



Crystalline mirrors for GW detectors

M. Leonardi on behalf of the KAGRA collaboration

Summary

- 1. Introduction to KAGRA
 - Performances during O3GK
 - Current status of upgrades
- 2. KAGRA's experience with crystalline mirrors (substrates)
 - Problem encountered
 - Substrate characterization
 - Lessons for 3G detectors

Part 1: introduction to KAGRA

KAGRA is located near Kamioka in Gifu prefecture, one hour drive from Toyama.

Unique features:

- 1. Underground site
- 2. Cryogenic operation



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Last news from KAGRA

- May 2020: O3GK finished (~1Mpc BNS range)
- July Oct. 2020: RSE trial
- Oct. 2020 -
 - Opened Vacuum Chambers
 - Upgrade of vibration isolation systems
 - IMC commissioning
 - IFO re-alignment
 - X-arm single arm commissioning
 - ALS commissioning
 - Pcal upgrade
 - PEM upgrade

O3GK noise budget





KAGRA Suspension Types Туре-А Туре-В Туре-Вр Type-C

0/0/2022

Fixed a number of issues with the room temperature part

- Some of the filters were stuck in O3GK
 - Now fixed
- ITM top GAS filter blades were replaced
- New accelerometers for the top stages
- LVDT driver coupling reduced
- Many other messy problems (cables touching, vacuum leaks, ...)



Contamination of the ETMY cryostat and mirror







Before

After





Surface Absorption Measurement of the Cleaned Mirror





Y [mm]

0.5

6/6/2022

GRASS 2022

O3GK noise budget





RSE lock trial

- Suspensions were not quiet enough
- Not enough time to tune feedback filters
- No WFS implemented

RSE lock ~ 2 sec



Alignment Sensing and Control

- ASC is much more difficult and messier than LSC (Universal Truth)
- It is especially hard for KAGRA because of the birefringence





Part 2: KAGRA experience with crystalline mirrors

Recap on sapphire

Sapphire substrate:

• Absorption:

well known problem from design phases

• Birefringence:

partially unexpected problem



Growth orientation impact on opt. absorption spatial distribution

C-axis growth:





	Company B		S1		S2	S3
	#0		167.6 (35.5)		126.8 (41.7)	26.6 (17.0)
	#1		96.1 (31.2)		93.1 (28.7)	99.3 (38.0)
	#2		27.1 (11.0)		31.5 (10.5)	24.8 (10.5)
	#3		73.3 (23.2)		137.6 (34.7)	143.1 (39.9)
	#4		130.9 (39.1)		99.4 (22.7)	139.5 (35.7)
	#5		85.3 (34.7)		56.5 (14.9)	119.6 (39.9)
	#6		184.9 (54.5)		140.9 (36.1)	201.6 (54.8)
-	#7		65.1 (14.9)		62.2 (16.6)	91.2 (28.7)
	#8		111.4 (34.1)		66.3 (18.6)	143.2 (30.3)
	#10		112.5 (51.7)		38.7 (17.6)	32.9 (13.7)
	#11		21.2 (12.3)		22.8 (18.2)	33.4 (23.4)
	#12		77.1 (24.2)		83.6 (30.2)	89.7 (31.1)
Company A		S 1	S1		2	S3
#1(F62-22)		41.1 (23.1)		64.3 (23.3)		59.9 (19.8)
‡2 (F47-21)		72.1 (31.1)		8	37.3 (38.2)	93.4 (36.5)
#3 (AC-179)		60.6 (76.0)		З	37.7 (17.3)	47.7 (34.1)
‡4 (F39-56)		94.0 (139.6)		З	805.3 (250.0)	160.2 (171.4)
‡5 (MMK-1)		106.98 (45.68)		e	9.02 (22.91)	76.0 (31.0)
‡6 (MMK-2)		216.47 (108.01)		8	32.82 (30.90)	99.4 (49.5)
‡7 (C14-11c)		114.2 (95.5)		7	9.3 (49.9)	72.4 (42.7)
#8 (OC-1)		86.6 (45.7)		e	9.5 (35.5)	58.0 (25.2)

Dislocation and optical absorption

Structural defects as preferential sites for the inclusion of absorbing centers

www.nature.com/scientificreports

Check for updates

scientific reports

OPEN 3D characterization of low optical absorption structures in large crystalline sapphire substrates for gravitational wave detectors

Manuel Marchiò^{1,2⊠}, Matteo Leonardi², Marco Bazzan³ & Raffaele Flaminio^{2,4}

https://doi.org/10.1038/s41598-020-80313-1



Discovery of birefringence in KAGRA



Constructing birefringence from TWE

$$\theta = -\frac{1}{2} \tan^{-1} \frac{TWE(45) - TWE(135)}{TWE(0) - TWE(90)}$$
$$\alpha_{-} = \frac{2\pi}{\lambda} \cdot \frac{TWE(0) - TWE(90)}{\cos 2\theta}$$

 θ (angle between input polarization direction and e-axis) looks very random due to the inhomogeneity of sapphire substrate.

The PV value of α_{-} (one-way differential phase) is around ±150nm.

There are many sharp features, which is bad.



6/6/2022



We calibrated rotation errors according to the markers in the TWE measurements.





Rotation angles and markers in HR map masurements are different from TWE (HRthAR) measurements.

500

400

300

200

ITMX single bounce

After careful calibrations for the TWE maps

P-pol shape





Calibration from rotation error of the 0° TWE map

Simulation



TEM00: <10% 1st mode: ~20% 2nd mode: ~10% Higher order: >60%

One last thing, the 0° TWE map may be not measured at 0°. If there are rotation errors (HR maps indicate the error is around 3°):



Preliminary ASC simulations (single arm)



Characterization setup at NAOJ



From birefringence to Δn

Δn_{RMS} is within original specs





From birefringence to Δn

Δn_{RMS} is within original specs



Correlation between absorption and internal stress leading to birefringence



From birefringence to Δn

Δn_{RMS} is within original specs

Correlation between absorption and internal stress leading to birefringence



Structural defects can be the cause of both absorption and birefringence*

*Paper submitted to Scientific Reports.

Closing remarks

Sapphire substrates are challenging:

- 1. Absorption and absorption distribution depends on manufacturing technique
- 2. Birefringence homogeneity is an issue for IFO control
- 3. Interaction with crystal makers (companies) is necessary but slow due to the lack of interest from commercial environment

Main take-away messages:

- 1. Characterization of full-size substrates is necessary since many issues seems to be size dependent
- 2. Whichever crystalline material will be selected for 3G detectors, dedicated facilities for substrate productions need to be established to approach those issues in a "scientific way".

Thank you!

ITMY