

Clustering phenomena in ¹⁸ O Melina Avila

Santa Tecla, September 22, 2011

α -structure of N=Z



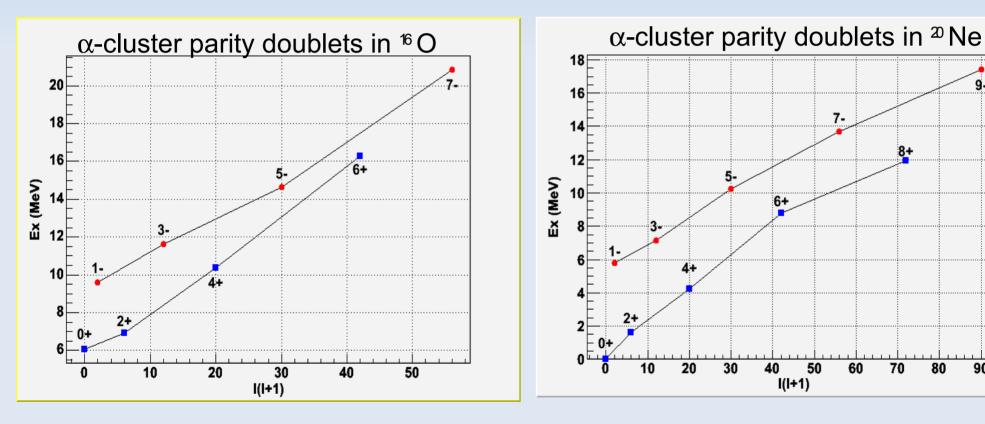
g.

8+

70

80

90



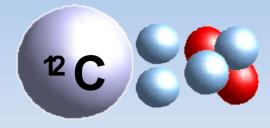
α-structure of N≠Z

- The α -clustering in N \neq Z nuclei is less studied
- Instrumental to understad the interplay between α-clustering and single particle degrees of freedom.
- Exotic Molecular type configurations have been predicted

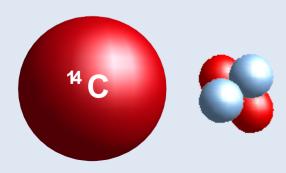


W. von Oertzen, Z. Phys. A354 (1996)

Why ¹⁸O is important?







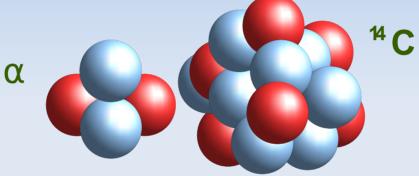
Molecular Type State

Shell model Structure

Cluster model Structure

¹⁴C(α,α)¹⁴C elastic scattering

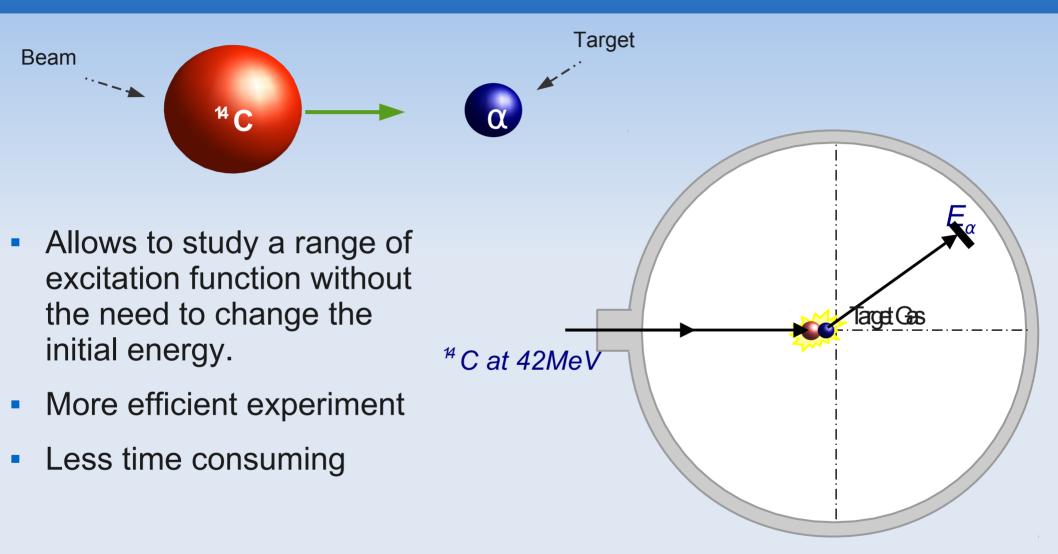
The structure of ¹⁸O at exitation energies above the α -decay threshold was studied using ¹⁴C+ α elastic scattering.



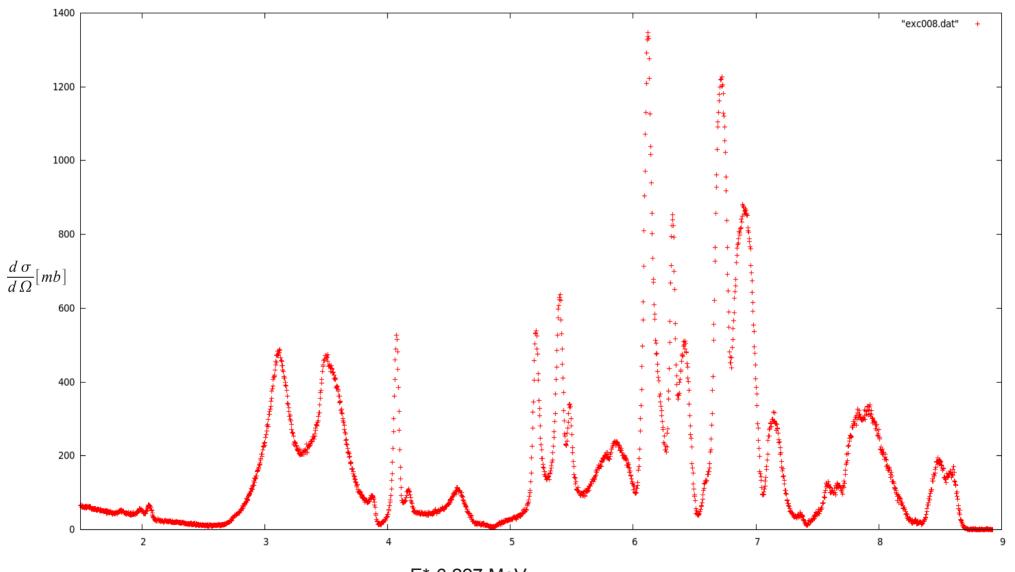
¹⁴C beam was produced by John D. Fox Superconducting Linear accelerator facility at Florida State University. $E(^{14}C) = 42 \text{ MeV}$.

Excitation function of resonance elastic scattering of ${}^{4}C+\alpha$ was measured using Method of Thick Target and Inverse Kinematics (TTIK).

Thick Target and Inverse Kinematics



Experimental data



E*-6.227 MeV

R-Matrix analysis

- multi-level, multi-channel approach

The elastic scattering data was fit using an R-Matrix analysis consisting of 3 open channels, ${}^{4}C(\alpha,\alpha)$, ${}^{4}C(\alpha,n)^{7}O$, and ${}^{4}C(\alpha,n)^{7}O^{*}$

> α + ¹⁴C(0⁺,g.s.) n + ¹⁷O(5/2⁺,g.s.) n + ¹⁷O(1/2⁺,0.87)

$$R_{cc'} = \sum_{\lambda} \frac{\gamma_{\lambda c} \gamma_{\lambda c'}}{E_{\lambda} - E}$$

$$\Gamma_{tot} = \Gamma_{\alpha} + \Gamma_{n}$$
$$\Gamma_{\alpha} = 2P\gamma_{\alpha}^{2}$$

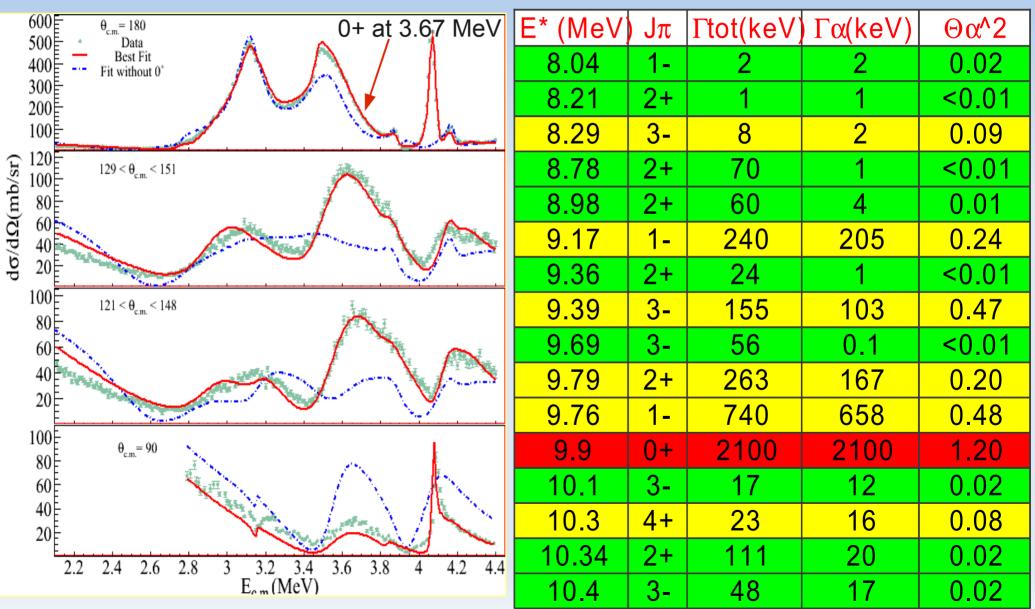
$$\theta_{\alpha}^{2} = \frac{\Gamma_{\alpha}}{\Gamma_{sp}}$$

 Γ_{sp} Was obtained using Potential model calculations [Robson, PRL 42 (1979)]

Resonances in ¹⁸O observed via ¹⁴C+α

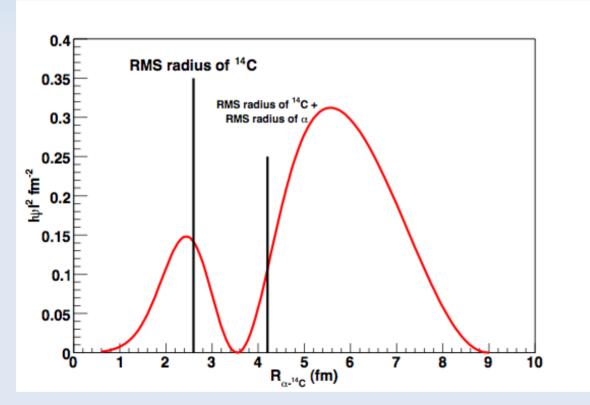
E.D. Johnson, et al., EPJA, 42 135 (2009)

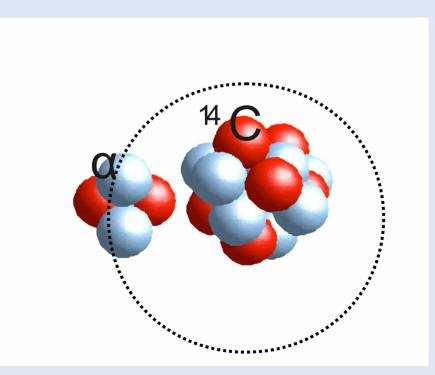
16 States were used to fit the data



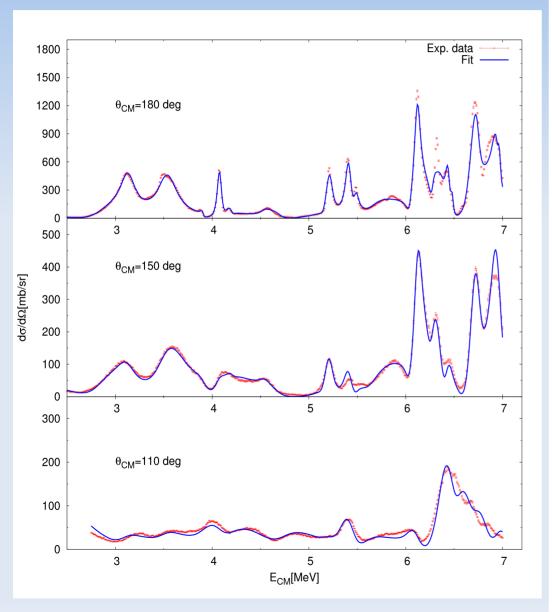
0⁺ α halo state in ¹⁸ O

0⁺ at 3.7+/-0.3 MeV (~10 MeV ¹⁸O excitation energy) with width of ~2 MeV is necessary to fit the α +¹⁴C data. This width corresponds to purely α particle state.



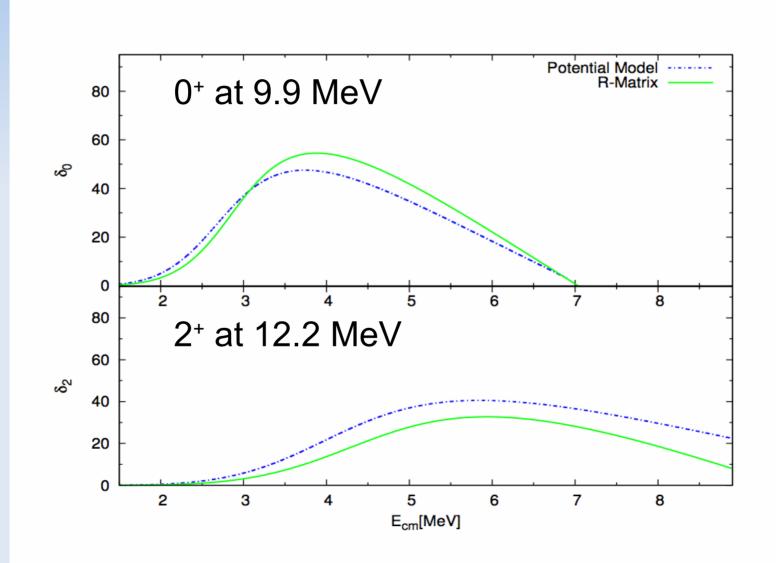


Experimental fit

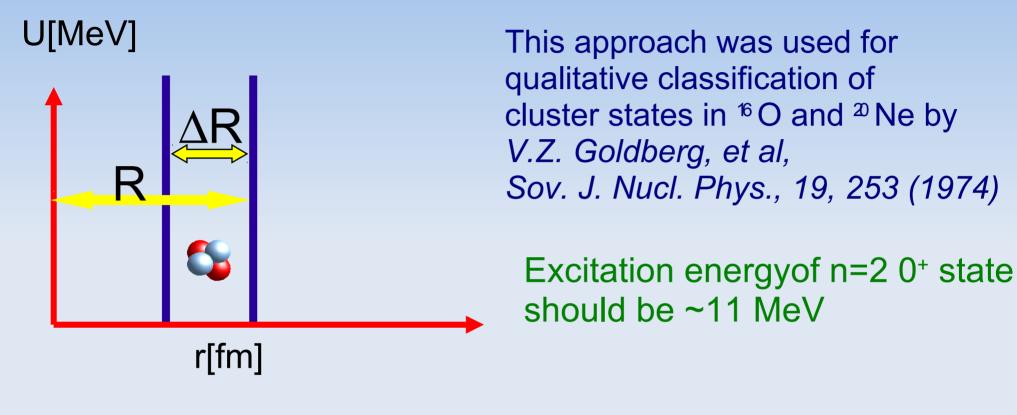


$E_{\rm exc}$	J^{π}	$\Gamma_{\rm tot}$	Γα	θ_{lpha}^2	$E_{\rm exc}$	J^{π}	$\Gamma_{\rm tot}$	Γα	θ_{α}^2
8.04	1-	2	2	0.02	11.47	1-	277	194	0.07
8.21	2^{+}	1	1	0.01	11.65	5^{-}	48	26	0.11
8.29	3^-	8	2	0.09	11.69	1^{-}	191	119	0.04
8.78	2^{+}	70	1	0.01	11.69	2^{+}	200	19	0.01
8.98	2^{+}	60	4	0.01	11.88	6^{+}	27	12	0.20
9.17	1^{-}	240	205	0.24	11.96	3^-	210	48	0.03
9.36	2^{+}	24	1	0.01	12.17	2^{+}	5700	5700	1.46
9.39	3^{-}	155	103	0.47	12.25	3^-	123	12	0.01
9.69	3^-	56	0.1	0.01	12.32	5^{-}	69	38	0.09
9.79	2^{+}	263	167	0.20	12.34	1-	235	154	0.05
9.90	0^{+}	2100	2100	1.20	12.38	6^{+}	82	50	0.46
10.10	3^-	17	12	0.02	12.43	2^{+}	482	472	0.17
10.30	4^{+}	23	16	0.08	12.47	1^{-}	361	355	0.11
10.34	2^{+}	111	20	0.02	12.71	3^{-}	660	624	0.28
10.40	3^-	48	17	0.02	12.72	1^{-}	110	105	0.03
10.73	3^-	185	13	0.02	13.03	5^{-}	99	70	0.10
10.89	1^{-}	225	201	0.09	13.10	5^{-}	60	43	0.06
11.39	3^-	197	101	0.08	13.12	1^{-}	127	56	0.02
11.41	4+	67	44	0.07	13.25	4+	259	258	0.16

Phase shift calculations

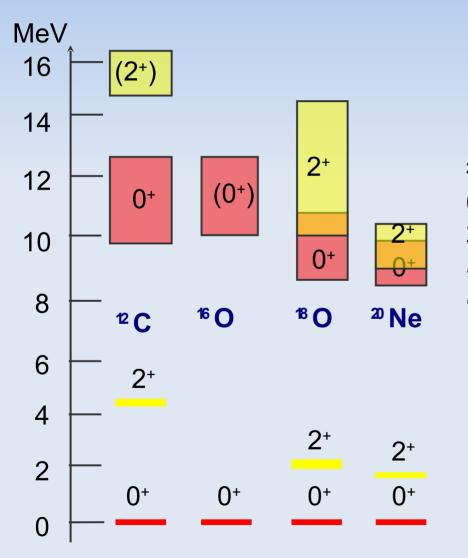


Infinite surface potential



$$E = \frac{n^2 \pi^2 \hbar^2}{2\mu (\Delta R)^2} + \frac{\hbar^2 \ell (\ell+1)}{2\mu R^2}$$

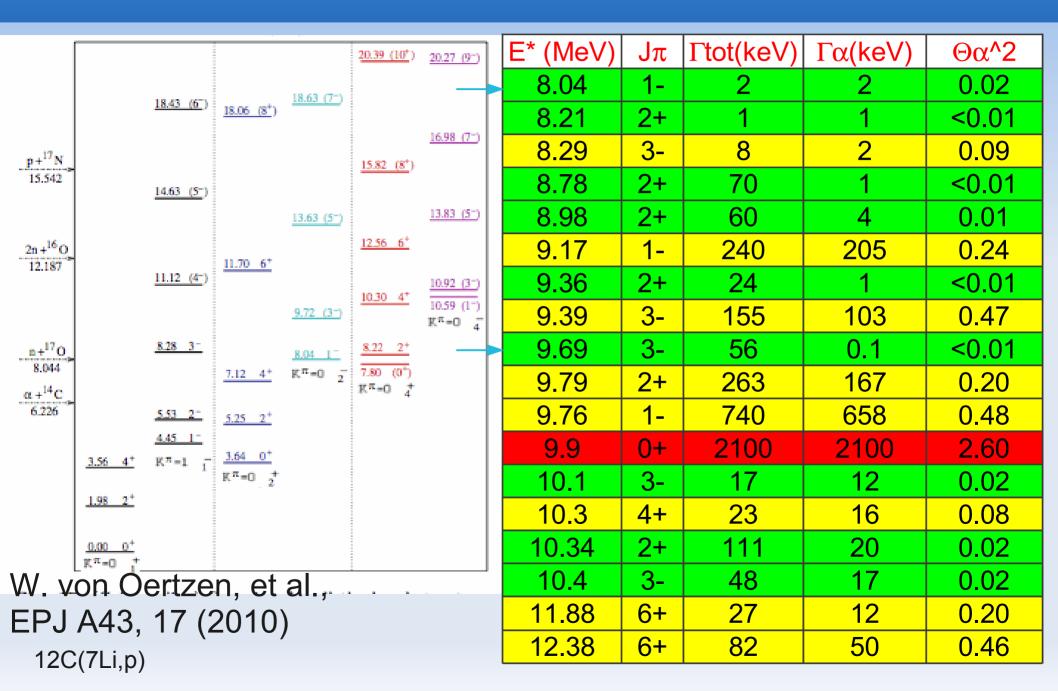
"α-halo" states in light nuclei



¹²C: 0⁺ at 11.25 MeV, Γ=2.5 MeV *H. Fynbo, et al., Nature 4*33 (2005) 136

²⁰Ne: 0⁺ at 8.6 MeV, Γ=1.47 MeV 2⁺ at 8.9 MeV, Γ=1.2 MeV *H. Shen, et al., NIM B 90 (1994) 593 L.C. McDermott, et al., PR 118 (1960) 175*

Negative parity rotational band



These states may play important role in (α, γ) ; (α ,p) and (α ,n) reactions relevant for various stellar nucleosynthesis processes.

Like in the mirror nucleus: ¹⁸ Ne(α ,p)

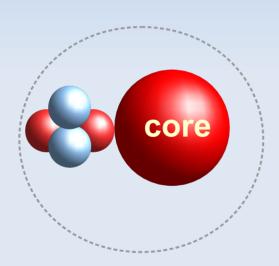
Summary

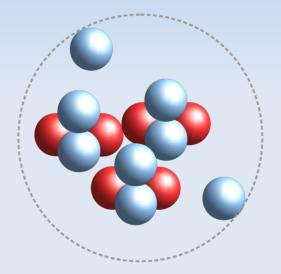
- Several α-cluster states were found in ¹⁸O including in the energy range of 2-7 MeV.
- Broad resonances that correspond to pure α-core structures ("α-halo" states) have been known to exist in ¹² C, ²⁰ Ne and now observed in ¹⁸ O.
- It is speculated that "α-halo" states are common phenomena for light nuclei.
- Potential model was able to predict these broad states.

Thank you!

Clusters in nuclear matter

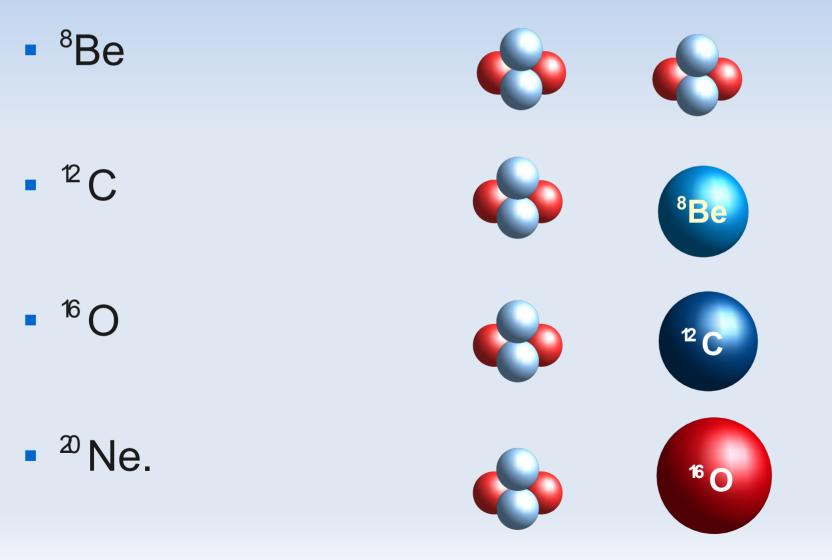
 Protons and neutrons inside the nucleus form into one or more clusters





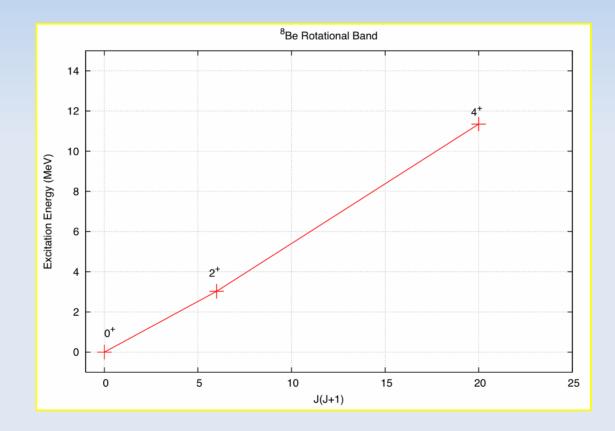
α -structure for N=Z

Well known alpha clusters in :

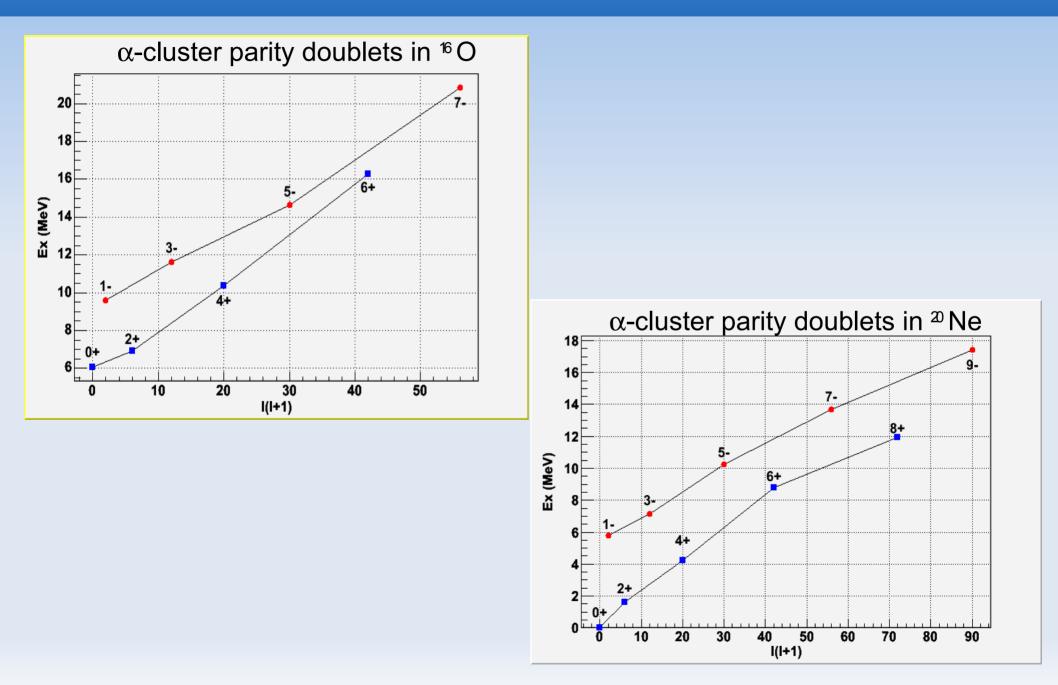


Rotational Bands

⁸Be Rotational Band



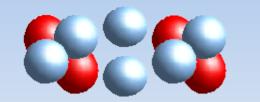
Rotational bands



α-structure of N≠Z

- The α -structure of N \neq Z in not well known
- Instrumental to understad the interplay between α-clustering and single particle degrees of freedom.
- Useful to understand Molecular type states

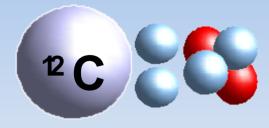
Clustering in N≠Z nuclei. Molecular type states



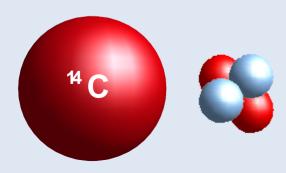


W. von Oertzen, Z. Phys. A354 (1996)

Why *o is important?







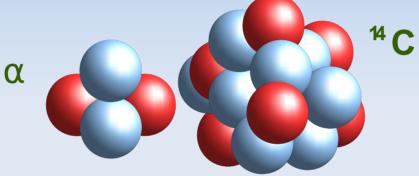
Molecular Type State

Shell model Structure

Cluster model Structure

¹⁴ C(α , α)¹⁴ C elastic scattering

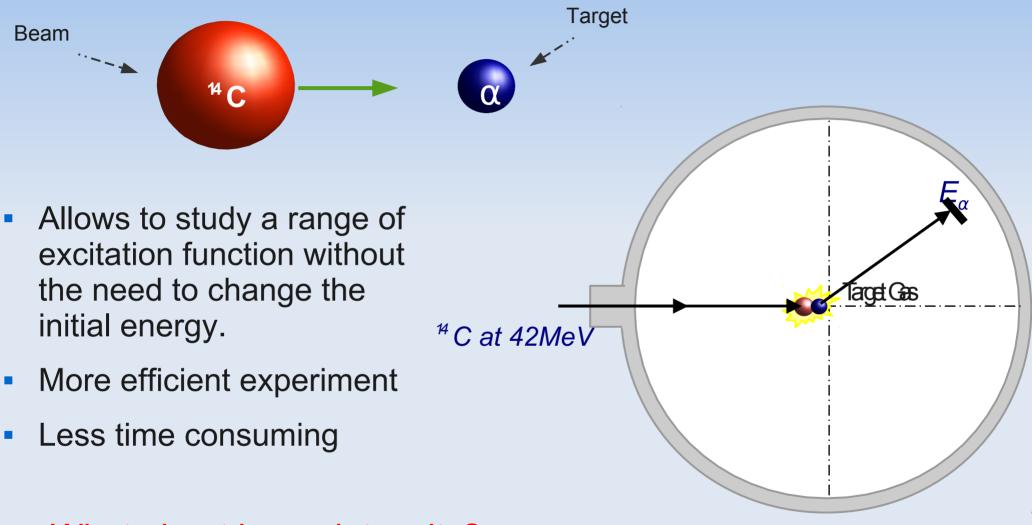
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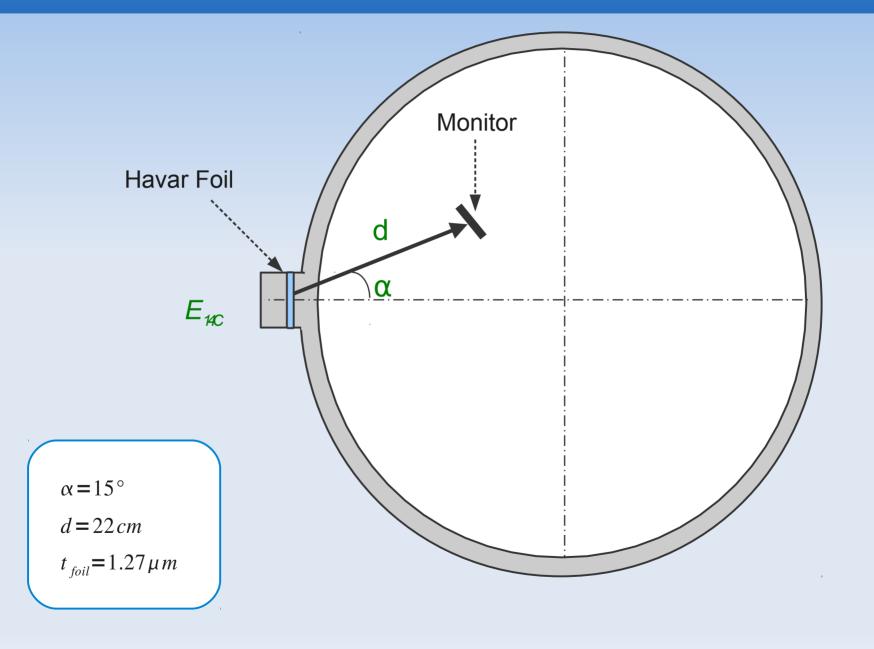
Excitation function of resonance elastic scattering of ${}^{4}C+\alpha$ was measured using Method of Thick Target and Inverse Kinematics (TTIK).

Thick Target and Inverse Kinematics



What about beam intensity?

Determination of the beam intensity



Determination of the beam intensity

$$\left(\frac{d\sigma}{d\Omega_{cm}}\right) = \left(\frac{\alpha \hbar c Z_1 Z_2}{4E_{cm}}\right)^2 \frac{1}{\sin^4(\theta/2)} \left(\frac{d\sigma}{d\Omega_{cm}}\right) = \frac{1}{I} \frac{N}{kt \Delta \Omega}$$

$$I = \frac{N}{k t \Delta \Omega} \left(\frac{4 E_{cm}}{\alpha \hbar c Z_1 Z_2} \right)^2 \sin^4(\theta/2)$$

R-Matrix analysis

The elastic scattering data was fit using an R-Matrix analysis consisting of 3 open channels, $14C(\alpha,\alpha)$, $14C(\alpha,n)17O$, and $14C(\alpha,n)17O^*$

 α + ¹⁴ C(0⁺,g.s.) n + ¹⁷ O(5/2⁺,g.s.) n + ¹⁷ O(1/2⁺,0.87)

R-Matrix analysis

External region: Channels

Internal Region: Compound Nucleus

Introduced parameters: E, l, γ_{α}

(Energy, angular momentum, reduced width)

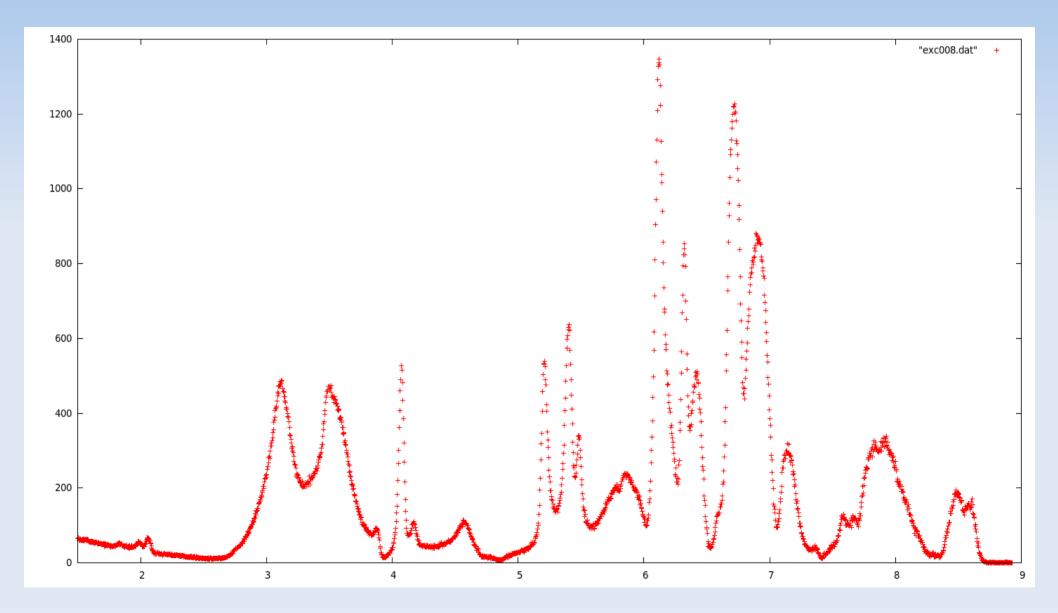
$$R_{cc'} = \sum_{\lambda} \frac{\gamma_{\lambda c} \gamma_{\lambda c'}}{E_{\lambda} - E}$$

$$\gamma_{sp}^{2} = \frac{\hbar^{2}}{\mu R^{2}}$$
$$\theta_{\alpha}^{2} = \frac{\gamma_{\alpha}^{2}}{\gamma_{sp}^{2}}$$

$$\gamma_{\lambda\alpha} \sim <\!\psi_{180} \, |(\psi_{14\mathrm{C}} \otimes \psi_{\alpha}) \!>$$

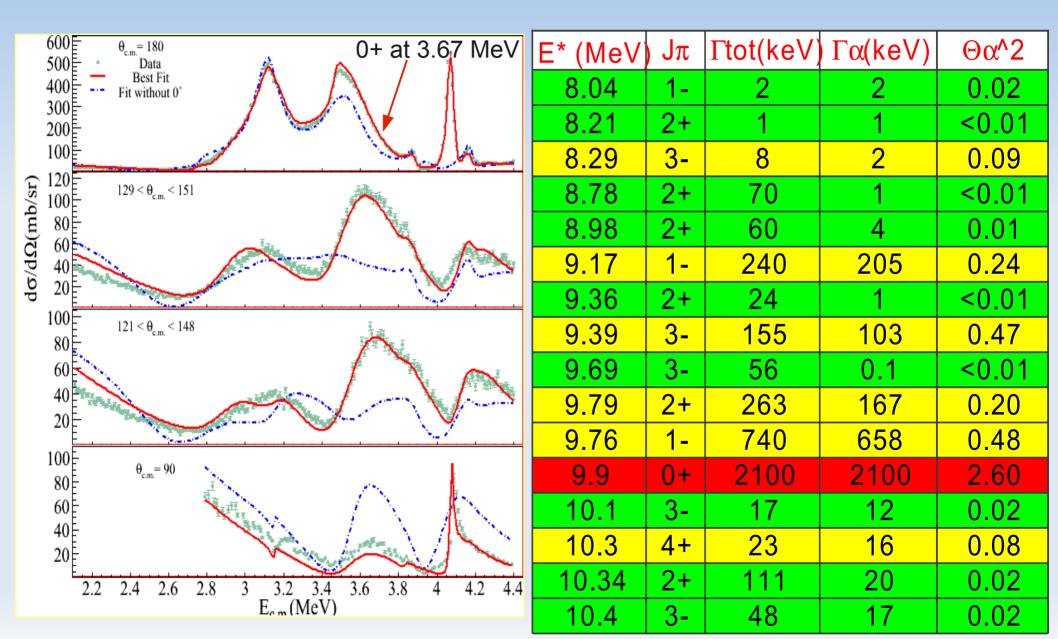
$$\Gamma_{tot} = \Gamma_{\alpha} + \Gamma_{n}$$
$$\Gamma_{\alpha} = 2P \gamma_{\alpha}^{2}$$

Experimental data

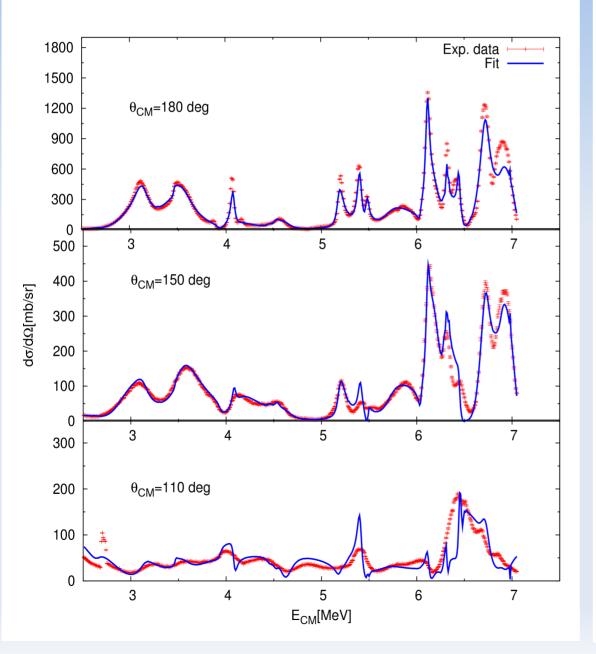


Resonances in ¹⁸ O observed via ¹⁴ C+ α

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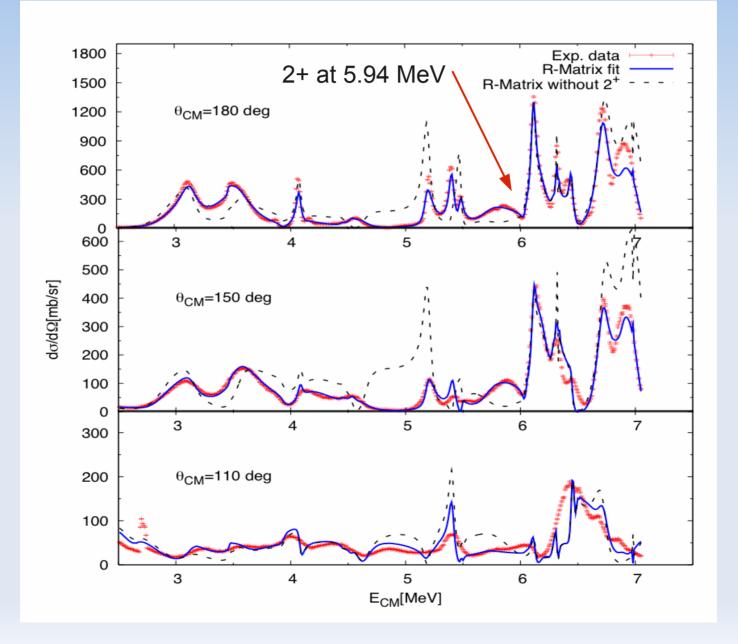


Experimental fit



	$E_{\rm exc}$	J^{π}	Γ_{tot}	Γ_{α}	θ_{α}^{2}
	10.73	3-	185	13	0.02
	10.89	1^{-}	225	201	0.09
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	13.25	4+	259	258	0.16

Broad 2⁺

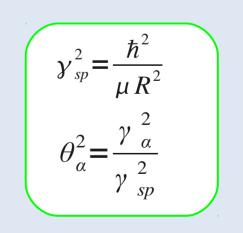


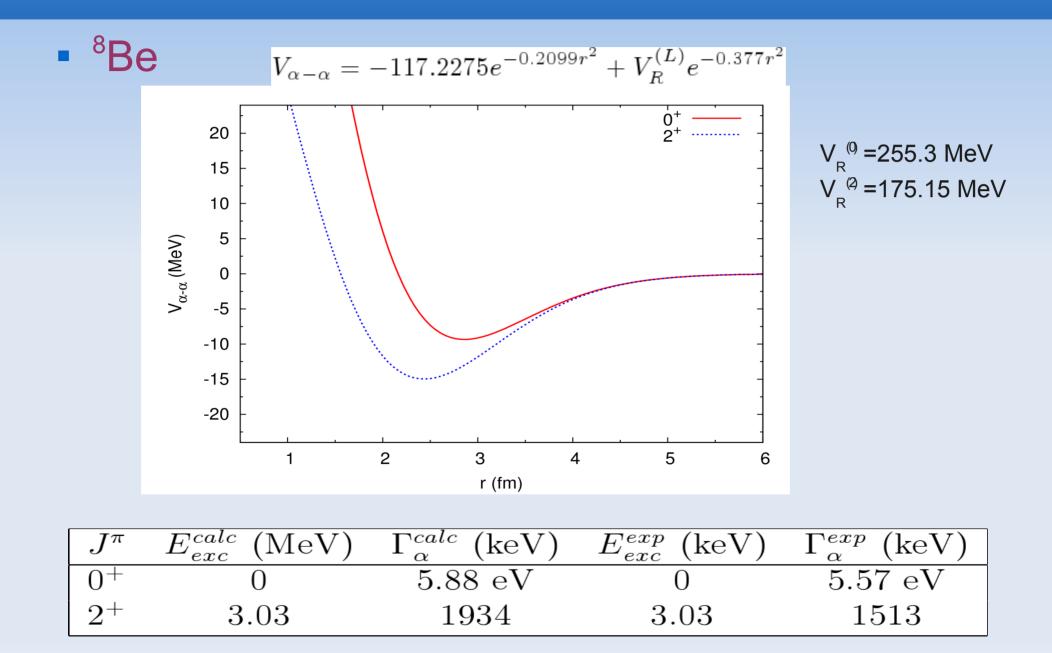
Broad States

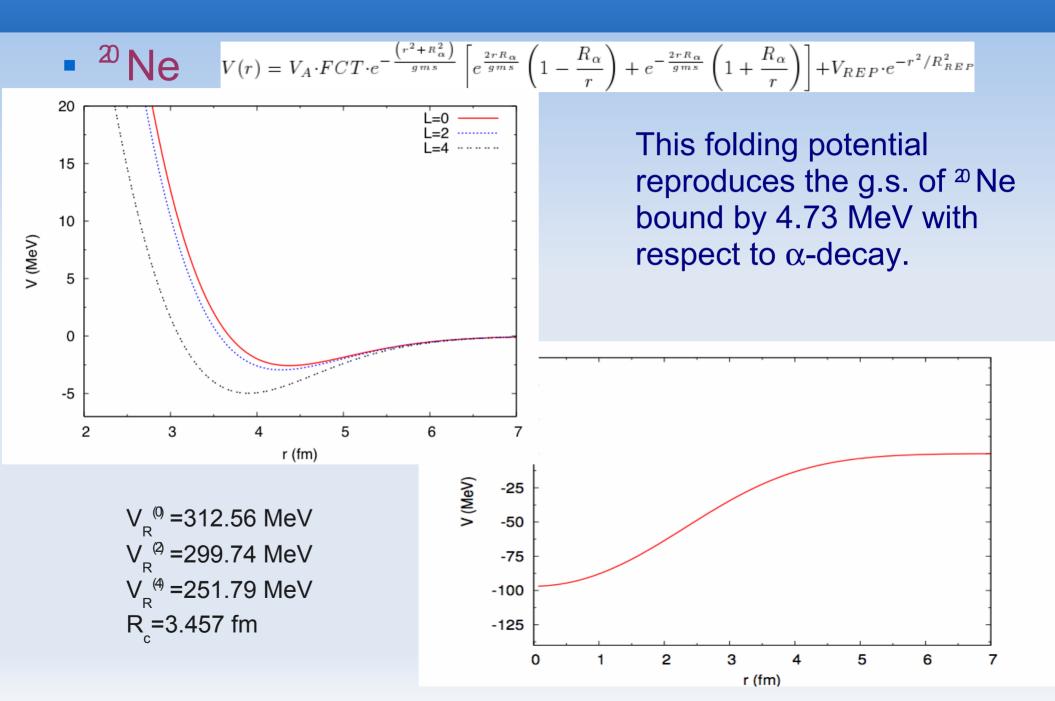
J ^π	E _{ec} (MeV)	r _α (MeV)	Θ^2_{α}
O+	9.9	2.1	2.6
2+	12.17	5.7	2.1

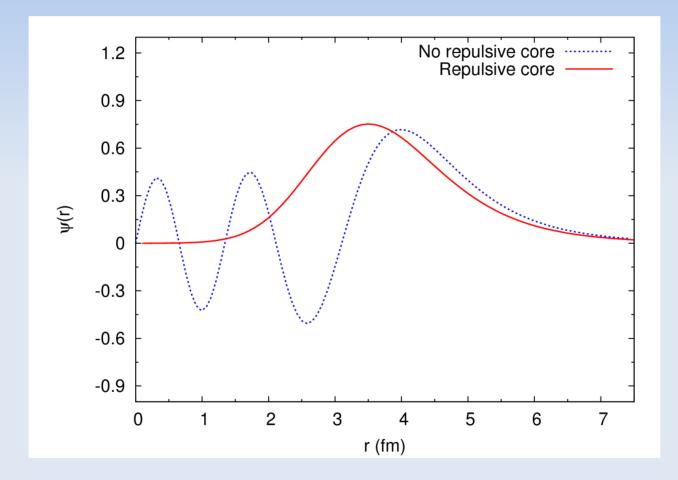
 $\Theta^2_{\alpha} > 1?$

More realistic potential is needed to describe α -core states









20Ne States using Potential model

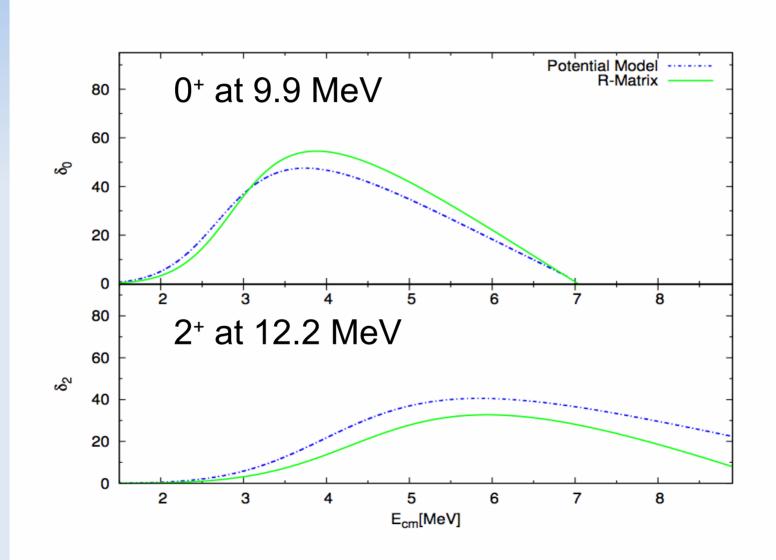
J^{π}	E_{exc}^{calc} (MeV)	$\Gamma^{calc}_{\alpha}({ m MeV})$	E_{exc}^{exp} (MeV)	$\Gamma^{exp}_{\alpha}(\mathrm{MeV})$
0^+_1	0	-	0	-
2^+_1	1.88	-	1.63	-
4_1^+	4.25	-	4.25	_
0^+_2	9.67	1.24	8.7	1.47
2^+_2	11.85	1.26	8.9	1.2

¹⁸ O

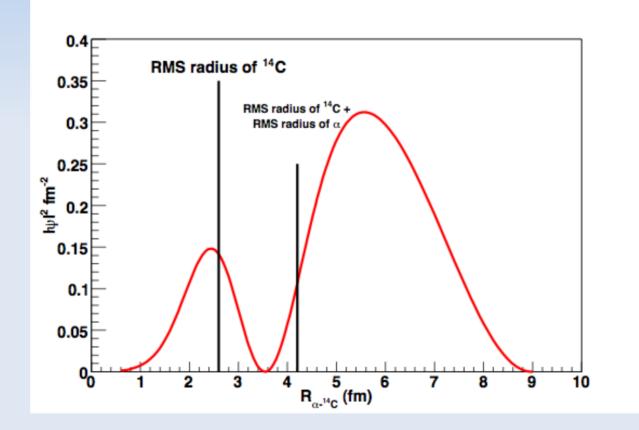
States for 18O using the same potential with Rc=3.379 fm

J^{π}	$E_{exc}^{calc}(\mathrm{MeV})$	$\Gamma^{calc}_{\alpha}({ m MeV})$	$E_{exc}^{exp}(MeV)$	$\Gamma^{exp}_{\alpha}(\mathrm{MeV})$
0^+_1	0	-	0	-
2^+_1	1.982	-	1.982	-
0^{+}_{2}	9.66	1.742	9.9	2.1
2^{+}_{2}	12.08	3.8	12.2	5.7

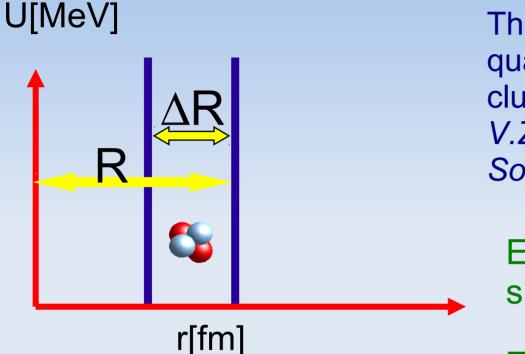
Phase shift calculations



$0^+ \alpha$ halo state in ${}^{18}O$



Infinite surface potential



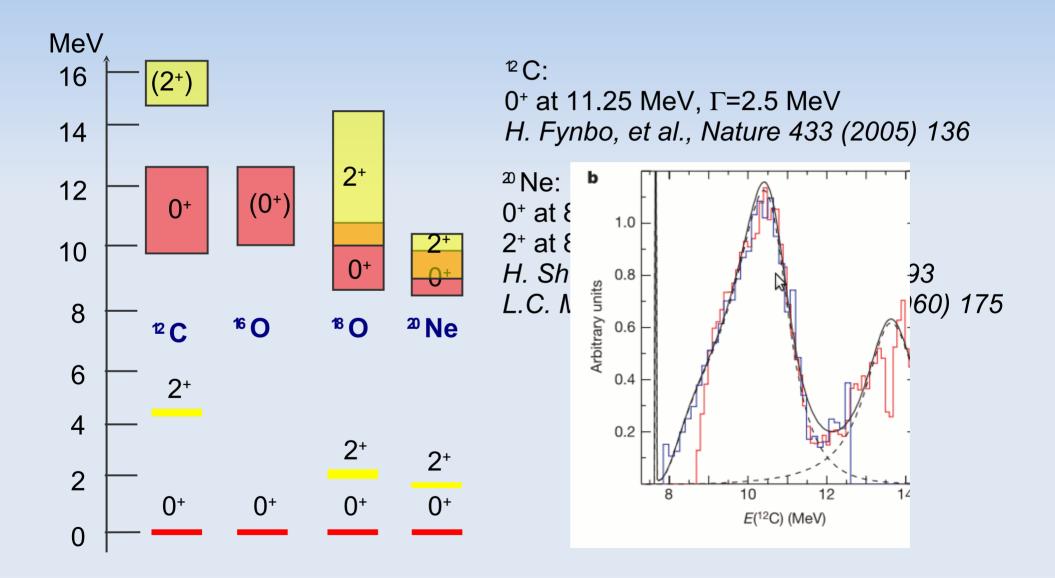
This approach was used for qualitative classification of cluster states in ¹⁶O and ²⁰Ne by *V.Z. Goldberg, et al, Sov. J. Nucl. Phys., 19, 253 (1974)*

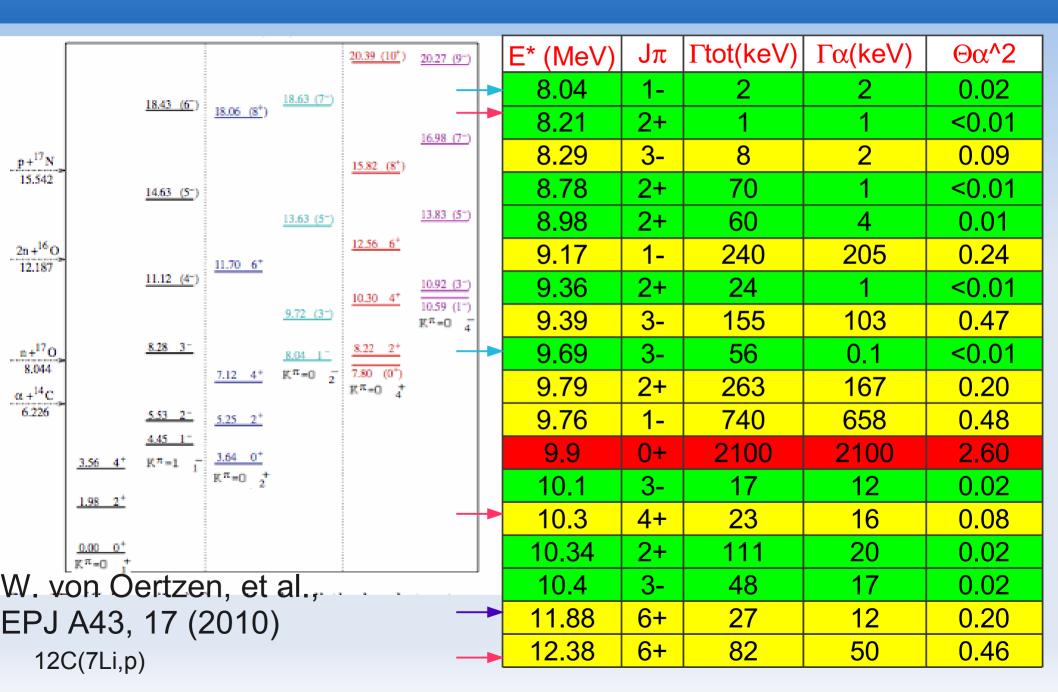
Excitation energyof n=2 0⁺ state should be ~11 MeV

There should be n=2 2⁺ state

$$E = \frac{n^2 \pi^2 \hbar^2}{2\mu (\Delta R)^2} + \frac{\hbar^2 \ell (\ell+1)}{2\mu R^2}$$

"α-halo" states in light nuclei







- Broad resonances that correspond to pure α-core structures ("α-halo" states) have been known to exist in ¹² C, ²⁰ Ne and now observed in ¹⁸ O.
- Several α-cluster states were found in ¹⁸O including in the energy range of 2-7 MeV.
- It is speculated that "α-halo" states are common phenomena for light nuclei.
- Potential model was able to predict these broad states.