
Radiation effects for the next generation of synchrotron radiation facilities

PhD Defence

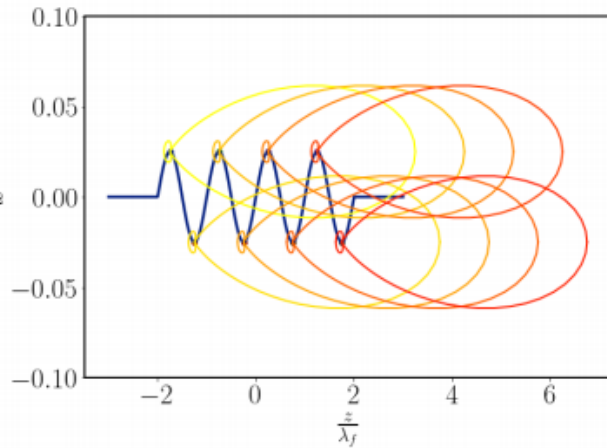
Marcel Ruijter, Supervisors: Luca Serafini
Matricola 1842024, Vittoria Petrillo
Cycle XXXIV



Introduction

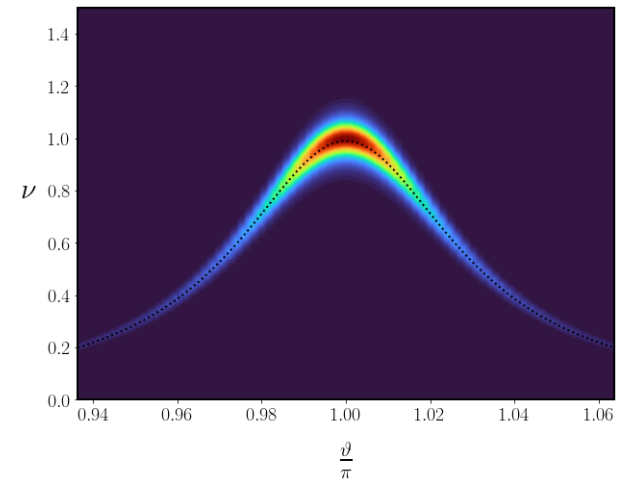
Generation of narrow bandwidth Synchrotron radiation

- Undulator / FEL
- Thomson Scattering

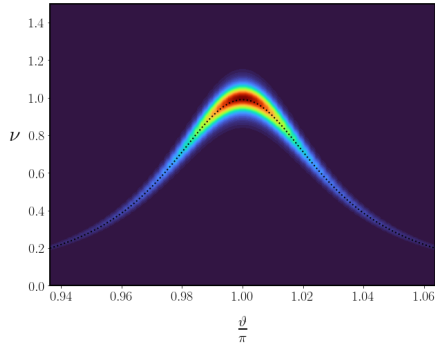


$$\langle P_{power} \rangle \propto \gamma^2 U_{field}$$

$$\langle \lambda \rangle \propto \frac{\lambda_f}{\gamma^2}$$



Introduction



$$\langle P_{power} \rangle \propto \gamma^2 U_{field}$$

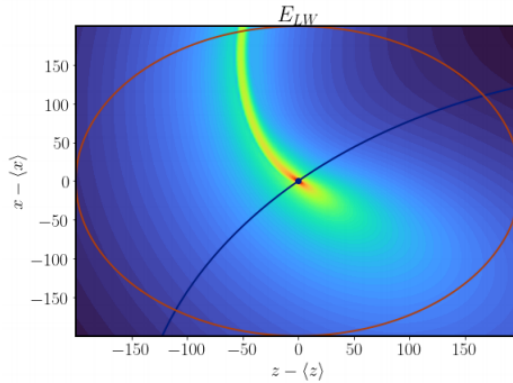
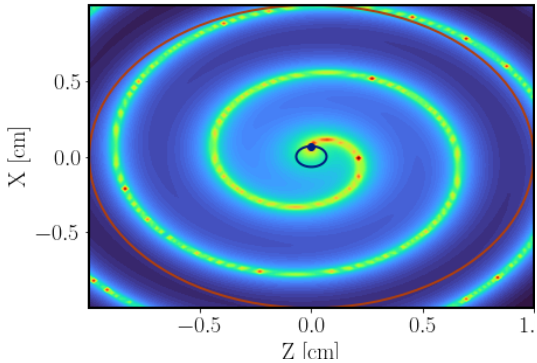
$$\langle \lambda \rangle \propto \frac{\lambda_f}{\gamma^2}$$

Electron Energy for FEL Thomson Scattering

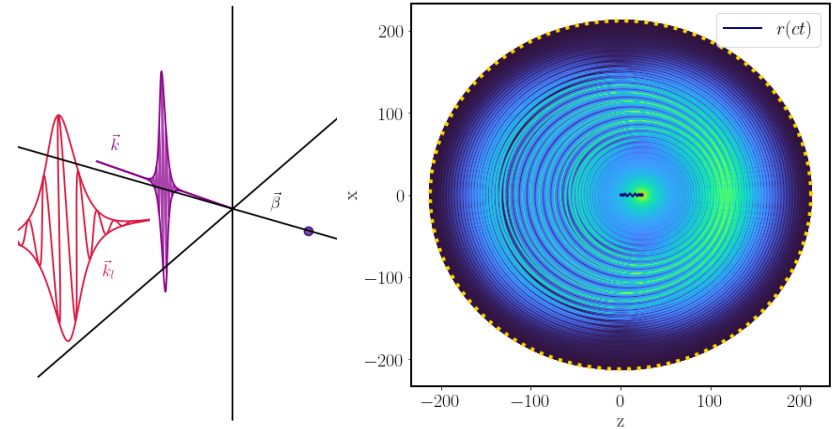
- | | | |
|--|------------------|------------------|
| ➤ eV – > 1 μm : Macroscale Material properties
Biological &
Chemical processes | ~50 | - |
| ➤ keV – <nm : Medical treatment, invivo
imaging, nuclear &
atomic research | ~10 ³ | ~ 50 |
| ➤ MeV - <pm : Astro-, Quantum &
Hadron physics | ~10 ⁵ | ~10 ³ |

Thesis Overview

I. Self-fields



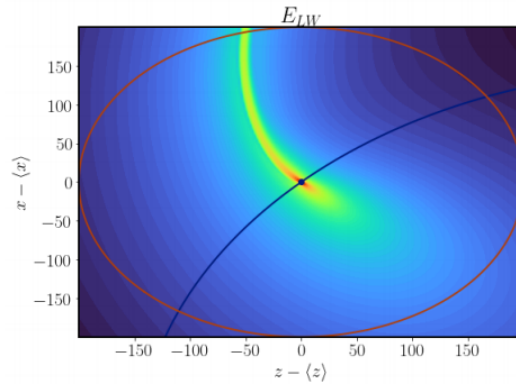
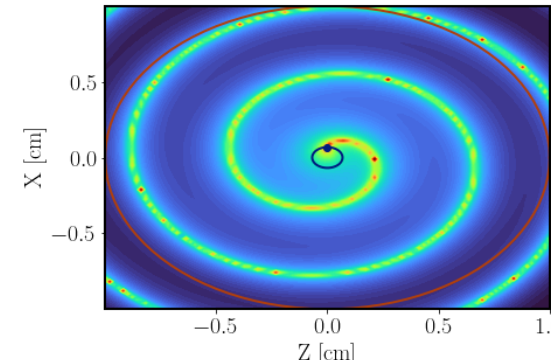
II. Thomson scattering



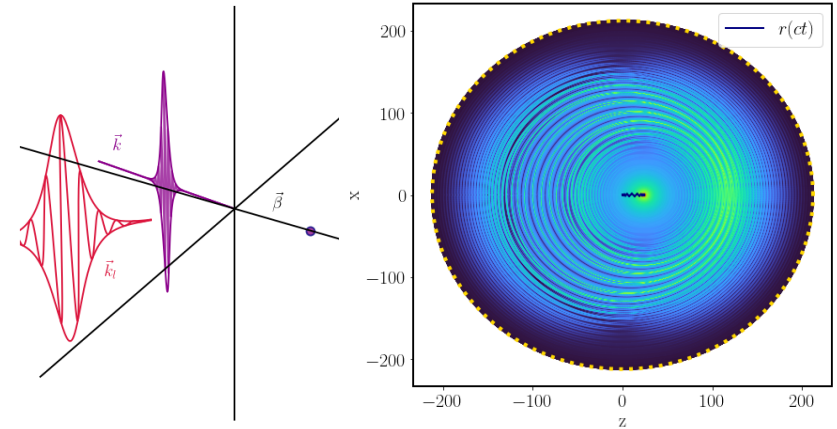
- Retarded time
- Models
 - 1D
 - 2D
 - 3D

- Degenerate Cavity
- Energy Compensation
- Carrier Envelope Phase

I. Self-fields



II. Thomson scattering



Lienard Wiechert Potentials

$$\vec{E}_{LW} = \vec{E}_{Coulomb} + \vec{E}_{Radiation}$$

C

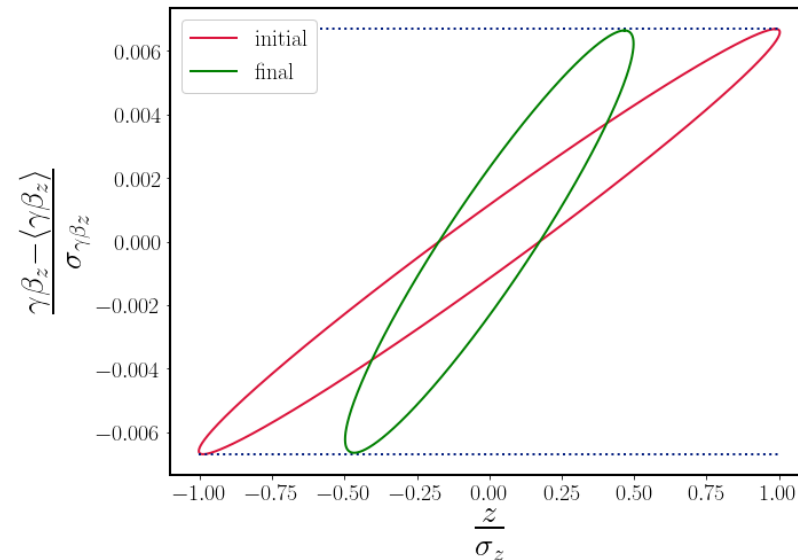
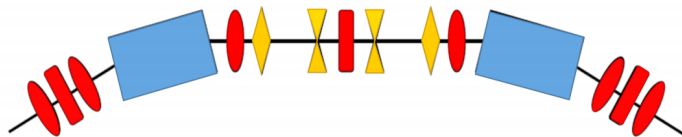
FEL

- Current km size – LINAC – to reach energy
- High brightness for radiation-bunch interaction
- High coherency \Leftrightarrow power $\propto N_e^{4/3}$

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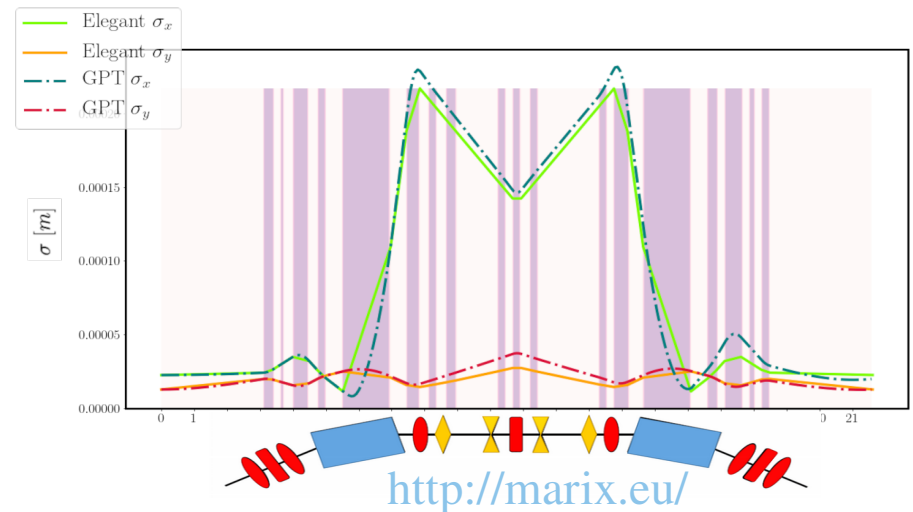
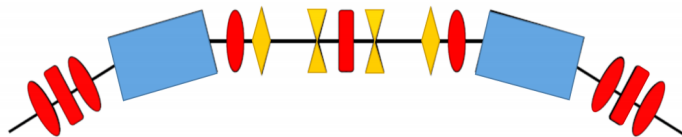
⇒ Compression
DBA



FEL

- Current km size – LINAC – to reach energy
- High brightness for radiation-bunch interaction
- High coherency \Leftrightarrow power $\propto N_e^{4/3}$

\Rightarrow Compression
DBA



I. Self fields

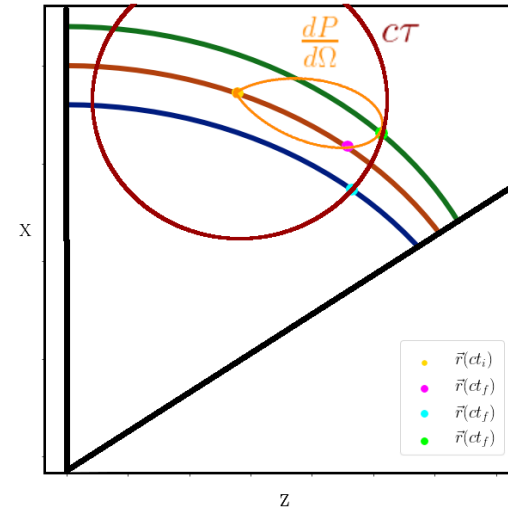
Retarded time: $c\tau$
1D Model

Derbenev 1995,
Saldin 1997

Focus on wakefield (total electric field)

Criteria of applicability

$$\frac{R}{\gamma^3} \ll \sigma_z \qquad \frac{\sigma_r}{\sigma_z} \ll \left(\frac{R}{\sigma_z}\right)^{\frac{1}{3}}$$



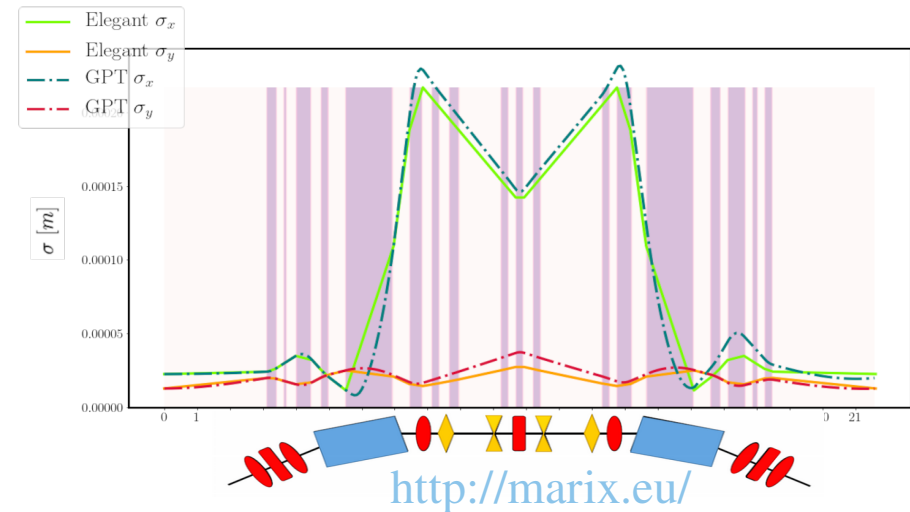
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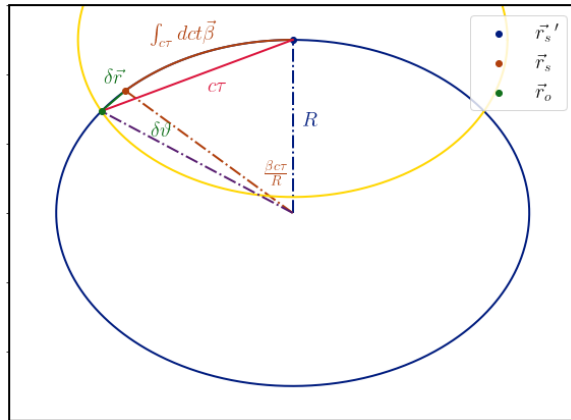
- Does it hold for large differences in transverse size bunch?
- What are the effects as the bunch focusses?



Revisited

$$R = \frac{\gamma \beta_{\perp} m c^2}{e B_0}$$

$$\frac{c\tau}{2R} - \sin\left(\frac{\beta c\tau}{2R} + \frac{\delta\vartheta}{2}\right) = 0$$

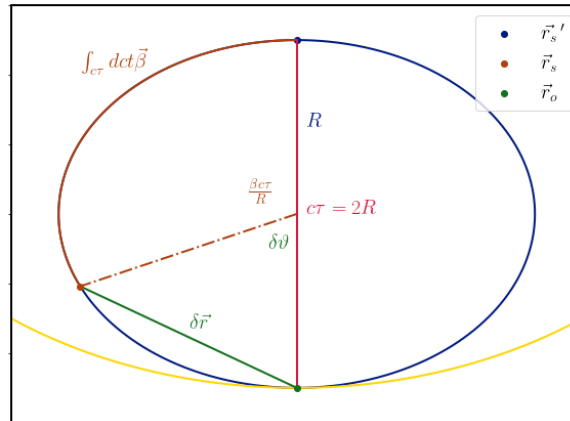
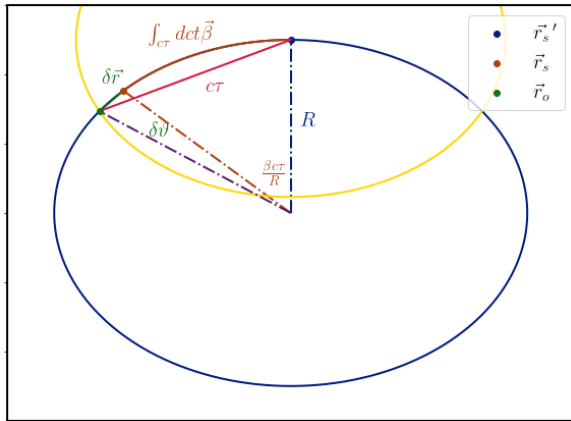


Revisited

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$$\frac{c\tau}{2R} - \sin\left(\frac{\beta c\tau}{2R} + \frac{\delta\vartheta}{2}\right) = 0$$

Argument sine $\leq \pi/2$
 \Rightarrow series up to 3rd
 order *always* suffice

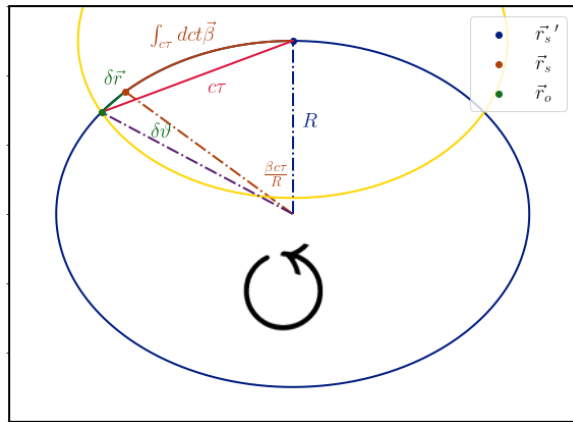


$$\delta\vartheta = \pi - 2\beta$$

Revisited

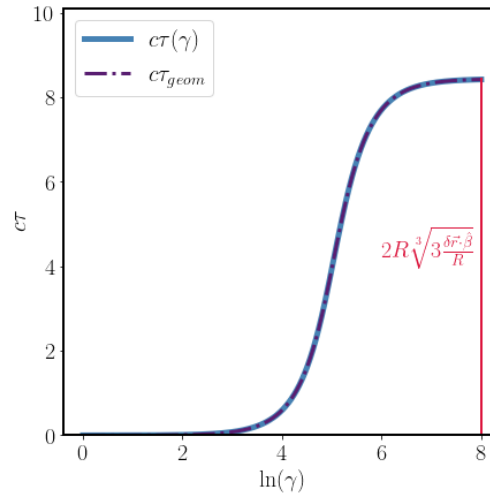
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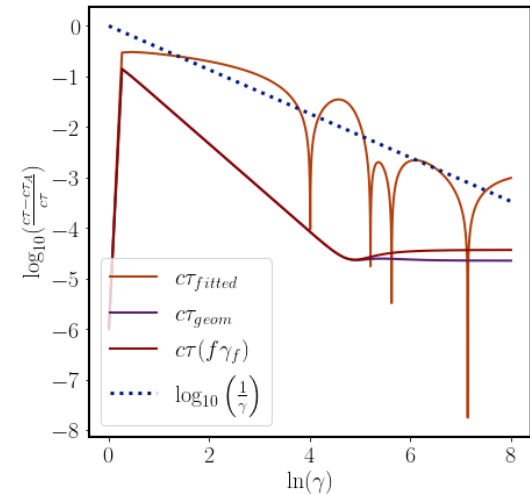


Behaviour of $c\tau$

Three approximated relations derived



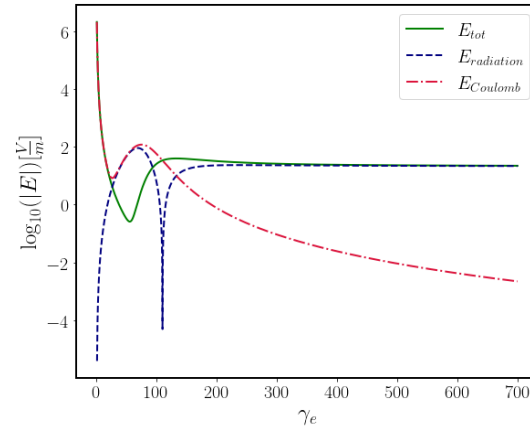
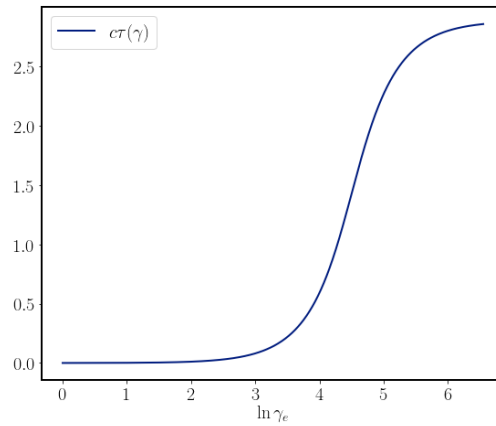
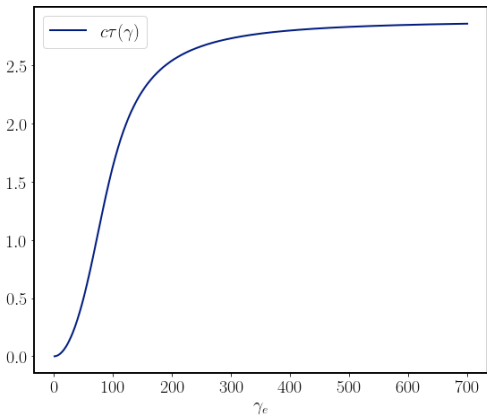
Relative error



I. Self fields

1D Model

$R = 1.00 \cdot 10^0 [\text{m}]$, $\delta S = 1.00 \cdot 10^{-6} [\text{m}]$



$$\vec{E}_{LW} = \vec{E}_{Coulomb} + \vec{E}_{Radiation}$$

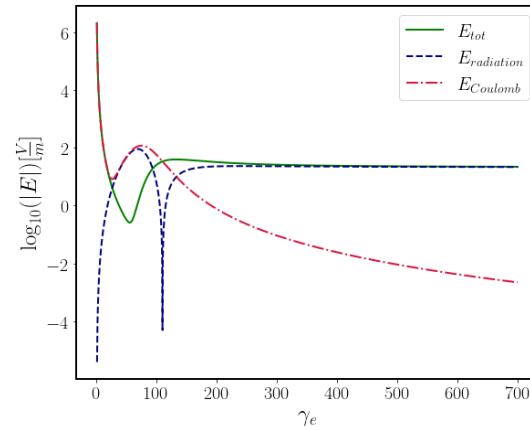
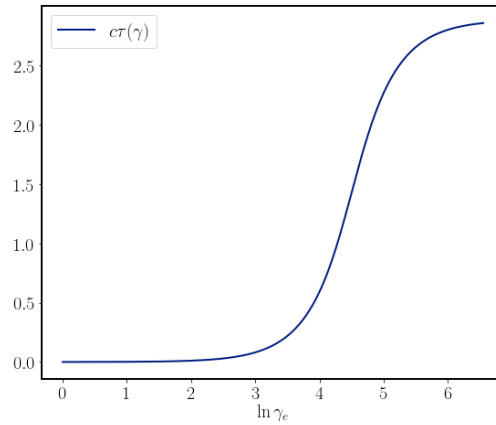
$$\vec{E}_{Coulomb} \propto \left(\frac{1}{\gamma c\tau} \right)^2$$

$$\vec{E}_{Radiation} \propto \left(\frac{1}{c\tau} \right)$$

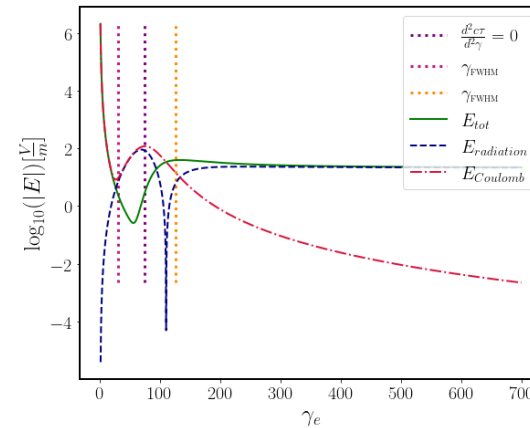
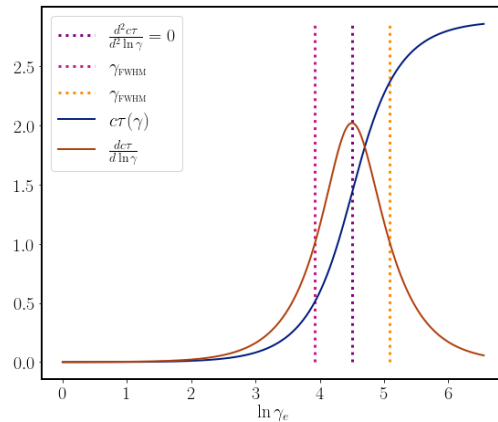
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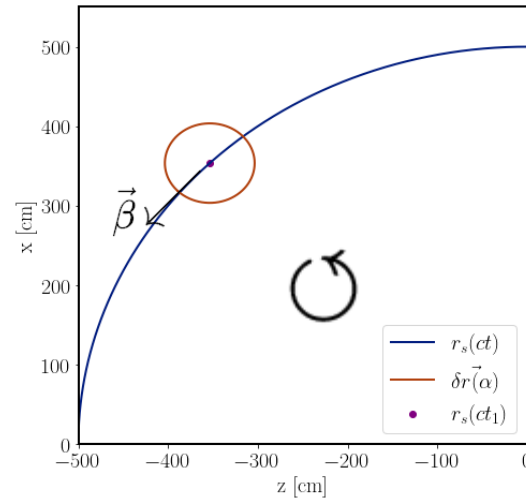
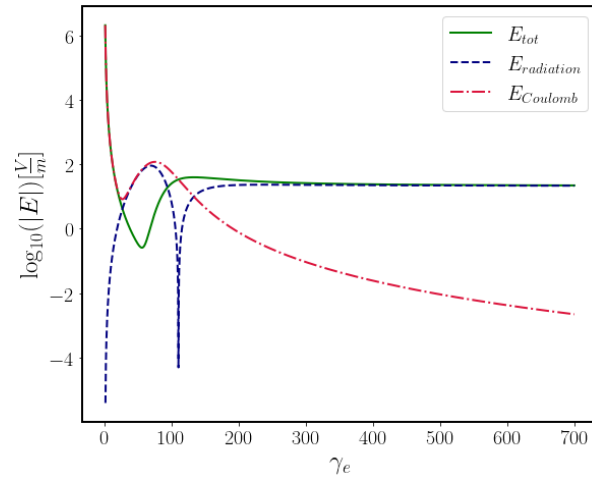


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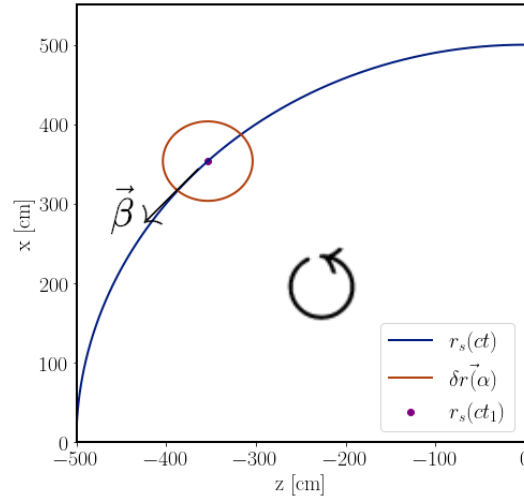
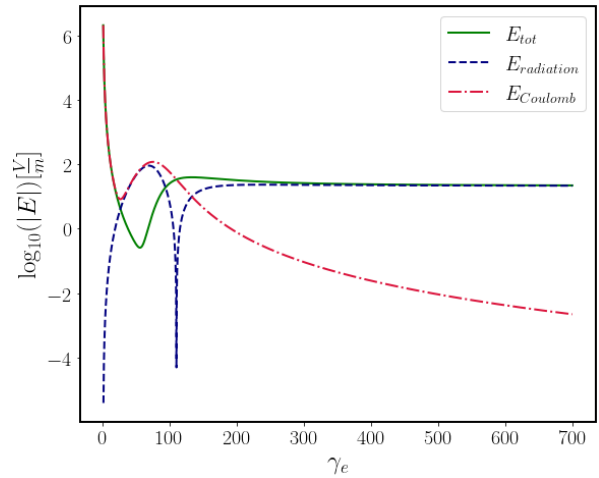
I. Self fields

2D Model



I. Self fields

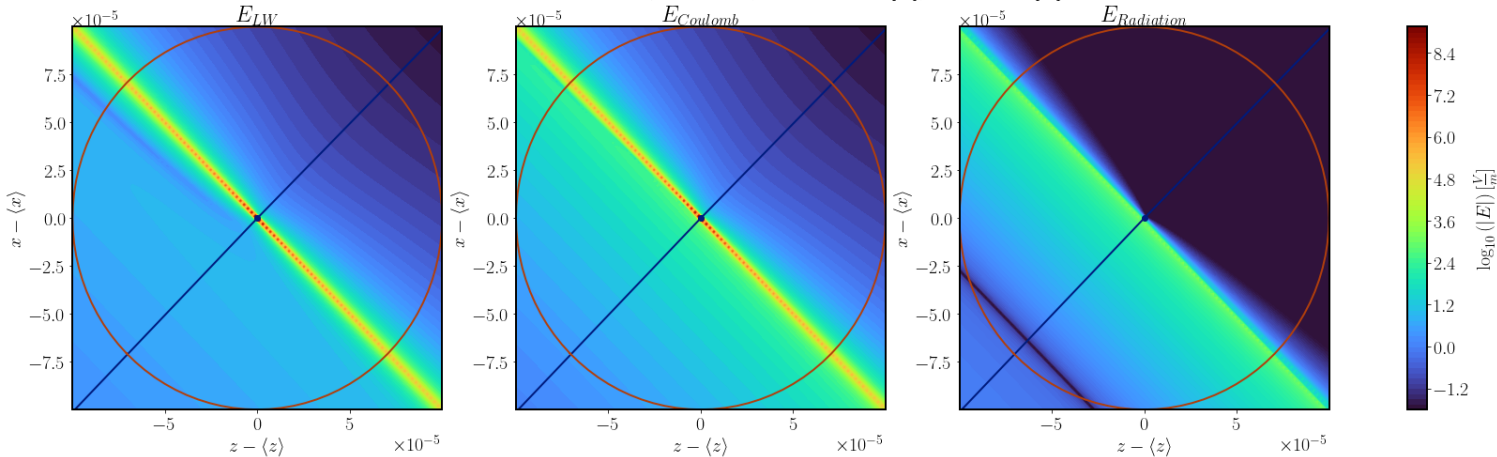
2D Model



$$\vec{E}_{Coulomb} \propto \left(\frac{1}{\gamma c\tau} \right)^2$$

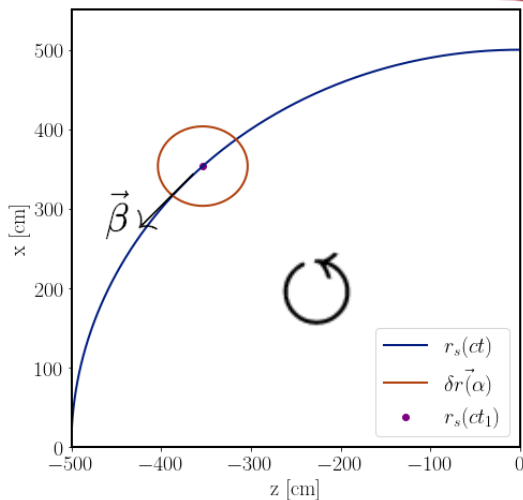
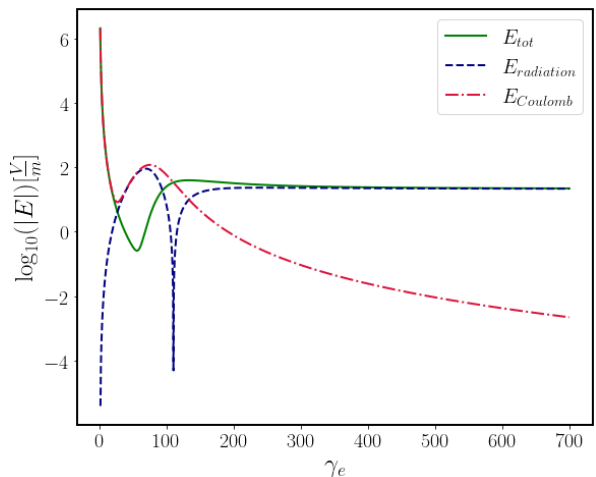
$$\vec{E}_{Radiation} \propto \left(\frac{1}{c\tau} \right)$$

III: $\gamma = 1.95 \cdot 10^2$, $\delta r = 1.00 \cdot 10^{-6}$ [m], $R = 5.00 \cdot 10^0$ [m]

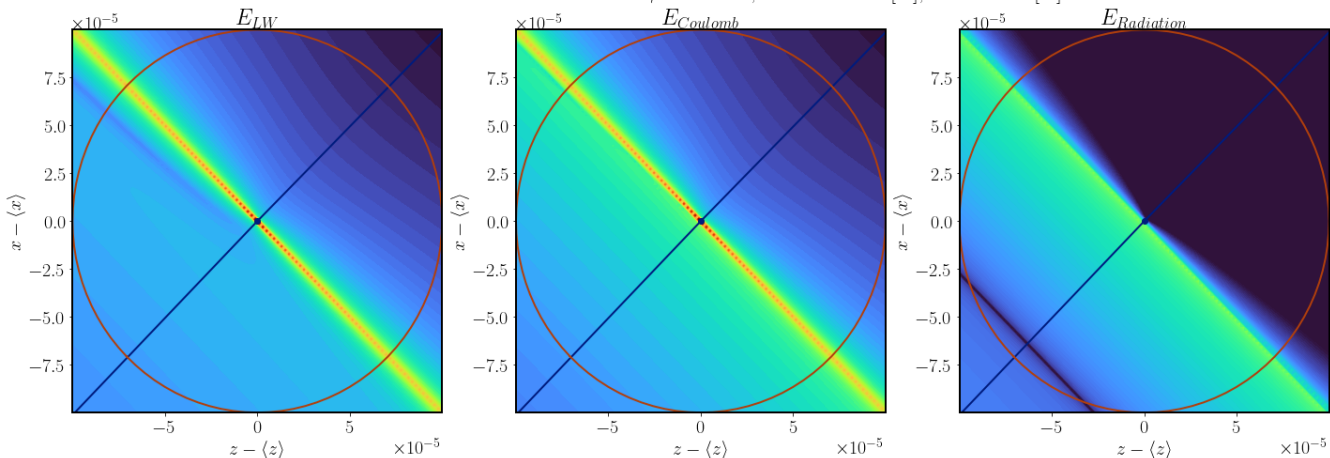


I. Self fields

2D Model



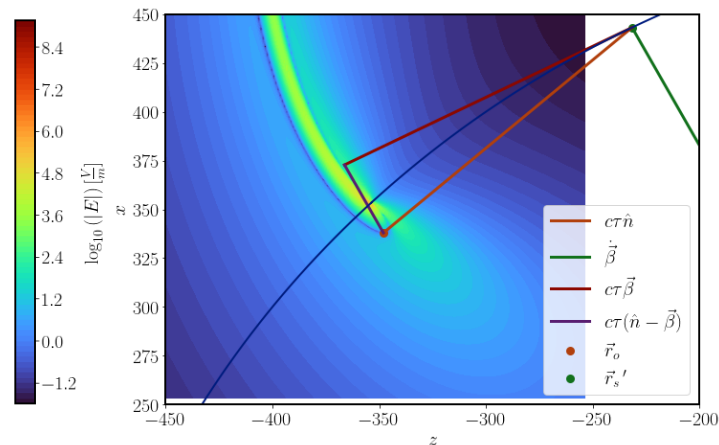
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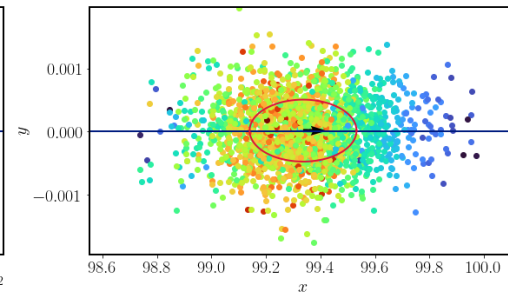
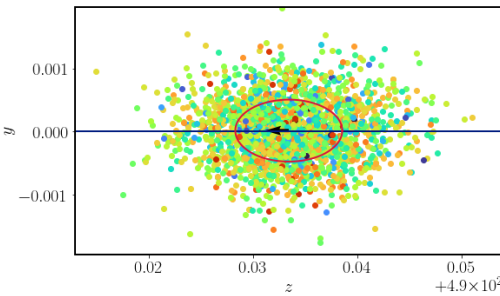
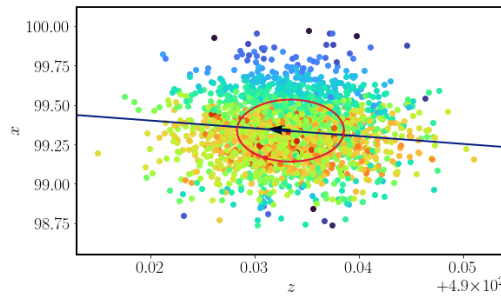
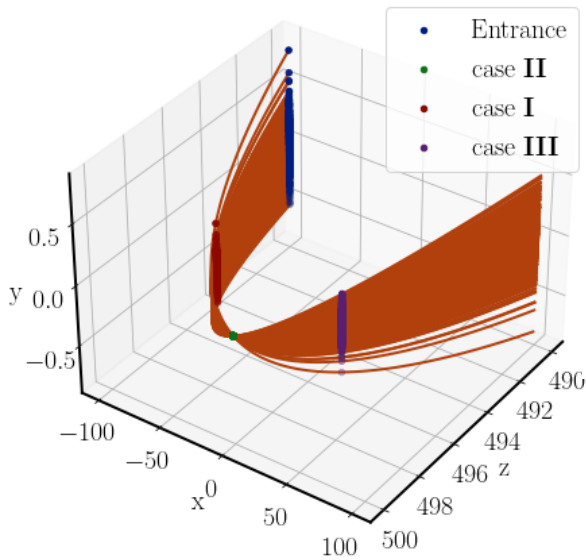
$$\vec{E}_{Coulomb} \propto \left(\frac{1}{\gamma c \tau} \right)^2$$

$$\vec{E}_{Radiation} \propto \left(\frac{1}{c \tau} \right)$$

$$\vec{E}_{Radiation} \propto \hat{n} \times \left((\hat{n} - \vec{\beta}) \times \frac{d\vec{\beta}}{dct} \right)$$

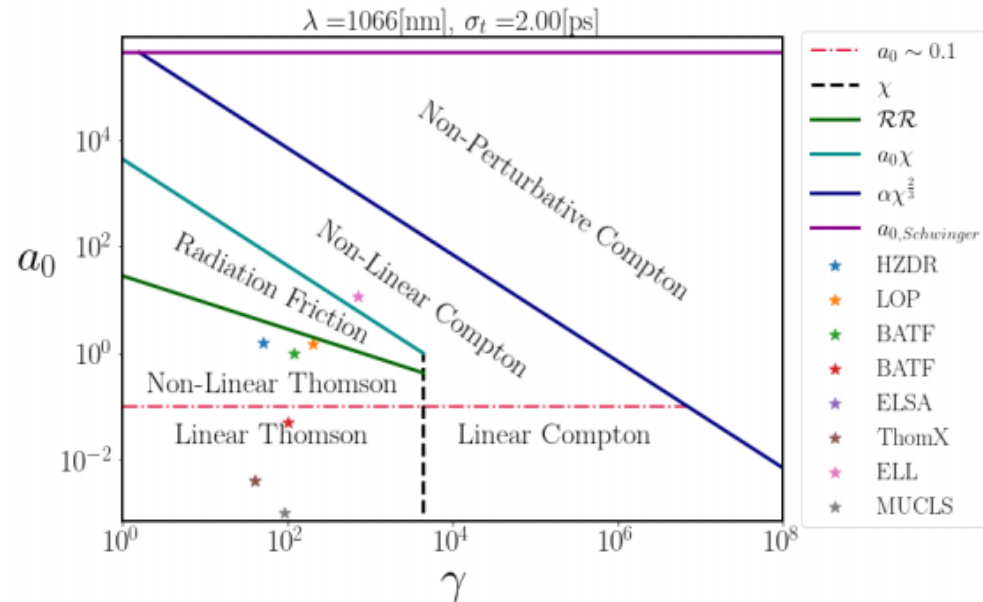
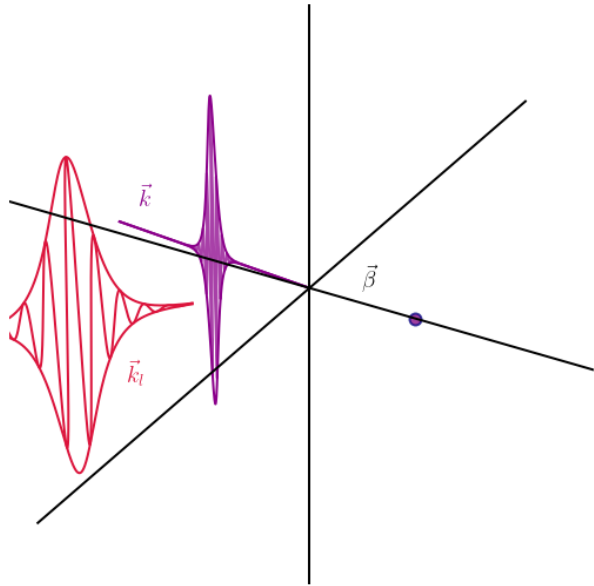


Model for a bunch traveling inside a dipole



For research developed own Thomson & Particle tracking code:

- Classical: Linear & Non-Linear regime



Scattering of High intensity lasers on Electrons

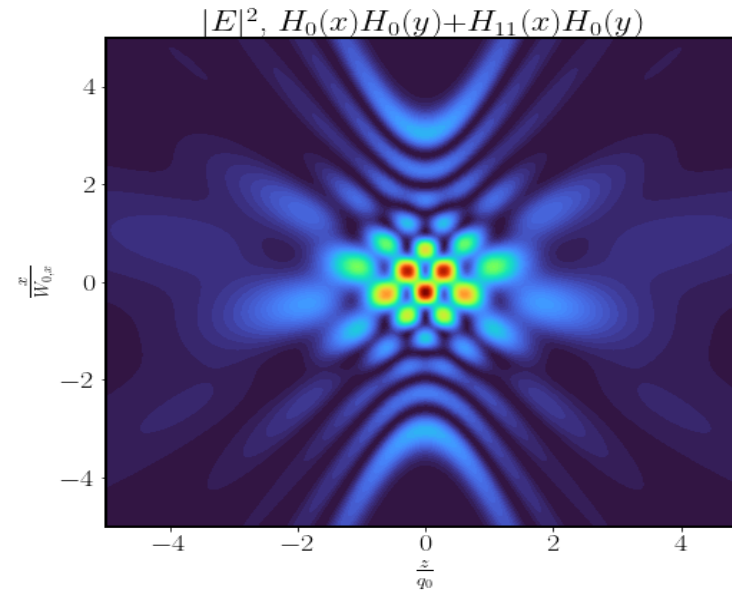
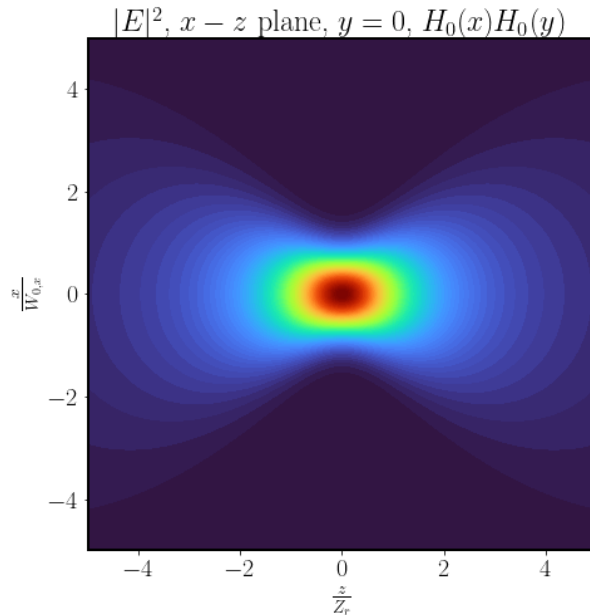
- Laser cavity: Fabry Perot $I[\text{W}/\text{cm}^2] \sim 10^{14} - a_0 < 10^{-2}$
 - Degenerate Cavity
- Chirped Pulse Amplification $I[\text{W}/\text{cm}^2] \sim 10^{18} - a_0 \sim 1$
 - Energy Compensation
 - Carrier Envelope Phase

II. Thomson scattering

Laser cavity: Fabry Perot

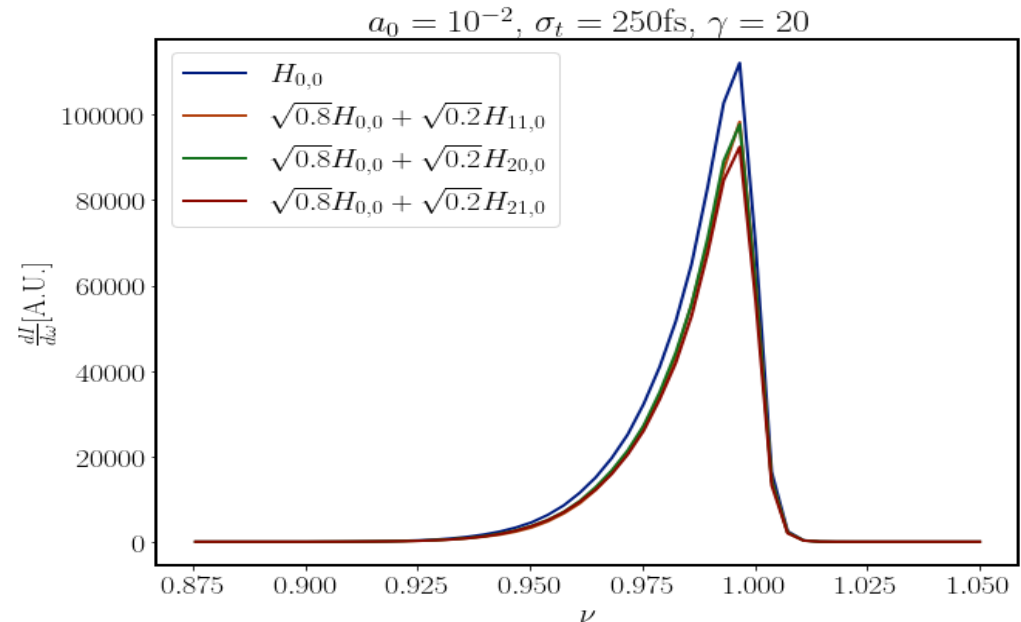
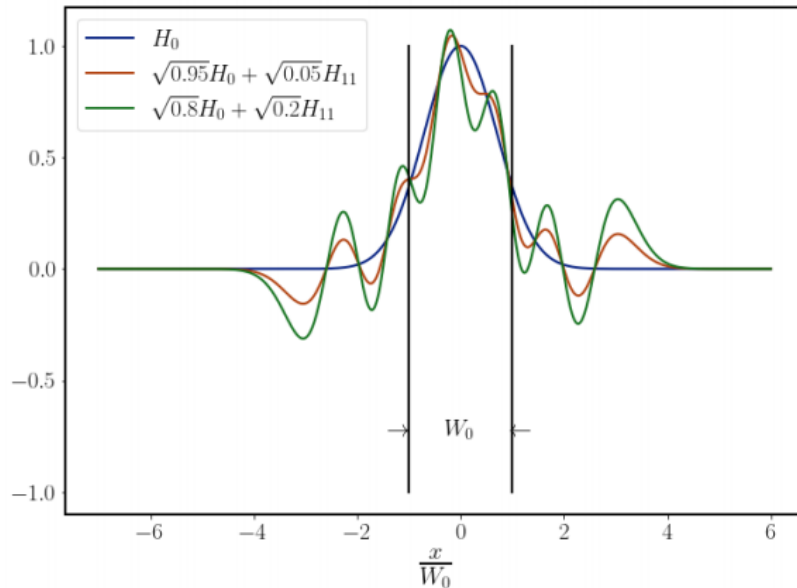
$I[\text{W}/\text{cm}^2] \sim 10^{14} - a_0 < 10^{-2} \Rightarrow$ Linear Thomson scattering

- Effect of Degenerate Cavity
- Higher modes $n > 10$



II. Thomson scattering

- Degenerate mode Power $\sim 20\%$
- Degenerate modes incoherent summation suffices. $W_e \sim W_0$
- No distinction between different modes in Thomson spectrum

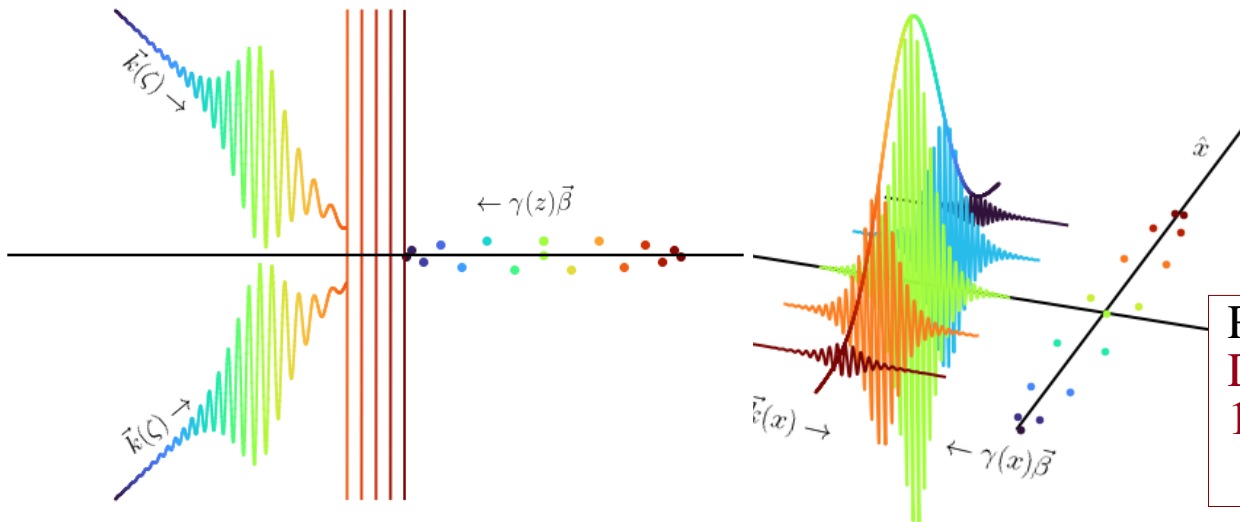


II. Thomson scattering

Linear Thomson scattering Compensation of electron energy by chirped laser

- Increase flux through larger bunch charge
- Two geometries

$$\langle \lambda \rangle \propto \frac{\lambda_f}{\gamma^2}$$



Ruijter et al, 2021

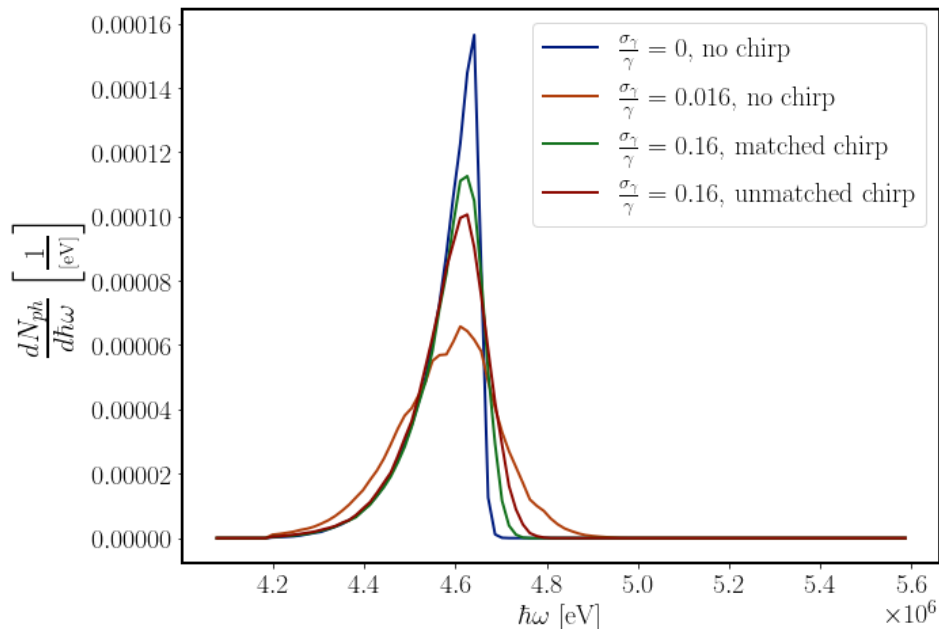
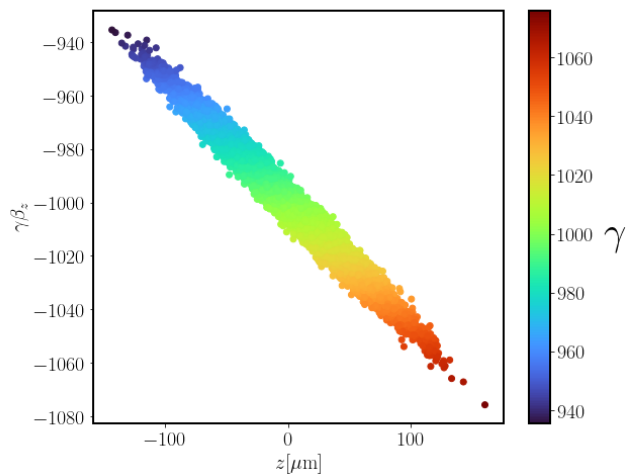
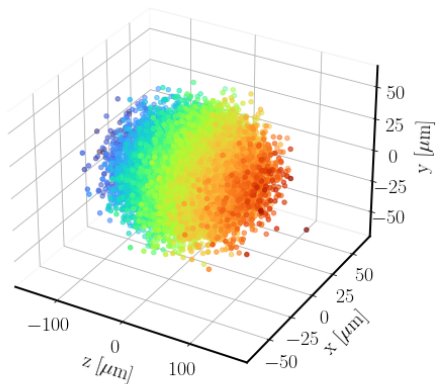
DOI:

[10.1103/PhysRevAccelBeams.24.020702](https://doi.org/10.1103/PhysRevAccelBeams.24.020702)

II. Thomson scattering

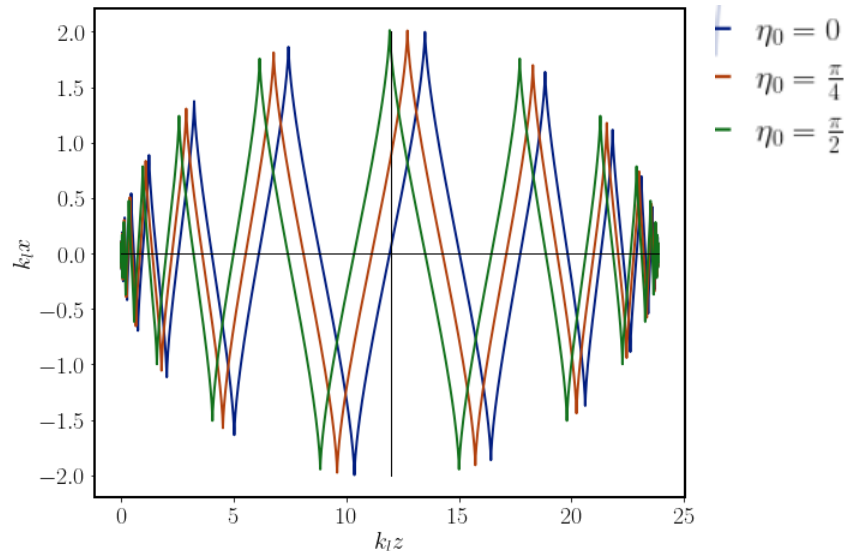
- Retrieval of ideal Thomson spectrum
- Partial compensation if
 - Mismatch in chirp & electron energy
 - Uncorrelated energy spread

Bunch Distribution, $\langle z\gamma\beta_z \rangle = 0.0583$



II. Thomson scattering

- $a_0 \geq 1$: Laser too intense to measure (matter turns into plasma)
- Carrier Envelope Phase : $L_{pulse} \sim \lambda_L$

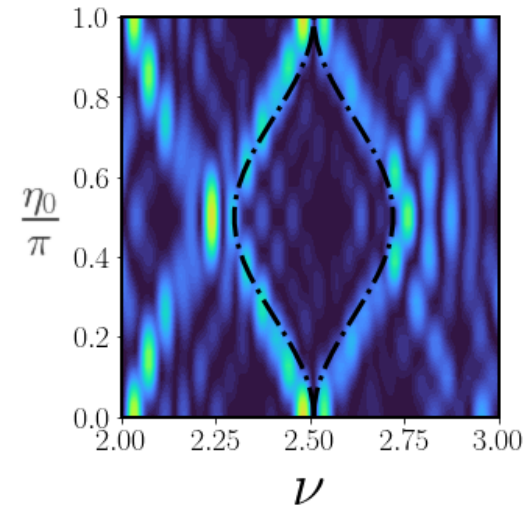
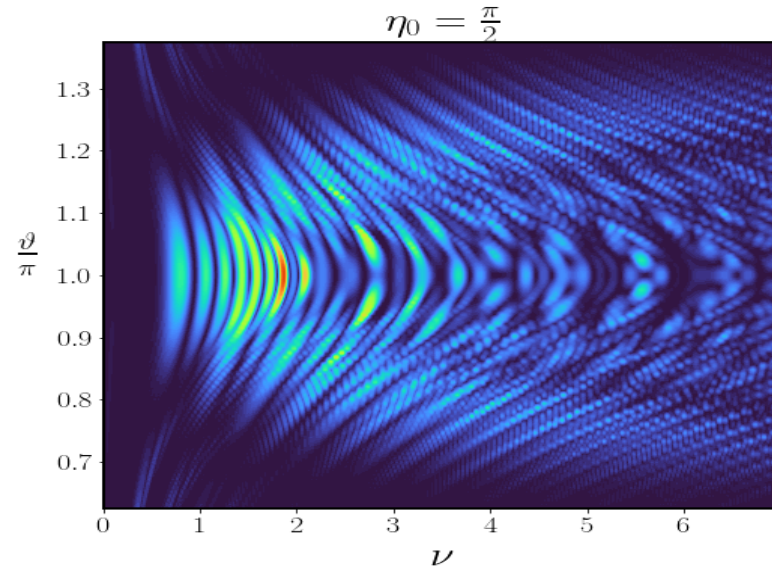
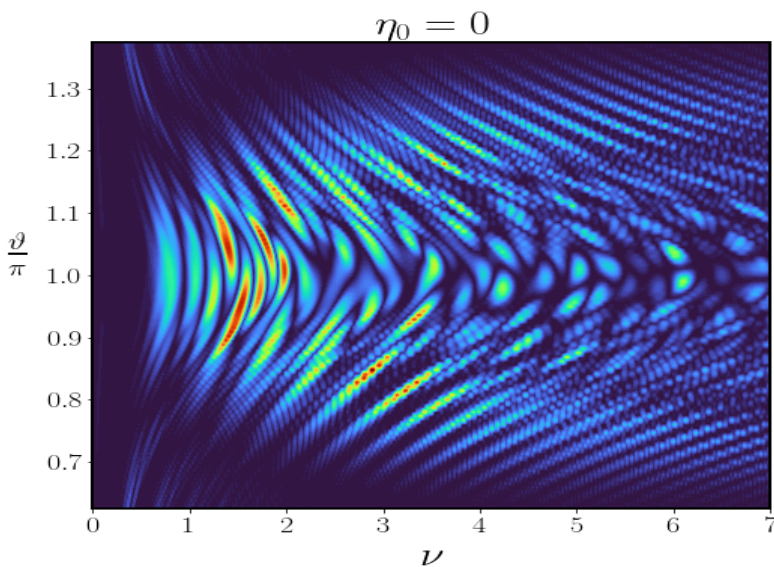


Ruijter et al, 2021
DOI:
[10.3390/cryst11050528](https://doi.org/10.3390/cryst11050528)

- Non-linear regime \Rightarrow emission of harmonics

II. Thomson scattering

- Possible to measure through Thomson scattering
 - Spectrum
 - Angular Emission



Diagnostic tool \Rightarrow low $\gamma \Rightarrow \lambda \sim 200$ nm & remain classical

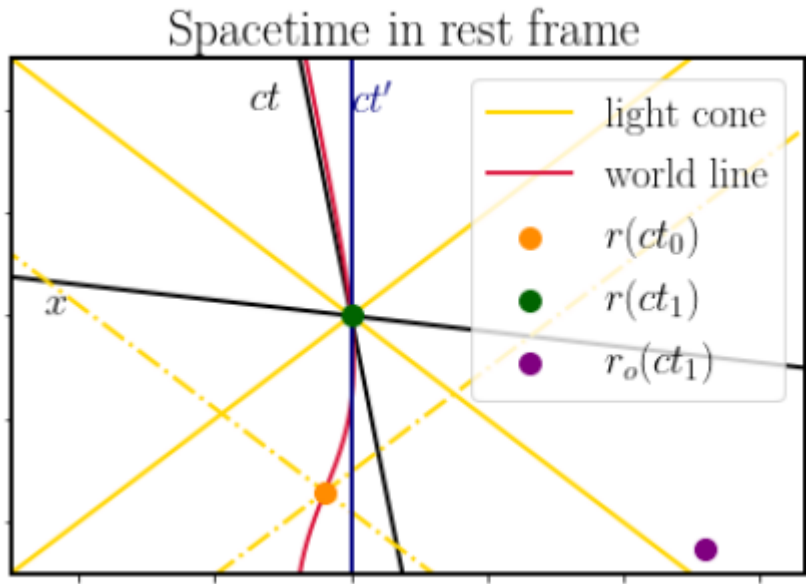
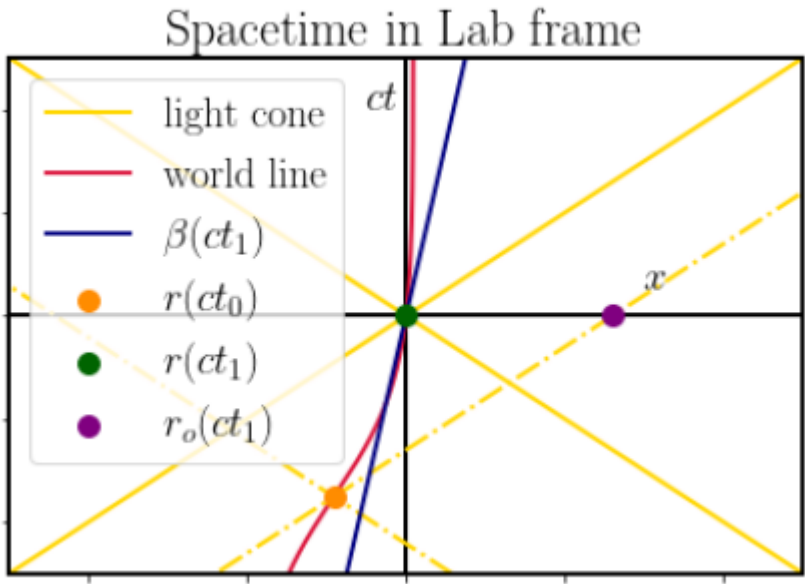
- 1D model
 - Extension
 - Relations Coulomb or Radiation field dominant
- 2D model
 - Asymmetry in behaviour of $c\tau$
 - Region where Radiation field is *zero*
- 3D model
 - Good agreement numerical and approximations
 - Optimize code & parallelisation

- Effect of Degenerate Cavity
 - Coherent summation might be required
 - For 20% power in higher modes incoherent summation suffices within W_0
 - No distinction between different modes in Thomson spectrum
- Energy Compensation
 - Two geometries
 - Could increase flux through larger bunch charge
 - Quadratic chirp
- Carrier Envelope Phase: Thomson scattering as diagnostic Tool
 - Requires $a_0 > 1$ for harmonics to overlap
 - Shift in energy peaks of harmonics



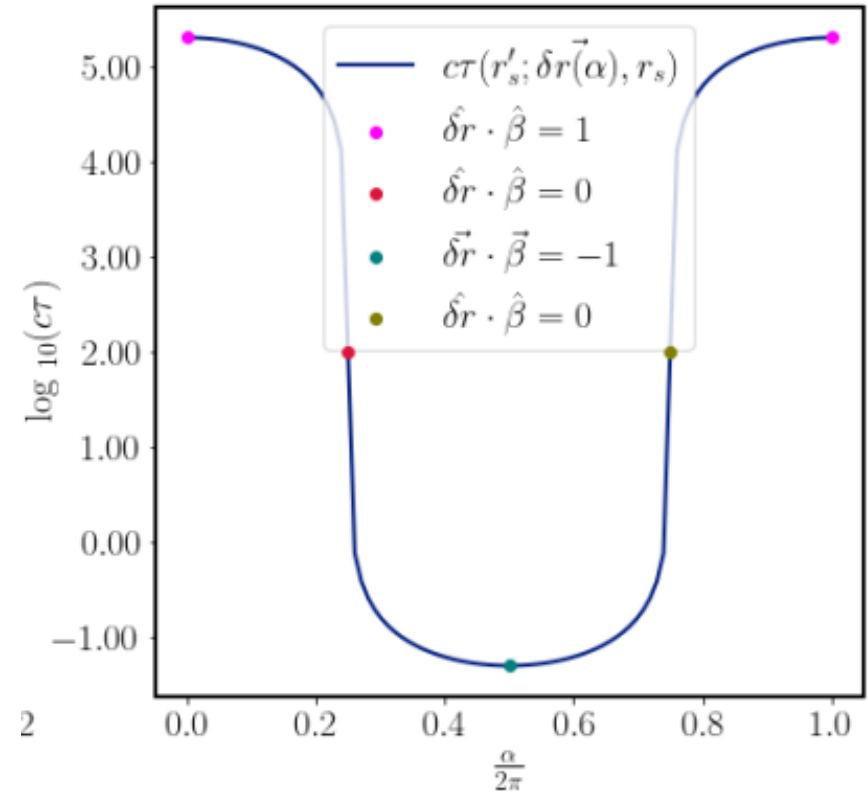
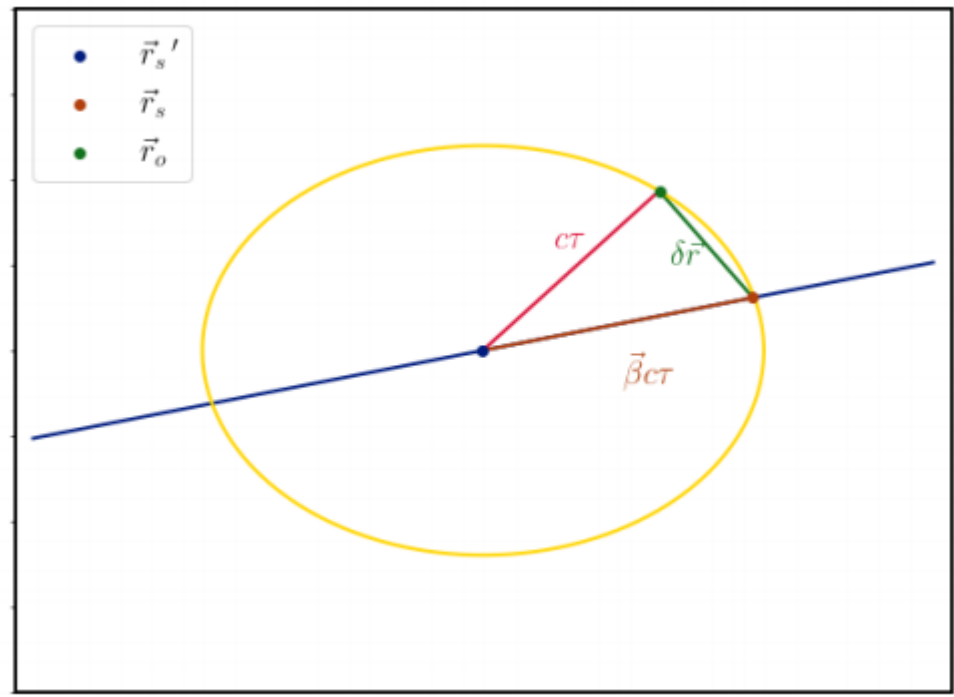
Back Up Slides

Lorentz Transformations with acceleration

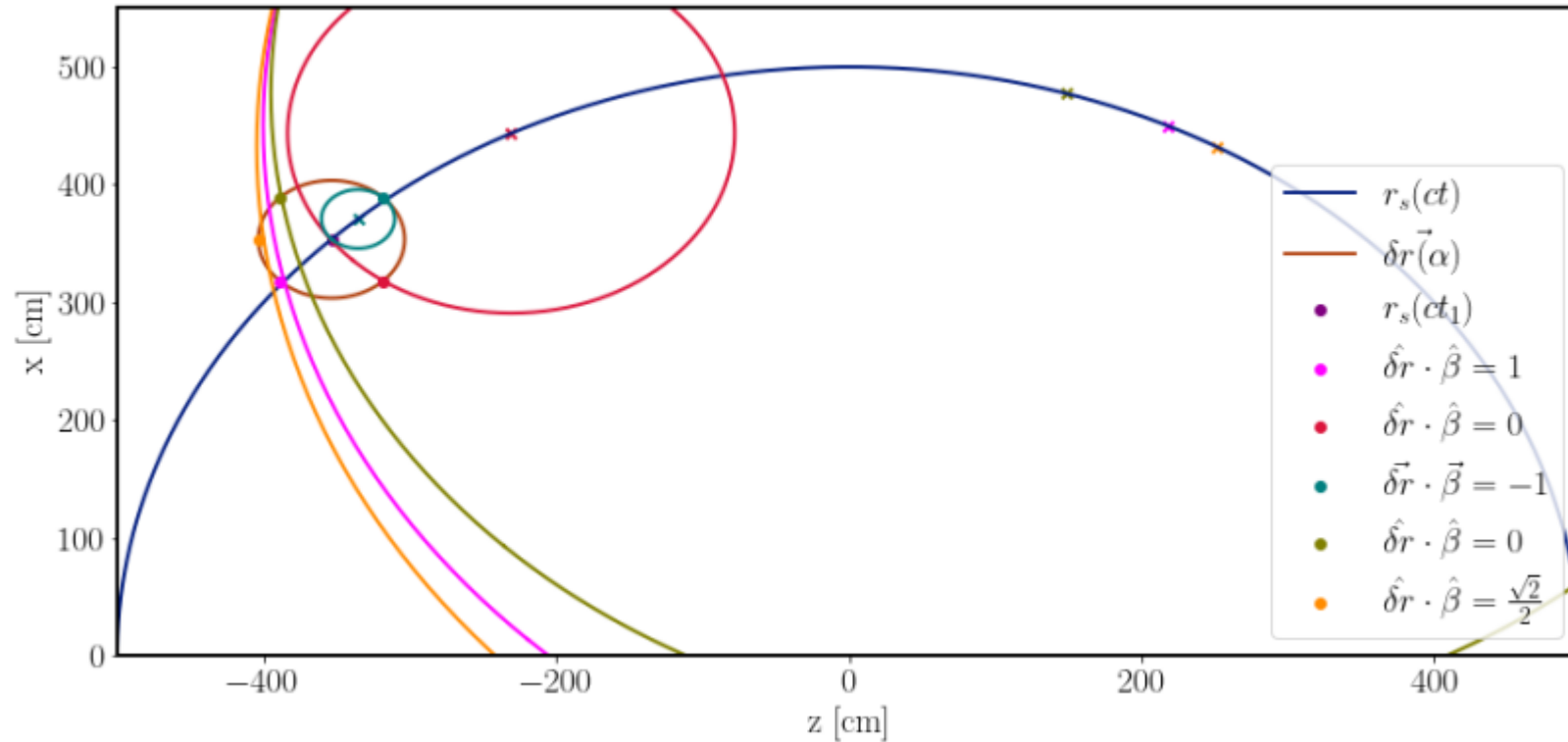


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Uniform linear motion motion (2D)



Retarded time 2D



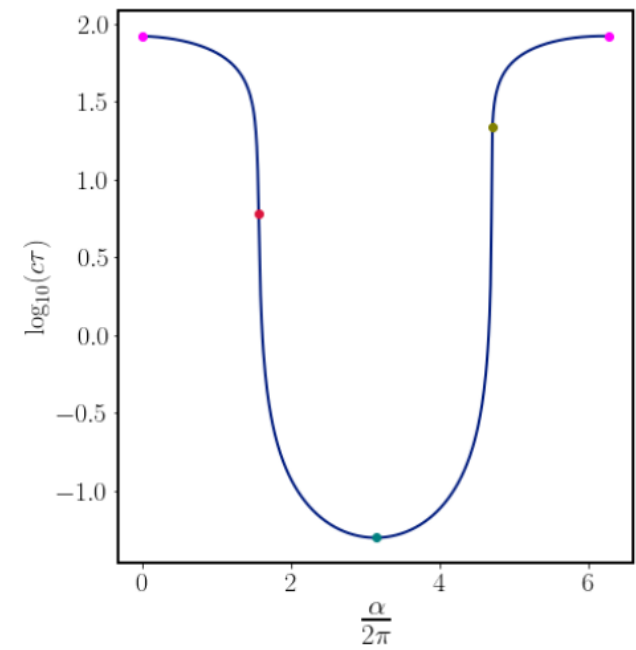
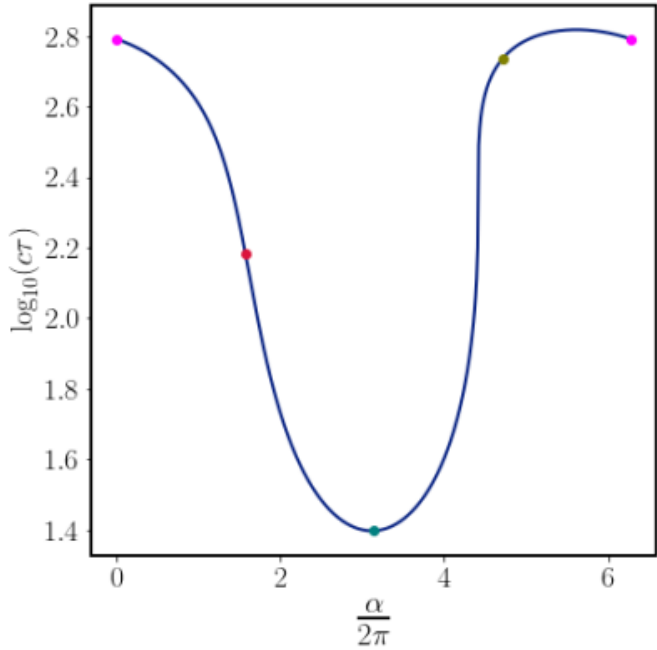
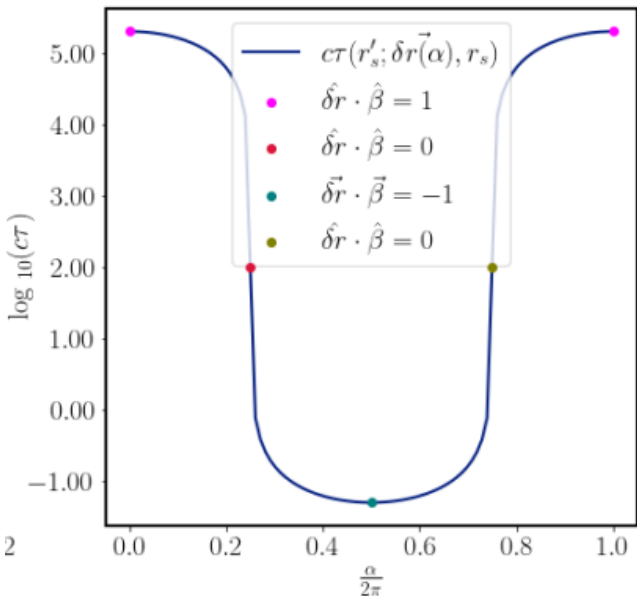
Back Up Slides

$\gamma = 10^3$, $R = 5$ [m], δr

Uniform linear

Circular large δr

Circular small δr



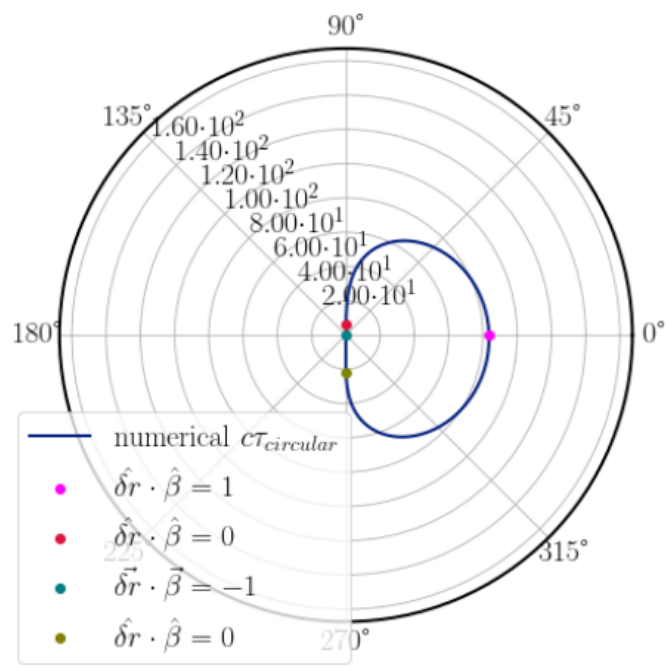
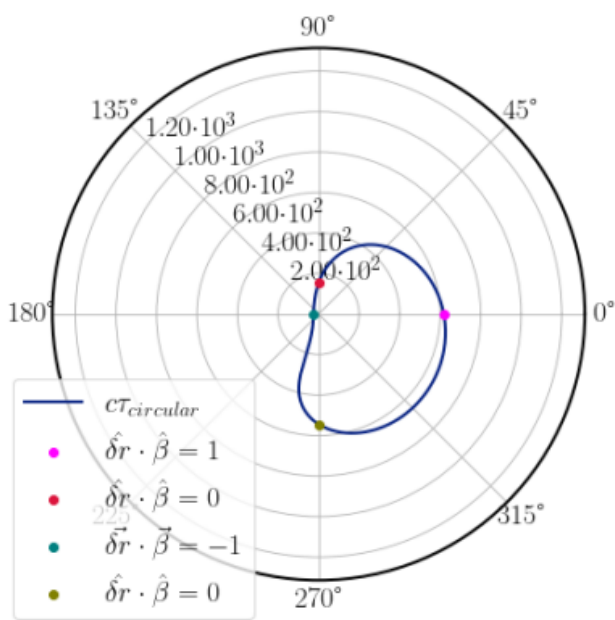
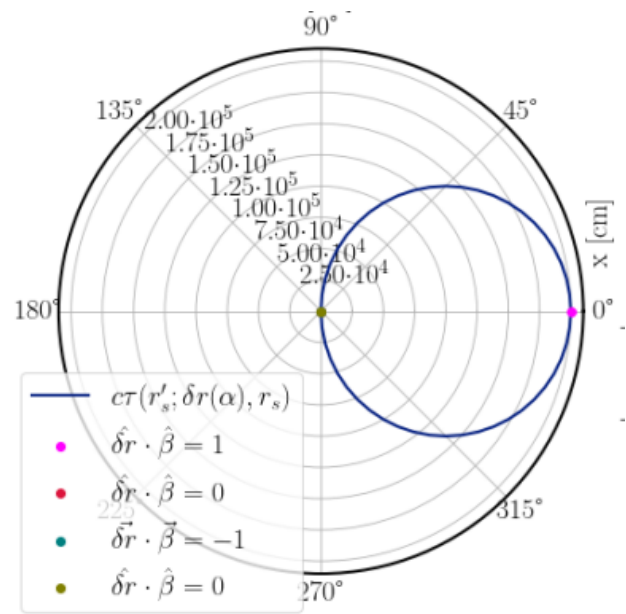
Back Up Slides

$\gamma = 10^3$, $R = 5$ [m], δr

Uniform linear
 $\delta r = 10^{-3}$ [m]

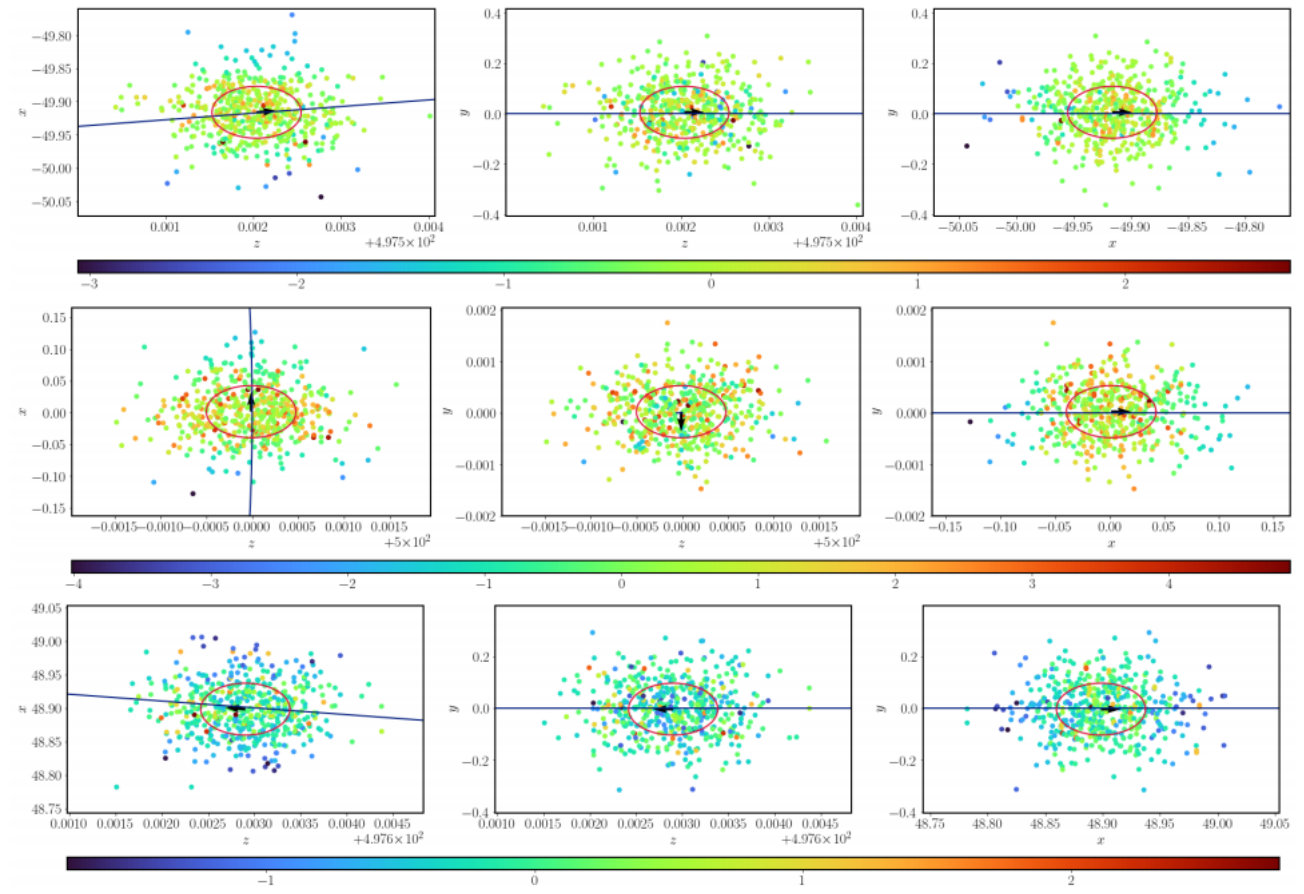
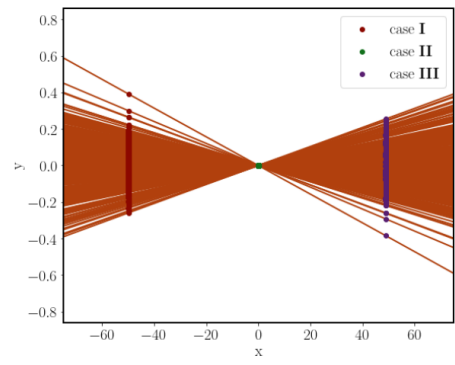
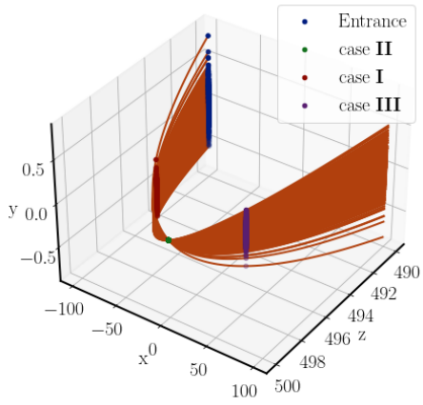
Circular large
 $\delta r = 0.5$ [m]

Circular small
 $\delta r = 10^{-3}$ [m]



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Focussing Effect E_{LW}



Meaning a_0

$$a_0 = \frac{eA_0}{mc^2} = \frac{eE_0}{mc\omega_{U_f}} = \frac{eB_0}{mc\omega_{U_f}} \quad (1.4)$$

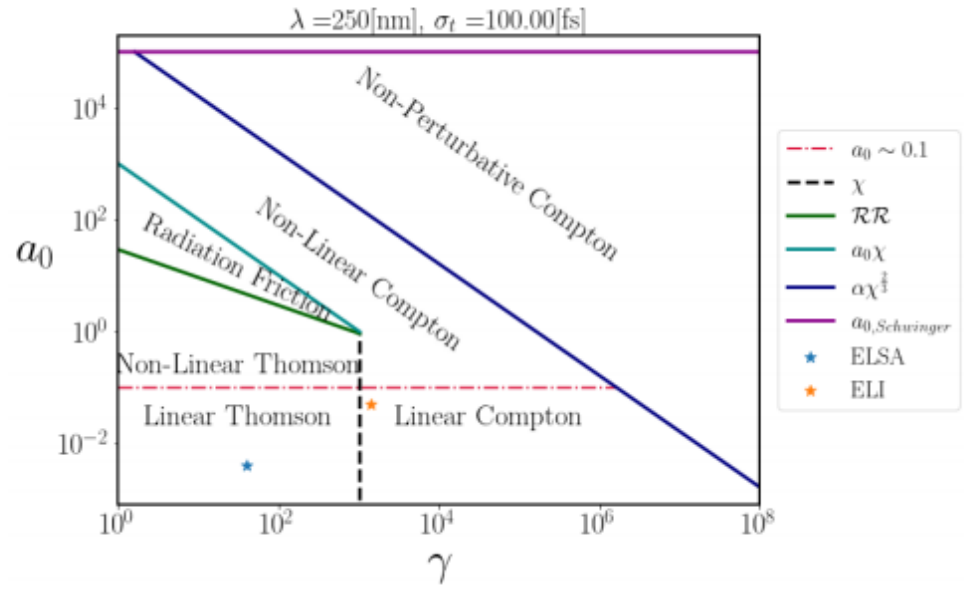
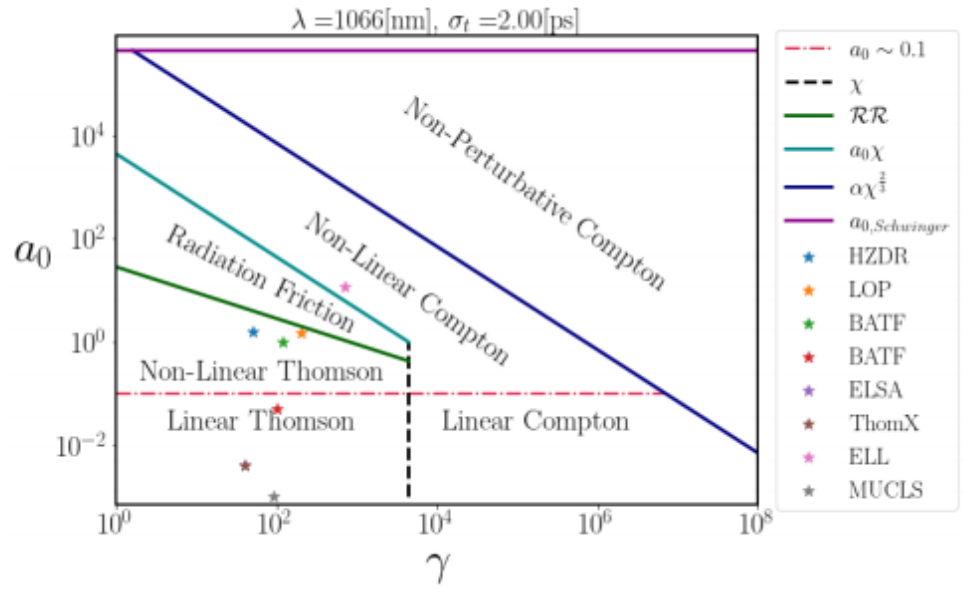
and is related to the energy density of the field ($U_f \propto (\omega_{U_f} a_0)^2$). It also describes when higher harmonics by an electron are emitted. For this we need the quantum picture of this parameter: a_0 represents the energy gain of an electron within one Compton wavelength per photon [27, 31, 32]. If $a_0 \geq 1$ then more than one photon is absorbed by the electron and emitted as one, thus giving the higher harmonics.

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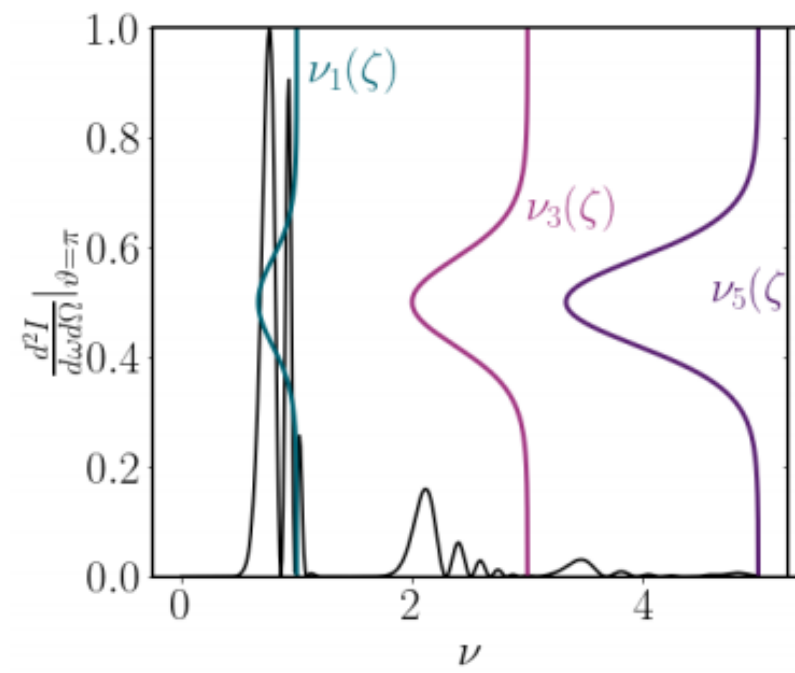
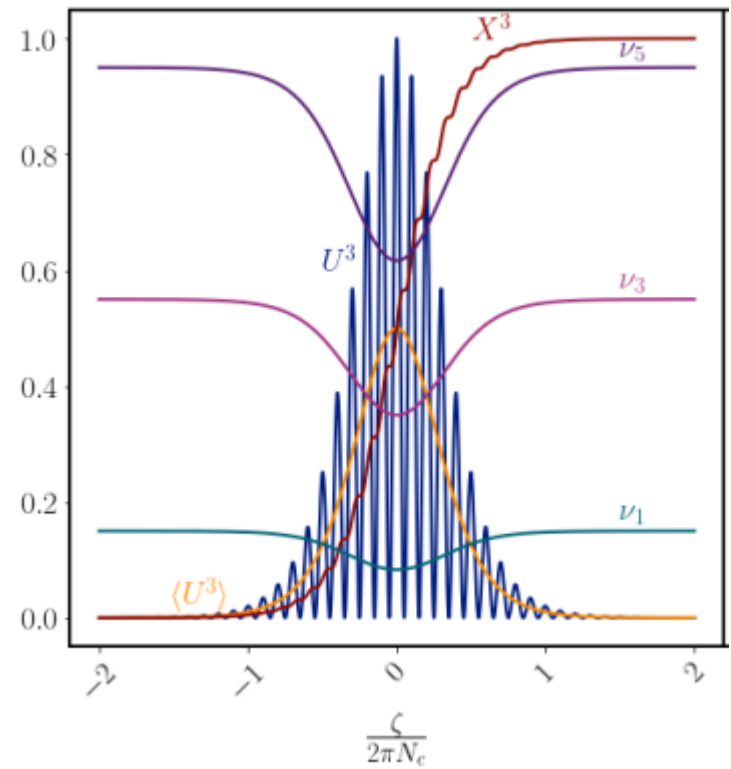
Scattering Classical vs Quantum

$$\omega \propto \frac{\omega_l}{f(\vartheta) + \chi}$$

$$\chi = 2 \frac{\hbar\gamma(1 + \beta)\omega_l}{mc^2} \approx 4 \frac{\hbar\gamma\omega_l}{mc^2}$$



Thomson Dynamics linearly polarized laser pulse



CEP – Harmonics & Interference

