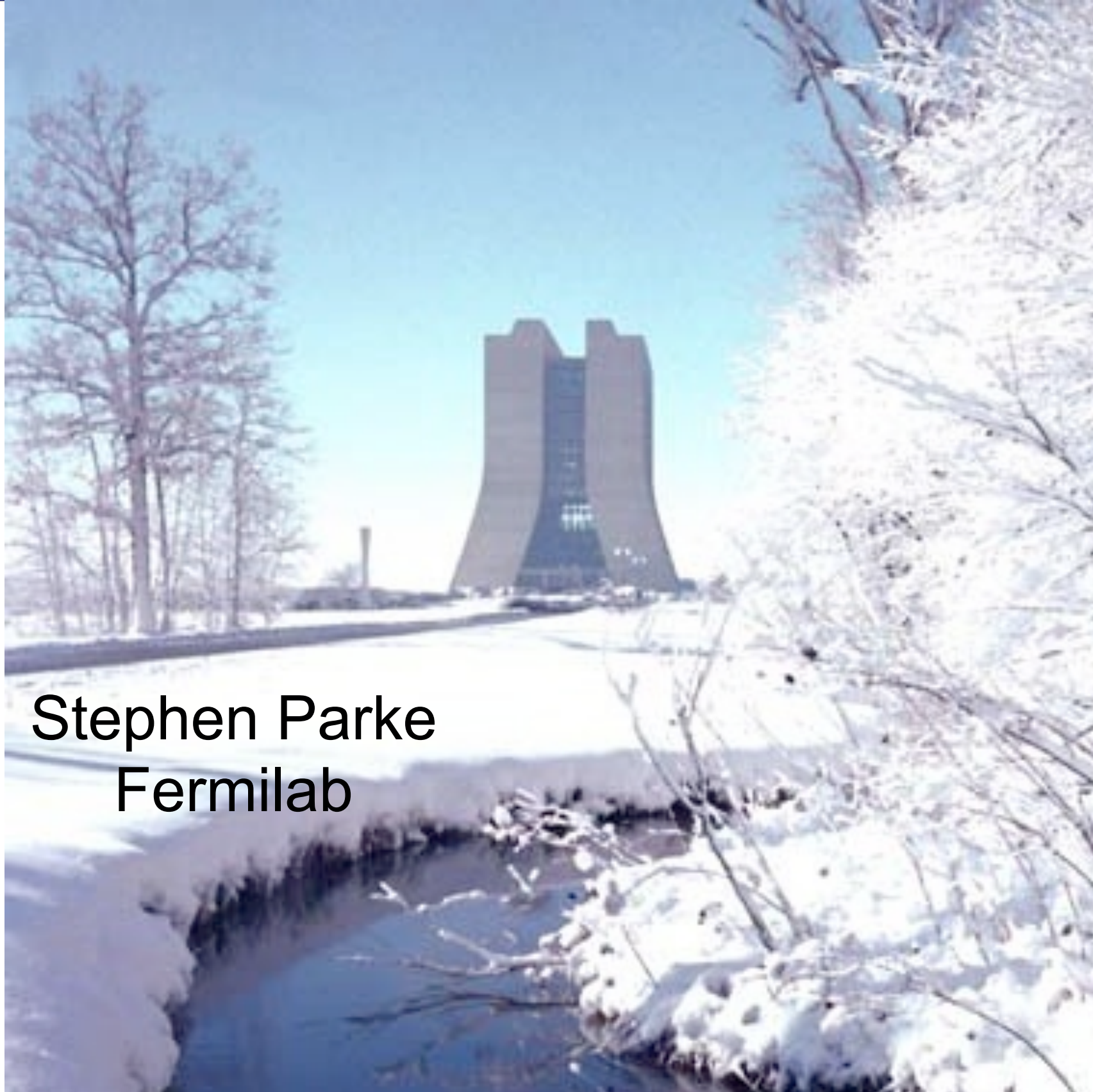
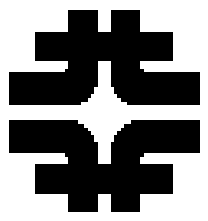


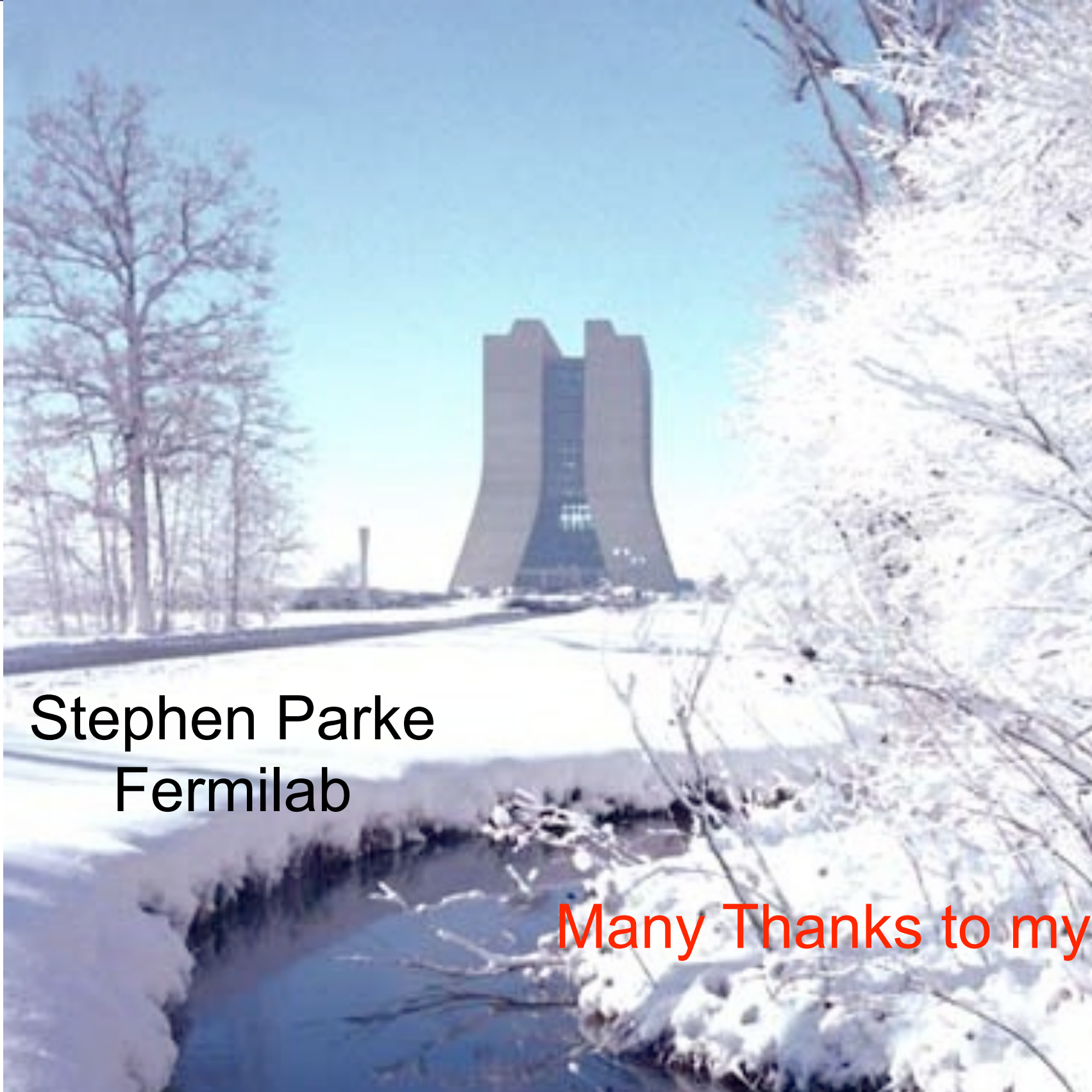
Neutrino SuperBeams at Fermilab



Stephen Parke
Fermilab



Neutrino SuperBeams at Fermilab



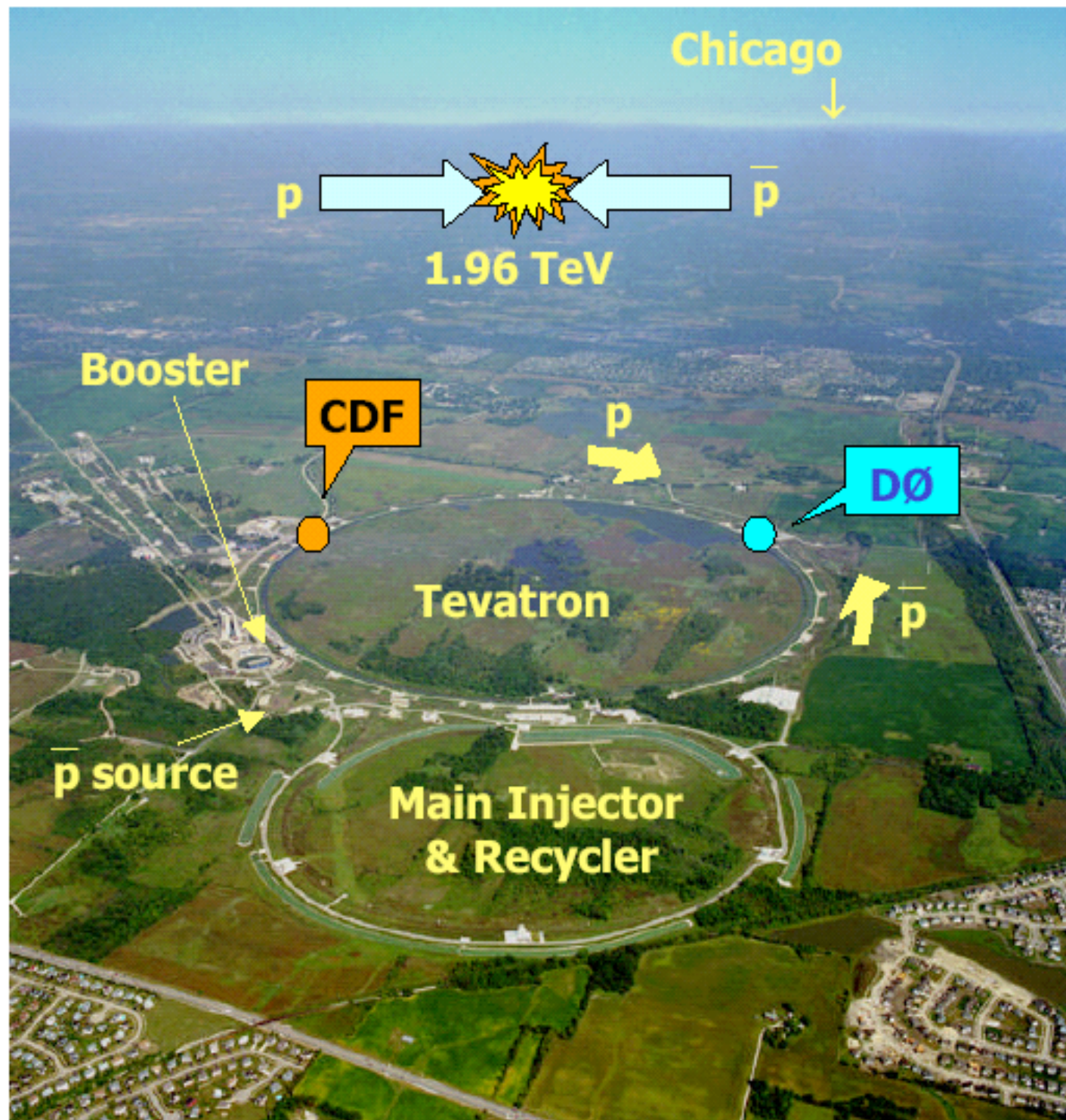
Stephen Parke
Fermilab

Many Thanks to my Colleagues



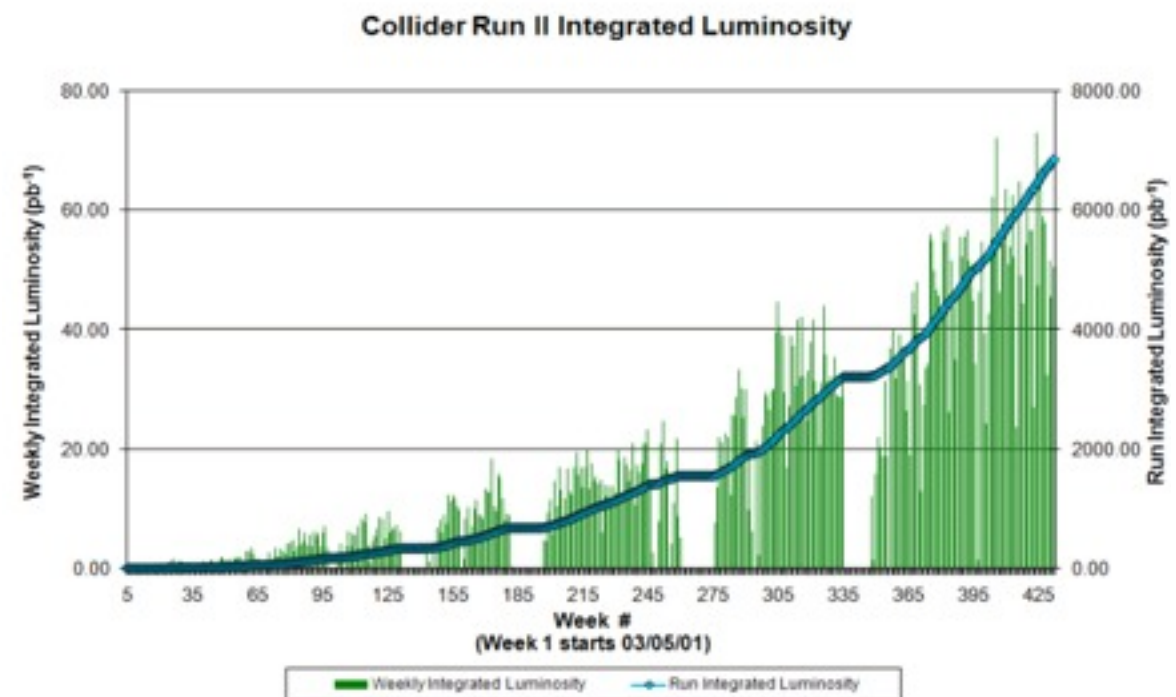
v's

Tevatron: CDF & D0



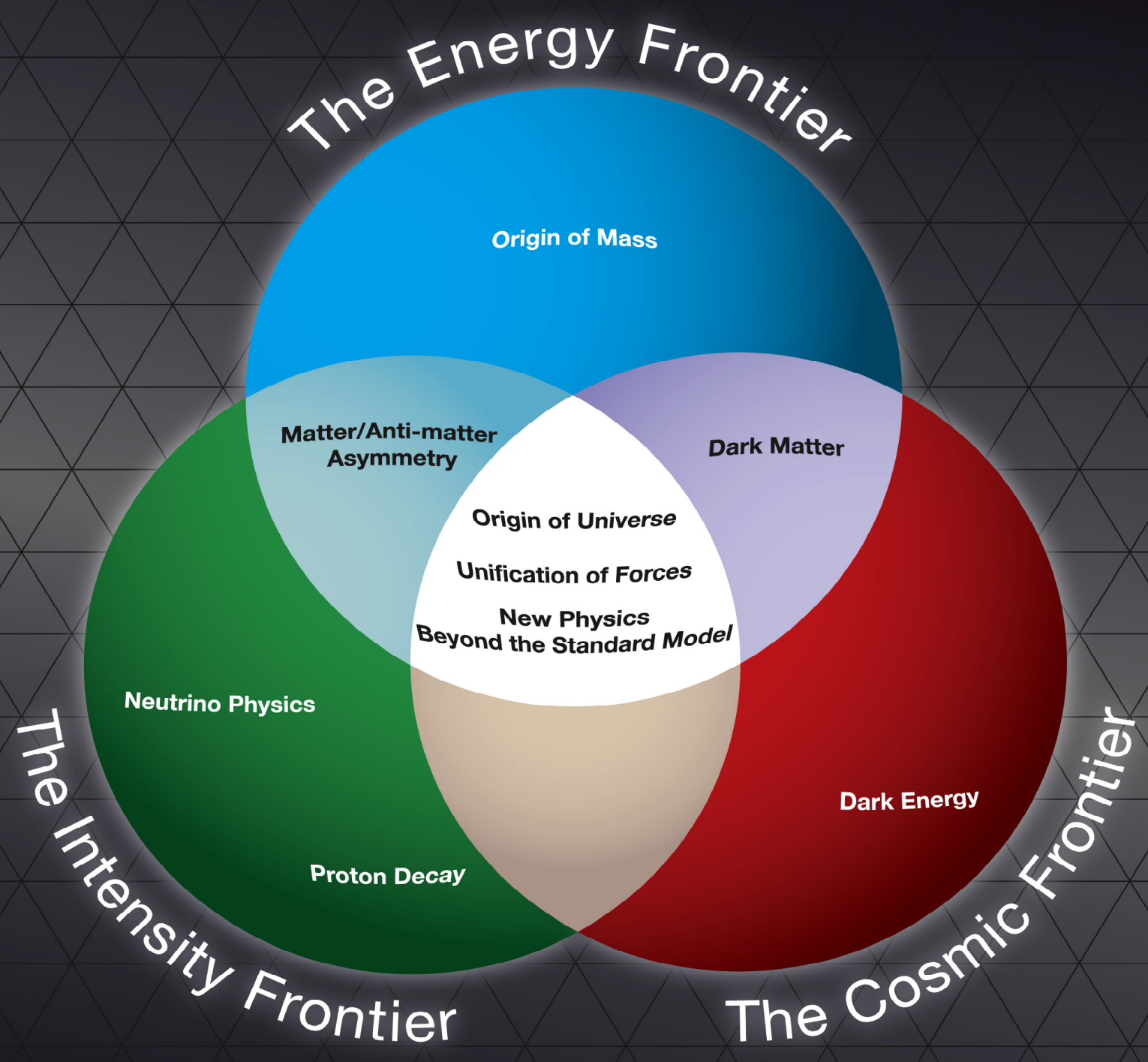
- 36x36 bunches
- bunch crossing 396 ns
- Run II started in March 2001
- Peak Luminosity: $3.5 \times 10^{32} \text{ cm}^{-2} \text{ sec}^{-1}$
- Run II delivered $\sim 10 \text{ fb}^{-1}$

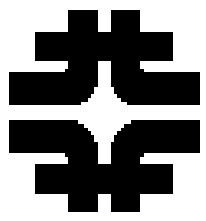
Peak Intergrated Luminosity:
 $3.5 \times 10^{32} \text{ cm}^{-2} \text{ sec}^{-1} \times 3 \times 10^7 \text{ sec} \approx 10 \text{ fb}^{-1} / \text{yr}$



Peak \bar{p} Collection Rate:
 $7 \times 10^7 / \text{sec}$

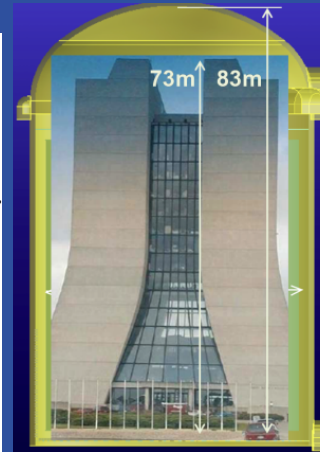
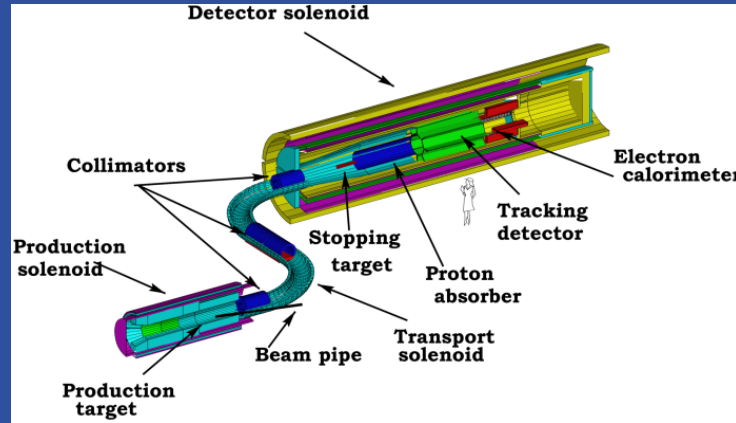
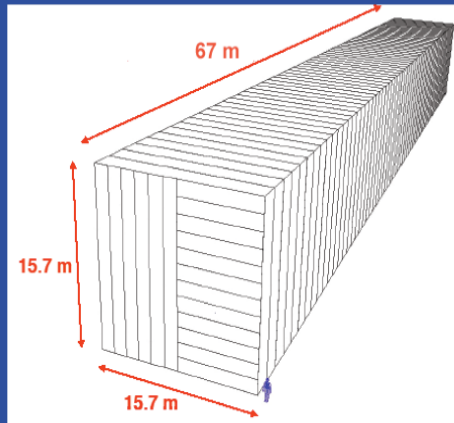
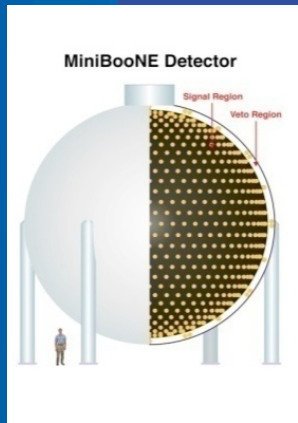
Peak \bar{p} Burn Rate:
 $3.5 \times 10^{32} \text{ cm}^{-2} \text{ sec}^{-1} \times 100 \text{ mb} = 3.5 \times 10^7 / \text{sec}$





Intensity Frontier:

Fermilab Intensity Frontier Experiments



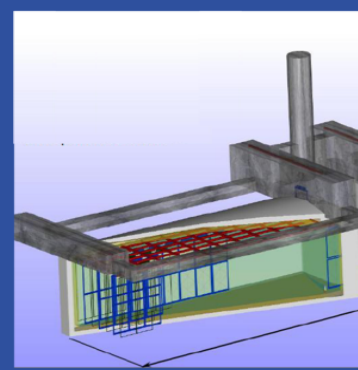
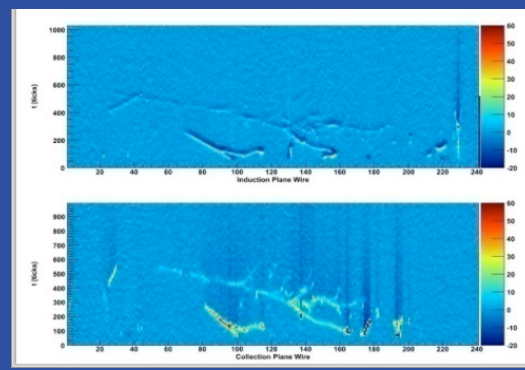
- ★ MINOS
- ★ MiniBooNE
- ★ MINERvA
- SeaQuest

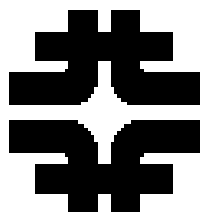
- ★ NOvA
- ★ MicroBooNE
- g-2?
- SeaQuest

- ★ LBNE
- Mu2e

- Project X+LBNE ★
- mu, K, nuclear, ...
- V Factory ??

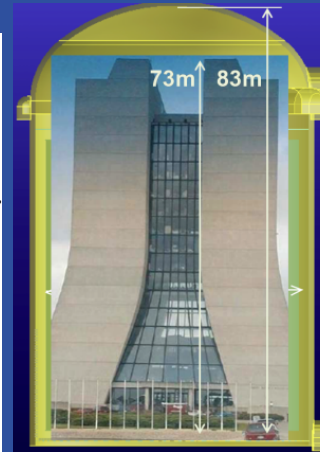
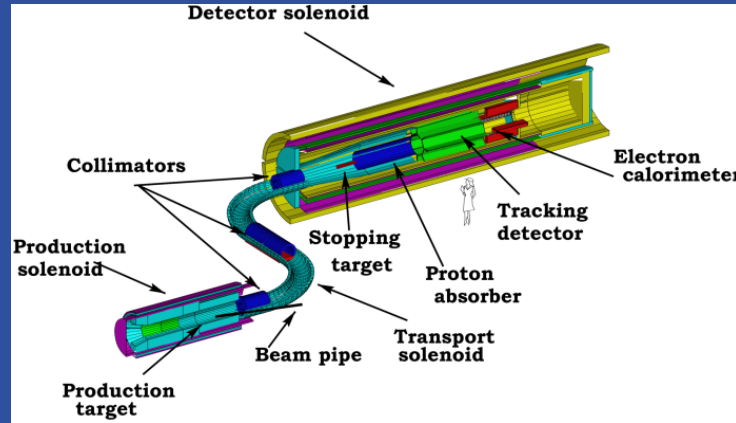
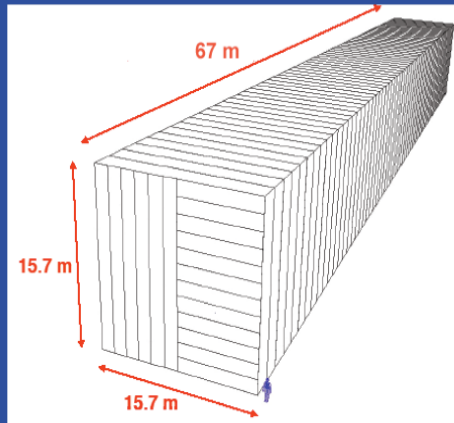
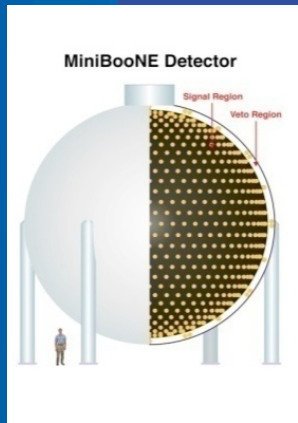
Now 2013 2016 2019 2022





Intensity Frontier:

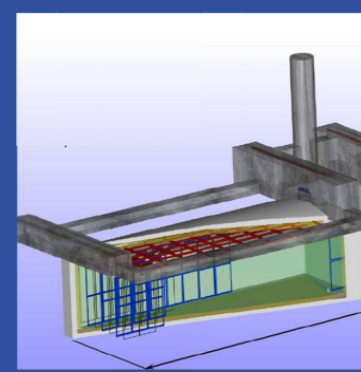
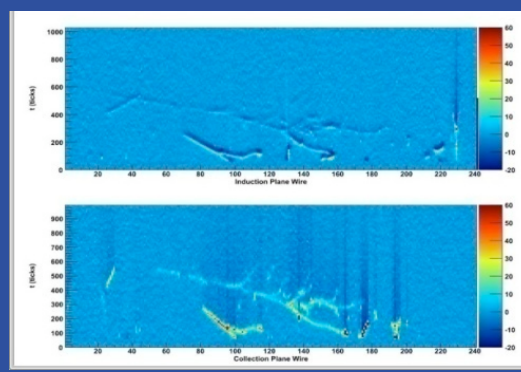
Fermilab Intensity Frontier Experiments



<ul style="list-style-type: none"> ★ MINOS ★ MiniBooNE ★ MINERvA SeaQuest 	<ul style="list-style-type: none"> ★ NOvA ★ MicroBooNE g-2? SeaQuest 	<ul style="list-style-type: none"> ★ LBNE Mu2e 	<ul style="list-style-type: none"> Project X+LBNE ★ mu, K, nuclear, ... √ Factory ??
---	--	--	---

Maybe back to Energy Frontier with Muon Collider

Now 2013 2016 2019 2022



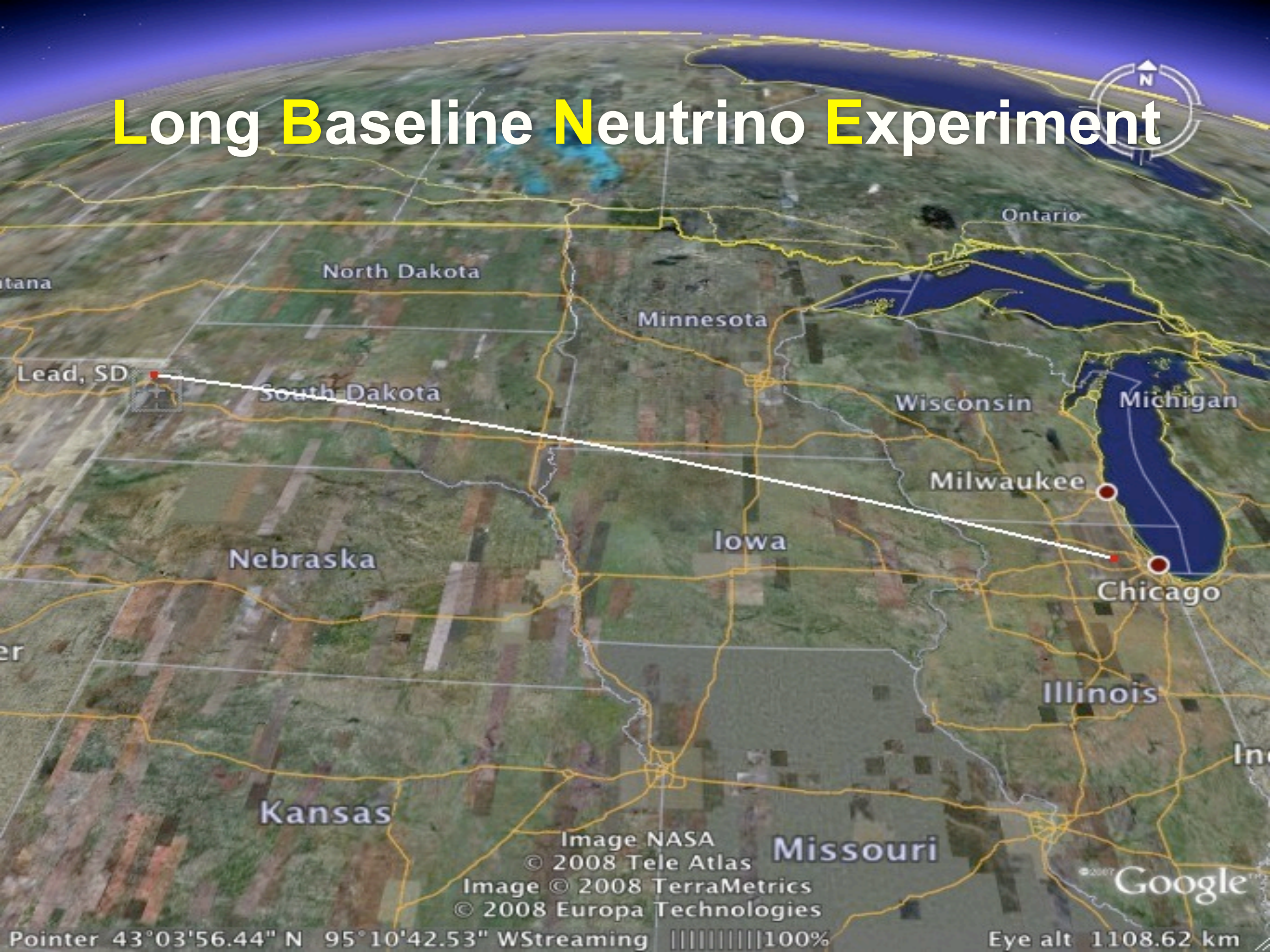
Long Baseline Neutrino Experiment



DUSEL
Deep Underground
Science and
Engineering Laboratory



Long Baseline Neutrino Experiment



Lead, SD

South Dakota

North Dakota

Minnesota

Ontario

Wisconsin

Michigan

Milwaukee

Iowa

Nebraska

Chicago

Illinois

Kansas

Missouri

Image NASA

© 2008 Tele Atlas

Image © 2008 TerraMetrics

© 2008 Europa Technologies

Google

Pointer 43°03'56.44" N 95°10'42.53" W Streaming ||||| 100%

Eye alt 1108.62 km

Long Baseline Neutrino Experiment



Lead, SD

New Neutrino Beam at Fermilab...



Image NASA

© 2008 Tele Atlas

Image © 2008 TerraMetrics

© 2008 Europa Technologies

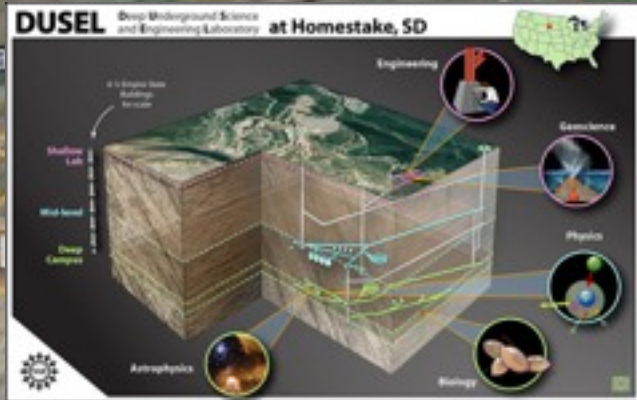
Missouri

© 2007 Google

Pointer 43°03'56.44" N 95°10'42.53" W Streaming ||||| 100%

Eye alt 1108.62 km

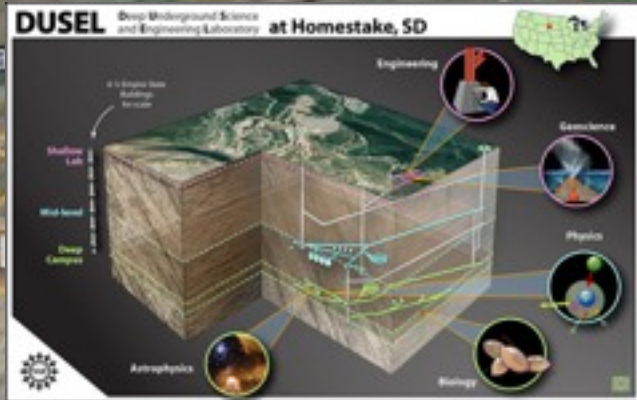
Long Baseline Neutrino Experiment



New Neutrino Beam at Fermilab...
...Directed towards NSF's proposed DUSEL



Long Baseline Neutrino Experiment



New Neutrino Beam at Fermilab...
...Directed towards NSF's proposed DUSEL
Precision Near Detector on the Fermilab site



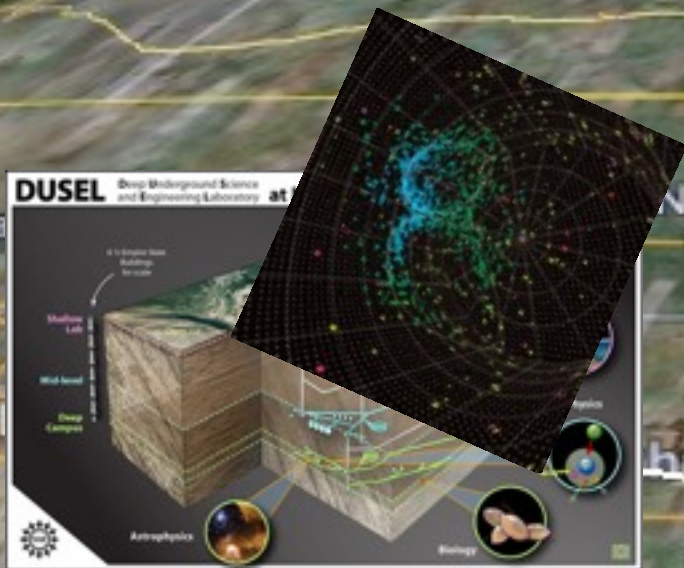
Image NASA
© 2008 Tele Atlas
Image © 2008 TerraMetrics
© 2008 Europa Technologies

Google

Pointer 43°03'56.44" N 95°10'42.53" W Streaming 100%

Eye alt 1108.62 km

Long Baseline Neutrino Experiment



**New Neutrino Beam at Fermilab...
...Directed towards NSF's proposed DUSEL
Precision Near Detector on the Fermilab site
100 kT Water Cherenkov Detector**



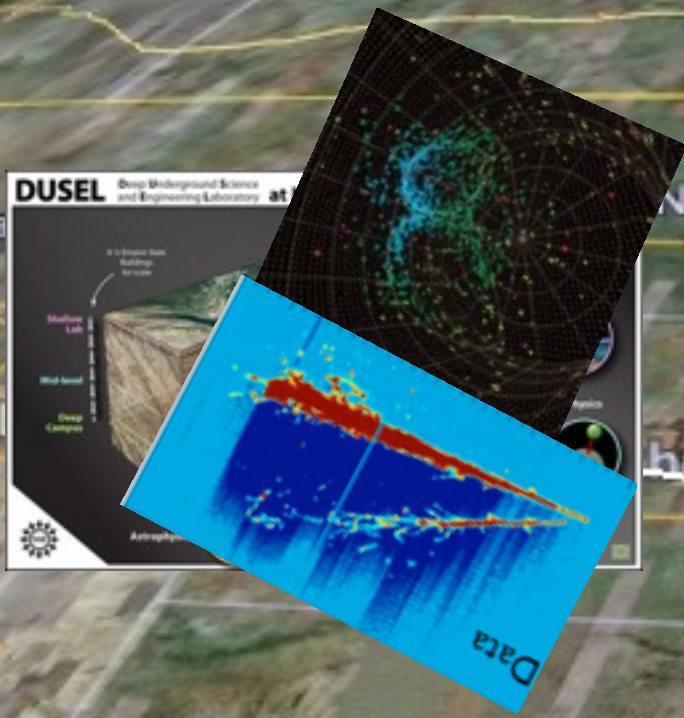
Image NASA
© 2008 Tele Atlas
Image © 2008 TerraMetrics
© 2008 Europa Technologies

Google

Pointer 43°03'56.44" N 95°10'42.53" W Streaming 100%

Eye alt 1108.62 km

Long Baseline Neutrino Experiment



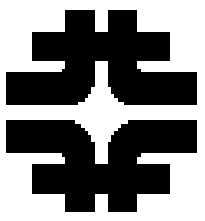
New Neutrino Beam at Fermilab...
...Directed towards NSF's proposed DUSEL
Precision Near Detector on the Fermilab site
100 kT Water Cherenkov Detector
20 kT Liquid Argon TPC Far Detector

Image NASA
© 2008 Tele Atlas
Image © 2008 TerraMetrics
© 2008 Europa Technologies

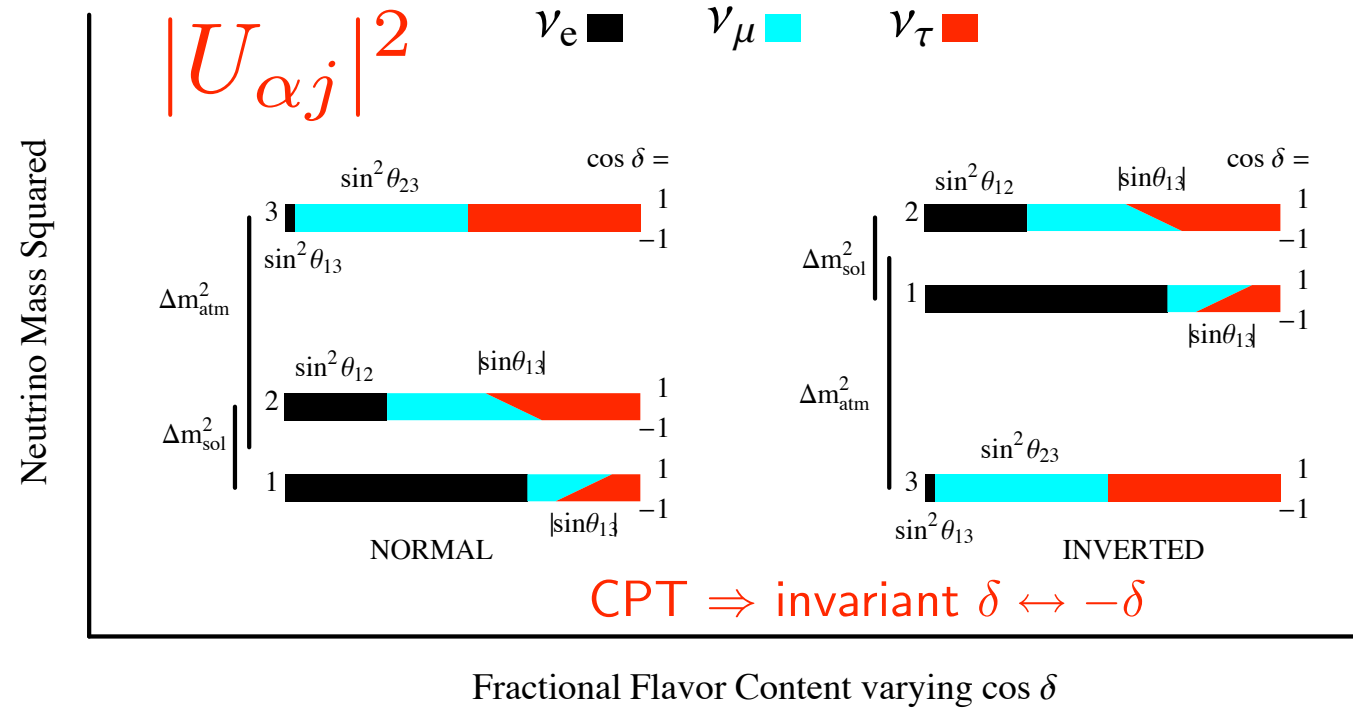
©2007 Google

Pointer 43°03'56.44" N 95°10'42.53" W Streaming ||||| 100%

Eye alt 1108.62 km



Masses & Mixings:



$$\delta m_{sol}^2 = +7.6 \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \theta_{12} \sim 1/3$$

$$|\delta m_{atm}^2| = 2.4 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} \sim 1/2$$

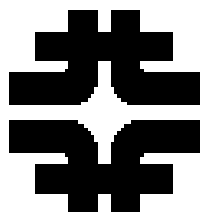
$$|\delta m_{sol}^2| / |\delta m_{atm}^2| \approx 0.03$$

$$\sin^2 \theta_{13} < 3\%$$

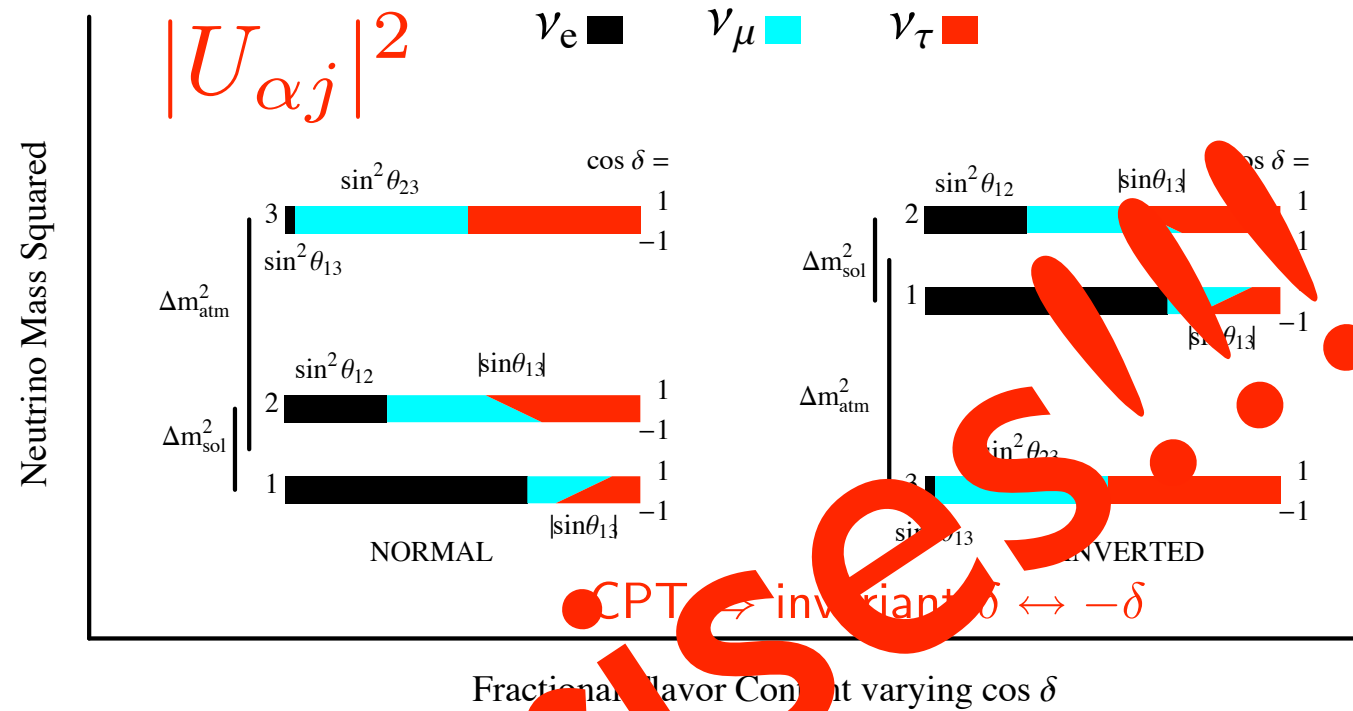
$$\sqrt{\delta m_{atm}^2} = 0.05 \text{ eV} < \sum m_{\nu_i} < 0.5 \text{ eV} = 10^{-6} * m_e$$

$$0 \leq \delta < 2\pi$$

$$\sin^2 \theta_{13} \equiv |U_{e3}|^2, \quad \sin^2 \theta_{12} \equiv \frac{|U_{e2}|^2}{(1 - |U_{e3}|^2)}, \quad \sin^2 \theta_{23} \equiv \frac{|U_{\mu 3}|^2}{(1 - |U_{e3}|^2)}$$



Masses & Mixings:



$$\delta m_{sol}^2 = +7.6 \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \theta_{12} \sim 1/3$$

$$|\delta m_{atm}^2| = 2.4 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} \sim 1/2$$

$$|\delta m_{sol}^2| / |\delta m_{atm}^2| \approx 0.03$$

$$\sin^2 \theta_{13} < 3\%$$

$$\sqrt{\delta m_{atm}^2} = 0.05 \text{ eV} < \sum m_{\nu_i} < 0.5 \text{ eV} = 10^{-6} * m_e$$

$$0 \leq \delta < 2\pi$$

$$\sin^2 \theta_{13} \equiv |U_{e3}|^2, \quad \sin^2 \theta_{12} \equiv \frac{|U_{e2}|^2}{(1 - |U_{e3}|^2)}, \quad \sin^2 \theta_{23} \equiv \frac{|U_{\mu 3}|^2}{(1 - |U_{e3}|^2)}$$

In Matter:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

$$\Delta = \frac{\delta m^2 L}{4E}$$

In Matter:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

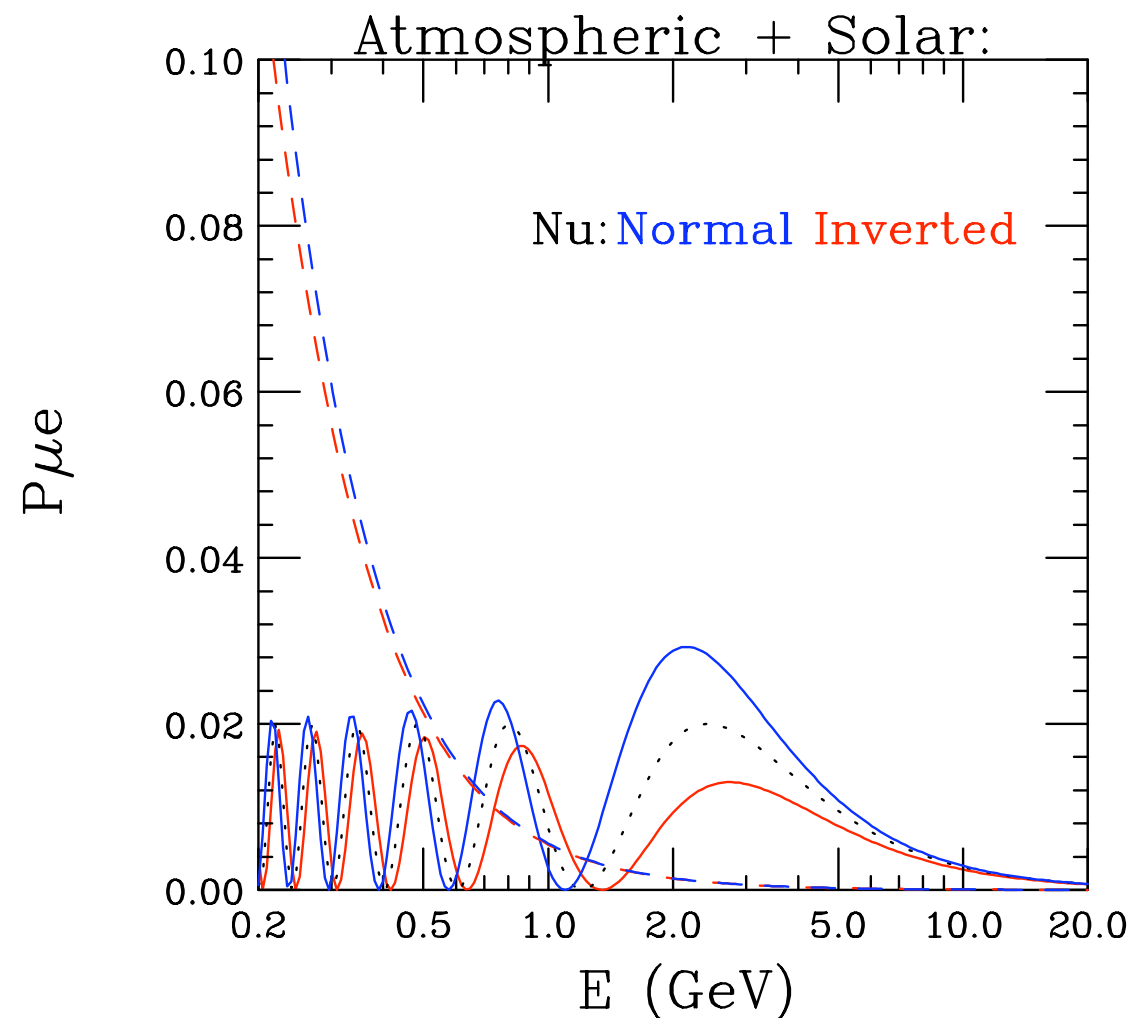
where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$

$$\Delta = \frac{\delta m^2 L}{4E}$$



In Matter:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

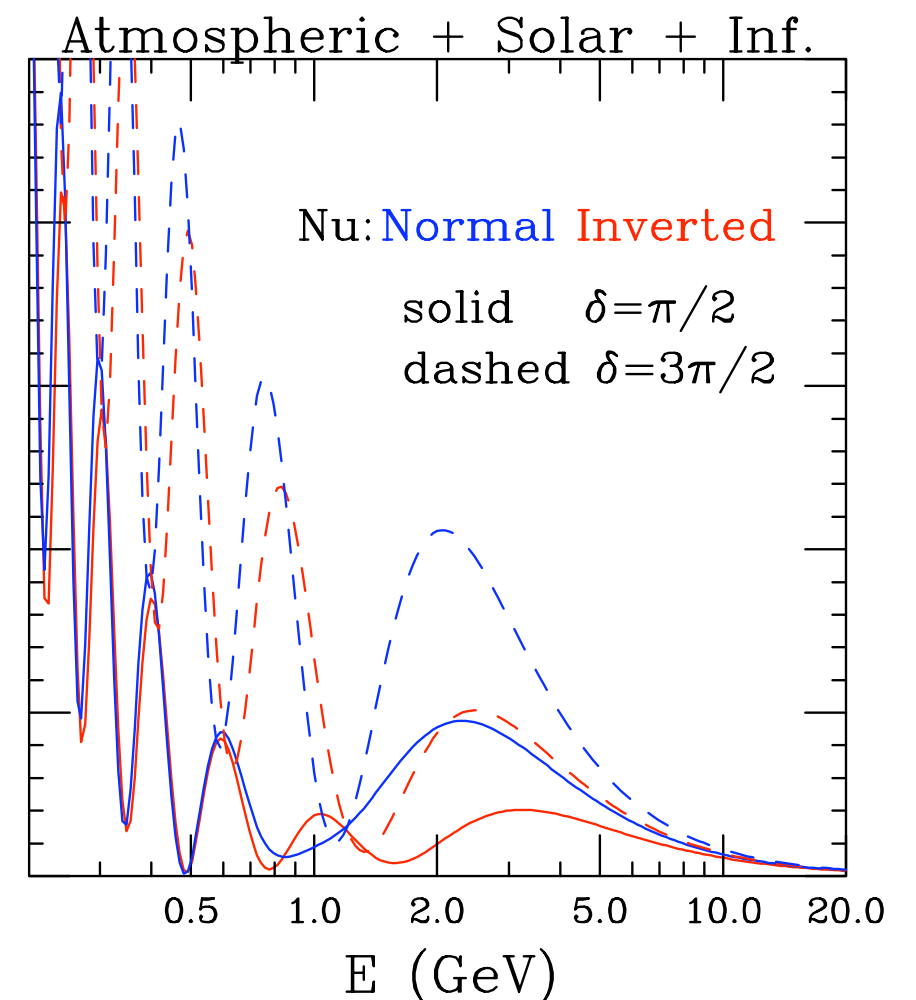
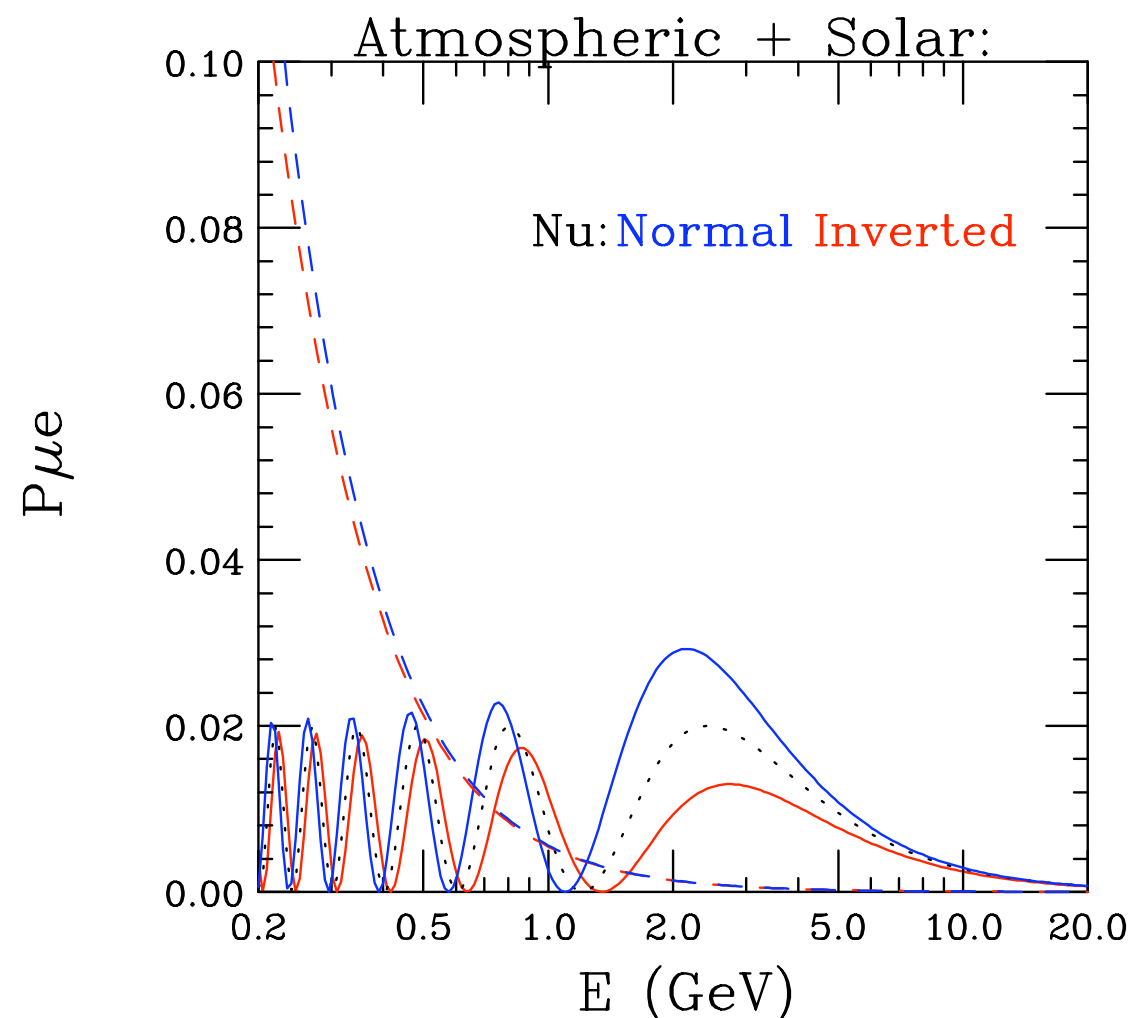
and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

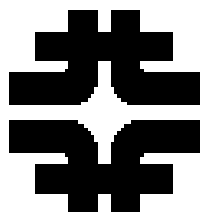
$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

For $L = 1200 \text{ km}$
and $\sin^2 2\theta_{13} = 0.04$

$$\Delta = \frac{\delta m^2 L}{4E}$$

Anti-Nu: Normal Inverted
dashes $\delta = \pi/2$
solid $\delta = 3\pi/2$

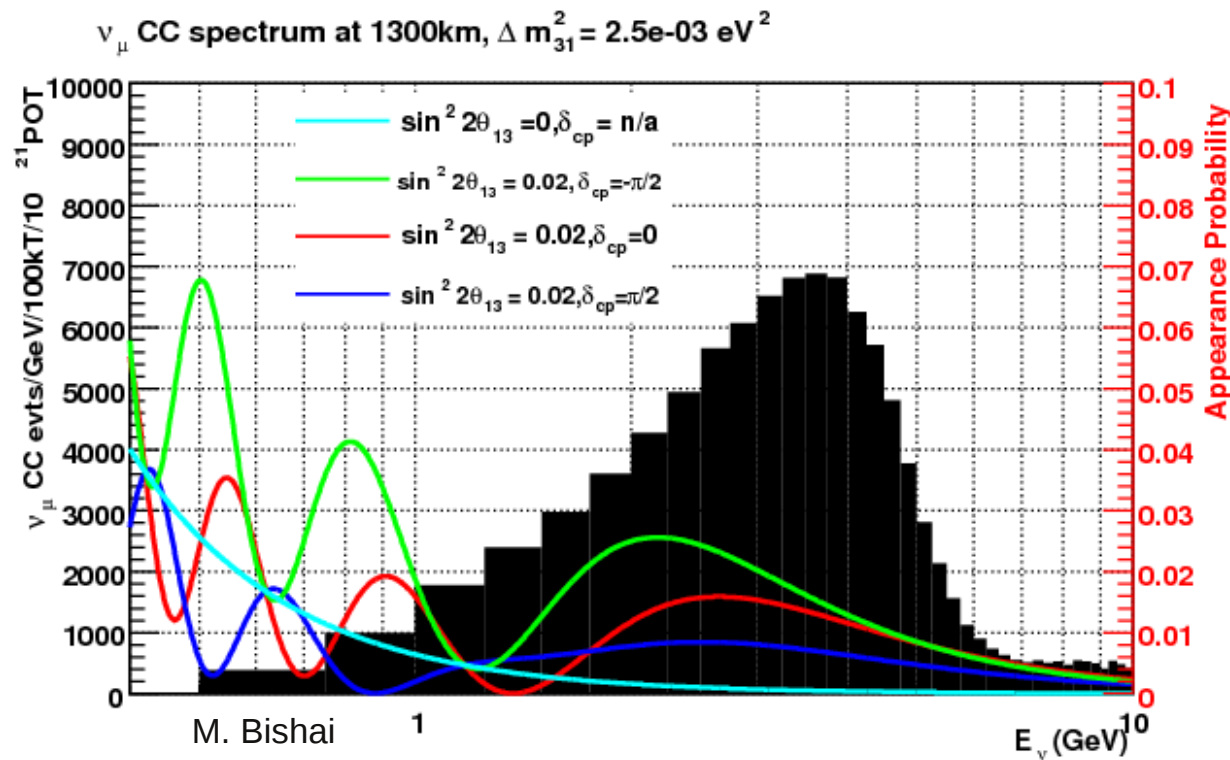




Neutrino Beam:

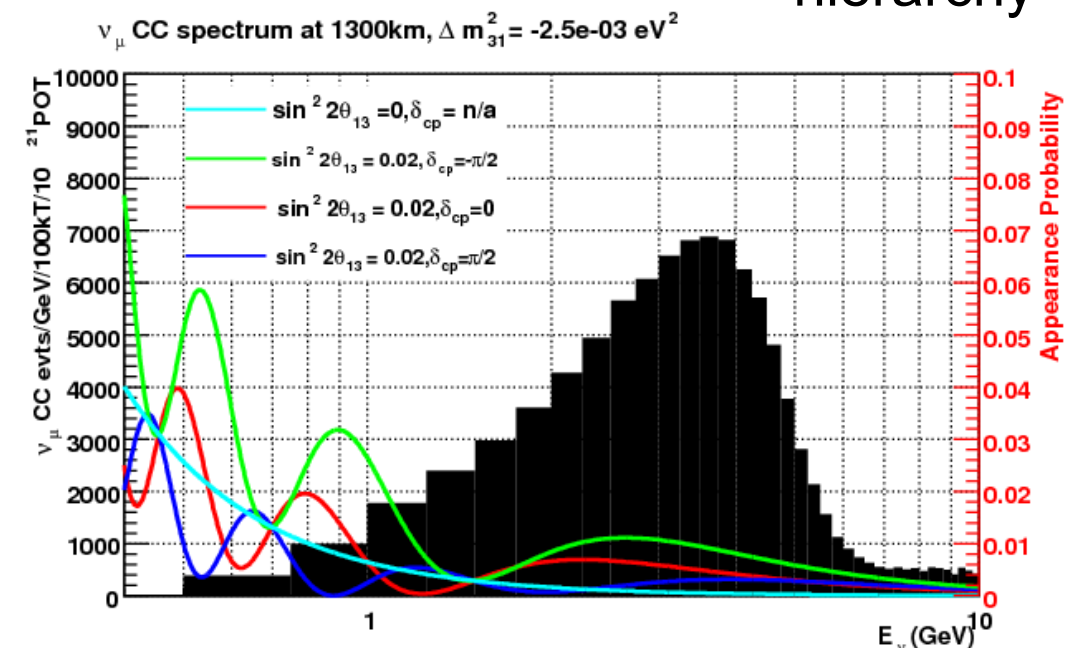
Beam requirements (driven by ν_e appearance):

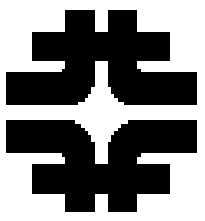
- wide band beam to cover both 1st and 2nd oscillation maxima
- minimize flux above 5 GeV to avoid feed down of NC background to low energy
- minimize beam ν_e contamination



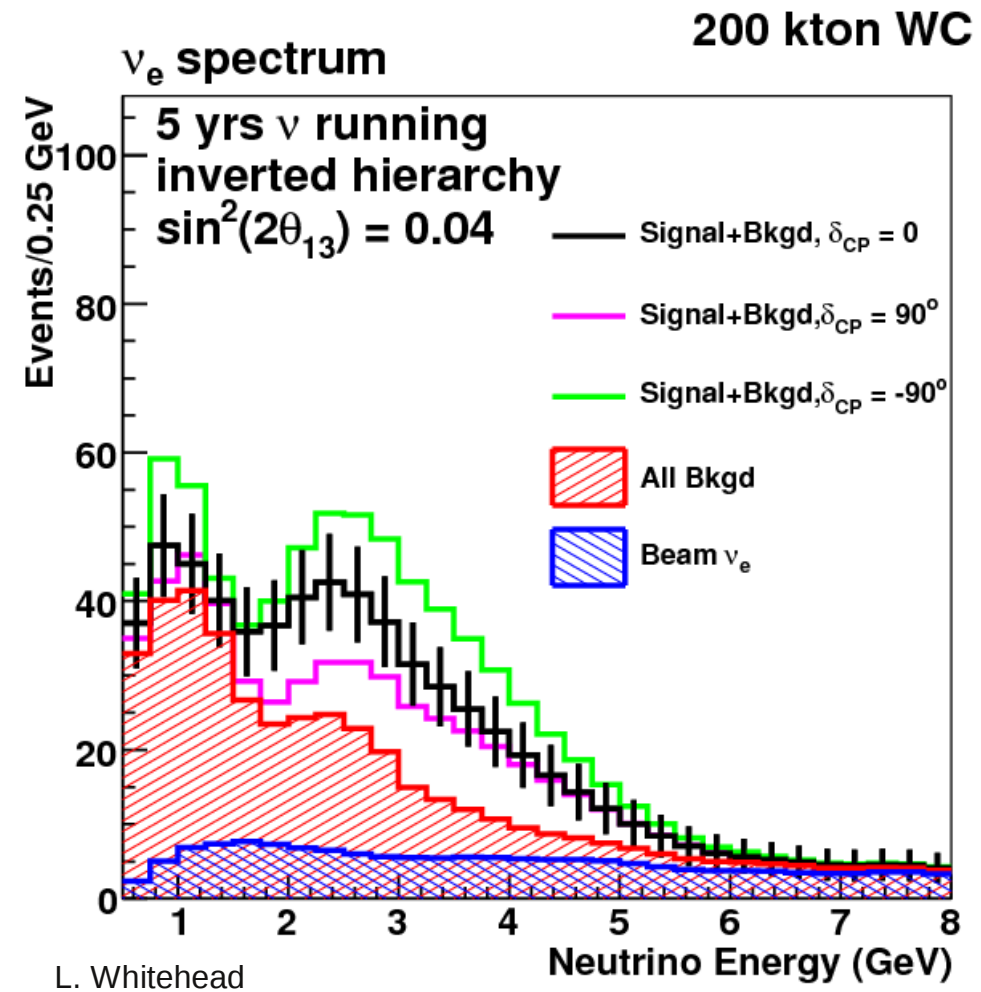
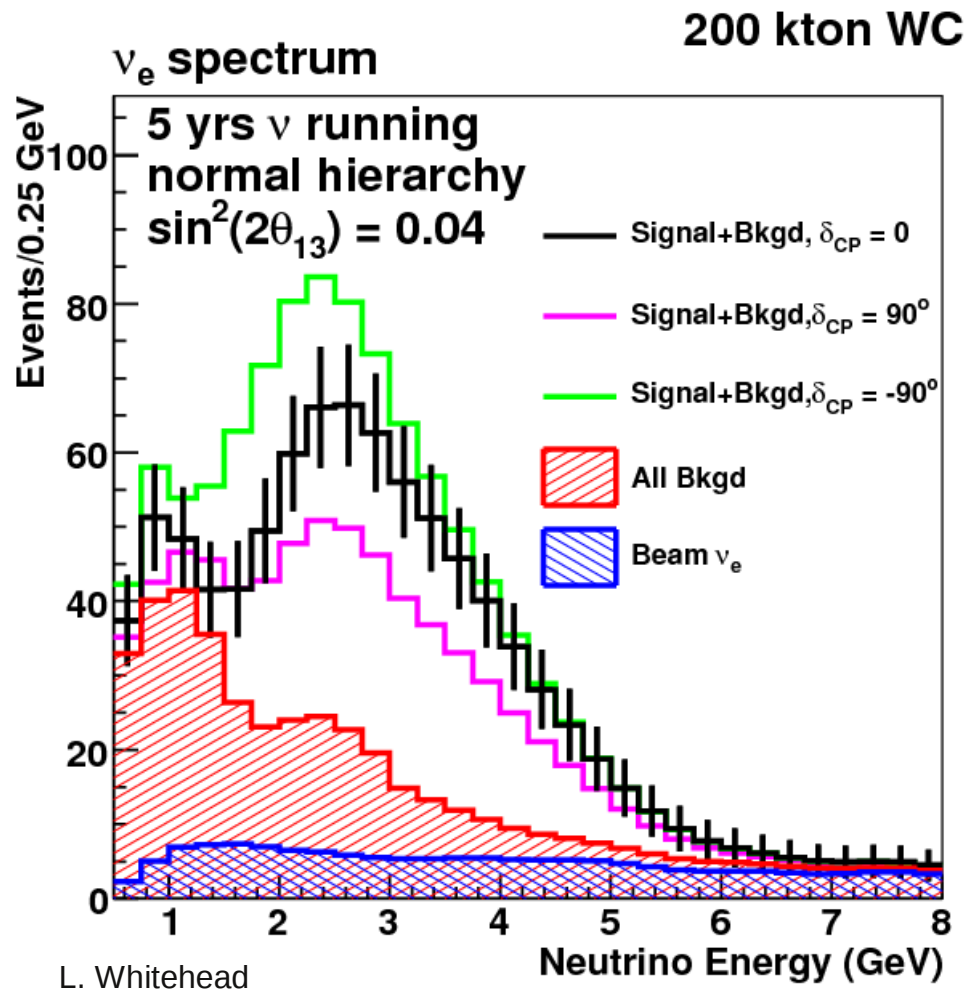
Normal hierarchy

Inverted hierarchy

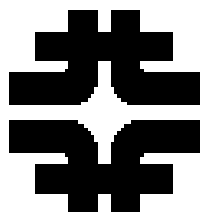




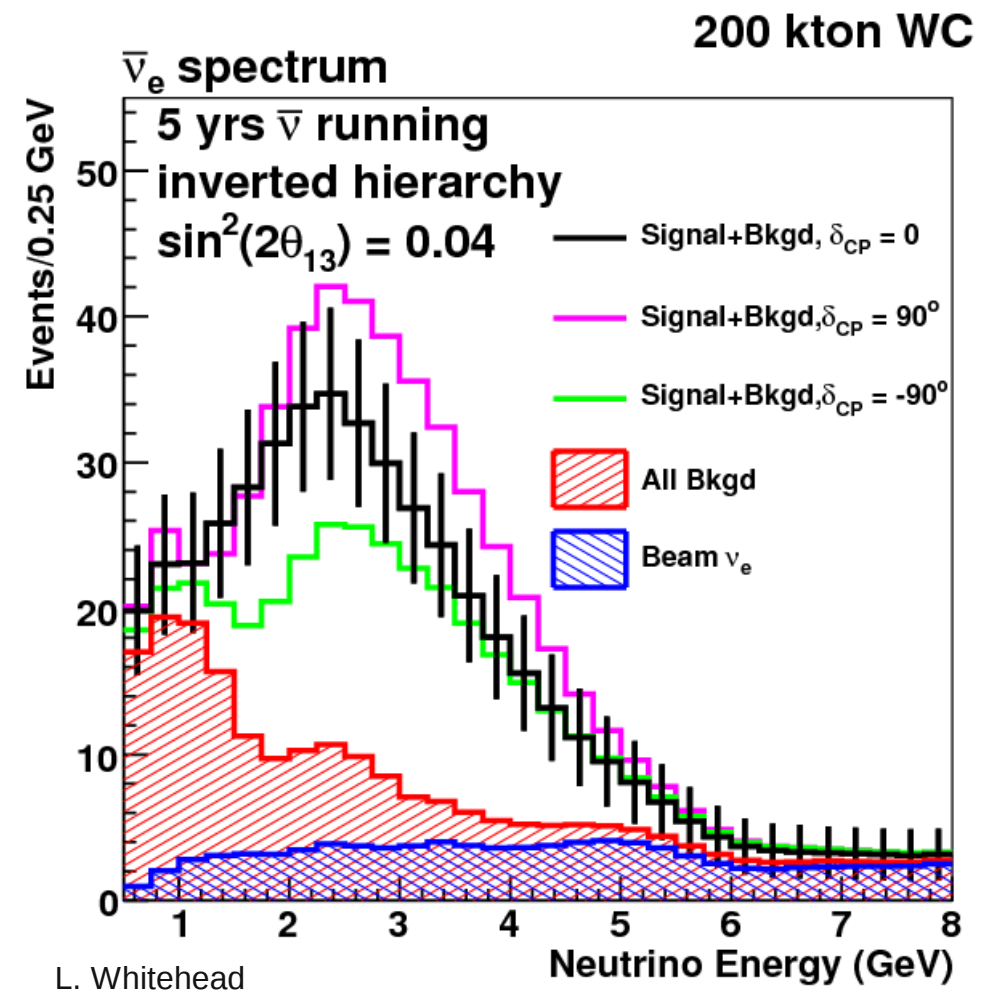
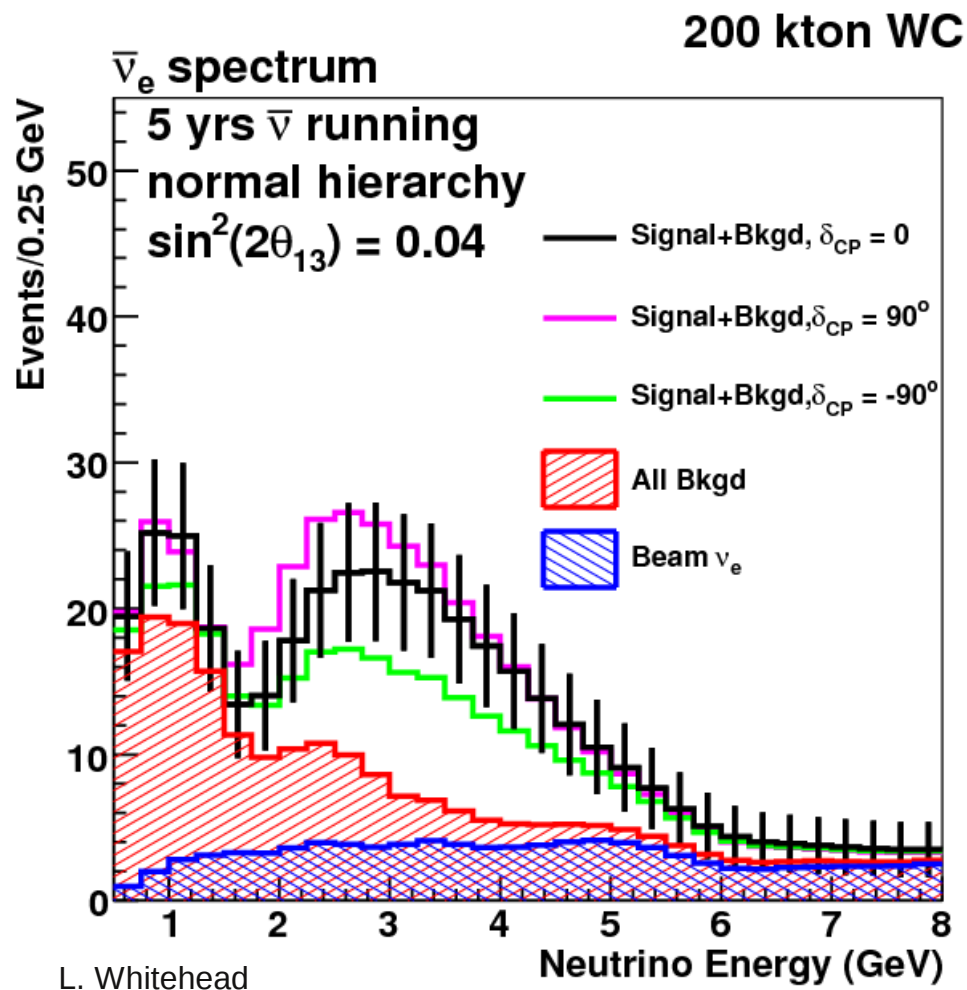
Neutrino



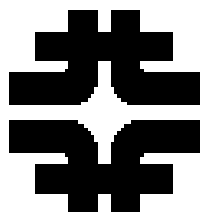
$\delta=0$	signal	beam ν_e	NC	$\nu\mu$ CC
Normal hier	484	218	276	15
Inverted hier	212	221	276	15



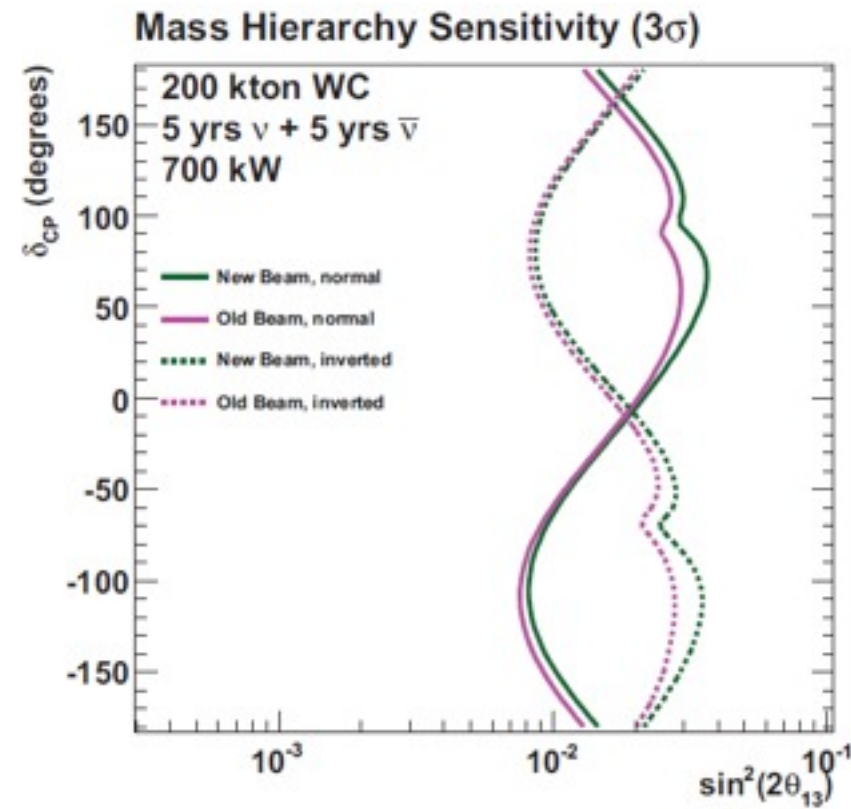
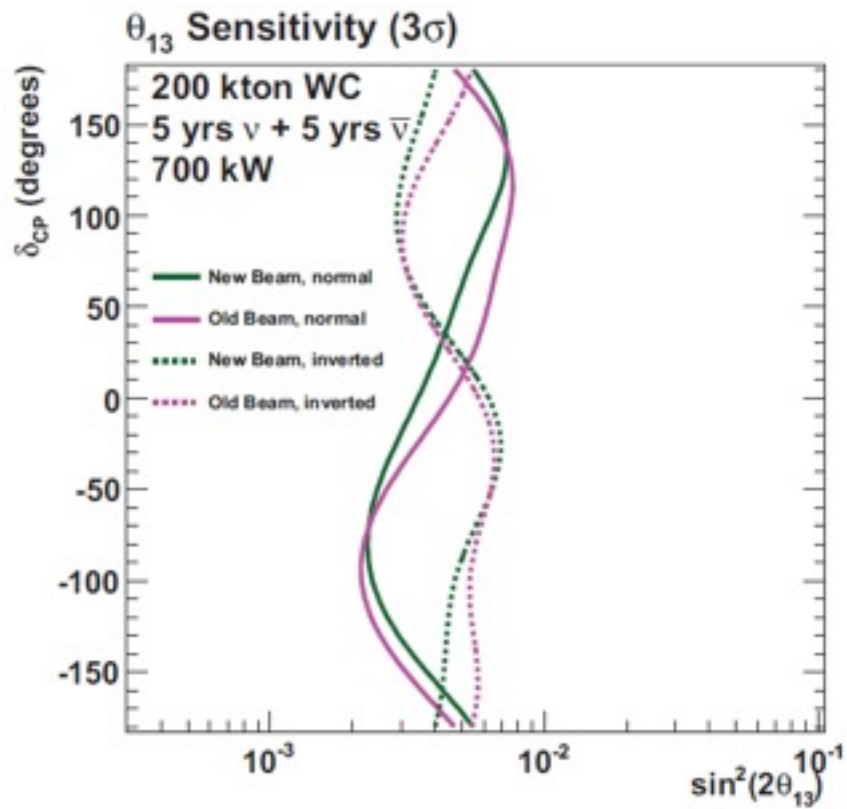
Anti-Neutrino



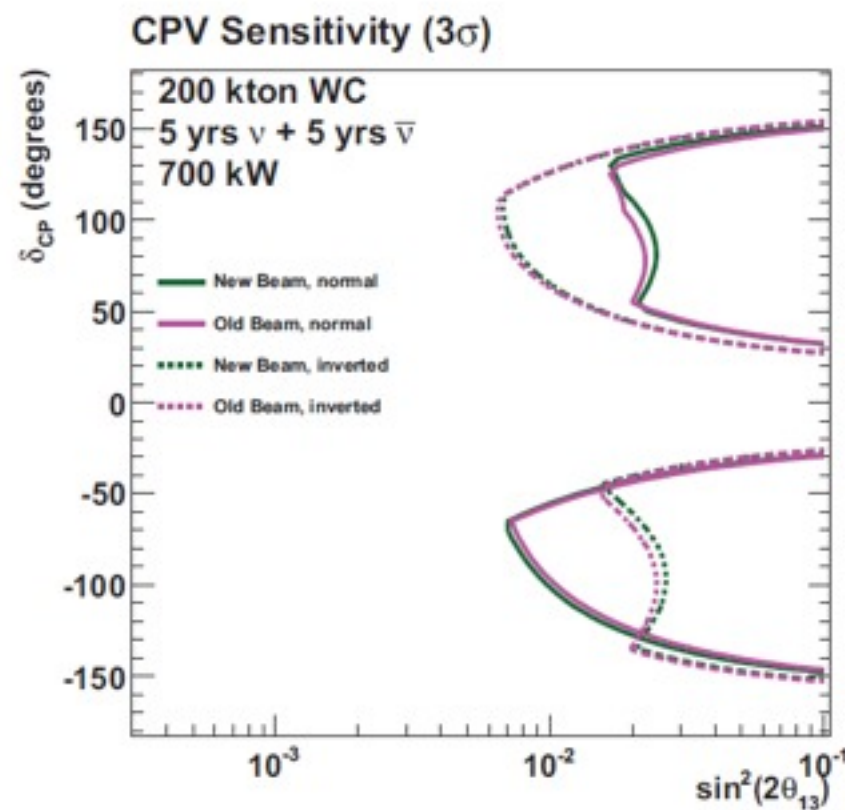
$\delta=0$	ν_e signal	$\bar{\nu}_e$ signal	beam ν_e	NC	ν_μ CC	Beam $\bar{\nu}_e$	$\bar{\nu}_\mu$ CC
Normal hier	35	145	88	118	2	27	5
Inverted hier	15	246	27	118	2	87	5



LBNE Sensitivities:

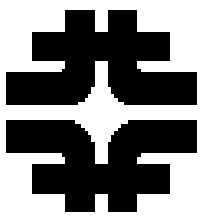


Sensitivities shown are for 200 ktons WCD.
With 34 ktons LAr, the sensitivities are equivalent.

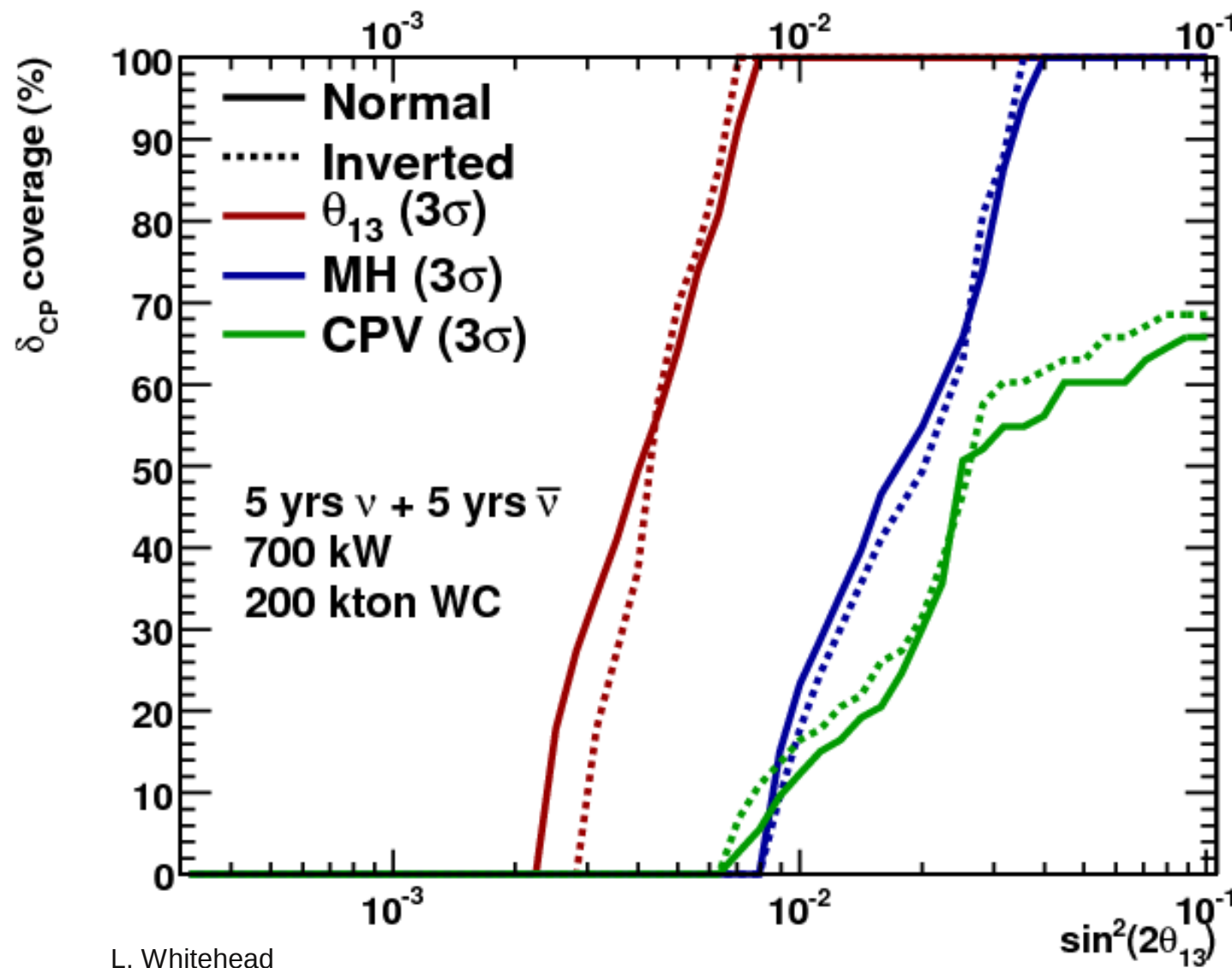


Sensitivities shown are for 10 years with 700 kW proton beam.
With Project X (≥ 2 MW), the same sensitivity would be reached in 1/3 the time ... or $\sim\sqrt{3}$ greater sensitivity for the same running period.

Fall 2010 Report from the LBNE Physics Working Group



Sensitivities:



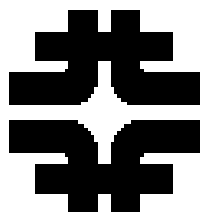
L. Whitehead

- 3σ non-zero θ_{13} for
100% of δ_{cp} values at
 $\sin^2 2\theta_{13} = 0.008$

- 3σ mass hier. for
100% of δ_{cp} values at
 $\sin^2 2\theta_{13} = 0.04$

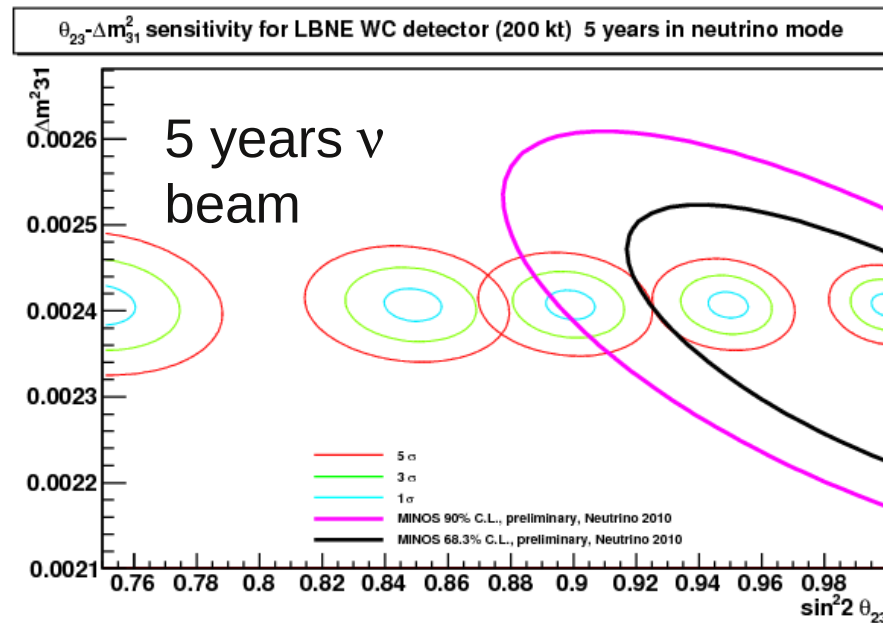
- 3σ CP violation for
50% of δ_{cp} values at
 $\sin^2 2\theta_{13} = 0.03$

(all assuming normal
hierarchy)



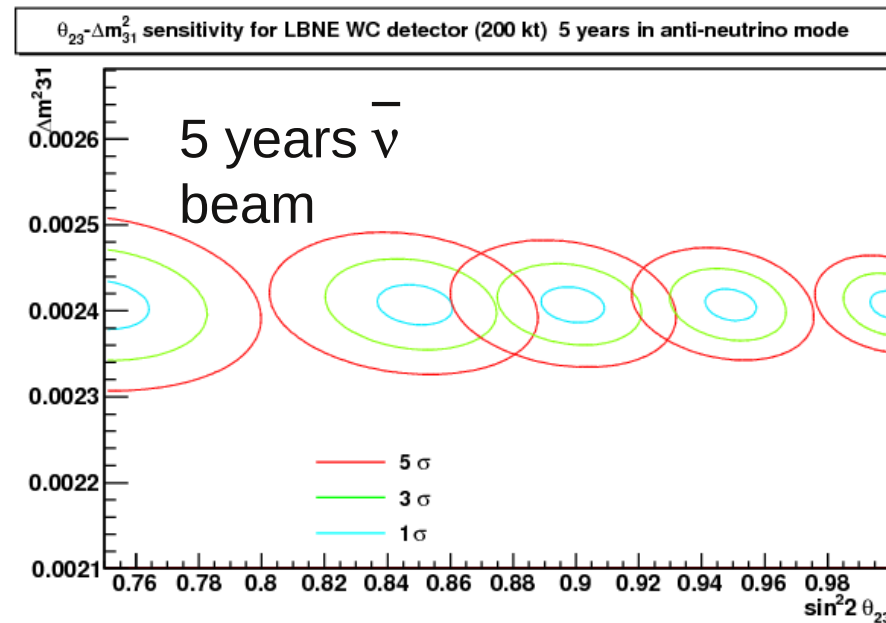
Disappearance

Δm^2 vs $\sin^2 2\theta_{23}$



R. Guenette

$\Delta \bar{m}^2$ vs $\sin^2 2\bar{\theta}_{23}$



R. Guenette

$$\begin{aligned} \delta m_{\mu\mu}^2 &= \frac{|U_{\mu 1}|^2 \delta m_{31}^2 + |U_{\mu 2}|^2 \delta m_{32}^2}{|U_{\mu 1}|^2 + |U_{\mu 2}|^2} \\ &= \delta m_{32}^2 + \\ &\quad [\sin^2 \theta_{12} + \mathcal{O}(\sin \theta_{13})] \delta m_{21}^2 \end{aligned}$$

$$|\delta m_{32}^2| = |\delta m_{\mu\mu}^2| (1 \mp 0.01)$$

**1% difference:
sign depends
on hierarchy**

- signal is CCQE events (97% efficiency)
- 5% normalization and 3% energy scale uncertainty for signal
- 10% normalization and 3% energy scale uncertainty for background

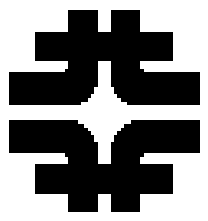
Which Octant?

At Vac. Osc. Max. ($\Delta_{31} = \frac{\pi}{2}$)

$$P(\nu_{\mu} \rightarrow \nu_e) + P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e) \approx 2 \sin^2 \theta_{23} \sin^2 2\theta_{13} + \mathcal{O}[(aL) \sin \delta]$$

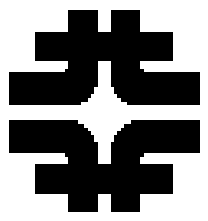
directly comparable to reactor

$$1 - P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = \sin^2 2\theta_{13}$$

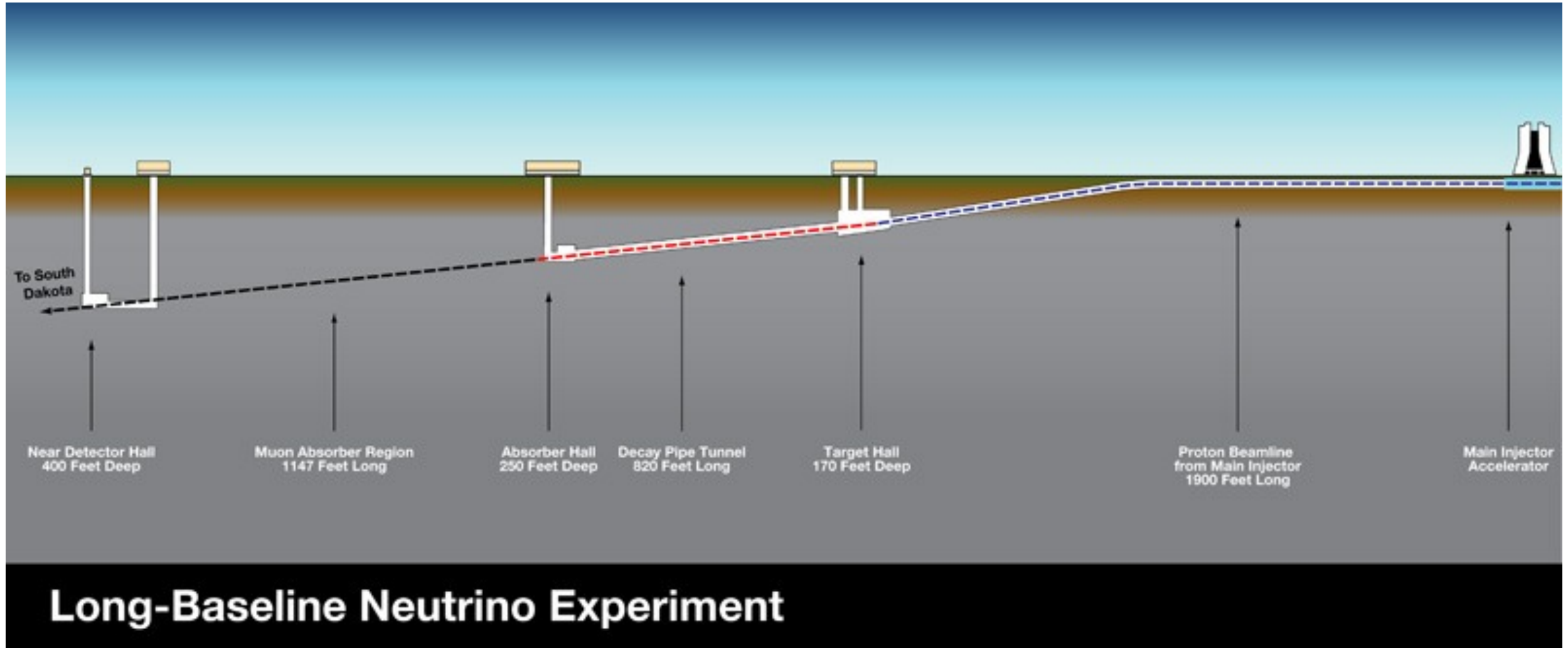


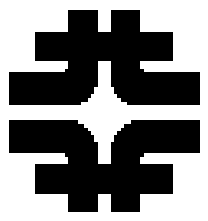
New Neutrino Beamline at Fermilab



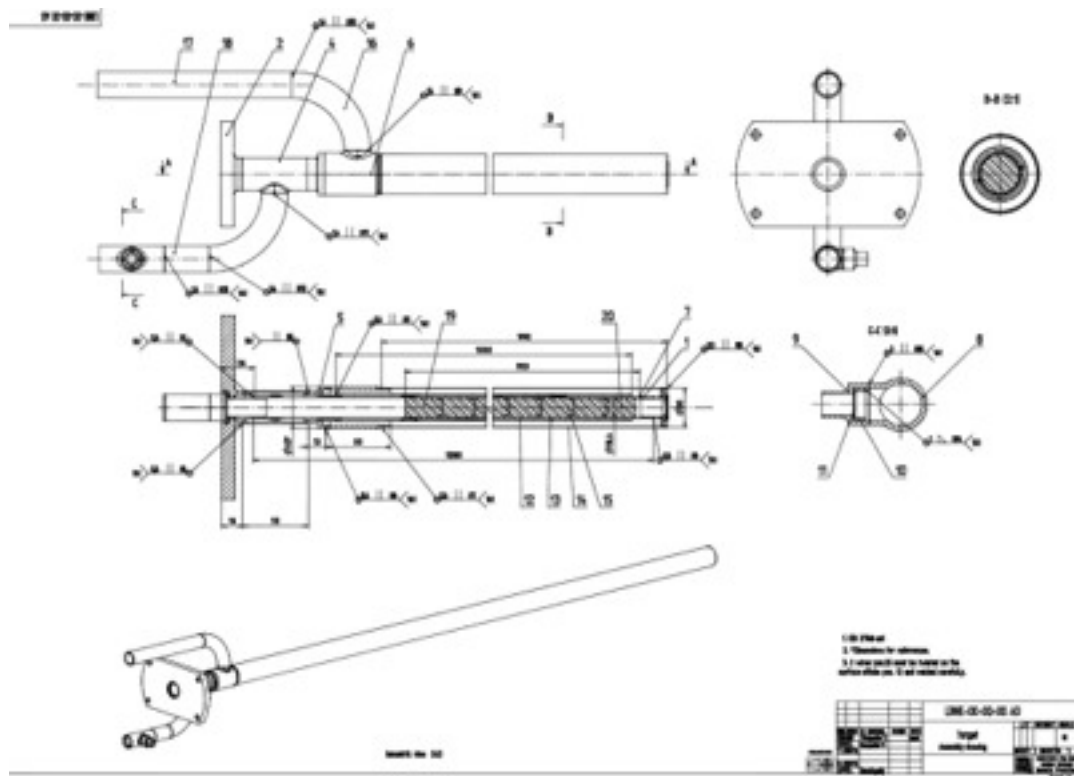


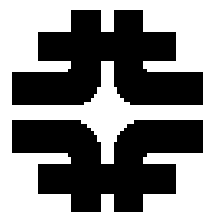
Aiming the Beam to DUSEL



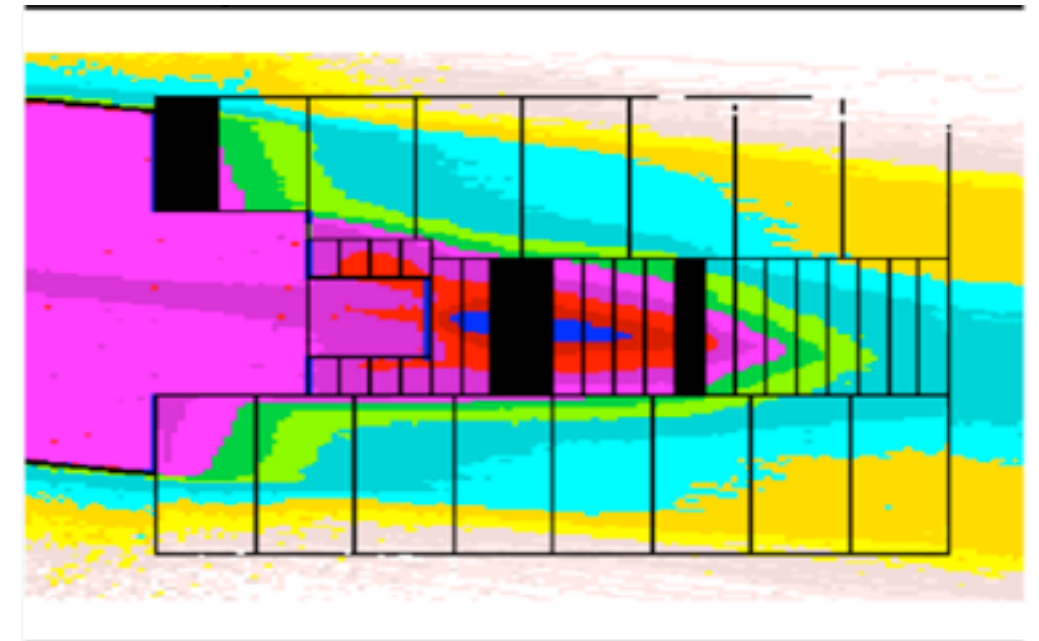
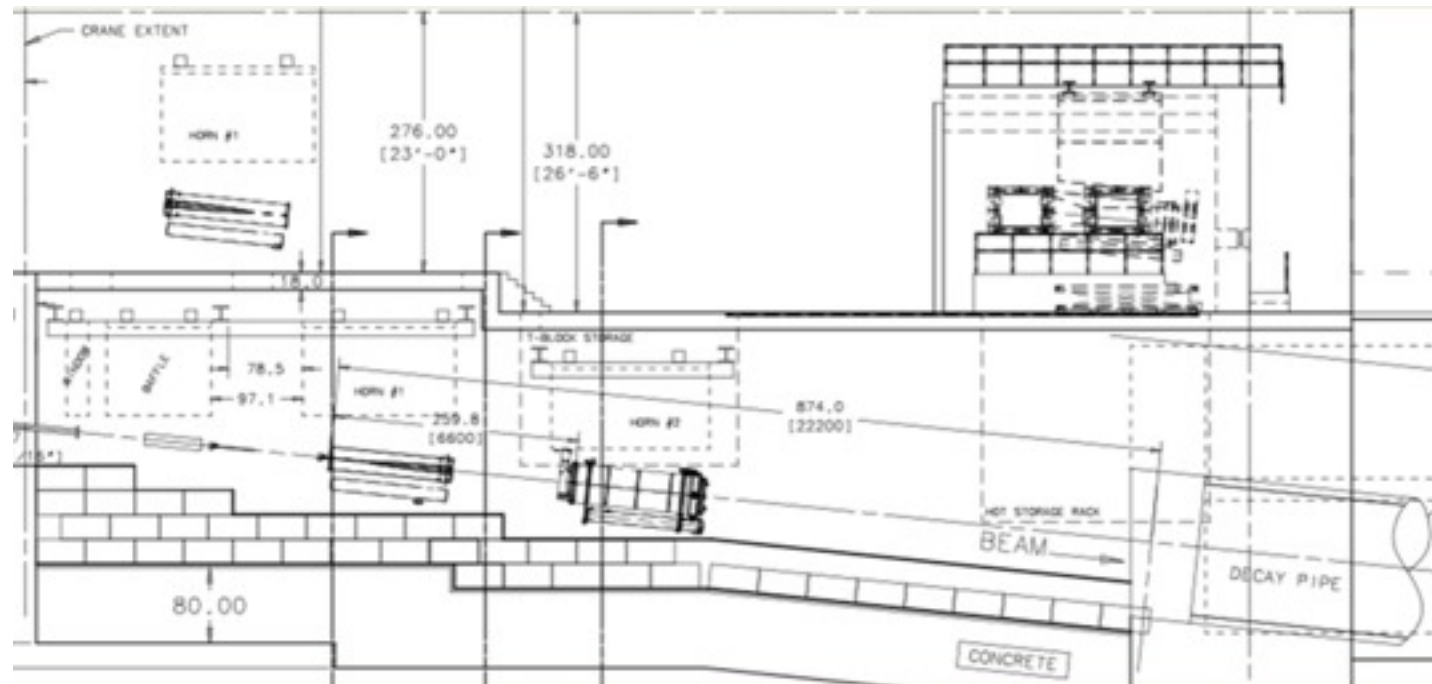


Target and Horns





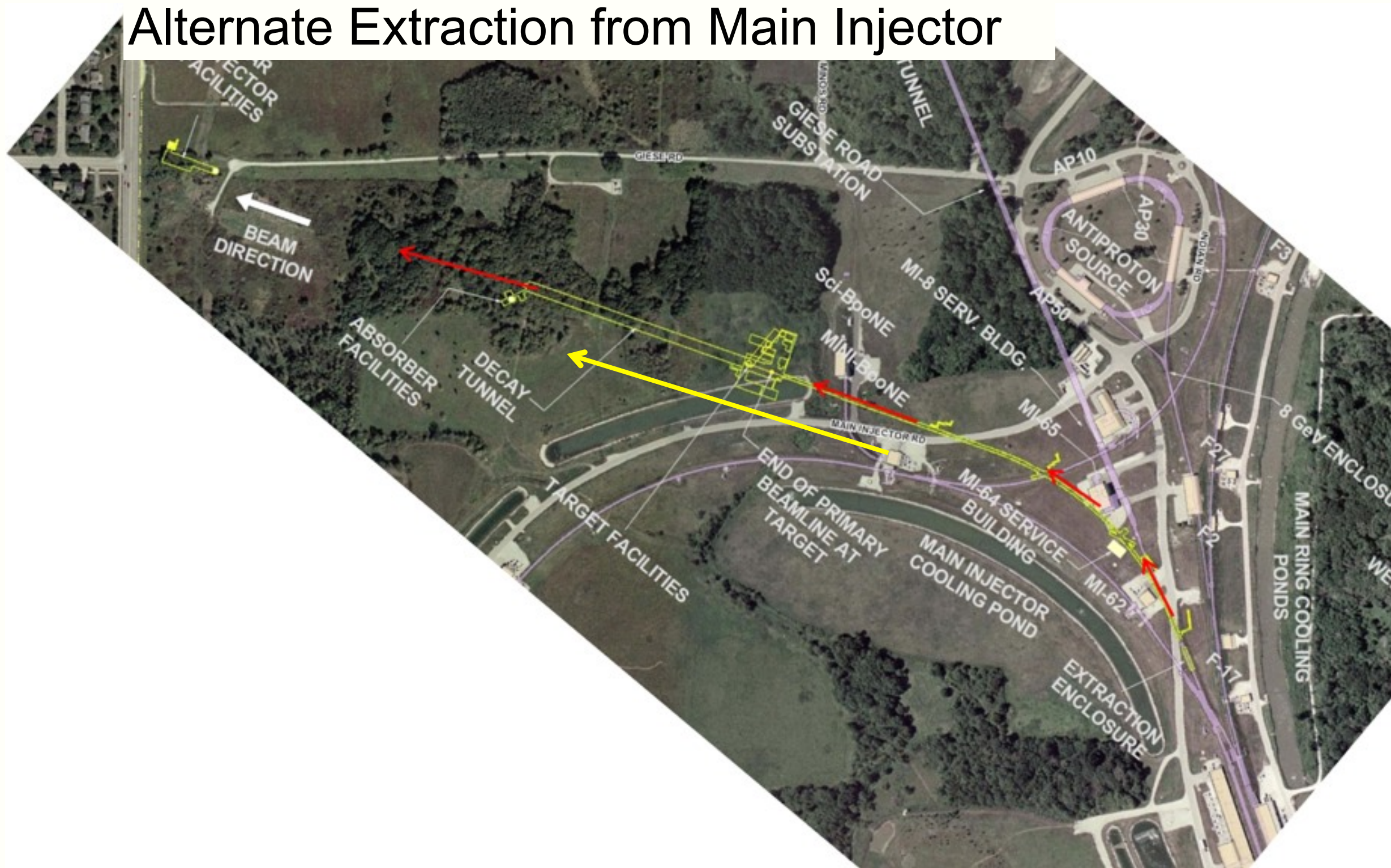
Target, Horn, Decay Pipe, and Absorber Engineering



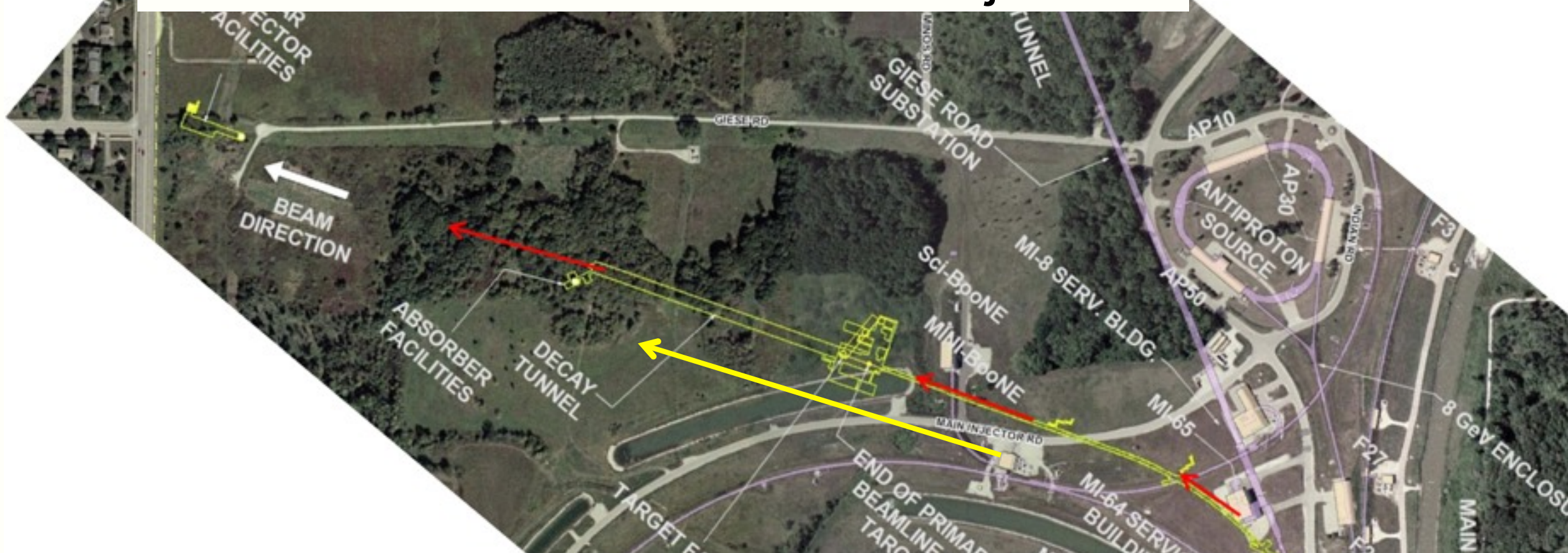
- Technical components designed for 700 kW beam (Main Injector with NOvA upgrades).
- Facilities designed to handle 2.3 MW beam (Project X).

Alternate Neutrino Beam Designs

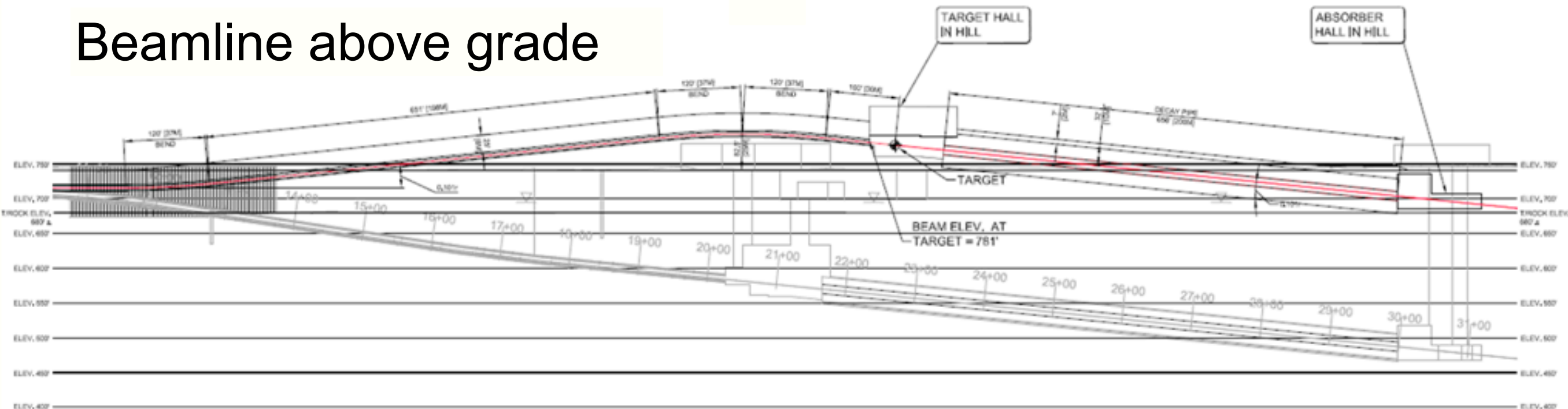
Alternate Extraction from Main Injector

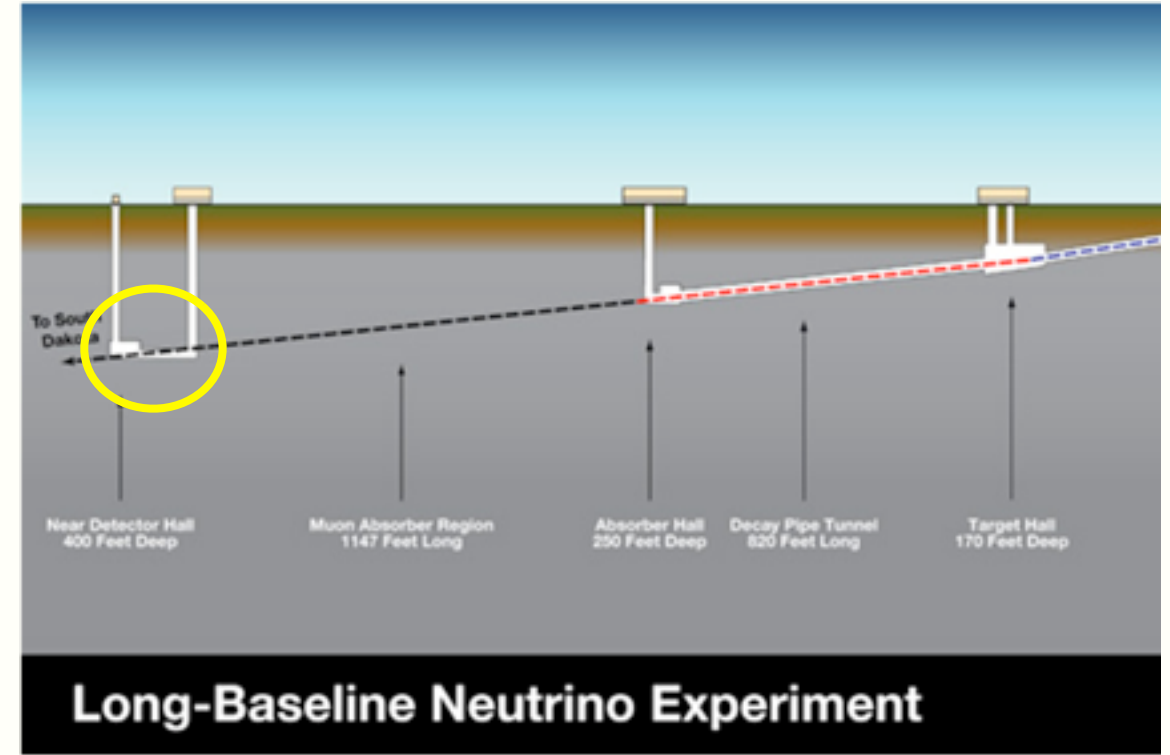


Alternate Extraction from Main Injector



Beamline above grade



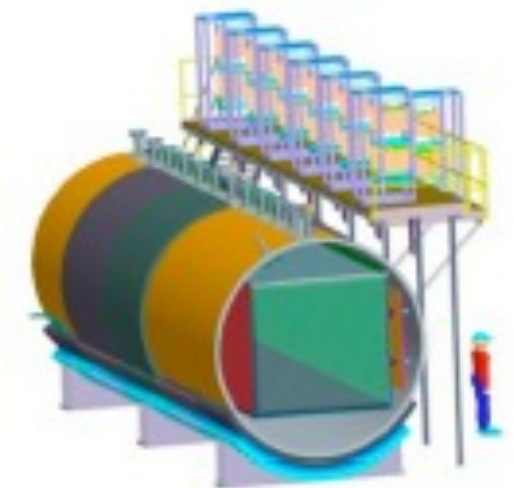
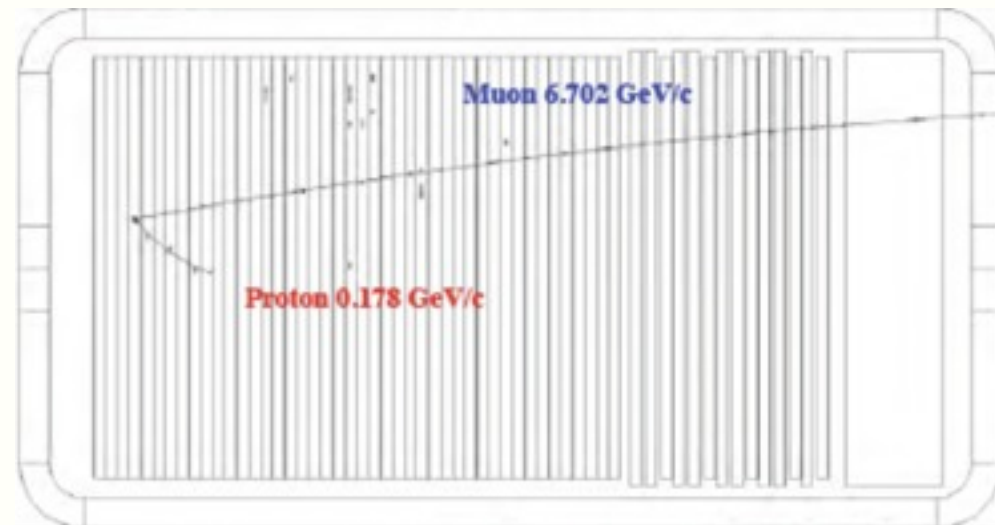
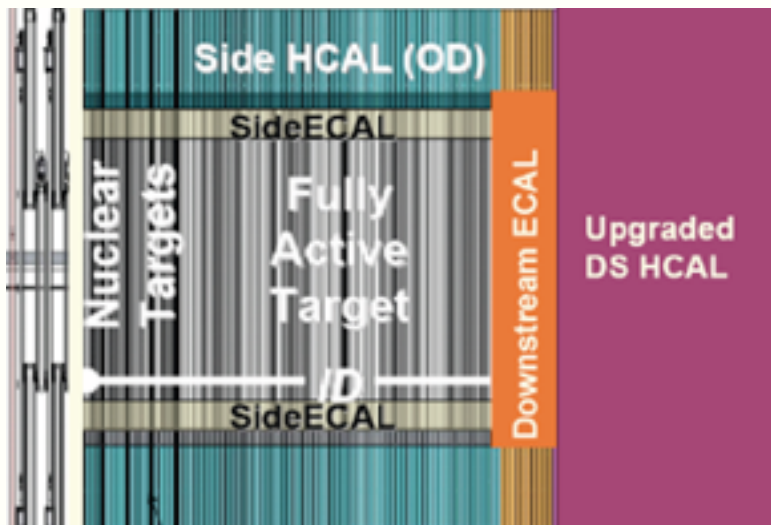


Several options under consideration:

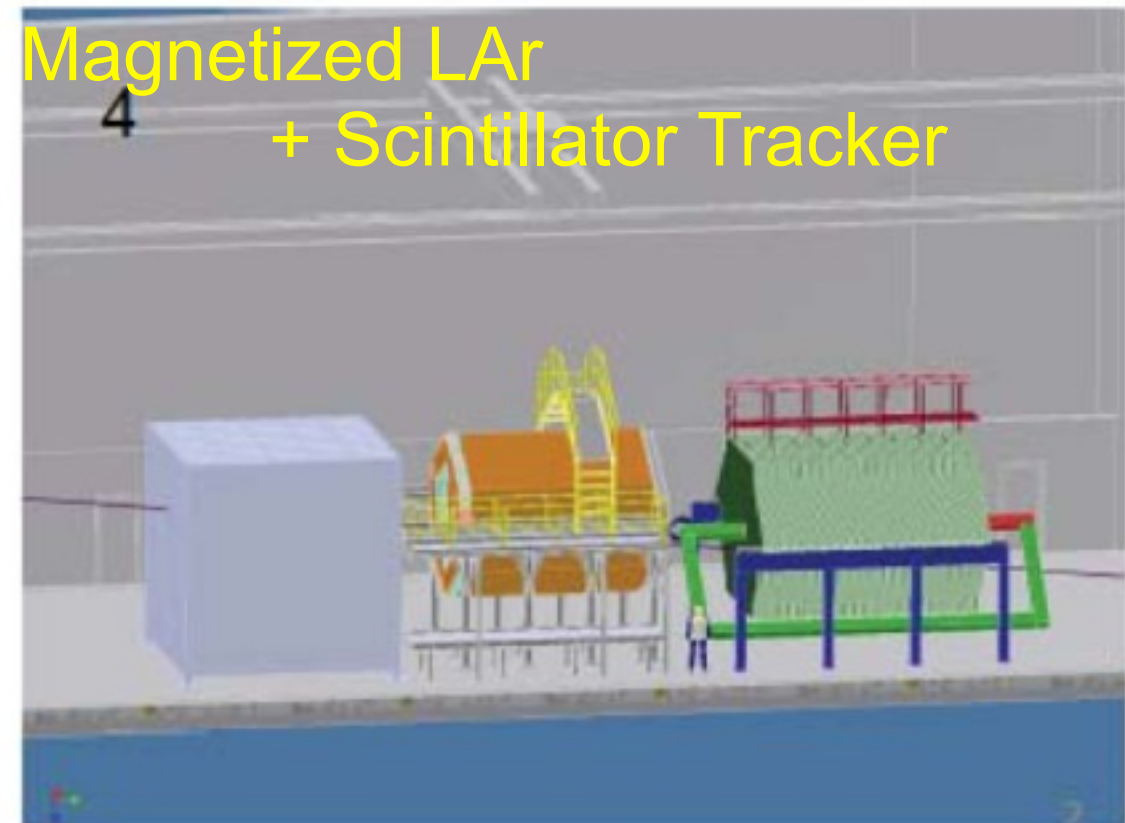
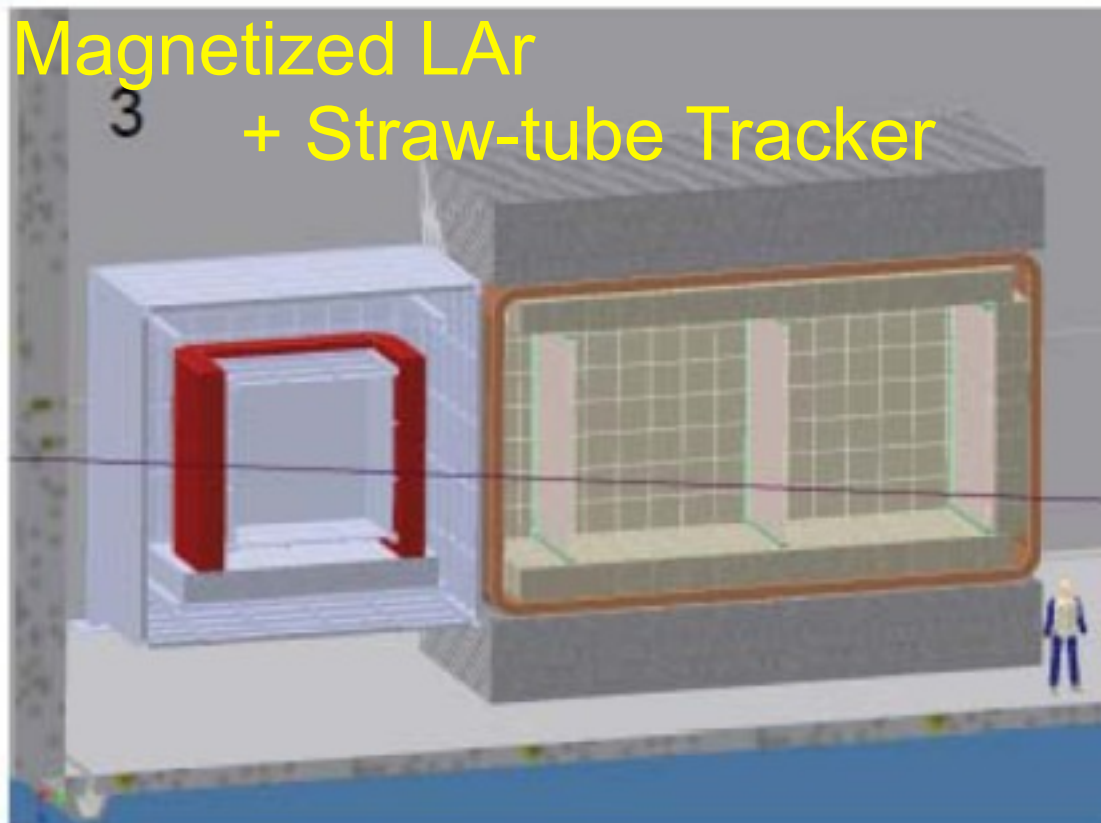
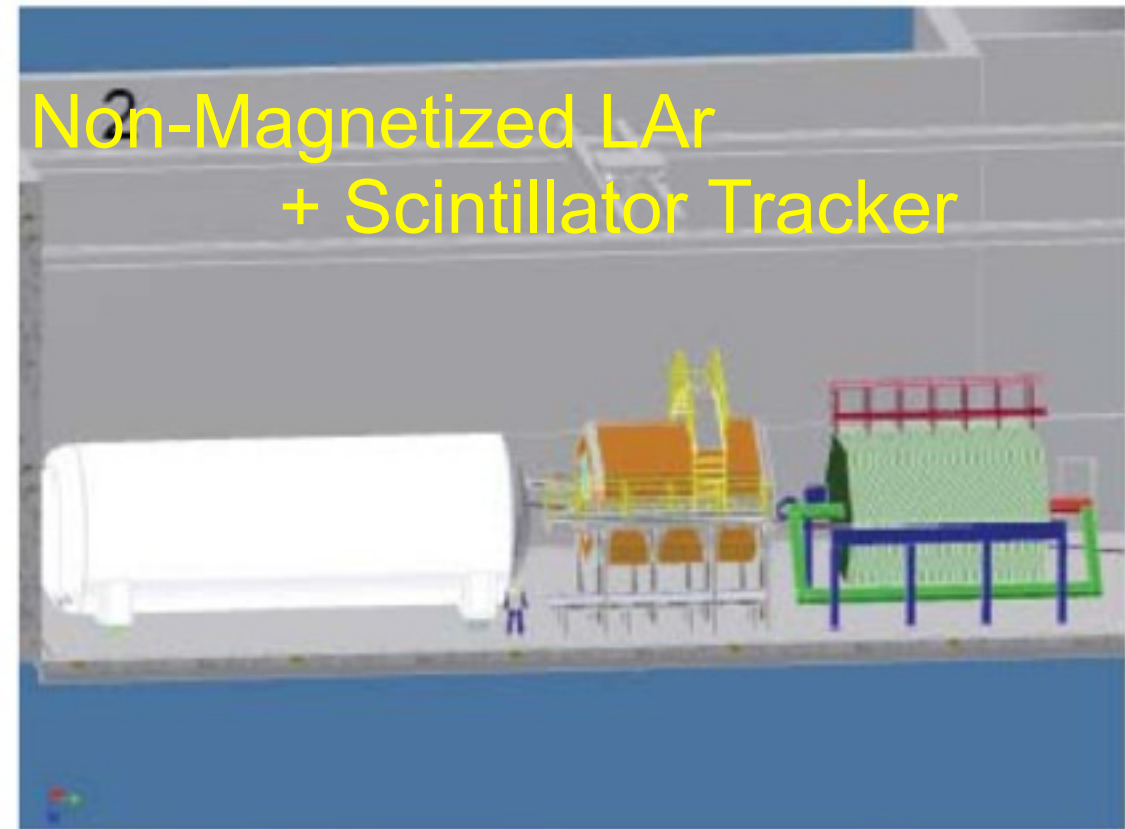
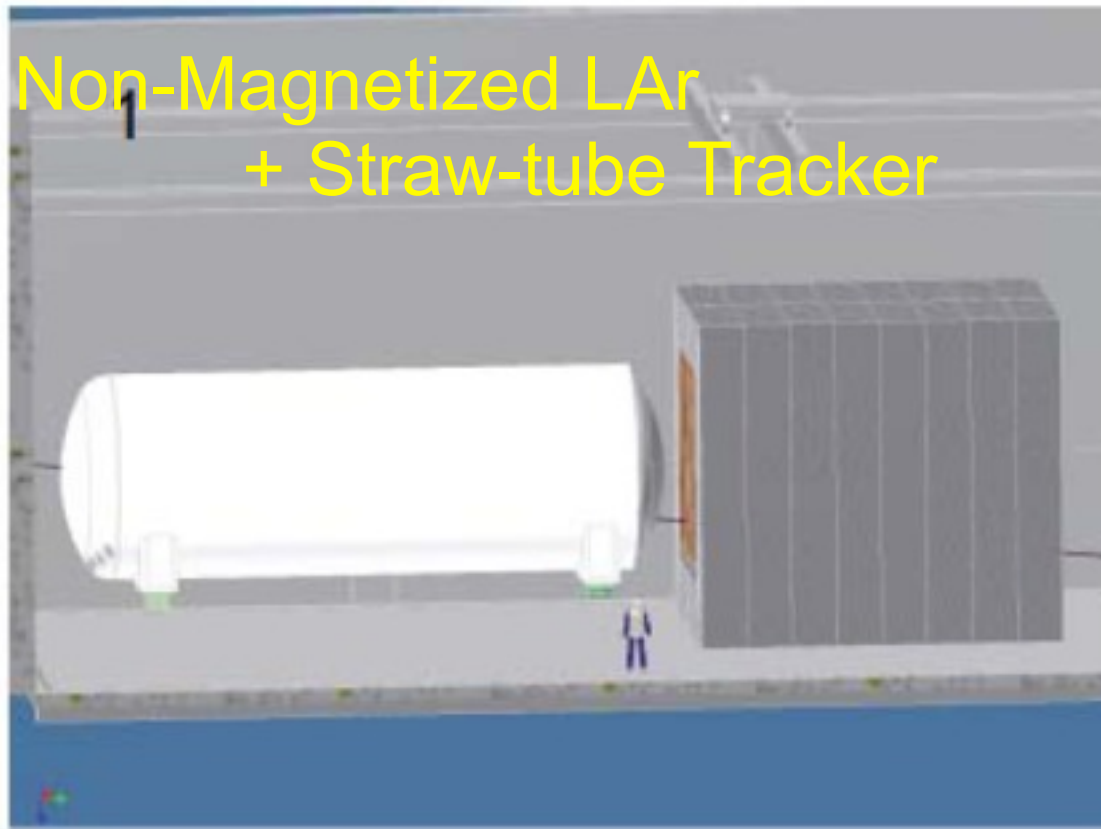
Scintillator tracker

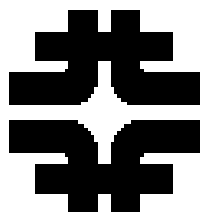
Straw-tube tracker

LAr TPC



Near Detector Options

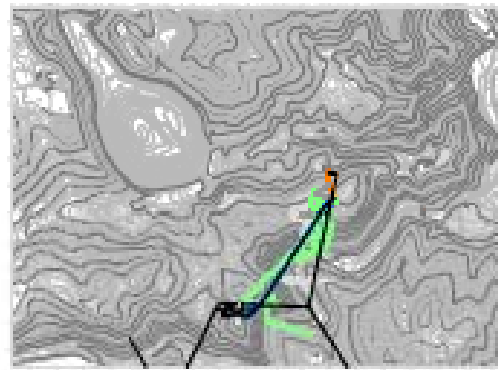




Water Cherenkov Far Detector

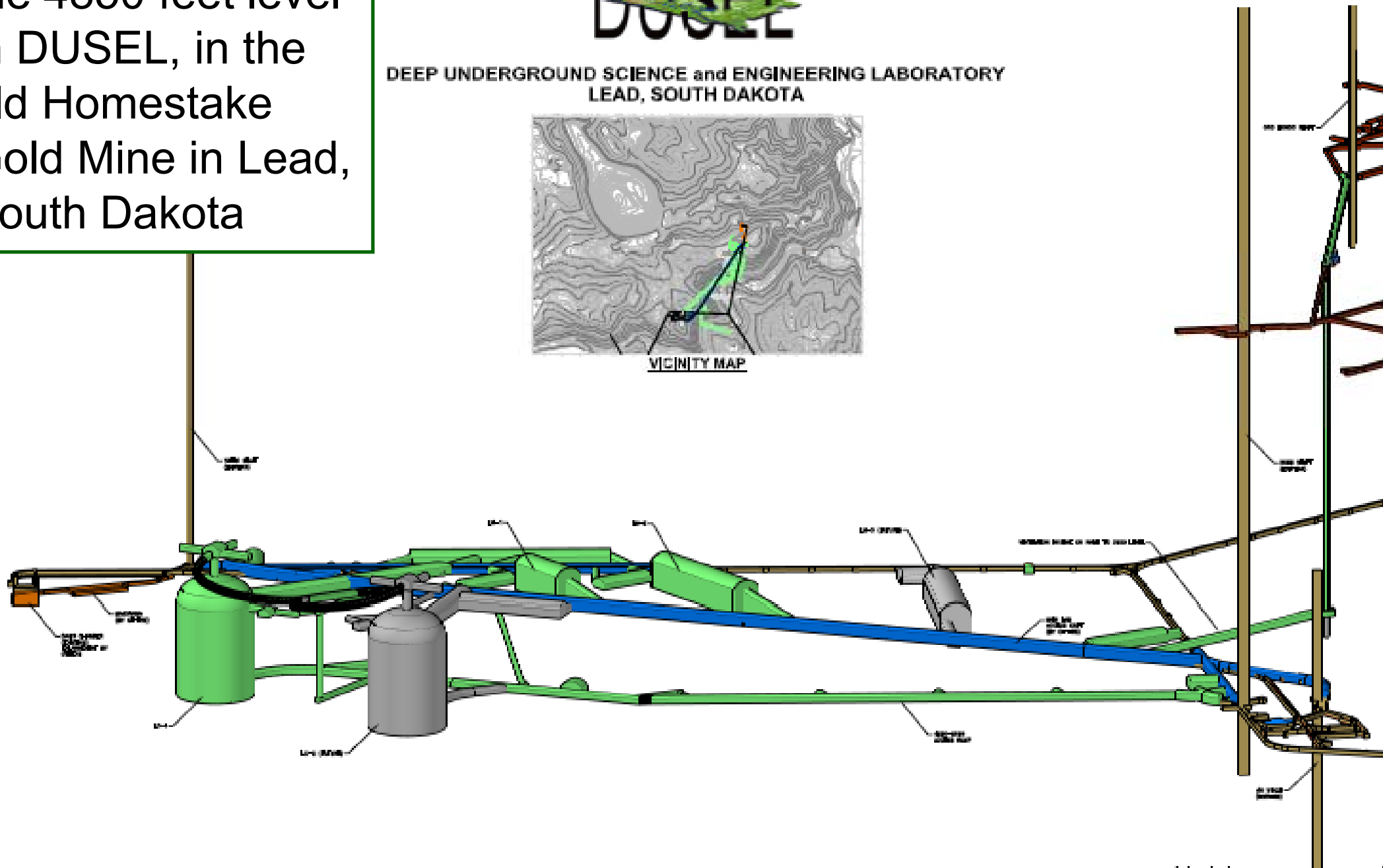
To be located at the 4850 feet level in DUSEL, in the old Homestake Gold Mine in Lead, South Dakota

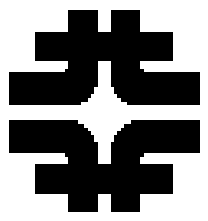
HOMESTAKE
DUSEL
DEEP UNDERGROUND SCIENCE and ENGINEERING LABORATORY
LEAD, SOUTH DAKOTA



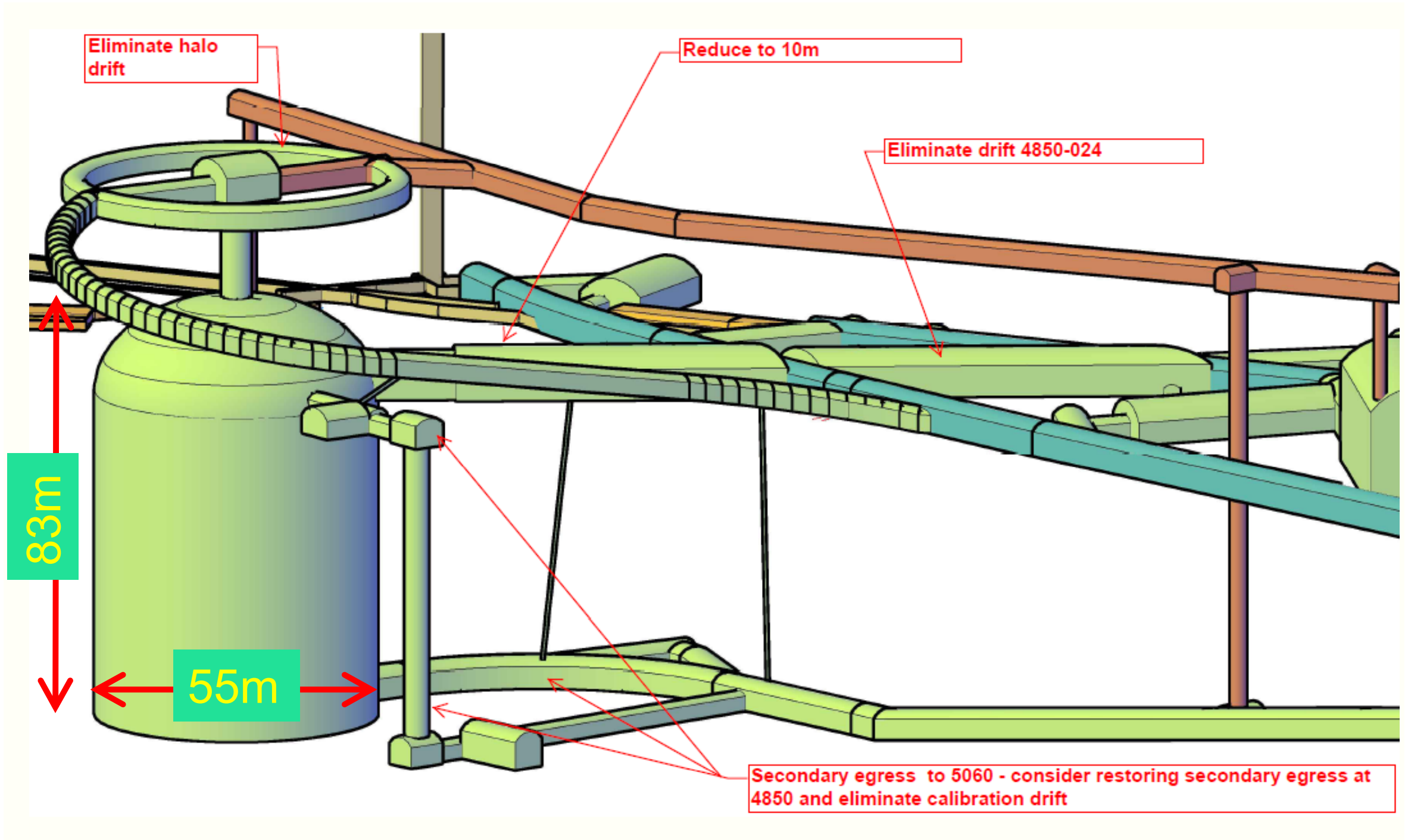
LEGEND

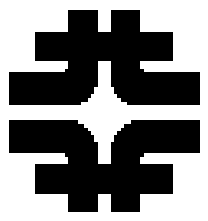
Blue line	NEW CONCRETE
Orange line	REPLACEMENT OF EXISTING STRUCTURE
Green line	EXISTING
Yellow line	PLUMB
Red line	IN PLACE





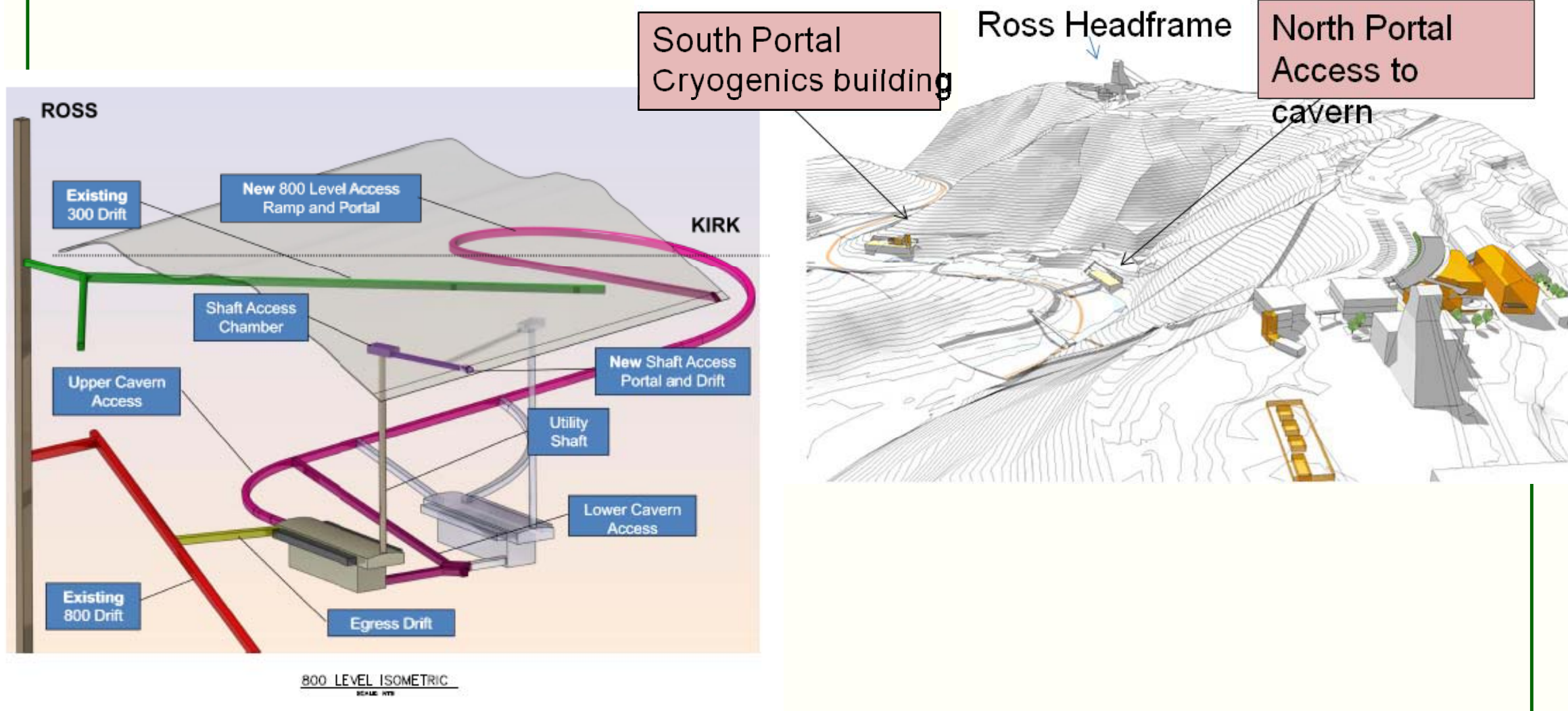
Water Cherenkov Detector

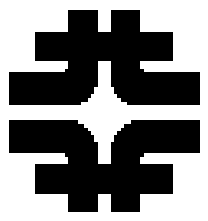




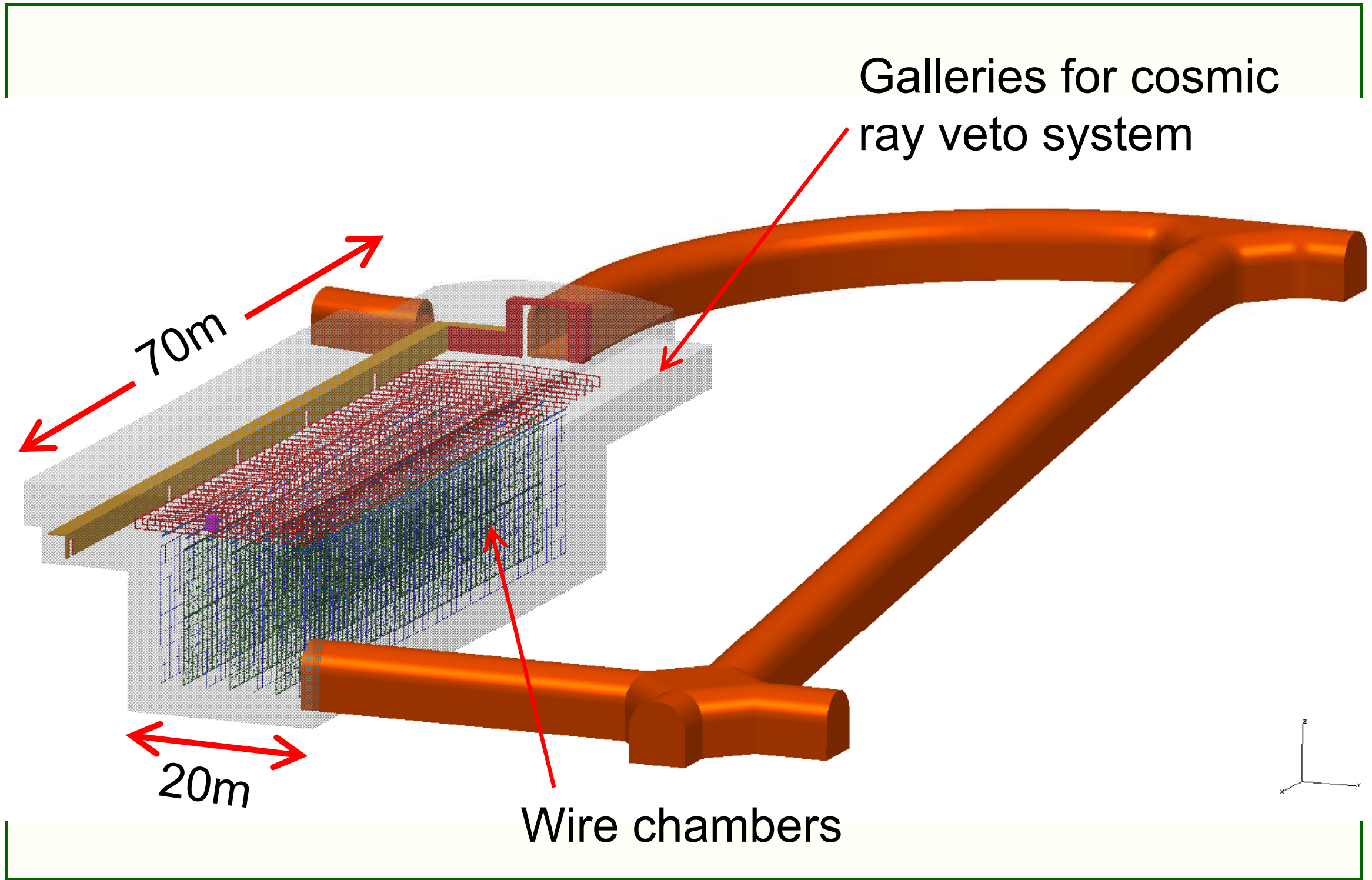
Liquid Argon TPC Far Detector

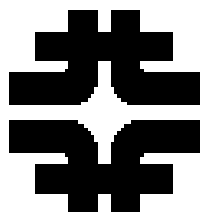
To be located 800 feet underground in DUSEL, in the old Homestake Gold Mine in Lead, SD



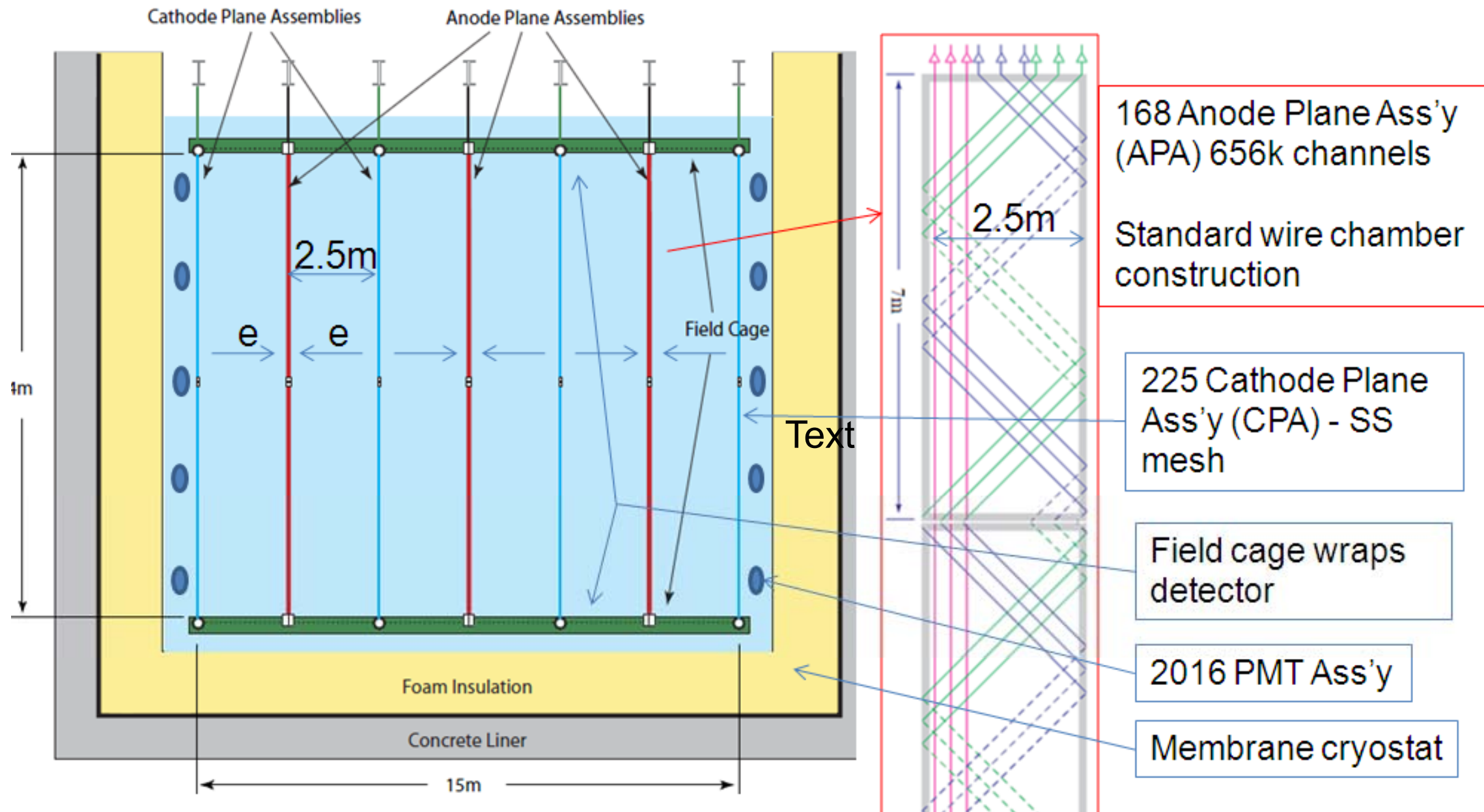


Liquid Argon TPC

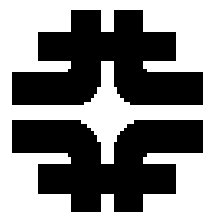




Liquid Argon TPC



Also considering designs: 3.75 m drift; two detectors in a common cavern.



Time Line

LBNE Project Time Line



- DOE CD-0 (Approve Mission Need)
January 2010
- CD-1 **2nd half CY2011***
- CD-2 (Project baseline) depending on funding . . . **mid/late FY2013**
- CD-3 (start construction) depending on funding . . . **2014 ~ 2015**
- Project complete : no sooner than **2020**

Currently dealing with TPC (total project funding) issues and DOE-NSF partnership issues

Alternative Sites



1300km



Alternative Sites



1300km

Henderson 1500km



Alternative Sites



1300km

Henderson 1500km

WIPP 1700km

Alternative Sites



1300km

Yucca Mtn 2300km

Henderson 1500km

WIPP 1700km

Alternative Sites

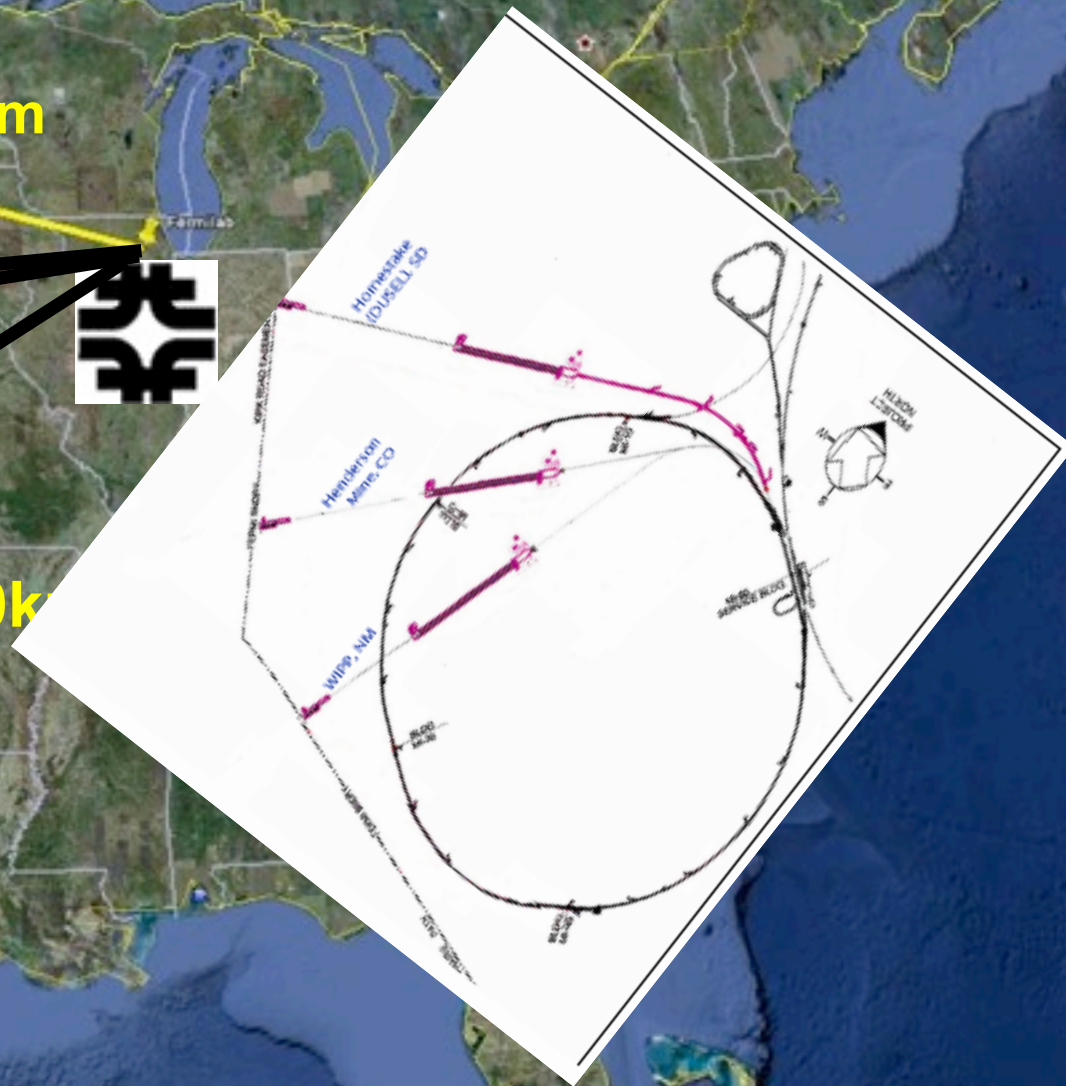


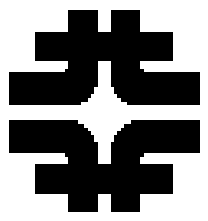
1300km

Yucca Mtn 2300km

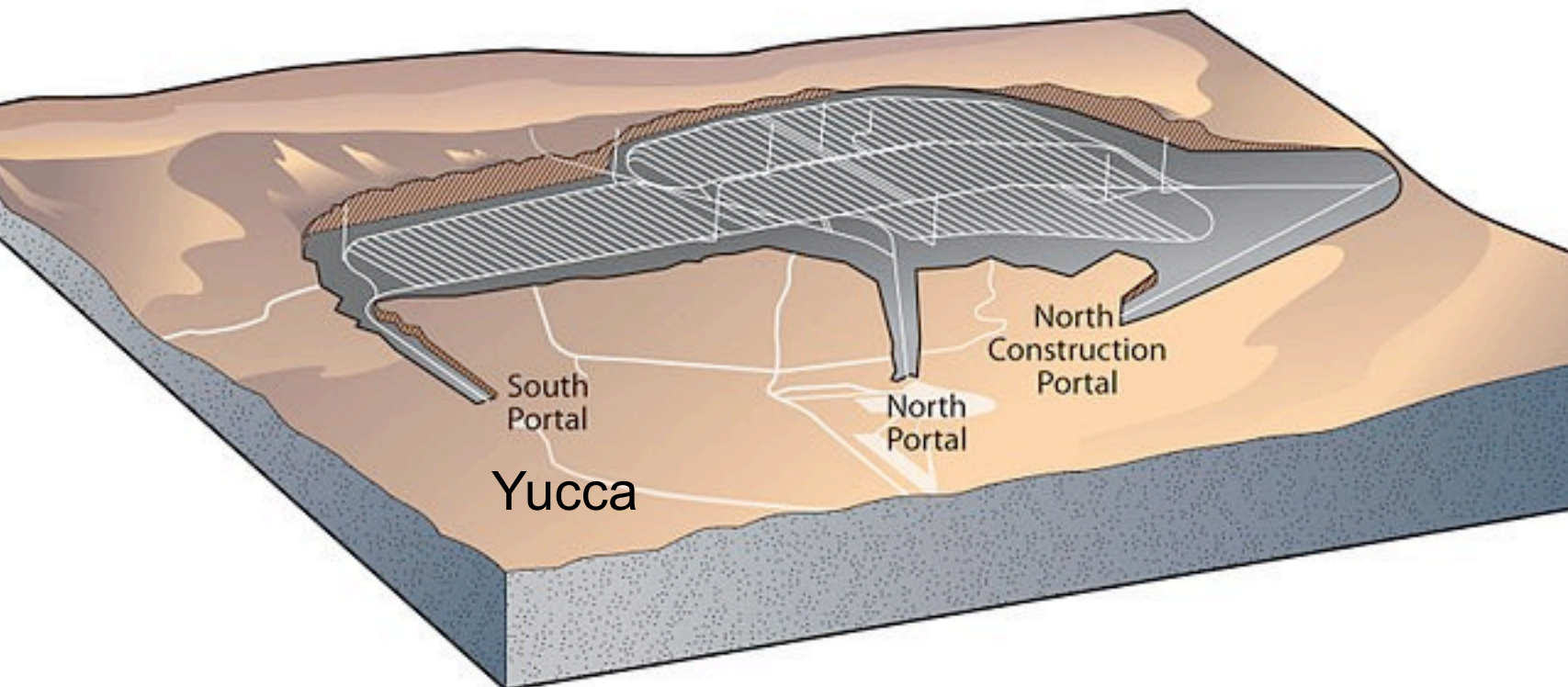
Henderson 1500km

WIPP 1700km





Henderson-WIPP-Yucca Mtn



In Matter:

Bi-Magic Baseline:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

$$\text{where } \sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$$

$$\text{and } \sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

In Matter:

Bi-Magic Baseline:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

$$\text{where } \sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$$

$$\text{and } \sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

In Matter:

Bi-Magic Baseline:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

$$\text{where } \sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$$

$$\text{and } \sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

In Matter:

Max for one Hierarchy and 0 other

Bi-Magic Baseline:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

In Matter:

Max for one Hierarchy and 0 other

Bi-Magic Baseline:

$$P_{\mu \rightarrow e} \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \right|^2$$

where $\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31}$

and $\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{(aL)} \Delta_{21}$

$$a = G_F N_e / \sqrt{2} = (4000 \text{ km})^{-1},$$

Sushant K. Raut, Ravi Shanker Singh, S.Uma Sankar [arXiv:0908.3741](#)

Amol Dighe, Srubabati Goswami, Shamayita Ray [arXiv:1009.1093](#)

“Bi-Magic” Baseline and Energy

Choose L such that

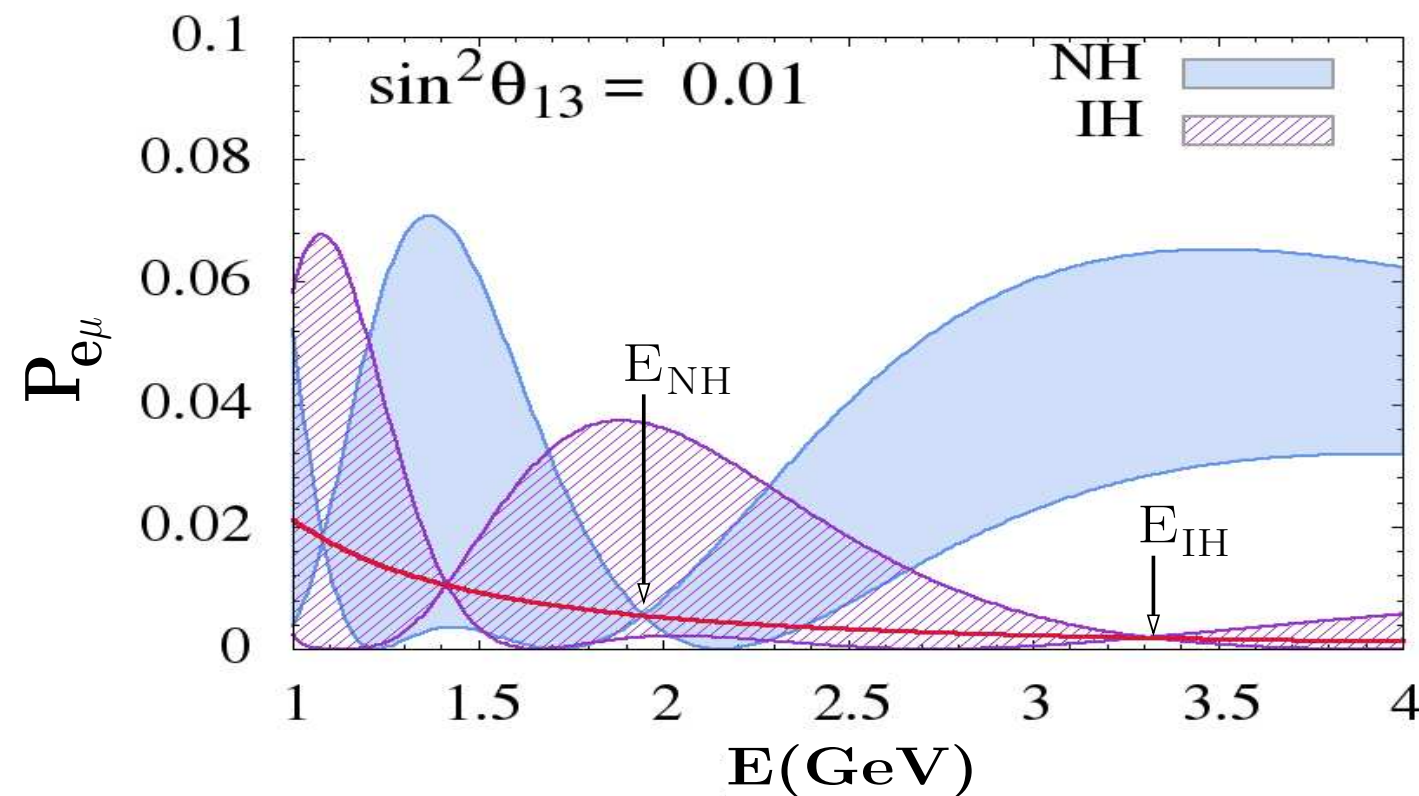
$$P_{atm}|_{IH} = 0 \text{ and } P_{atm}|_{NH} \text{ is max. at } E_{IH}$$

and

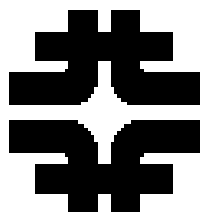
$$P_{atm}|_{NH} = 0 \text{ and } P_{atm}|_{IH} \text{ is max. at } E_{NH}$$

$L=2540 \text{ km}$ and $E_{IH}=3.3 \text{ GeV}$ and $E_{NH}=1.9 \text{ GeV}$

flip when ν and $\bar{\nu}$ interchange



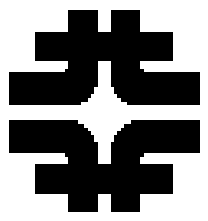
Approx. Fermilab to Yucca Mtn:



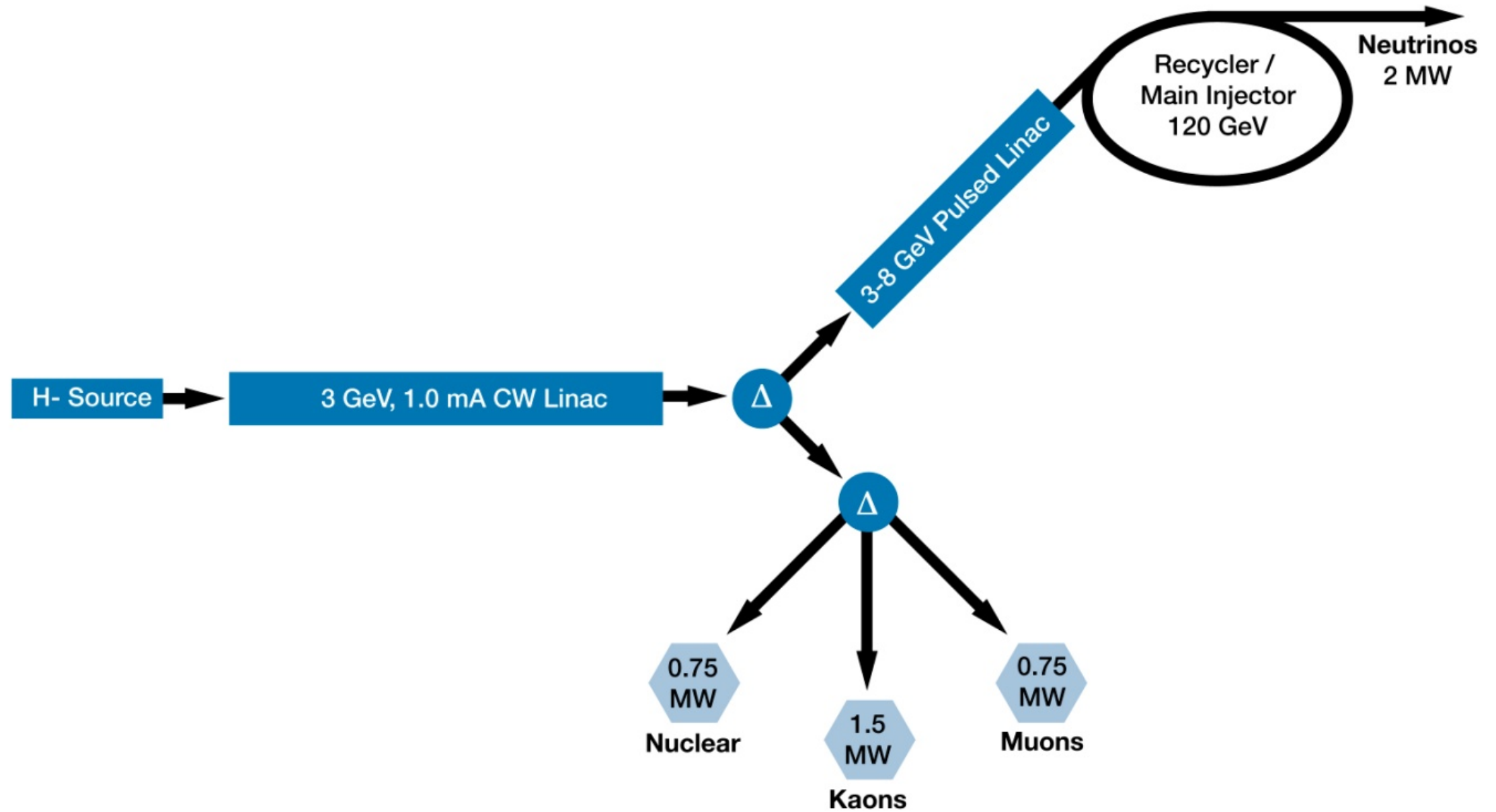
Project X

- A neutrino beam for long baseline neutrino oscillation experiments
 - 2 MW proton source at 60-120 GeV
- High intensity, low energy protons for kaon and muon based precision experiments
 - Operations simultaneous with the neutrino program
- A path toward a muon source for possible future Neutrino Factory and/or a Muon Collider
 - Requires ~4 MW at ~5-15 GeV
- Possible missions beyond P5
 - Standard Model Tests with nuclei and energy applications



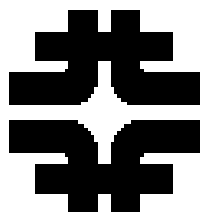


Reference Design

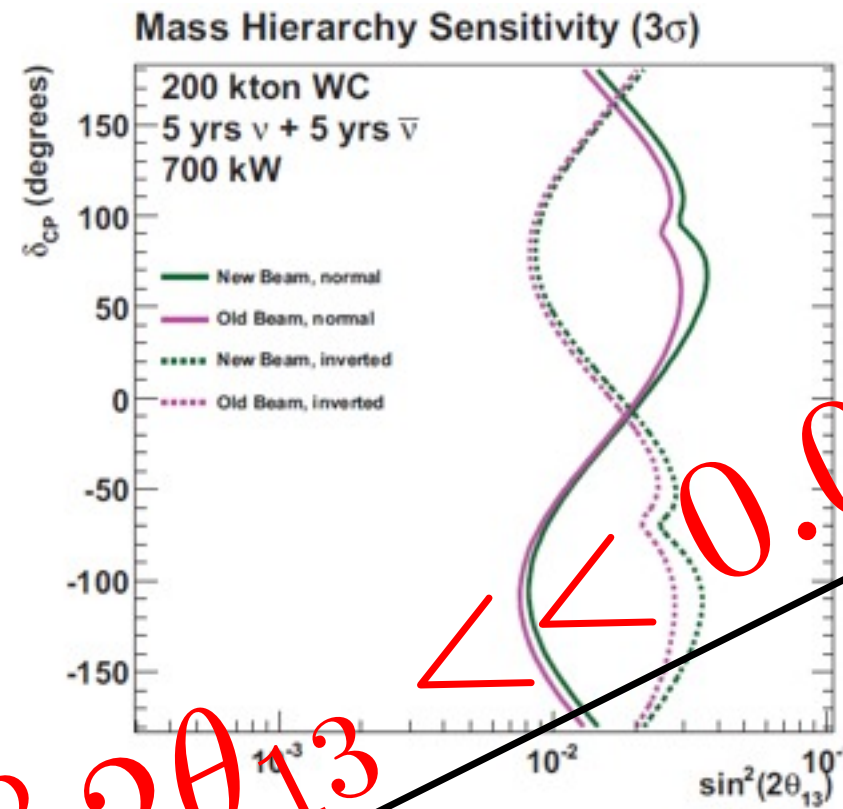
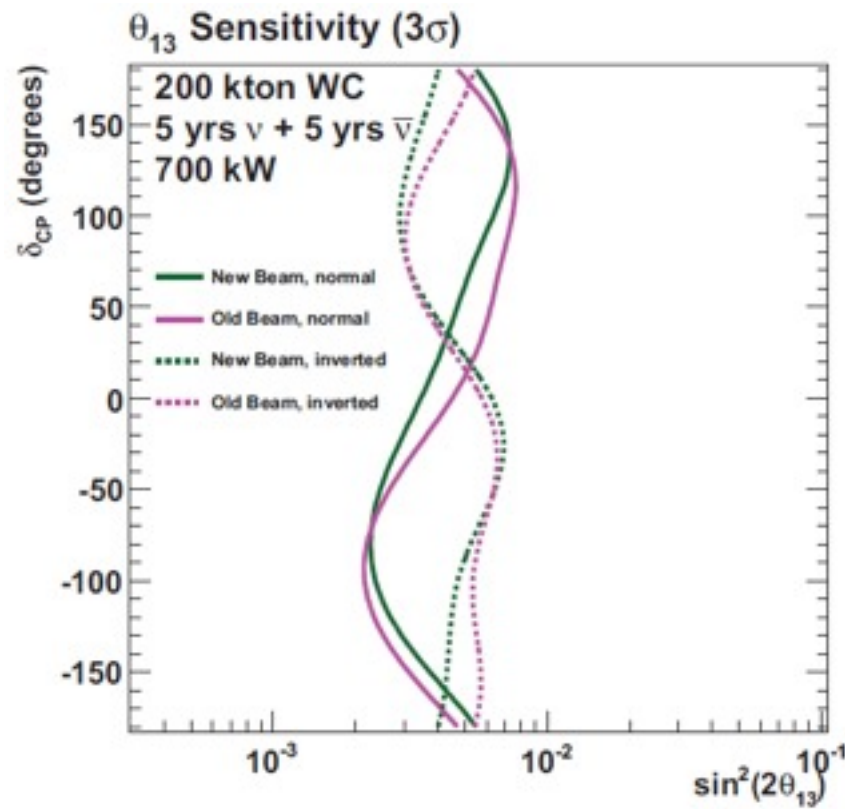




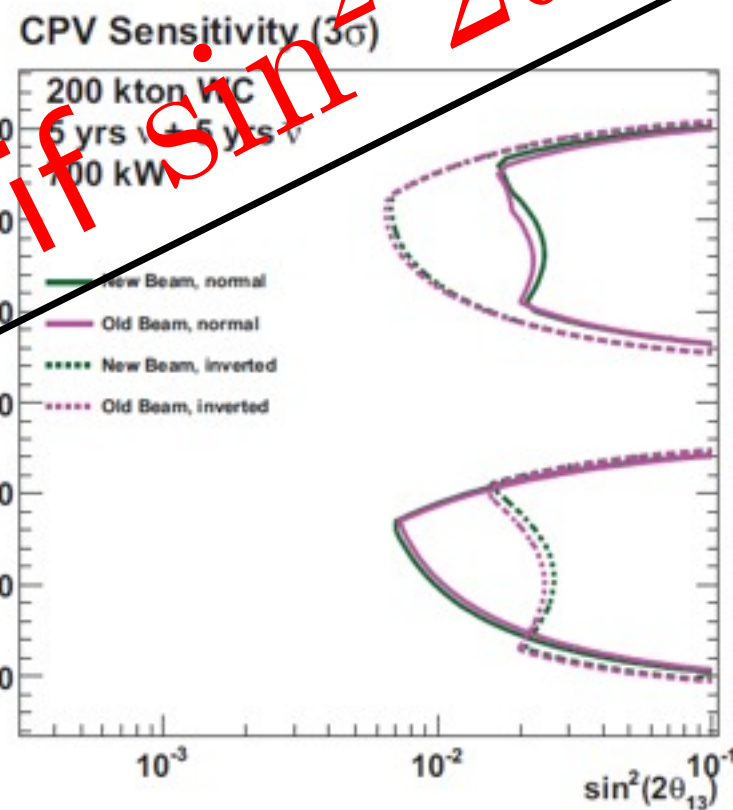
- 3 GeV CW superconducting H- linac with 1 mA average beam current.
 - Flexible provision for variable beam structures to multiple users
 - CW at time scales $>1 \mu\text{sec}$, 10% DF at $<1 \mu\text{sec}$
 - Supports rare processes programs at 3 GeV
 - Provision for 1 GeV extraction for nuclear energy program
 - 3-8 GeV pulsed linac capable of delivering 300 kW at 8 GeV
 - Supports the neutrino program
 - Establishes a path toward a muon based facility
 - Upgrades to the Recycler and Main Injector to provide ≥ 2 MW to the neutrino production target at 60-120 GeV.
- ⇒ Utilization of a CW linac creates a facility that is unique in the world, with performance that cannot be matched in a synchrotron-based facility.



LBNE Sensitivities:



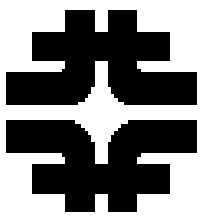
Sensitivities shown are for 200 ktons WCD.
With 34 ktons LAr, the sensitivities are equivalent.



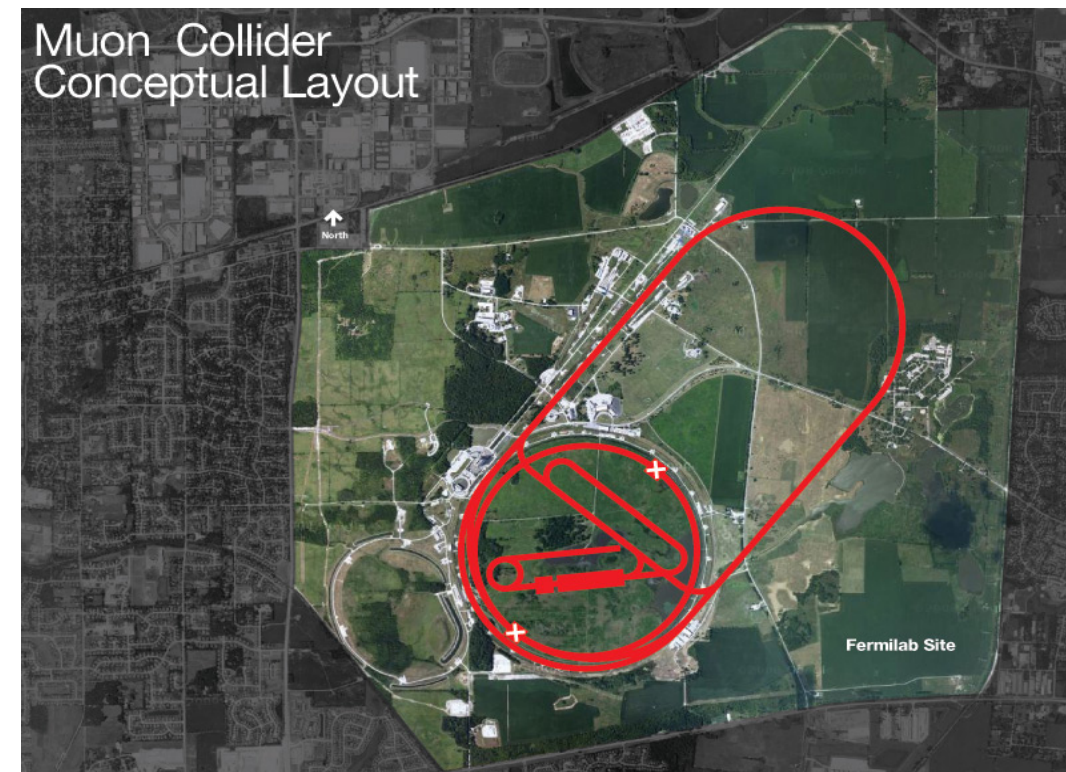
Sensitivities shown are for 10 years with 700 kW proton beam.
With Project X (≥ 2 MW), the same sensitivity would be reached in 1/3 the time ... or $\sim\sqrt{3}$ greater sensitivity for the same running period.

What if $\sin^2 2\theta_{13} < 0.01$????

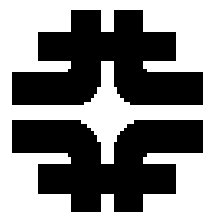
Fall 2010 Report from the LBNE Physics Working Group



- Project X shares many features with the proton driver required for a Neutrino Factory or Muon Collider
 - NF and MC require ~ 4 MW @ 10 ± 5 GeV
 - Primary issues are related to beam “format”
 - NF wants proton beam on target consolidated in a few bunches; Muon Collider requires single bunch
 - Project X linac is not capable of delivering this format



⇒ It is inevitable that a new ring(s) will be required to produce the correct beam format for targeting.



Project X: Time Line

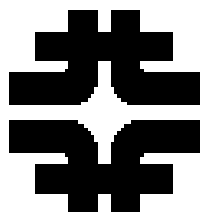
Project X Timeline

- Assumed Critical Decision dates
 - CD-0: January 2011
 - CD-1: July 2012
 - CD-2: August 2013
 - CD-3: September 2014
 - CD-4: September 2019

5 yr construction

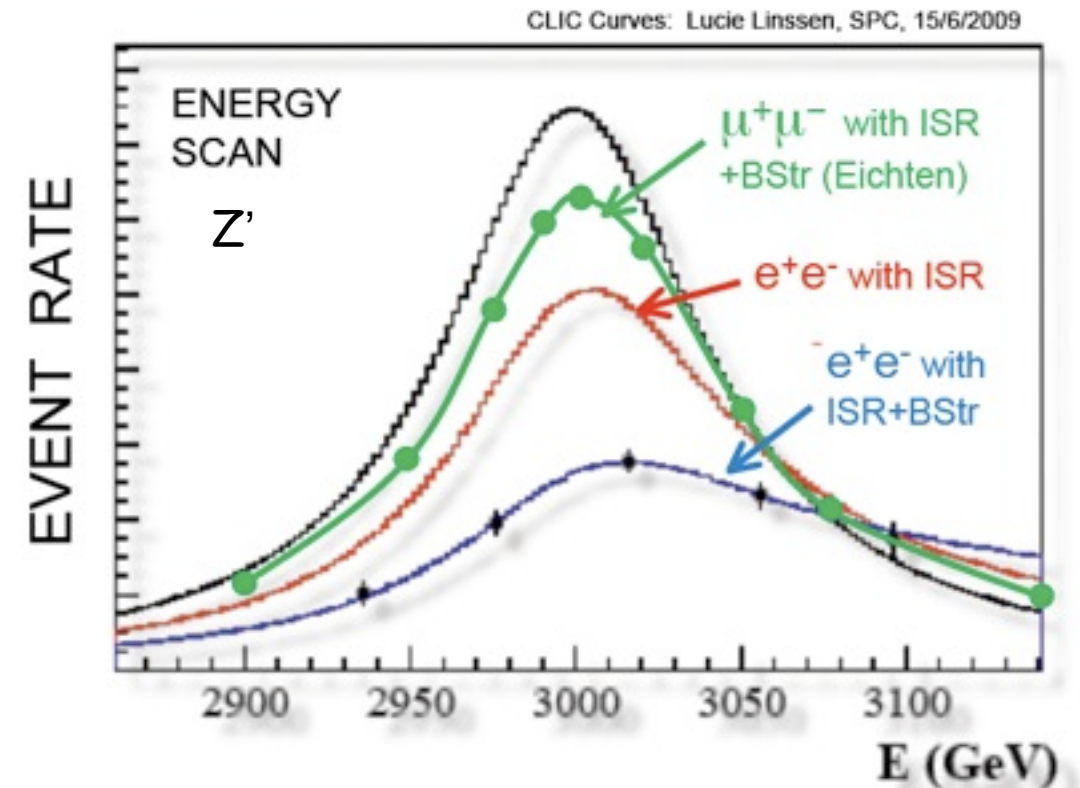
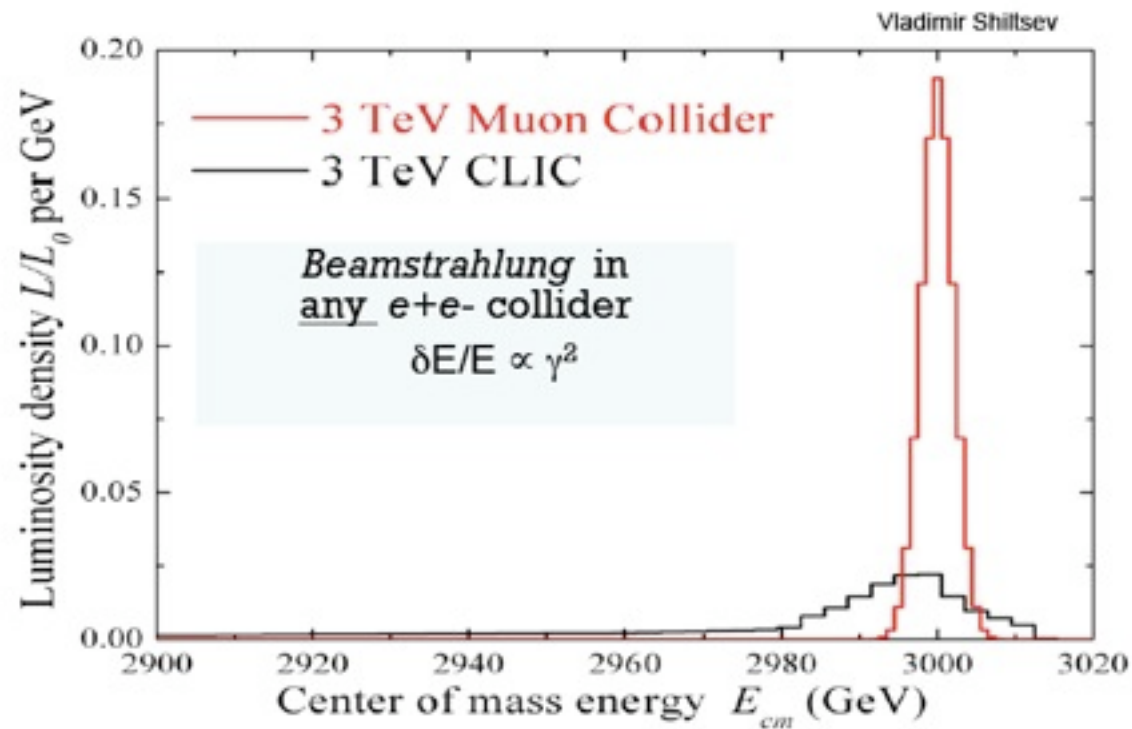
Project X could also be up and running in ~2020 :

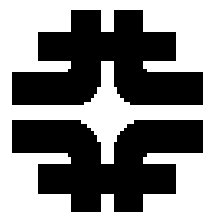
All depends on funding profiles...



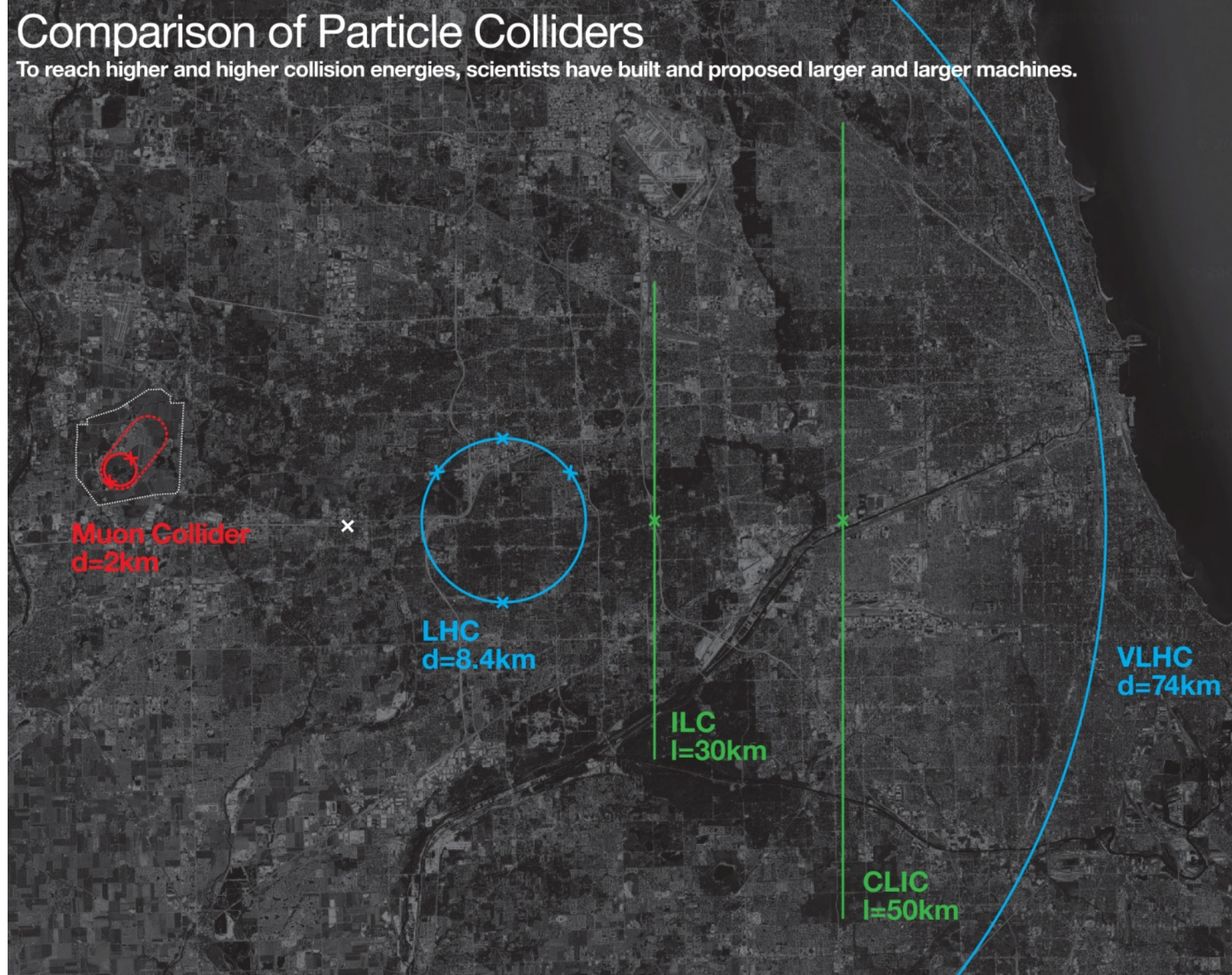
Muon Collider

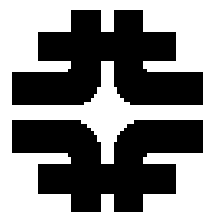
- $\mu^+\mu^-$ Collider:
 - Center of Mass energy: 1.5 - 5 TeV (focus 3 TeV)
 - Luminosity $> 10^{34} \text{ cm}^{-2} \text{ sec}^{-1}$ (focus 400 fb^{-1} per year)
- Compact facility
- Superb Energy Resolution
 - MC: 95% luminosity in $dE/E \sim 0.1\%$
 - CLIC: 35% luminosity in $dE/E \sim 1\%$





Muon Collider:





Minimum Luminosity for Muon Collider

□ Universal behavior for s-channel resonance

$$\sigma(E) = \frac{2J+1}{(2S_1+1)(2S_2+1)} \frac{4\pi}{k^2} \left[\frac{\Gamma^2/4}{(E-E_0)^2 + \Gamma^2/4} \right] B_{in} B_{out}$$

Convolute with beam resolution ΔE .

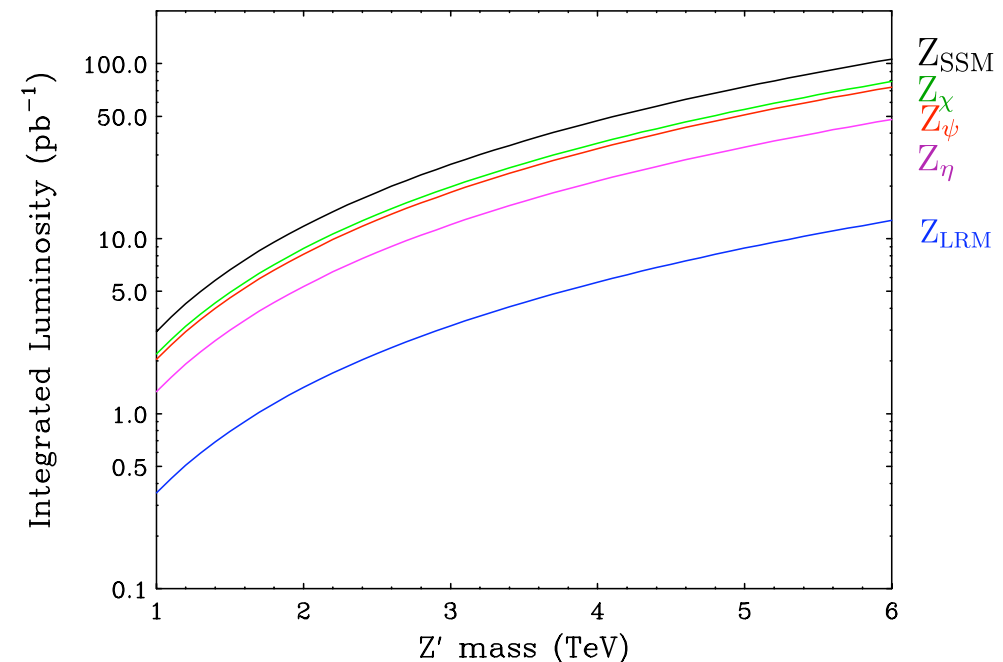
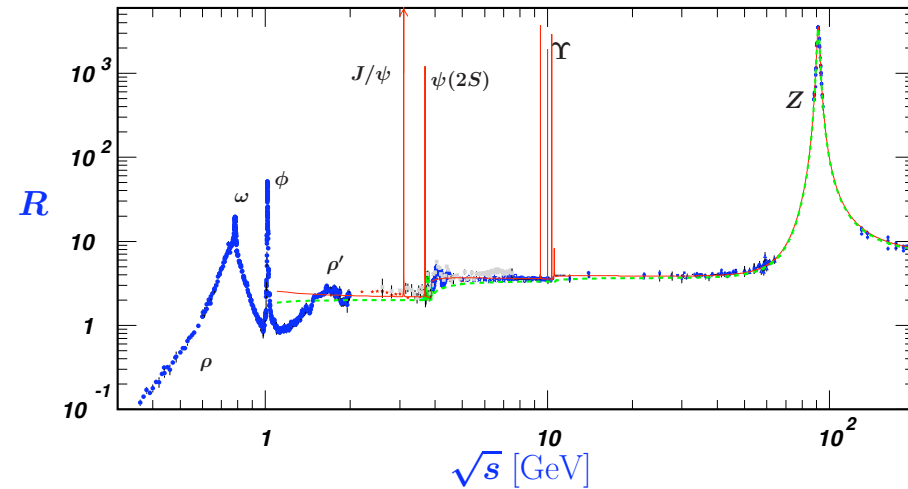
If $\Delta E \ll \Gamma$

$$R_{\text{peak}} = (2J+1) 3 \frac{B(\mu^+\mu^-) B(\text{visible})}{\alpha_{\text{EM}}^2}$$

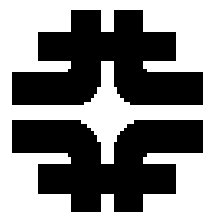
□ Can use to set minimum required luminosity

- Likely new physics candidates:
 - scalars: h, H^0, A^0, \dots
 - gauge bosons: Z'
 - new dynamics: bound states
 - ED: KK modes
- Example - new gauge boson: Z'
 - SSM, E6, LRM
 - 5σ discovery limits: 4-5 TeV at LHC (@ 300 fb⁻¹)

Minimum luminosity at Z' peak:
 $\mathcal{L} = 0.5-5.0 \times 10^{30} \text{ cm}^{-2} \text{ sec}^{-1}$
 for $M(Z') \rightarrow 1.5-5.0 \text{ TeV}$

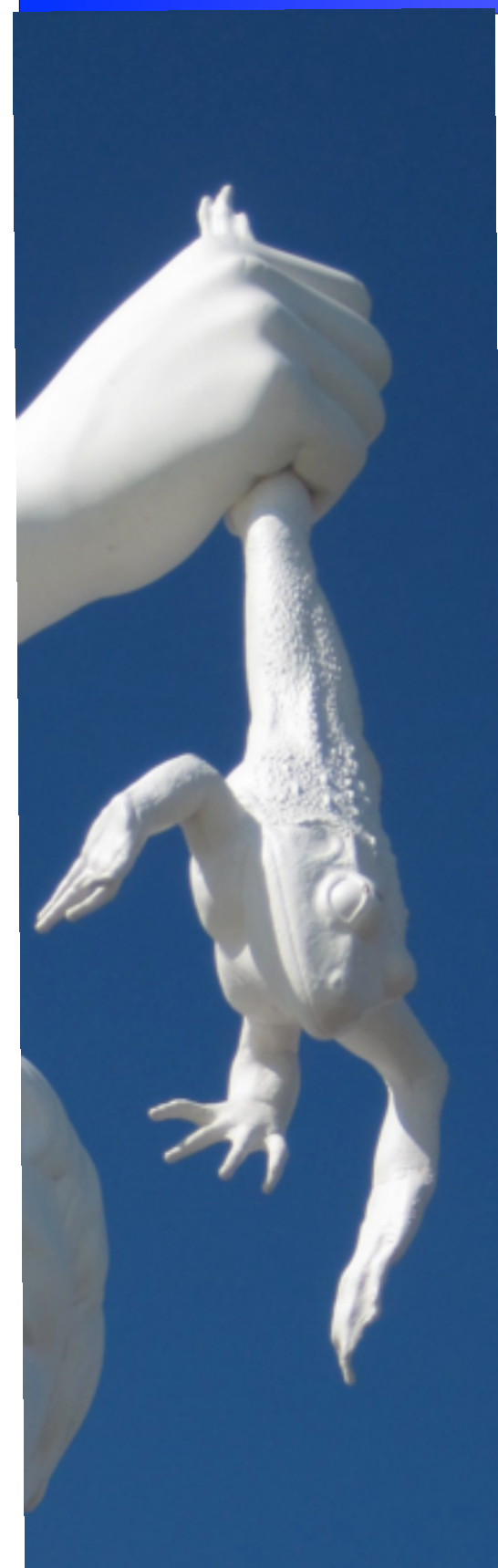


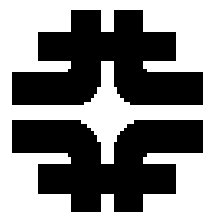
The integrated luminosity required to produce 1000 $\mu^+\mu^- \rightarrow Z'$ events on the peak



Summary:

- LBNE: Fermilab to DUSEL (or ?) will probe leptonic CPV and determine the mass hierarchy down to $\sin^2 2\theta_{13} \sim 0.01$
- Project X: Increases the beam power from 700kW to 2MW and provide another 3MW at 3 GeV for K's, μ/s ,
- If $\sin^2 2\theta_{13} < 0.01$ a Neutrino Factory is a possible option. Maybe Low Energy NuF.
- This could be a natural stepping stone to a Muon Collider. New physics such as a Z' would make this very compelling.

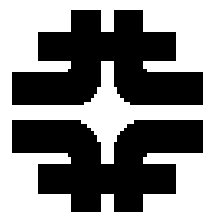




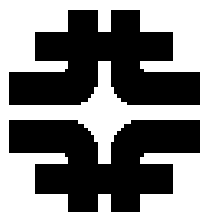
Summary:

- LBNE: Fermilab to DUSEL (or ?) will probe leptonic CPV and determine the mass hierarchy down to $\sin^2 2\theta_{13} \sim 0.01$
- Project X: Increases the beam power from 700kW to 2MW and provide another 3MW at 3 GeV for K's, μ/s ,
- If $\sin^2 2\theta_{13} < 0.01$ a Neutrino Factory is a possible option. Maybe Low Energy NuF.
- This could be a natural stepping stone to a Muon Collider. New physics such as a Z' would make this very compelling.

funding ?



extras



LBNE Collaboration

288 Scientists and
Engineers
57 Institutions
... And still growing!

Alabama: J. Goon, I Stancu

Argonne: M. D'Agostino, G. Drake, Z. Djurcic, M. Goodman, X. Huang, V. Guarino, J. Paley, R. Talaga, M. Wetstein

Boston: E. Hazen, E. Kearns, J. Raaf, J. Stone

Brookhaven: M. Bishai, R. Brown, H. Chen, [M. Diwan](#), J. Dolph, G. Geronimo, R. Gill, R. Hackenberg, R. Hahn, S. Hans, D. Jaffe, S. Junnarkar, J. S. Kettell, F. Lanni, L. Liffenberg, J. Ling, D. Makowiecki, W. Marciano, W. Morse, Z. Parsa, C. Pearson, V. Radeka, S. Rescia, T. Russo, N. Samios, R. Sharma, N. Simos, J. Sondericker, J. Stewart, H. Tanaka, C. Thorn, B. Viren, Z. Wang, S. White, L. Whitehead, M. Yeh, B. Yu

Caltech: R. McKeown, X. Qian, C. Zhang

Cambridge: A. Blake, M. Thomson

Catania/INFN: V. Bellini, G. Garilli, R. Potenza, M. Trovato

Chicago: E. Blucher

Colorado: R. Johnson, A. Marino, M. Tzanov, E. Zimmerman

Colorado State: M. Bass, B. Berger, J. Brack, N. Buchanan, J. Harton, V. Kravtsov, W. Toki, D. Warner, R. Wilson

Columbia: R. Carr, L. Camillieri, C.Y. Chi, G. Karagiorgi, C. Mariani, M. Shaevitz, W. Sippach, W. Willis

Crookston: D. Demuth

Dakota State: B. Szczerbinska

Davis: R. Breedon, T. Classen, J. Felde, P. Gupta, [M. Tripanthi](#), [R. Svoboda](#)

Drexel: C. Lane, J. Maricic, R. Milincic, K. Zbiri

Duke: J. Fowler, J. Prendki, K. Scholberg, C. Walter, R. Wendell

Duluth: R. Gran, A. Habig

Fermilab: D. Allspach, B. Baller, D. Boehnlein, M. Campbell, A. Chen, S. Childress, B. DeMaat, A. Drozhdin, T. Dykhuis, A. Hahn, S. Hays, J. Howell, P. Huhr, J. Hylan, M. Johnson, J. Johnstone, T. Junk, B. Kayser, G. Koizumi, T. Lackowski, P. Lucas, B. Lundberg, T. Lundin, P. Mantsch, N. Mokhov, C. Moore, J. Morfin, B. Norris, V. Papadimitriou, R. Plunkett, C. Polly, S. Pordes, O. Prokofiev, G. Rameika, B. Rebel, D. Reitzner, K. Riesselmann, R. Rucinski, R. Schmidt, D. Schmitz, P. Shanahan, J. Strait, S. Striganov, K. Vaziri, G. Velev, G. Zeller, R. Zwaska

Hawaii: S. Dye, J. Kumar, J. Learned, S. Matsuno, S. Pakvasa, M. Rosen, G. Varner

Indian Universities: V. Singh (BHU); [B. Choudhary](#), S. Mandal (DU); B. Bhuyan [IIT(G)]; V. Bhatnagar, A. Kumar, S. Sahijpal (PU)

Indiana: W. Fox, C. Johnson, M. Messier, S. Mufson, J. Musser, R. Tayloe, J. Urheim

Iowa State: M. Sanchez

IPMU/Tokyo: M. Vagins

Irvine: G. Carminati, W. Kropp, M. Smy, H. Sobel

Kansas State: T. Bolton, G. Horton-Smith

LBL: R. Kadel, B. Fujikawa, D. Taylor

Livermore: A. Bernstein, R. Bionta, S. Dazeley, S. Ouedraogo

London-UCL: J. Thomas

Los Alamos: S. Elliott, V. Gehman, G. Garvey, T. Haines, D. Lee, W. Louis, C. Mauger, G. Mills, A. Norrick, Z. Pavlovic, G. Sinnis, W. Sondheim, R. Van de Water, H. White

Louisiana State: W. Coleman, T. Kutter, W. Metcalf, M. Tzanov

Maryland: E. Blaufuss, R. Hellauer, T. Straszheim, G. Sullivan

Michigan State: E. Arrieta-Diaz, C. Bromberg, D. Edmunds, J. Huston, B. Page

Minnesota: M. Marshak, W. Miller

MIT: W. Barletta, J. Conrad, T. Katori, R. Lanza, L. Winslow

NGA: S. Malys, S. Usman

New Mexico: B. Becker, J. Mathews

Notre Dame: J. Losecco

Oxford: G. Barr, J. DeJong, A. Weber

Pennsylvania: J. Klein, K. Lande, A. Mann, M. Newcomer, R. vanBerg

Pittsburgh: D. Naples, V. Paolone

Princeton: Q. He, K. McDonald

Rensselaer: D. Kaminski, J. Napolitano, S. Salon, P. Stoler

Rochester: R. Bradford, K. McFarland

SDMST: X. Bai, R. Corey

SMU: T. Liu, J. Ye

South Carolina: H. Duyang, S. Mishra, R. Petti, C. Rosenfeld

South Dakota State: B. Bleakley, K. McTaggart

Texas: S. Kopp, K. Lang, R. Mehdiev

Tufts: H. Gallagher, T. Kafka, W. Mann, J. Schnepfs

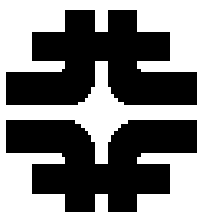
UCLA: K. Arisaka, D. Cline, K. Lee, Y. Meng, F. Sergiampietri, H. Wang

Virginia Tech: E. Guarnaccia, J. Link, D. Mohapatra, R. Raghavan

Washington: S. Enomoto, J. Kaspar, N. Tolich, H.K. Tseung

Wisconsin: B. Balantekin, F. Feyzi, K. Heeger, A. Karle, R. Maruyama, D. Webber, C. Wendt

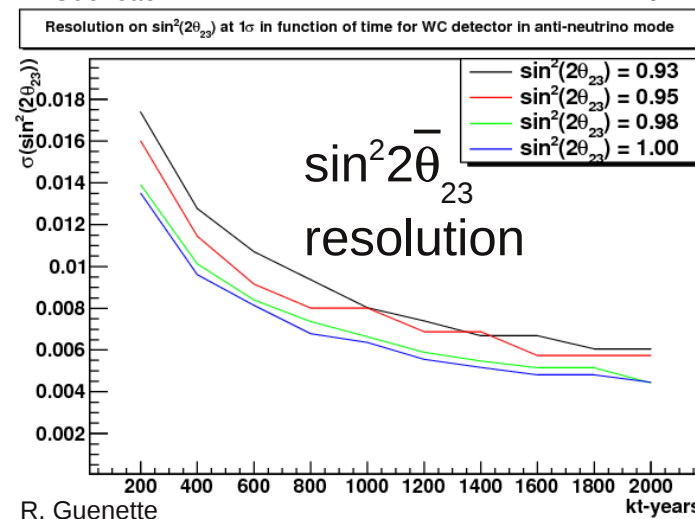
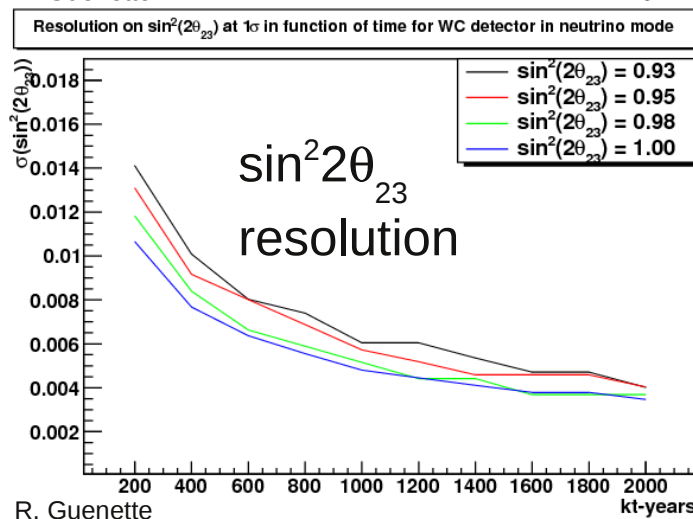
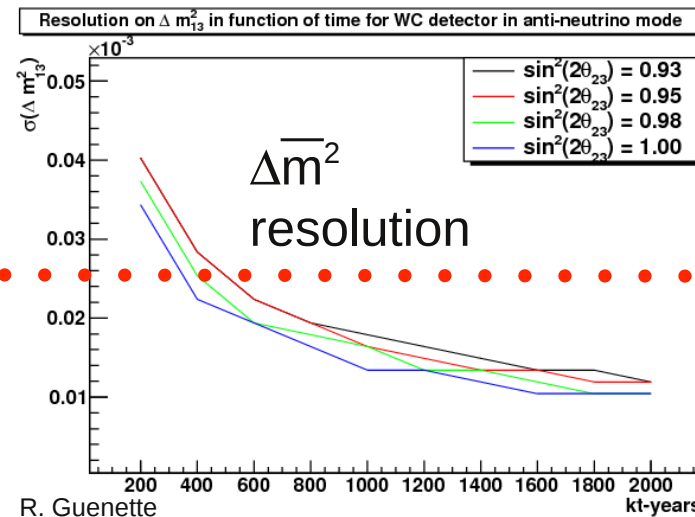
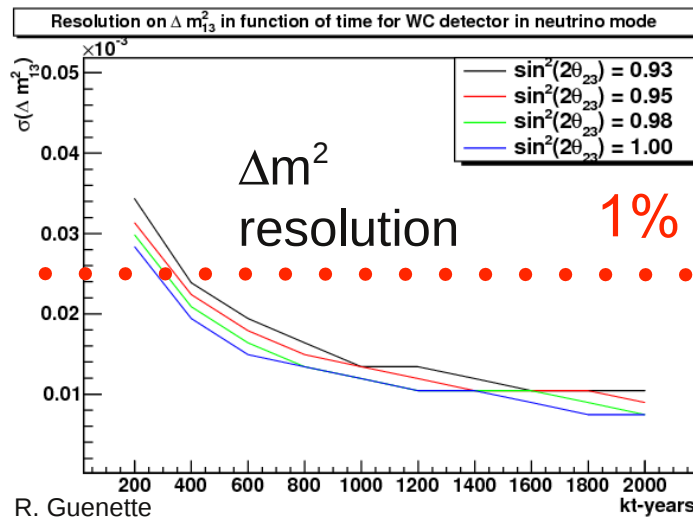
Yale: E. Church, B. Fleming, R. Guenette, M. Soderberg, J. Spitz



Resolution:

What's measured

hep-ph/0503283



$$\delta m_{\mu\mu}^2 = \frac{|U_{\mu 1}|^2 \delta m_{31}^2 + |U_{\mu 2}|^2 \delta m_{32}^2}{|U_{\mu 1}|^2 + |U_{\mu 2}|^2}$$

$$= \delta m_{32}^2 + [\sin^2 \theta_{12} + \mathcal{O}(\sin \theta_{13})] \delta m_{21}^2$$

$$|\delta m_{32}^2| = |\delta m_{\mu\mu}^2| (1 \mp 0.01)$$

1% difference:
sign depends on hierarchy