## Ab initio calculations of complex nuclei

Fiera di Primiero
26-30 September 2022

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## Programme

1. Ab initio description of nuclei
2. Exact many-body methods
3. Expansion many-body methods for closed-shell nuclei
4. Expansion many-body methods for open-shell nuclei

## Part 1

## Ab initio description of nuclei

## Diversity of nuclear phenomena

## Ground state

Mass, size, superfluidity, ...


Strongly-correlated systems
Angular corr. $\rightarrow$ Deformation
Pairing corr. $\rightarrow$ Superfluidity
Quartet corr. $\rightarrow$ Clustering

## Spectroscopy

Excitation modes


Radioactive decays
$\beta, 2 \beta, \alpha, p, 2 p$, fission, ...


Exotic structures
Clusters, halos, ...


Several scales at play Nucleon momenta $\boldsymbol{\sim} \mathbf{1 0 0} \mathrm{MeV}$ Separation energies $\boldsymbol{\sim} \mathbf{1 0} \mathbf{~ M e V}$
Vibration modes $\boldsymbol{\sim} \mathbf{1} \mathrm{MeV}$
Rotation modes $\sim 0.01$-few MeV

## Reaction processes

Fusion, transfer, knockout, ...


## Which is the most appropriate theoretical description?

$\odot$ Richness of nuclear phenomena propelled the formulation of many models


## Which is the most appropriate theoretical description?

๑ Modern view: effective (field) theories


1. Separation of scales $\rightarrow$ Definition of d.o.f.
2. Most general dynamics $\rightarrow$ All allowed terms
3. Organisation $\rightarrow$ Power counting
4. Truncation \& fit of interaction strengths
$\Rightarrow$ Systematically improvable
$\Rightarrow$ Internal consistency check

Possible choices as d.o.f.

Quarks \& gluons

## Nuclei from lattice QCD

$\odot$ First option: compute directly nuclear observables
$\times$ Noise-to-signal ratio of $A$-nucleon correlation functions scales as $e^{A\left(M_{N}-\frac{3}{2} m_{\pi}\right) t}$
$\checkmark$ Could provide highly useful benchmarks

$\odot$ Second option: compute NN (\& NNN) potential
$\times$ Unphysical pion masses
$\times$ Difficult to extend to 3-body forces
$\checkmark$ Extremely useful if extended to hyperons
[Ishii et al. 2007]


## Which is the most appropriate theoretical description?

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## Ab initio nuclear many-body problem

Goal: solve $A$-body Schrödinger equation (for any $A=Z+N$ )
$A$-body wave function
many-nucleon Hamiltonian


A-body energies of ground and excited states

1. Model interactions between nucleons
a) Model the form of $H$
b) Fit coupling constants in $H$
c) Pre-process $H$
input -----------
$\qquad$
feedback
2. Solve many-body Schrödinger eq.
a) Formulate many-body approach
b) Implement, benchmark, optimise
c) Run calculations
$\Rightarrow$ Difficult formal and computational tasks

- Automatised algebraic derivations
- Techniques from applied maths
- High-performance computing



## One-boson exchange potentials

$\bigcirc$ Yukawa potential: nuclear force mediated by massive spin-0 boson (the "mesotron" $\rightarrow$ later, pion)

Yukawa potential

$$
V(r) \propto \frac{e^{-m r}}{r}
$$

$$
\mathrm{m} \sim 100 \mathrm{MeV} \leftarrow \mathrm{r} \sim 2 \mathrm{fm}
$$

Range $\sim$ Compton wavelength of exchanged boson $\sim 1 / \mathrm{m}$
$\odot$ OBE potentials: mesons with larger masses $(\rho, \omega, \sigma)$ can model ranges smaller than $1 / \mathrm{m}_{\pi}$

- Different spin/isospin structures generated
- Additional phenomenological terms

$\checkmark$ High precision $\rightarrow \chi^{2} \approx 2$ in the 1980's, $\chi^{2} \approx 1$ in the 1990's $x$ Hard repulsive core $\rightarrow$ strong (short-range) correlations $x$ Phenomenological component $\rightarrow$ model dependence



## Chiral effective field theory

$\odot$ Chiral EFT: a systematic framework to construct $A \mathrm{~N}$ interactions ( $A=2,3, \ldots$ )

- Expansion around $\mathrm{Q} \sim \mathrm{m}_{\pi} \rightarrow$ d.o.f.: nucleons and pions
- Interactions organised according to power counting
- Many-body forces / currents consistently derived
- Theoretical error assigned to each order

Apply to the many-nucleon system (and propagate the theoretical error)

[Epelbaum et al. 2015, 2020]

## Accuracy of chiral potentials

$\odot$ Accuracy of chiral potentials steadily improving


Rms deviations approaching phenomenological approaches
$\circ$ Ground-state energies $\rightarrow$ rms deviation around $3 \mathrm{MeV}(\sim \mathbf{1 - 1 . 5 \%}$ ) (cf. $\sim 1 \mathrm{MeV}$ in energy density functionals)
$\circ$ Charge radii $\rightarrow$ rms deviation around $0.02 \mathrm{fm}(\sim \mathbf{0 . 5 - 1} \%)$
(similar in energy density functionals)

## Part 2

Exact many-body methods

## Many-body Schrödinger equation

© Goal: solve $A$-body Schrödinger equation (for any $A$ )

$A$-body energies of ground and excited states

Other observables $\leftarrow$ Expectation value of any operator


○ Only input

$$
H=H_{\mathrm{int}}=T_{\mathrm{int}}+V_{\mathrm{NN}}+V_{3 \mathrm{~N}}+\ldots
$$

- Given as a sum of many operators in momentum space ( $\otimes$ spin \& isospin)
- Transformed into basis of choice (e.g. harmonic oscillator)
- Typically truncated at $3 N$ level


## Coordinate-space vs configuration-space methods

## ○ Coordinate-space methods

- Directly work with many-body wave function (e.g. Monte Carlo sampling)



## Coordinate-space vs configuration-space methods

## © Coordinate-space methods

- Directly work with many-body wave function (e.g. Monte Carlo sampling)
$\circ$ Discretise the problem on a lattice $\rightarrow$ Nuclear Lattice Effective Field Theory



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© Coordinate-space methods

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$\checkmark$ Flexible (any spatial configuration is accessible) + no intensive memory requirement
$x$ Sign problem $\rightarrow$ constrained choice of $H+$ expensive in processor time


## Coordinate-space vs configuration-space methods

© Coordinate-space methods

- Directly work with many-body wave function (e.g. Monte Carlo sampling)
- Discretise the problem on a lattice $\rightarrow$ Nuclear Lattice Effective Field Theory
$\checkmark$ Flexible (any spatial configuration is accessible) + no intensive memory requirement
$\times$ Sign problem $\rightarrow$ constrained choice of $H+$ expensive in processor time
$\bigcirc$ Configuration-space methods
- Expand eigenstates on a basis of known many-body states
$\checkmark$ Universally applicable to any $H+$ amenable to controlled approximations
$x$ Expensive in memory usage + constrained by the properties of basis states


## One-body (= single-particle) basis

- Basic constituents: nucleons characterised by position, spin and isospin
- Single-nucleon states expressed as

$$
\left|\varphi_{k}\right\rangle=\left[\left|\varphi_{k}^{\text {space }}\right\rangle \otimes\left|\varphi_{k}^{\text {spin }}\right\rangle\right] \otimes\left|\varphi_{k}^{\mathrm{isospin}}\right\rangle
$$

- Standard choice for nuclear structure approaches

$$
\begin{array}{ll}
\left|\varphi_{k}^{\text {space }}\right\rangle=\left|n \ell m_{\ell}\right\rangle & \text { e.g., solutions of one-body harmon } \\
\left|\varphi_{k}^{\text {spin }}\right\rangle=\left|s m_{s}\right\rangle=\left|\frac{1}{2} m_{s}\right\rangle & \text { eigenstates of } \mathrm{s}^{2} \text { and } \mathrm{s}_{z} \text { with } \mathrm{s}=1 / 2 \\
\left|\varphi_{k}^{\text {isospin }}\right\rangle=\left|t m_{t}\right\rangle=\left|\frac{1}{2} m_{t}\right\rangle & \text { eigenstates of } \mathrm{t}^{2} \text { and } \mathrm{t}_{\mathrm{z}} \text { with } \mathrm{t}=1 / 2
\end{array}
$$

$\odot$ Orbital angular momentum and spin are typically coupled

$$
\left|\varphi_{k}\right\rangle=\left|n\left(\ell \frac{1}{2}\right) j m ; \frac{1}{2} m_{t}\right\rangle=\sum_{m_{l}, m_{s}} c\left(\begin{array}{cc|c}
\ell & \frac{1}{2} & j \\
m_{l} & m_{s} & m
\end{array}\right)\left|n \ell m_{\ell}\right\rangle \otimes\left|\frac{1}{2} m_{s}\right\rangle \otimes\left|\frac{1}{2} m_{t}\right\rangle
$$

## Many-body basis

$\odot$ When dealing with fermions, many-body states have to be explicitly antisymmetrised
Antisymmetrisation operator $\mathcal{A}=\frac{1}{A!} \sum_{\pi} \operatorname{sgn}(\pi) P_{\pi}$
Direct product of A 1-body states

$$
\begin{aligned}
\left|\Phi^{A}\right\rangle & =\mathcal{A}\left\{\left|\varphi_{k_{1}}\right\rangle \otimes\left|\varphi_{k_{2}}\right\rangle \otimes \cdots \otimes\left|\varphi_{k_{A}}\right\rangle\right\} \\
& =\frac{1}{\sqrt{A!}} \sum_{\pi} \operatorname{sgn}(\pi) P_{\pi}\left(\left|\varphi_{k_{1}}\right\rangle \otimes\left|\varphi_{k_{2}}\right\rangle \otimes \cdots \otimes\left|\varphi_{k_{A}}\right\rangle\right) \\
& \equiv\left|k_{1} k_{2} \cdots k_{A}\right\rangle
\end{aligned}
$$

$\circ$ Antisymmetric under exchange $P_{i j}\left|\cdots k_{i} \cdots k_{j} \cdots\right\rangle=\left|\cdots k_{j} \cdots k_{i} \cdots\right\rangle=-\left|\cdots k_{i} \cdots k_{j} \cdots\right\rangle$

- Encodes Pauli principle $\left|\cdots k_{i} \cdots k_{i} \cdots\right\rangle=0 \quad \rightarrow$ minimal intrinsic correlations
$\odot$ Any antisymmetric state can be expanded in the Slater determinant basis

$$
\left|\Psi^{A}\right\rangle=\sum_{k_{1}>k_{2} \cdots>k_{A}} c_{k_{1} k_{2} \ldots k_{A}}\left|k_{1} k_{2} \cdots k_{A}\right\rangle \equiv \sum_{i} c_{i}\left|\Phi_{i}\right\rangle
$$

## Configuration interaction

○ The strategy is the following

1. Select a one-body basis

$$
|\alpha\rangle \equiv\left|n \ell j m m_{t}\right\rangle
$$

2. Construct $A$-body basis of Slater determinants

$$
\left|\Phi_{i}\right\rangle \equiv\left|\left\{\alpha_{1} \alpha_{2} \ldots \alpha_{A}\right\}_{i}\right\rangle
$$

3. Convert Schrödinger equation into a matrix eigenvalue problem

$$
\begin{aligned}
& H\left|\Psi_{k}\right\rangle=E_{k}\left|\Psi_{k}\right\rangle \quad \rightarrow \text { expand }\left|\Psi_{k}\right\rangle=\sum_{i} C_{i}^{(k)}\left|\Phi_{i}\right\rangle \\
& \left\langle\Phi_{j}\right| \times\left[H \sum_{i} C_{i}^{(k)}\left|\Phi_{i}\right\rangle=E_{k} \sum_{i} C_{i}^{(k)}\left|\Phi_{i}\right\rangle\right] \\
& \sum_{i} \underbrace{\left\langle\Phi_{j}\right| H\left|\Phi_{i}\right\rangle}_{\equiv H_{j i}} C_{i}^{(k)}=E_{k} \sum_{i} C_{i}^{(k)} \underbrace{\left\langle\Phi_{j} \mid \Phi_{i}\right\rangle}_{=\delta_{i j}}
\end{aligned}>\left[\begin{array}{c}
\vdots \\
\ldots H_{j i} \\
\vdots \\
\vdots
\end{array}\right]\left[\begin{array}{c}
\vdots \\
C_{i}^{(k)} \\
\vdots
\end{array}\right]=E_{k}\left[\begin{array}{c}
\vdots \\
C_{j}^{(k)} \\
\vdots
\end{array}\right]
$$

## Model space truncations

© Expansion on Slater determinants involves an infinite number of basis states

$$
\left|\Psi_{k}\right\rangle=\sum_{i=1}^{\infty} C_{i}^{(k)}\left|\Phi_{i}\right\rangle \quad \Rightarrow \quad\left|\Psi_{k}(D)\right\rangle=\sum_{i=1}^{D} C_{i}^{(k)}\left|\Phi_{i}\right\rangle
$$

obviously cannot store an infinite vector...
$\Rightarrow$ truncations have to be necessarily introduced
© Two main ways of truncating the basis

- Full CI: truncate the one-body basis (at some maximum single-particle energy $\mathbf{e}_{\max }$ )
- No-core shell model: cut the many-body basis (total number of HO excitation quanta $N_{\max }$ )


$$
\text { Example: } N_{\max }=6
$$

## Computational strategy

© Involved computational problem as A increases
© Key features

- One is only interested in a few low-lying eigenstates
- Hamiltonian matrix is sparse ( $<0.01 \%$ of non-zeros at working values of $N_{\max }$ )

© Computational solutions \& limitations
- Lanczos-type algorithms employed to extract first few eigenstates and associated eigenvalues
- Fast storage of non-zero matrix elements sets the limits of matrix dimensions
- Extensive use of parallelisation, matrix transformations, optimisation techniques, ...


## CI dimensionality

$\bigcirc$ "Back-of-the-envelope" estimate of matrix dimensions

- Case of Full CI (recall: truncation acts on the single-particle basis)
© How many Slater determinants can be built from a given number of single-particle states?
- Take $A$ nucleons and $n$ single-particle states
$\Rightarrow$ Number of different possible Slater determinants $\quad\binom{n}{A}=\frac{n!}{(n-A)!A!}$
© Example: ${ }^{16} \mathbf{O}(Z=8, N=8)$ in 40 single-particle states
$\binom{40}{8}=\frac{40!}{(40-8)!8!} \approx 8 \cdot 10^{7} \quad$ for protons $\quad x \quad\binom{40}{8}=\frac{40!}{(40-8)!8!} \approx 8 \cdot 10^{7} \quad$ for neutrons
c) Total of $\mathrm{D}=6 \cdot 10^{15}$ Slater determinants
$\Rightarrow$ Number of non-zero matrix elements ( $N N$ only!) scales as D ${ }^{1.2} \rightarrow \sim 10^{18}$ non-zero entries
$\Rightarrow$ Size in memory beyond EB $\rightarrow$ well beyond current capabilities
$\odot$ Current computational limits for the storage and diagonalisation of a large matrix
- Petascale machines: D ~ $10^{10}$ / / Exascale machines: D ~ $10^{12}$


## NCSM dimensionality

## © No-core shell model

- More gentle scaling (recall: truncation $N_{\max }$ acts on the many-body basis)


Convergence w.r.t. $N_{\text {max }}$

$\Rightarrow$ Computational limits quickly reached

## NCSM dimensionality



## Short-range correlations \& "low-momentum" interactions

$\odot$ Why do we need to include such high values of $N_{\max } /$ large matrix dimensions?
$\odot$ Nuclear interactions generate short-range correlations in many-body states

- Traditionally linked to "hard core" of one-boson exchange potentials
- Weaker but present in modern chiral interactions
$\circ$ Short distance / high momenta / high energy $\rightarrow$ large Hilbert space needed
$\bigcirc$ Idea: use unitary transformations on $\boldsymbol{H}$ to suppress these correlations

$\rightarrow$ Decouple low- and high-momenta
$\rightarrow$ Can work in small Hilbert space
$\rightarrow$ Less correlated wave functions
$\rightarrow$ Observables unchanged!

$$
U^{\dagger} H U U^{\dagger}|\Psi\rangle=E U^{\dagger}|\Psi\rangle
$$

$$
\ldots \quad \tilde{H}|\tilde{\Psi}\rangle=E|\tilde{\Psi}\rangle
$$

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$\rightarrow$ Many-body forces generated
$\leftrightarrows$ Similarity renormalisation group (SRG) transformation


## A matter of resolution


[figures from K. Hebeler]

## SRG transformation



SRG transformation
[Figures: R. Roth]


## chiral NN

Entem \& Machleidt. N ${ }^{3}$ LO, 500 MeV

$$
J^{\pi}=1^{+}, T=0
$$

deuteron wave-function


SRG transformation
[Figures: R. Roth]


$$
\underset{\substack{\wedge=1.58 \mathrm{fm}^{-1}}}{\alpha=0.160 \mathrm{fm}^{4}}
$$

deuteron wave-function


## SRG transformation

3B-Jacobi HO matrix elements

$(E, i)$

## chiral NN+3N

$$
J^{\pi}=\frac{1}{2}^{+}, T=\frac{1}{2}, \hbar \Omega=28 \mathrm{MeV}
$$

## NCSM ground state ${ }^{\mathbf{3}} \mathrm{H}$



## SRG transformation




## SRG in A-body systems

$\bigcirc$ Effect of induced many-body forces is non-negligible already in small systems


Initial ("genuine") 4-body forces assumed to be very small

- $\lambda$-dependence provides estimate of neglected induced 4-body contributions in ${ }^{4} \mathrm{He}$


## SRG in A-body systems

$\bigcirc$ Example: no-core shell model calculations of ${ }^{4} \mathbf{H e}$ and ${ }^{6} \mathrm{Li}$ ground-state energies

Flow parameters [ $\mathrm{fm}^{-1}$ ]



## NCSM dimensionality



## Normal-ordered two-body approximation

© From original Hamiltonian (normal-ordered w.r.t. the particle vacuum)...

$$
\begin{gathered}
H=\sum_{p q} t_{p q} c_{p}^{\dagger} c_{q}+\frac{1}{(2!)^{2}} \sum_{p q r s} v_{p q r s} c_{p}^{\dagger} c_{q}^{\dagger} c_{s} c_{r}+\frac{1}{(3!)^{2}} \sum_{p q r s t u} w_{p q r s t u} c_{p}^{\dagger} c_{q}^{\dagger} c_{r}^{\dagger} c_{u} c_{t} c_{s} \\
\\
\text { introduce Slater determinant }\left|\phi_{0}\right\rangle=\prod_{i=1}^{A} a_{i}^{\dagger}|0\rangle
\end{gathered}
$$

... to a Hamiltonian normal-ordered w.r.t. to a reference Slater determinant

$$
H=h^{(0)}+\sum_{p q} h_{p q}^{(1)}: a_{p}^{\dagger} a_{q}:+\frac{1}{2!} \sum_{p q r s} h_{p q r s}^{(2)}: a_{p}^{\dagger} a_{q}^{\dagger} a_{s} a_{r}:+\frac{1}{6!} \sum_{p q r s t u} h_{p q r s t u}^{(3)}: a_{p}^{\dagger} a_{q}^{\dagger} a_{r}^{\dagger} a_{u} a_{t} a_{s}:
$$

© Define density matrix \& occupation numbers

$$
\rho_{p q} \equiv\left\langle\phi_{0}\right| a_{p}^{\dagger} a_{q}\left|\phi_{0}\right\rangle=n_{p} \delta_{p q} \quad \rightarrow \quad \begin{cases}n_{i}=1 & \text { holes } \\ n_{a}=0 & \text { particles }\end{cases}
$$

## Normal-ordered two-body approximation

$\odot$ Normal-ordered matrix elements

$$
\begin{aligned}
& h^{(0)}=\sum_{i} t_{i i} n_{i}+\frac{1}{2} \sum_{i j} v_{i j i j} n_{i} n_{j}+\frac{1}{6} \sum_{i j k} w_{i j k i j k} n_{i} n_{j} n_{k} \\
& h_{p q}^{(1)}=t_{p q}+\sum_{i} v_{p i q i} n_{i}+\frac{1}{2} \sum_{i j} w_{p i j q i j} n_{i} n_{j} \\
& h_{p q r s}^{(2)}=v_{p q r s}+\sum_{i} w_{p q i r s i} n_{i} \longleftarrow \begin{array}{c}
\text { Large part of the original 3N transferred } \\
\text { into effective lower-rank operators }
\end{array} \\
& h_{p q r s t u}^{(3)}=w_{\text {pqrstu }} \\
& \begin{array}{c}
\text { Normal-ordered 2-body approximation (NO2B) } \\
\rightarrow \text { Discard residual 3N operator }
\end{array}
\end{aligned}
$$

## Normal-ordered two-body approximation

$\odot$ Normal-ordered matrix elements

$$
\begin{aligned}
& h^{(0)}=\sum_{i} t_{i i} n_{i}+\frac{1}{2} \sum_{i j} v_{i j i j} n_{i} n_{j}+\frac{1}{6} \sum_{i j k} w_{i j k i j k} n_{i} n_{j} n_{k}= \\
& h_{p q}^{(1)}=t_{p q}+\sum_{i} v_{p i q i} n_{i}+\frac{1}{2} \sum_{i j} w_{p i j q i j} n_{i} n_{j}=,
\end{aligned}
$$

$$
h_{p q r s t u}^{(3)}=\text { wpqrstu }^{\text {pa }}
$$

## Normal-ordered 2-body approximation (NO2B)

$\rightarrow$ Discard residual 3N operator

- Benchmarked in light nuclei
- 1-3\% error
- Comparable to other errors
[Roth et al. 2012]



## NCSM dimensionality



## Importance truncation

© Not all matrix elements of $H$ are equally relevant

- $N_{\max }$ cuts might not be the most efficient way of selecting important entries
$\circ$ Is there a way of discarding a priori the most irrelevant entries for a given $N_{\text {max }}$ ?
$\odot$ Importance truncation: prior to diagonalisation

1. Estimate the size of each entry upon a given criterion
2. Discard irrelevant entries (i.e., make the matrix even more sparse)
$\Rightarrow$ Construct importance-truncated space from all basis states having $\left|\kappa_{\nu}\right| \geq \kappa_{\text {min }}$
$\odot$ Required features:

- Estimate has be done with a cheap method
- Typical tool of choice: many-body perturbation theory
- In the limit of null threshold one must recover the original (exact) problem
- Smooth behaviour desirable in order to perform extrapolations


## Importance truncation

๑ Example: no-core shell model calculation of ${ }^{16} \mathrm{O}$


- Smooth threshold dependence
- Extrapolation to un-truncated result
- Uncertainty quantification from fit
- Benchmarks possible for for small $N_{\max }$



## Applications: oxygen isotopes

$\odot$ First ab initio calculations with $\mathrm{NN}+3 \mathrm{~N}$ chiral interactions along the oxygen chain


- Converged results achieved up to ${ }^{24} \mathrm{O}$
- Unbound ${ }^{26} \mathrm{O}$ harder to compute in HO basis
- Role of "genuine" 3N forces evident

[Hergert et al. 2013]


## Part 3

Expansion many-body methods for closed-shell nuclei

## Correlation expansion methods: the idea

© The goal is always to solve $\quad H\left|\Psi_{k}^{A}\right\rangle=E_{k}^{A}\left|\Psi_{k}^{A}\right\rangle$
© Idea: write the exact ground-state wave function as

then expand and truncate $\Omega_{0}$
$\Rightarrow$ Before truncation, the expansion is exact
$\Rightarrow$ After truncation, cost reduced from $e^{N}$ to $N^{\alpha}$ with $\alpha \geq 4$

- Reference state
- Must be simple enough (such that it can be computed easily and exactly)
- Must be rich enough (such that it is a suitable starting point for the expansion)
- Obtained by

1) Splitting $H=H_{0}+H_{1}$
2) Solving for $H_{0}$ (one-body operator) $H_{0}\left|\phi_{k}\right\rangle=\epsilon_{k}\left|\phi_{k}\right\rangle$

## Mean field

$\bigcirc$ Independent-particle picture
$\circ$ One-body potential: $H_{0}=\sum_{i=1}^{A} h_{0}(i) \quad \rightarrow \quad H_{0}\left|\phi_{k}\right\rangle=\epsilon_{k}\left|\phi_{k}\right\rangle \quad \Rightarrow \quad \begin{gathered}h_{0}|\alpha\rangle=\varepsilon_{\alpha}|\alpha\rangle \quad \forall i \\ A \text {-body problem }\end{gathered}$

- Build Slater determinant $\left|\phi_{0}\right\rangle=\prod_{i=1}^{A} a_{\alpha_{i}}^{\dagger}|0\rangle \quad A$-body problem $\quad$ A one-body problems
- Nucleons move independently inside a (one-body) potential well or mean field
$\odot$ Does an independent-particle picture make any sense at all?
- Range of nuclear interaction $\approx$ Inter-particle distance in nuclei $\sim 2 \mathrm{fm}$
- However, it looks like it actually does make sense


$\checkmark$ Success of nuclear shell model


## Effective or phenomenological models

## Energy density functionals

$$
H^{\mathrm{eff}}\left|\Psi^{\mathrm{eff}}\right\rangle=E\left|\Psi^{\mathrm{eff}}\right\rangle
$$

Simplified w.f. Compensate for correlations in H
(Beyond) mean field Phenomenological fit
$\checkmark$ Low cost $\rightarrow$ Access whole nuclear chart
$x$ Unclear how to improve (systematically)


## Interacting shell model

$$
H^{\mathrm{eff}}\left|\Psi^{\mathrm{eff}}\right\rangle=E\left|\Psi^{\mathrm{eff}}\right\rangle
$$

Compensate for correlations in H
Full (CI) w.f., but in valence space
Phenomenological fit
$\checkmark$ Very accurate locally in the nuclear chart
$\times$ Limited predictive power + scaling


[Lenzi et al. 2010]

## Hartree-Fock with ab initio interactions

## OBE potentials




Expansion problematic: full diagonalisation needed

Chiral potentials


Expansion possible, but problem non-perturbative

SRG potentials


Expansion simple: even perturbation theory works!

## Correlation expansion: perturbative approach

$\odot$ Expansion of the exact wave function

$$
\begin{aligned}
& \text { Ref } \\
& \text { 1p1h } \\
& \text { 2p2h } \\
& \text { 3p3h }
\end{aligned}
$$

C) Perturbative methods: expansion coefficients computed independently
$\odot$ Standard many-body perturbation theory (MBPT)

- Simple expressions for $E$ at low orders
- Non-iterative calculation
- Polynomial scaling $O\left(N^{\alpha}\right) \rightarrow O\left(N^{4}\right)$ at MBPT(2) level

$$
E^{(2)}=\frac{1}{4} \sum_{a b}^{>\epsilon_{\mathrm{F}}} \sum_{i j}^{<\epsilon_{\mathrm{F}}} \frac{\langle a b| W|i j\rangle\langle i j| W|a b\rangle}{\left(\epsilon_{a}+\epsilon_{b}-\epsilon_{i}-\epsilon_{j}\right)}
$$

## Many-body perturbation theory

## $\odot$ Convergence of MBPT series

- Convergence of the series can be tested up to high orders in small basis (recursive scheme)

$\Rightarrow$ Importance of using the right reference
$\Rightarrow$ Resummation schemes possible (e.g. Padé, eigenvector continuation, ...)


## Many-body perturbation theory

## $\odot$ Reach

- Calculations currently possible up to mass $A \sim 100$ (and beyond)
© Benchmark
[Tichai et al. 2016]
- Accuracy competitive with coupled cluster calculations (non-perturbative and more costly)

$$
\operatorname{MBPT} E_{0}^{(2)}(\mathrm{O}) \quad E_{0}^{(2)}+E_{0}^{(3)}(\bullet)
$$



## Correlation expansion: non-perturbative approach

$\odot$ Expansion of the exact wave function

$$
\begin{aligned}
& \text { Ref } \\
& \text { 1p1h } \\
& \text { 2p2h } \\
& \text { 3p3h }
\end{aligned}
$$

$\Rightarrow$ Perturbative methods: expansion coefficients computed independently
$\Rightarrow$ Non-perturbative methods: expansion coefficients computed self-consistently

- Examples of non-perturbative approaches
- Coupled-cluster theory (CC)
$\Rightarrow$ Exponential ansatz for the wave function $\quad\left|\Psi_{C C}\right\rangle=e^{T}|\Phi\rangle$
- In-medium similarity renormalisation group (IMSRG) $\Rightarrow$ SRG evolution for $H$ normal-ordered w.r.t. to a reference Slater determinant
$\circ$ Self-consistent Green's function (SCGF) [next slide]


## Green's function techniques

© The goal is to solve the $A$-body Schrödinger equation

$$
H\left|\Psi_{k}^{A}\right\rangle=E_{k}^{A}\left|\Psi_{k}^{A}\right\rangle
$$

○ Instead of working with the full $A$-body wave function $\left|\Psi_{k}^{A}\right\rangle$, rewrite the Schrödinger equation in terms of 1-, 2-, .... A-body objects $G_{1}=G, G_{2}, \ldots G_{\mathrm{A}}$ (Green's functions)
$\xrightarrow{\prime \prime} \rightarrow A-1$ coupled equations
© 1-, 2-, .... A-body Green's functions yield expectation values of 1-, 2-, .... A-body operators
$\xrightarrow{\prime} \rightarrow$ In practice, one usually needs 1- and / or 2-body GFs ( $\sim 1-\& 2$-body density matrices)
© One-body Green's function obtained by solving Dyson equation (derived from Schrödinger eq.)

$$
G=G^{(0)}+G^{(0)} \Sigma G
$$

unperturbed Green's function
many-body effects contained in the self-energy $\Sigma$
© Bonus: one-body Green's function contains information about $A \pm 1$ excitation energy spectra
${ }^{\prime \prime} \rightarrow$ Spectral or Lehmann representation of the Green's function

## Benchmarks

Oxygen binding energies

© Convergence of many-body results

- Different strategies to solve $\mathrm{H} \Psi=\mathrm{E} \Psi$
- Same input Hamiltonian (except lattice EFT)
- All methods agree within 5\%


## $\odot$ Physics of oxygen isotopes

- Energy trend reproduced by $2 \mathrm{~N}+3 \mathrm{~N}$ results
- Correct drip line only with 3N forces


## Part 4

Expansion many-body methods for open-shell nuclei

## Closed- vs. open-shell systems

$\odot$ In practice: expand on Slater determinant basis $\rightarrow$ particle-hole ( ph ) expansion


Ref. state varies with $N \& Z$



Closed-shell systems

Larger energy gap, excitations hindered, enhanced stability

Weakly correlated, clear ph hierarchy, expansion well defined


Nucleons partially fill levels below a magic number


Open-shell systems

Smaller ( $\rightarrow 0$ ) energy gap, excitations enabled, lesser stability

Strongly correlated, no ph hierarchy, expansion ill defined

## Breakdown of ph expansion

## Closed-shell



- Breakdown of ph expansion evident already in MBPT(2) expressions
- Can be explicitly demonstrated by artificially decreasing the gap in ${ }^{16} \mathrm{O}$
gap


## Open-shell



$$
\Delta E_{\mathrm{MBPT}}^{(2)}=-\frac{1}{4} \sum_{i j a b} \frac{\left|h_{i j a b}^{(2)}\right|^{2}}{e_{a}+e_{b}-e_{i}-e_{j}}=0
$$



## Symmetry breaking

$\bigcirc$ Idea: reopen gap via symmetry breaking

$\bigcirc$ Which symmetries?
$\circ G_{\text {Ham }} \rightarrow$ symmetries of $H$ usually dictated by QCD + general principles
$\circ G_{\mathrm{wf}} \rightarrow$ symmetries of w.f. depend on a given ansatz
$\circ G_{\mathrm{bas}} \rightarrow$ eigenfunctions of a given operator with certain symmetries (e.g. HO Hamiltonian)

Usually one chooses $G_{\text {Ham }}=G_{\text {wf }}=G_{\text {bas }}$ Symmetry breaking $\rightarrow G_{\text {Ham }} \neq G_{\text {wf }}$
$\odot$ Why should it help?

- Variational space of w.f. is enlarged
- Degeneracy is lifted by deformation $\rightarrow$ Particle-hole expansion again well defined
- We know it works from experience (collective model, energy density functionals)


## Symmetry breaking

© Allowing w.f. to break symmetries is an efficient way to account for strong correlations
Order parameter $\quad\left\langle\Phi_{0}\right| Q\left|\Phi_{0}\right\rangle=q \equiv|q| e^{i \arg (q)}$


Which symmetry for which type of correlation?

| Physical symmetry | Group | Correlations |
| :---: | :---: | :--- | :--- |
| Rotational inv. | $\mathrm{SU}(2)$ | Deformation |
| Particle-number | $\mathrm{U}(1)_{\mathrm{N}} \times \mathrm{U}(1)_{\mathrm{Z}}$ | Superfluidity |$\quad$| Singly open-shell $\Rightarrow$ Sufficient to break U(1) |
| :--- |
| Doubly open-shell $\Rightarrow$ Necessary to break SU(2) |

$\checkmark$ Advantage: polynomially-scaling $\left(N^{\alpha}\right)$ method that can tackle strongly correlated systems
$X$ Prices to pay:

1) $N_{\text {sym-breaking }}>N_{\text {sym-conserving }}$
2) Symmetries must be eventually restored in finite systems

## Symmetry breaking

© Example: U(1)-breaking SCGF calculations

- Description deteriorates when going away from singly open-shell
- Correlation with (expected) deformation observed




## Partition, expand, project

$\odot$ Partition, then expand \& project

1. Compute symmetry-breaking ref. state

$$
\left|\Theta^{0}\right\rangle=\left|\Phi\left(q_{\min }\right)\right\rangle \quad \longrightarrow \quad H=H_{0}+H_{1}
$$

2. Expand in $H_{1}$
3. Restore broken symmetries
$\bigcirc$ Partition \& project, then expand
4. Compute symmetry-breaking states [at many $q$ ]
5. Restore symmetries [+ $q$-mixing (PGCM)]

$$
\left|\Theta^{0}\right\rangle=\sum_{q} f(q) P|\Phi(q)\rangle \longrightarrow H=H_{0}+H_{1}
$$

3. Expand in $H_{1}$


Each step scales polynomially!

## 1. Constrained HFB



- Constrained HFB calculations
- Maps total energy surface (TES)
- Minimum at strongly deformed configuration
- TES soft along the octupole direction


## 1. Constrained HFB


2. Projected HFB

© Projected HFB calculations

- Projections favour deformed configurations
- Negative parity states accessed
- Provide input for computing PGCM state


## ${ }^{20} \mathrm{Ne}$

## 1. Constrained HFB



## 2. Projected HFB


3. PGCM


## ${ }^{20} \mathrm{Ne}$

## 1. Constrained HFB





## 2. Projected HFB


3. PGCM


## ${ }^{20} \mathrm{Ne}$



## 2. Projected HFB


3. PGCM


## Excitation spectrum

Binding energy

$\circ$ (Rotational) excitation spectrum emerges in both (symmetry-breaking and -conserving) approaches

- Symmetry-breaking approach achieves it at a much smaller cost
- Relative energies reproduced at PGCM level
- Dynamical correlations (PT correction) needed for absolute energies


## Revisiting EDF and shell model

## Energy density functionals

$$
H^{\mathrm{eff}}\left|\Psi^{\mathrm{eff}}\right\rangle=E\left|\Psi^{\mathrm{eff}}\right\rangle
$$

Simplified w.f.
Derive ab initio effective $H$
(Beyond) mean field

[Duguet et al. 2022]

## Interacting shell model

$$
H^{\mathrm{eff}}\left|\Psi^{\mathrm{eff}}\right\rangle=E\left|\Psi^{\mathrm{eff}}\right\rangle
$$

Derive ab initio effective $H$
Full (CI) w.f., but in valence space

[Stroberg et al. 2021]

## Ab initio nuclear chart


[Figure: B. Bally]
■ Stable
$\square$ Atomic mass evaluation 2020
$\square$ Ab initio 2010
$\square$ Energy density functional (Gogny D1M)

Data taken from:
S. Hilaire and M. Girod, EPJA 33, 237 (2007)
M. Wang et al., Chin. Phys. C 45, 030003 (2021)
H. Hergert (private communications)

## Ab initio nuclear chart


$\begin{array}{ll}\square & \text { Stable } \\ \square \text { Atomic mass evaluation } 2020 \\ & \text { Ab initio 2020 } \\ \square \text { Energy density functional (Gogny D1M) }\end{array}$

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