



**Laboratory for
Nuclear Technologies**
Applied to the Environment



**Università
degli Studi
di Ferrara**



Geophysical and geochemical data for antineutrinos from reactors and the Earth

Andrea Maino

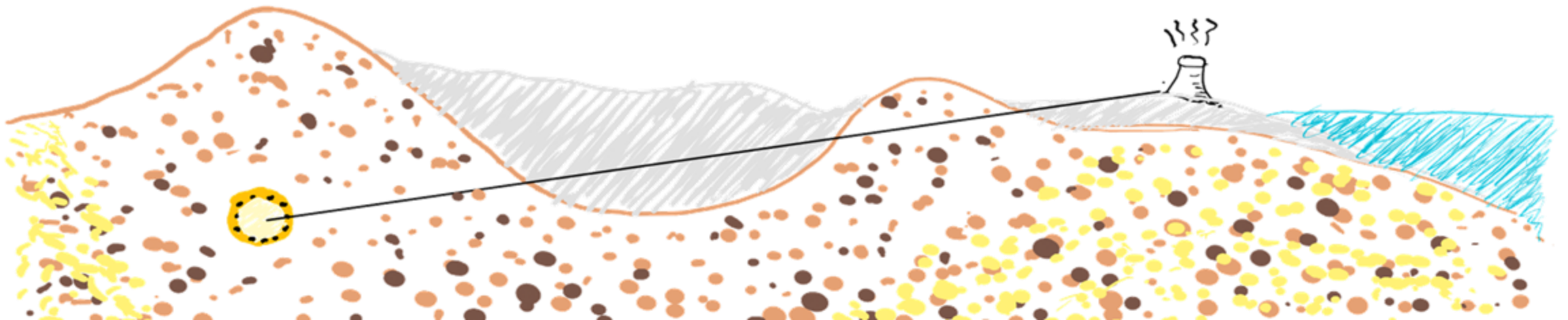
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Together with
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Summary

- Neutrino oscillations in matter
- Rock samples from JUNO: density and composition
- Electron density from a reference rock
- Implications on reactor antineutrino spectrum



How does electron density affect neutrino oscillations?

$$P_{\text{mat}}^{3\nu} \simeq c_{13}^4 P_{\text{mat}}^{2\nu} + s_{13}^4 + 2s_{13}^2 c_{13}^2 \sqrt{P_{\text{mat}}^{2\nu}} w \cos(2\Delta_{ee} + \alpha\varphi)$$

$$P_{\text{mat}}^{2\nu} = 1 - 4\tilde{s}_{12}^2 \tilde{c}_{12}^2 \sin^2 \tilde{\delta}$$

$$\bullet \quad \tilde{s}_{12}^2 \tilde{c}_{12}^2 = \frac{1}{4} \sin^2 2\tilde{\theta}_{12} \simeq \frac{1}{4} \sin^2 2\theta_{12} (1 - \mu_{12} \cos 2\theta_{12})^2$$

$$\bullet \quad \tilde{\delta} = \frac{\delta \tilde{m}^2 L}{4E} \simeq \delta m^2 (1 + \mu_{12} \cos 2\theta_{12}) \frac{L}{4E}$$

$$\mu_{12} = \frac{2\sqrt{2}G_F N_e E}{\Delta m_{12}^2}$$

$$s_{ij}^n = \sin^n \theta_{ij}, \quad c_{ij}^n = \cos^n \theta_{ij}$$

$$\Delta_{ee} = \frac{[\Delta m^2 + \alpha/2(c_{12}^2 - s_{12}^2)\delta m^2] L}{4E}$$

$$w \simeq 1 - 2\Delta_{ee}^2 \sum_n w_n \lambda_n^2$$

$$\sum_n w_n \lambda_n^2 = 2.16 \cdot 10^{-5}$$

$$\Delta m^2 = \frac{1}{2} |\Delta m_{31}^2 + \Delta m_{32}^2|$$

$$\delta m^2 = \Delta m_{21}^2, \quad \delta = \frac{\delta m^2 L}{4E}$$

$$\alpha = 1 \text{ if NH}, \quad \alpha = -1 \text{ if IH}$$

$$\varphi \simeq 2s_{12}^2 \delta \left(1 - \frac{\sin \delta}{2\delta \sqrt{P_{\text{vac}}^{2\nu}}} \right)$$

$$G_F = 1.16637 \cdot 10^{-5} \text{ GeV}^{-2}$$

Electron density N_e

$$N_e(c, \rho) = \frac{\sum_i \rho_{Ox_i} f_i \frac{e^-_{Ox_i}}{N_{Ox_i}}}{\sum_i \rho_{Ox_i} f_i} \rho_{nat}$$

i : oxide index

ρ_{Ox_i} : density of the i -th oxide

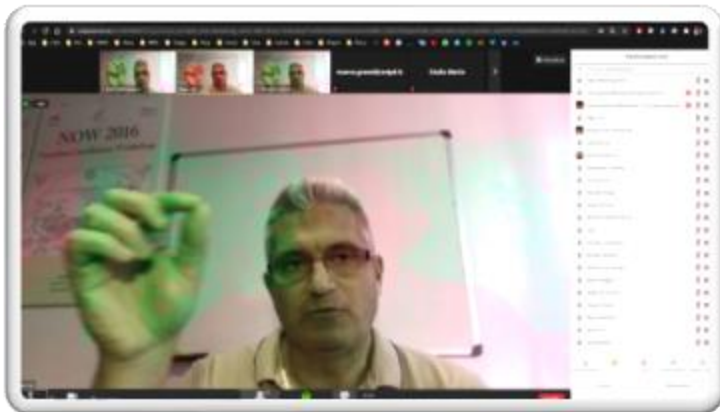
f_i : weight fraction of the i -th oxide

$e^-_{Ox_i}$: number of electrons in the i -th oxide

N_{Ox_i} : number of nucleons in the i -th oxide

ρ_{nat} : average rock density

- Elements are considered in terms of **oxides**, compounds of one or more metallic elements combined with oxygen
- Generally, the ratio $e^-_{Ox_i}/N_{Ox_i}$ is **approximated** with 0.5. This gives the **simplified** formula **$N_e(\rho) \approx 0.5 \rho_{nat}$**



“JUNO will make the first **precise** measurement of neutrino oscillations”

	Composition	ρ [g/cm ³]	N_e [mol/cm ³]
Capozzi et al., 2014	Not specified	2.6	1.3
Li et al., 2016	Not specified	2.6	1.3
Khan et al., 2020	Not specified	2.6	1.3
Our study	We need more precise values		

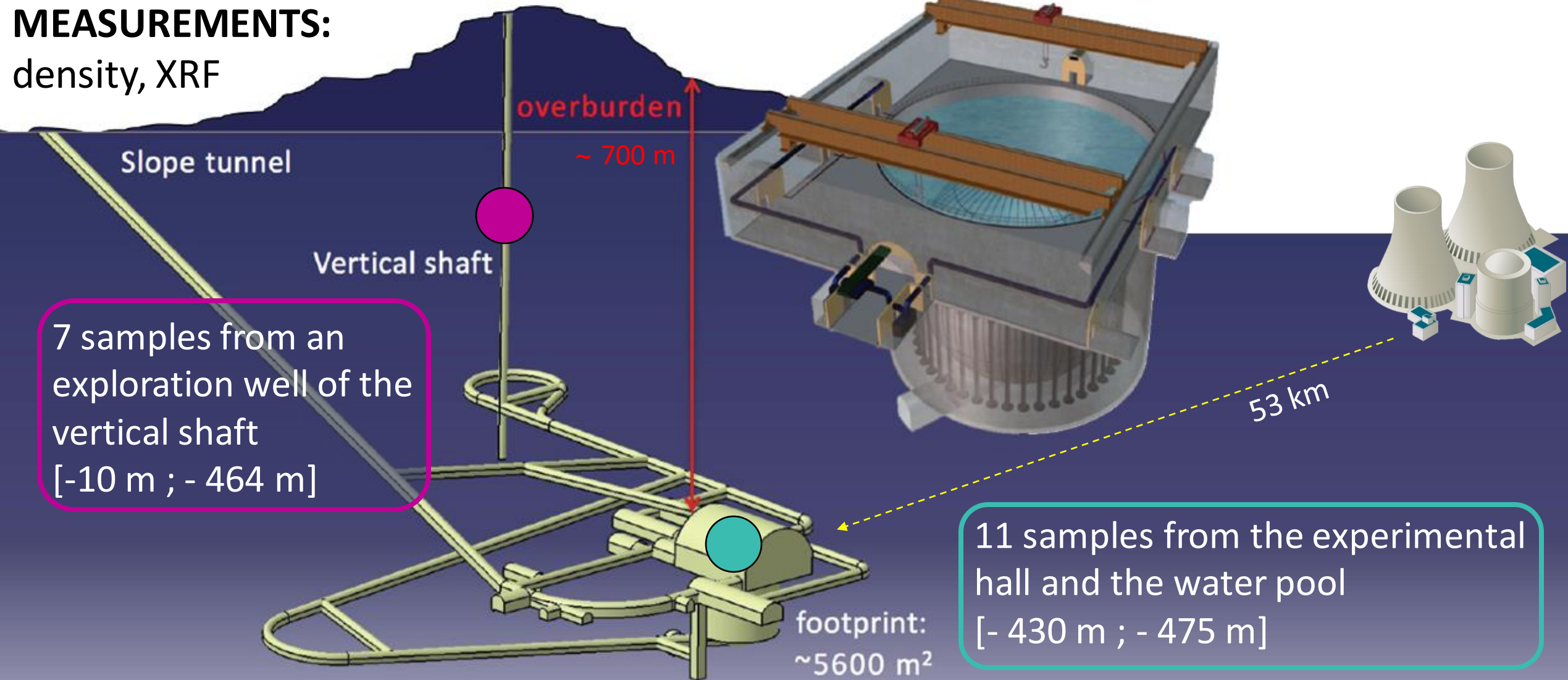
Rock samples from JUNO site

TOTAL NUMBER OF SAMPLES: 18

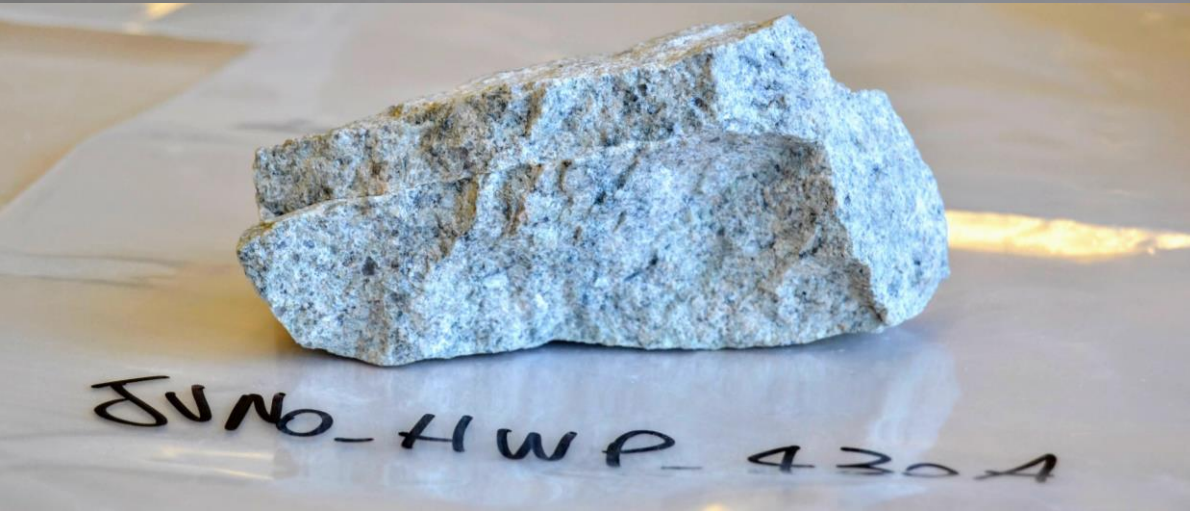
TYPE OF ROCK: GRANITES

MEASUREMENTS:

density, XRF



Rock samples' location



ID	Sample	Depth (m)	Notes
1	JUNO FRESH	10	Exploration well
2	JUNO ZK1 100	36	Exploration well
3	JUNO ZK1 200	64	Exploration well
4	JUNO ZK1 300	164	Exploration well
5	JUNO ZK1 400	264	Exploration well
6	JUNO ZK1 500	364	Exploration well
7	JUNO 2 HWP 430 A	430	Hall/water pool
8	JUNO 2 HWP 430 B	430	Hall/water pool
9	JUNO HWP 430 A	430	Lower part of the water pool
10	JUNO HWP 430 B	430	Above the water pool
11	JUNO HWP 460	460	Hall/water pool
12	JUNO 2 HWP 464	464	Exploration well
13	JUNO ZK1 600	464	Lower part of the water pool
14	JUNO HWP 465	465	Hall/water pool
15	JUNO 2 HWP 467	467	Hall/water pool
16	JUNO HWP 467	467	Lower part of the water pool
17	JUNO HWP 474	474	Hall/water pool
18	JUNO 2 HWP 475	475	Hall/water pool

Results of density measurements

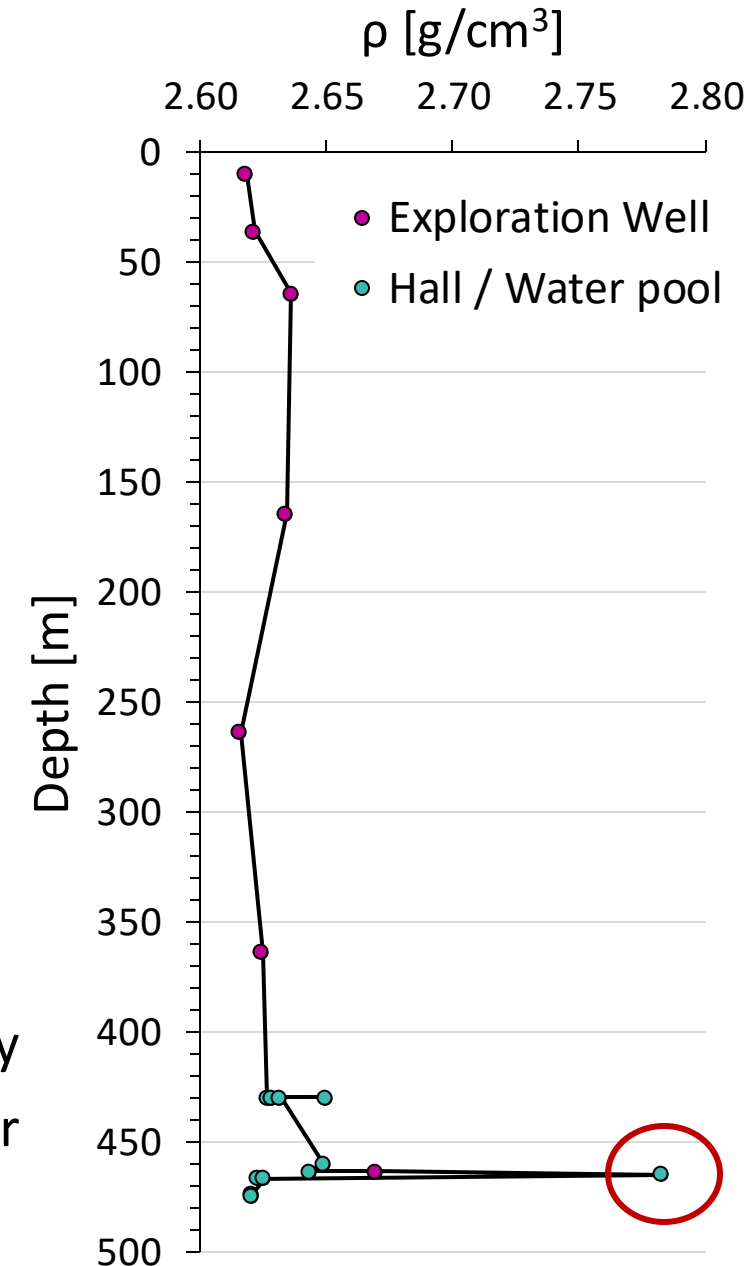
ID	Depth [m]	ρ [g/cm ³]
1	10	2.6185
2	36	2.6218
3	64	2.6363
4	164	2.6344
5	264	2.6162
6	364	2.6251
7	430	2.6267
8	430	2.6284
9	430	2.6500
10	430	2.6319
11	460	2.6489
12	464	2.6438
13	464	2.6695
14	465	2.7832
15	467	2.6233
16	467	2.6257
17	474	2.6209
18	475	2.6205

- Experimental relative uncertainty $\sim 10^{-4}$ g/cm³
- $\langle \rho \rangle = 2.6403 \pm 0.0382$ [g/cm³]
- Std. dev. $\sim 1.4\%$ -> homogeneity

JUNO HWP 465



- The appearance of the rock is certainly different from the others. The darker color suggests that more **mafic minerals** are present



XRF technique for compositional analysis

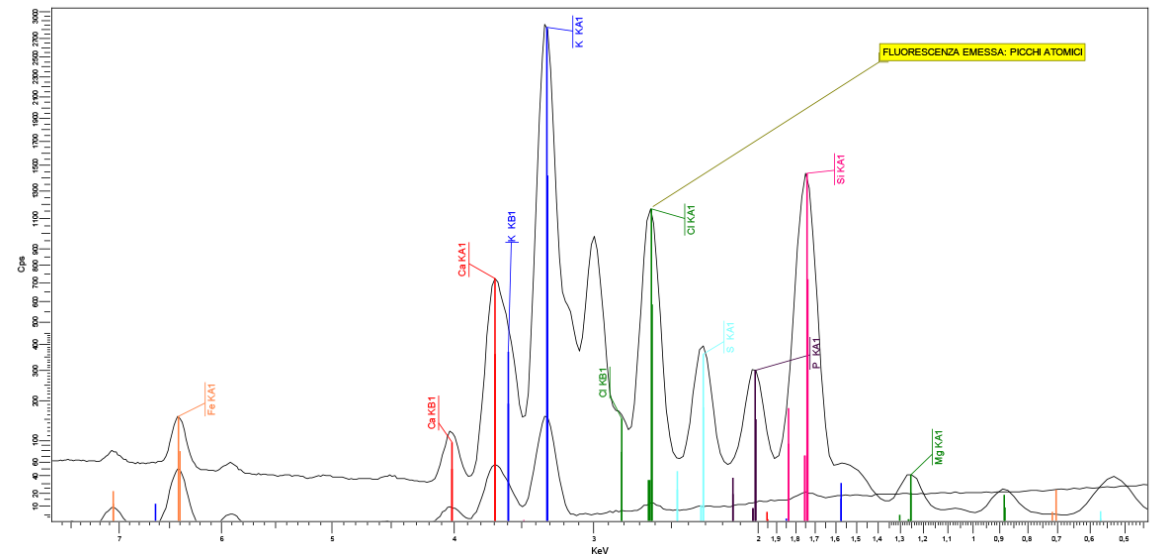
- The XRF technique (X-Ray Fluorescence) permits to identify the chemical elements constituting a sample thanks to the emitted X-rays following an induced atomic excitation.
- The **qualitative analysis** is made by **identifying** the X-emission lines associated with each element.
- The **quantitative analysis** requires a study of the **intensity** of the X-emission lines.
- The XRF technique permits to analyze both major and trace elements.

Major elements, expressed as equivalent oxides:

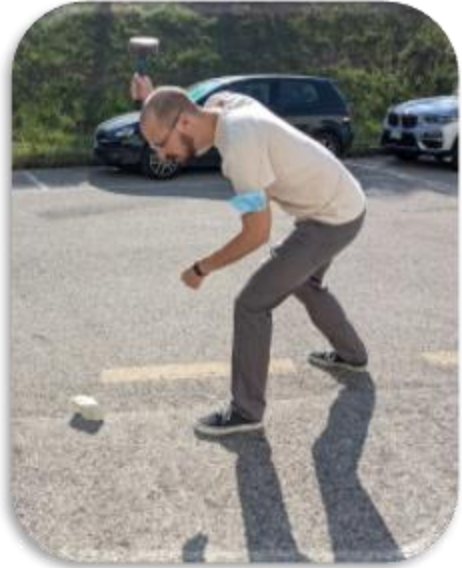
SiO₂, TiO₂, Al₂O₃, Fe₂O₃, MnO, MgO, CaO, Na₂O, K₂O, P₂O₅

Trace elements (ppm or µg/g):

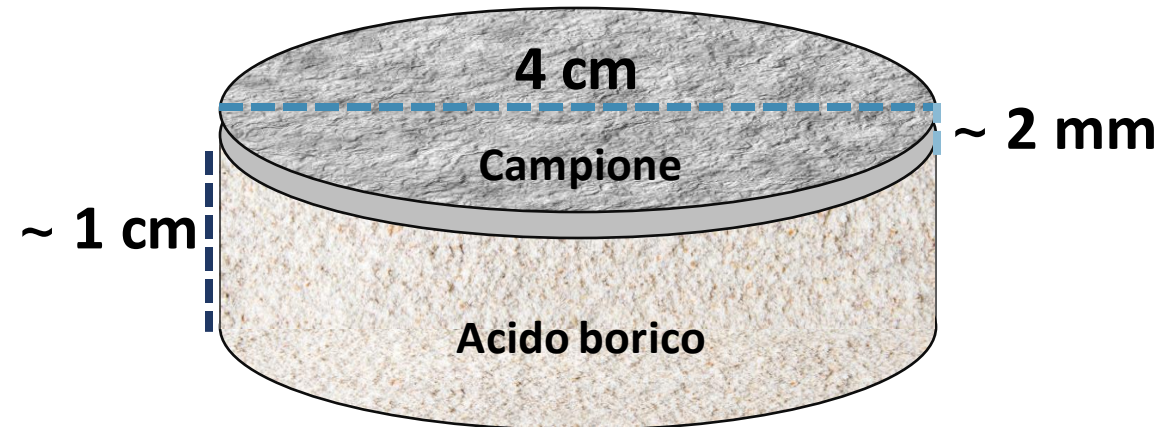
Ba, Ce, Co, Cr, Cu, Ga, Hf, La, Nb, Nd, Ni, Pb, Rb, Sc, Sr, Th, V, Y, Zn, Zr



Sample preparation



- **Shattering of the sample** by means of a hammer
- **Grinding of the sample** into small pieces
- **Redundant crushing** (grain size < 0.15 mm)
- **Paste preparation:** ~ 5 g of sample with organic glue
- **Press filling** with an additional ~ 20 g of boric acid
- **Compression** with a hydraulic press
- **Drying**



Results of XRF measurements

	Sample	Depth [m]	Si $\pm \sigma$ Si [%]	Al $\pm \sigma$ Al [%]	K $\pm \sigma$ K [%]	Ca $\pm \sigma$ Ca [%]	Fe $\pm \sigma$ Fe [%]	S $\pm \sigma$ S [ppm]	Mn $\pm \sigma$ Mn [ppm]	Zn $\pm \sigma$ Zn [ppm]	Pb $\pm \sigma$ Pb [ppm]
	JUNO_FRESH	10	35 \pm 2	6.8 \pm 0.4	3.7 \pm 0.1	0.62 \pm 0.02	0.8 \pm 0.1	N/A	507 \pm 60	30 \pm 9	55 \pm 12
	JUNO_ZK1_100	36	34 \pm 2	7.0 \pm 0.4	3.8 \pm 0.1	0.76 \pm 0.02	0.9 \pm 0.1	441 \pm 51	358 \pm 58	27 \pm 9	46 \pm 12
	JUNO_ZK1_200	64	34 \pm 2	7.1 \pm 0.4	4.2 \pm 0.1	0.60 \pm 0.02	1.0 \pm 0.1	723 \pm 51	1574 \pm 76	101 \pm 12	115 \pm 12
	JUNO_ZK1_300	164	34 \pm 2	7.1 \pm 0.4	3.9 \pm 0.1	1.18 \pm 0.03	0.9 \pm 0.1	466 \pm 51	565 \pm 60	25 \pm 9	51 \pm 12
	JUNO_ZK1_400	264	34 \pm 2	6.8 \pm 0.4	4.0 \pm 0.1	0.86 \pm 0.02	0.7 \pm 0.1	N/A	511 \pm 56	44 \pm 9	46 \pm 12
	JUNO_ZK1_500	364	34 \pm 2	7.2 \pm 0.4	4.0 \pm 0.1	0.56 \pm 0.02	1.2 \pm 0.1	1529 \pm 57	643 \pm 64	30 \pm 12	55 \pm 12
	JUNO_2_HWP_430_A	430	34 \pm 2	7.2 \pm 0.4	3.7 \pm 0.1	0.64 \pm 0.02	1.1 \pm 0.1	532 \pm 51	444 \pm 58	27 \pm 9	46 \pm 12
	JUNO_2_HWP_430_B	430	34 \pm 2	7.1 \pm 0.4	3.8 \pm 0.1	0.54 \pm 0.02	1.0 \pm 0.1	917 \pm 51	474 \pm 56	22 \pm 9	44 \pm 8
	JUNO_HWP_430_A	430	33 \pm 2	7.7 \pm 0.4	3.7 \pm 0.1	0.77 \pm 0.02	1.3 \pm 0.1	463 \pm 47	681 \pm 56	103 \pm 12	44 \pm 8
	JUNO_HWP_430_B	430	34 \pm 2	7.7 \pm 0.4	3.9 \pm 0.1	0.61 \pm 0.02	1.0 \pm 0.1	614 \pm 49	494 \pm 54	34 \pm 9	48 \pm 8
	JUNO_HWP_460	460	34 \pm 2	7.5 \pm 0.4	3.8 \pm 0.1	0.73 \pm 0.02	1.3 \pm 0.1	594 \pm 49	736 \pm 62	35 \pm 9	38 \pm 8
	JUNO_2_HWP_464	464	34 \pm 2	7.1 \pm 0.4	3.7 \pm 0.1	0.83 \pm 0.02	1.3 \pm 0.1	1506 \pm 55	846 \pm 62	182 \pm 15	91 \pm 12
	JUNO_ZK1_600	464	33 \pm 2	7.6 \pm 0.4	2.7 \pm 0.1	0.64 \pm 0.02	1.5 \pm 0.1	5257 \pm 70	3145 \pm 98	78 \pm 12	33 \pm 12
	JUNO_HWP_465	465	32 \pm 2	7.0 \pm 0.4	3.4 \pm 0.1	0.64 \pm 0.02	4.2 \pm 0.1	7584 \pm 74	2779 \pm 88	375 \pm 21	88 \pm 12
	JUNO_2_HWP_467	467	34 \pm 2	7.4 \pm 0.4	3.9 \pm 0.1	0.63 \pm 0.02	1.0 \pm 0.1	662 \pm 49	484 \pm 58	24 \pm 9	40 \pm 8
	JUNO_HWP_467	467	34 \pm 2	7.0 \pm 0.4	3.7 \pm 0.1	0.58 \pm 0.02	1.0 \pm 0.1	426 \pm 49	496 \pm 56	25 \pm 9	39 \pm 8
	JUNO_HWP_474	474	34 \pm 2	6.9 \pm 0.4	3.9 \pm 0.1	0.57 \pm 0.02	1.1 \pm 0.1	482 \pm 51	717 \pm 62	30 \pm 9	43 \pm 8
	JUNO2_HWP_475	475	34 \pm 2	7.0 \pm 0.4	3.8 \pm 0.1	0.79 \pm 0.02	1.2 \pm 0.1	451 \pm 49	558 \pm 60	27 \pm 9	44 \pm 12
Average \pm st.dev			34 \pm 1	7.2 \pm 0.3	3.8 \pm 0.3	0.70 \pm 0.16	1.3 \pm 0.8	1258 \pm 1960	890 \pm 801	68 \pm 87	54 \pm 22

- Si, Al and K contents are **homogeneous**: relative standard deviation < 8 %
- Ca and Fe are less abundant and are characterized by higher variabilities
- Trace elements (S, Mn, Zn, Pb) present some relevant **outliers**

Anomalous samples

	Sample	Depth [m]	Si \pm σ Si [%]	Al \pm σ Al [%]	K \pm σ K [%]	Ca \pm σ Ca [%]	Fe \pm σ Fe [%]	S \pm σ S [ppm]	Mn \pm σ Mn [ppm]	Zn \pm σ Zn [ppm]	Pb \pm σ Pb [ppm]
	JUNO_FRESH	10	35 \pm 2	6.8 \pm 0.4	3.7 \pm 0.1	0.62 \pm 0.02	0.8 \pm 0.1	N/A	507 \pm 60	30 \pm 9	55 \pm 12
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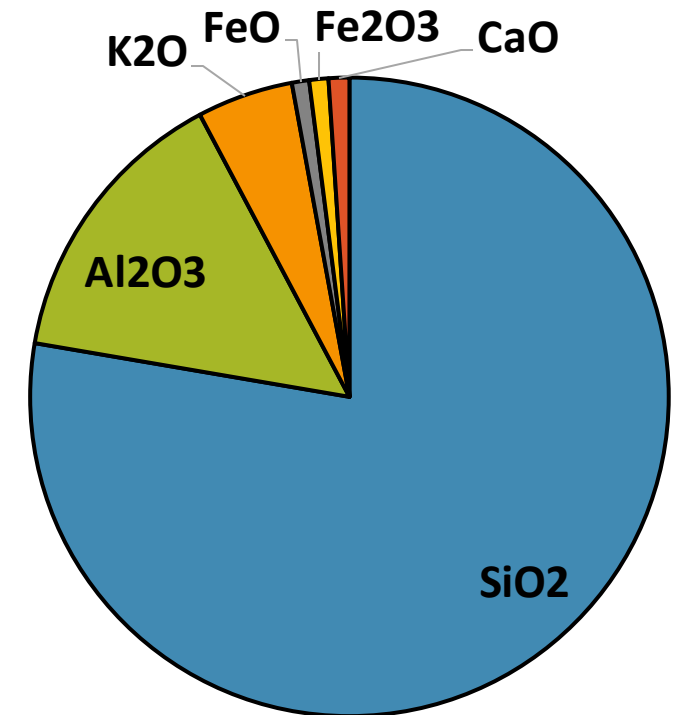
- JUNO_2_HWP_464, JUNO_ZK1_600 and JUNO_HWP_465 (outliers) come from the same depth
- JUNO_HWP_465 is the already spotted **density outlier**
- JUNO_ZK1_600 is characterized by the presence of a **quartz vein**



Building a reference rock

- **Major elements'** average concentrations have been used to build a **reference rock** representative of the local rock composition
- The composition **in terms of equivalent oxides** have been obtained using the element-to-stoichiometric oxide **conversion factors (CF)**
- Fe concentration has been divided evenly between FeO and Fe₂O₃ oxides
- The sum of the obtained oxides equate to ~ 94 % so oxide concentrations have been **rescaled** to sum at 100%
- Trace elements have been **neglected**

Element	Element wt%	CF	Oxide	Oxide wt%	Normalized Oxide wt%
Si	34 ± 1	2.139	SiO ₂	72	77
Al	7.2 ± 0.3	1.889	Al ₂ O ₃	14	15
K	3.8 ± 0.3	1.205	K ₂ O	5	5
Ca	0.70 ± 0.16	1.399	CaO	1	1
Fe	0.65 ± 0.40	1.286	FeO	0.8	1
Fe	0.65 ± 0.40	1.430	Fe ₂ O ₃	0.9	1

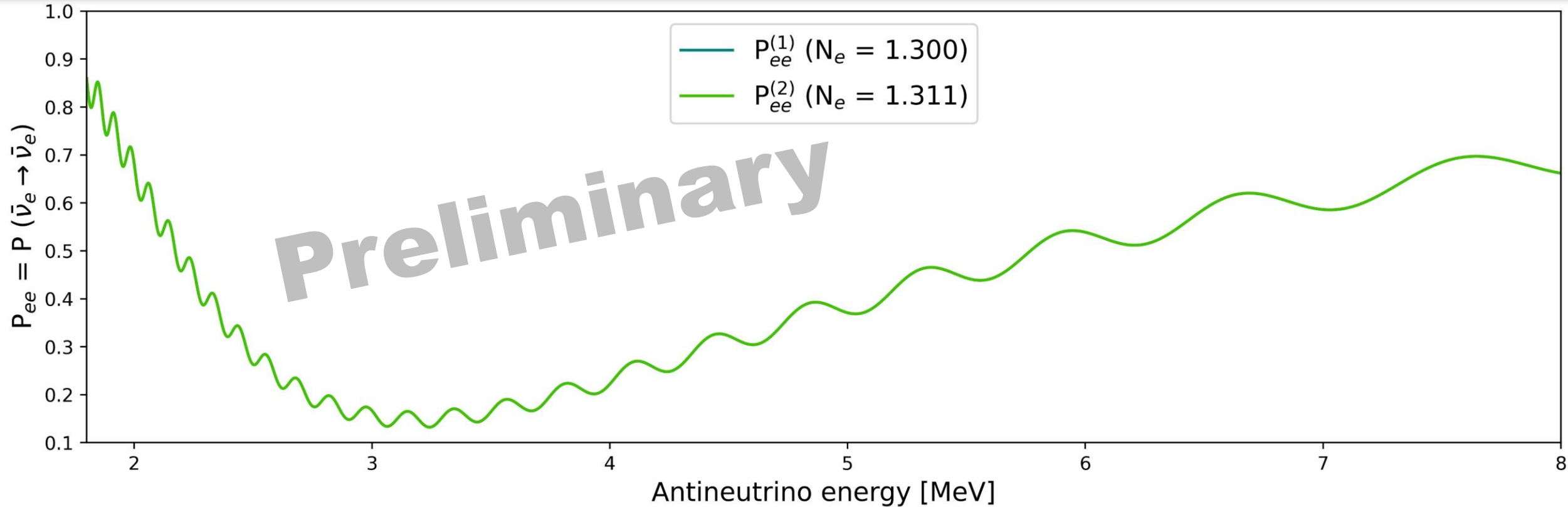


Electron density from geophysical and geochemical data

	Composition	ρ [g/cm ³]	N_e [mol/cm ³]
Capozzi et al., 2014	Not specified	2.6	1.3
Li et al., 2016	Not specified	2.6	1.3
Khan et al., 2020	Not specified	2.6	1.3
Our study	Reference rock	2.640 ± 0.038	1.311

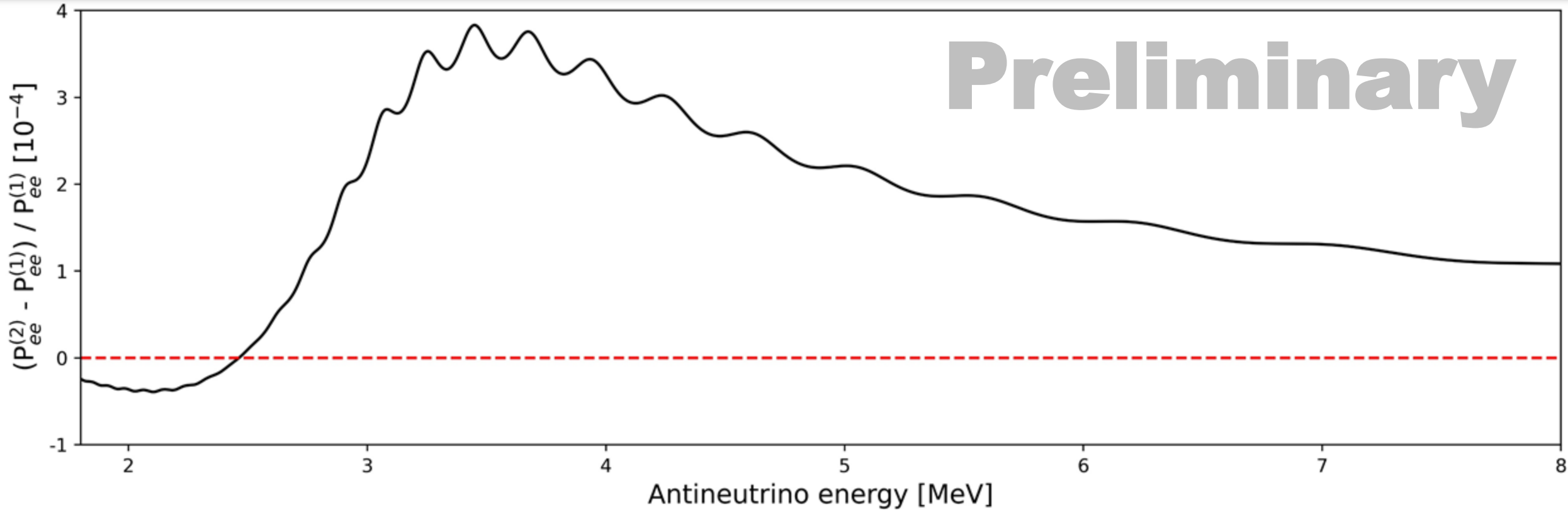
- Our estimations are based on geophysical and geochemical analysis of local rock samples
- Our density estimation differs from literature by 1.5%, **comparable with standard deviation**
- Estimation on N_e differs by 0.8% from literature
- With the complete measurement dataset, we will estimate the **uncertainty on N_e**

N_e and $\bar{\nu}_e$ survival probability in JUNO



- $\bar{\nu}_e$ survival probability as a function of antineutrino energy has been calculated for a **baseline $L = 52.5$ km** with **mixing parameters from ParticleDataGroup 2021**
- At this scale the effects of a different electron density are **not visible**

Impact of a precise N_e estimation



- The **relative difference** between $P_{ee}^{(2)}$ (obtained with $N_e = 1.311 \text{ mol/cm}^3$) and $P_{ee}^{(1)}$ (obtained with $N_e = 1.300 \text{ mol/cm}^3$) is plotted as a function of antineutrino energy
- The maximum relative difference is of order of $4 \cdot 10^{-4}$, at $\sim 3.5 \text{ MeV}$
- The **average** relative difference is $\sim 1.7 \cdot 10^{-4}$
- The effect of N_e obtained from geochemical and geophysical local data is small, but **should be considered for precise spectral calculation**

Conclusions and next steps

- We've made **density and compositional measurements** on 18 rock samples from JUNO site
- From measurement results we've estimated an **electron density of 1.311 mol/cm³**, to be compared with the value 1.300 mol/cm³ from literature
- This estimation modifies the $\bar{\nu}_e$ survival probability **at the order of 10⁻⁴**
- Next steps are the **evaluation of N_e uncertainties** with a full dataset, the **characterization of rock and sediment types** crossed by antineutrinos and the **reconstruction of reactor $\bar{\nu}$ spectrum** in JUNO

Thank you