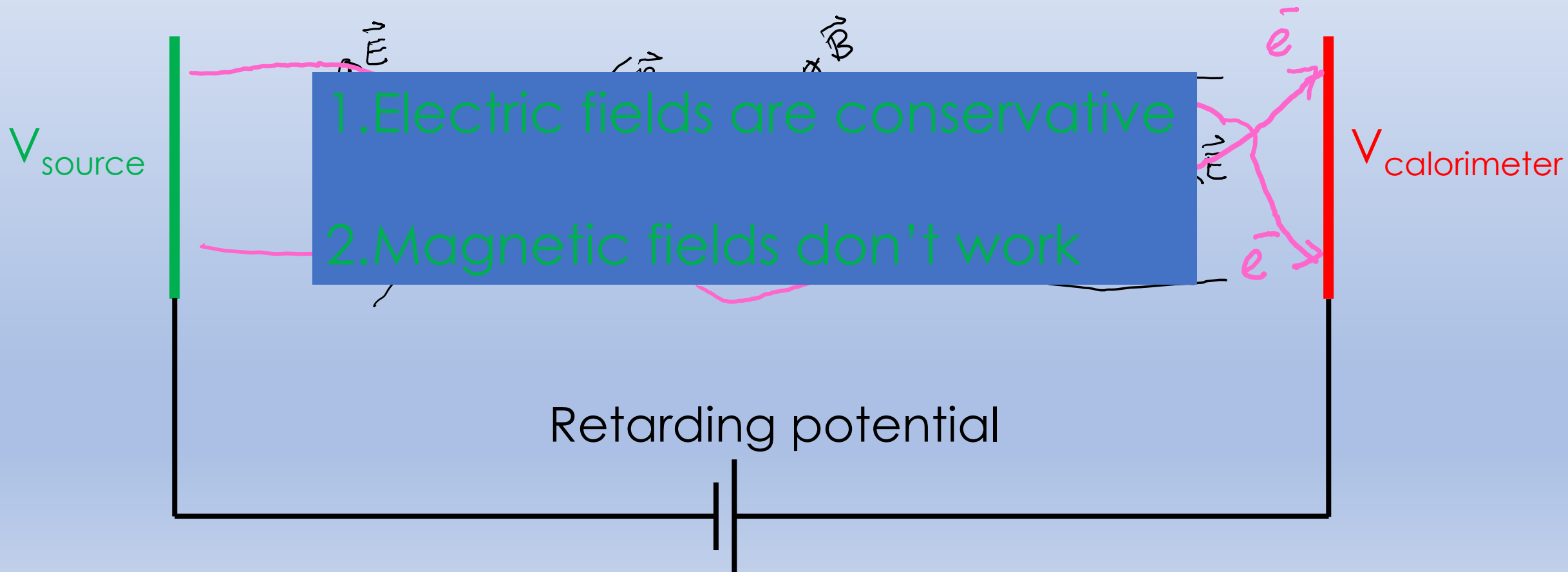
# PTOLEMY experiment - concept

$$E_e = e (V_{calorimeter} - V_{source}) + E_{calorimeter}$$



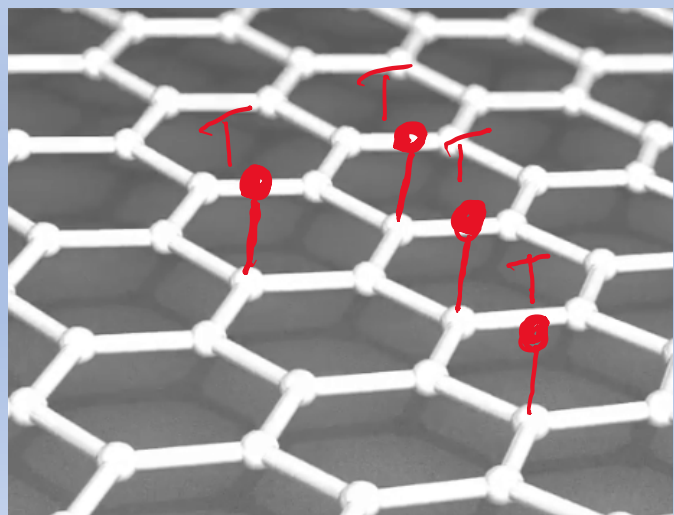
# PTOLEMY experiment - concept

$$E_e = e (V_{calorimeter} - V_{source}) + E_{calorimeter}$$

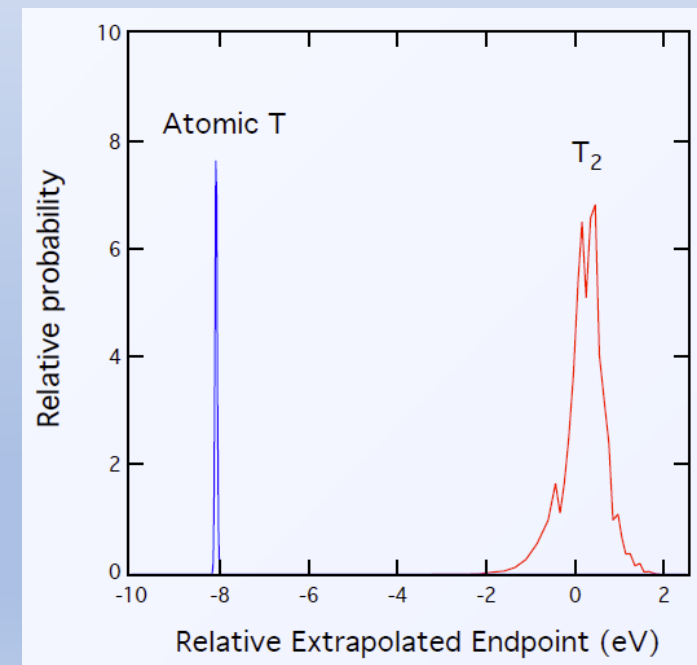
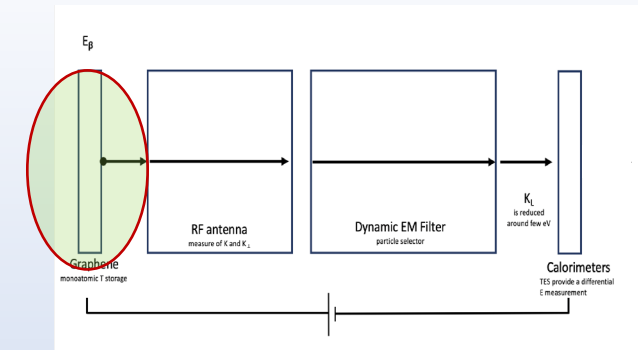


# PTOLEMY: tritium target

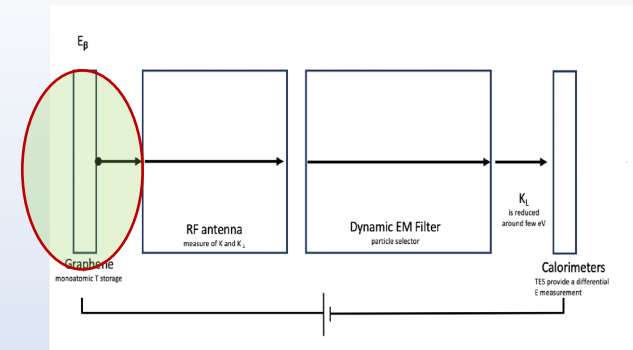
- Use **atomic  $^3\text{T}$** 
  - No ro-vibrational modes in final state like for  $^3\text{He}$ - $^3\text{T}$  final state.
  - Limit to energy resolution not determined by target itself
- $dE/dx$  of electrons requires extremely **thin targets**
- We investigate  $^3\text{T}$  loosely bound to graphene
  - Theoretical maximum is about 0.2 mg tritium per  $\text{m}^2$



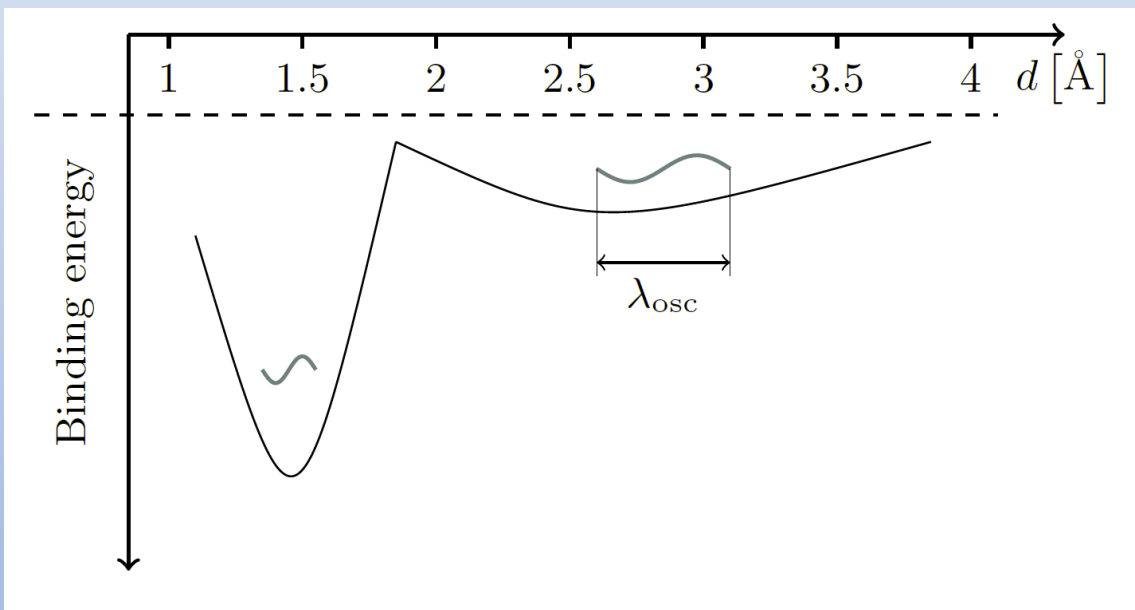
Max 1-tritium for every C



# Trouble with Heisenberg?



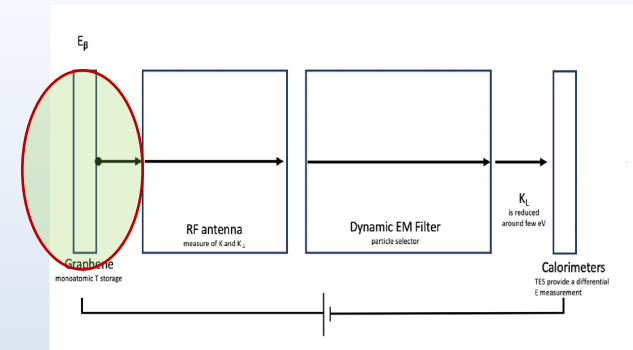
Binding  ${}^3\text{T}$  to graphene = localizing  ${}^3\text{T} \rightarrow \Delta p \Delta x \geq \frac{\hbar}{2} \rightarrow$  energy spread



Potential	Source	$\kappa$ , [eV/Å <sup>2</sup> ]	$\lambda$ , [Å]	$\Delta E$ , [eV]
Chemisorption	[15]	2.15	0.16	0.60
	[13], GGA	4.62	0.13	0.73
	[13], vdW-DF	4.9	0.13	0.75
Physisorption	[16]	0.08	0.37	0.26
	[15]	0.09	0.34	0.28
	[13], GGA	0.18	0.29	0.33
	[13], vdW-DF	0.13	0.32	0.3
	[14], GGA	0.04	0.43	0.22
	[14], LDA	0.01	0.55	0.17
Migration	[18]	0.283	0.264	0.37

$\Delta E \approx 100 - 400$  meV (?)

# Ways to solve?

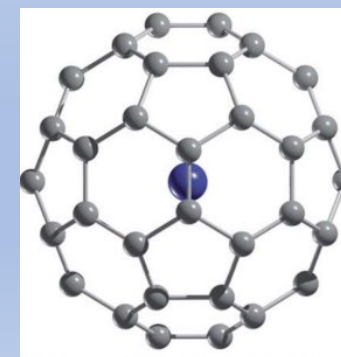
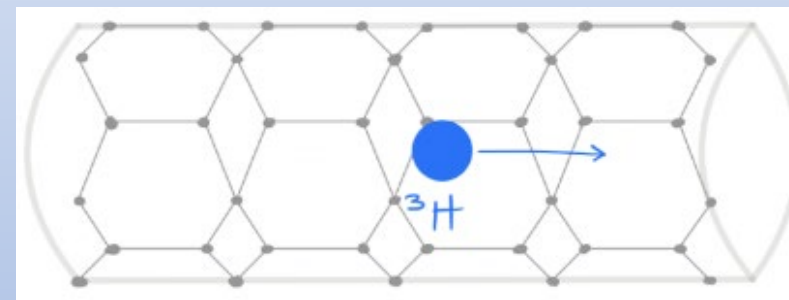


**Strategy 1:** find other target material

1 H Hydrogen																	2 He Helium																												
3 Li Lithium	4 Be Beryllium											5 B Boron	6 C Carbon	7 N Nitrogen	8 O Oxygen	9 F Fluorine	10 Ne Neon																												
11 Na Sodium	12 Mg Magnesium											13 Al Aluminium	14 Si Silicon	15 P Phosphorus	16 S Sulfur	17 Cl Chlorine	18 Ar Argon																												
19 K Potassium	20 Ca Calcium	21 Sc Scandium	22 Ti Titanium	23 V Vanadium	24 Cr Chromium	25 Mn Manganese	26 Fe Iron	27 Co Cobalt	28 Ni Nickel	29 Cu Copper	30 Zn Zinc	31 Ga Gallium	32 Ge Germanium	33 As Arsenic	34 Se Selenium	35 Br Bromine	36 Kr Krypton																												
37 Rb Rubidium	38 Sr Strontium	39 Y Yttrium	40 Zr Zirconium	41 Nb Niobium	42 Mo Molybdenum	43 Tc Technetium	44 Ru Ruthenium	45 Rh Rhodium	46 Pd Palladium	47 Ag Silver	48 Cd Cadmium	49 In Indium	50 Sn Tin	51 Sb Antimony	52 Te Tellurium	53 I Iodine	54 Xe Xenon																												
55 Cs Cesium	56 Ba Barium	57 La Lanthanum	72 Hf Hafnium	73 Ta Tantalum	74 W Tungsten	75 Re Rhenium	76 Os Osmium	77 Ir Iridium	78 Pt Platinum	79 Au Gold	80 Hg Mercury	81 Tl Thallium	82 Pb Lead	83 Bi Bismuth	84 Po Polonium	85 At Astatine	86 Rn Radon																												
87 Fr Francium	88 Ra Radium	89 Ac Actinium	104 Rf Rutherfordium	105 Db Dubnium	106 Sg Seaborgium	107 Bh Bohrium	108 Hs Hassium	109 Mt Meitnerium	110 Ds Darmstadtium	111 Rg Roentgenium	112 Cn Copernicium	113 Nh Nihonium	114 Fl Flerovium	115 Mc Moscovium	116 Lv Livermorium	117 Ts Tennessine	118 Og Oganesson																												
<table border="1"> <tbody> <tr> <td>58 Ce Cerium</td> <td>59 Pr Praseodymium</td> <td>60 Nd Neodymium</td> <td>61 Pm Promethium</td> <td>62 Sm Samarium</td> <td>63 Eu Europium</td> <td>64 Gd Gadolinium</td> <td>65 Tb Terbium</td> <td>66 Dy Dysprosium</td> <td>67 Ho Holmium</td> <td>68 Er Erbium</td> <td>69 Tm Thulium</td> <td>70 Yb Ytterbium</td> <td>71 Lu Lutetium</td> </tr> <tr> <td>90 Th Thorium</td> <td>91 Pa Protactinium</td> <td>92 U Uranium</td> <td>93 Np Neptunium</td> <td>94 Pu Plutonium</td> <td>95 Am Americium</td> <td>96 Cm Curium</td> <td>97 Bk Berkelium</td> <td>98 Cf Californium</td> <td>99 Es Einsteinium</td> <td>100 Fm Fermium</td> <td>101 Md Mendelevium</td> <td>102 No Nobelium</td> <td>103 Lr Lawrencium</td> </tr> </tbody> </table>																		58 Ce Cerium	59 Pr Praseodymium	60 Nd Neodymium	61 Pm Promethium	62 Sm Samarium	63 Eu Europium	64 Gd Gadolinium	65 Tb Terbium	66 Dy Dysprosium	67 Ho Holmium	68 Er Erbium	69 Tm Thulium	70 Yb Ytterbium	71 Lu Lutetium	90 Th Thorium	91 Pa Protactinium	92 U Uranium	93 Np Neptunium	94 Pu Plutonium	95 Am Americium	96 Cm Curium	97 Bk Berkelium	98 Cf Californium	99 Es Einsteinium	100 Fm Fermium	101 Md Mendelevium	102 No Nobelium	103 Lr Lawrencium
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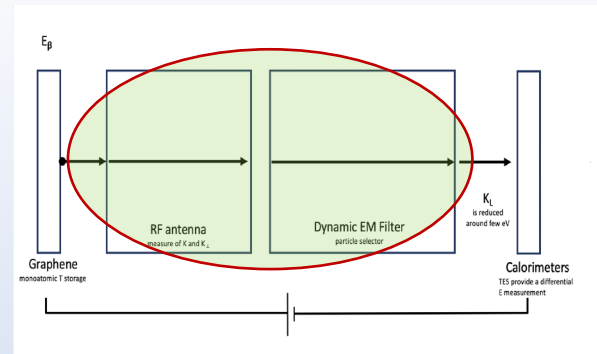
from Mendeleev et al

**Strategy 2:** alternative  $^3\text{H}$  storage

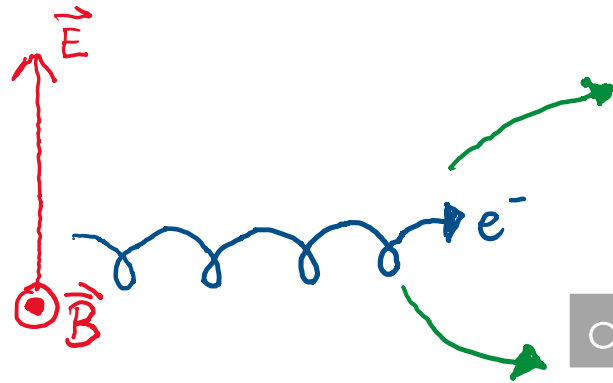


from A. Esposito

# PTOLEMY: two types of drift



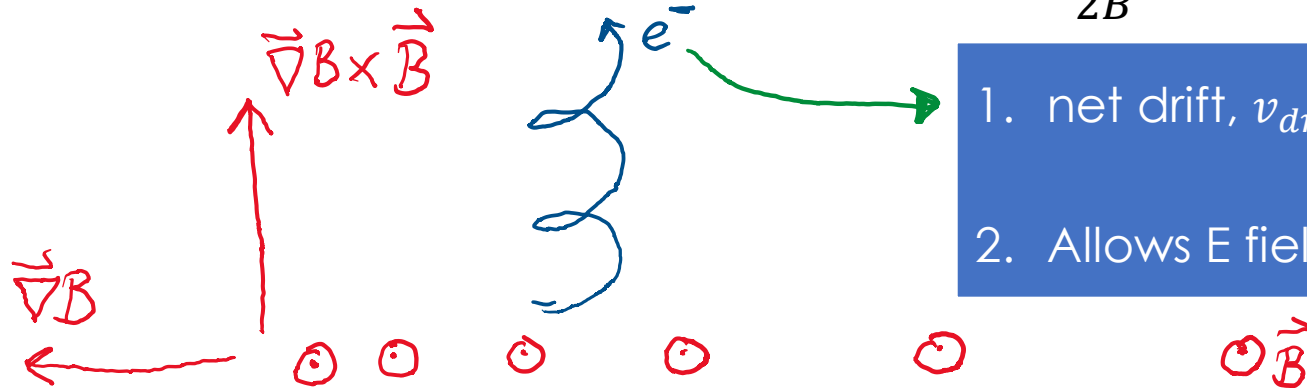
I:  $\vec{E} \times \vec{B}$  drift



1. net drift,  $v_{drift} = E/B$
2. no work, drift along equipotential planes

cyclotron motion – detectable RF

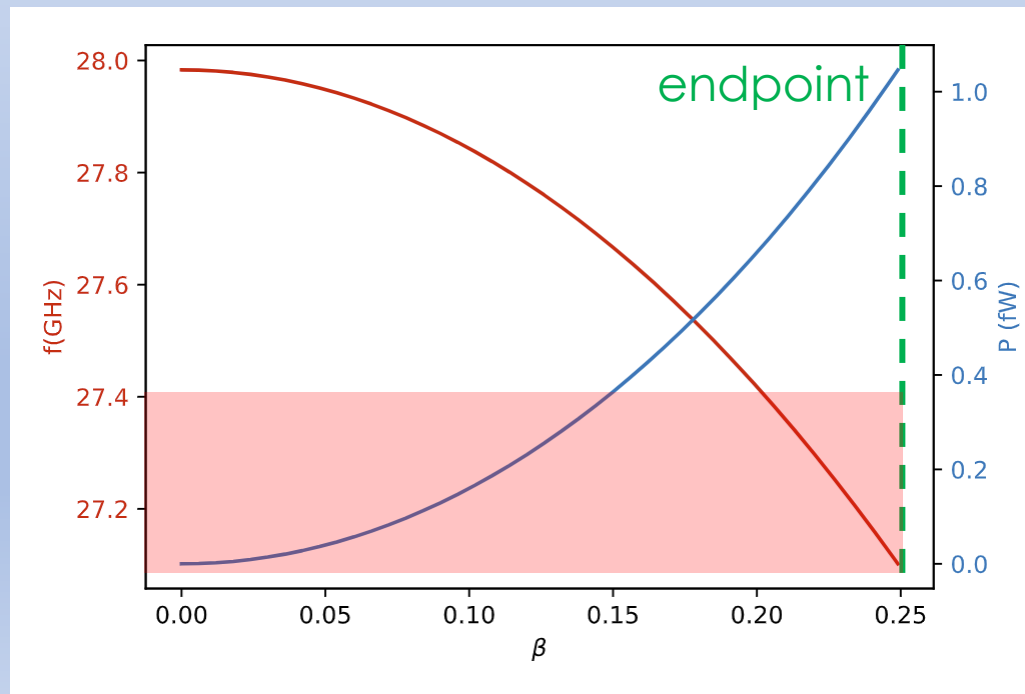
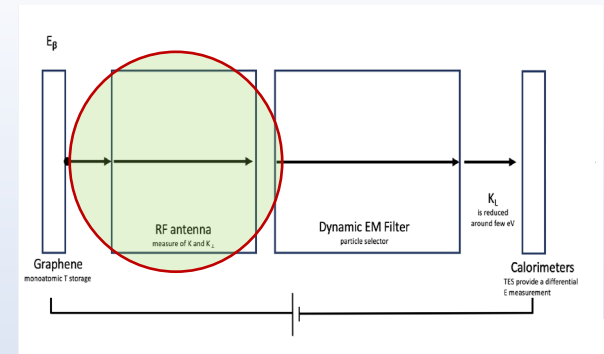
II:  $\frac{\mu}{B^2} \vec{\nabla} B \times \vec{B}$  drift, with magnetic moment  $\mu = \frac{m_e v_{\perp}^2}{2B}$



1. net drift,  $v_{drift} = \mu \frac{|\vec{\nabla} B|}{B}$
2. Allows E field to work (!):  $\frac{dT_{\perp}}{dt} = e\vec{E} \cdot \vec{v}_{drift}$

# PTOLEMY: RF pickup

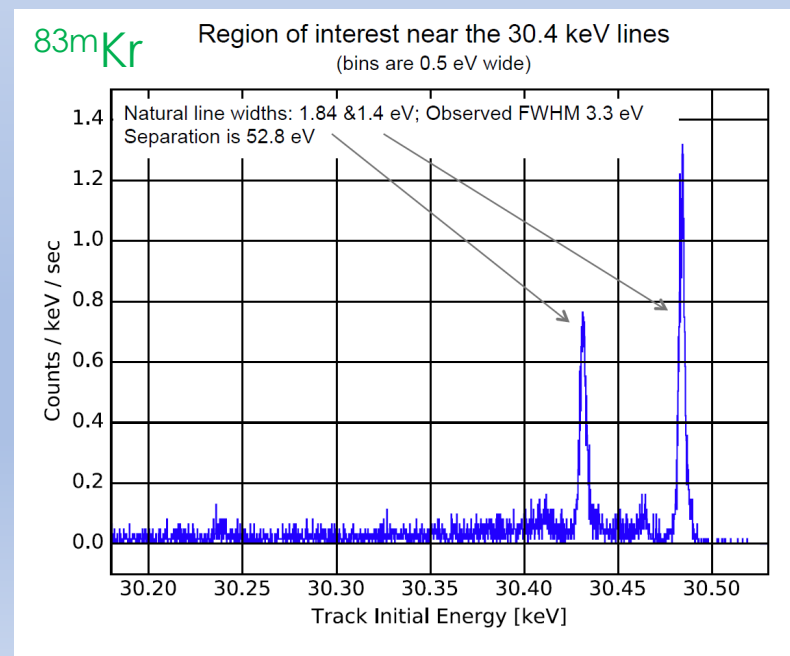
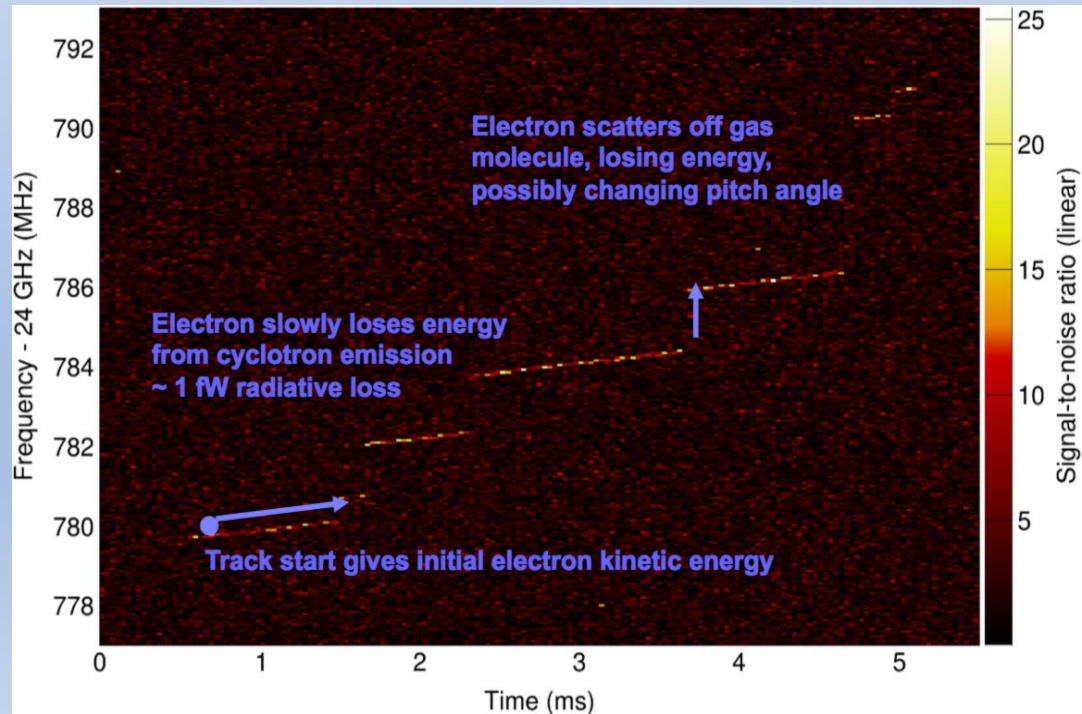
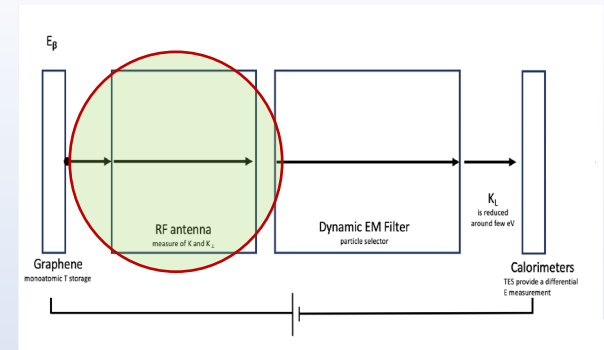
- Transport electrons through  **$\mathbf{E} \times \mathbf{B}$**  field
- RF emission with  $f = \frac{1}{2\pi} \frac{eB}{m_e \gamma} \approx 27 \text{ GHz}$
- Power  $P = \frac{1}{4\pi\epsilon_0} \frac{2e^4 B^2}{3m_e^2 c} (\gamma^2 - 1) \sin^2 \theta \approx 1 \text{ fW} \rightarrow \theta$  is angle between  $\mathbf{B}$  and  $\vec{\beta}$





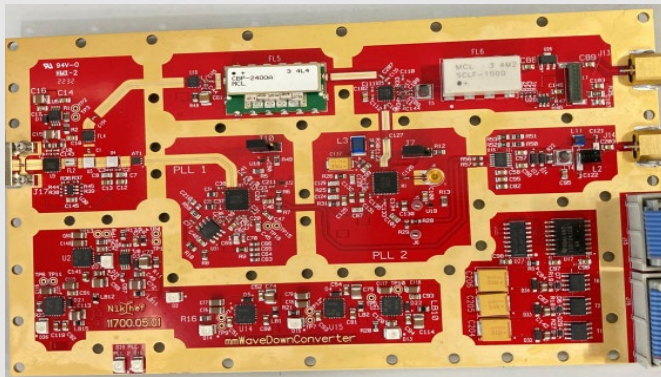
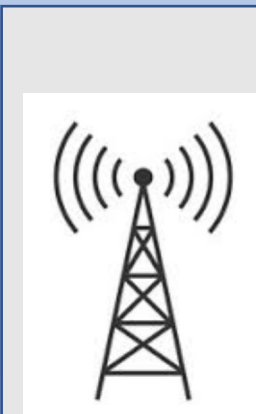
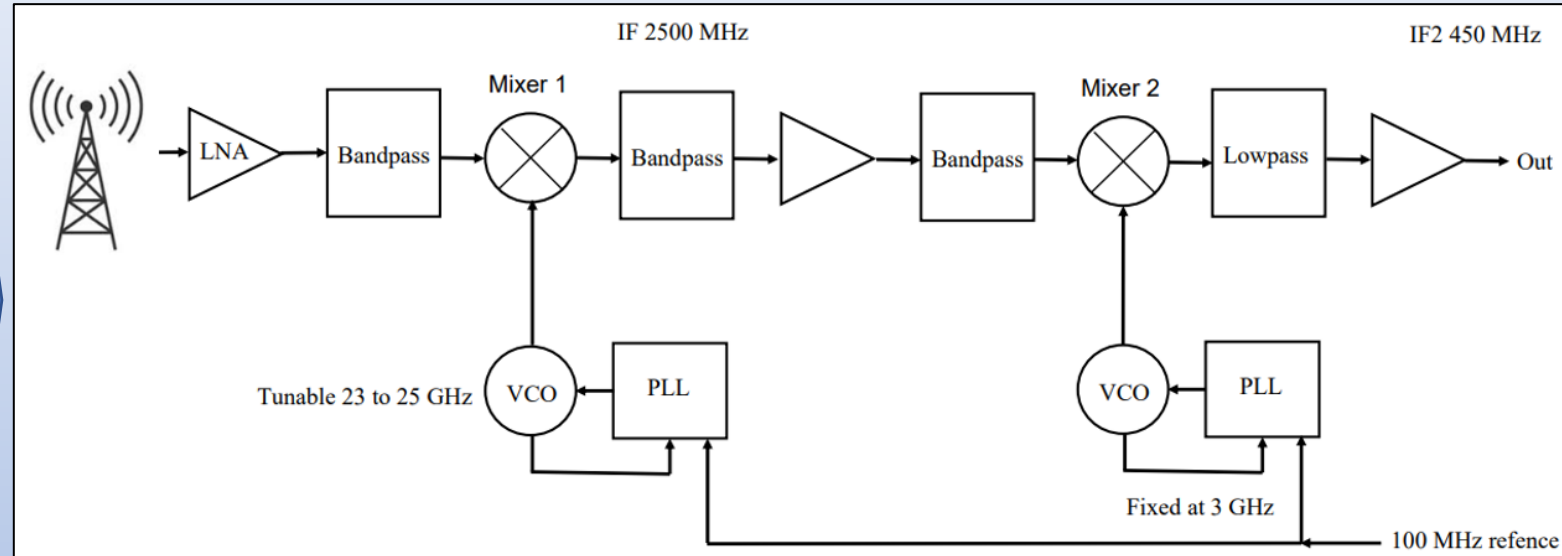
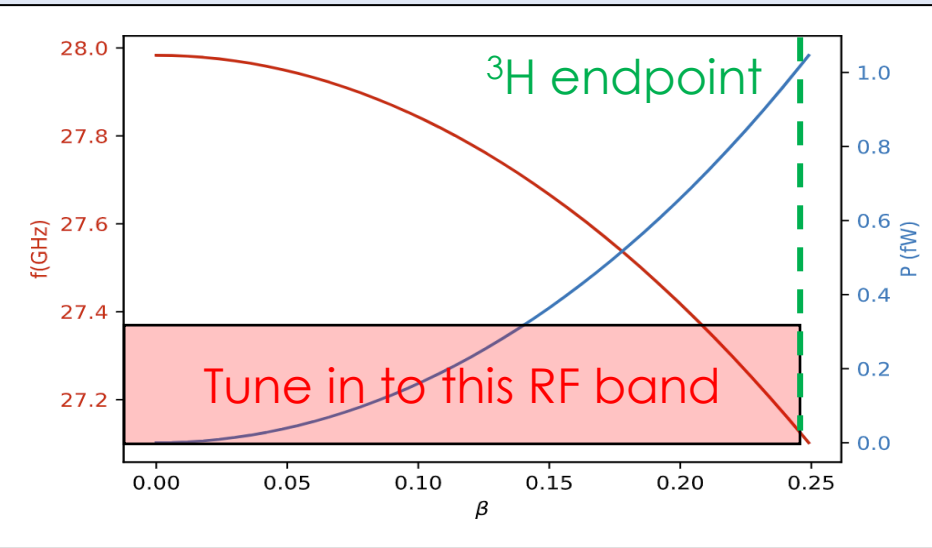
# PTOLEMY: RF pickup

- Transport electrons through  $\mathbf{ExB}$  field
- RF emission with  $f = \frac{1}{2\pi} \frac{eB}{m_e \gamma} \approx 27 \text{GHz} \rightarrow \gamma = \text{Energy}$
- Power  $P = \frac{1}{4\pi\epsilon_0} \frac{2e^4 B^2}{3m_e^2 c} (\gamma^2 - 1) \sin^2 \theta \approx 1 \text{fW} \rightarrow \theta$  is angle between  $\mathbf{B}$  and  $\vec{\beta}$



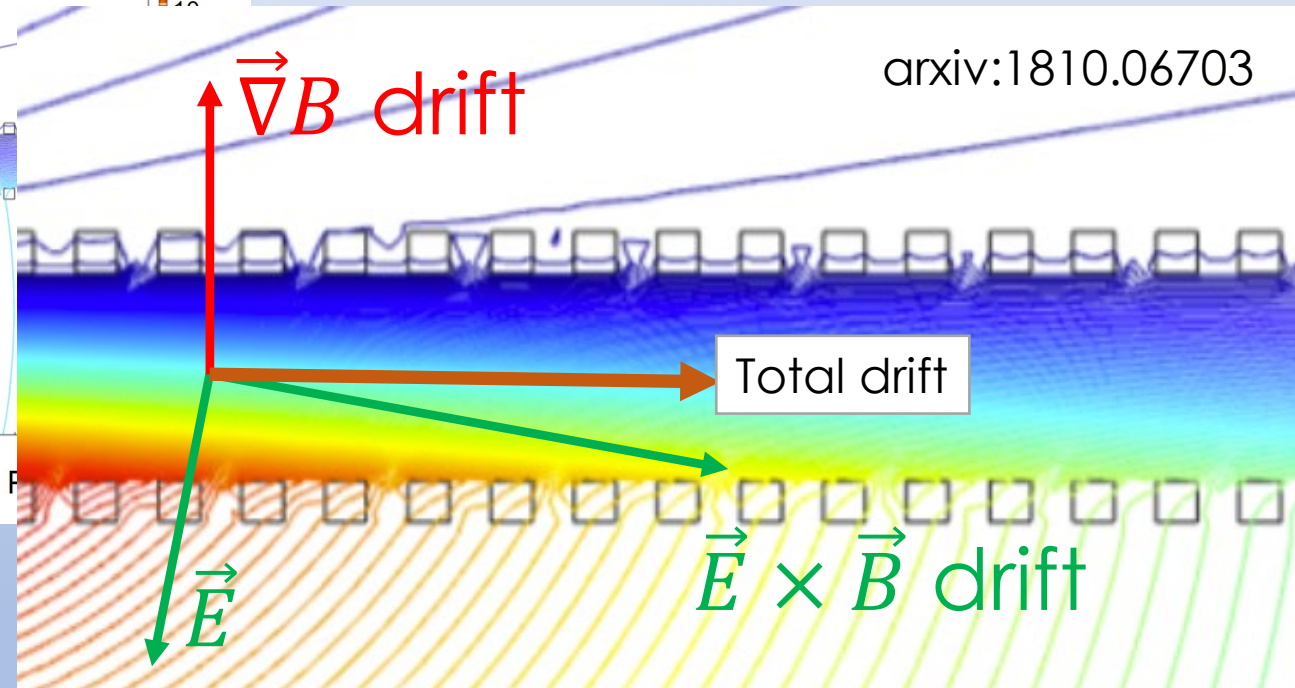
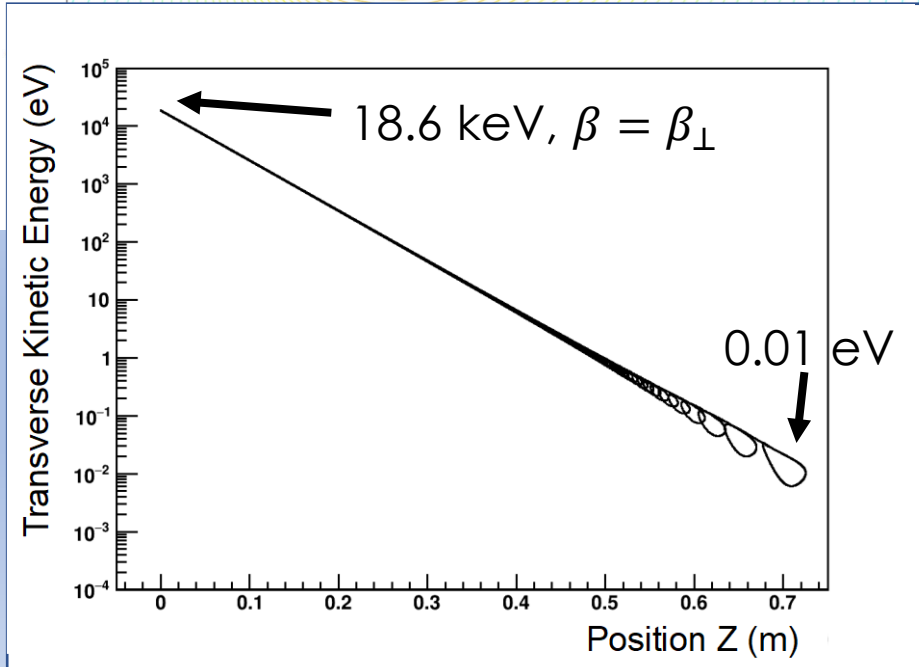
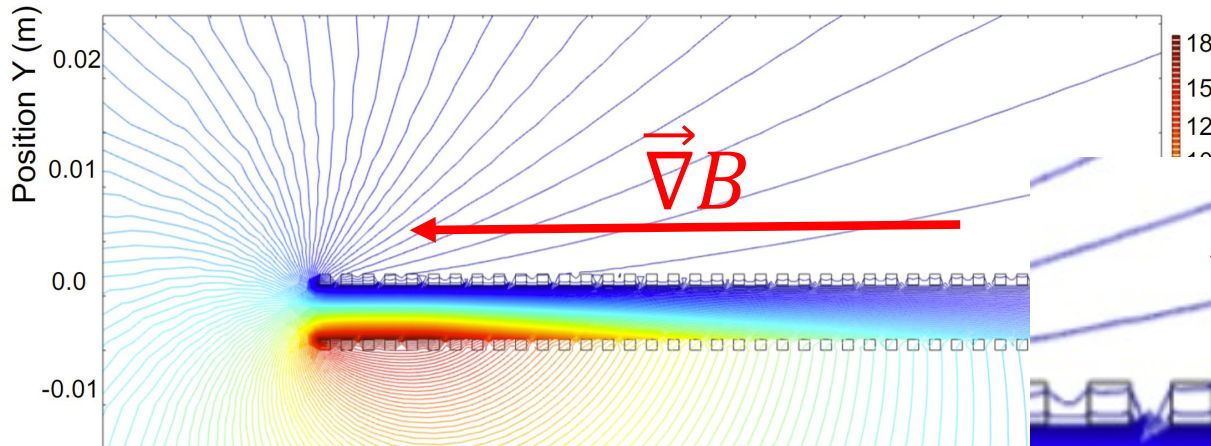
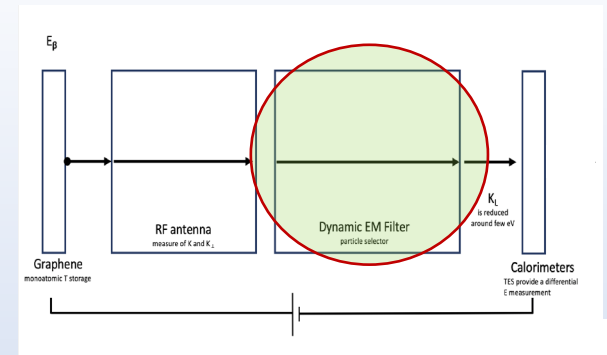
Pioneering work by Project8 @ FNAL  
 A.Esfahani et. al, JPG Vol44,#5, 2017  
<https://arxiv.org/abs/1703.02037v1>

# PTOLEMY: RF pickup @ Nikhef



Trigger HV

# PTOLEMY: Transverse drift filter



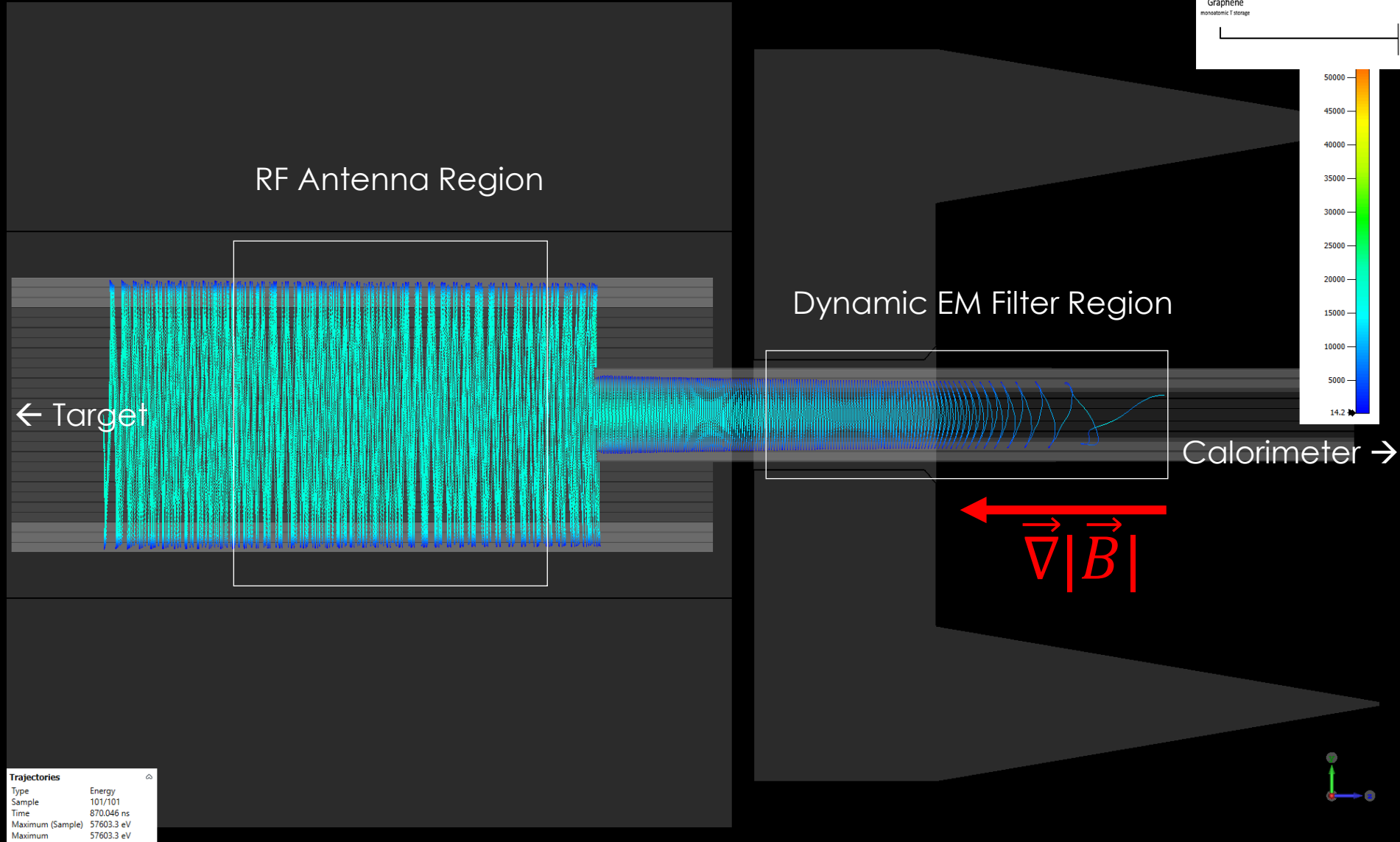
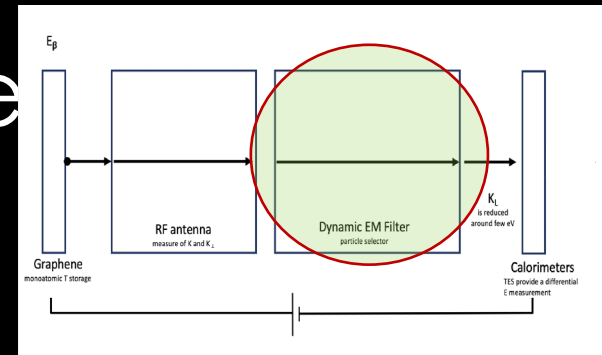
arxiv:1810.06703

After filter, only component of  $\beta$  parallel to B is left.

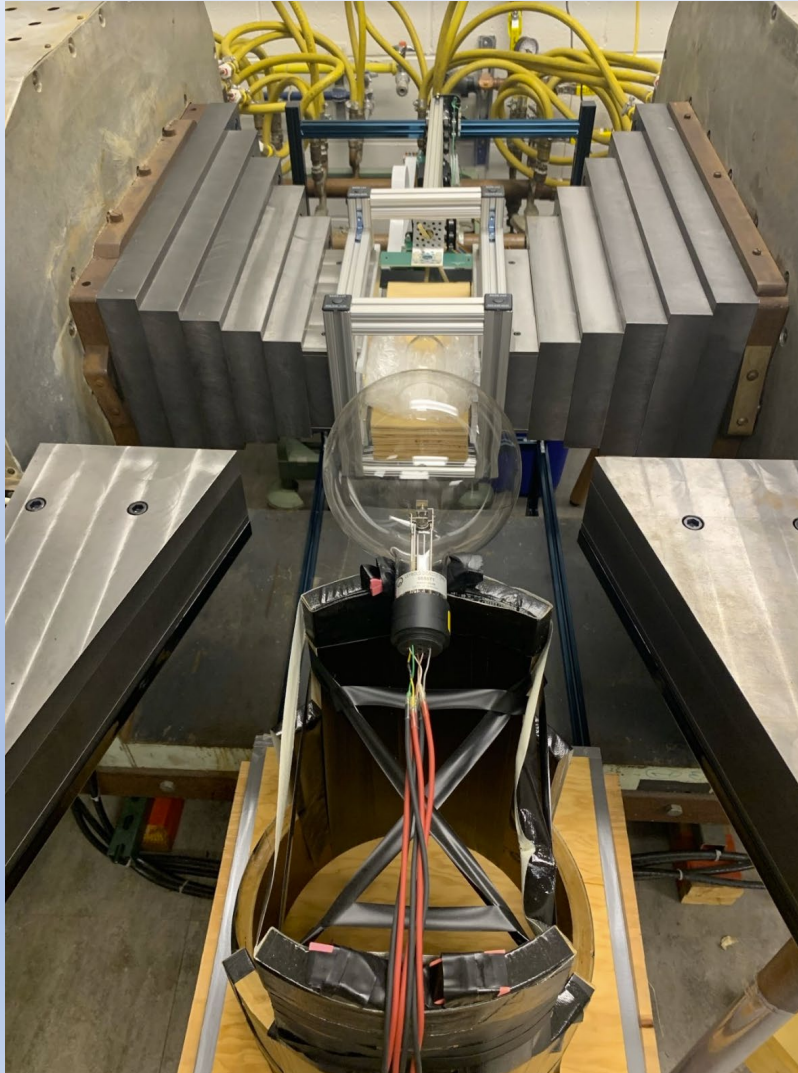
Can be reduced by retarding potential



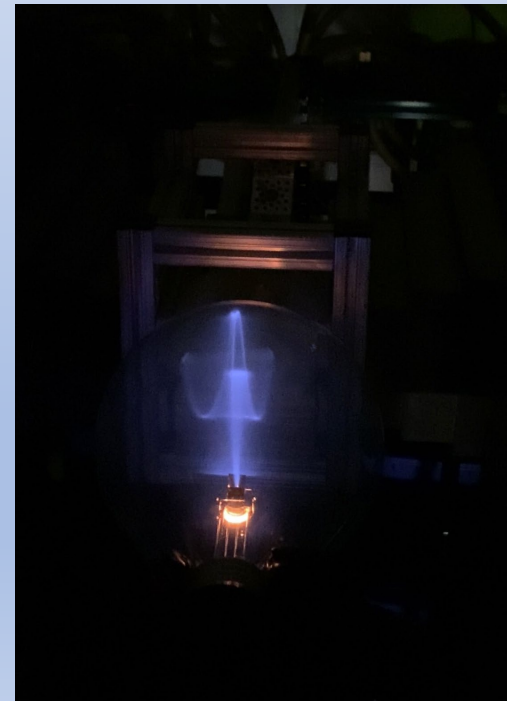
# Electron Transport: RF pickup & Filter



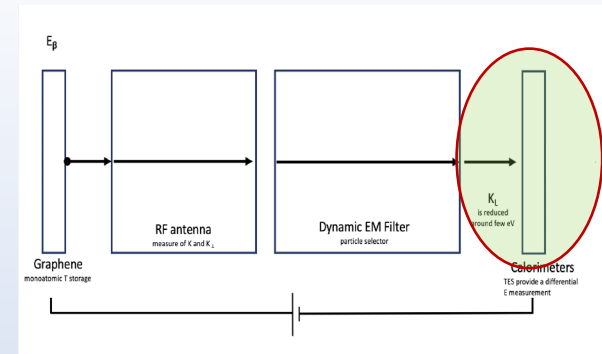
# PTOLEMY: RF + transverse drift filter



- US – Italian – Dutch enterprise to make demonstrator setup
- Expect 'something' under construction at LNGS in 2023



# PTOLEMY: Energy



- Energy measurement from  $\Delta V$  and calorimeter:

$$E_e = e (V_{cal} - V_{source}) + E_{RF} + E_{cal}$$

- Calorimeter energy resolution must be  $O(50\text{meV})$ 
  1. Transition Edge Sensors
  2. State-of-the-art 202x  $O(100\text{meV}@100\text{eV})$
- Voltage stability over experiment better than 10-20mV
- NOTE: internal voltages are actively adjusted for each interesting electron

# PTOLEMY: “expected” performance

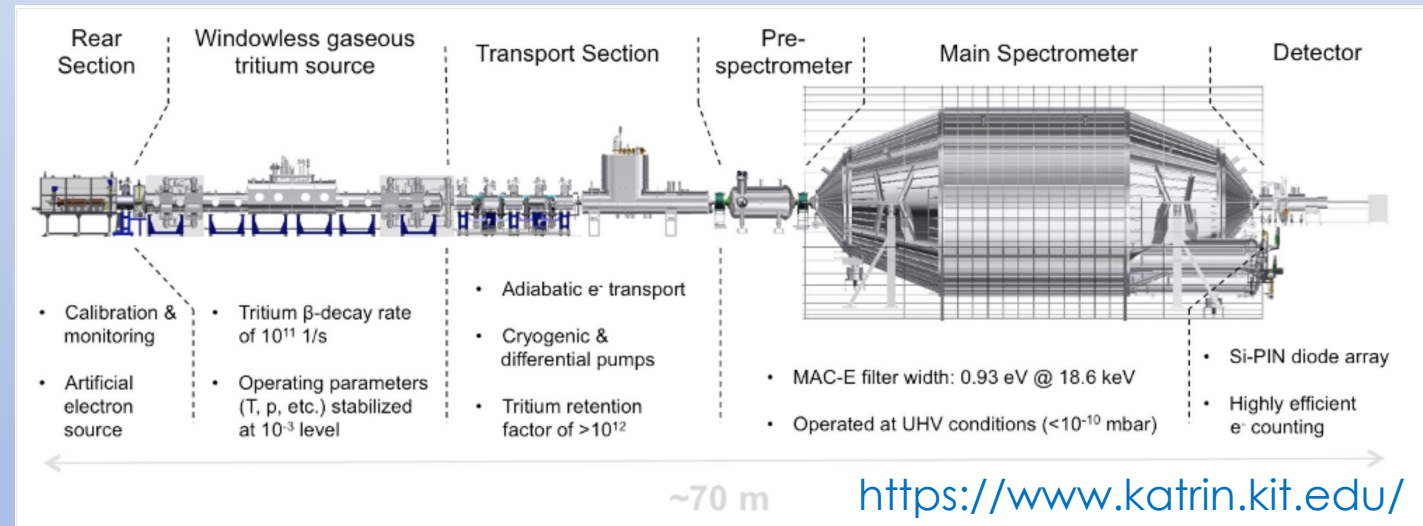
1. Sensitivity to  $m_\nu$

2. Sensitivity to  $C\nu B$

3. Astronomy with  $C\nu B$

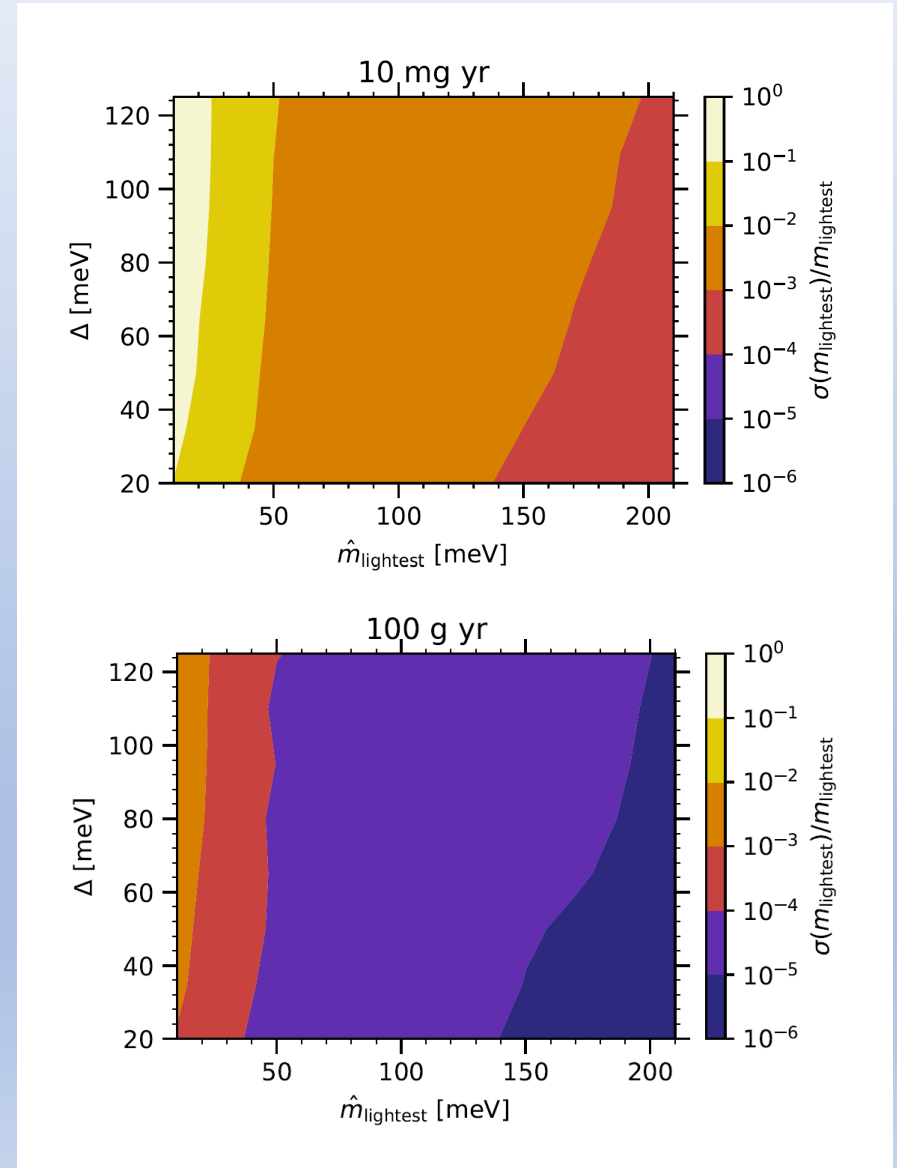
**PLEASE NOTE**

KATRIN = experiment with street credibility



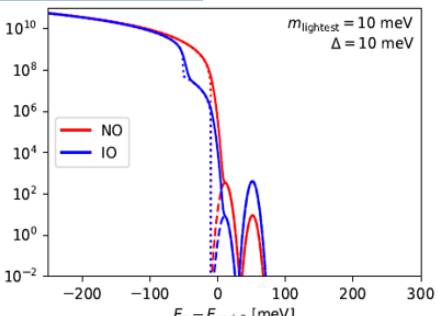
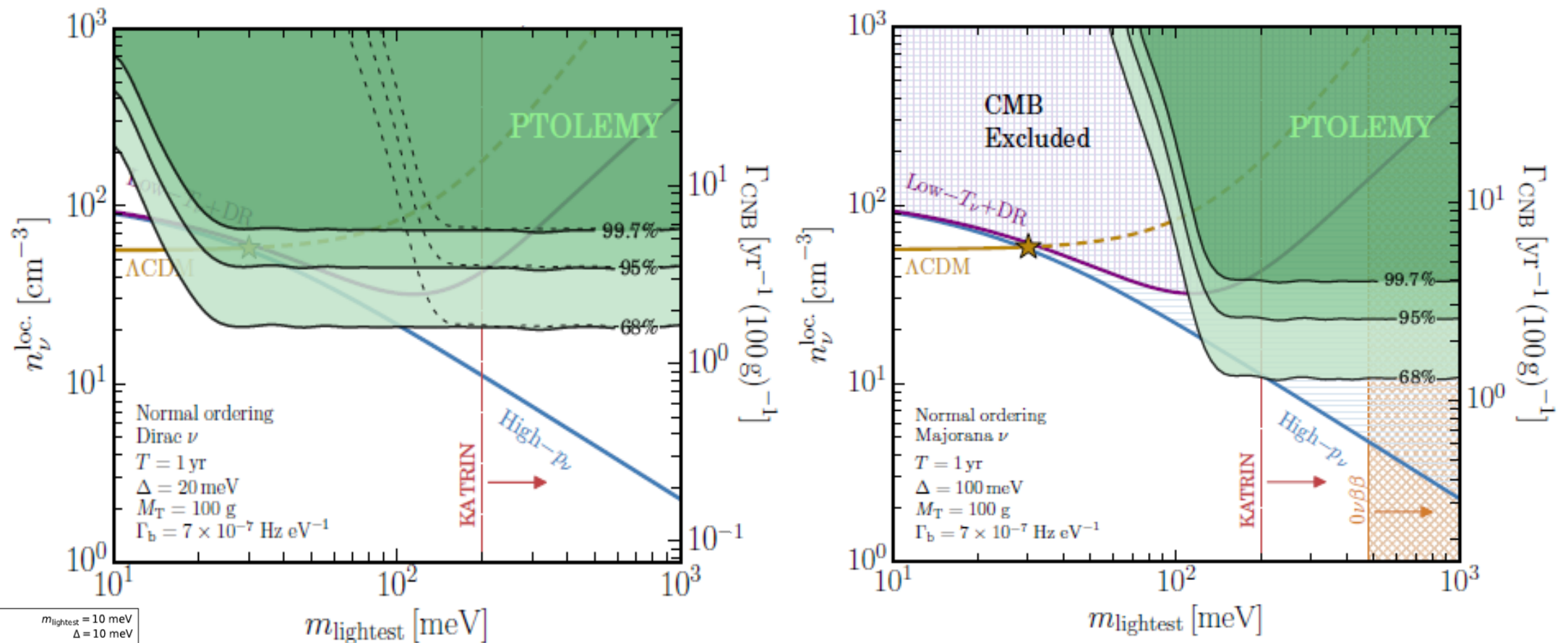
# PTOLEMY: $m_\nu$ “expected” performance

- Neutrino mass as first result
  1. Small exposure already gives sensitivity to  $O(10\text{meV}) m_\nu$
  2. Crucial for design of full scale  $C\nu B$  PTOLEMY with 100g tritium
- Mass hierarchy
  1. Clearly decided with 100g yr exposure
  2. Up to masses  $<100\text{meV}$





# PTOLEMY: $C\nu B$ expected performance



# $C\nu B$ : Astronomy (SF)

science fiction noun

 Save Word

**Definition of *science fiction***

: fiction dealing principally with the impact of actual or imagined science on society or individuals or having a scientific factor as an essential orienting component

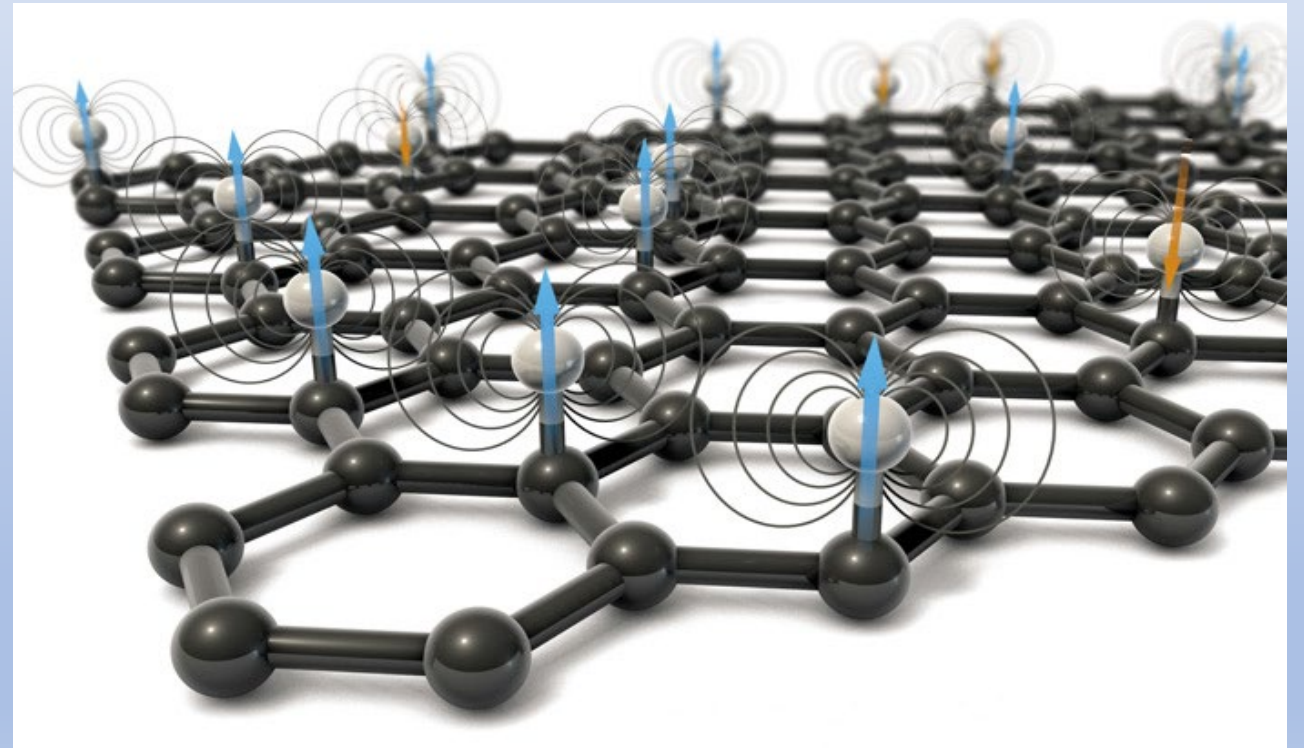
1. Suppose you have discovered  $C\nu B$
2. Suppose you have a polarized target

1+2. Localization of neutrinos:

$$\frac{d\sigma}{d\cos\theta} \propto 1 + \cos\theta$$

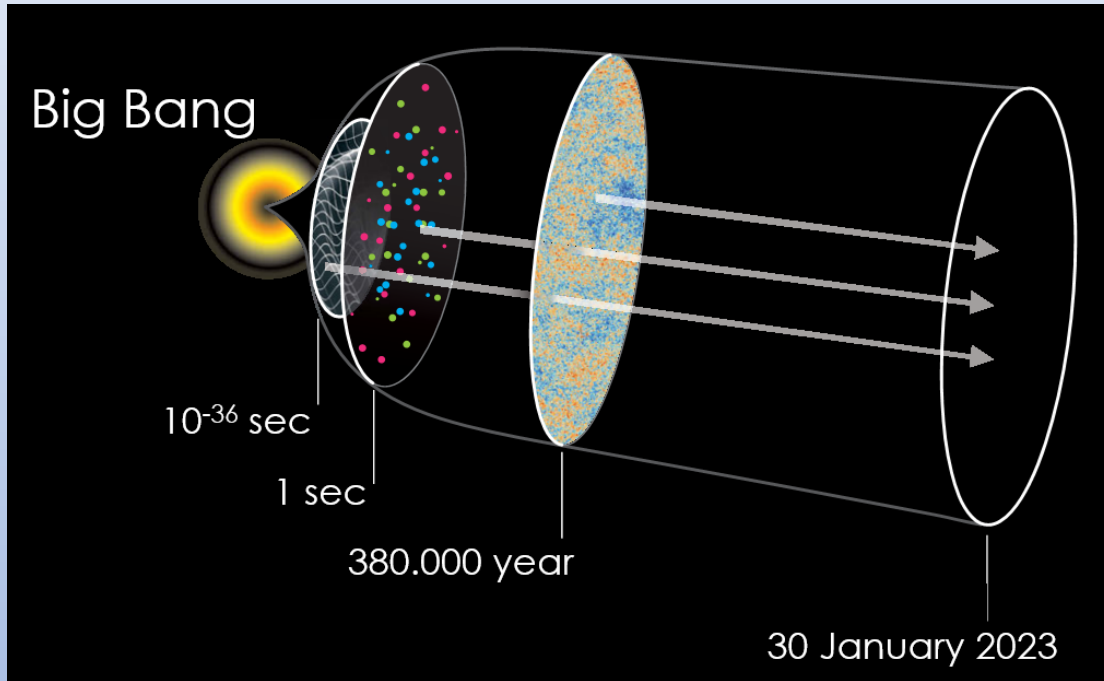
Why interesting?

Graphene with polarized tritium nuclear spin



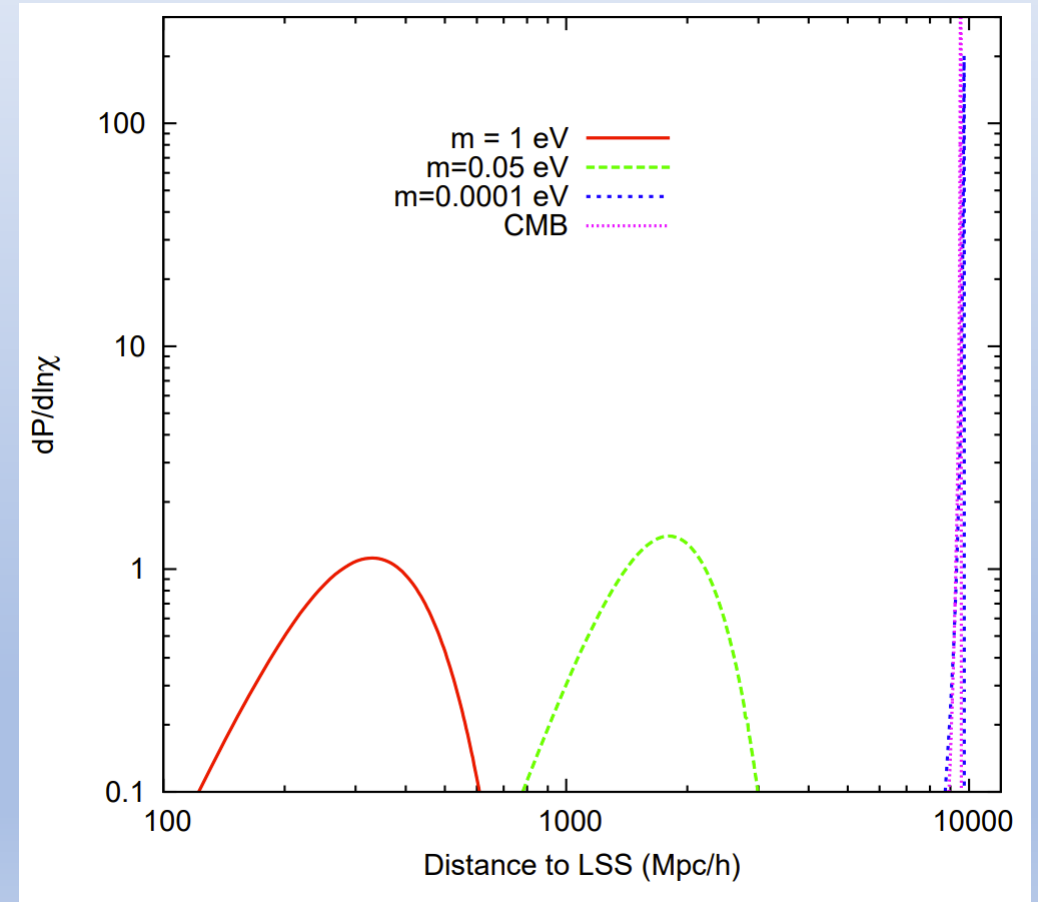
# $C\nu B$ : Astronomy

Big Bang in "time"



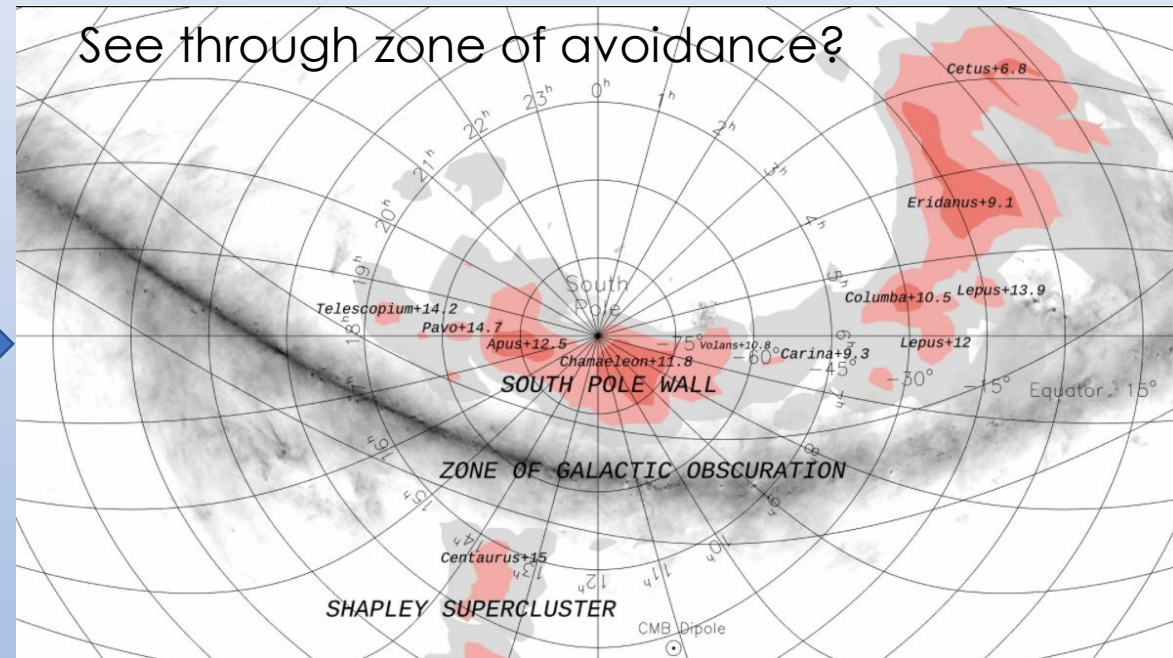
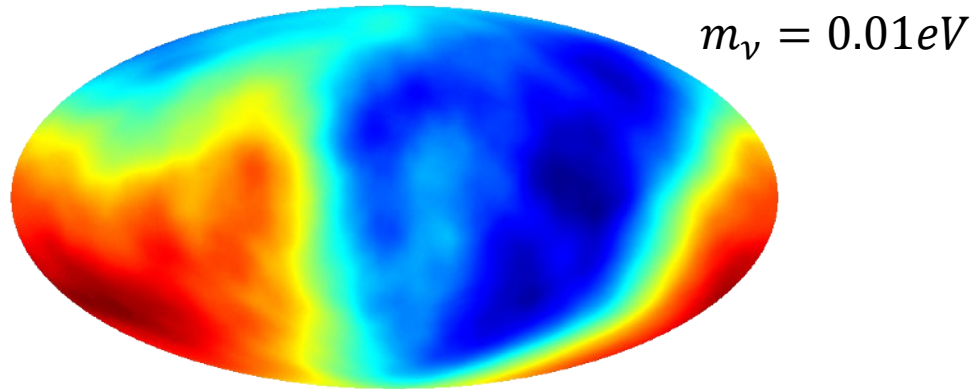
1.  $C\nu B$  has large non-relativistic component
2. It comes from 'nearby' compared to CMB

Big Bang in "distance"



# $C\nu B$ : Astronomy

1. Neutrinos 'feel' large scale structure at distances of  $O(5\text{Gly})$
2. Fluctuations hugely amplified because  $\nu$  non-relativistic  $\rightarrow 1/v^2$
3. Maybe even  $O(1-10\%)$



Could be connected to optical sky surveys....

Multi-messenger astroparticle physics?



