High-energy emission from a magnetar giant flare in the Sculptor galaxy

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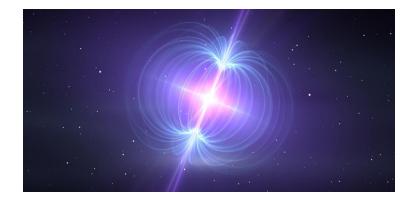
On behalf Fermi-LAT Collaboration

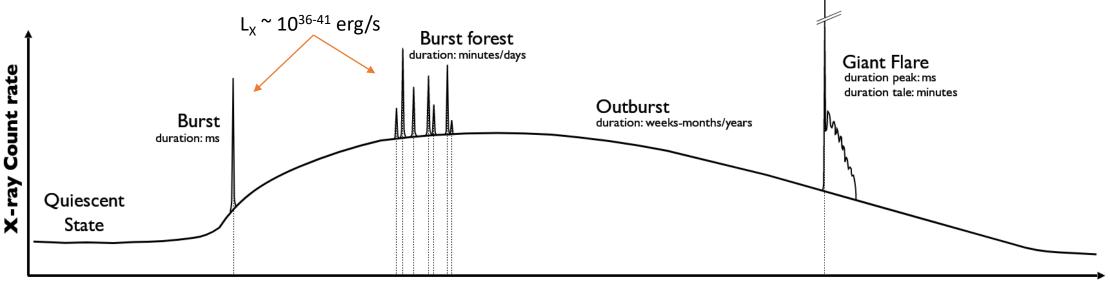




Magnetars

- Strongly magnetized neutron stars:
 - Magnetic field ~ 10^{13-15} G
 - Rotation period of 0.1-10 s
 - Steady X-ray luminosity $L_{\chi} \sim 10^{31-36}$ erg/s

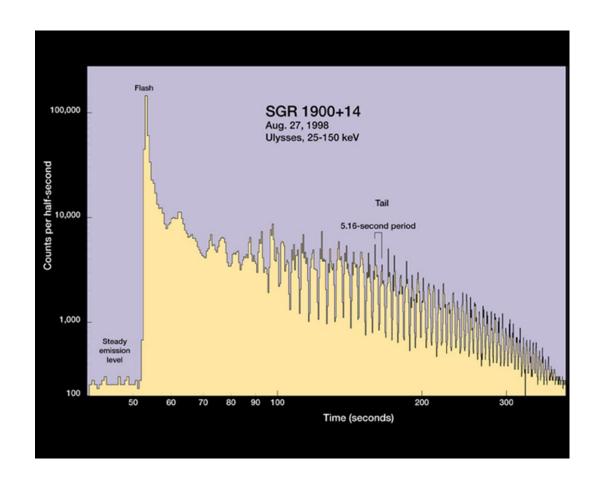




Time

Magnetar Giant Flares

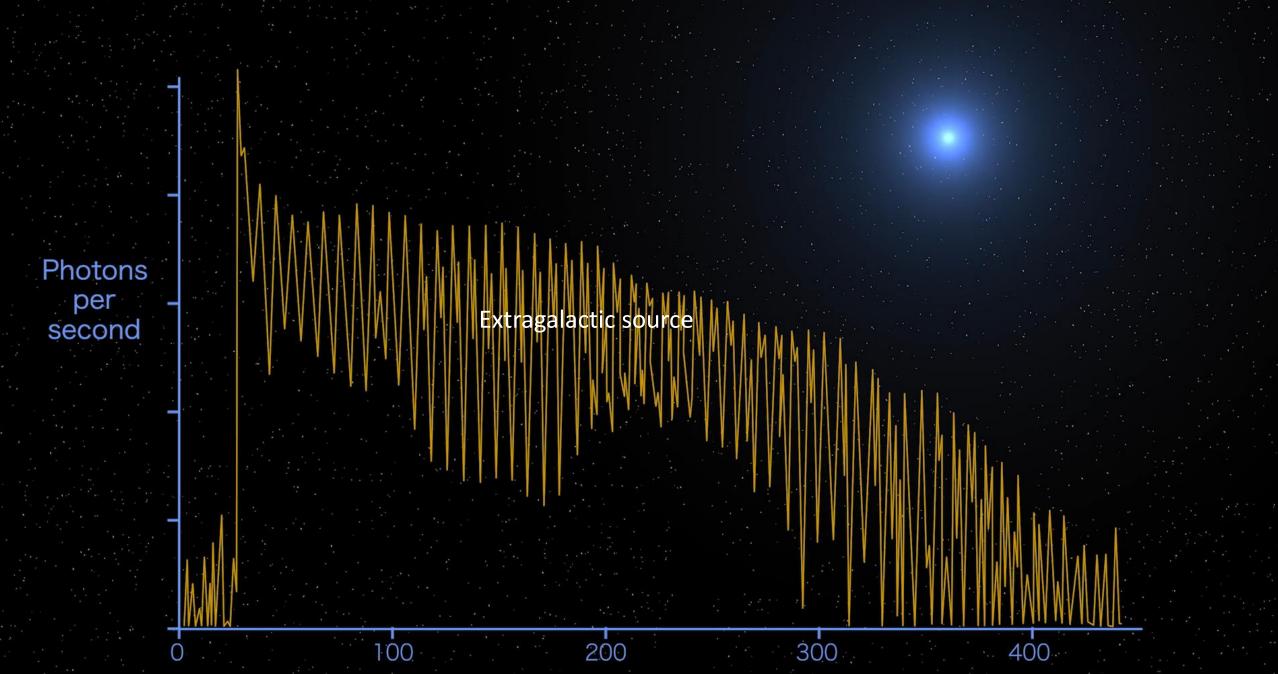
- Rare, short duration bursts of hard X-rays and soft gamma-rays with luminosity ~ 10⁴⁴⁻⁴⁷ erg/s:
 - Bright and variable initial spike lasting a few tenths of a second
 - Dimmer pulsating tail lasting a few hundred of seconds
- Triggered by extreme starquakes:
 - Induced by the extreme magnetic field which causes crustal fractures and the release of hot plasma





Galactic source

Video credit: NASA Goddard



Time in seconds

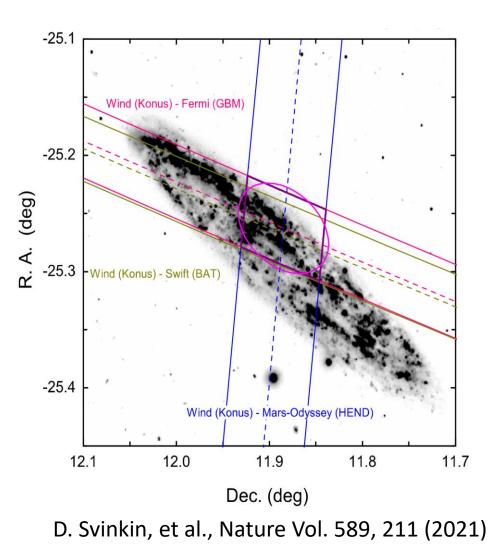
Video credit: NASA Goddard



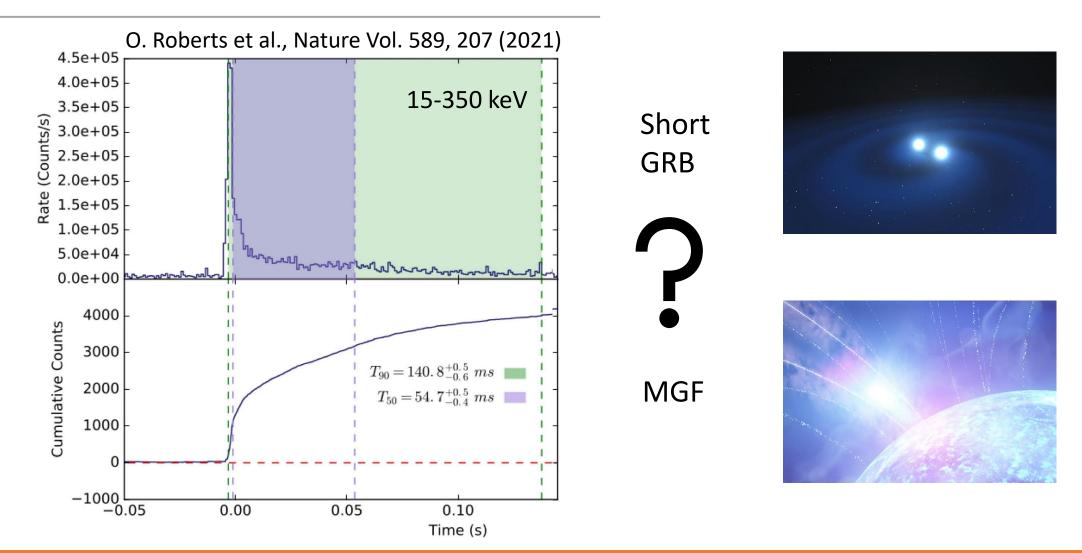
Video credit: NASA Goddard

GRB 200415A

- Bright transient on April 15th 2020:
 - GBM triggered at 08:48:05.56 UTC
 - O. Roberts et al., Nature Vol. 589, 207 (2021)
 - Localized by the Inter-Planetary Network in a 17 arcmin² region overlapping with NGC 253,
 - D. Svinkin, et al., Nature Vol.589, 211 (2021)
 - Active star-bursting spiral galaxy at a distance of 3.5 Mpc
- Chance coincidence with NGC 253: 1 in 230,000
 - E. Burns et al., ApJL 907 L28 (2021)

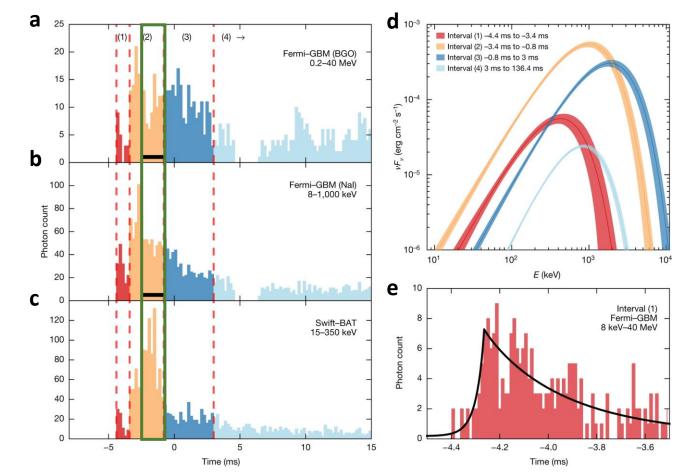


Swift BAT light curve



Fermi GBM observation

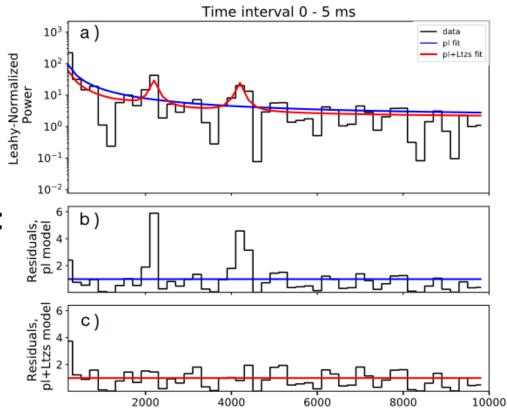
- Flux and spectral shape unusual for a short GRB:
 - 77 µs rise time
 - Sub-ms spectral evolution
 - Flat spectrum ($\alpha \sim 0$), Ep \sim MeV
 - 180 Hz QPO at 2.5σ in the burst decay
- Very bright:
 - Liso = $1.1 \times 10^{47} \text{ erg/s}$
 - Eiso = $1.5 \times 10^{46} \text{ erg}$
 - Highest energy photon: 3 MeV
- No radio counterpart (VLA) or GW emission (KAGRA)



O. Roberts et al., Nature Vol. 589, 207 (2021)

ASIM observation

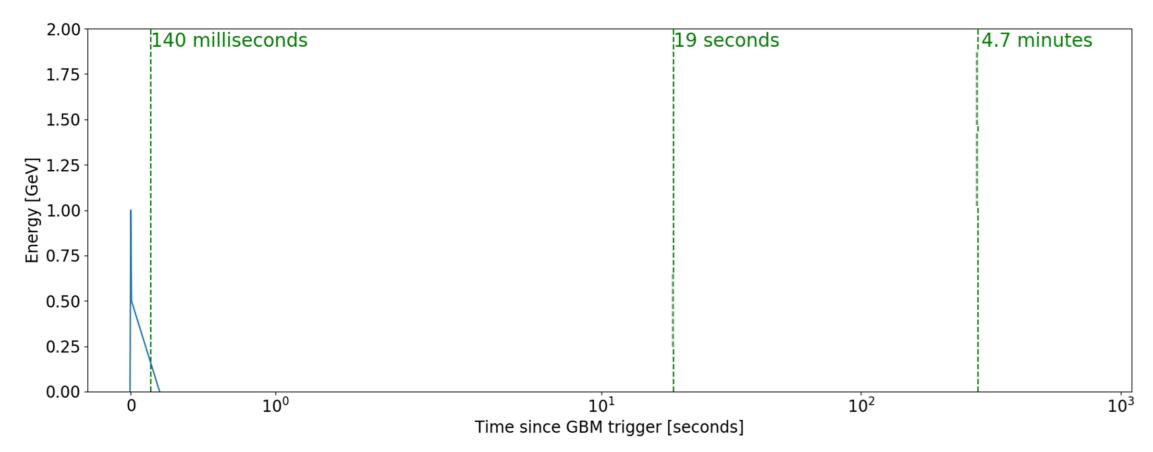
- ASIM Modular X- and Gamma-ray Sensor:
 - LED (50-400 keV) and HED (400 keV-40 MeV)
- Not suffering from saturation effects:
 - Recorded fine structure of the first burst (0.8-3.2 ms) with μs time resolution
- Two significant QPOs during the first peak:
 - 2156 ± 150 Hz (chance probability ~ 10⁻⁹)
 - 4256 ± 323 Hz (likely first harmonic)
 - Confirmed using SWIFT BAT data
- Alfven waves in the magnetosphere during the magnetic reconnection



A. J. Castro-Tirado et al., Nature Vol. 600 (2021)

Fermi LAT light curve

GRB200415A was well within the LAT FoV until 500s after the GBM emission



A peculiar LAT GRB?

Fermi LAT Collaboration, Nature Astronomy Vol. 5 (2021) 170127C 160702A RB200415A short 104 103 Т_{LAT,0} (0.1-100 GeV) [s] 10² 101 10⁰ 10-10-1 10° 10¹ 10² 10³ TGBM. 95 (50-300 keV) [s]

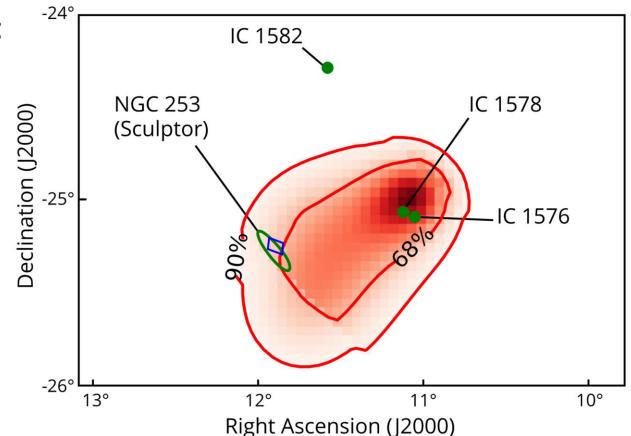
GRB 200415A is the only LAT sGRB within the FoV that was detected much later

the end of the GBM prompt emission

LAT Localization map

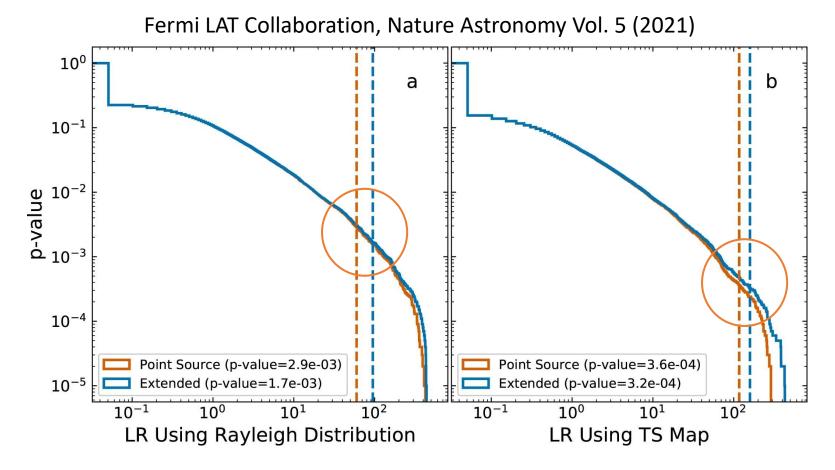
- Likelihood analysis and TS map:
 - Maximum TS = 29 at RA = 11.13°, dec = -24.97°
 - 4 NGC 2000 galaxies in the 3° x 3° ROI
 - NGC 253 at 72% localization CL
- L.R. association results:

Galaxy	IC	IC	IC	NGC
	1576	1578	1582	253
L.R.	2.1	2.9	0.3	60



Fermi LAT Collaboration, Nature Astronomy Vol. 5 (2021)

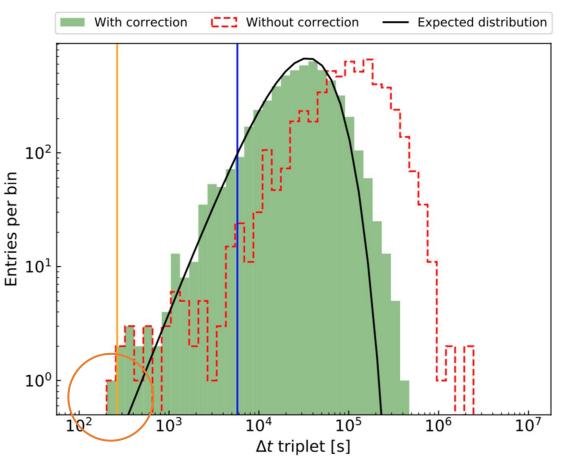
Spatial association



False Alarm Rate ~ 1 chance coincidence in 200-1800 years

Temporal association

- Significance 5.3 σ :
 - 3 events in ~ 300 seconds in the Sculptor region [Li&Ma, 1983]
- $\Delta t_i = t_{i+2} t_i$
- Expected distribution:
 - Obtained using the Poisson statistics
- In 12 years of LAT data:
 - 3 years of livetime
 - Only 1 triplet has a smaller Δt (TS=16)
- Probability of chance coincidence with GBM signal: 1 in million years



Fermi LAT Collaboration, Nature Astronomy Vol. 5 (2021)



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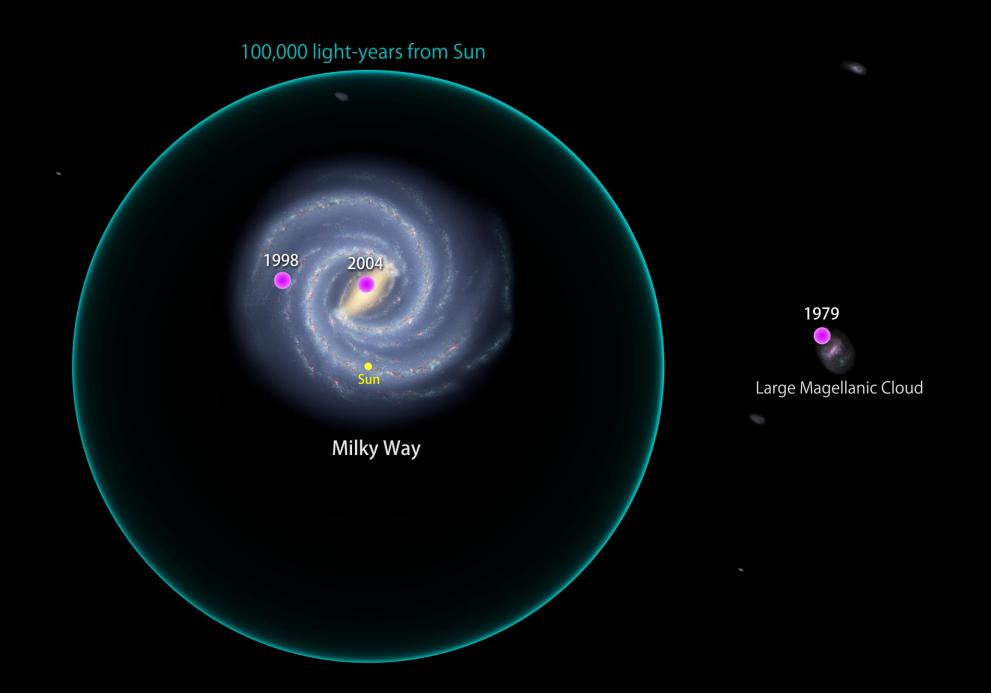
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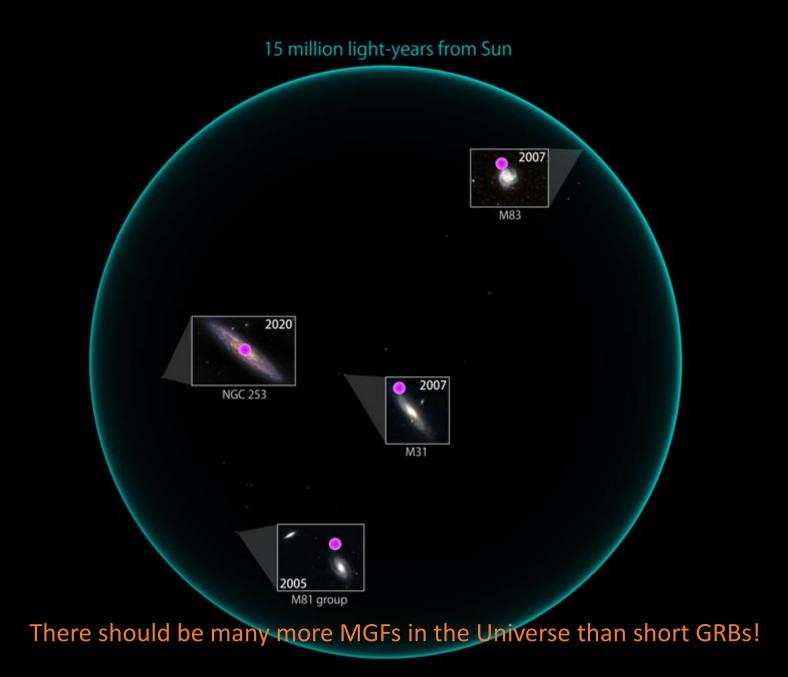
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Case closed?

- Not yet:
 - Clear detection of tale pulsations is needed to absolutely prove magnetar origin
 - Repeating burst would ensure a not cataclysmic event
- But clear fingerprint at the crime scene:
 - Burst morphology (in star-forming galaxy):
 - Initial spike, spectral evolution and properties, QPOs
 - Absence of gravitational waves yet so nearby
- With an unexpected discovery:
 - Delayed high energy gamma rays seen by the LAT
 - Do all MGF produce GeV emission? Is the delay a constant?

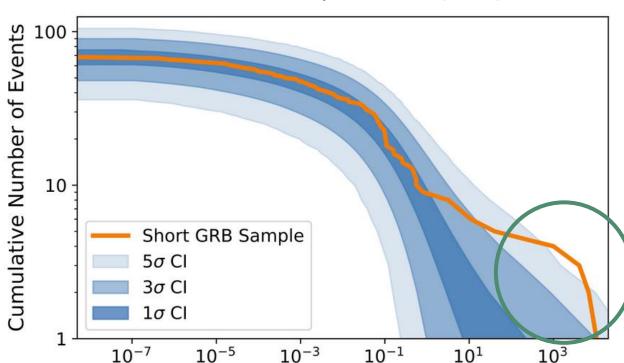






Local population of extragalactic GRBs

- New class of local GRBs:
 - 4 events deviate from the background at more than 5σ
- Common characteristics:
 - Very short rise time (a few ms)
 - Intrinsic energetic (10⁴⁶ erg), much fainter than cosmological GRBs (~10⁵⁰ erg)
 - Very hard spectrum (α ~ 0) and high peak energy (> MeV)
 - Alignment to nearby starforming galaxies



Likelihood that a given sGRB has an MGF origin

E. Burns et al., ApJL 907 L28 (2021)

MGF intrinsic rate

- MGFs are the dominant gammaray transient detected from extragalactic sources:
 - Some magnetars could produce multiple MGFs
 - Repeating GRBs?
- Additional MGF candidates may be identified in the GBM and LAT dataset:
 - Analysis still ongoing!

Event	Local Rates (Gpc ⁻³ yr ⁻¹)	Identified events
Magnetar Giant Flares	380,000	7
Neutron Star Mergers (short GRBs)	320ª	~ 2000
Collapsars (long GRBs)	~100 ^b	~10,000
Type la Supernovae	30,100 ^d	~15,000 ^e
Core-Collapse Supernovae	~70,000 ^d	~ 8000 ^e

E. Burns et al., ApJL 907 L28 (2021)

Integral rates				
< 10 Mpc	< 25 Mpc			
Predicted MGF Detections over GBM Lifetime				
$5.6^{+5.8}_{-4.5}$	15^{+16}_{-12}			

Selected Fermi GI proposal

Conclusions

- Fermi LAT Collaboration reported the high-energy detection of a magnetar giant flare coming from NGC 253:
 - The first detection at GeV energies!
- Simple physical model can explain the observations
- MGFs may constitute a fraction of current short GRB samples:
 - Search for additional MGF candidates in Fermi data still ongoing
- Further details:
 - LAT observation, Fermi LAT Collaboration, Nature Astronomy [link]
 - IPN localization, Svinkin et al., Nature [link]
 - GBM/Swift results, Roberts et al. Nature [link]
 - ASIM observation, A.J. Castro-Tirado et al., Nature [link]
 - Population of Magnetar Giant Flares, Burns et al., ApJL [link]