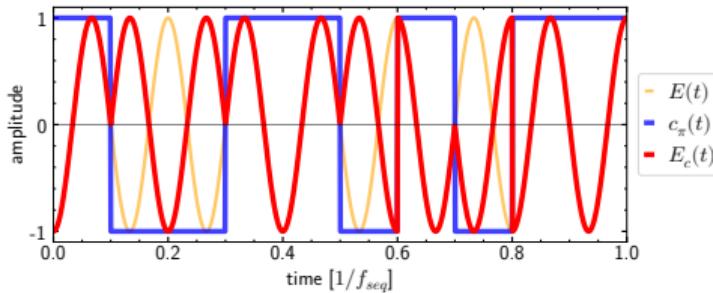


# Straylight suppression with tunable coherence in high precision interferometers

GWADW Elba 2023



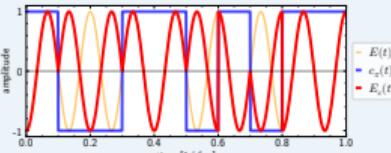
Daniel Voigt, André Lohde, Oliver Gerberding

Institute for Experimental Physics  
Universität Hamburg

23<sup>rd</sup> May 2023

# Outline

## Motivation & Concept

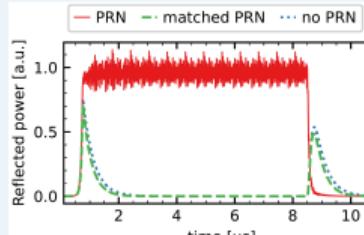


The graph plots amplitude against time in units of  $1/f_{\text{cav}}$ . It shows three oscillating signals:  $E(t)$  (blue),  $c_x(t)$  (orange), and  $E_x(t)$  (red). The blue signal has a higher frequency than the others.

*scattered light*

- limitations at low frequency
  - characteristics
  - resulting problems
- PRN phase modulation
  - concept
  - expectations

## Cavity response



The graph plots reflected power in arbitrary units (a.u.) against time in microseconds (μs). Three curves are shown: PRN (solid red), matched PRN (dashed green), and no PRN (dotted blue). The PRN curve shows a sharp initial drop followed by a noisy plateau around 1.0. The matched PRN curve follows a similar path but with less noise. The no PRN curve drops to zero almost immediately.

simulation of cavity fields

- lengths matching
- detuning
- power build-up
- locking

*related paper in review*

## Outlook



A photograph of a complex optical setup on a light-colored metal frame. The setup includes various lenses, mirrors, and a central beam splitter cube. A camera is mounted on the frame to capture images of the interference patterns.

- time-domain simulation
  - plans for improvement
  - recycling cavities
- experimental setup
  - Michelson interferometer
  - adding cavities

## Motivation

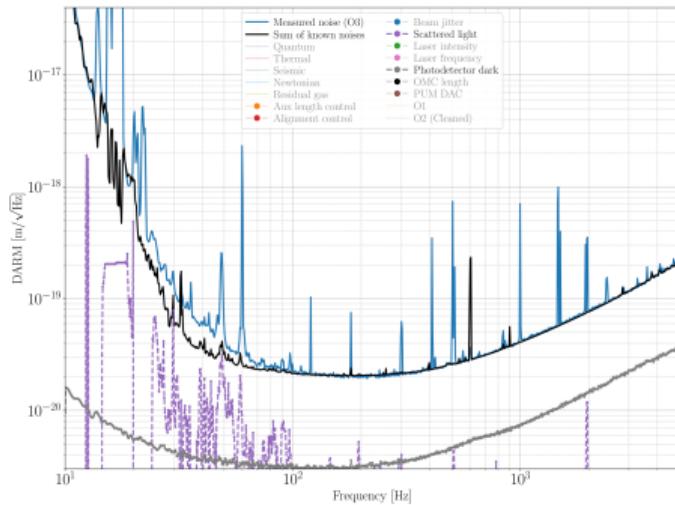


Figure: LIGO Hanford noise budget [1]

- scattered light is major limitation at low frequencies
- non stationary noise

## Motivation

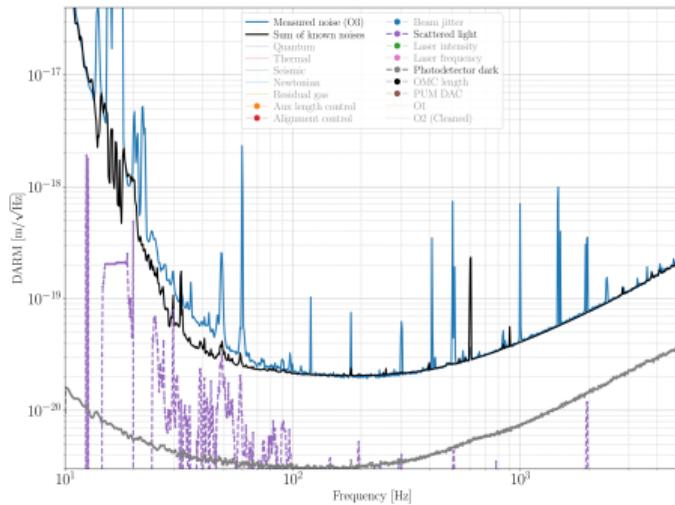
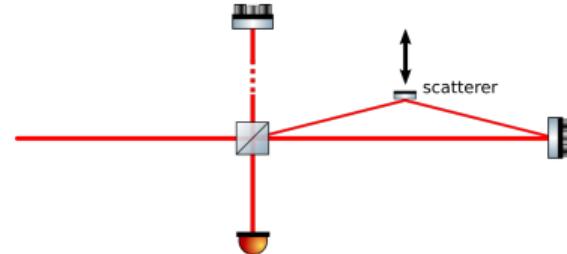
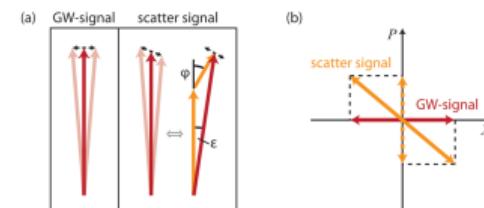


Figure: LIGO Hanford noise budget [1]

- scattered light is major limitation at low frequencies
- non stationary* noise



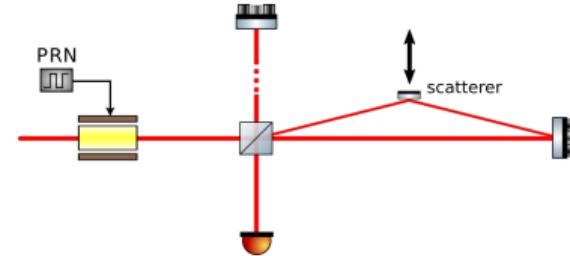
- non-linear* coupling
- frequency *up-conversion*
- amplitude- *and* phase- modulation



## Concept

### Tunable Coherence

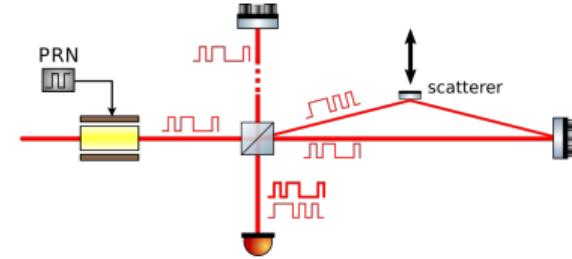
- phase modulation at **GHz**
- "random" noise as modulation sequence  $c_0$   
→ **pseudo white-light interferometer**



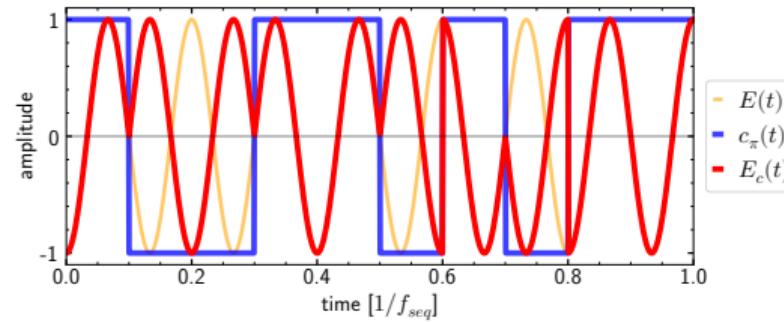
## Concept

### Tunable Coherence

- phase modulation at **GHz**
- "random" noise as modulation sequence  $c_0$   
→ **pseudo white-light interferometer**
- **chips** of  $c_0$  generated as 0 or 1  
→ modulation  $c_{\pm 1}(t)$  of 0 or  $\pi$



$$E(t) = E_0 e^{i(\omega t + \varphi + c_0(t)\pi)} = c_{\pm 1}(t) E_0 e^{i(\omega t + \varphi)}$$



## Concept

### modulation sequence

- pseudo-random-noise (**PRN**) sequence as modulation input  $c_0$
- $m$ -sequence of length  $l_{seq}$   
 $\dots \underline{1011100} \underline{1011100} 1011100 \underline{\underline{1011100}} 101\dots$   
"random"  $l_{seq}$

## Concept

### modulation sequence

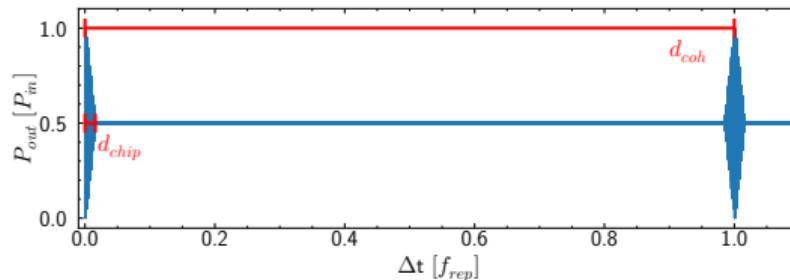
- pseudo-random-noise (**PRN**) sequence as modulation input  $c_0$
- $m$ -sequence of length  $l_{seq}$

...1011100101110010111001011100101...

"random"  $l_{seq}$

### influence on MI output

- small scale dependent on chip length  $d_{chip}$
- large scale dependent on autocorellation of PRN sequence
  - **tunable coherence**
  - **re-coherence** length  $d_{coh}$



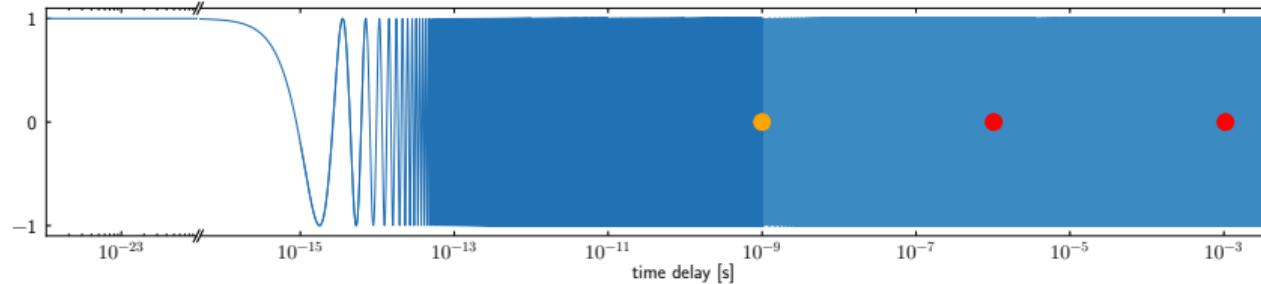
$$d_{coh} = \frac{l_{seq}}{f_{mod}} \cdot c$$

$$d_{chip} = \frac{c}{f_{mod}}$$

## Concept

### Distance examples

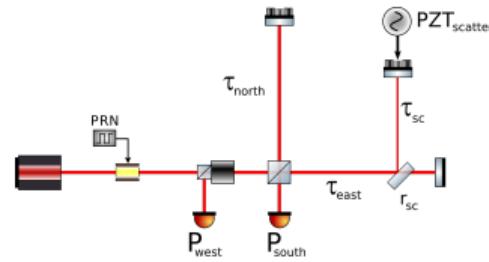
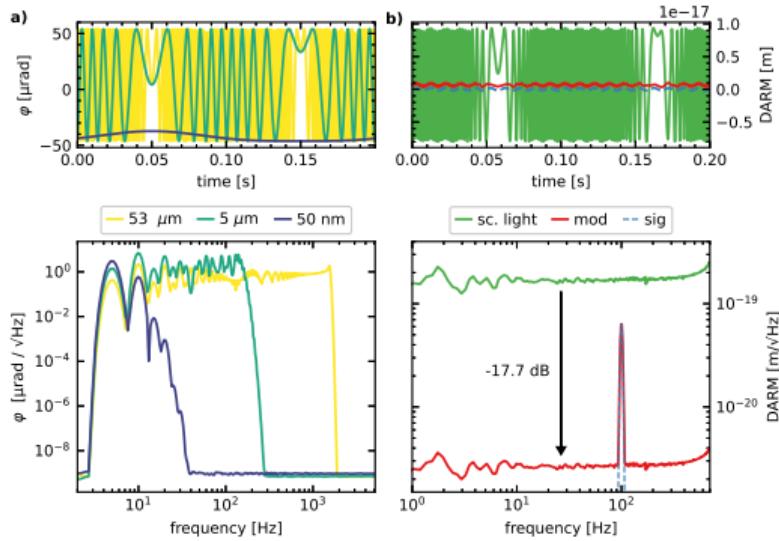
		$f_{\text{mod}} = 1 \text{ GHz}$		$f_{\text{mod}} = 10 \text{ GHz}$	
laser frequency	1064 nm	$l_{\text{seq}}$ :	31 chips	16 383 chips	31 chips
● PRN chip	$d_{\text{chip}}$	[cm]	29.9 cm	29.9 cm	2.99 cm
● PRN sequence	$d_{\text{coh}}$	[m]	9.29 m	4911.50 m	0.93 m





## Simulation results

### Michelson Interferometer

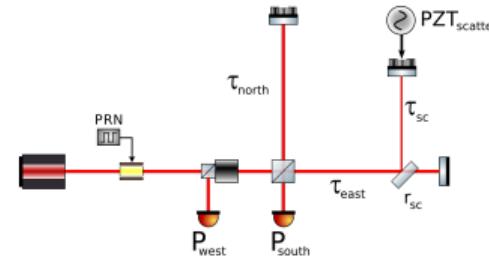
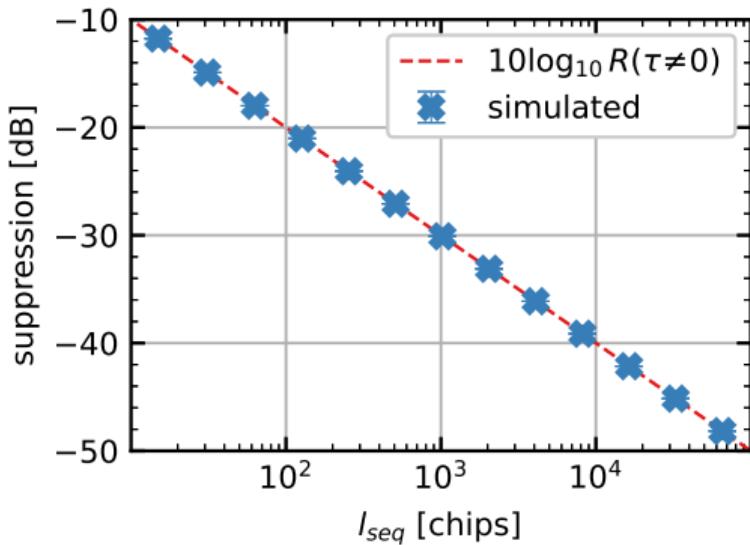


$$\begin{aligned}
 P_S = & \frac{P_0}{2} \left[ (1 - r_{sc}^2) [1 - R(\tau) \cos(\omega\tau)] \right. \\
 & + r_{sc}^2 [R(\tau + \tau_{sc}) \cos(\omega(\tau + \tau_{sc})) - R(\tau_{sc}) \cos(\omega\tau_{sc})] \\
 & \left. + r_{sc}^4 [1 + R(\tau_{sc}) \cos(\omega\tau_{sc})] \right]
 \end{aligned}$$



## Simulation results

Michelson Interferometer

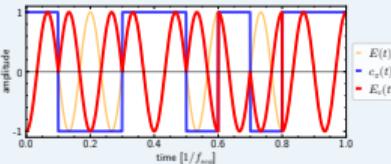


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## Outline

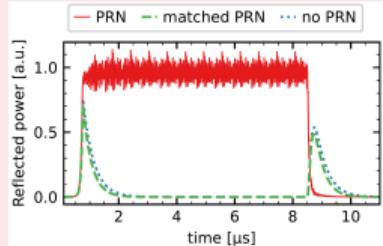
### Motivation & Concept



*scattered light*

- limitations at low frequency
  - characteristics
  - resulting problems
- PRN phase modulation
  - concept
  - expectations

### Cavity response



*simulation of cavity fields*

- lengths matching
- detuning
- power build-up
- locking

### Outlook



- time-domain simulation
  - plans for improvement
  - recycling cavities
- experimental setup
  - Michelson interferometer
  - adding cavities

## Cavity response

### Equations for fields

modulated input field

$$E_{in}(t) = E_0 e^{i(\omega t + \varphi + c_0(t)\pi)} = c_{\pm 1}(t) E_0 e^{i(\omega t + \varphi)}$$

random behavior for e.g. reflected field

$$\begin{aligned} E_{\text{refl}}(t) = & r_1 E_0 e^{i(\omega t + \varphi)} \left[ c_{\pm 1}(t) \right. \\ & \left. - \frac{t_1^2}{r_1^2} \sum_{n=1}^{\infty} c_{\pm 1}(t-n\tau) (r_1 r_2 e^{-i\omega\tau})^n \right] \end{aligned}$$

and equivalently for other fields



## Cavity response

### Equations for fields

#### modulated input field

$$E_{in}(t) = E_0 e^{i(\omega t + \varphi + c_0(t)\pi)} = c_{\pm 1}(t) E_0 e^{i(\omega t + \varphi)}$$

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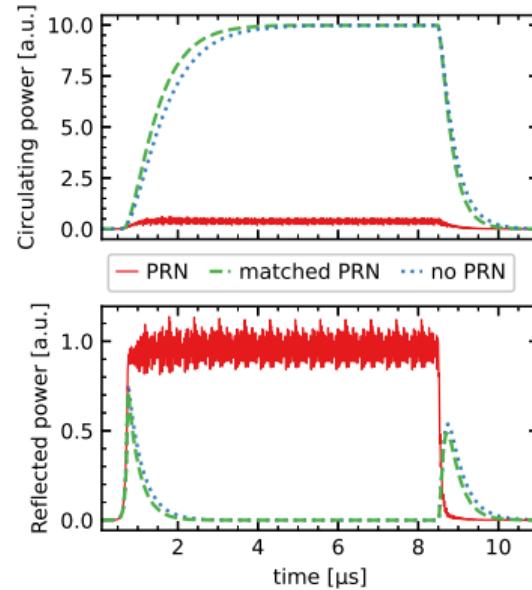
and equivalently for other fields

#### recoherence for sequence matching cavity

if  $c_{\pm 1}(t-n\tau) = c_{\pm 1}(t) \forall n, t$ :

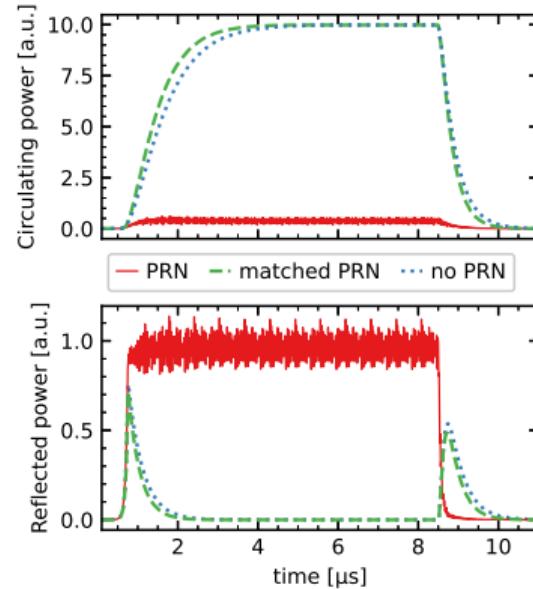
$$\begin{aligned} E_{\text{refl}}(t) &= r_1 c_{\pm 1}(t) E_0 e^{i(\omega t + \varphi)} \left[ 1 - \frac{t_1^2}{r_1^2} \sum_{n=1}^{\infty} (r_1 r_2 e^{-i\omega\tau})^n \right] \\ &= c_{\pm 1}(t) E_0 e^{i(\omega t + \varphi)} \left[ r_1 - t_1^2 r_2 e^{-i\omega\tau} \sum_{n=0}^{\infty} (r_1 r_2 e^{-i\omega\tau})^n \right] \\ &= E_{in}(t) \left[ r_1 - \frac{r_2 t_1^2 e^{-i\omega\tau}}{1 - r_1 r_2 e^{-i\omega\tau}} \right] \end{aligned}$$

## Length matching

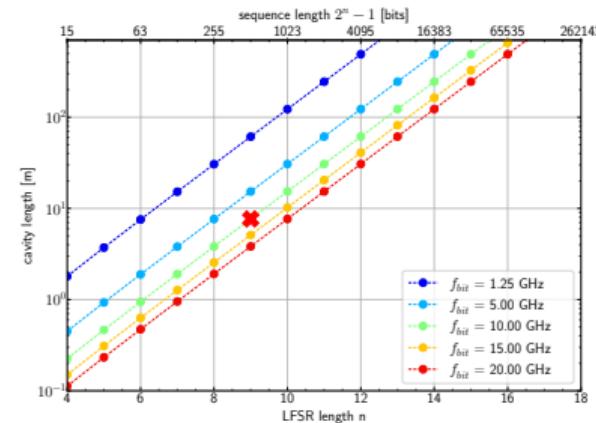


- random behavior if sequence does not fit
- cavity length locked to recohorence length
- modulation frequency locked to laser frequency

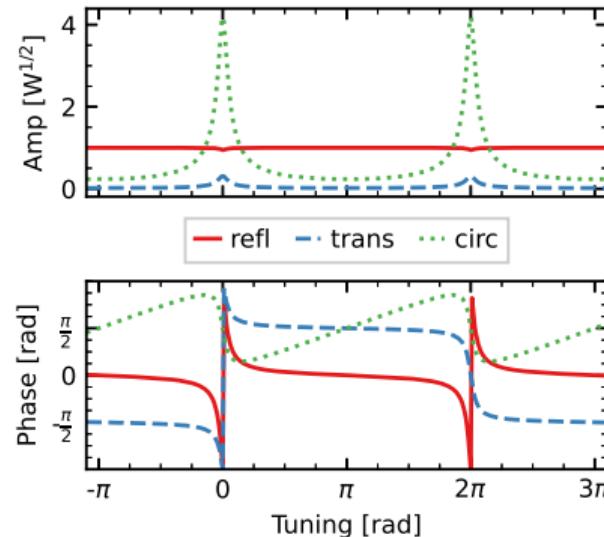
## Length matching



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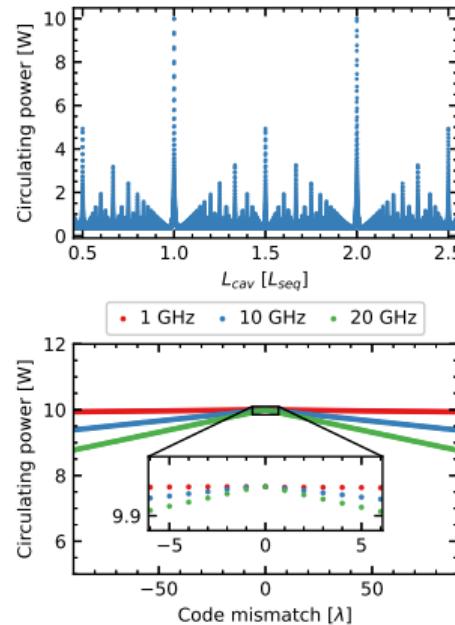
## Detuning



- no difference if cavity is locked to sequence length
- shown left: over-coupled
- same for under-coupled & impedance matched
- tuning range limited due to modulation
- modulation frequency should be locked to laser frequency

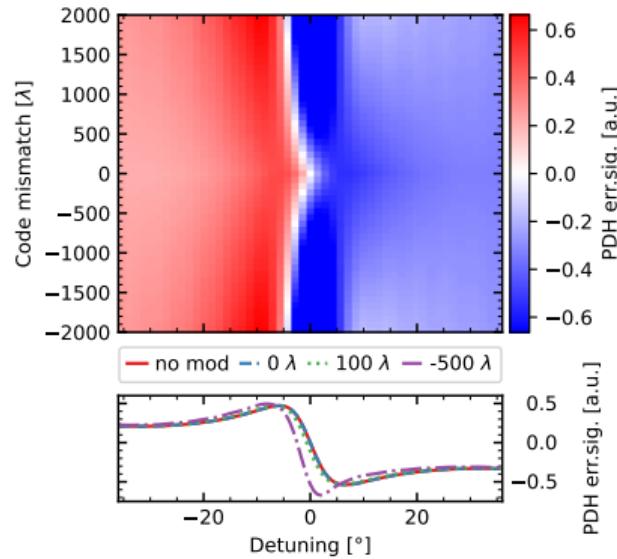


## Power build-up



- full build-up for integer multiple of sequence length
- halved build-up for e.g. 1.5 sequence lengths
- FWHM depending on modulation frequency
  - in  $\mu\text{m}$  to  $\text{mm}$  range
  - e.g. for 10 GHz around 1776 wavelengths
- sensitivity depending on cavity finesse

## PDH - error signal



- normal error signal for matched cavity
- local oscillator phase not adjusted for each simulation
  - probably cause for observed steeper flanks
- locking might require additional unmodulated laser



## Outline

### Motivation & Concept

amplitude

time [ $1/f_{\text{osc}}$ ]

$E(t)$   
 $c_d(t)$   
 $E_d(t)$

*scattered light*

- limitations at low frequency
  - characteristics
  - resulting problems
- PRN phase modulation
  - concept
  - expectations

### Cavity response

Reflected power [a.u.]

time [ $\mu\text{s}$ ]

PRN  
 matched PRN  
 no PRN

*simulation of cavity fields*

- lengths matching
- detuning
- power build-up
- locking

### Outlook

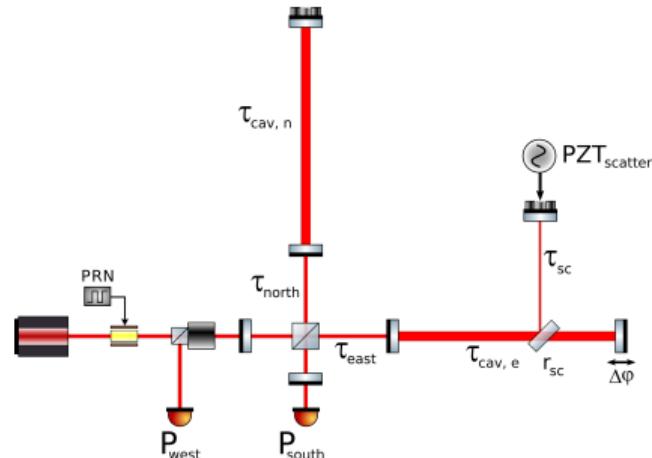
- time-domain simulation
  - plans for improvement
  - recycling cavities
- experimental setup
  - Michelson interferometer
  - adding cavities

# Outlook

## Simulation

### future needs

- recycling cavities & scattered light
- more generalized setup
- higher precision without performance penalty
- performance increase



# Outlook

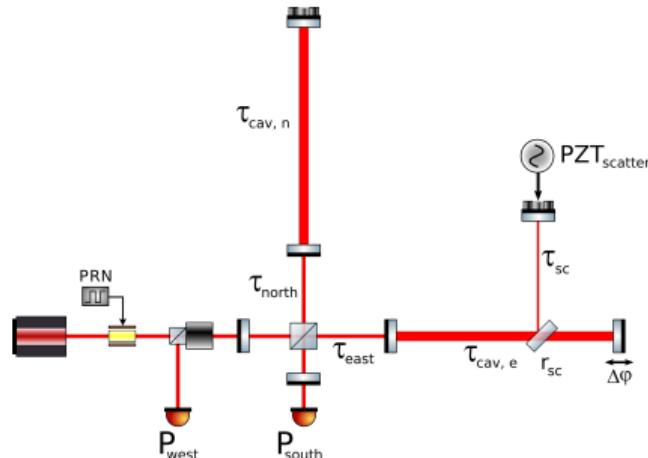
## Simulation

### future needs

- recycling cavities & scattered light
- more generalized setup
- higher precision without performance penalty
- performance increase

### conceptual ideas

- use modular setup
  - define each optic and its "connections"
  - iterate through setup and propagate fields in time-steps
- for FINESSE compatibility, use similar input method
  - run both simulations parallel



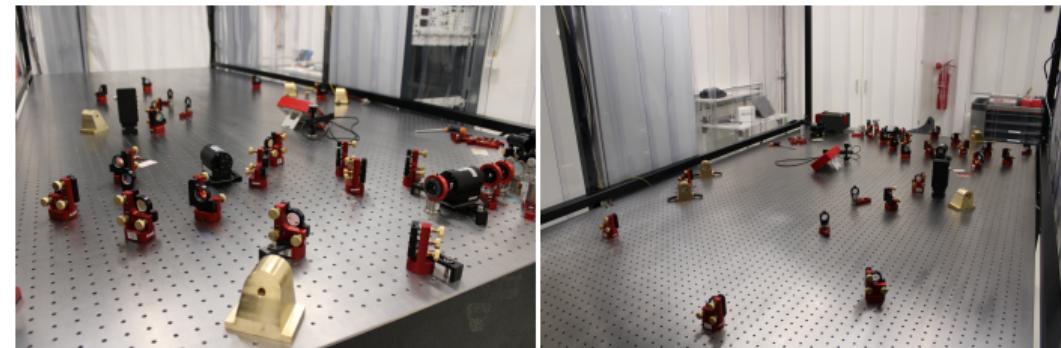
## Outlook Experiment

moved into the lab

prepared supporting infrastructure

started setup for testing

adjust interferometer and measure



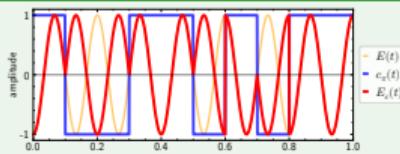
experiments in preparation

tunable coherence & dual-port quadrature read-out



## Conclusion

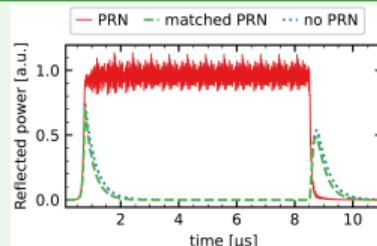
### Motivation & Concept



*scattered light*

- non-stationary noise
  - limiting at low frequency
  - even more problematic in future detectors
- PRN phase modulation
  - GHz modulation
  - high suppression possible
  - length matching crucial

### Cavity response

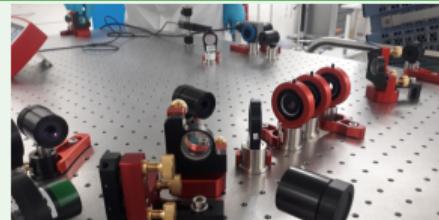


*simulation of cavity fields*

- lengths matching is crucial
- detuning is possible
- power build-up works normal if lengths match
- locking probably needs an unmodulated laser

*related paper in review*

### Outlook



- time-domain simulation
  - recycling cavities
  - more general setup
  - FINESSE compatibility
- experimental setup
  - moved into lab
  - Michelson interferometer setup in preparation

## Bibliography

-  **Craig Cahillane.** <https://ccahilla.github.io/>. URL: <https://ccahilla.github.io/> (visited on 03/10/2022).
-  **Melanie Ast.** "Quantum-dense metrology for subtraction of back-scatter disturbances in gravitational-wave detection". PhD thesis. Universität Hannover, 2017.