



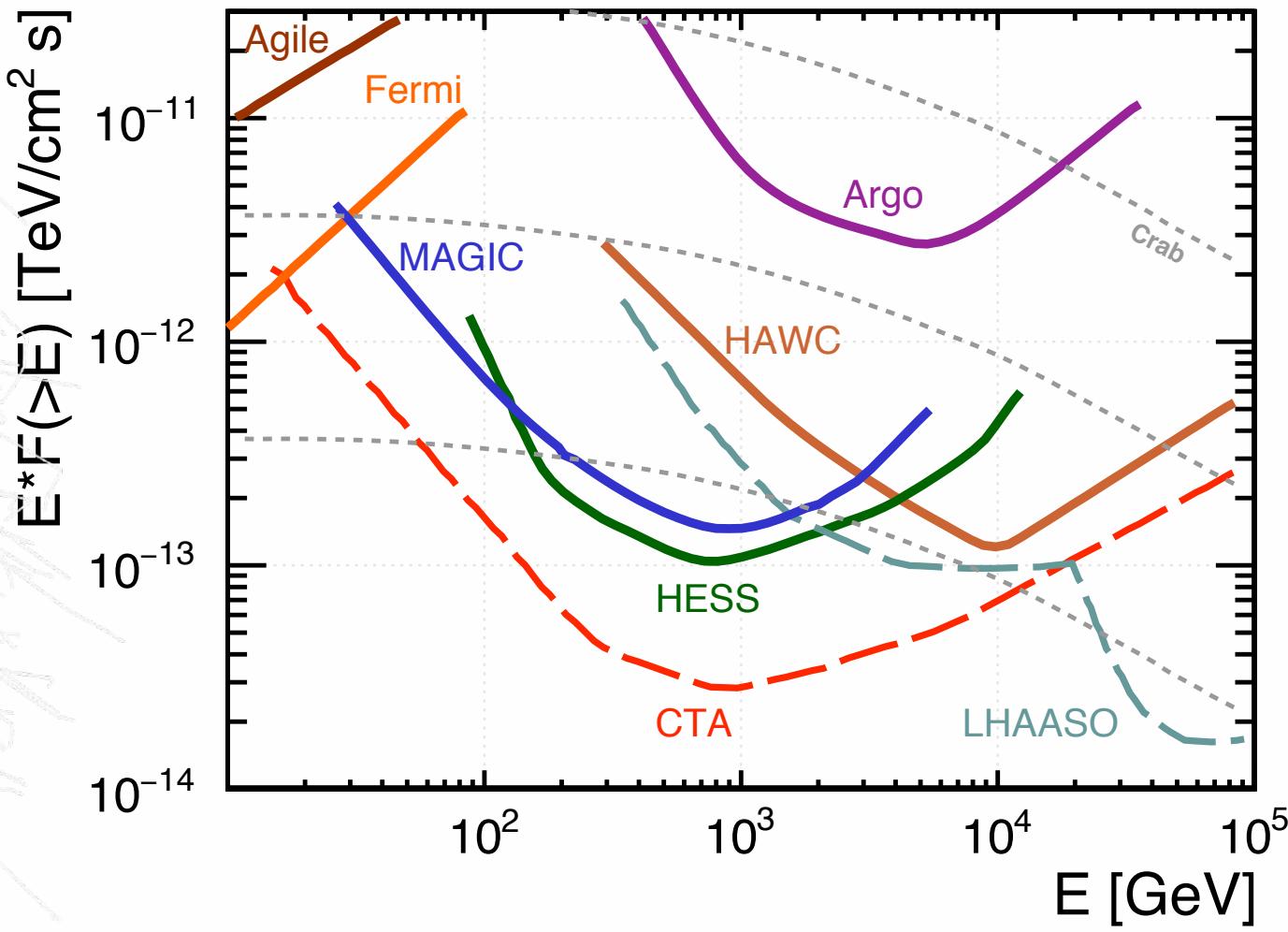
LATTES: a new gamma-ray detector concept for South America

Ruben Conceição

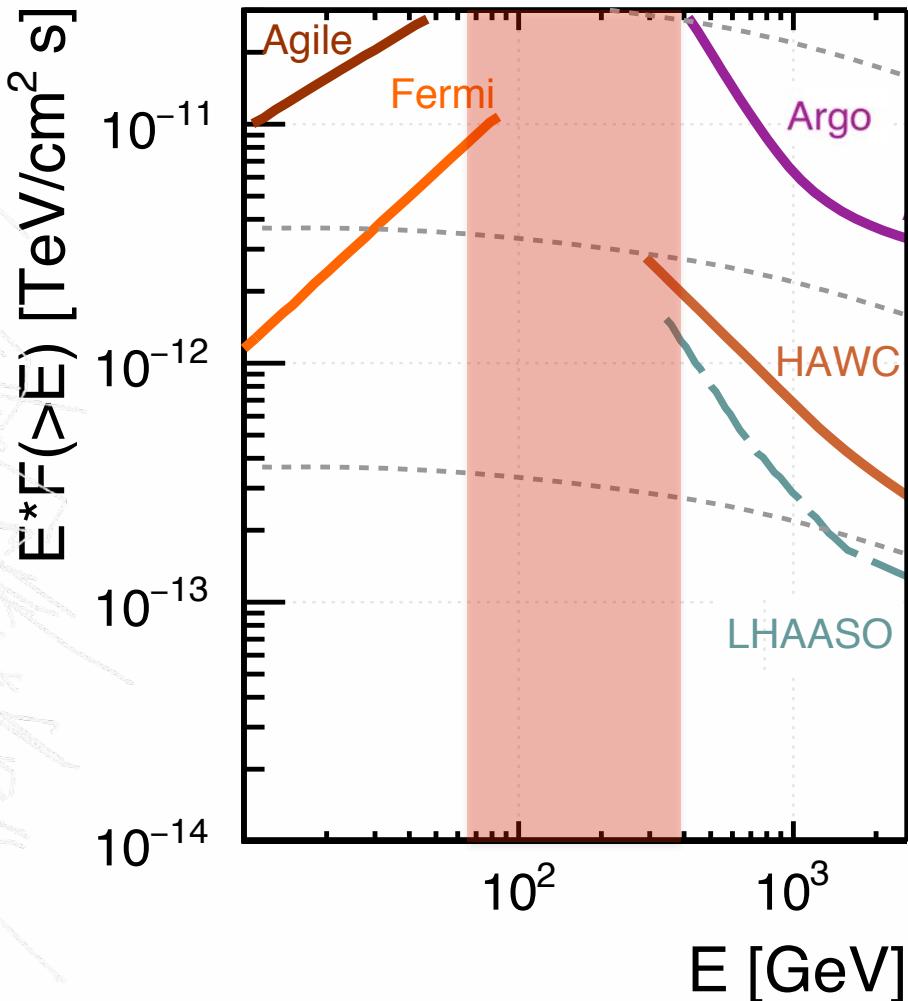
*P. Assis, U. Barres de Almeida, A. Blanco, A. De Angelis,
P. Fonte, L. Lopes, G. Matthiae, M. Pimenta, R. Shellard, B. Tomé*



Current experimental status

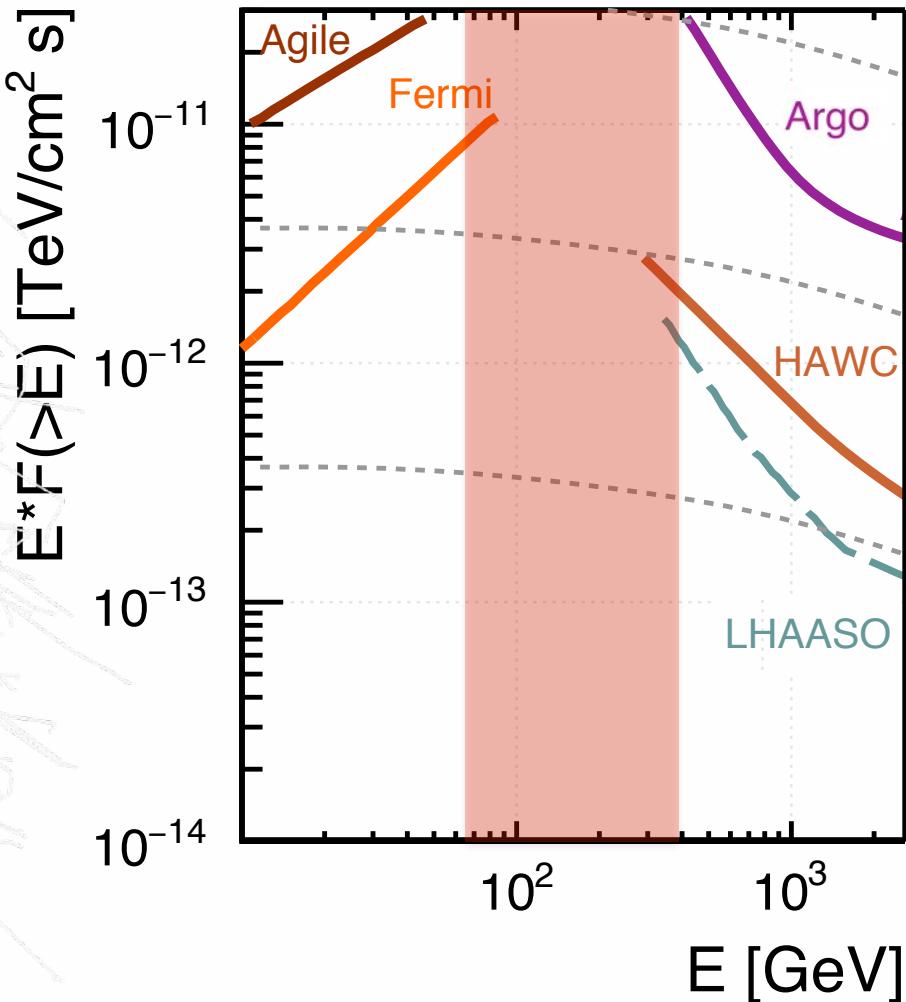


Current Situation



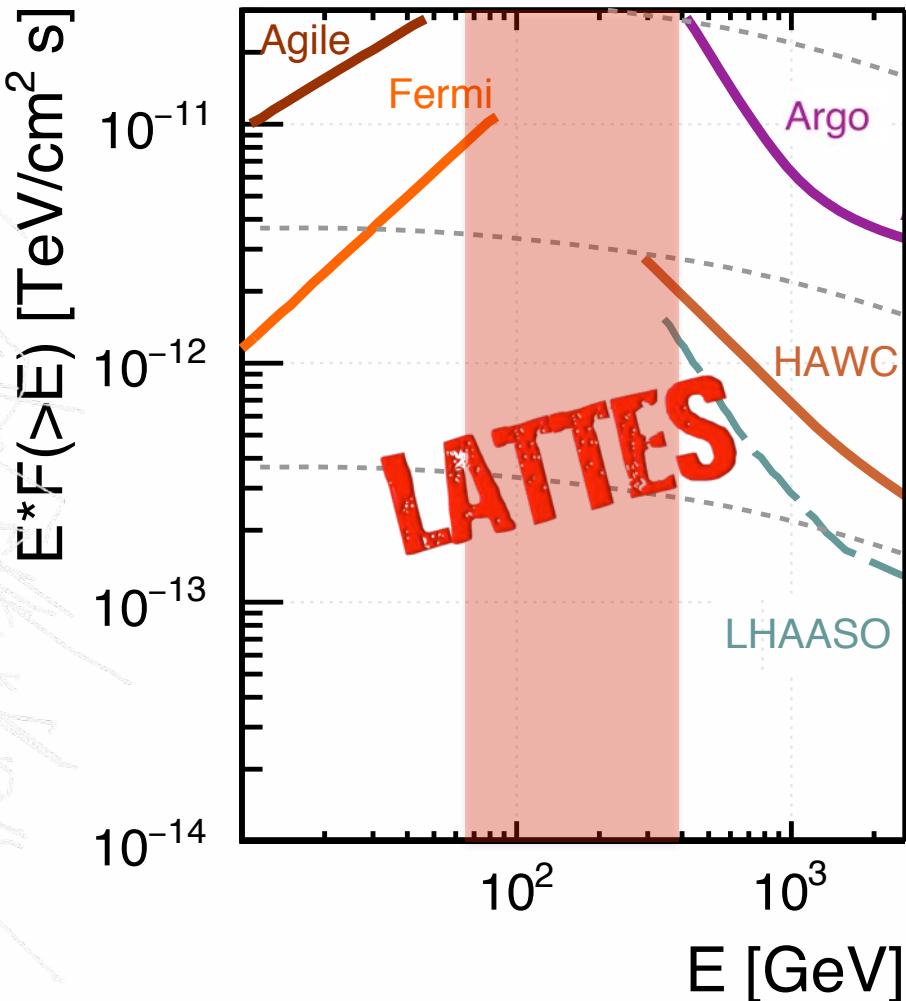
- No wide FoV experiment to:
 - Survey the Galactic Center (GC)
 - Explore the energy region of 100 GeV

Requirements



- Build an **EAS array experiment**:
 - Located in the **South Hemisphere**
 - Low energy threshold:
 - High altitude
 - Next generation detector concept

Solution



- Build an EAS array experiment:
 - Located in the South Hemisphere
 - Low energy threshold:
 - High altitude
 - Next generation detector concept

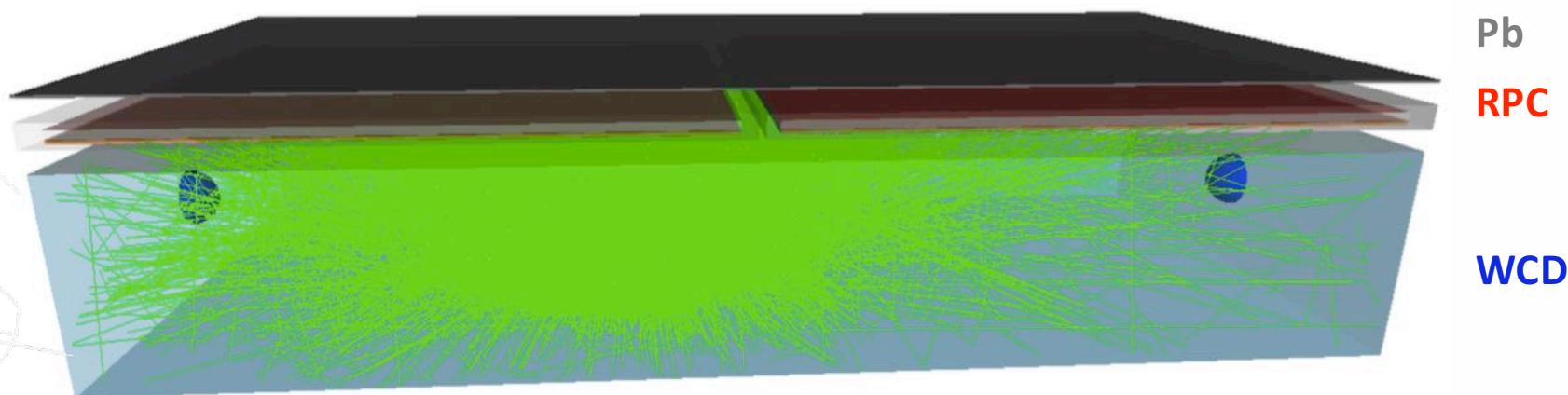
LATTES @ ALMA site

Large Array Telescope for Tracking Energetic Sources



LATTES concept

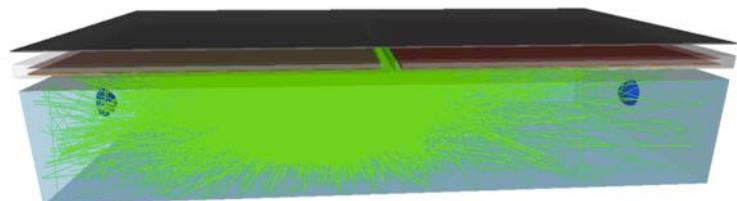
LATTES STATION



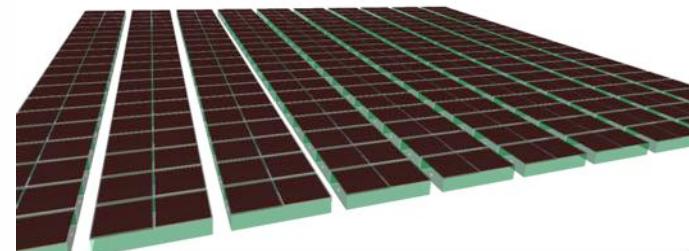
- Thin lead plate (**Pb**)
 - 5.6 mm (one radiation length)
- Resistive Plate Chambers (**RPC**)
 - 2 RPCs per station
 - Each RPC with 4x4 readout pads
- Water Cherenkov Detector (**WCD**)
 - 2 PMTs (diameter: 15 cm)
 - Dimensions: 1.5 m x 3 m x 0.5 m

LATTES concept

- Hybrid detector:
 - Thin lead plate
 - To convert the secondary photons
 - Improve geometric reconstruction
 - Resistive Plates Chamber
 - Sensitive to charged particles
 - Good time and spatial resolution
 - Improve geometric reconstruction
 - Explore shower particle patterns at ground
 - Water Cherenkov Detector
 - Sensitive to secondary photons and charged particles
 - Measure energy flow at ground
 - Improve trigger capability
 - Improve gamma/hadron discrimination



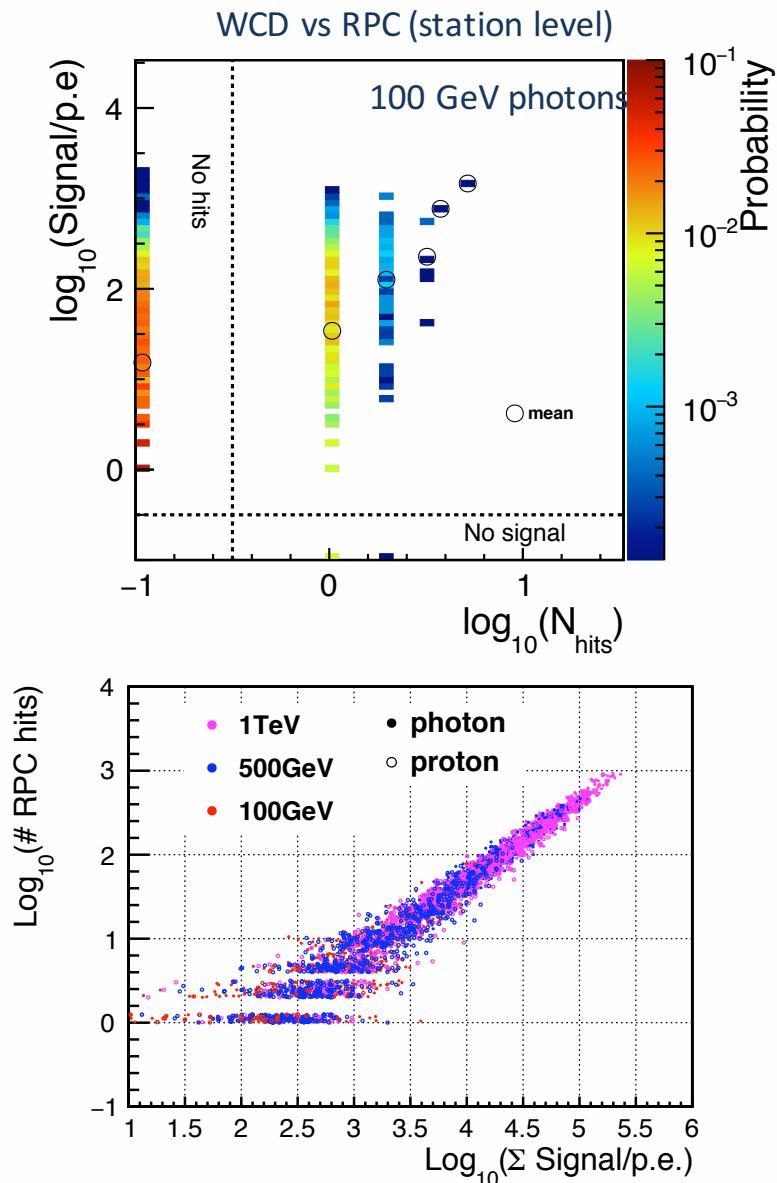
LATTES station
 $1.5 \text{ m} \times 3 \text{ m} \times 0.5 \text{ m}$



LATTES core array
 30×60 stations
 $100 \times 100 \text{ m}^2$

LATTES: complementary

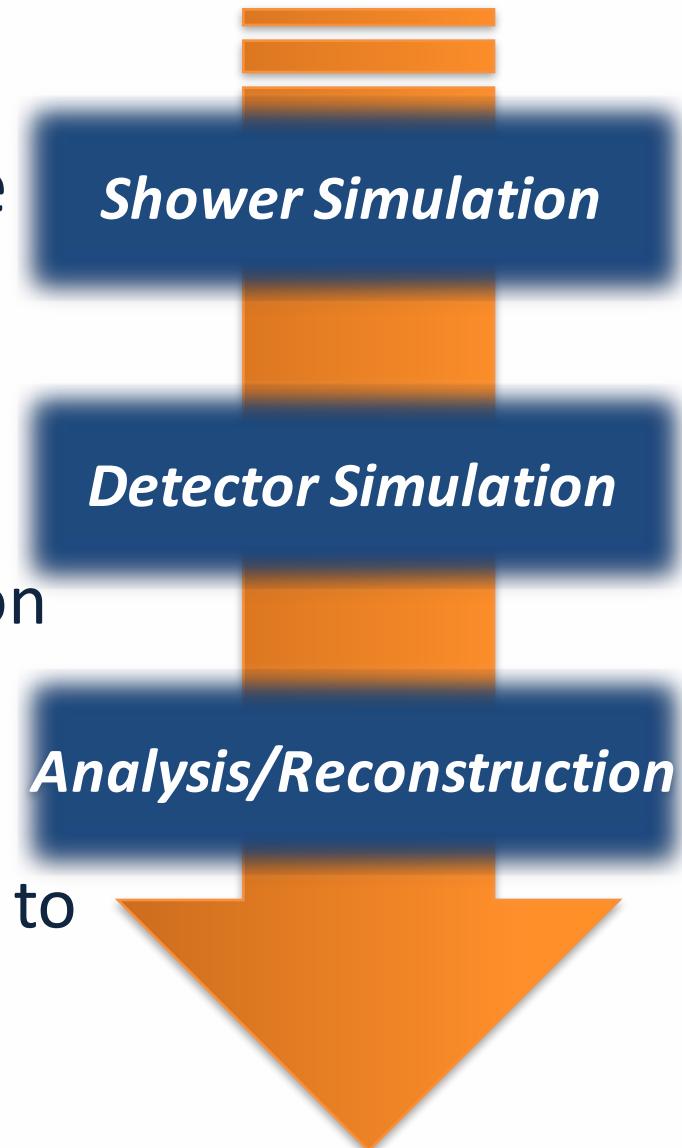
- Combined detection:
 - Lower the **energy threshold**
 - Improve the trigger conditions (WCD)
 - Enable detector **inter-calibrations**
 - Energy calibration can be used to control detector systematic uncertainties
 - Check Monte Carlo simulations performance
 - Enhance **gamma/hadron discrimination**
 - Explore shower characteristics
 - Access to Argo/HAWC discrimination techniques



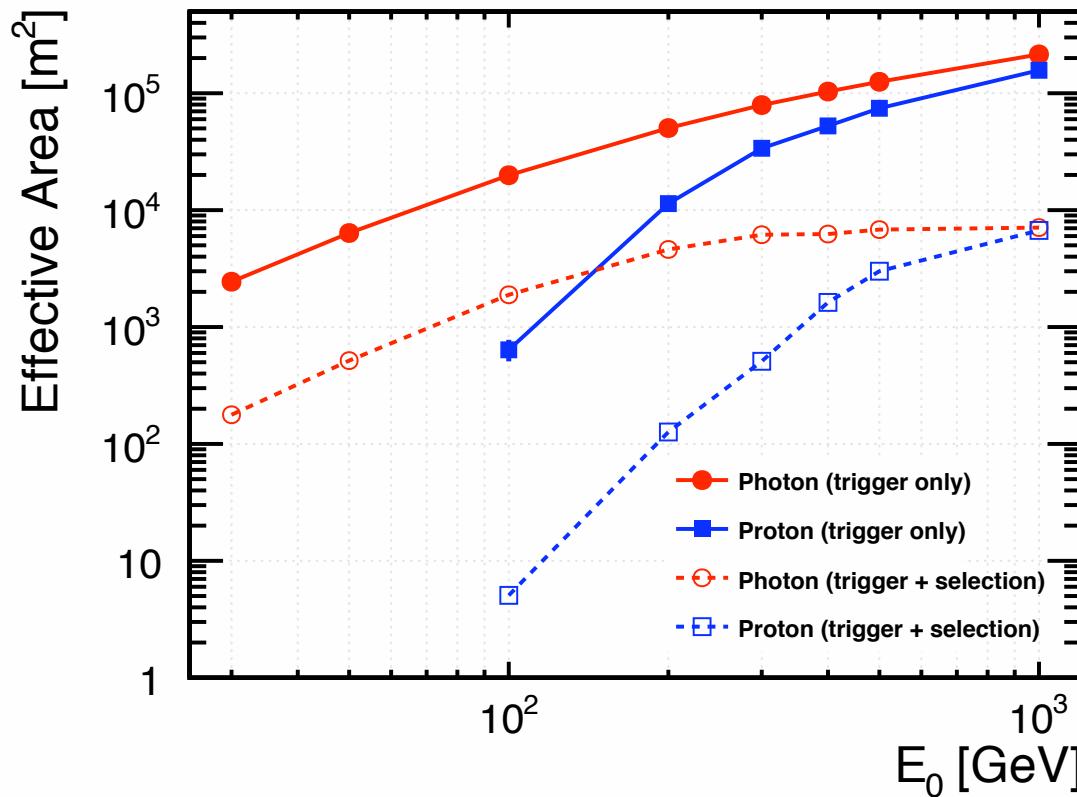
- LATTES performance:
 - Trigger efficiency
 - Energy Reconstruction
 - Geometric Reconstruction
 - Gamma-hadron discrimination
- LATTES sensitivity

Simulation Framework

- Complete **end-to-end simulation** chain to evaluate LATTES performance
 - Showers simulated using **CORSIKA**
 - Detector layout and simulation performed by **Geant4**
 - **LATTESSim**: Integrated toolkit to study and optimize LATTES performance



Trigger efficiency

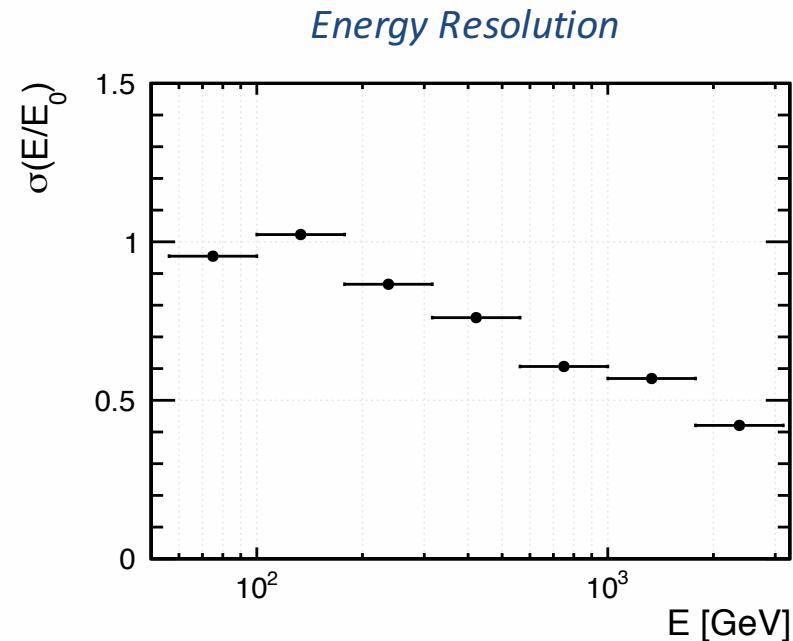
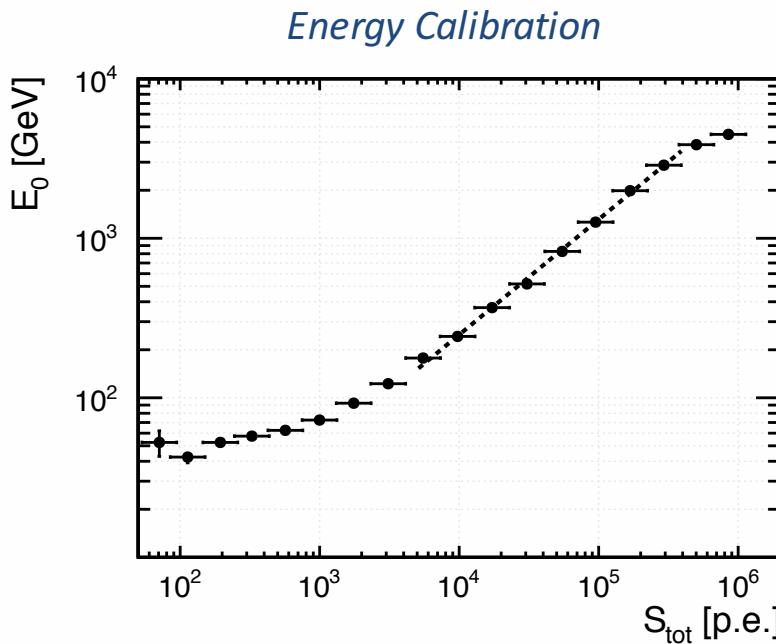


- Use WCD stations to trigger at low energies
 - Trigger condition
 - Station: require more than 5 p.e. in each PMT
 - Event: require 3 triggered stations
 - *Effective Area of 1000 m² at 100 GeV! (after quality cuts)*

Energy reconstruction

$E_0 \rightarrow$ Simulated energy

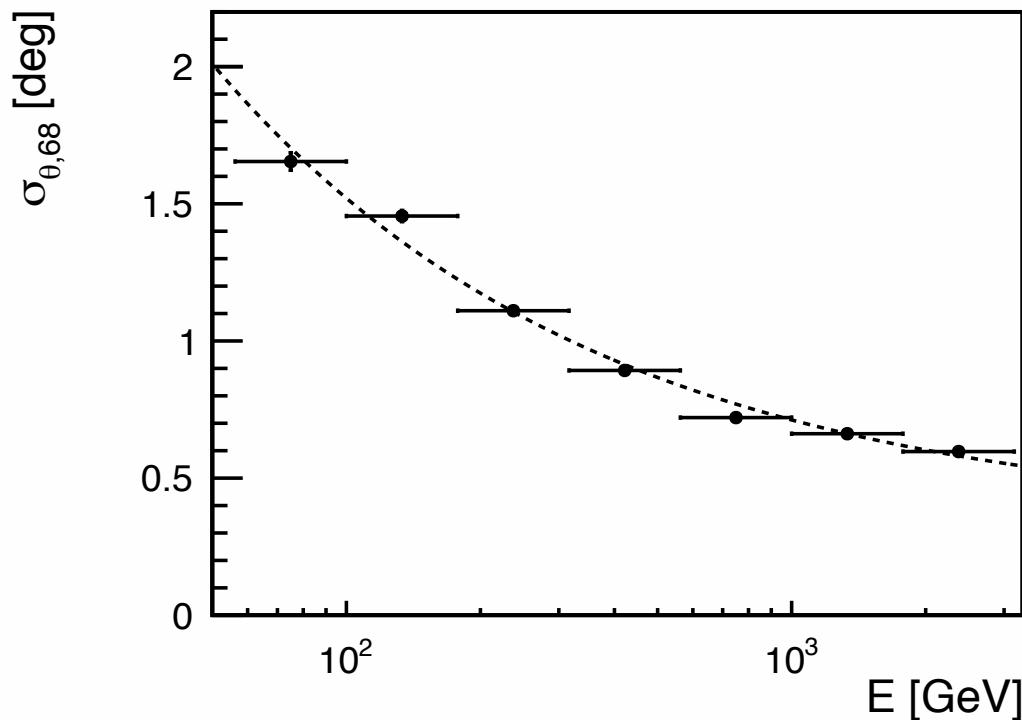
$E \rightarrow$ Reconstructed energy



- Use as **energy estimator** the **total signal** recorded by WCDs
- Energy resolution below 100% even at 100 GeV
 - Dominated by the shower fluctuations

Geometric reconstruction

γ – showers; $\theta = 10^\circ$

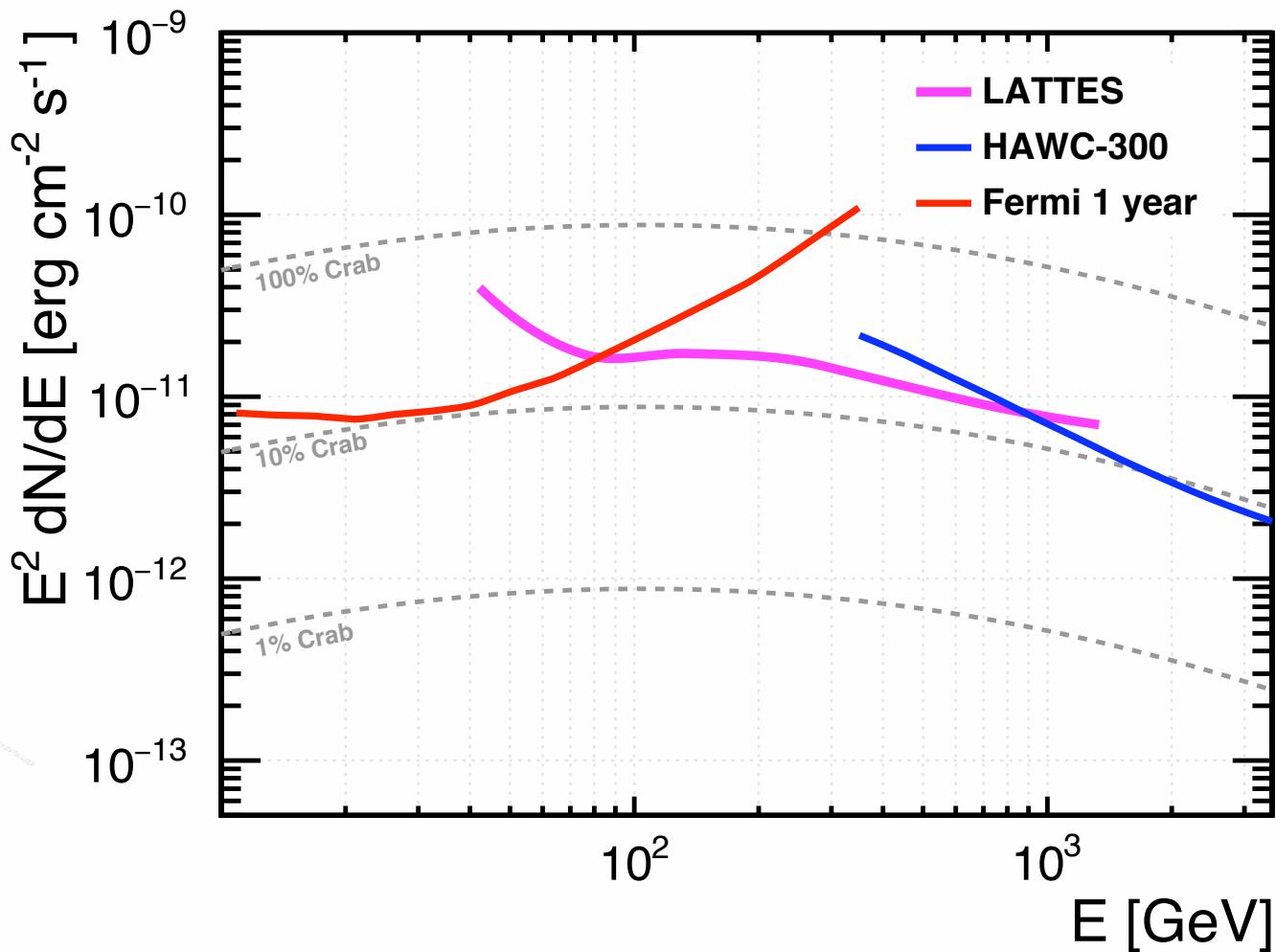


- Shower **geometry reconstruction** done using **RPC hit time**
 - Take advantage of RPCs **high spatial and time resolution**
 - Consider a time resolution of 1 ns
 - Use shower front plane approximation
 - Require more than 10 hits in the RPCs
- *Angular resolution below 2 deg even for 50 GeV showers*

- LATTES performance:
 - Trigger efficiency ✓
 - Energy Reconstruction ✓
 - Geometric Reconstruction ✓
 - Gamma-hadron discrimination
 - For now use a conservative approach:
 - Below 300 GeV don't consider any discrimination
 - Above 300 GeV use HAWC discrimination curve
- LATTES sensitivity

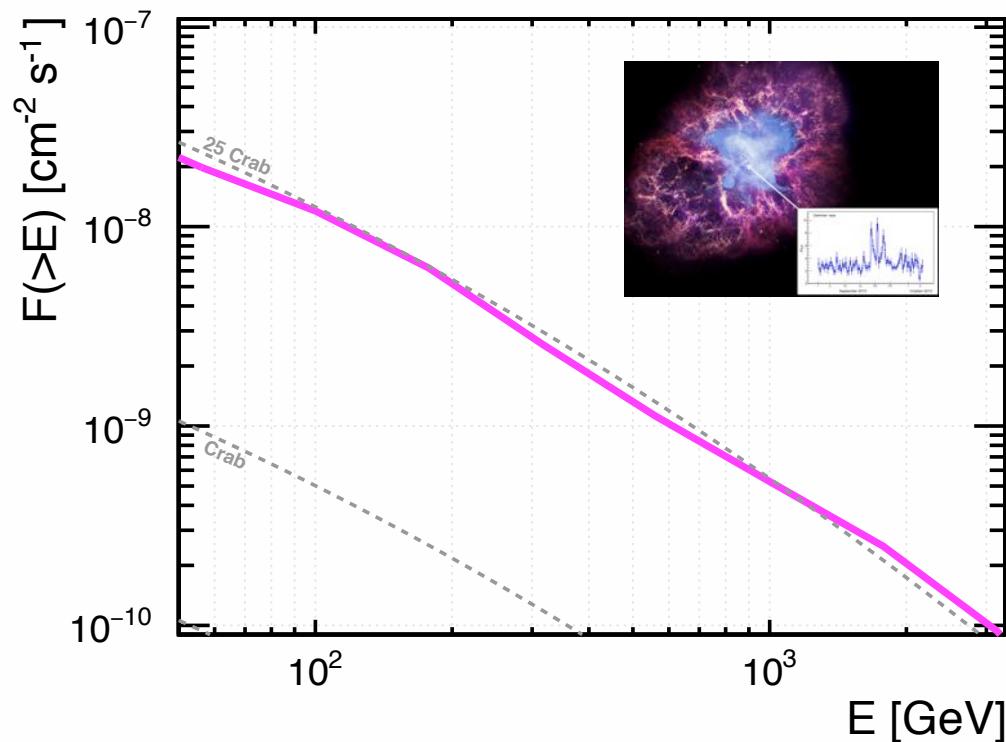
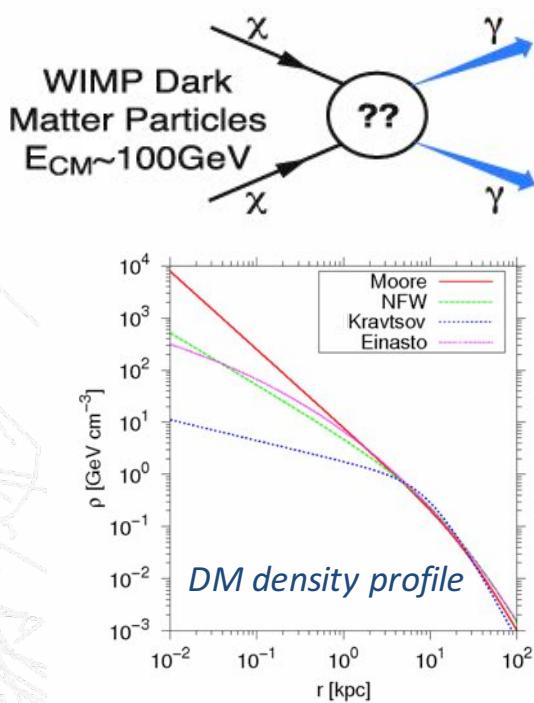


LATTES sensitivity



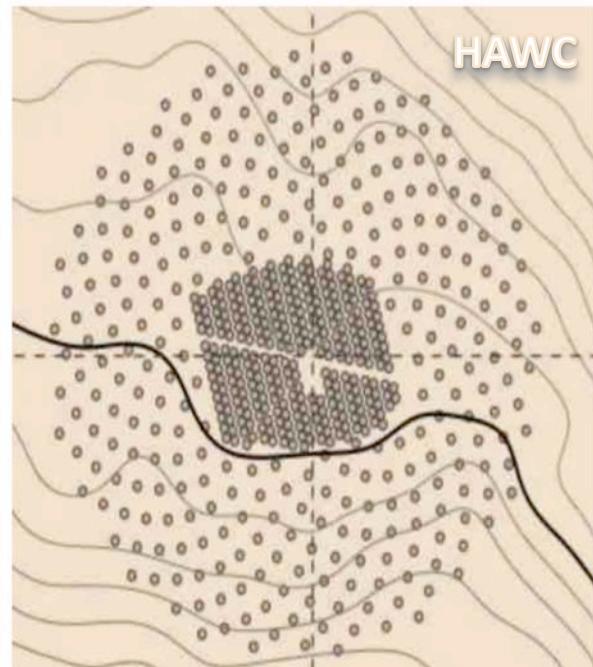
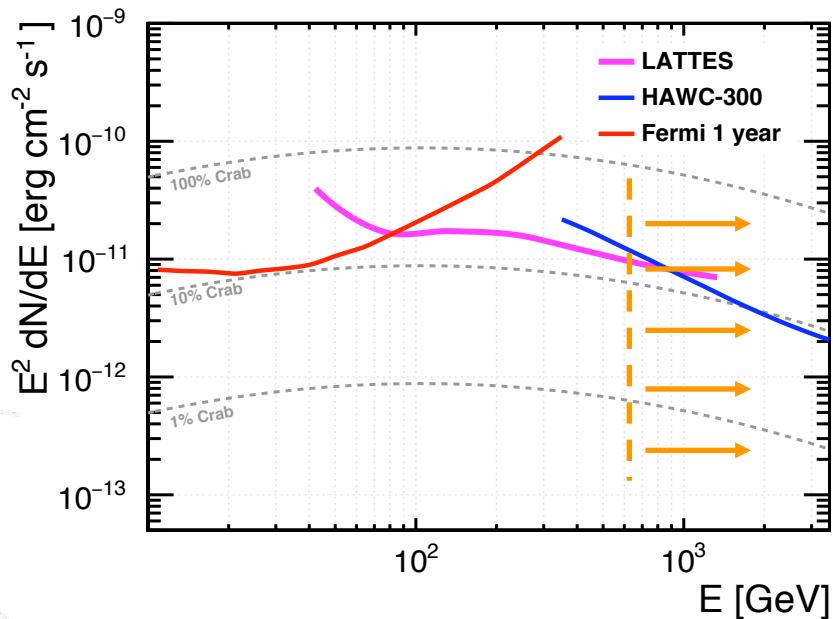
*Differential sensitivity to steady
sources in one year*

LATTES physics opportunities



- Many interesting scientific goals:
 - Dark matter searches at the **center of the galaxy**
 - Study **transient phenomena**
 - LATTES can detect a 25 Crab source at 3 sigma in 1 minute

LATTES at higher energies



- The sensitivity scales with the array area
- It could be extended to reach higher energies with an external corona of sparse detectors

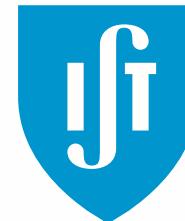
- LATTES: gamma ray wide field of view experiment at South America
 - Complementary project to CTA to survey the center of the galaxy
 - Next generation gamma-ray experiment (**hybrid**)
 - Good sensitivity at **low energies** (100 GeV)
 - Cover the gap between satellite and ground based measurements
 - Powerful tool to trigger observations **of variable source** and to **detect transients**



Acknowledgments



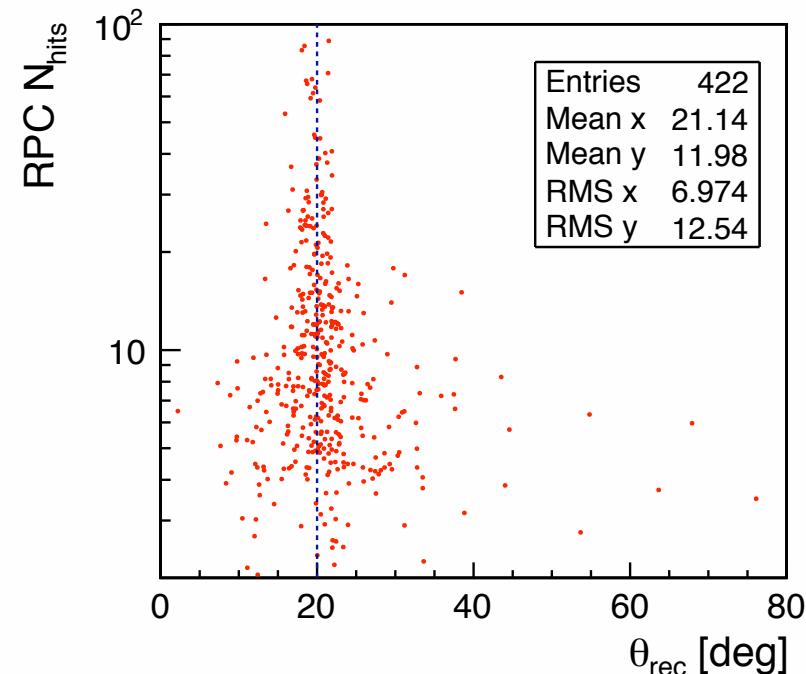
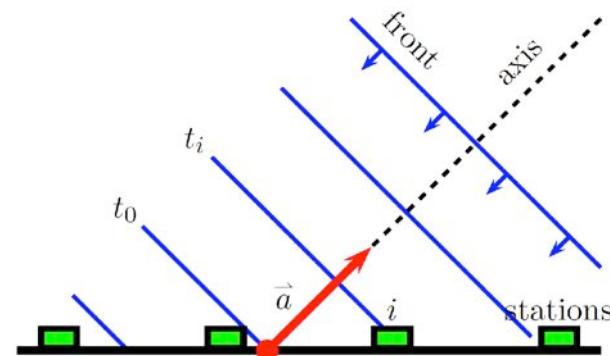
Fundação para a Ciência e a Tecnologia
MINISTÉRIO DA EDUCAÇÃO E CIÊNCIA



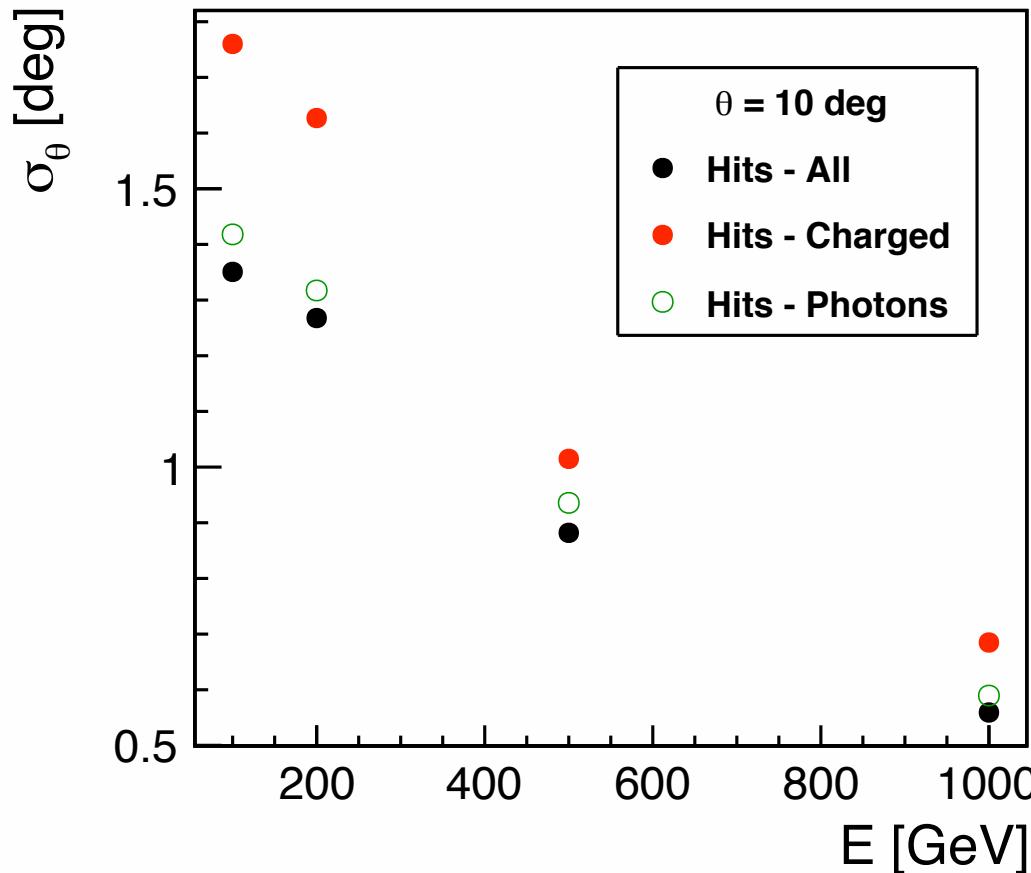
BACKUP SLIDES

Reconstruction of shower geometry

- **Use RPC hit time** information to reconstruct the shower
 - Take advantage of **high spatial and time resolution**
- Shower geometry reconstruction:
 - Use **shower front plane approximation**
 - Analytical procedure
 - Apply trigger conditions
 - Apply **cut on the number of registered hits** by the RPCs

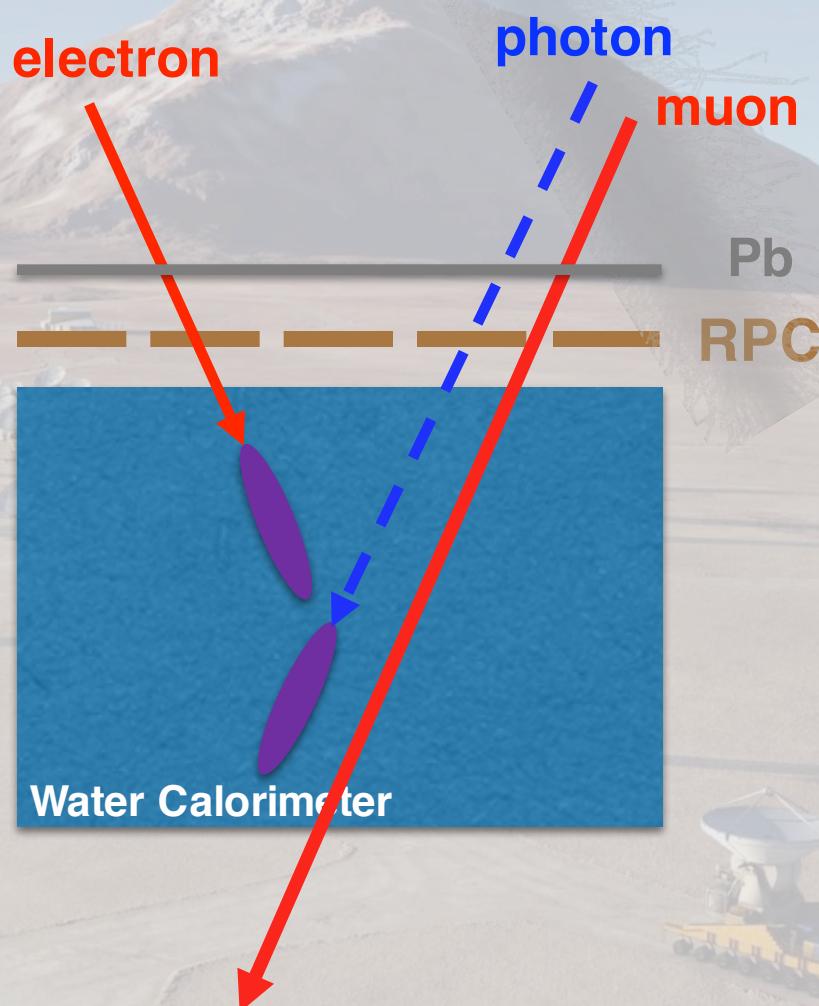


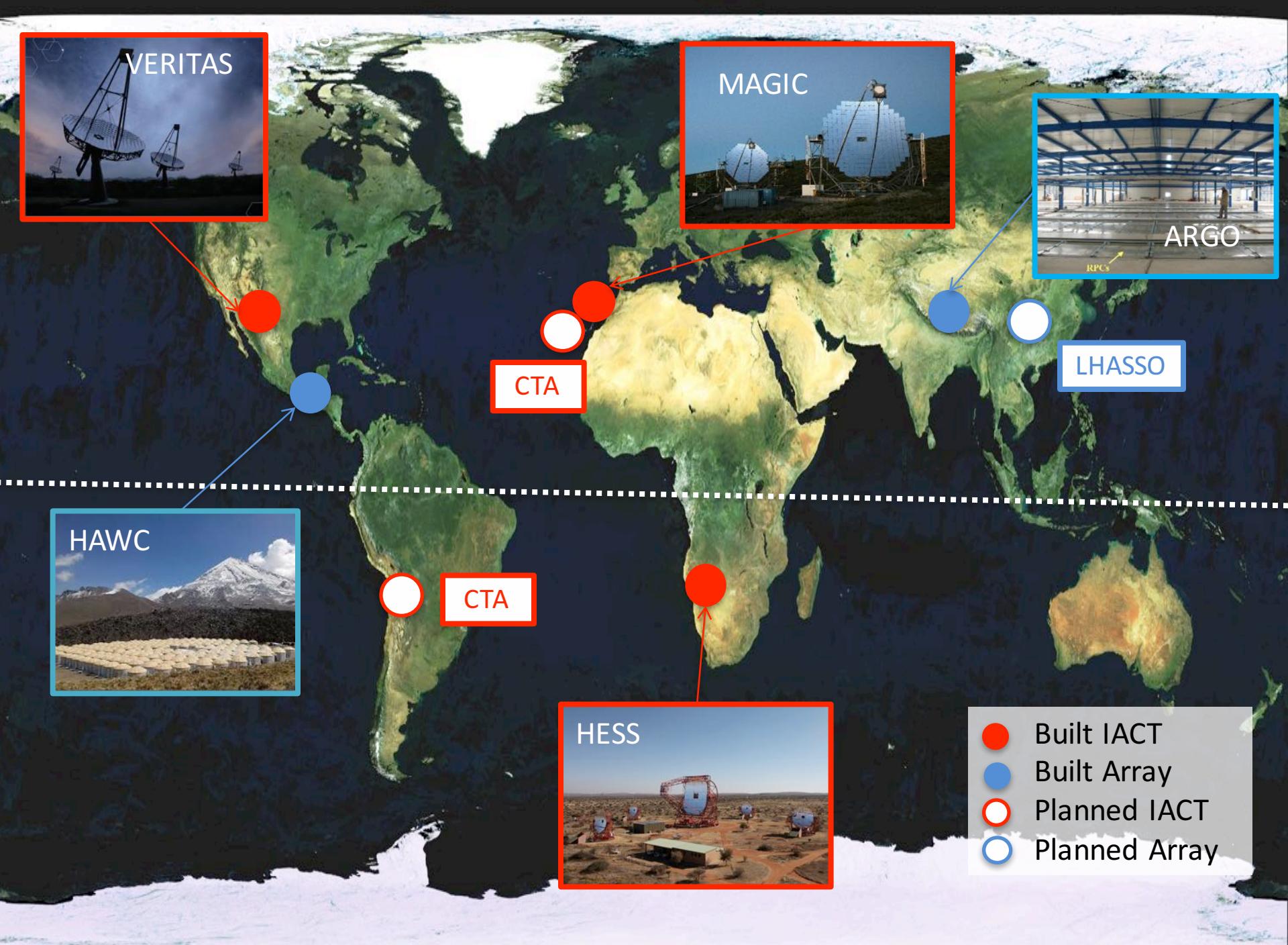
Contributions to the geometric reconstruction

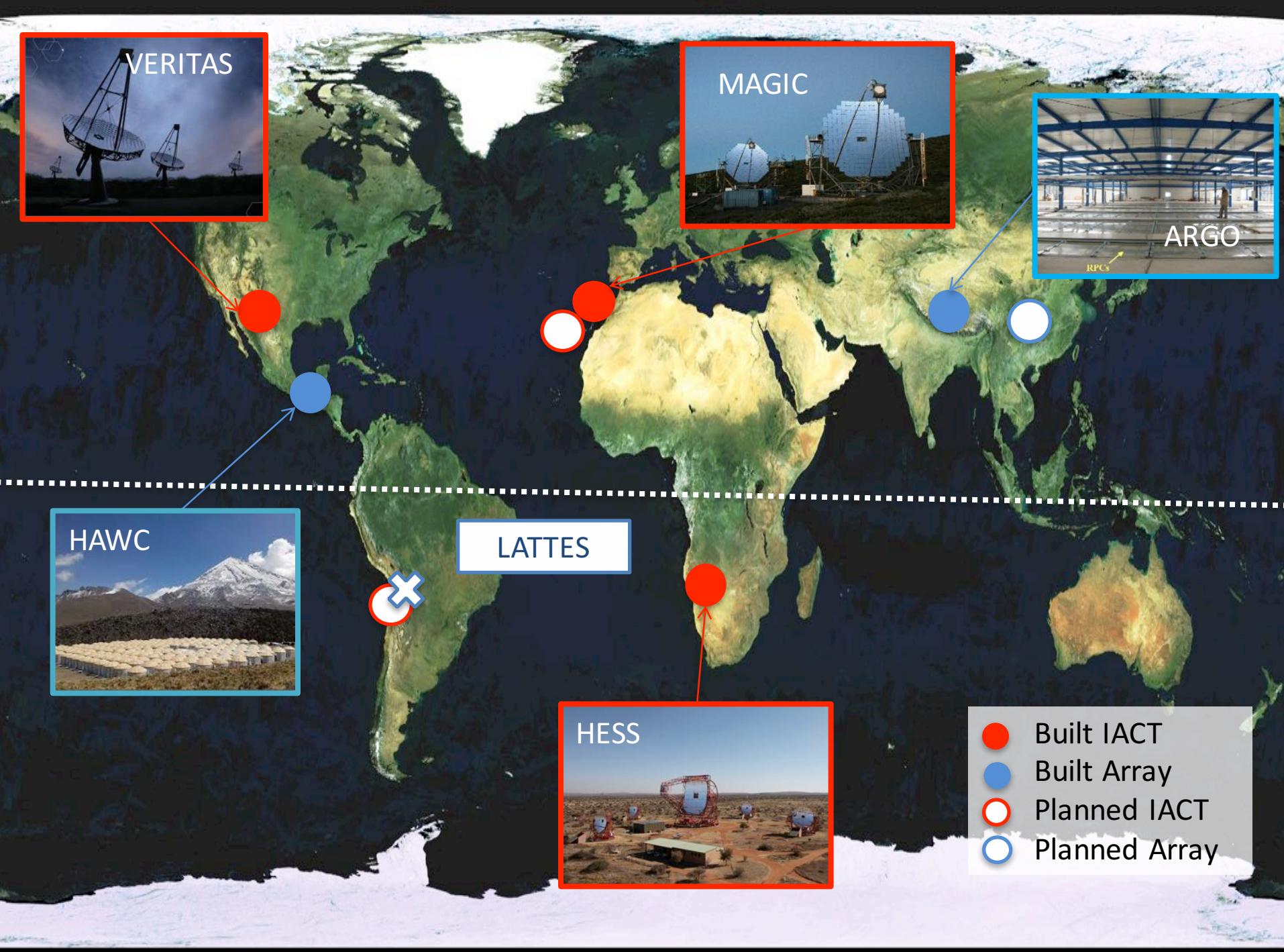


- **Photons** retain a **higher correlation** with the **shower geometry** than charged particles
- Could we measure photons with the RPC instead?

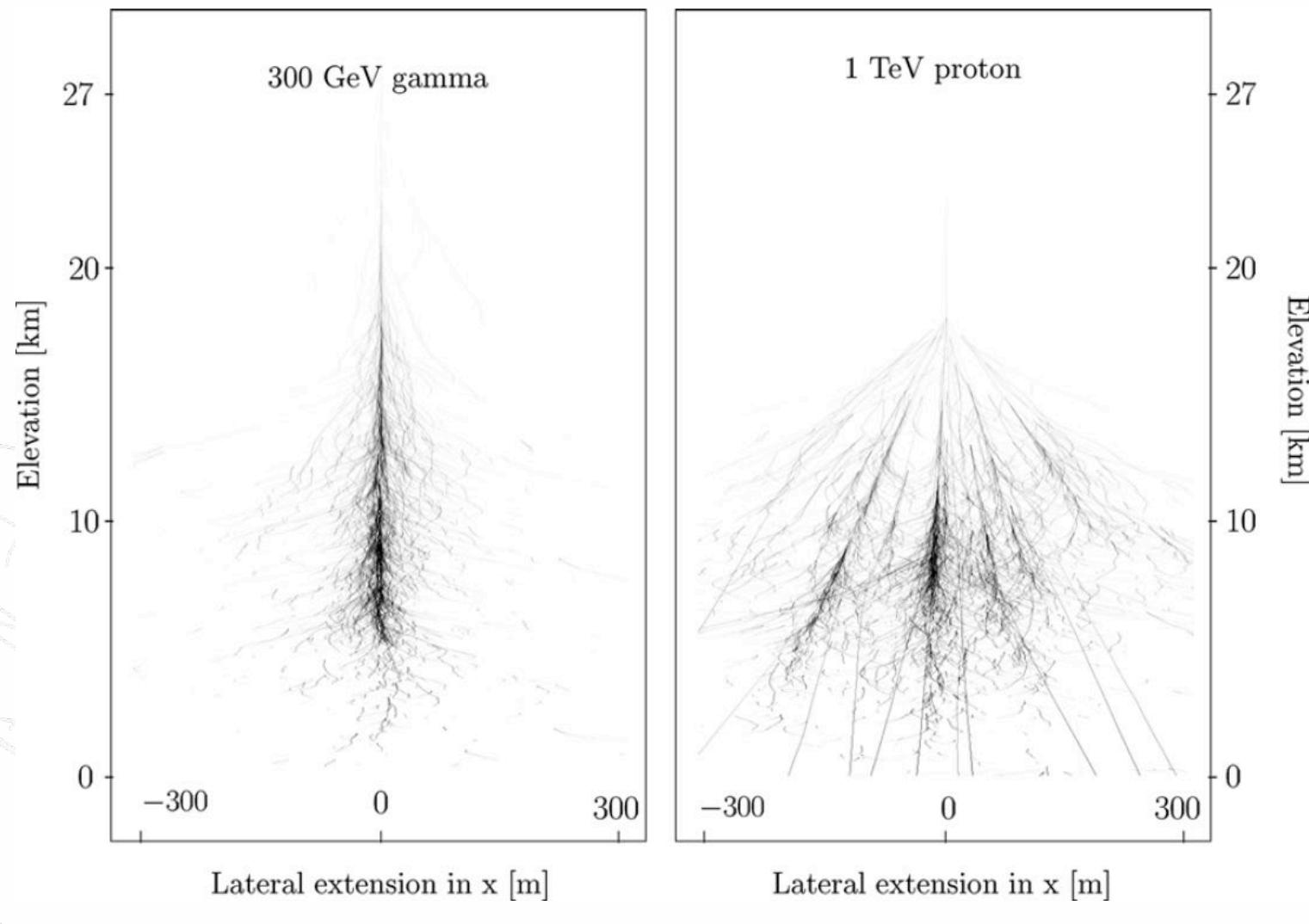
LATTES station baseline concept





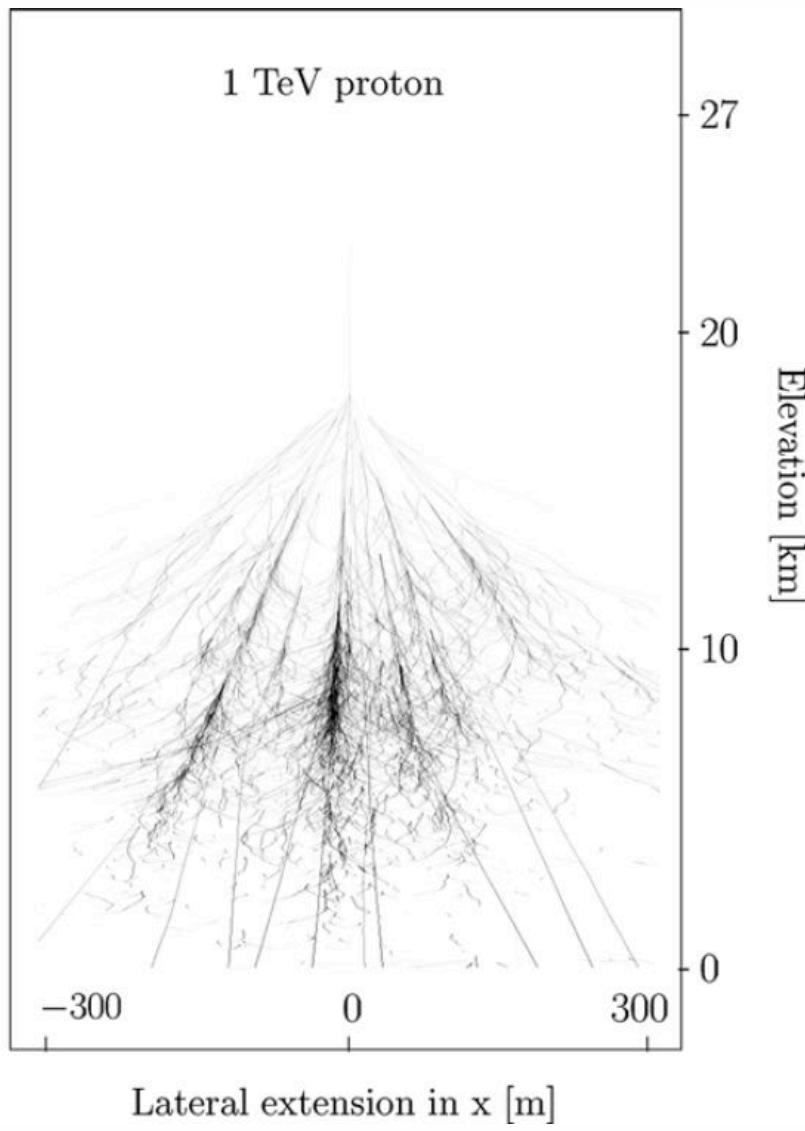


Strategies for primary discrimination



Explore differences in shower development

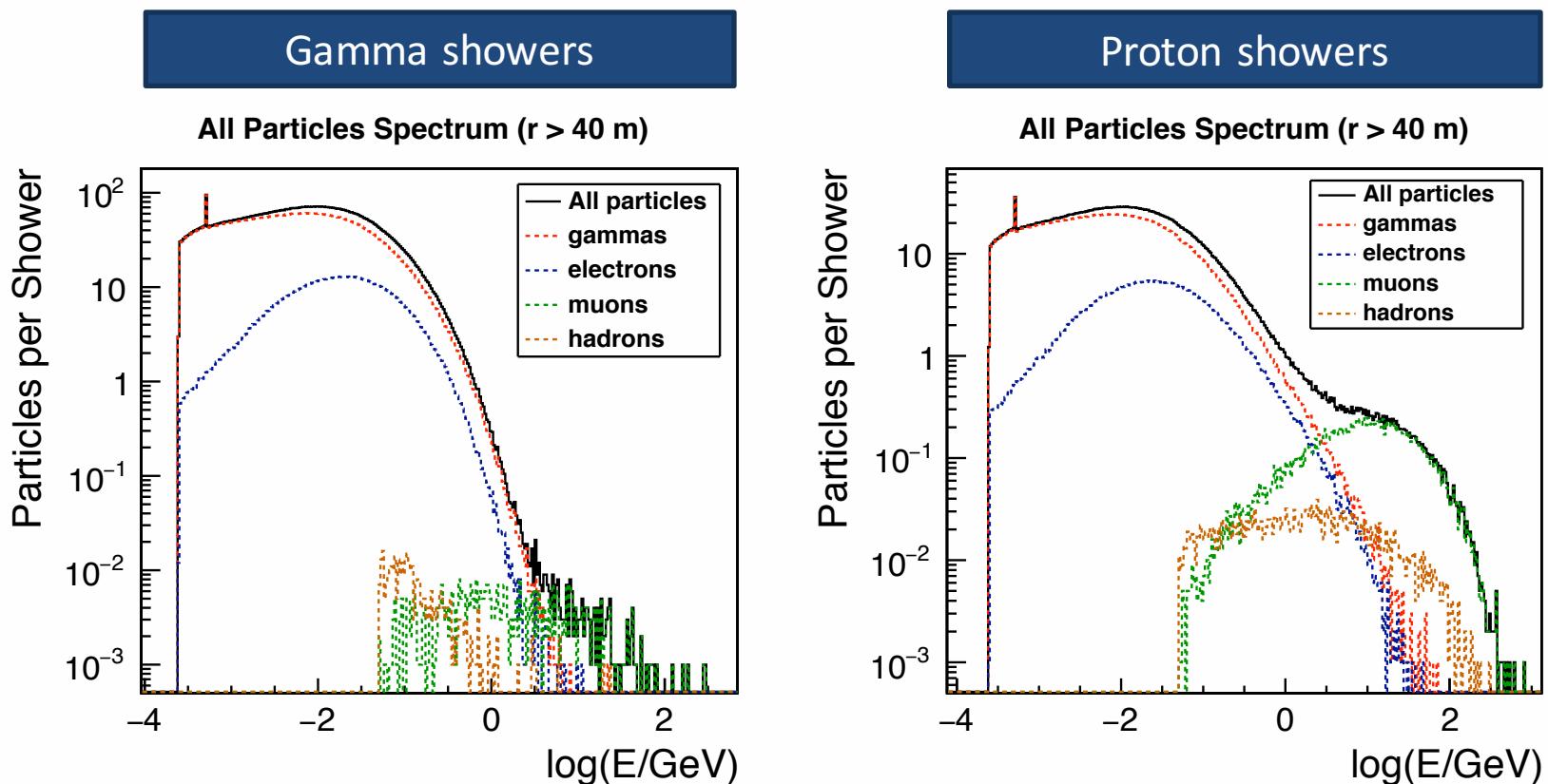
Strategies for primary discrimination



- **Hit pattern at ground**
 - Hits from hadronic showers are more sparse than in gamma induced showers
 - RPC detectors
 - Explored by the ARGO collaboration
- **Search for energetic clusters far from the shower core**
 - Present only in hadronic showers
 - Water Cherenkov Detectors
 - Explored by the HAWC collaboration
- **Combine both strategies using an hybrid detector: LATTES**
 - Work on-going...

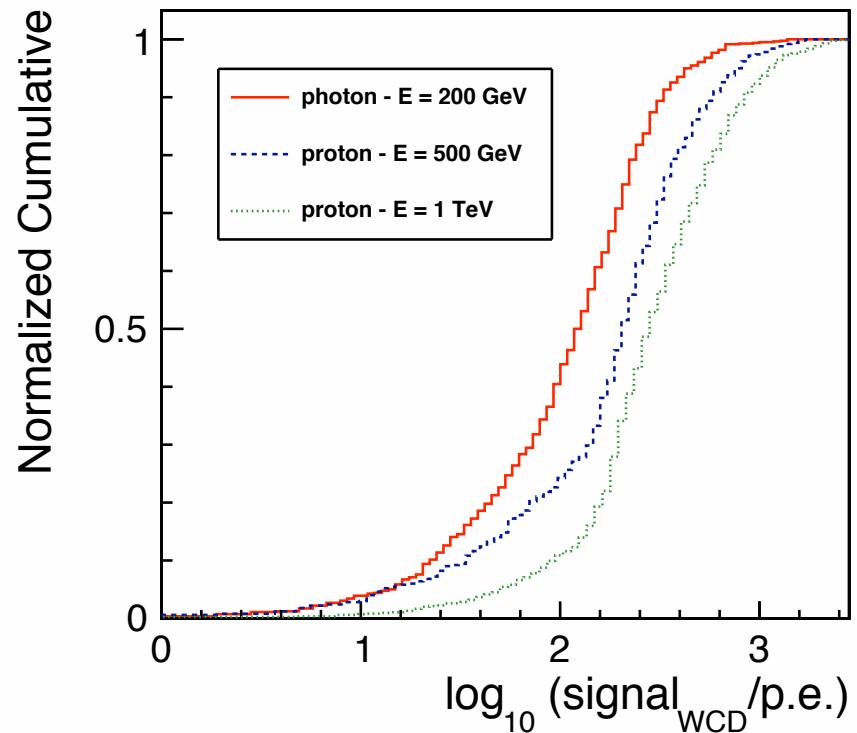
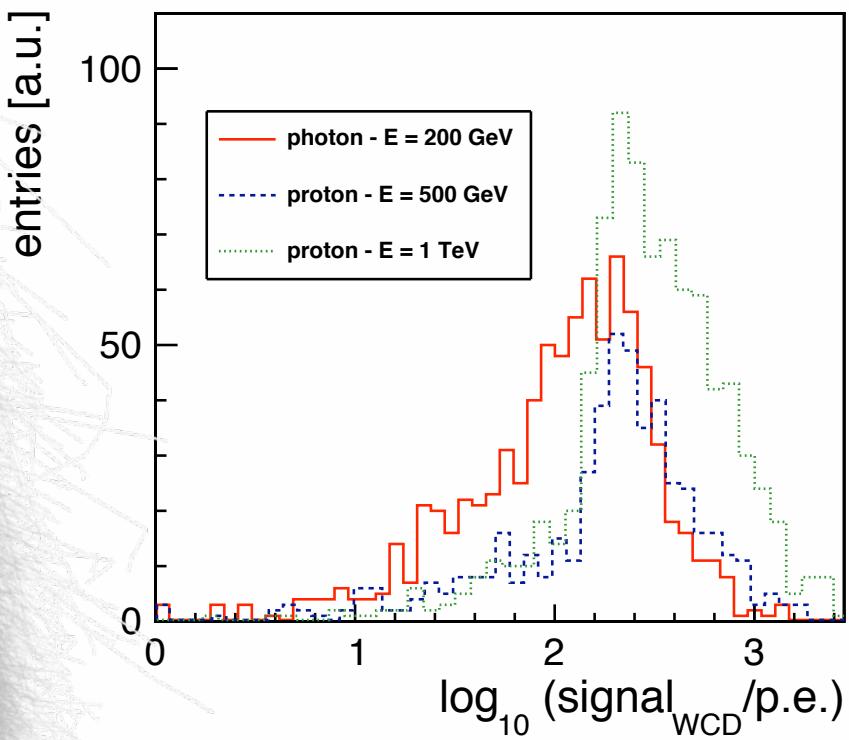
Exploring the WCD

5 TeV



- What should we look for?
 - Look for energetic clusters far from the shower core
 - Above 40 m

LATTES hottest station



- Signal of **the hottest WCD station**
 - above 40 m from the shower core
 - with **only one hit in the RPC**

LATTES integrated sensitivity

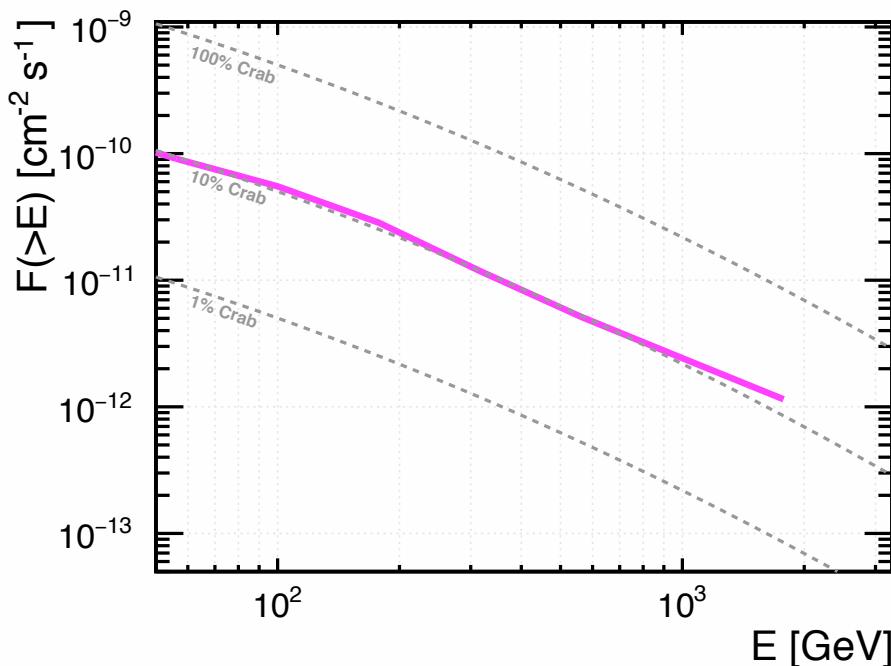
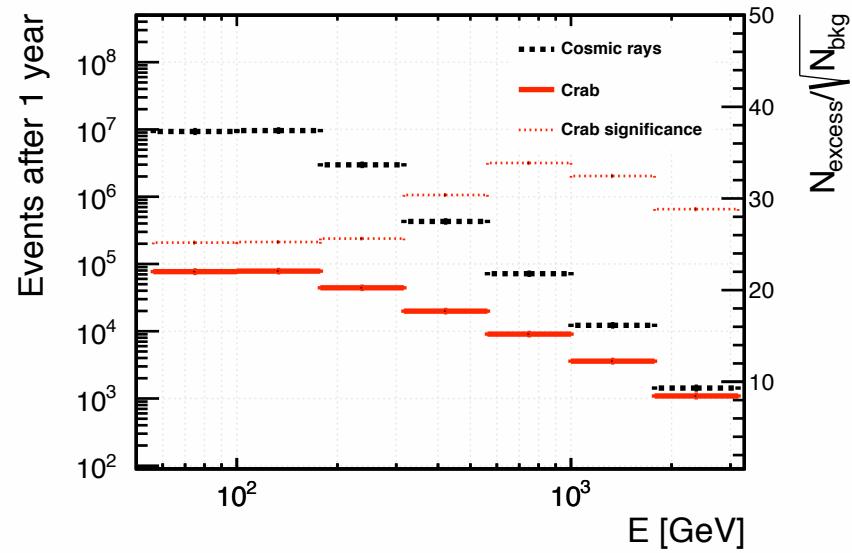
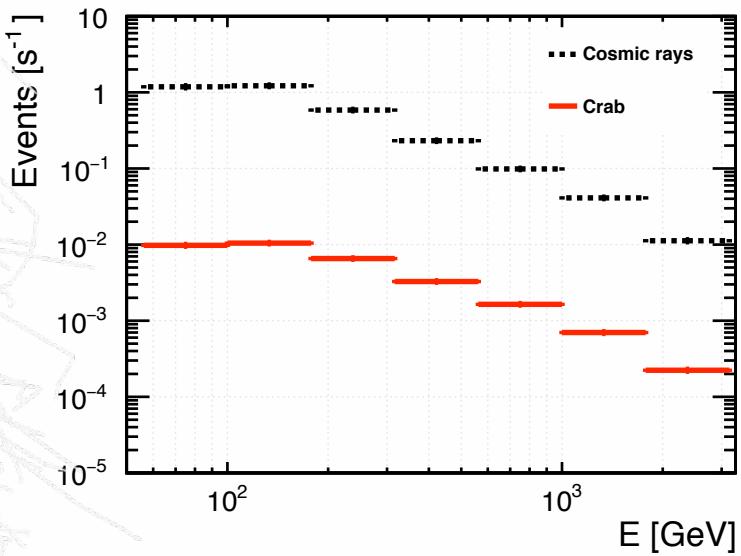


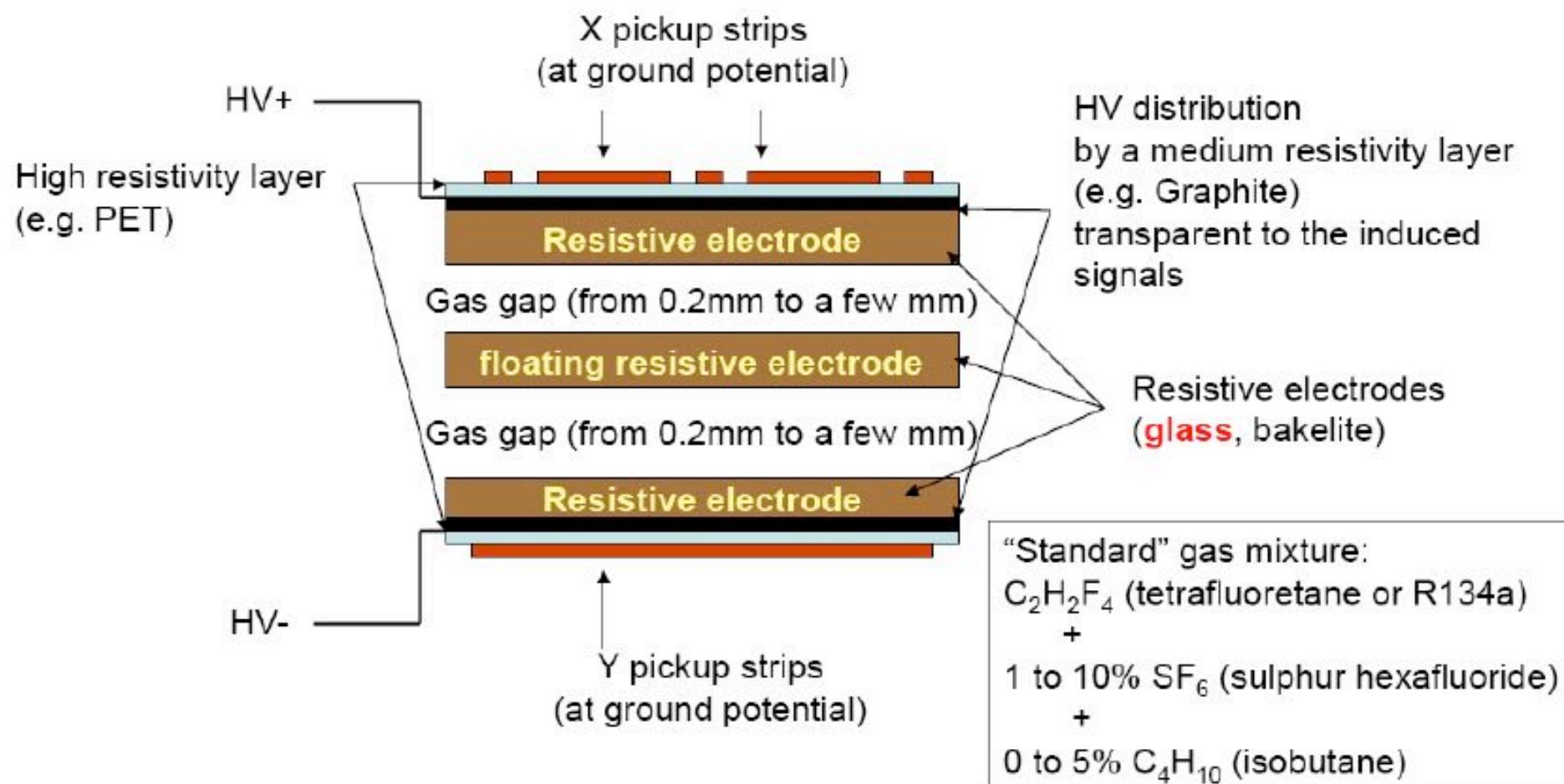
Figure 10: Integral sensitivity, defined as the flux of a source above a given energy for which $N_{\text{excess}} / \sqrt{N_{\text{bkg}}} = 5$ after 1 year; it is assumed that the SED is proportional to the SED of Crab Nebula. For comparison, fractions of the integral Crab Nebula spectrum are plotted with the thin, dashed, gray lines.

LATTES expect events from Crab



RPCs – basic structure

Many variations allowed



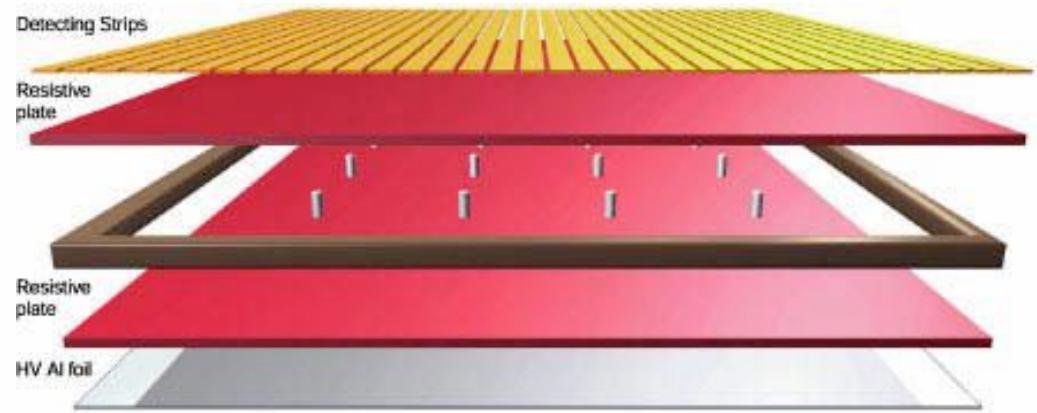
The current is limited by the resistive electrodes: no sparks by construction

- ▷ very safe detector, although limited to low particle rates ($\sim 2\text{kHz}/\text{cm}^2$)
- ▷ excellent efficiency (99%), time ($\sim 50\text{ ps}$) and position resolution ($\sim 100\mu\text{m}$)

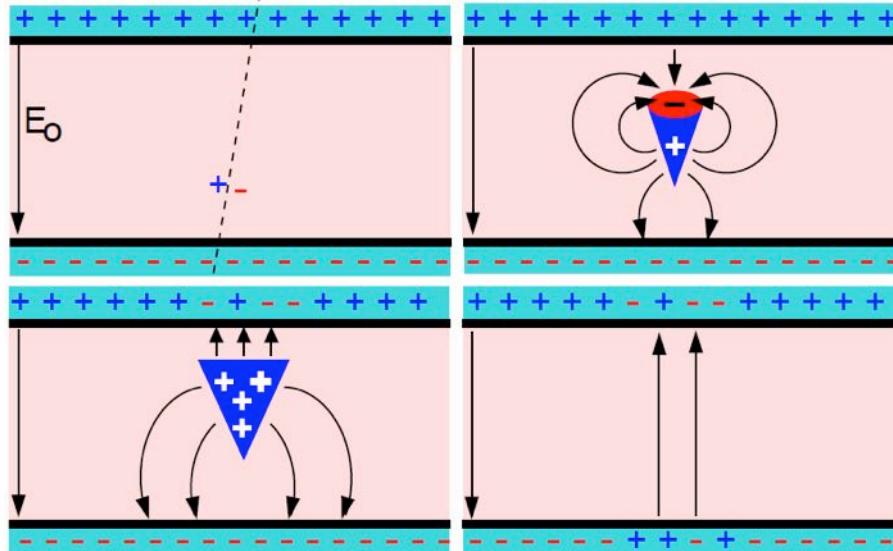
RPCs

Resistive Plate Chamber

- Gaseous detector
- Planar geometry
- uniform electrical field imposed.
- High resistive plates in between the electrodes limit the avalanche current.
- Signal is picked up by the induction of the avalanche in the readout pads.



Avalanche mode



Streamer mode

