

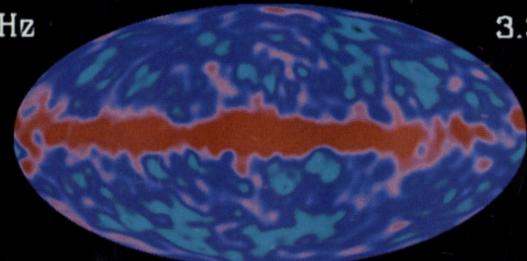
# CMB@TOV Future Perspectives

**Nicola Vittorio, 11 Maggio 2022**

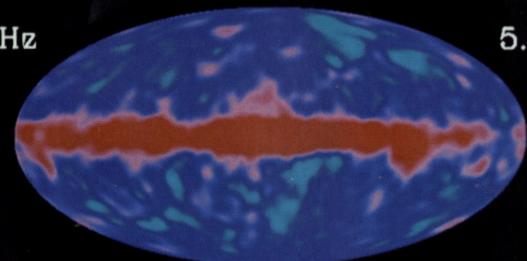
# COBE/DMR, 1992:

COBE DMR - FULL SKY MAPS

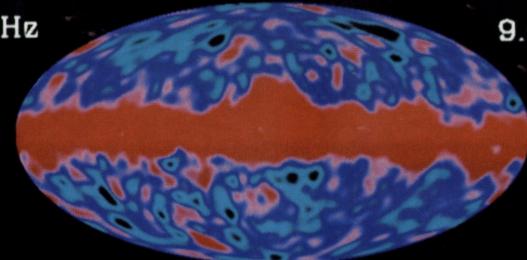
90 GHz    3.3 mm



53 GHz    5.7 mm

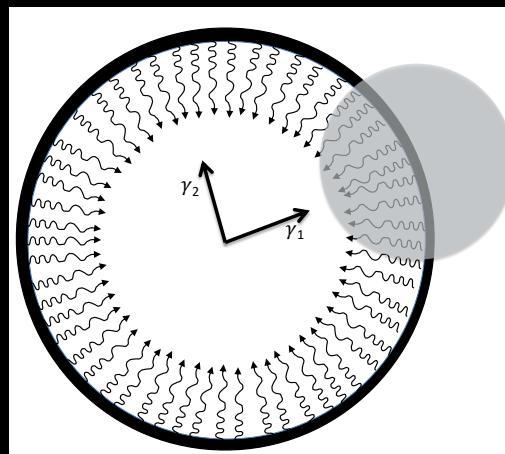
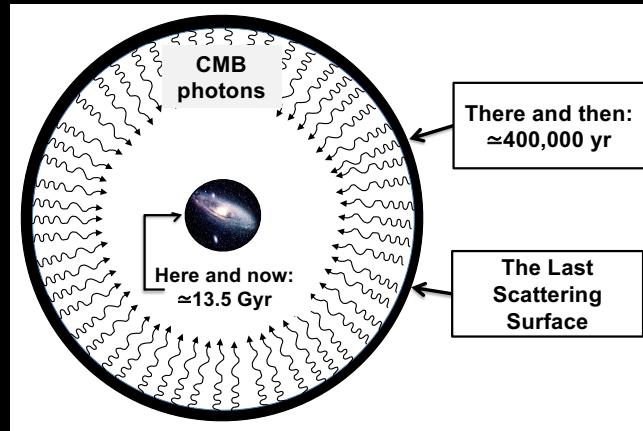


31 GHz    9.5 mm



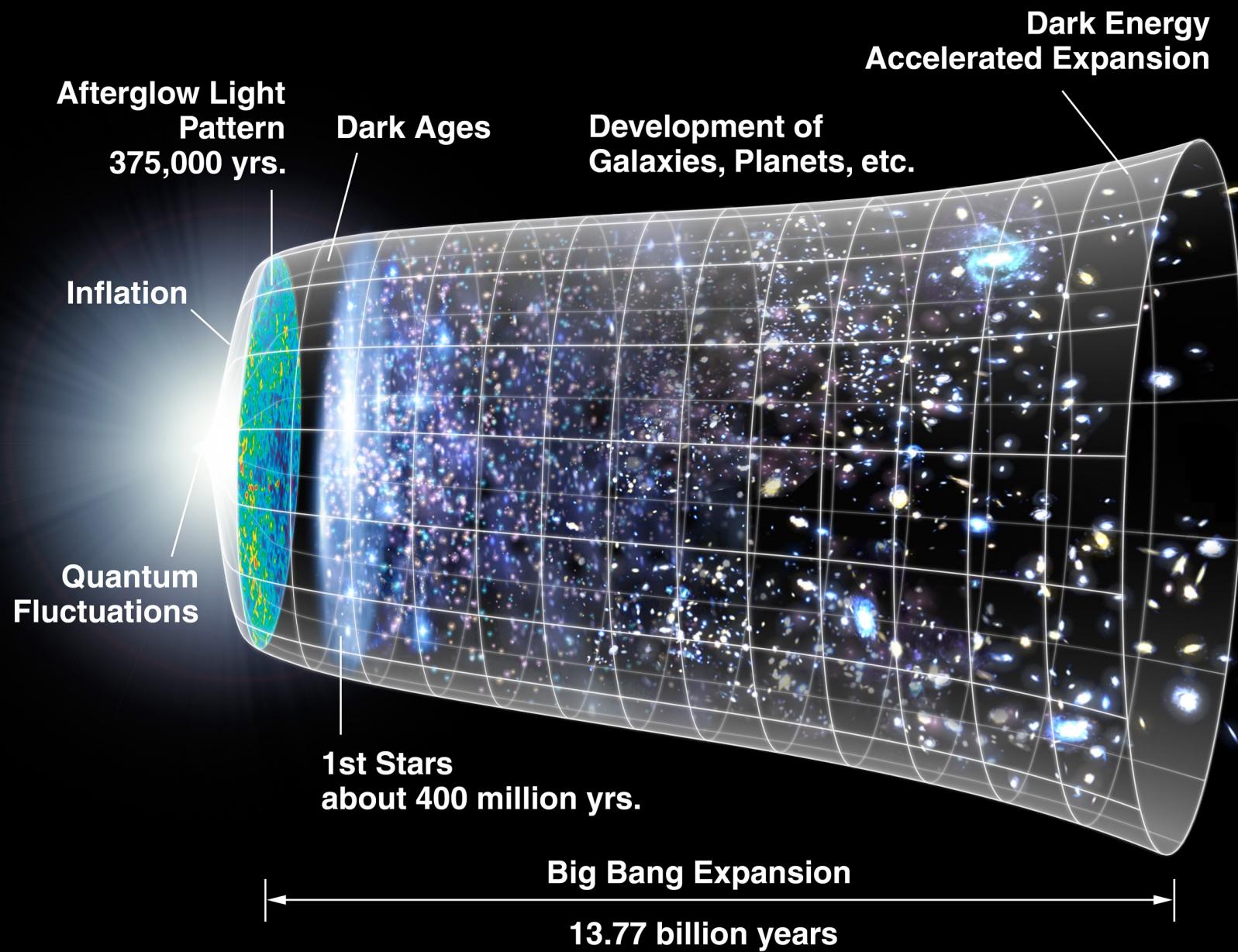
12/89 - 12/90

$$\Delta T|_{rms} = (30 \pm 5)\mu K$$



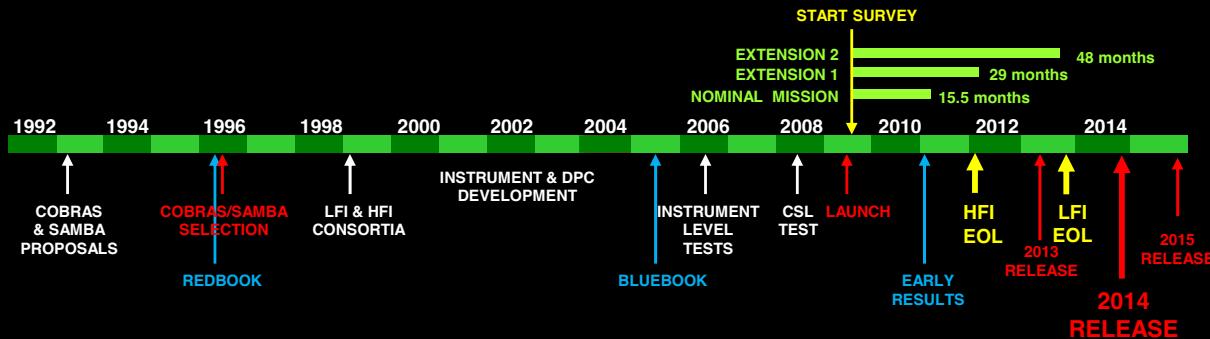
Sachs and Wolfe effect  
(1967)

$$\frac{\delta T}{T} \simeq \frac{\delta \phi}{c^2} \simeq \frac{H_0^2}{c^2} R_0^2 \frac{\delta \rho}{\rho} \Big|_{t_0}$$



# COBRAS/SAMBA

## The Planck Collaboration



## COBRAS/SAMBA

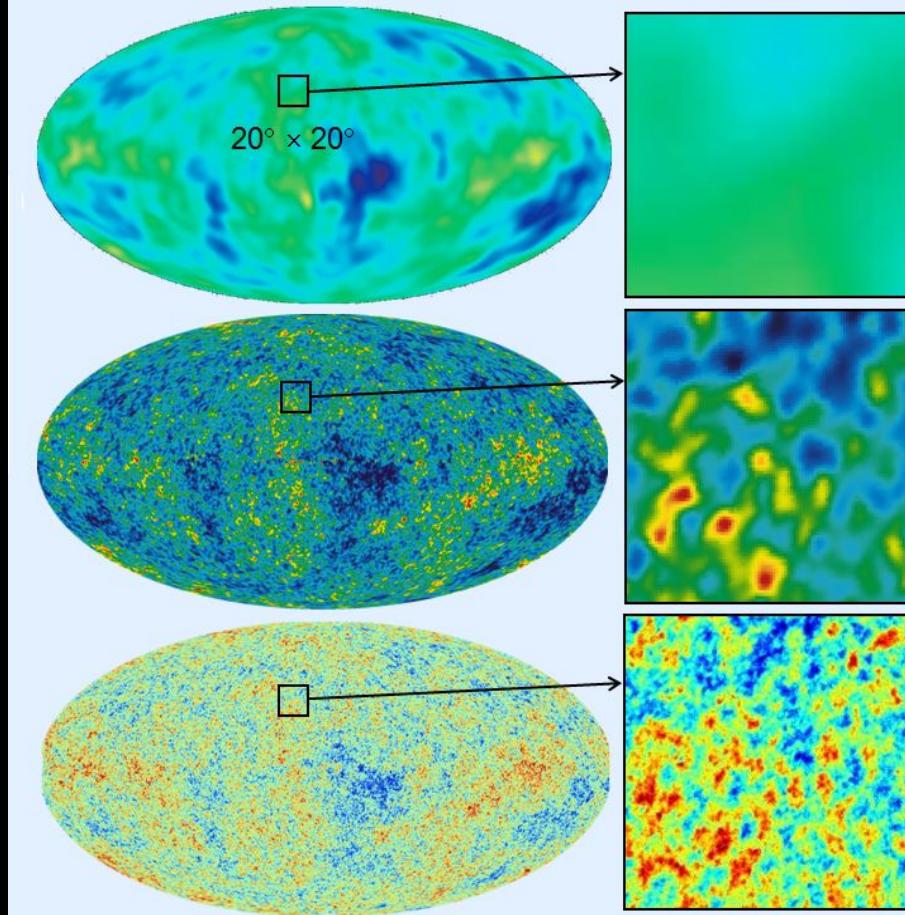
A Mission Dedicated to Imaging the  
Anisotropies of the Cosmic Microwave Background

### REPORT ON THE PHASE A STUDY

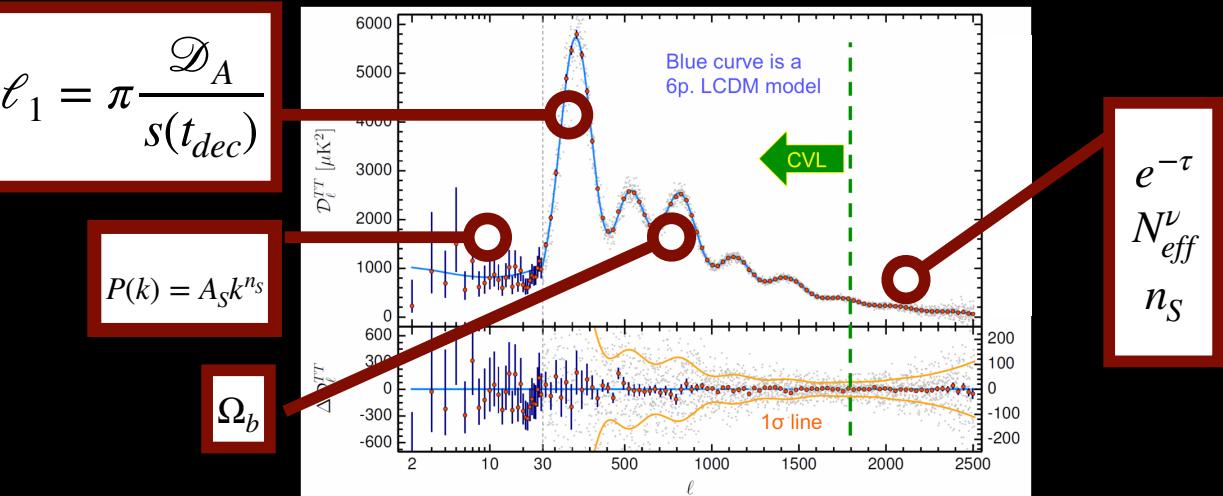
M. Bersanelli, F.R. Bouchet, G. Efstathiou, M. Griffin, J.M. Lamarre,  
N. Mandolati, H.U. Norgaard-Nielsen, O. Pace, J. Polny, J.L. Puget,  
J. Tauber, N. Vittorio, S. Volonté

# 3 CMB Space Missions

- Increasing sensitivity
- Increasing angular resolution
- Increasing control of
  - *Systematic*
  - *Foreground*
- Better parameter estimation



# Planck 2018 TT power spectrum

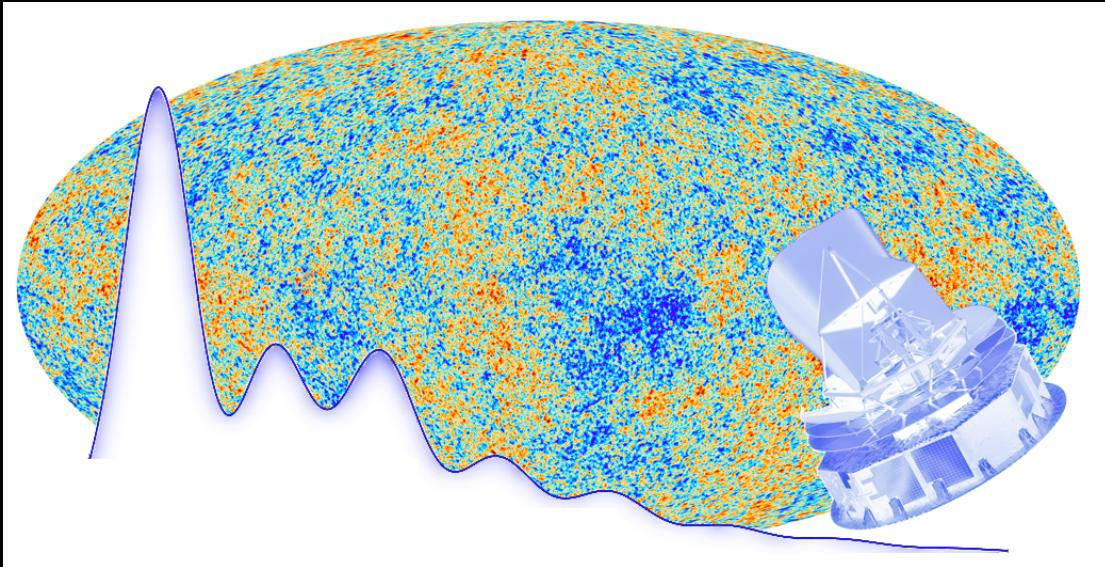


	Mean	Stdev	Rel. err.
Baryon density	0.02237	0.00015	0.7%
DM density	0.1200	0.0012	1%
Acoustic horizon	1.04092	0.00031	0.03%
Optical depth	0.0544	0.0073	13%
Primordial amplitude	3.044	0.014	0.7%
Primordial spectral index	0.9649	0.0042	0.4%

# Precision vs. accuracy?

- **LCDM extension**

- *Running spectral index:*  $dn_S/d \ln k$
- *Primordial non Gaussianity:*  $f_{NL}$
- *Non adiabatic (isocurvature) primordial perturbations*
- *Dark energy equation of state,*  $p = w(z)\rho$
- *Spatial curvature*  $\Omega_k = 1 - \Omega_0 - \Omega_\Lambda$
- *Neutrino total masses:*  $\sum_1^{N_{eff}} m_\nu$
- *Effective number of relativistic species*  $N_{eff}$



# InDark (INFN/CNS IV)

From 2014

- **8 Sezioni INFN partecipanti:**

- *Bologna, Ferrara, Lab. Naz. del Gran Sasso, Padova, Roma II, Roma III, Torino, Trieste*
- *Coordinatore locale @ Tor Vergata: prof. Marina Migliaccio (dal 2020)*

- **2 years INFN Fellowship @ TOV (48 applications for 1 fellowship)**

- *Adrià Gomez-Valent (From September 2021)*
- *Dark Energy and Modified Gravity Models in the light of Low-Redshift Observation*

# Accordo ASI/INFN per lo Space Science Data Center

- **Accordo 2014-2017, esteso al 2021**

- *Cosmic rays, Gamma Astronomy, Cosmology,*
  - *WP CMB : Responsible*
    - *prof. Paolo Natoli (2016-2017)*
    - *prof. Marina Migliaccio (2017-2021)*
- 

- **Accordo 2022-2025**

- *KO Meeting February 2022*
- *Cosmic rays, Gamma Astronomy, Cosmology, Space Weather, Multimessenger Astronomy and GW, IT & AI*
- *WP Cosmology (Responsible Prof. Marina Migliaccio)*
  - *CMB (Planck Legacy, LSPE, LiteBIRD) + cross-correlation with LSS (Euclid).*

# ASI/COSMOS

## KO Meeting - 21/12/2016

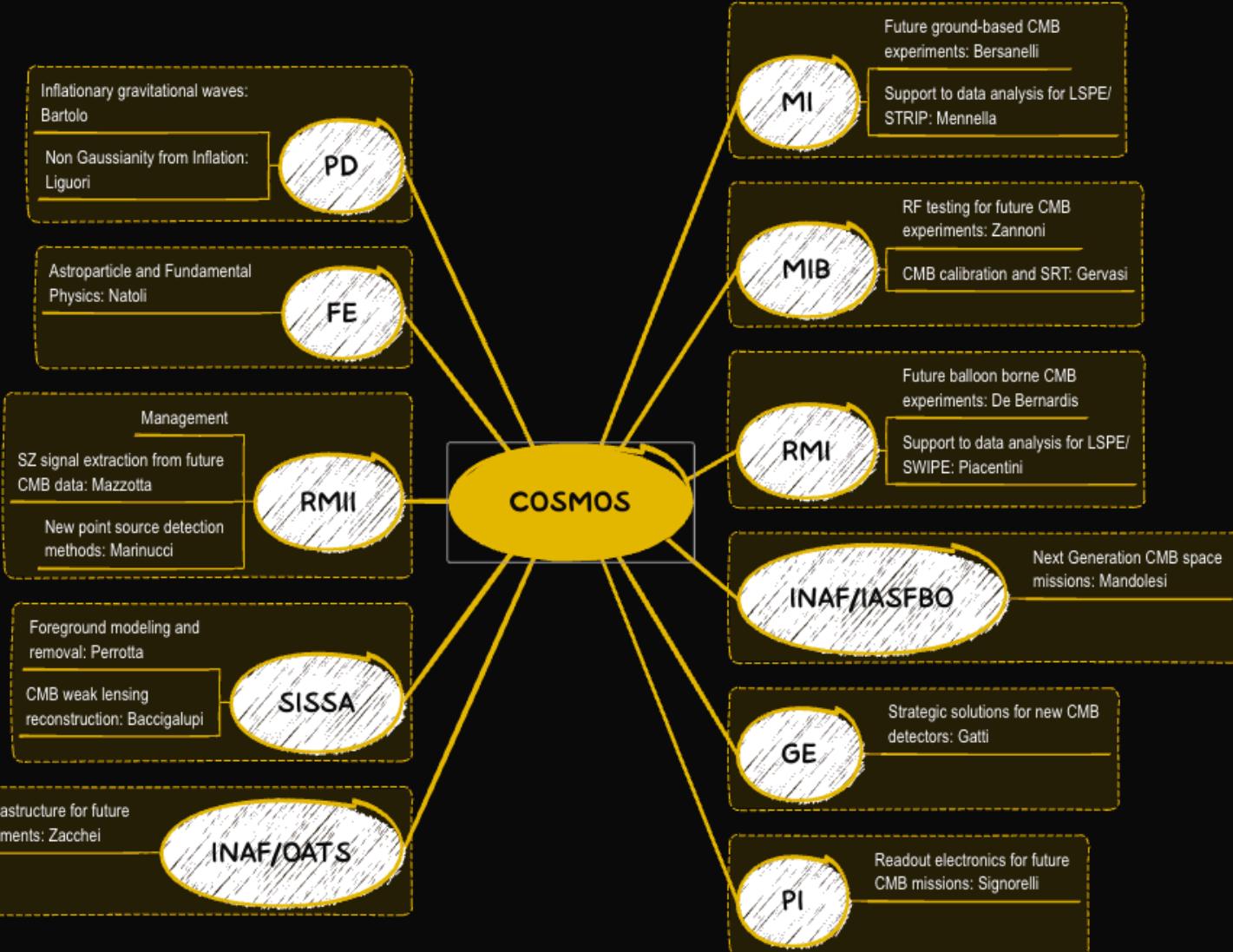
- Agreement with Tor Vergata



The challenges of the Italian Cosmic Microwave Background community: the ASI/COSMOS Project

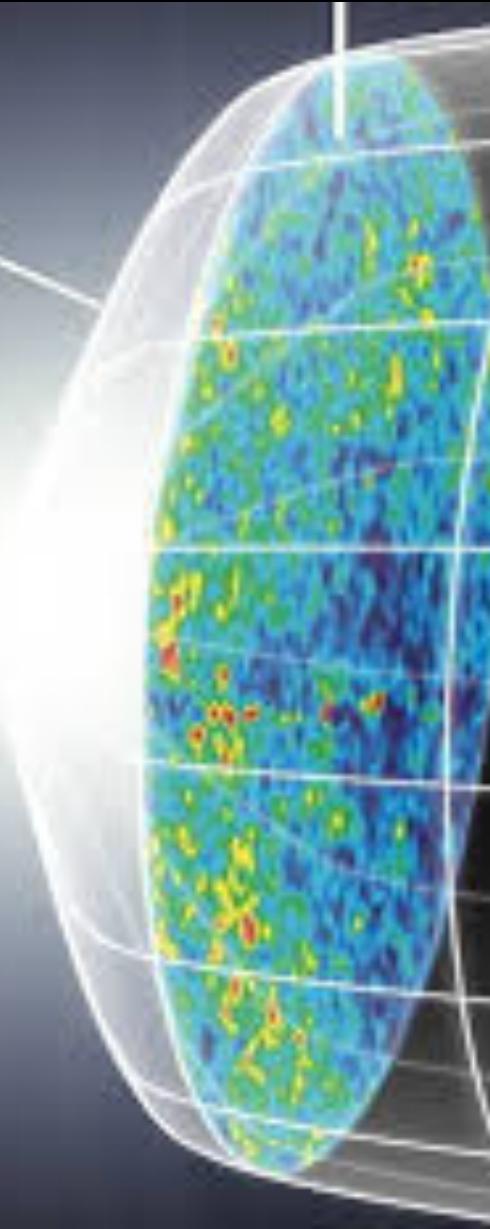
Objectives of the proposal

HOME  
MEETINGS  
INSTITUTIONS / PEOPLE  
ASI/COSMOS WPs



Inflation

Quantum  
Fluctuations



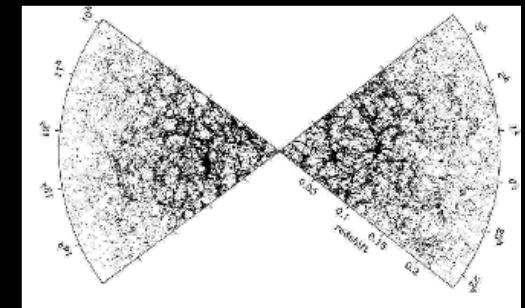
# Gauge-independent cosmological perturbations

- The metric of a perturbed, spatially flat FLRW universe

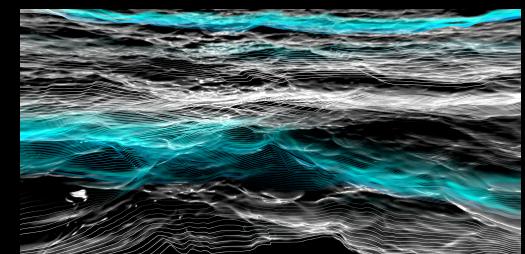
$$ds^2 = a^2(t) \left\{ [1 + 2\Phi(x^k)] c^2 d\mathcal{T}^2 - \left( \delta_{ij} [1 - 2\Phi(x^k)] - h_{ij}^{TT}(x^\mu) \right) dx^i dx^j \right\}$$

- Field equations:

→ Scalar modes:  $\nabla^2 \Phi = \frac{8\pi G}{c^4} \epsilon \Rightarrow$  Large Scale Structure formation and evolution



→ Tensor modes:  $\square h_{ij}^{(TT)} = 0 \Rightarrow$  Stochastic background of primordial GW



# Polarization of the CMB

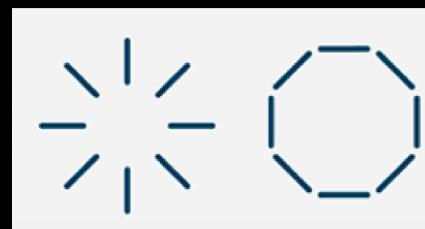
- In terms of the Stokes parameters,

$$I = \frac{31}{16\pi} \sigma_T \left[ \frac{8}{3}\sqrt{\pi}a'_{00} + \frac{4}{3}\sqrt{\frac{\pi}{5}}a'_{20} \right]$$

$$Q + iU = \frac{3}{4\pi} \sigma_T \sqrt{\frac{2\pi}{15}} a'_{22}$$

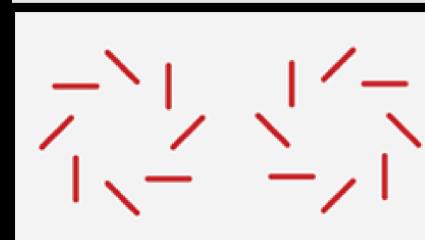
- Q and U are combined in

► E field (even under parity)



↔ Scalar modes

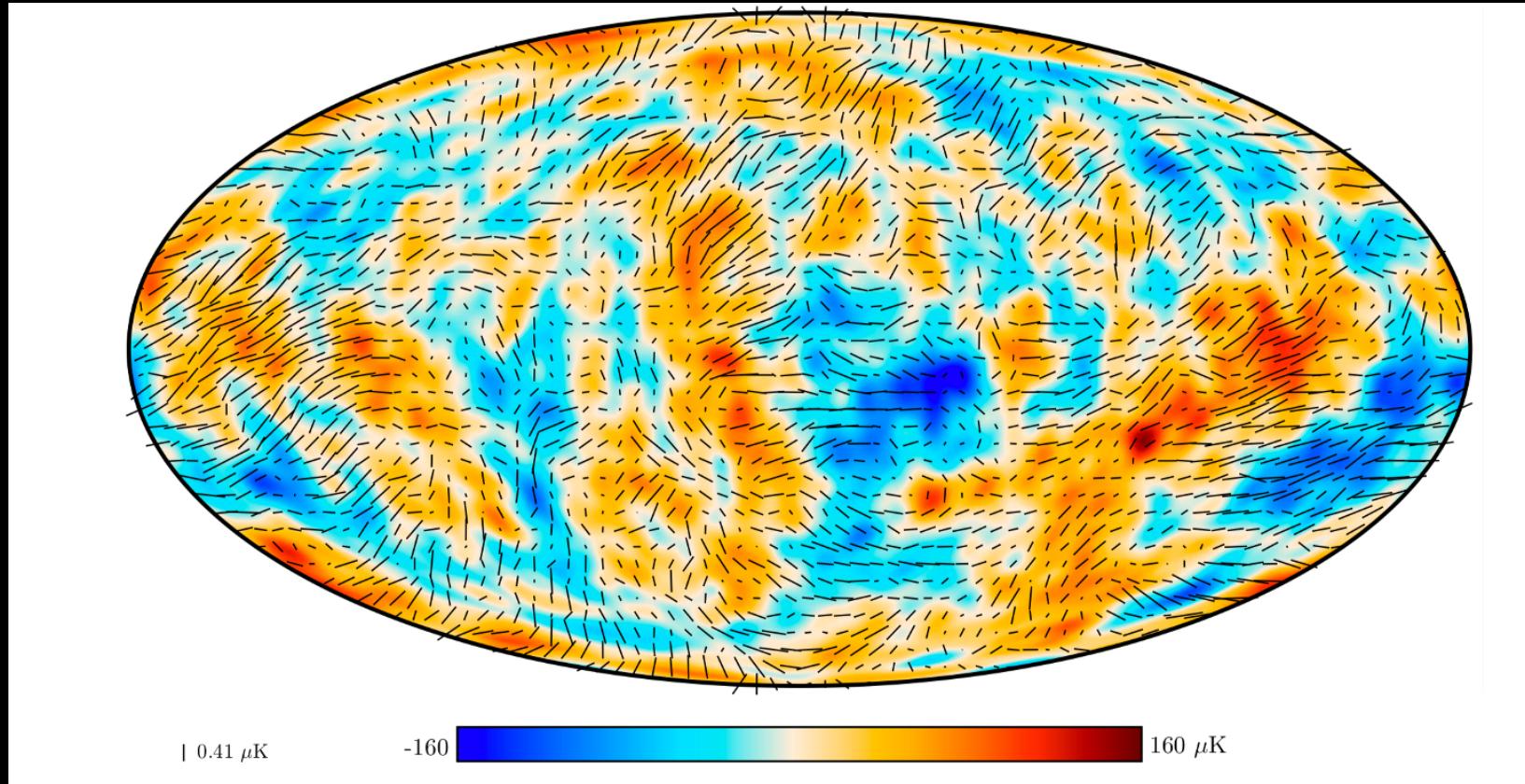
► B field (odd under parity)



↔ Tensor modes ↔ GW from inflation

# PLANCK 2018

## CMB E-modes



# LiteBIRD

Lite(Light) B-mode Inflation Radiation Detection



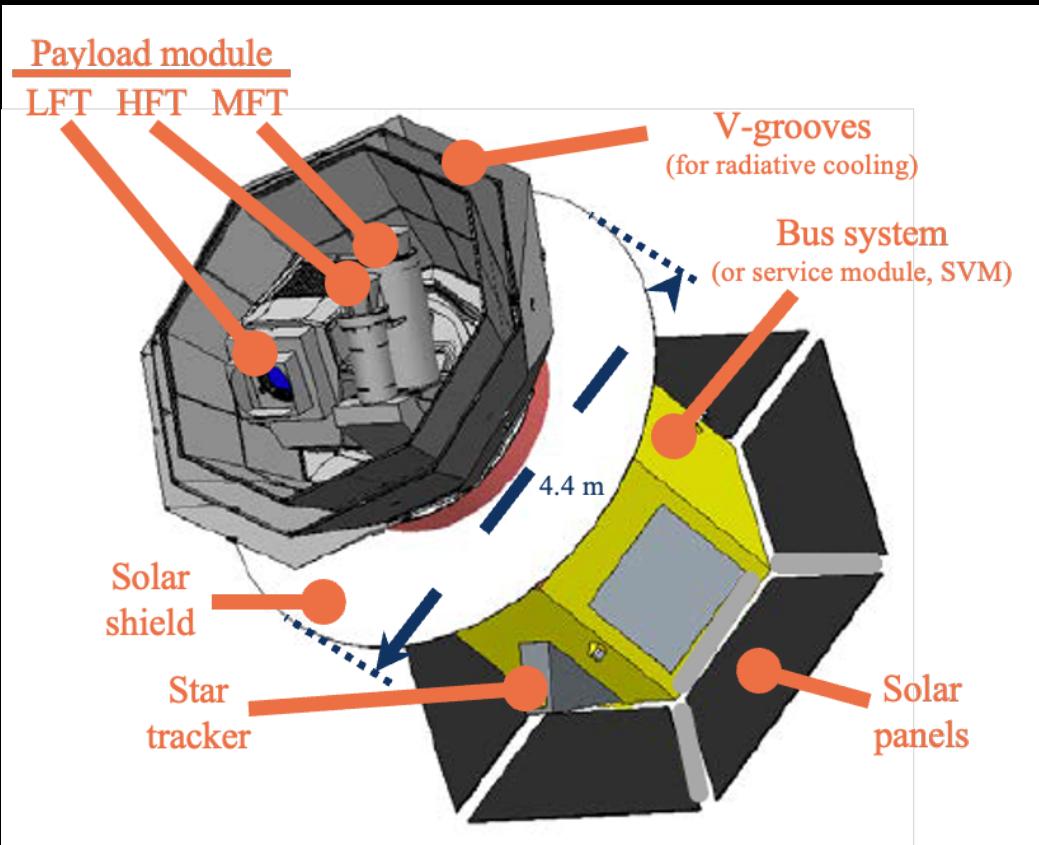
- Post-Planck space mission for CMB polarization
  - Selected by JAXA as strategic L-class mission (05/2019)
  - With participation from US and Europe
- Expected launch
  - in late 2029
  - JAXA's H3 rocket
- All-sky 3-year survey
  - from Sun-Earth Lagrangian point L2





# LiteBIRD spacecraft overview

- **3 telescopes**
  - *LFT (low frequency telescope)*
  - *MFT (middle frequency telescope)*
  - *HFT (high frequency telescope)*
- **MFT and HFT are European Instruments**
- **Large frequency coverage**
  - *40–402 GHz, 15 bands*
  - *70–18 arcmin angular resolution*
  - *Final combined sensitivity:  $2.2 \mu\text{K}\cdot\text{arcmin}$*





# LiteBIRD

Lite(Light) B-mode Inflation Radiation Detection

- **Science outcome**

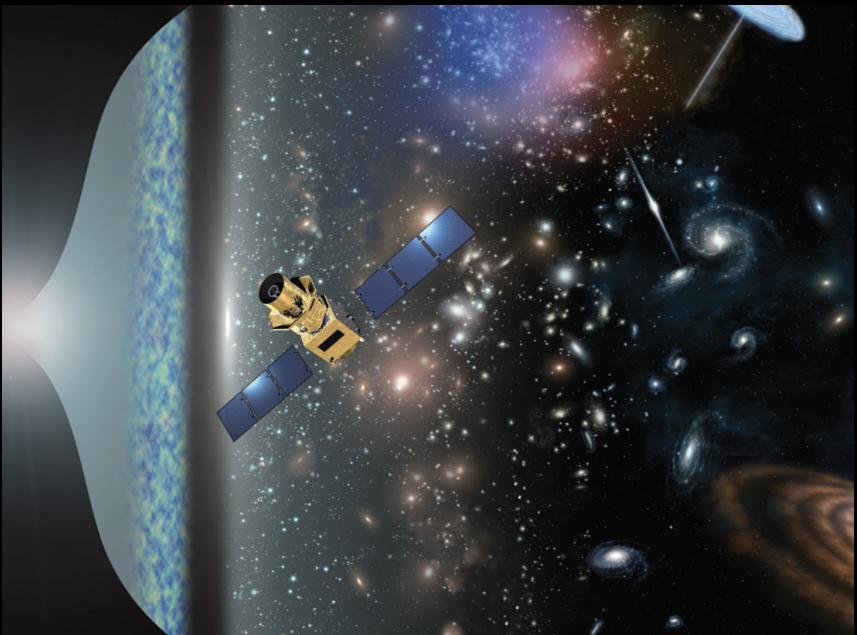
- *Detection of CMB B-mode from primordial gravitational waves*

$$r = \frac{A_T}{A_S} \text{ @ } k = 0.05 \text{ Mpc}^{-1}$$

- *Direct evidence for inflation*
- *Knowledge on the inflation energy scale*

$$V^{1/4} \simeq 10^{16} \left( \frac{r}{0.01} \right)^{1/4} \text{ GeV}$$

- *First evidence for quantum fluctuation of space-time*

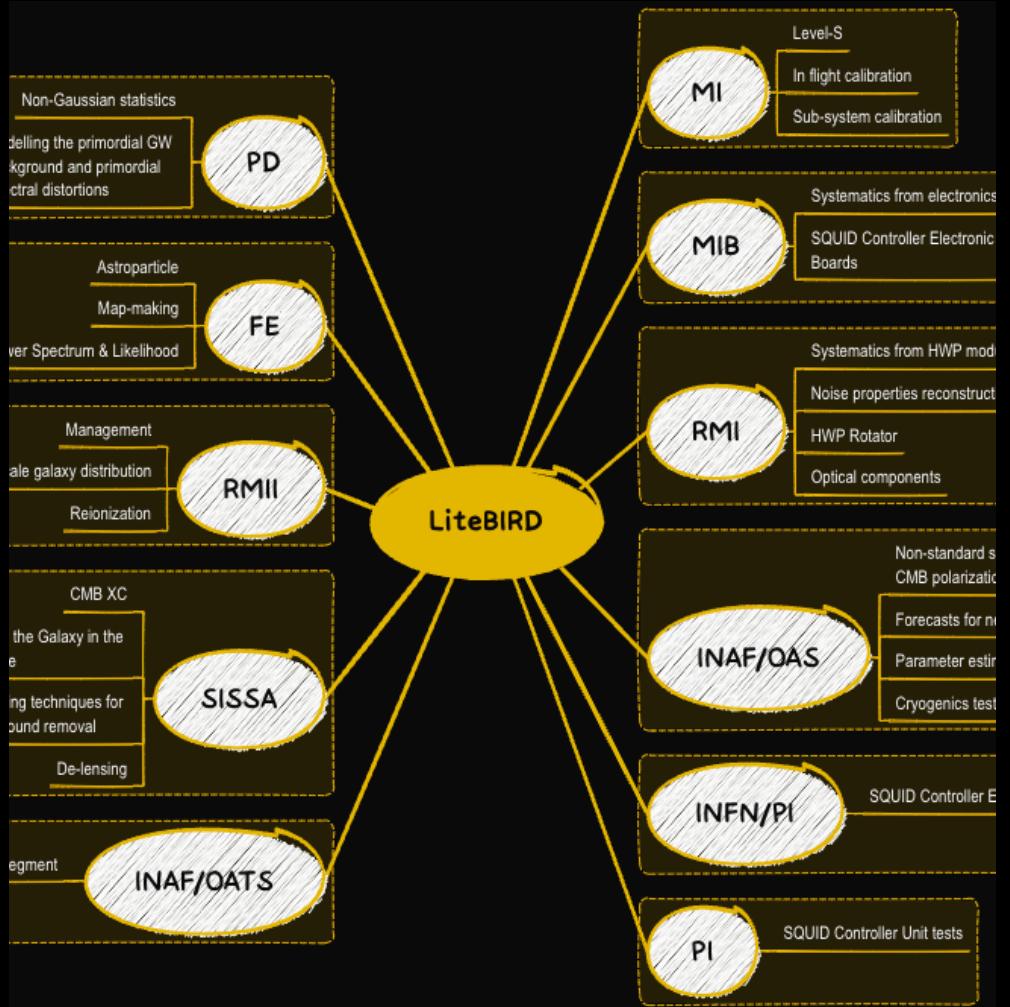
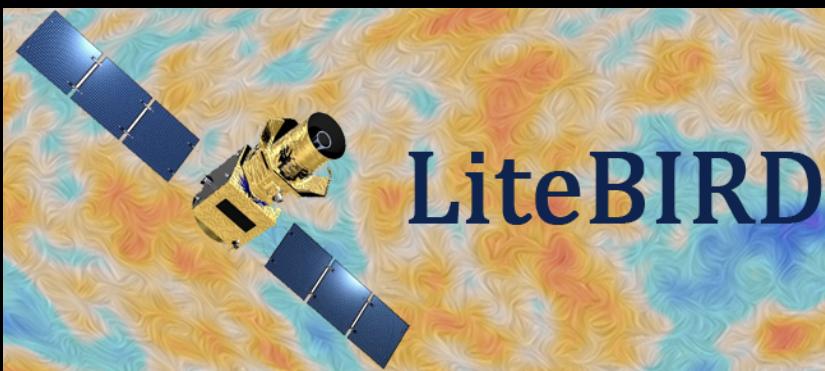




# ASI/LiteBIRD

## KO Meeting - 12/05/2020

- Agreement with Tor Vergata
- Support to the Italian community for participating to the Phase A study

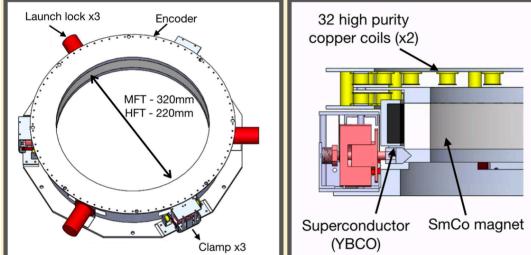




# Il contributo italiano sull'HW

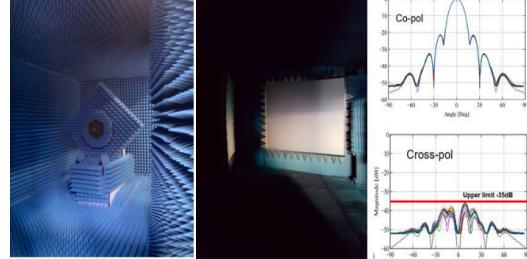


## Rotatore di polarizzazione MHFT (Sapienza)



- HWP rotante a levitazione
- Componenti dell'ottica di MHFT

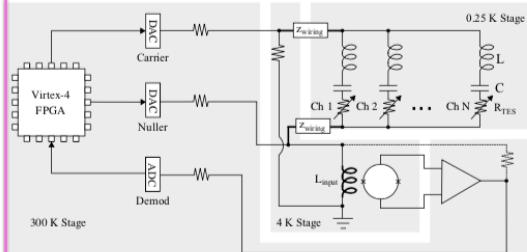
## MHFT System-level calibrations (Milano Statale)



- Esperienza di Milano nella calibrazione di Planck



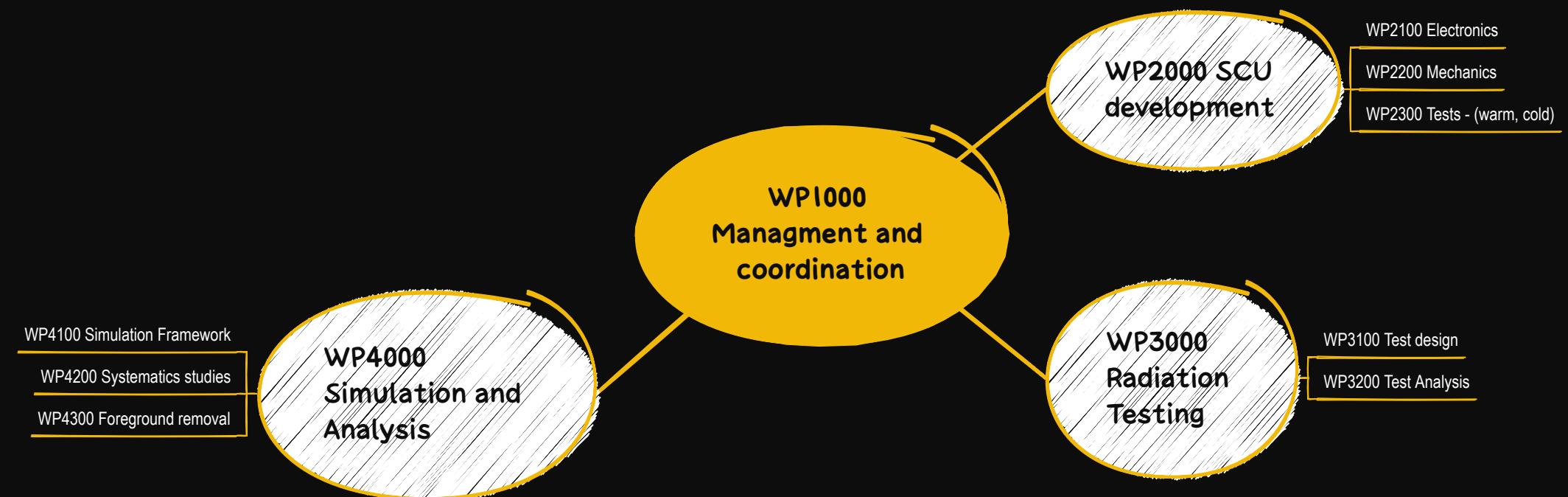
## SQUID readout (INFN)



- Read-out dei bolometri criogenici

# LiteBIRD (CSN II)

- FE: P. Natoli; MI: M. Tomasi; MIB: M. Zannoni; INFN/PI: G. Signorelli (PI); RM1: L. Lamagna; RM2: M. Migliaccio; TS: N. Krachmalnicoff



# Conclusions

- **Hunting for B-modes of the CMB polarization**
  - *the forefront of the experimental research*
- **Why to go from space?**
  - *Rule of thumb: 1,000 detectors in space  $\Leftrightarrow$  100,000 detectors on ground*
  - *No statistical/systematic uncertainty due to atmosphere, no ground pickup*
  - *The only way to access to the lowest multipoles*
  - *A wide ancillary science: neutrinos, Galactic and extragalactic astrophysics*

# Conclusions

- LiteBIRD is
  - ▶ *approved*
  - ▶ *the only space-born CMB mission in this decade*
- The Italian community is heavily involved
- The Sezione-INFN and the Physics Dept. are heavily involved

