

# New source model

Maxi Horst

# The new source model

## Basic idea

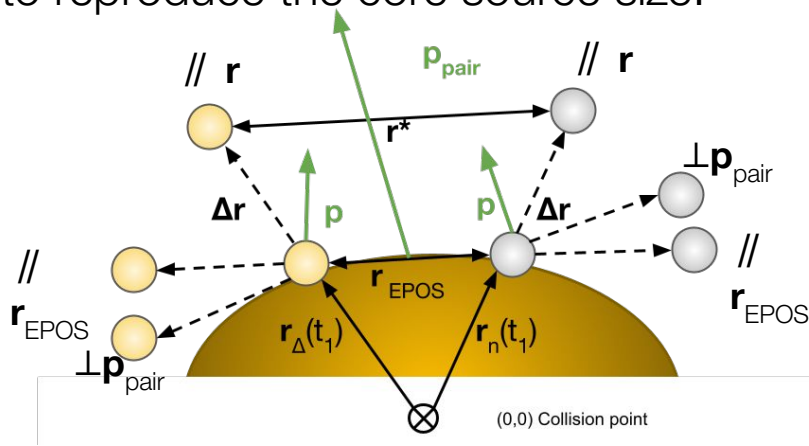
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- Create a source model in EPOS which will reproduce the size and  $m_T$  scaling of the measurement
- Basic idea: force Primordials to reproduce the core source
- Propagate resonance decays to reproduce effective source

# The new source model

## Basic idea

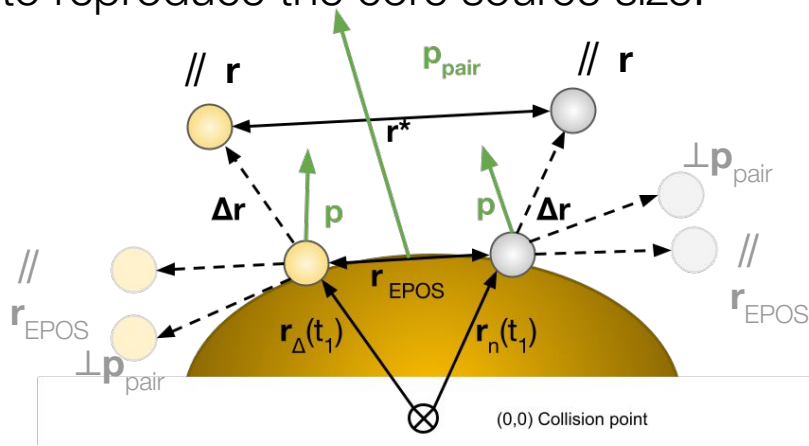
- Create a source model in EPOS which will reproduce the size and  $m_T$  scaling of the measurement
- Basic idea: force Primordials to reproduce the core source
- Propagate resonance decays to reproduce effective source
- Multiple ways to reproduce the core source size:



# The new source model

## Basic idea

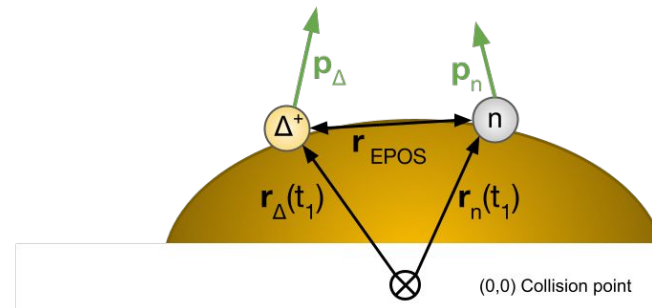
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- Propagate resonance decays to reproduce effective source
- Multiple ways to reproduce the core source size:



# The new source model

## Step-by-step

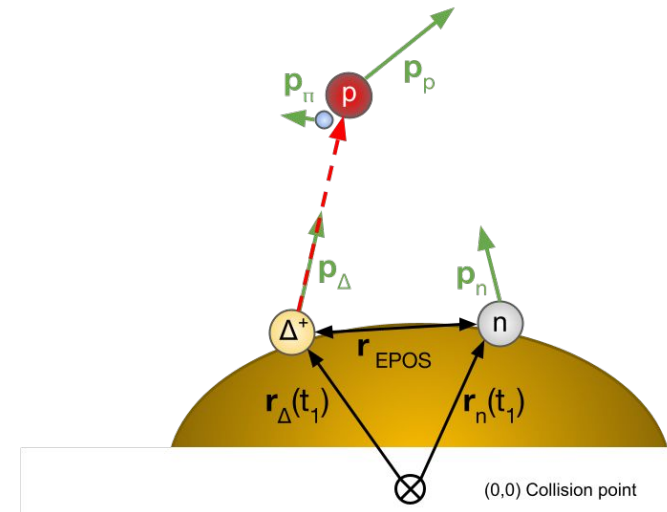
1. Propagate primordials (resonances **and** nucleons) to equal times



# The new source model

## Step-by-step

1. Propagate primordials (resonances **and** nucleons) to equal times
2. Save the vector connecting a resonance and its product (“decay vector”)

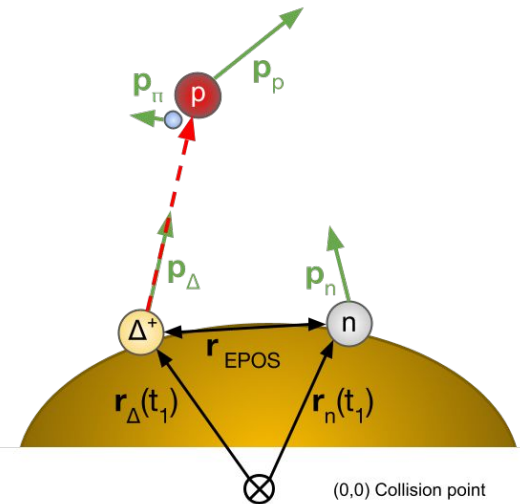
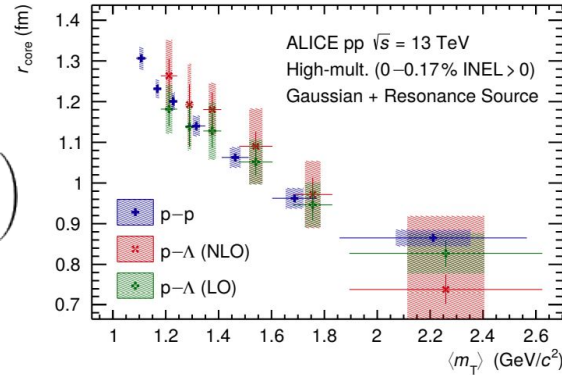


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2. Save the vector connecting a resonance and its product (“decay vector”)
3. Source size in EPOS is not correct! -> use measured  $r_{\text{core}}$

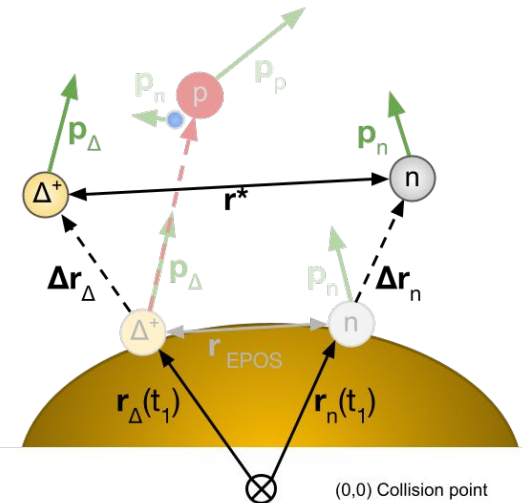
$$S(r^*) = \frac{4\pi r^{*2}}{(4\pi r_{\text{core}}^2)^{3/2}} * \exp\left(-\frac{r^{*2}}{4r_{\text{core}}^2}\right)$$



# The new source model

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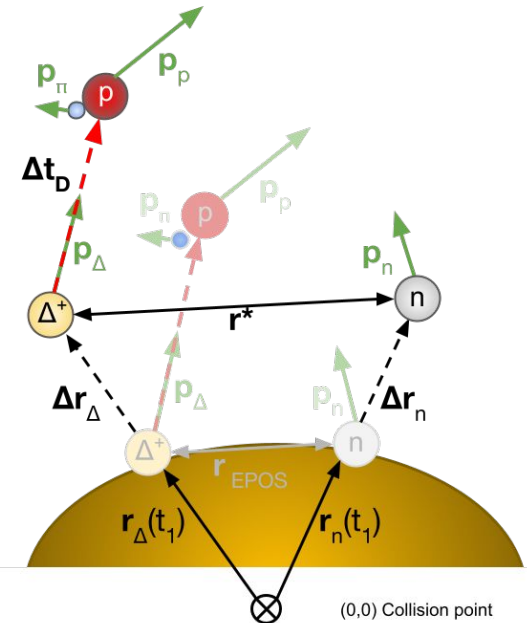
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4. Move primordials along their position vector until we reach a distance of  $r^*$



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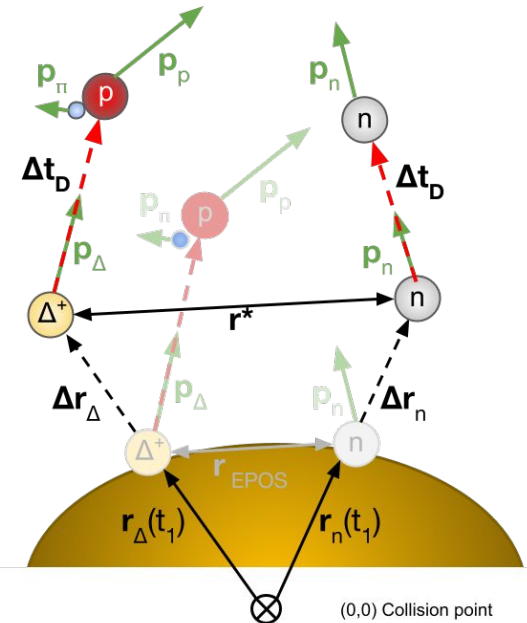
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## Step-by-step

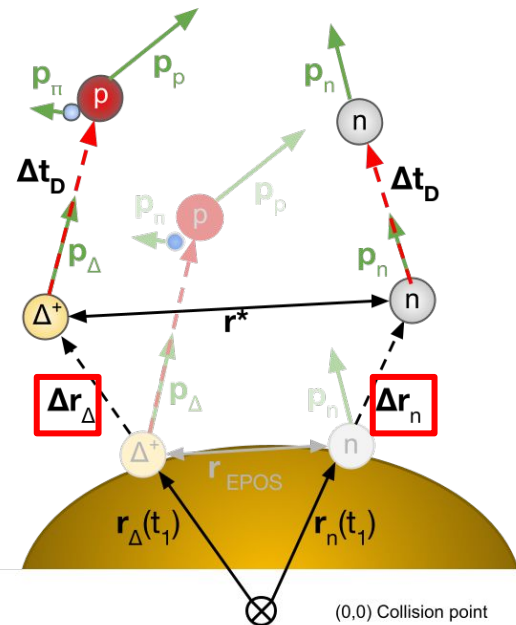
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6. Propagate to equal times



# The new source model

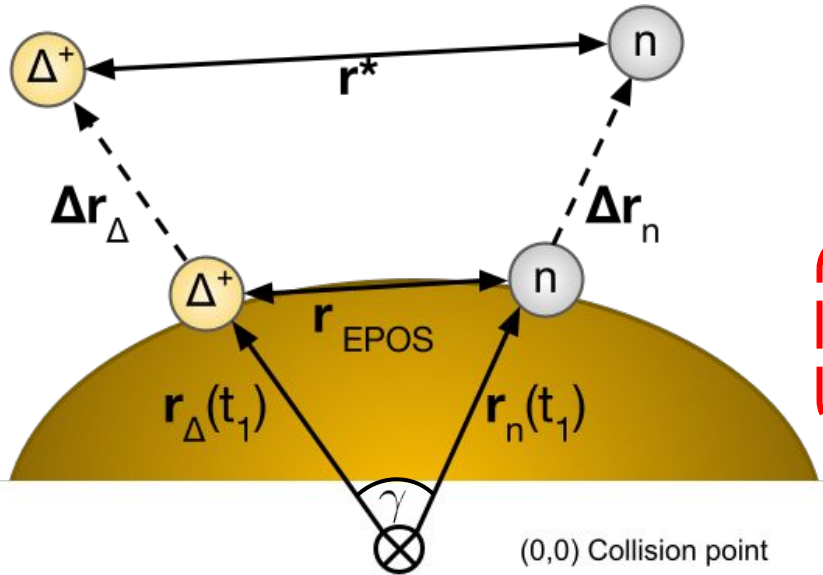
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# The new source model

## Calculating $\Delta r$



$$R = r_n / r_\Delta, \quad \Delta r^* = r^* - r_E \quad \Delta r_n = R \cdot \Delta r_\Delta^*$$

$$\text{using: } c^2 = a^2 + b^2 + 2ab \cos \gamma$$

$$\Leftrightarrow \Delta r^{*2} = \Delta r_n^2 + \Delta r_\Delta^2 + 2\Delta r_n \Delta r_\Delta \cos \gamma$$

$$\Rightarrow \Delta r_\Delta = \frac{\Delta r^*}{\sqrt{R^2 + 1 - 2R \cos \gamma}} = \Delta r_n / R$$

# Shao et al. Coalescence paper

Production of light antinuclei in pp collisions by dynamical coalescence and their fluxes in cosmic rays near earth  
(<https://arxiv.org/pdf/2205.13626.pdf>)

## Coalescence model used

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- They don't describe their Coalescence model in detail
- Ref [38]: (same model, but for  ${}^3_{\Lambda}\text{H}$  and  ${}^3\text{He}$  <https://arxiv.org/pdf/0908.3357.pdf>)

tonic and hadronic effect. The overlap Wigner phase-space density of the three-hadron cluster,  ${}^3_{\Lambda}\text{H}(\text{p}, \text{n}, \Lambda)$  and  ${}^3\text{He}(\text{p}, \text{p}, \text{n})$ , is then calculated as discussed above, and a Monte-Carlo sampling is employed to determine if the cluster is to form a nucleus or not. A nucleus emerges if the current sample value  $\rho$  is less than  $\rho_{\Lambda}^W({}^3\text{H}({}^3\text{He}))$ . Figure 1

$$\rho_{\Lambda}^W({}^3\text{H}({}^3\text{He})) = 8^2 \exp\left(-\frac{\rho^2 + \lambda^2}{b^2}\right) \exp\left(-(\mathbf{k}_p^2 + \mathbf{k}_\lambda^2)b^2\right)$$

- A little unclear, but it looks like they use the Wigner function as a probability and use a statistical rejection method

BACKUP

## Coalescence model used

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- They use a “dynamical Coalescence model” which has been in use in the chinese community for a while (oldest reference ~2010) <https://arxiv.org/pdf/0908.3357.pdf>
- The multiplicity of a cluster of M nucleons (M-Cluster) can be expressed as

$$N_M = G \int d\mathbf{r}_{i_1} d\mathbf{q}_{i_1} \cdots d\mathbf{r}_{i_{M-1}} d\mathbf{q}_{i_{M-1}} \times \left\langle \sum_{i_1 > i_2 > \cdots > i_M} \underbrace{\rho_i^W(\mathbf{r}_{i_1}, \mathbf{q}_{i_1} \cdots \mathbf{r}_{i_{M-1}}, \mathbf{q}_{i_{M-1}})}_{\text{Single nucleon phase-space density}} \right\rangle$$

- A M-Cluster forms a nucleus when the phase-space density of an M-cluster is smaller than the Wigner function of the d/<sup>3</sup>He/<sup>3</sup>H/...

$$\rho_d^W(\mathbf{r}, \mathbf{k}) = \int \psi(\mathbf{r} + \frac{\mathbf{R}}{2}) \psi^*(\mathbf{r} + \frac{\mathbf{R}}{2}) \times \exp(-i\mathbf{k} \cdot \mathbf{R}) d^3\mathbf{R} = 8 \exp\left(-\frac{\rho^2}{\sigma_d^2} - \sigma_d^2 \mathbf{k}^2\right)$$

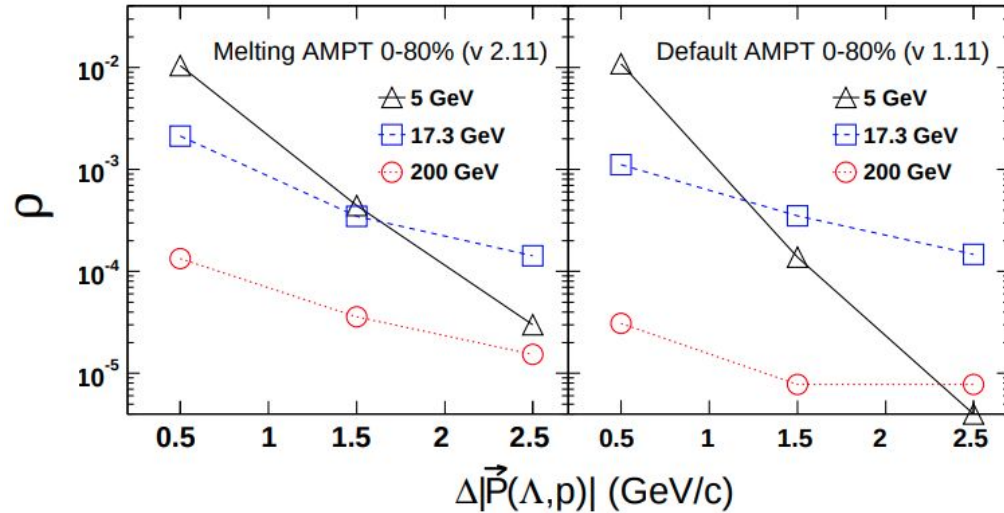


Figure 1: (color online) The Wigner phase-space density  $\rho$  for  ${}^3_{\Lambda}\text{H}$  from melting AMPT (left panel) and default AMPT (right panel) as a function of  $(\Lambda, p)$  pair momentum. Densities are shown for  $\sqrt{s_{\text{NN}}} = 5 \text{ GeV}$ ,  $17.3 \text{ GeV}$  and  $200 \text{ GeV}$ . The distributions have been normalized by the number of events at each collision energy.

## Timeline

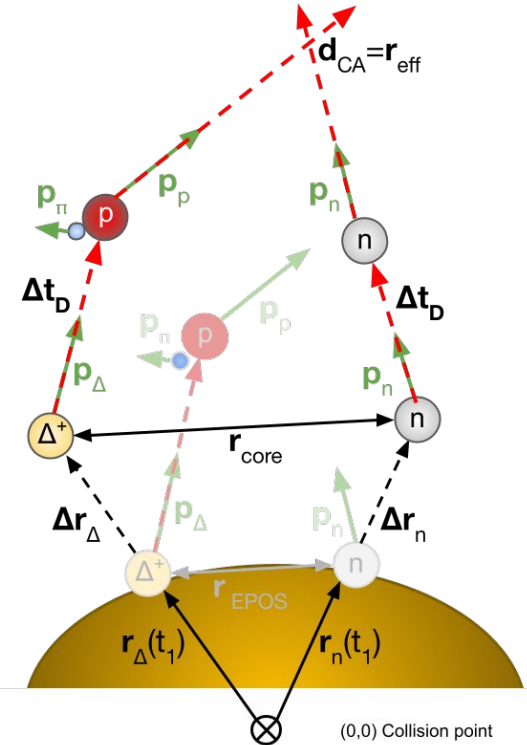
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- Currently there are ~50M events done by the Virgo (GSI) cluster
- Speed is ~10M/day
- We will reach 300M by the end of the month
- LRZ will contribute very few events for technical reason running many jobs does not scale well
- Locally also some events are running at about 2M events/day
- This will contribute ~50M events by the end of the month
- With our existing 100M events we will have 450-500M events by the end of the month

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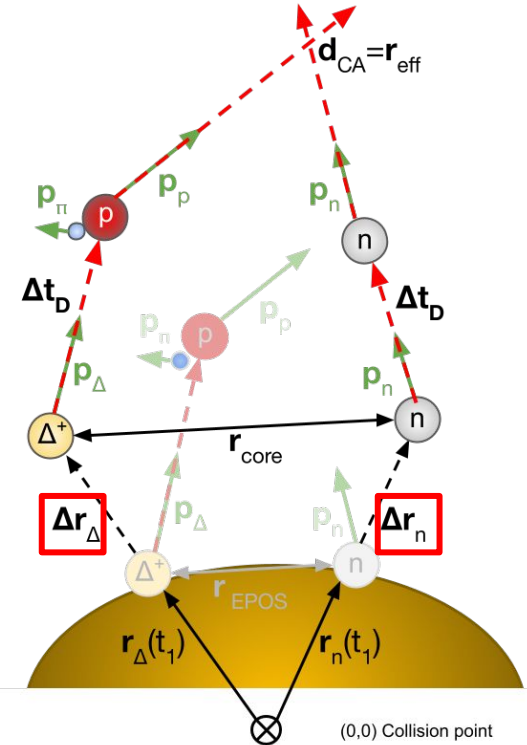
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5. Add the decay vector to moved resonance
6. Propagate to equal times
7. Calculate  $d_{\text{CA}}$  to get  $r_{\text{eff}}$



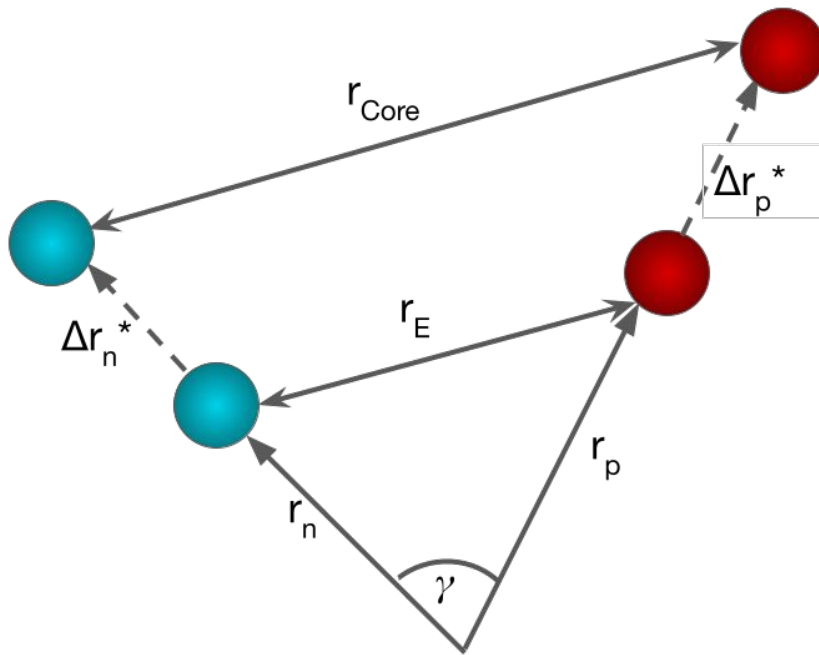
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using:  $c^2 = a^2 + b^2 + 2ab \cos \gamma$

$$\Leftrightarrow \Delta r_{\text{core}}^2 = \Delta r_n^{*2} + \Delta r_p^{*2} + 2\Delta r_n^* \Delta r_p^* \cos \gamma$$

$$\Rightarrow \Delta r_p^* = \frac{\Delta r_{\text{core}}}{\sqrt{R^2 + 1 - 2R \cos \gamma}} = \Delta r_n^* / R$$