G2301005



Parametric Instability pre-O4 at LIGO

How TCS saves the day

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Vladimir Bossilkov Detector Engineer LIGO Livingston

Overview

• Graphical overview of PI in LIGO cavities

• How PI changes with all Future Detector Proposals

References O



Parametric Instability Intro 0000

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LIGO

Optical Beat notes and Mechanical Modes



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Parametric Instability

Parametric Instability Criterion, Danilishin et al.[1]:

$$\frac{8 \Lambda_{01} \omega_{1} \mathsf{P}_{\text{in}}}{m \omega_{m} L^{2} \gamma_{m} \gamma_{0} \gamma_{1}} \left(\frac{\gamma_{1}^{2}}{\gamma_{1}^{2} + 4\Delta_{m}^{2}}\right) \geq 1$$
(1)

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PI in action at LHO

Play with External Player Alt. Link: LHO aLog

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More Power



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Longer Cavity Arms



$$\frac{8 \Lambda_{01} \omega_1 \mathsf{P}_{\text{in}}}{m \omega_m L^2 \gamma_m \gamma_0 \gamma_1} \left(\frac{\gamma_1^2}{\gamma_1^2 + 4\Delta_m^2}\right) \ge 1$$

Optical Mode Spacing Mechanical Mode Spacing

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Bigger Test Masses



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References I

[1] Stefan L. Danilishin et al. "Time evolution of parametric instability in large-scale gravitational-wave interferometers". In: *Physical Review D* 90.11 (2014), p. 122008.

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References

Optical Mode Structure of Fabry-Perot Cavity



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Shapes of Optical Modes



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Higher test mass Q-factors



Effect on PI Gain

$$\frac{8 \Lambda_{01} \omega_{1} \mathsf{P}_{\mathsf{in}}}{m \omega_{m} L^{2} \frac{\gamma_{m}}{\gamma_{0}} \gamma_{0} \gamma_{1}} \left(\frac{\gamma_{1}^{2}}{\gamma_{1}^{2} + 4\Delta_{\mathsf{m}}^{2}}\right) \geq 1$$

Optical Mode Spacing Mechanical Mode Spacing

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