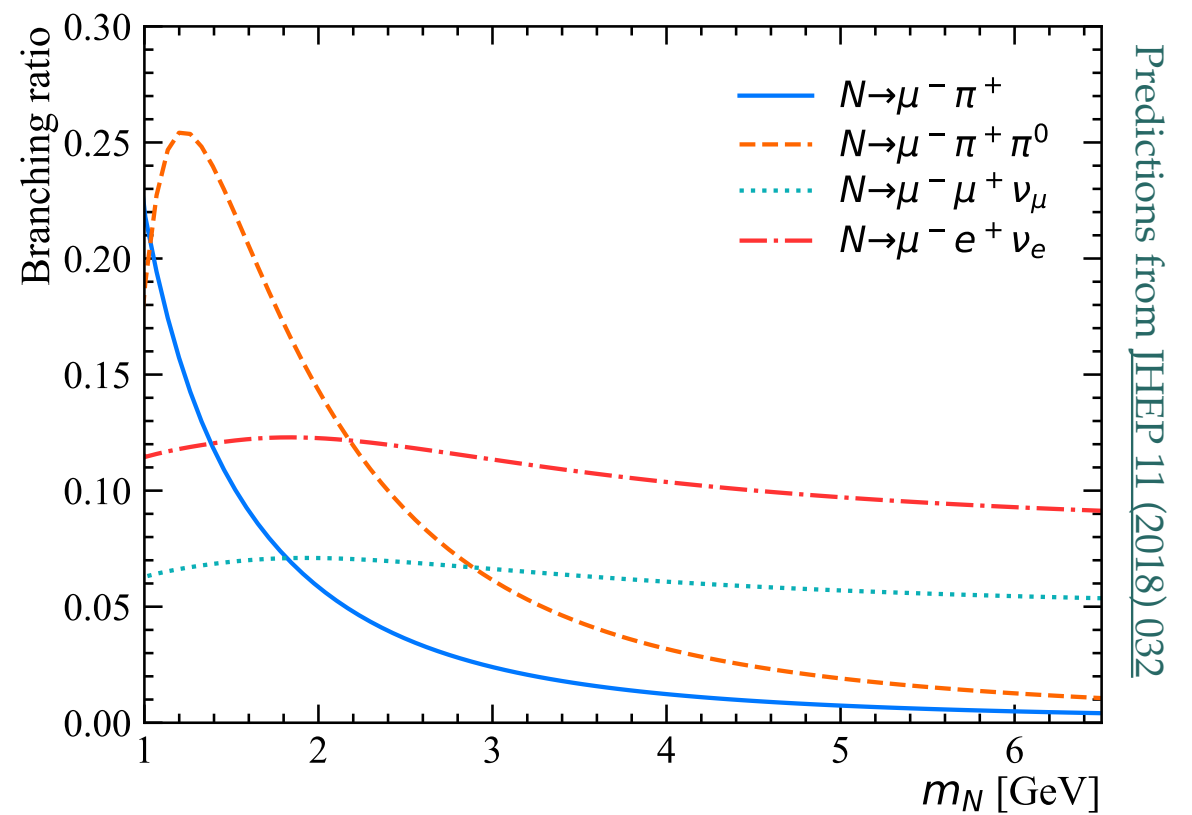
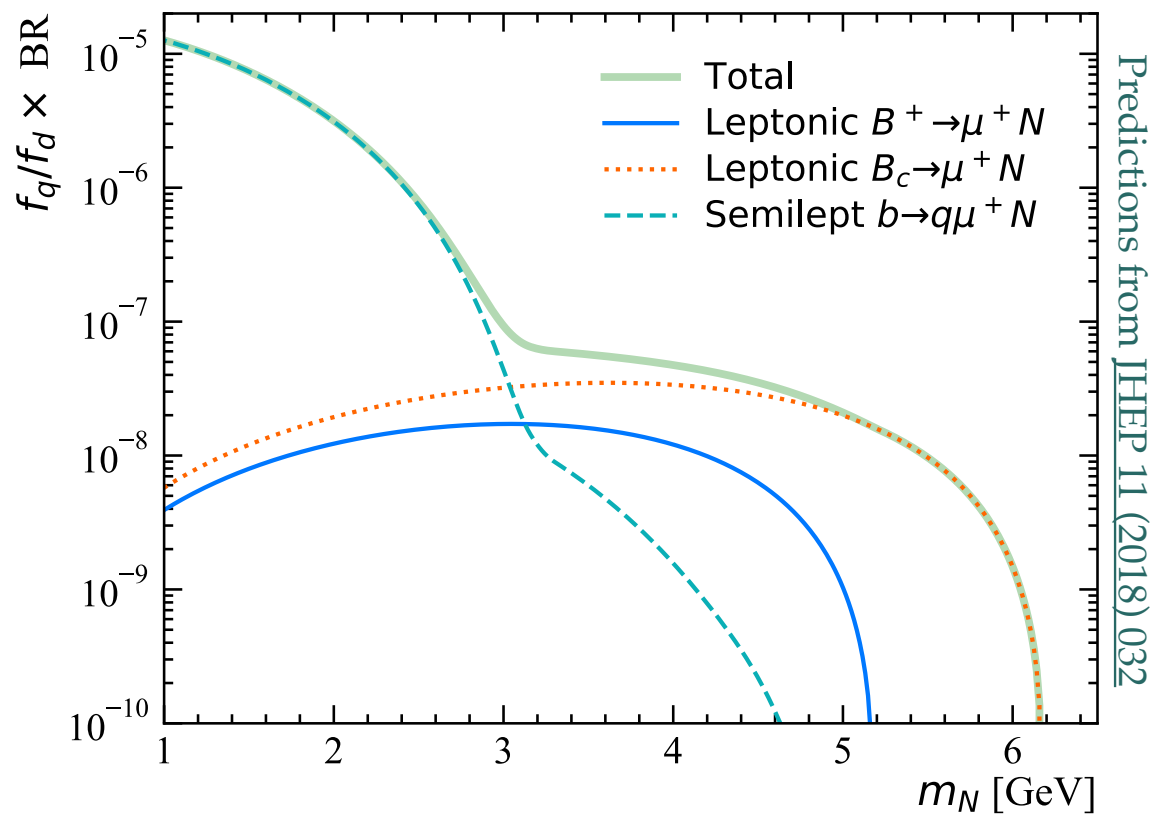
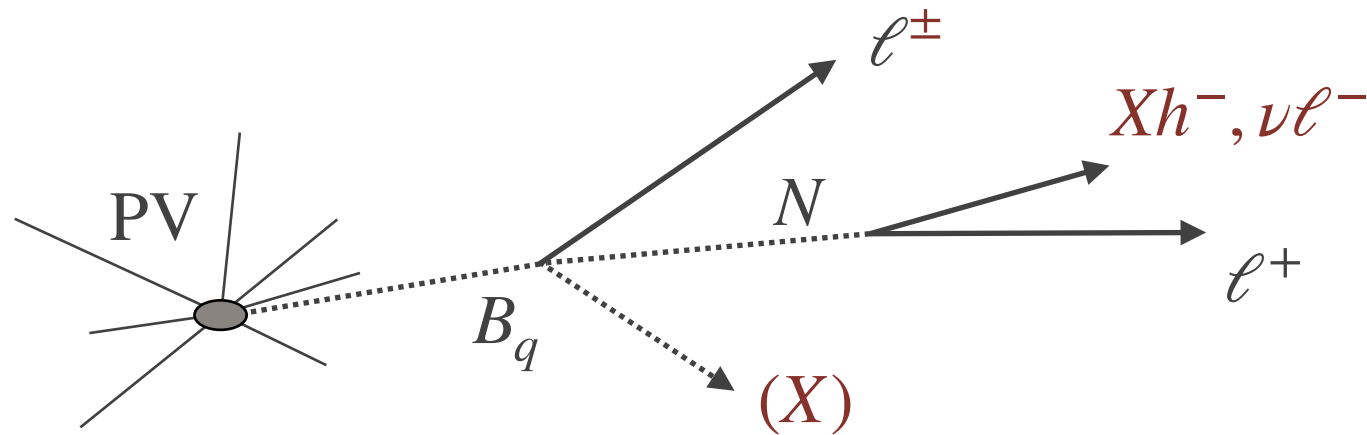

Dark sector → LHCb

- **Dark sector** = light masses, small couplings, displaced vertices, invisible signatures
- Why searching at LHCb?
 - Huge production rates in forward LHC region
 - Precise low- p_T trigger for soft signatures
 - Real-time identification of displaced vertices (at first trigger level in the upgrade!)
 - Huge boost in forward region → increased flight distance

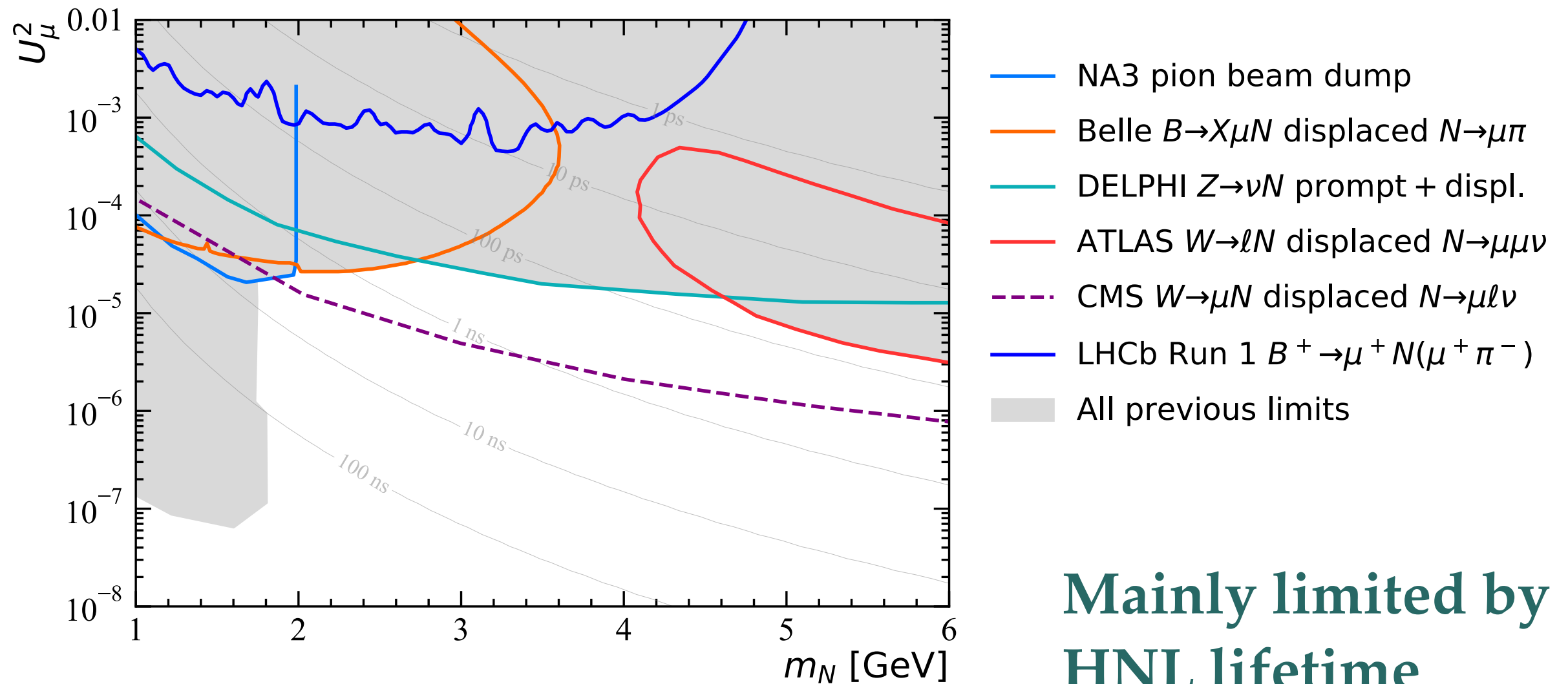
- **Heavy** neutral leptons (**GeV**)
- **Light** scalars (**GeV**)
- **Visible dark** photons

Heavy neutral leptons

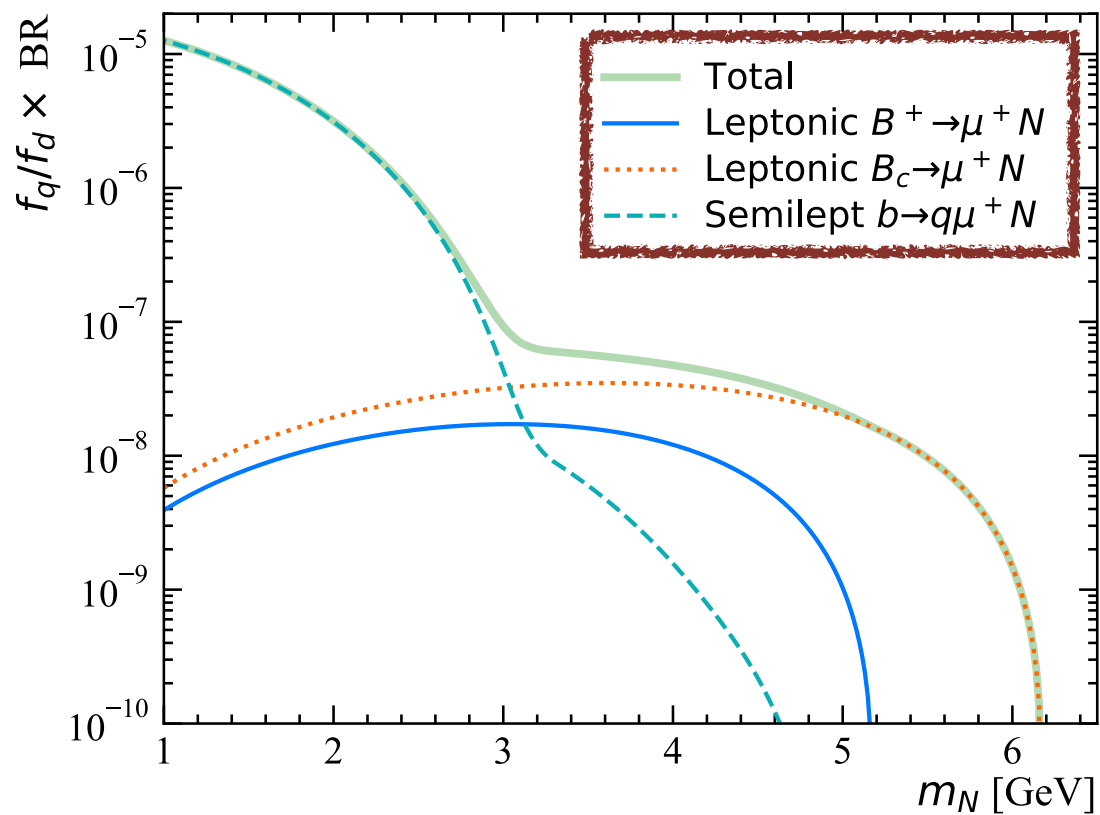
HNL in B decays



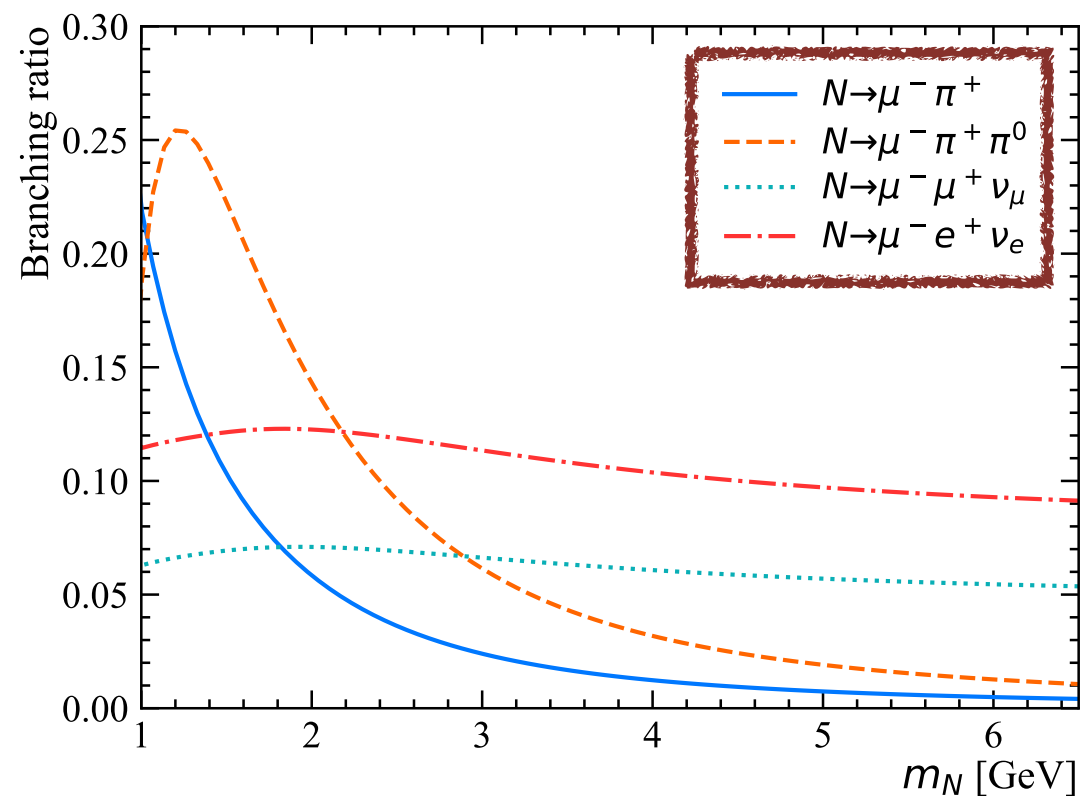
HNL searches in GeV range



B → HNL in future LHCb



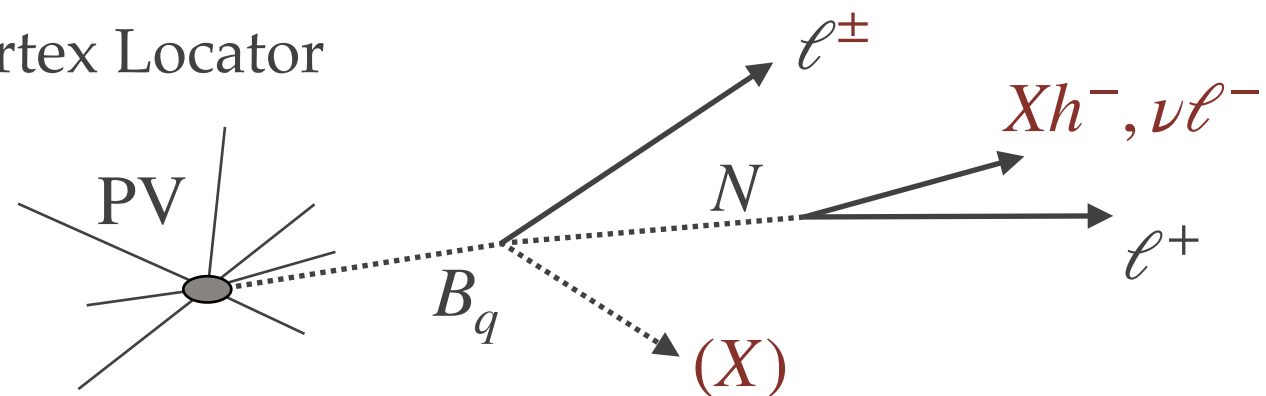
Predictions from [JHEP 11 \(2018\) 032](#)



Predictions from [JHEP 11 \(2018\) 032](#)

Future LHCb strategy:

- Include $B_c \rightarrow \ell N$ and $B_q \rightarrow X \ell N$
- Include partially reconstructed N decays
- Include N decays downstream of the Vertex Locator (10x longer decay time)
- Search in all lepton flavours (also τ ?)
- Search both LNC and LNV decays

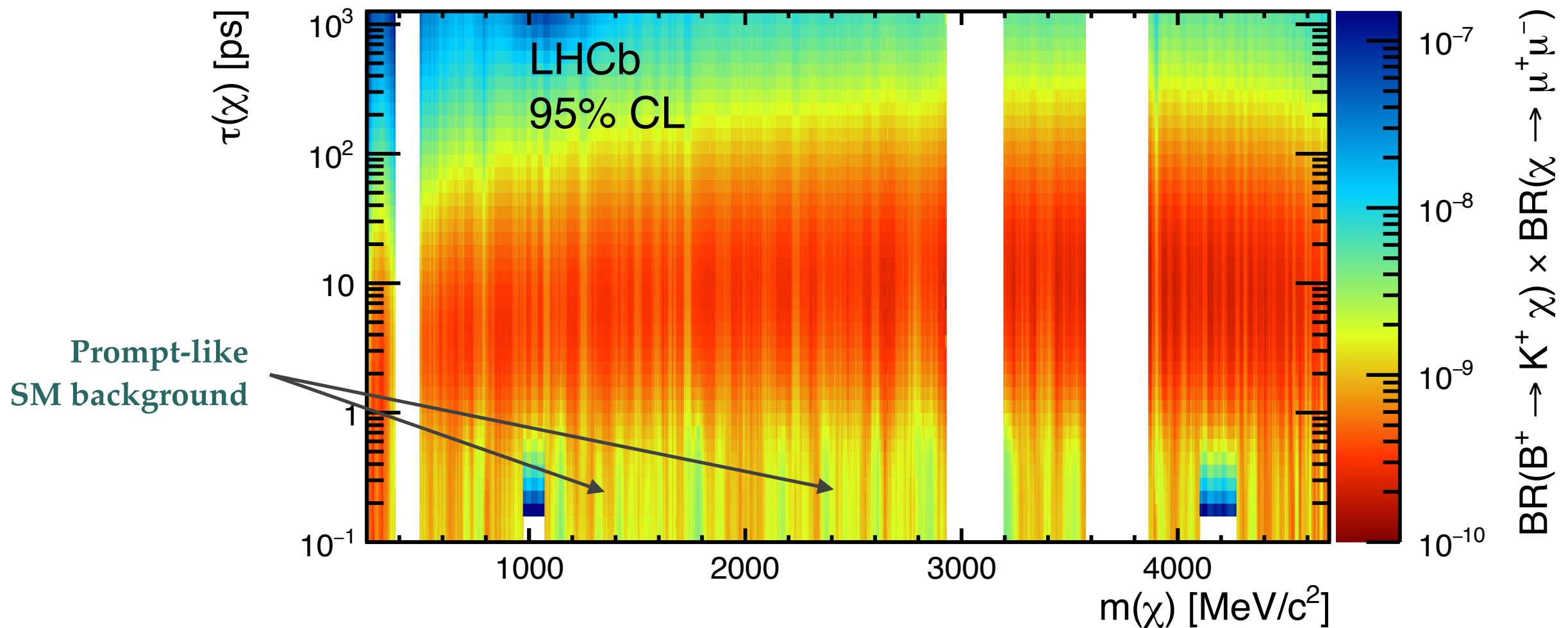


Light scalars

Light scalar from $b \rightarrow s$

[Phys Rev Lett 115 161802 \(2015\)](#)

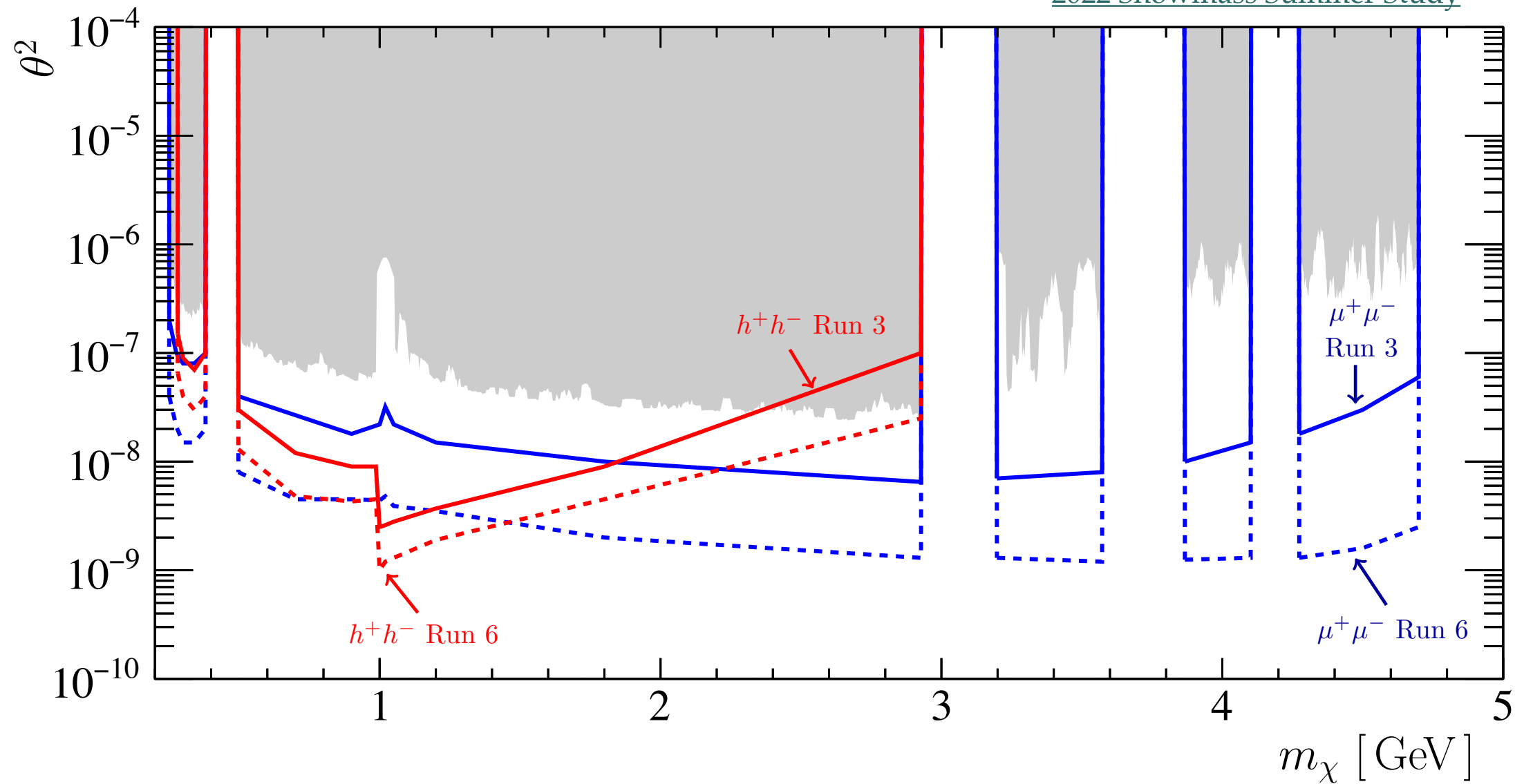
[Phys Rev D 95, 071101\(R\) \(2017\)](#)



- Use peaks in reconstructed m_B and $m(\mu\mu)$ to reduce background
- Upper limits down to 10^{-10} on $\text{BR}(B^+ \rightarrow K^+ \chi) \times \text{BR}(\chi \rightarrow \mu\mu)$

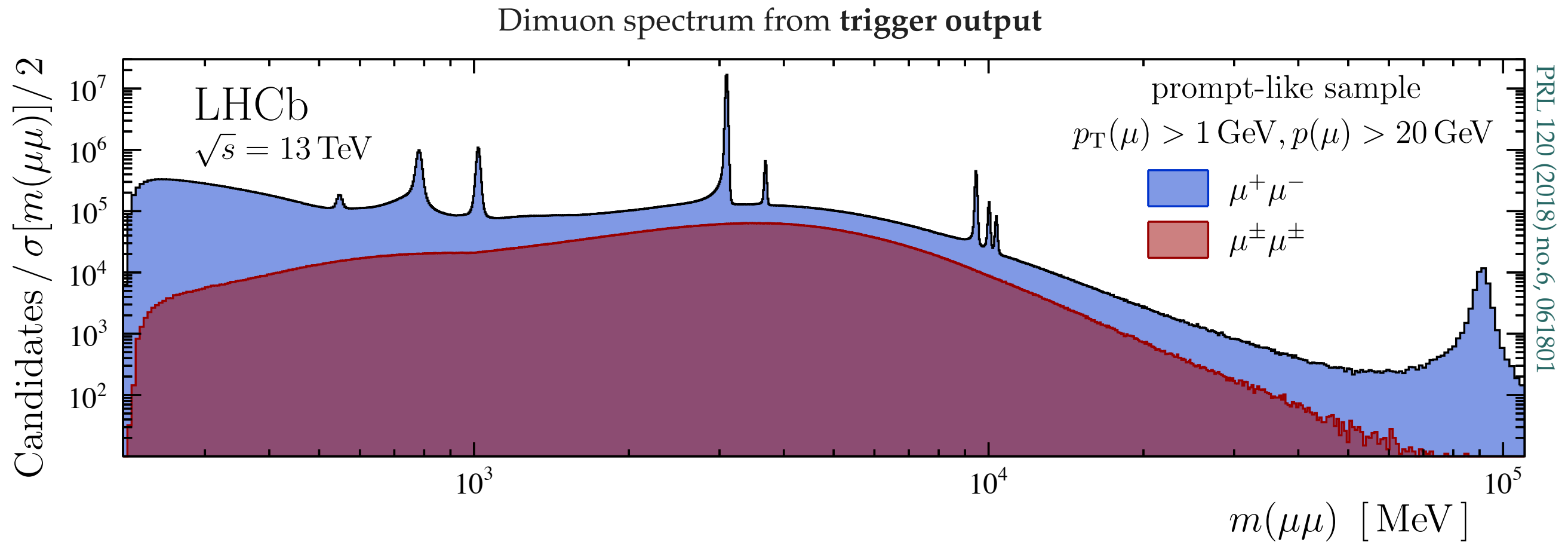
Light scalar from $b \rightarrow s$

2022 Snowmass Summer Study



Dark photons

Visible $A' \rightarrow$ LHCb



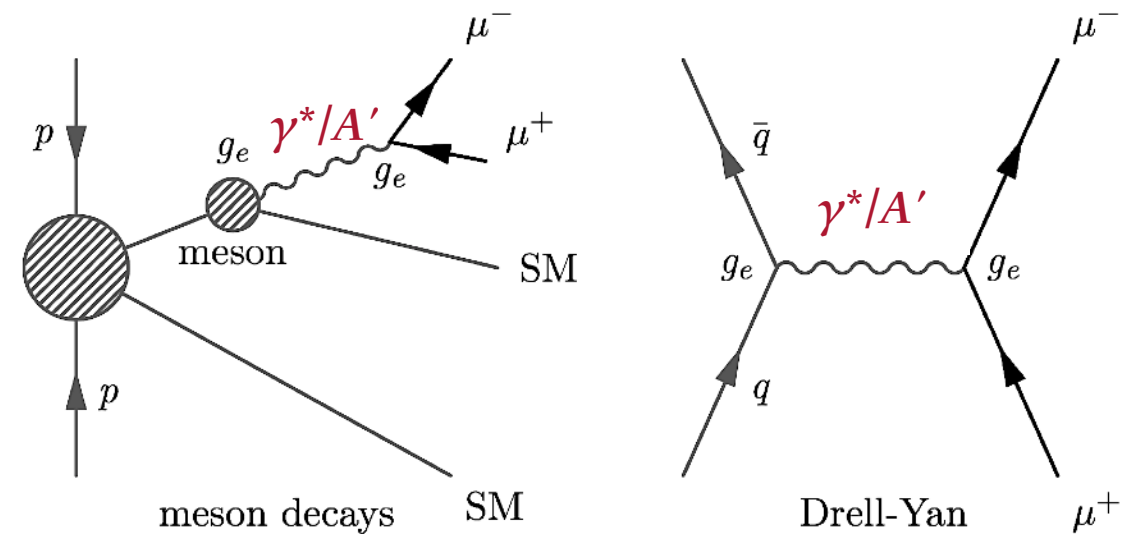
- Started inclusive $A' \rightarrow \mu^+\mu^-$ searches in Run 2
 - Leveraging the online-analysis capabilities introduced in 2015
 - no pre-scale down to threshold $2m_\mu$
 - Great prospects for upcoming upgrade

Visible $A' \rightarrow$ LHCb

PRL 120 (2018) no.6, 061801
and PRL 124 (2020) 041801

Analysis strategy:

- inherits production mode of off-shell photon
 - Can normalise to $\gamma^* \rightarrow \mu\mu$ continuum
 - just need to separate non γ^* background
 - No need for efficiencies from simulation (only if displaced vertex)



$$n_{\text{ex}}^{A'}[m(A'), \epsilon^2] = \epsilon^2 \left[\frac{n_{\text{ob}}^{\gamma^*}[m(A')]}{2\Delta m} \right] \mathcal{F}[m(A')] \epsilon_{\gamma^*}^{A'}[m(A'), \tau(A')]$$

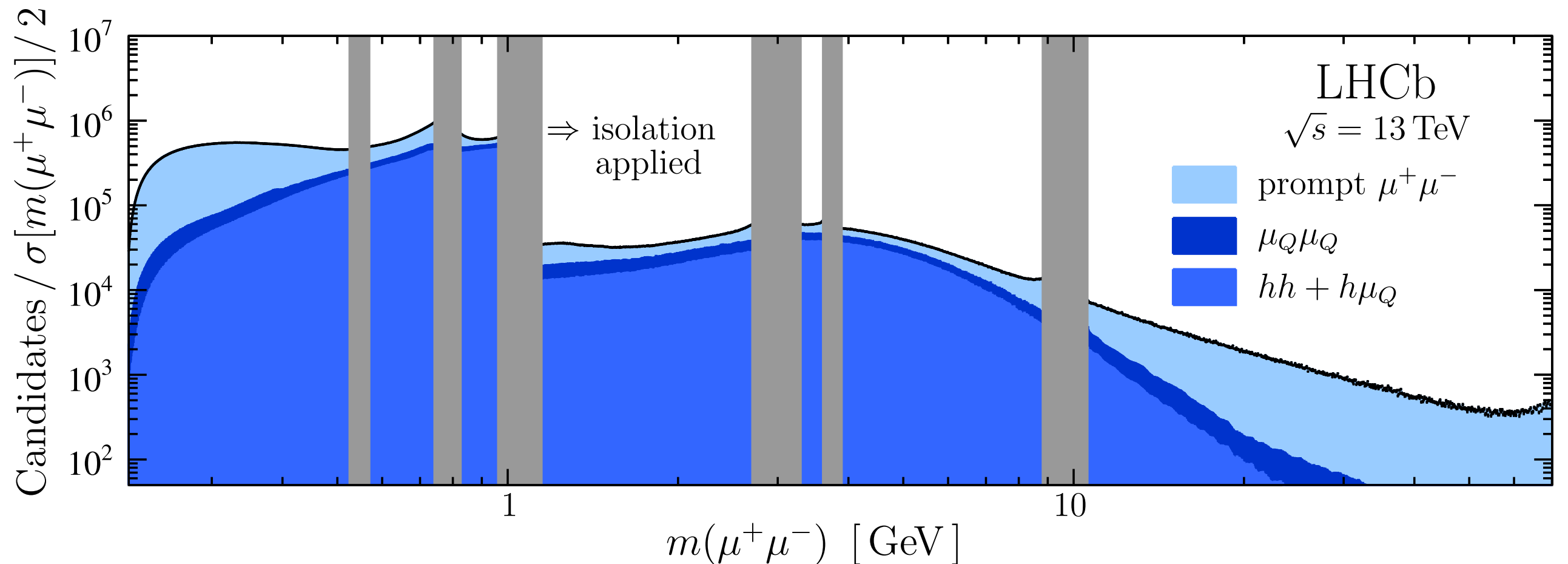
off-shell photon

phase-space

A'/γ^* eff ratio,
 $\epsilon=1$ for prompt

Prompt-like $A' \rightarrow \mu^+ \mu^-$

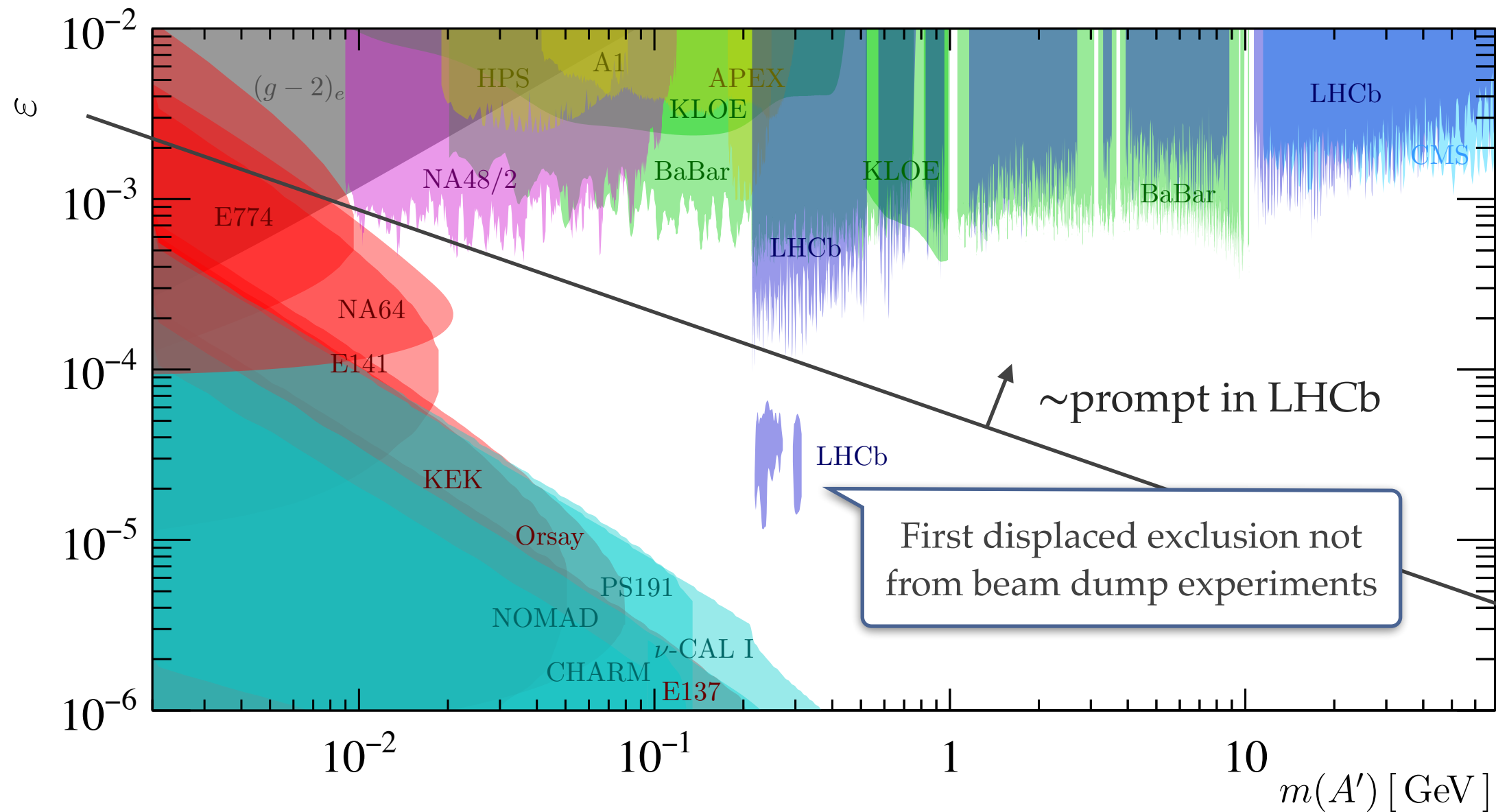
PRL 120 (2018) no.6, 061801
and PRL 124 (2020) 041801



- $O(10^7 - 10^{11}) \times \epsilon^2$ dark photons expected
 - Peak hunt on top of large background
 - Remove regions with QCD resonances
- Also displaced search at low mass (low background)

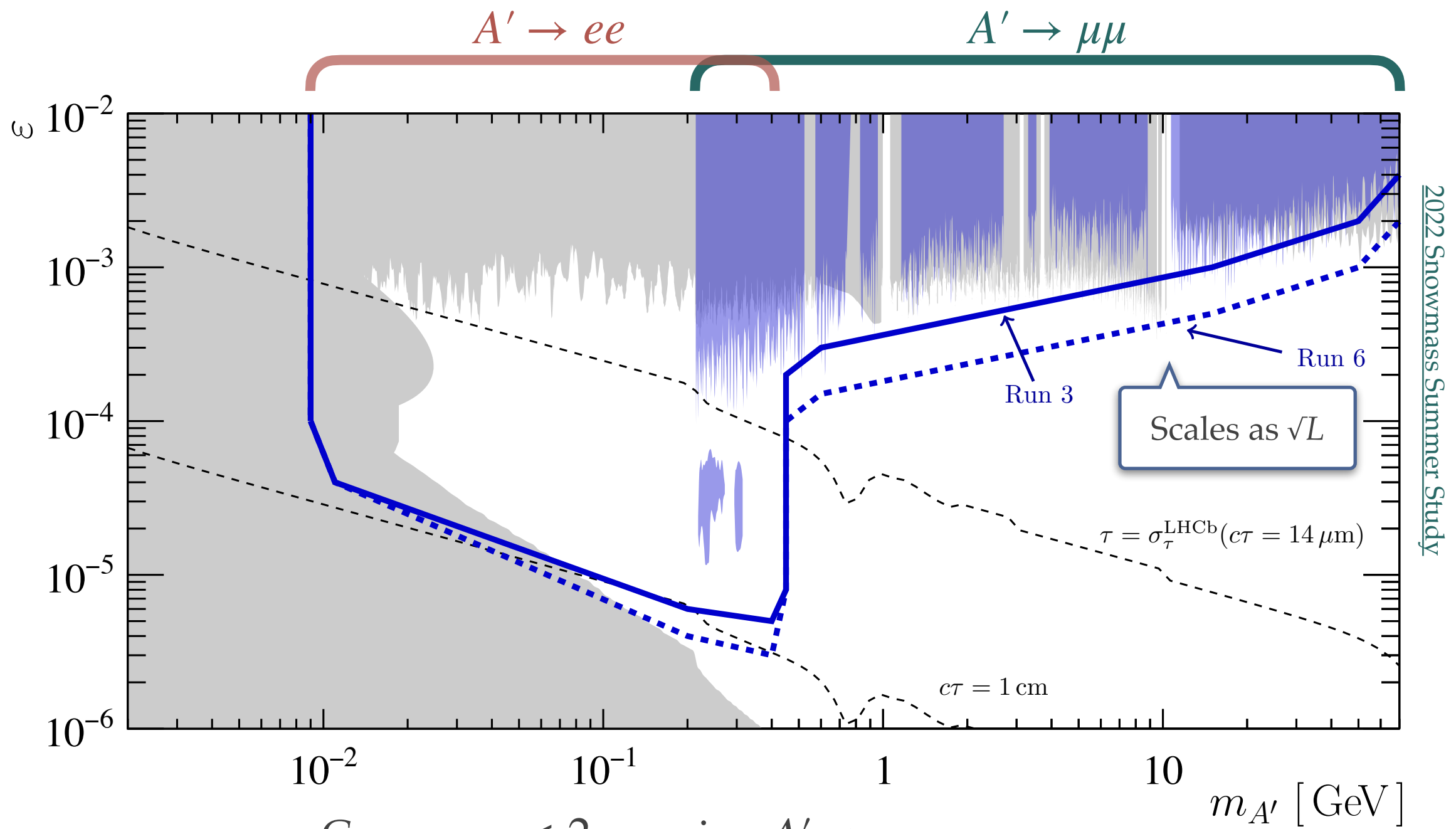
Visible A' limits

PRL 124 (2020) 041801



- Easy to recast to other vector models [JHEP 06 \(2018\) 004](#)

Visible A' prospects



Cover $m_{A'} < 2m_\mu$ using $A' \rightarrow ee$:

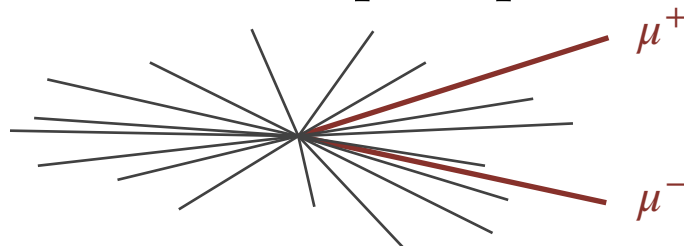
- $D^* \rightarrow D^0 A'$ (lower bkg) [PRD 92 no.11, 115017 \(2015\)](#)
- $\pi^0/\eta \rightarrow A'\gamma$ (higher rate) [2022 Snowmass Summer Study](#)

Inclusive $X \rightarrow \mu^+ \mu^-$ search

LHCb [JHEP10\(2020\)156](#)

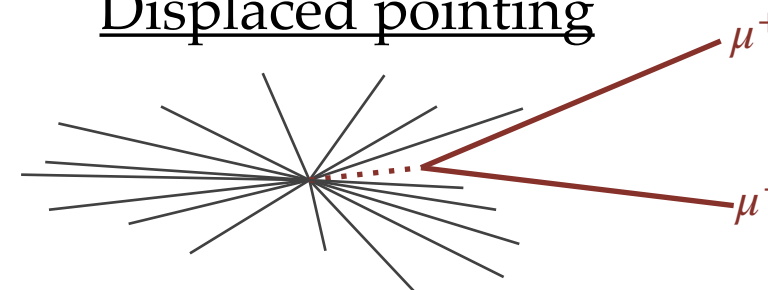
- Probe additional dark sectors in $\mu\mu$
 - Drop assumption of kinetic mixing with γ^*
 - Minimise assumptions on production mechanism (tight fiducial regions and results in kinematic bins)

Inclusive prompt

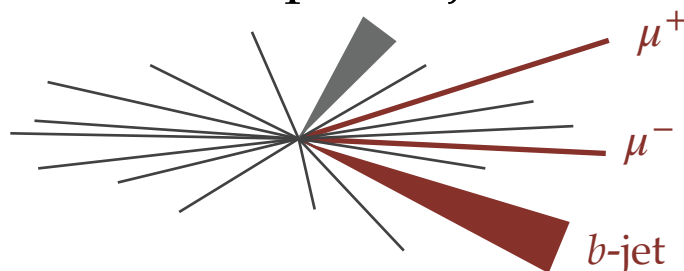


- No isolation requirements
- Non-zero width considered

Displaced pointing

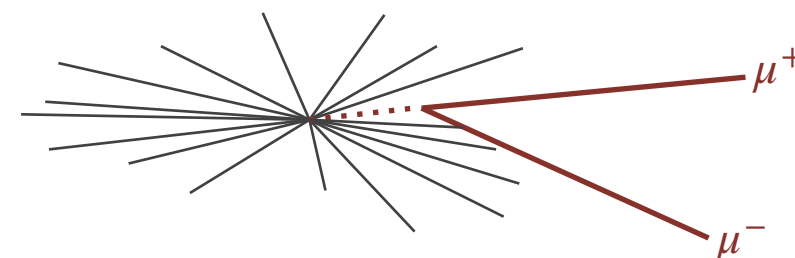


Prompt + b -jet



- Non-zero width considered

Displaced non-pointing



Inclusive $X \rightarrow \mu^+ \mu^-$ search

LHCb [JHEP10\(2020\)156](#)

