2022/10/13 Kick-off meeting Spoke 2 WP-3

#### Landscape and computational challenges for the LISA global fit

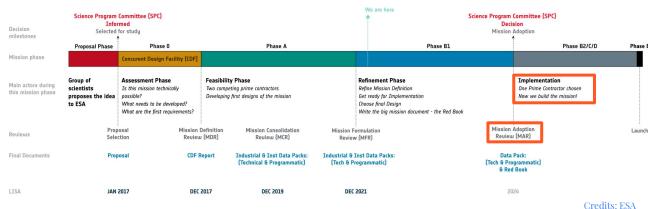
**R. Buscicchio** collaborators: M.Colpi, A.Sesana, D.Gerosa & al.

WAVE ASTRONOMY





- Mission Formulation Review 🗸
- Towards Mission Adoption:
  - Reshaping the collaboration: Performance Experts & Data Analysis Experts Groups
  - Red book
- Launch expected in 2034





Credits: NASA/JPL-Caitech/NASAEA/E CXC/STScl/GSFCSVS/S.Barke



Question: Are we ready to process LISA data? Answer: Data Challenges



- Waveforms
- Instrument response
- Likelihoods
- Global fit
- Population inference
- Get in touch for references and discussion!

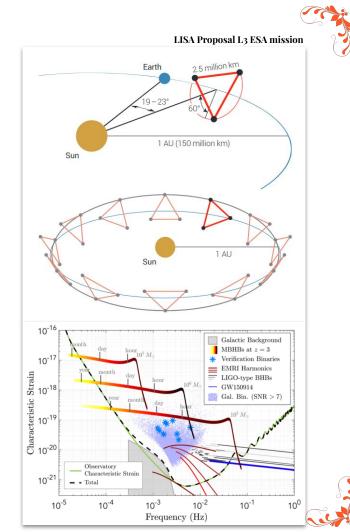








- Long baseline: 2.5 Mkm
- Sensitivity bucket ~ mHz
- Data on a stick: 4 to 10 years, 3 "science" channels = 15 GB
- Source-rich sky: persisting & overlapping sources
- Time dependent response: that's new wrt ground-based detectors
- Noise/Signal distinction is blurry: "confusion" noise
- Dominating laser phase noise: synthetic interferometry





**Waveform evaluations Key computational features** 

Bayesian parameter estimation: ~a few long-term, incremental, codebase developments worldwide (including UniMiB)

 $10^{-16}$ 

10-17

Characteristic Strain <sub>61-</sub>01 Characteristic Strain <sub>61-</sub>01 Characteristic Strain

10<sup>-19</sup>

10<sup>-21</sup>

 $10^{-5}$ 

Observatory Characteristic Strain

10-4

Total

- **Monochromatic (DWDs)**
- **Drifting sources (BBHs)**
- **Chirping sources (SMBBHs)**
- **Polichromatic (EMRIs)**
- Many unresolved ones (SGWBs)
- **Instrumental artifacts**
- Multiband sources »

#### computationally challenging

**Question:** Are we ready to search for all? Answer: Yes for some sources, other under development

10<sup>-3</sup>

Frequency (Hz)

10-2



+noise

Galactic Background

Verification Binaries

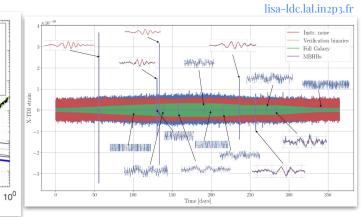
MBHBs at z = 3

EMRI Harmonics

— LIGO-type BHBs GW150914

Gal. Bin. (SNR > 7)

 $10^{-1}$ 





**Waveform evaluations Key computational features** 

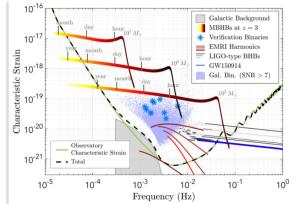
Bayesian parameter estimation: ~a few long-term, incremental, code developments worldwide (including UniMiB)

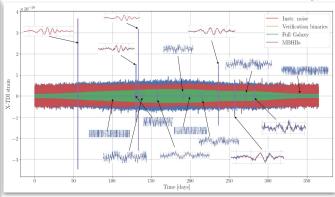




- Fast multi-sources waveforms >> slow
- Slow single-source bright waveforms»slow
- Only heterogeneous modelling is possible

**>>** Towards the global fit **<**<





Question: Are we ready to search for all? Answer: Yes for some sources, other under development





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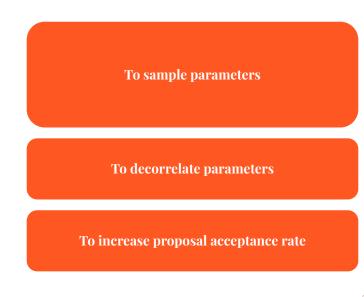
- Some source classes are GPU accelerated.
- $\circ$  Some other are going to be
  - Chirping signals are inherited from ground-based detectors (+eccentricity)
  - Drifting signals are slowly inspiralling. Must track ~10<sup>5-6</sup> signal cycles
- Surrogates
  - Leverage exact solutions from numerical relativity. Useful for massive sources. Require massive simulations to "train"
  - Reduced order models trained on neural networks: suitable for gradient based inferences





# Computational challenges Likelihoods

- highly multimodal (multiple source inference):
  - issues are solvable on a statistical basis (label switching problem)
  - other just require massive computation
- highly curved likelihood surfaces
- $\circ \quad$  a zoo of stochastic samplers:
  - Nested sampling
  - Markov-Chain Montecarlo
  - Hamiltonian samplers
  - multithreading is mostly out-of-the-box
- Sometimes enhanced with
  - normalizing flows
  - genetic programming (not algorithms)
- gaussian process augmentation
- neural amortized proposals
- particle swarm optimizers

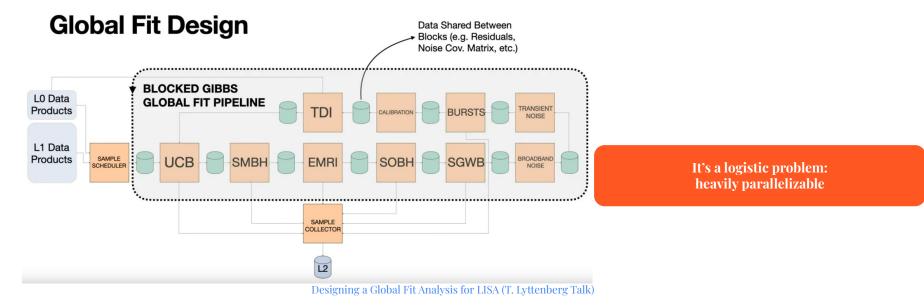






## **Computational challenges** Global fit: the elephant in the room





#### **Population inference:**

Given the number of sources, established hierarchical likelihoods might be too slow. Not covered in this talk.



### **The end** Questions?



