



Mitigating low-frequency noise

Role of tilt, software tools, controller optimization

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Outline

Low-frequency noise

Tilt & Sensors

Seismic isolation & Control

Modelling Software

Low-frequency noise





Low-frequency noise

• Limiting sensitivity of current GW detectors







Low-frequency noise

- Even bigger problem for next generation: several orders of magnitude improvement required
- Can't make the "noise wall" arbitrarily steep → need good improvement even below detection band



Tilt & sensors





Low-frequency noise: ground motion and tilt

- Imperfect compensation of seismic motion in the isolation chain
- Limited by sensor performance
- Platform tilt produces spurious translational readings
- This and other noises get **imprinted onto the motion** of isolated platform by actuation

Effect of tilt on inertial sensor:







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Some solutions:

- **Tilt sensors** (on the ground, e.g. BRS @ LIGO¹, or suspended beam²)
- Better translational local sensors based on interferometry, e.g. HoQi³
- See also: <u>yesterday's session</u>

1. <u>https://doi.org/10.1063/1.4862816</u>

- 2. https://doi.org/10.1063/5.0118606
- 3. <u>https://doi.org/10.1088/1361-6382/aab2e9</u>

Effect of tilt on inertial sensor:







Our group

Compact Balanced Readout Interferometer (COBRI) sensors

- Local displacement readout or inertial sensors
- Based on Deep Frequency Modulation Interferometry (DFMI) [poster #105 by Oliver]
- Very compact: fits into half inch mirror mount
- Expected sensitivity ~10⁻¹⁴ m/ VHz









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- Very compact: fits into half inch mirror mount
- Expected sensitivity ~10⁻¹⁴ m/ vHz
- Two other readout techniques currently investigated to go below 10⁻¹⁵/VHz:
 - Resonantly enhanced DFMI
 - Heterodyne cavity readout [poster #24 by Shreevathsa]





Seismic isolation & Control





Our group

VATIGrav: seismically isolated platform for testing sensor designs in vacuum

Vacuum goal <10⁻⁶ mbar

• Reached: 1.9*10⁻⁶ mbar

Passive isolation for optical table inside

Active seismic isolation from the ground (STACIS III)







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Currently: Translation/tilt measurement





Measurements of translation and tilt



Active isolation





Measurements of translation and tilt



Active isolation





Path to improvement







Path to improvement







Path to improvement

COBRIs on suspensions + test mass = inertial sensor







Path to improvement







Path to improvement

Need to model:

- Effects of couplings between different • DOFs in sensing and actuation
- Effects of sensor noise; optimize sensor • placement
- Optimize the controller

COBRIs on suspensions + test mass = inertial sensor



Big(ger) picture



Same modelling needed + more for big IFOs

- For Einstein Telescope, LIGO, LISA..
- Many algorithms developed by the community over the years
 - "behind closed doors" of internal repos and private computers
- We need better organized modelling software:
 - Compatible interfaces
 - Provenance

See some thoughts on that in my/Oliver's talk at ET Symposium





Common modelling tool(s)?

- We started a project to try to translate some of "community knowledge" to python and open source it
- Eager to learn about/collaborate on similar projects!

My inspirations: <u>GWpy</u>, <u>python-control</u>, <u>Kontrol</u>, <u>Finesse3</u>, ... (incomplete list)

Spicypy project



Combining several tools to facilitate **signal processing**, **control systems modelling**, and the **interface between the two**.



Laser physics, material science, ...

- Open-source, transparent, collaborative development (hosted on <u>gitlab.com</u>, Apache License, contributors from 6 institutes and 4 countries...)
- Based on well-tested GWpy and python-control
- "Modern" approach: <u>unit tests</u>, <u>examples</u> and <u>documentation</u>





Summary

Low-frequency noise: difficult and important Tilt affects Sensors \rightarrow LF noise Seismic isolation & Control => Intertwined, need to model & optimize Modelling Software => Should be a common, open effort











Thank you!







Backup





DFMI Concept



Main features:

- Signal is inherently non-linear and linearized by the phasemeter algorithm/estimator
- Each interferometer (optical head) has only one input beam and can be very compact
- Laser frequency noise is common mode (can be suppressed actively or in post-processing)
- Provides wide-range sensing of displacement & absolute ranging

See Oliver's talk at ET Symposium

O. Gerberding, Optics Express, 23, 11, (2015) G. Heinzel et al., Optics Express, 18, 19, (2010) K.-S. Isleif et al., Optics Express, 24, 2, (2016) K.-S. Isleif et al., PRApplied 12, 034025 (2019)

displacement





ReDFMI and Heterodyne cavity

<10⁻¹⁵m/sqrt(Hz) dozens of sensors per laser absolute ranging

See Oliver's talk at ET Symposium

Heterodyne cavity tracking readout

<10⁻¹⁵m/sqrt(Hz) one sensor per laser best possible performance

Caveats of tilt measurement

- Transfer function of seismometers is ignored, constant scaling in time domain (effect expected only in very low frequencies, otherwise TF ~1)
- Still working on calibration, absolute amplitudes may be off... Comparison with Blueseis 3A rotational seismometer:
- Due to limited number of seismometers can't measure at the same time on the ground and on the table, therefore using different measurements on the same plot
- Plan to use PPSD to generate more "averaged" result

- Transfer functions of STS 2.5 vs TH120:
- Almost identical in LF
- Constant = 1 in wide freq range