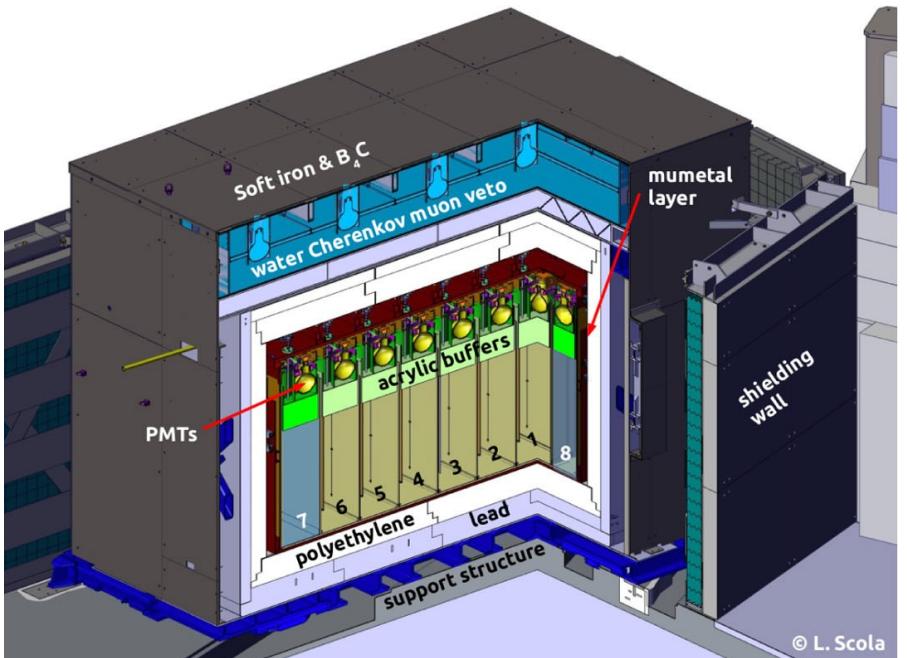




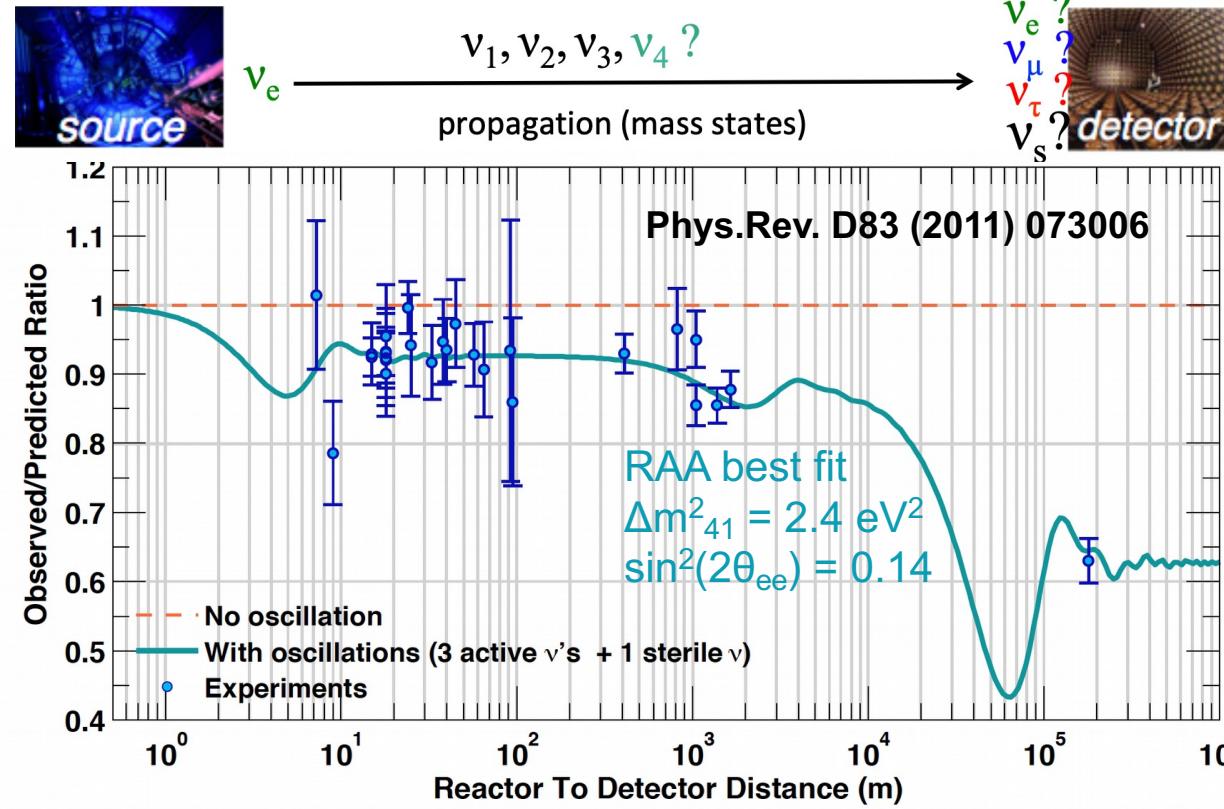
# final results

September 7<sup>th</sup> 2022

Pablo DEL AMO SANCHEZ  
on behalf of the STEREO collaboration

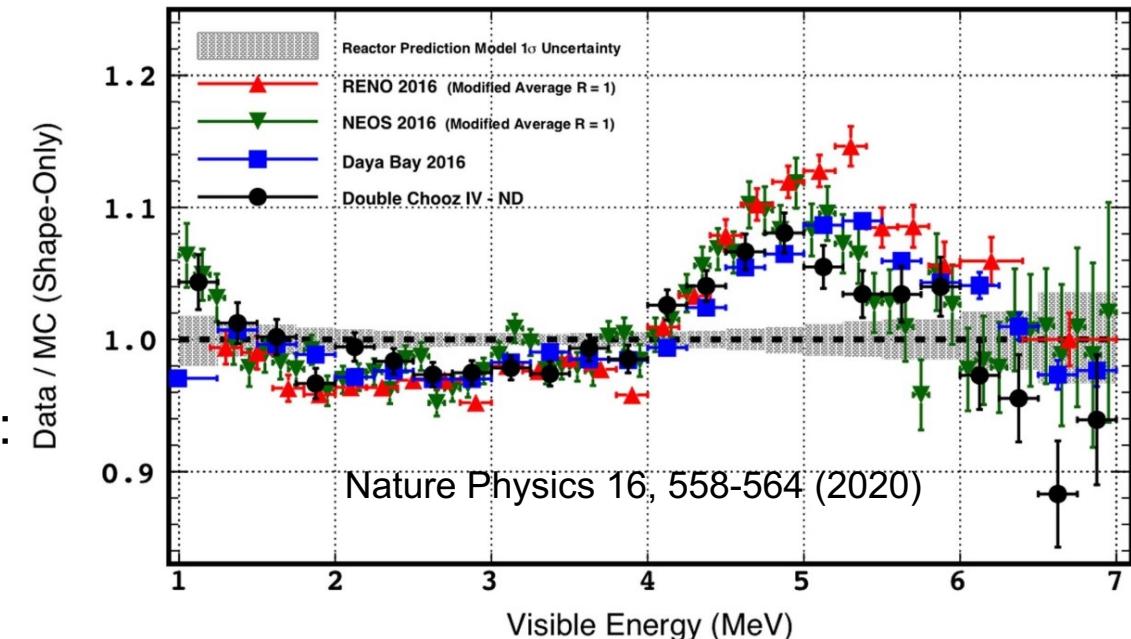


# Reactor Antineutrino Anomaly(ies)



Daya Bay, RENO, Double Chooz (2014):  
 unexpected excess around 5 MeV wrt  
 spectrum predictions (“5 MeV bump”)

Reactor Antineutrino Anomaly (2011):  
 ~6% deficit observed in measured  
 reactor antineutrino fluxes when  
 compared with latest predictions.  
 Sterile neutrino with  $\sin^2(2\theta_{ee}) \sim 0.17$ ,  
 $\Delta m_{41}^2 \sim 2.3 \text{ eV}^2$  would explain RAA and  
 Gallium anomalies



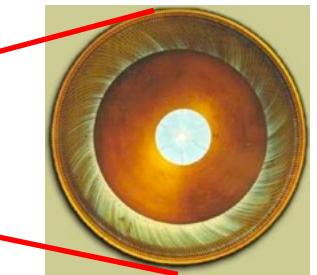
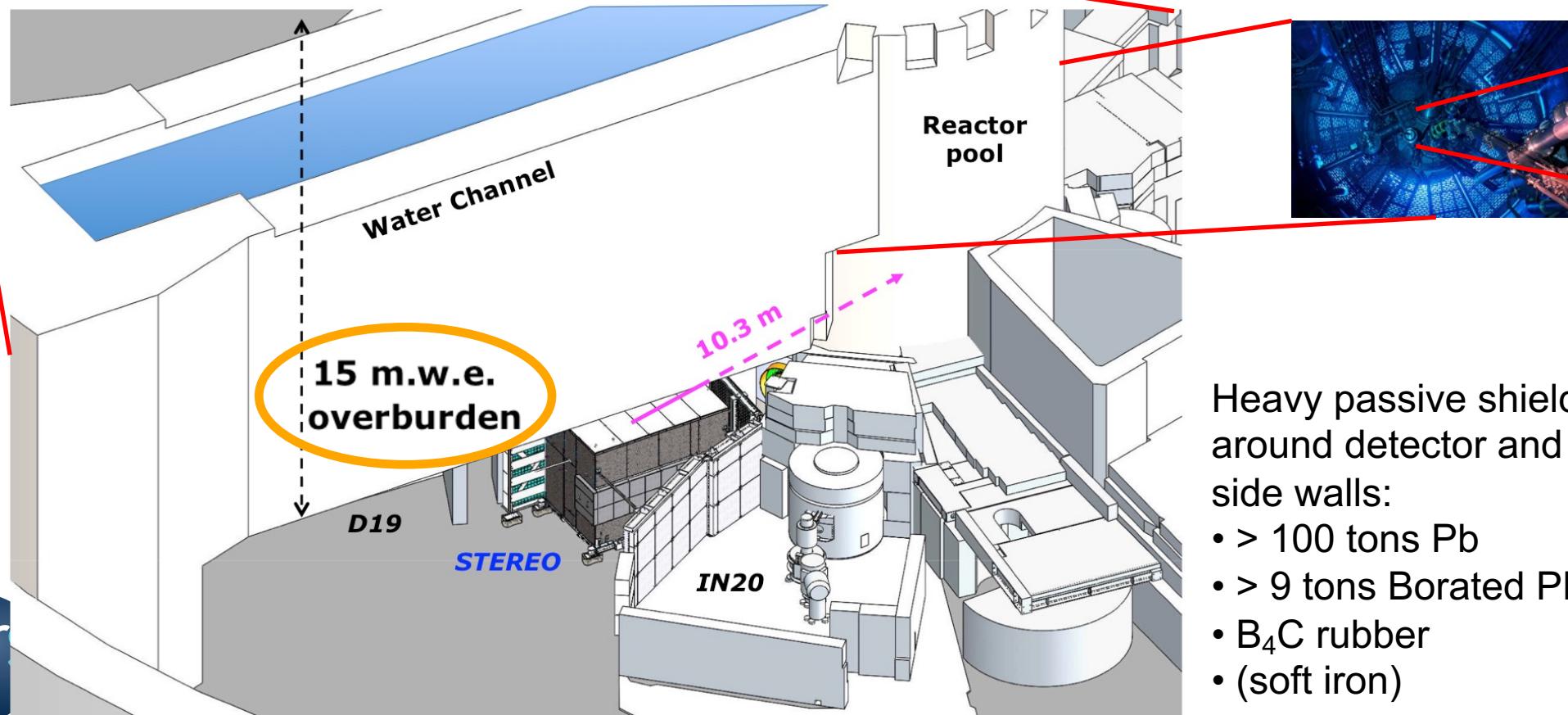


# The ILL site

Challenging backgrounds induced by:

- Cosmic rays
- Neighbouring experiments

- 58.3 MW<sub>thermal</sub>
- Ø40 cm × 80 cm
- Highly enriched:  
93% <sup>235</sup>U  
(fissions > 99% <sup>235</sup>U)
- 3-4 cycles/yr x  
50 days/cycle
- 10<sup>19</sup> s<sup>-1</sup> pure  $\bar{\nu}_e$  flux



Compact core

Heavy passive shielding added around detector and on front and side walls:

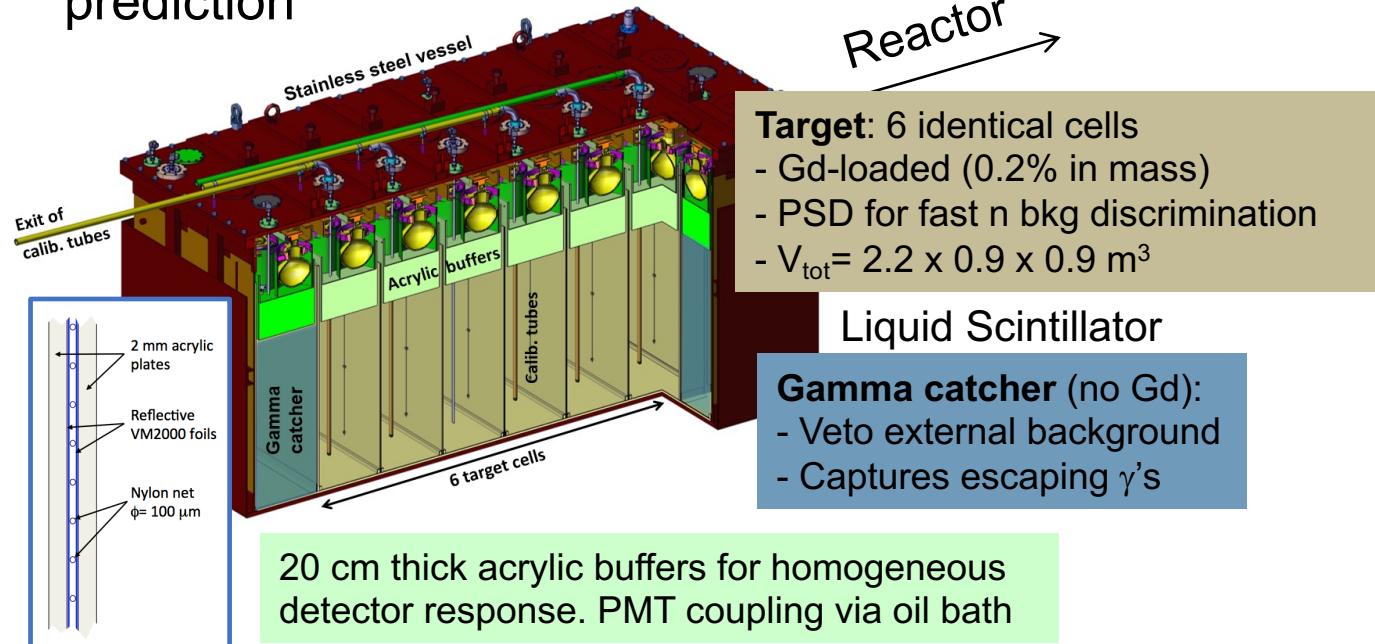
- > 100 tons Pb
- > 9 tons Borated PE
- B<sub>4</sub>C rubber
- (soft iron)

# The STEREO detector

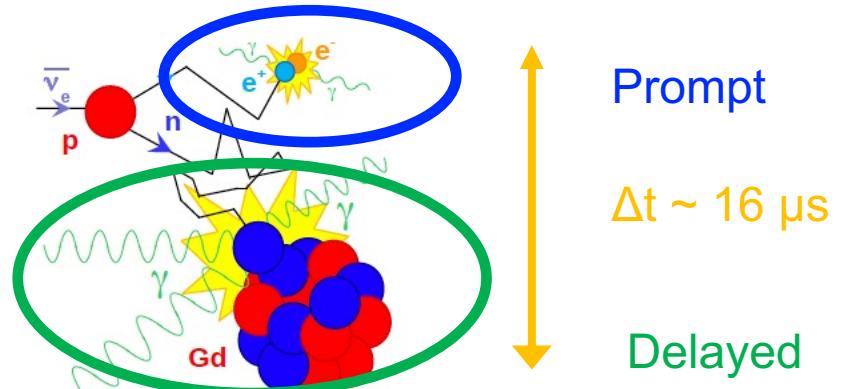
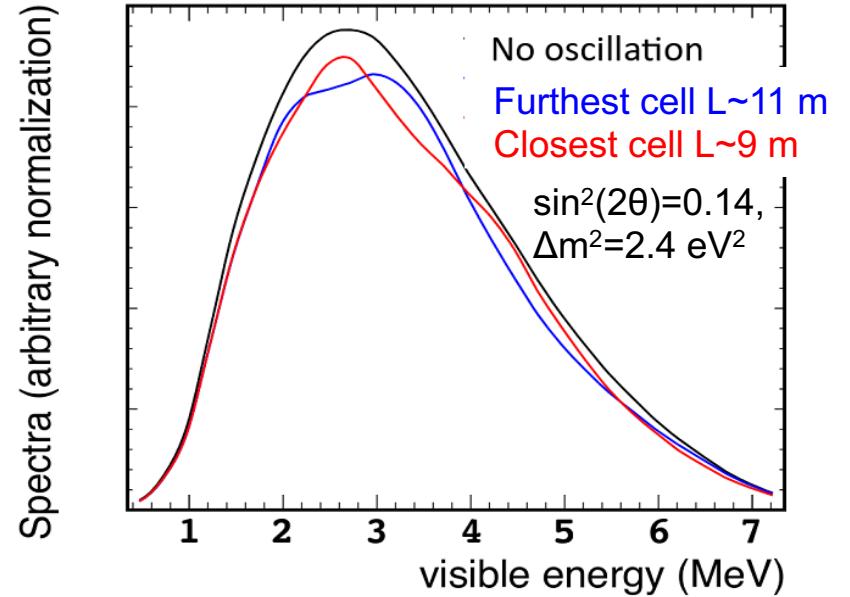
JINST 13, 07 (2018): P07009

Compare of 6 target cells looking for oscillation-like distortions in  $E_\nu$  spectra

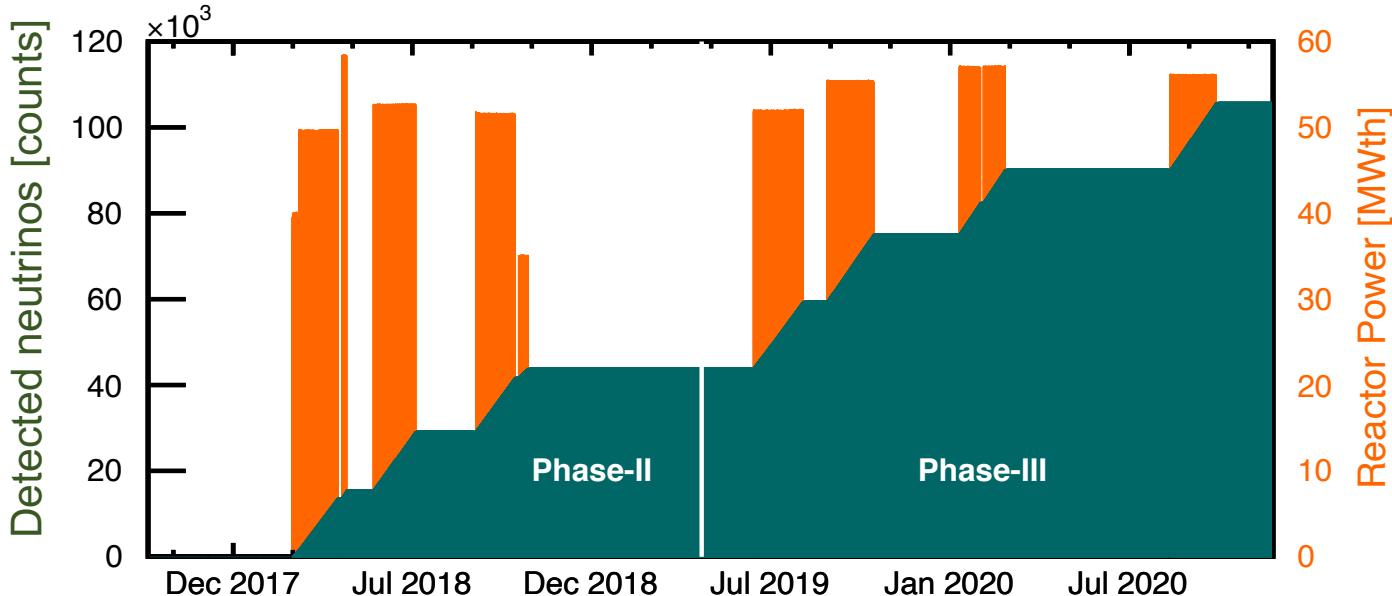
⇒ Reduce dependence on spectrum prediction



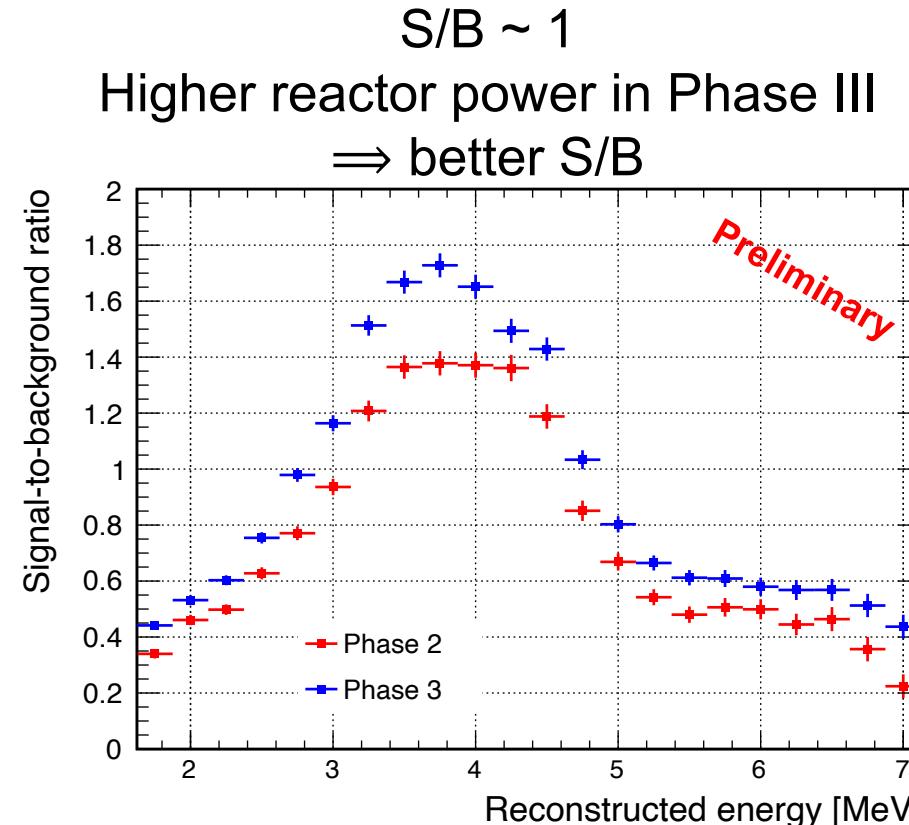
$\nu$  detection through Inverse Beta Decay  
 (prompt + delayed coincidence)



# Dataset

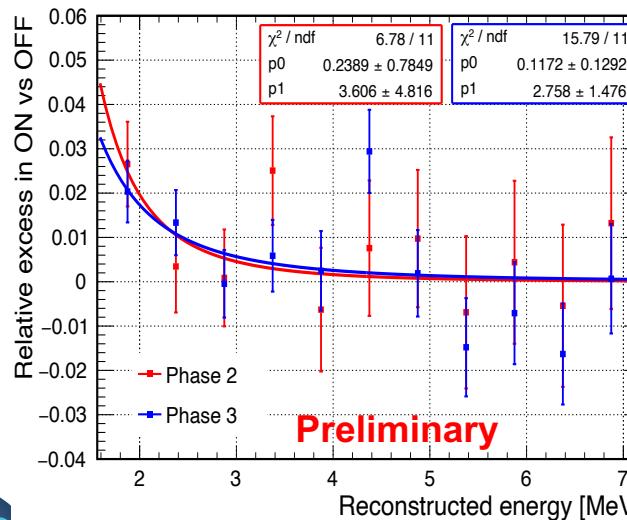


- Data-taking ended Nov 2020
- For better data quality, do not use Phase I
- Phase II + Phase III: 107k nu  
+ calibration runs (hourly LEDs, weekly  $^{54}\text{Mn}$ , monthly AmBe, bi-annual  $^{68}\text{Ge}$ ,  $^{137}\text{Cs}$ ,  $^{60}\text{Co}$ ,  $^{24}\text{Na}$ )
- Reactor OFF data for background subtraction

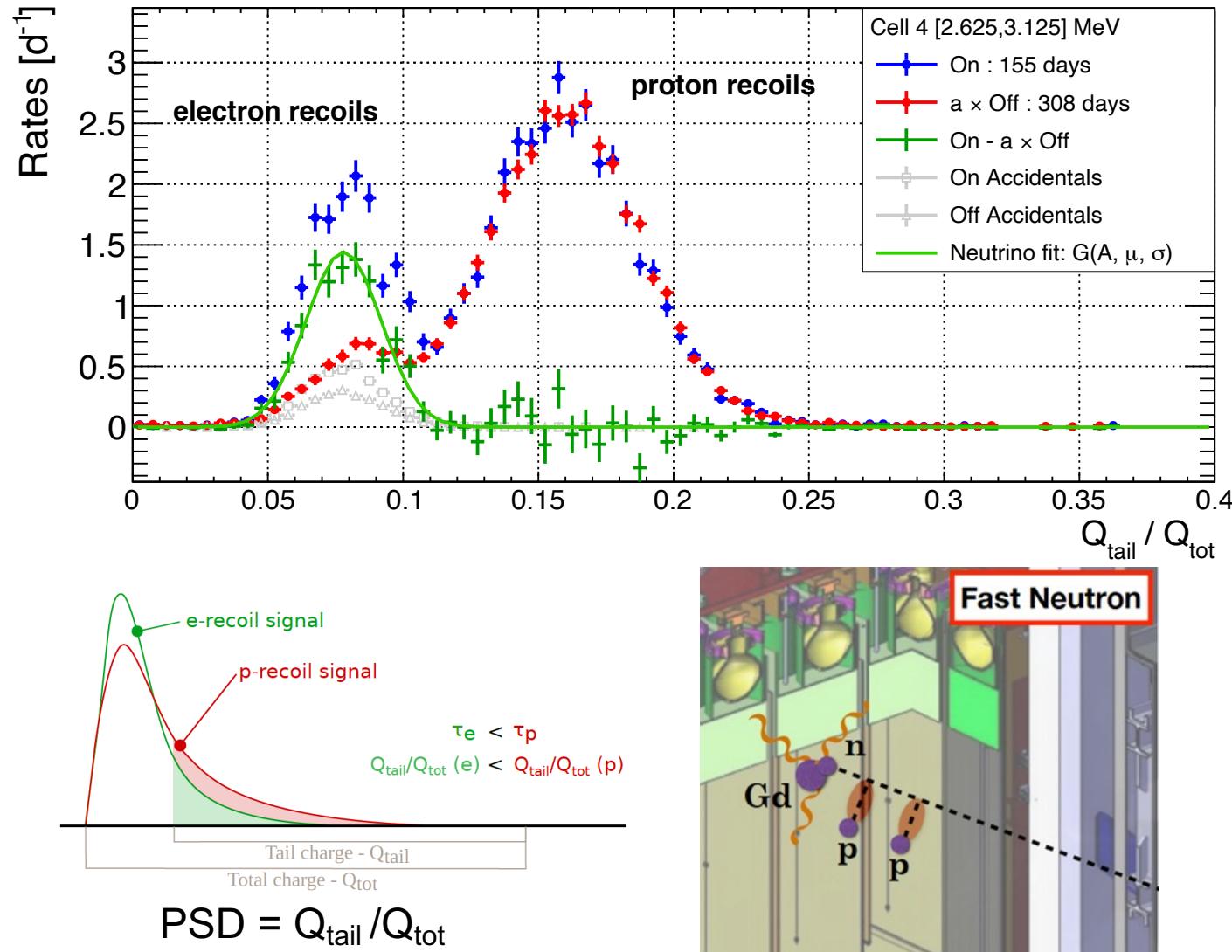


# Neutrino extraction

- Fit Pulse Shape Discrimination (PSD) spectra to extract neutrino signal from correlated backgrounds (e.g. fast neutron from spallation by cosmics) in each cell and energy bin
- Background shape from OFF data
- Accidental from displaced-time window

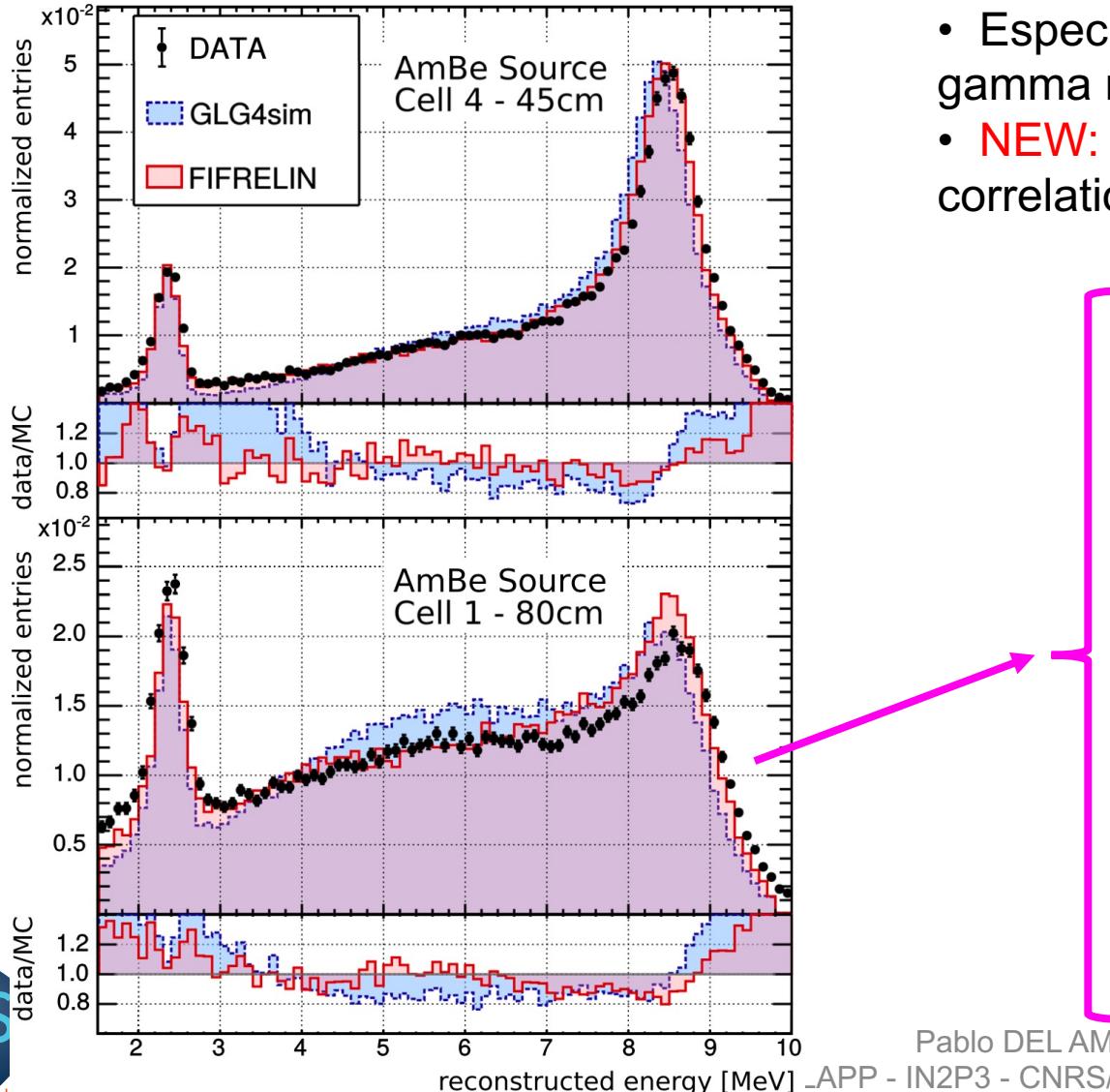


Look for reactor prompt background in proton recoils

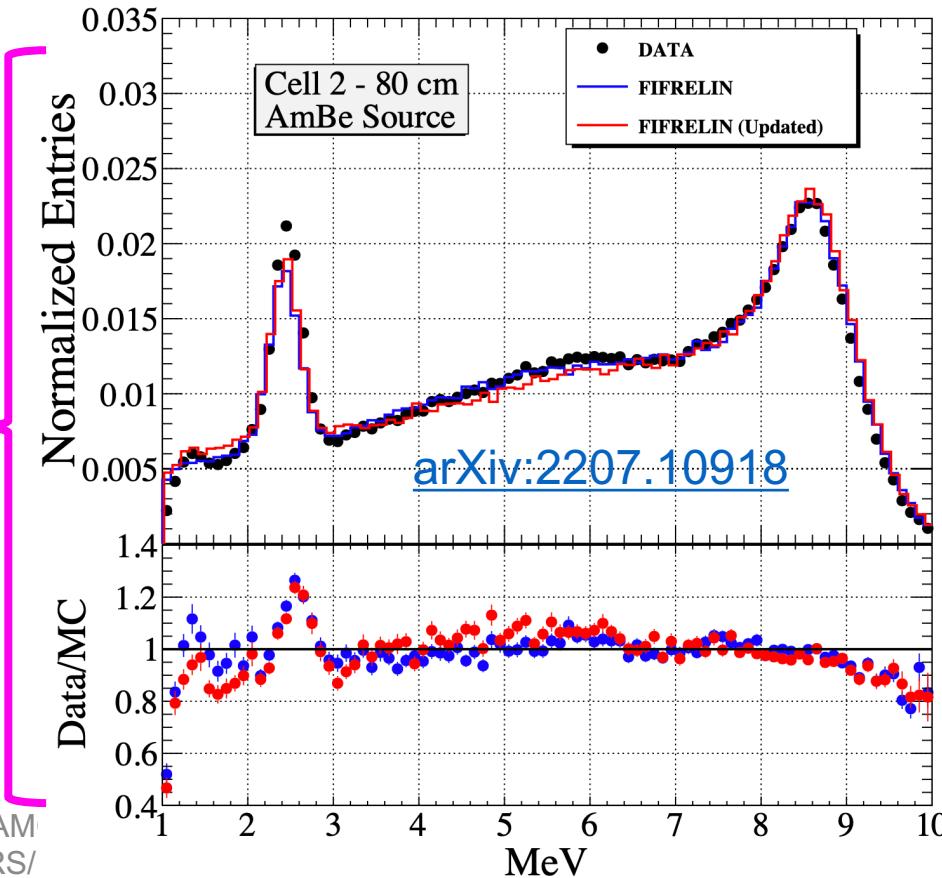


# New n-Gd $\gamma$ -cascade model

O. Litaize et al., [Eur. Phys. J. A \(2015\) 51: 177](#)



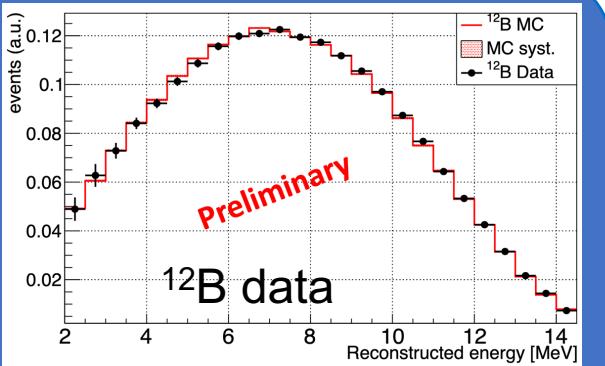
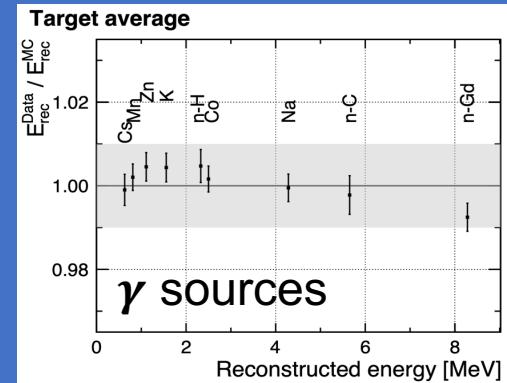
- Especially important for small detectors, where high energy gamma more likely to escape
- NEW:** FIFRELIN code improved: EGAF exp data, angular correlations, better conversion electrons [arXiv:2207.10918](#)



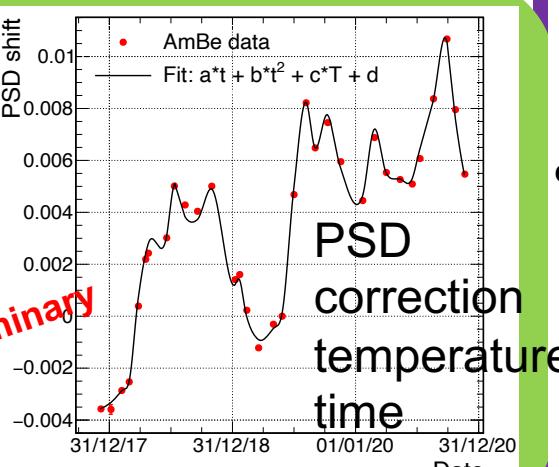
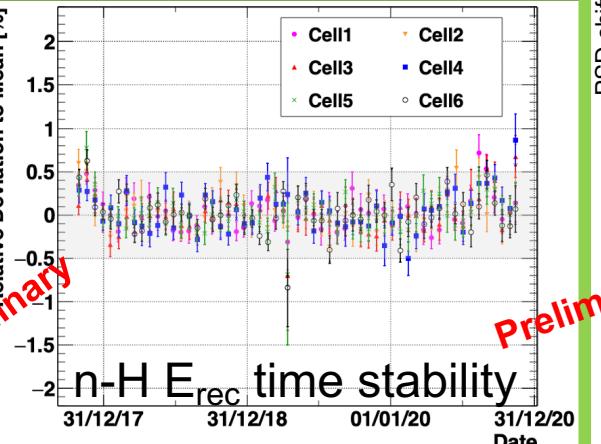
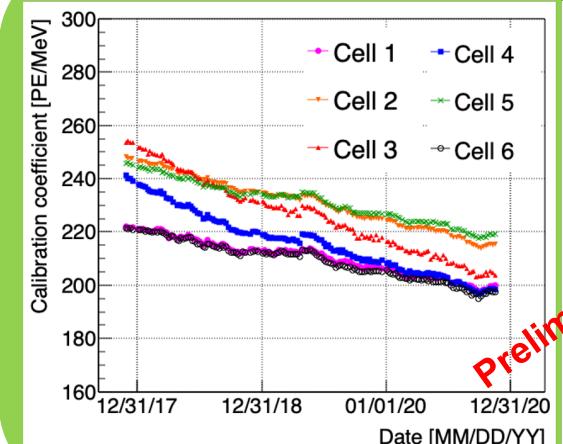
Better  
n-Gd peak  
data/MC  
agreement  
in borders  
of detector

# Detector response

## Energy scale



## Time stability

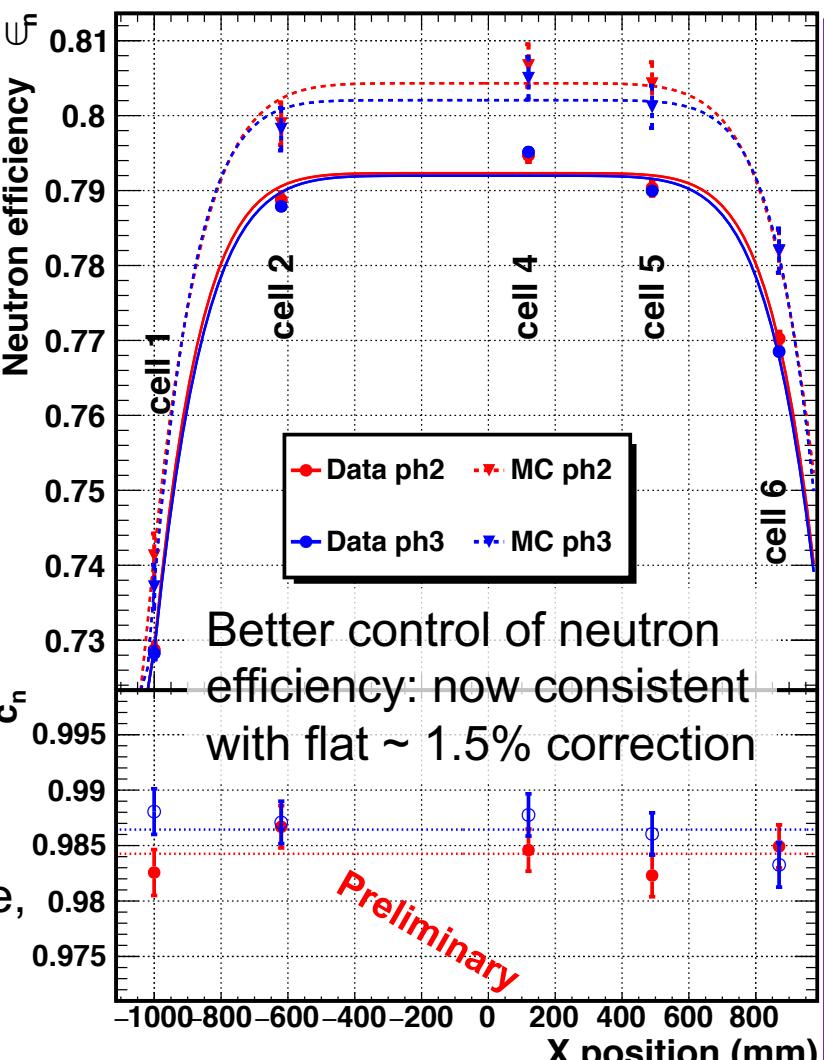


## Time stability

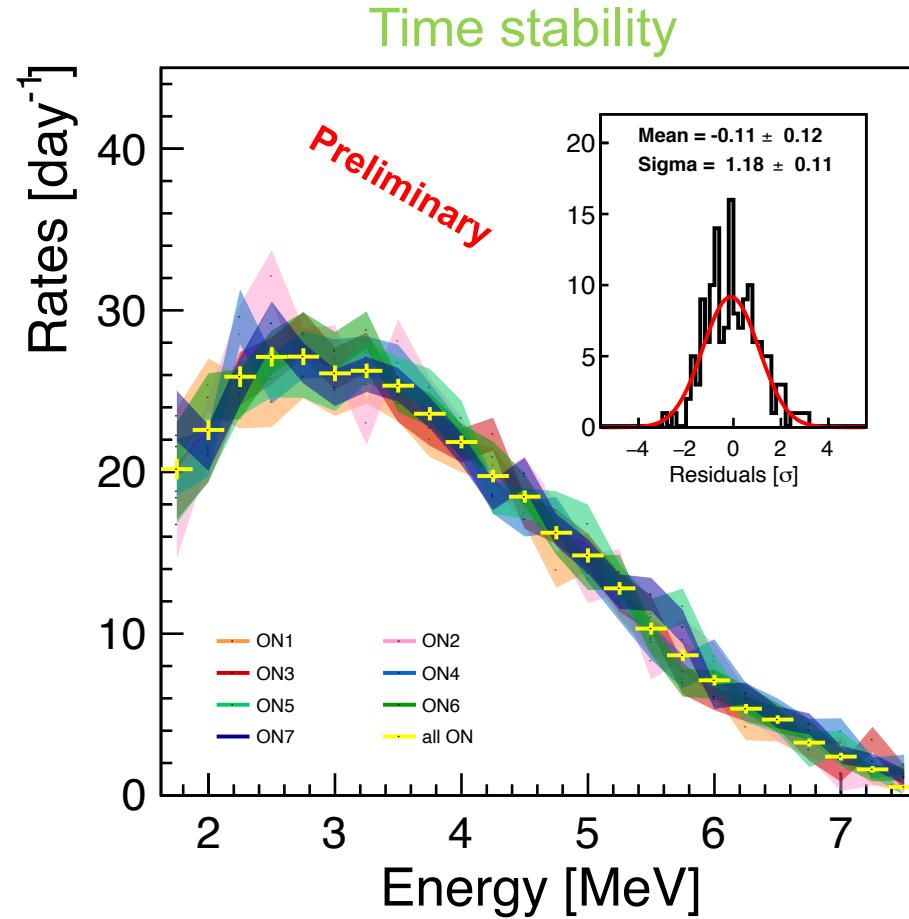
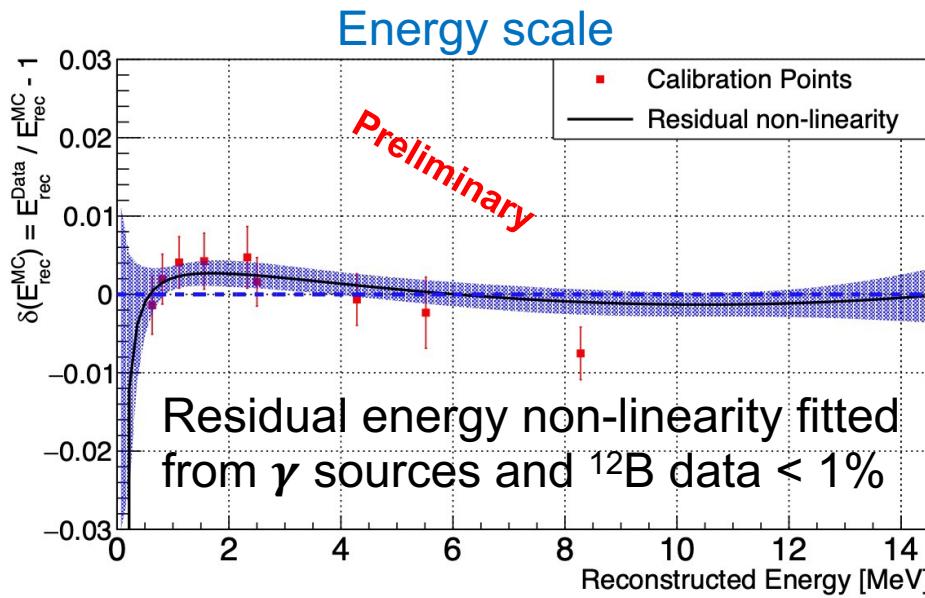
07/09/2022

Pablo DEL AMO SANCHEZ,  
LAPP - IN2P3 - CNRS/ U. Savoie Mont Blanc

## Neutron efficiency



# Validation and systematics



Stability of correlated background subtraction: comparison of IBD spectra extracted from every ON period using adjacent OFF data. Spectra are compatible with each other within statistical uncertainties.

For final neutrino signal extraction, all ON and OFF runs within same Phase are merged

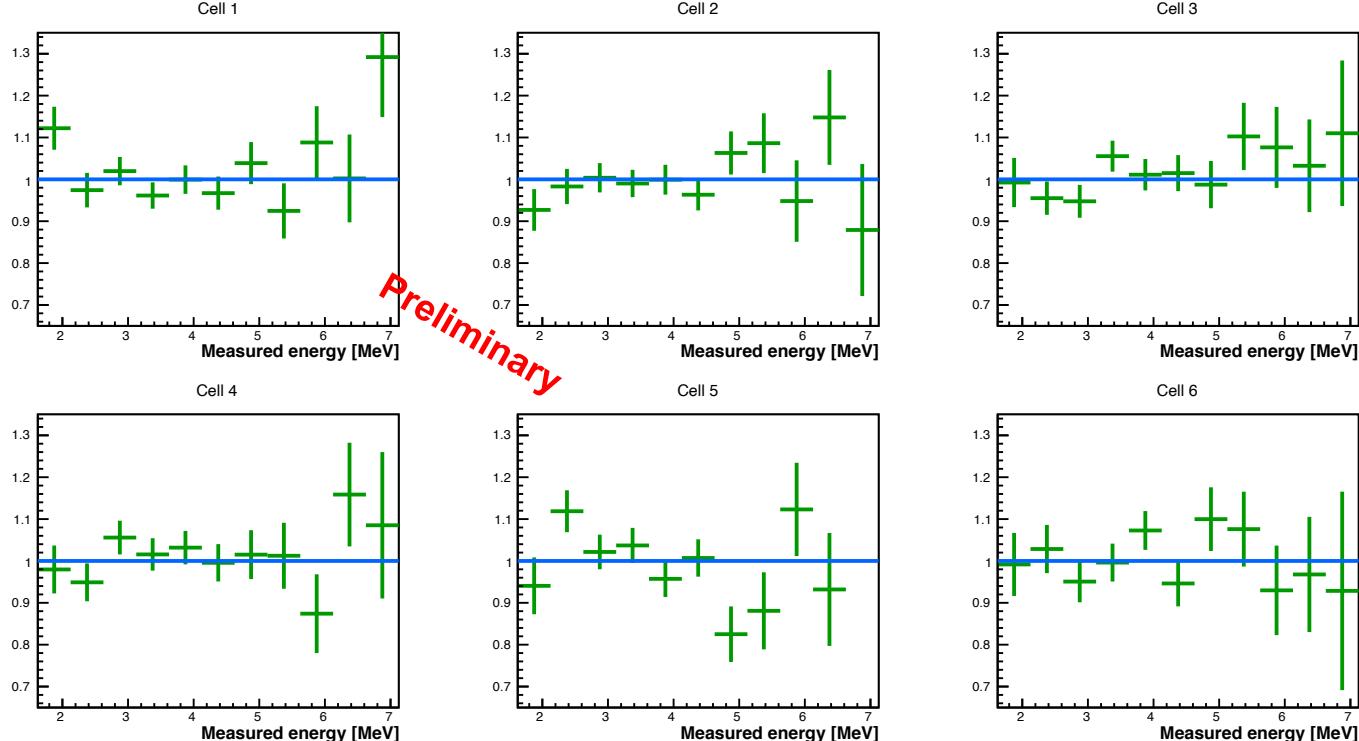
# Oscillation analysis

Compare data energy spectra of each cell  $D_{l,i}$  to prediction  $M_{l,i}$  for oscillation parameter values  $\sin^2(2\theta_{ee})$  and  $\Delta m^2_{41}$  ( $\alpha_j$ , nuisance parameters):

$$\chi^2(\sin^2(2\theta), \Delta m^2, \phi_i, \alpha_j) = \sum_l^{N_{cells}} \sum_i^{N_{Ebins}} \left( \frac{D_{l,i} - \phi_i \times M_{l,i}(\sin^2(2\theta), \Delta m^2, \alpha_j)}{\sigma_{l,i}} \right)^2 + \sum_j \left( \frac{\alpha_j}{\sigma_j} \right)^2$$

## Spectrum prediction-independent (free $\phi_i$ )

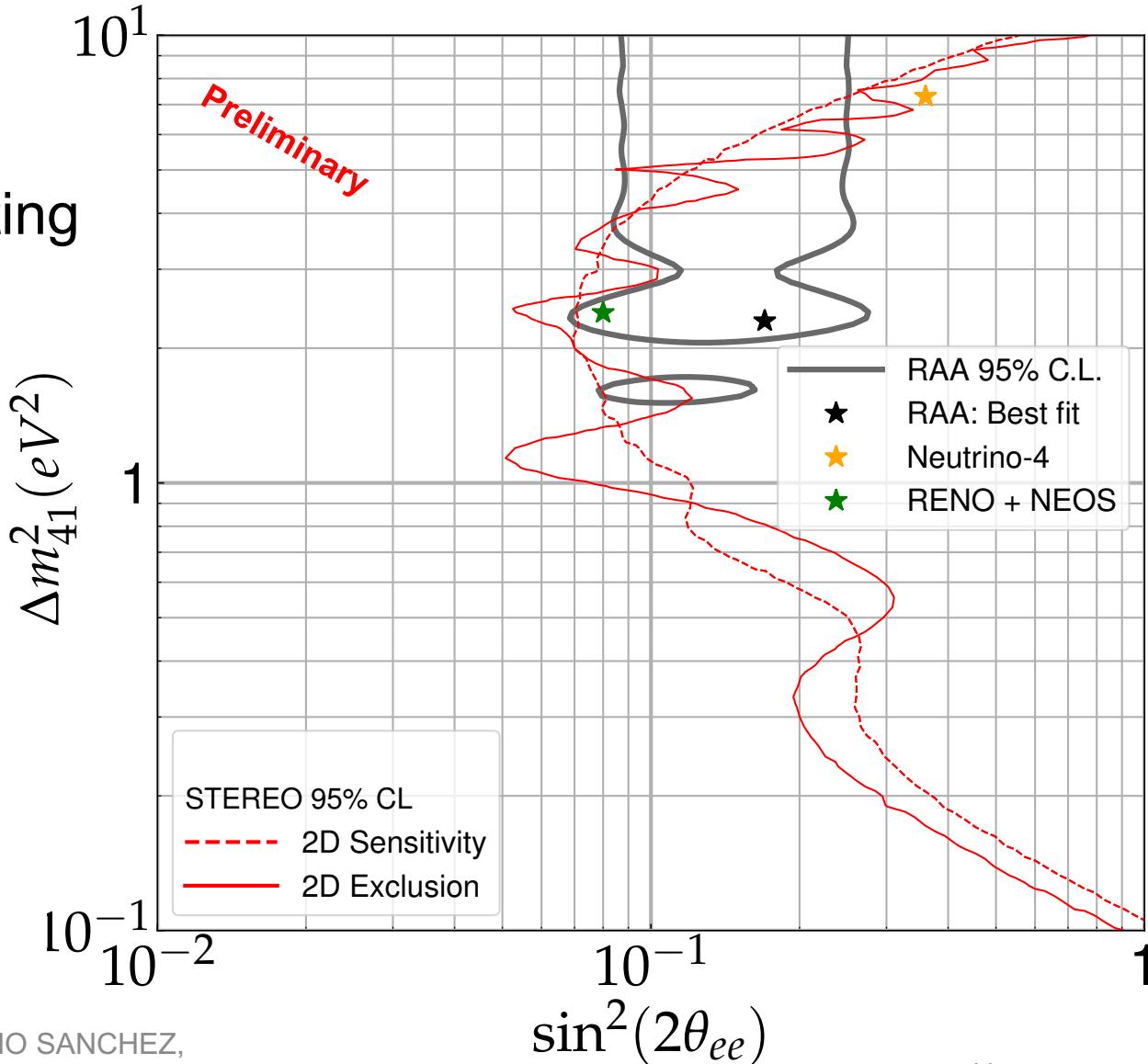
## Phase III data/No oscillation prediction ( $\phi_i, \alpha_j$ optimised)



Uncertainty type	Value (relative)
E scale (cell uncorrelated): cell-to-cell deviations	1.0%
E scale (corr): time stability	0.25%
Norm (uncor): cell volume $\oplus$ n efficiency correction	0.83% $\oplus$ 0.63%
Reactor background	0.7% (P-II: 118 days) 0.6% (P-III: 156 days)
Norm Phases II VS III ( $\sim \sigma(P_{th})$ )	1.5%

# Oscillation results

- Reactor spectrum prediction-independent
- $\Delta\chi^2$  pdf determined at each point in  $(\sin^2(2\theta), \Delta m^2)$  plane by generating and fitting pseudoexperiments (“Feldman-Cousins”)
- **Exclude most RAA allowed param space** at  $> 95\%$  CL for  $\Delta m_{41}^2 < 4 \text{ eV}^2$ 
  - No oscillation ***not*** excluded (p-value=0.54)
  - RAA best fit excluded at  $\gtrsim 4 \sigma$
  - Neutrino-4 best fit excluded at  $3.1 \sigma$
  - Neos-RENO best fit excluded at  $2.8 \sigma$
- Higher  $\Delta m_{41}^2$  in strong tension with cosmology, will be probed by KATRIN



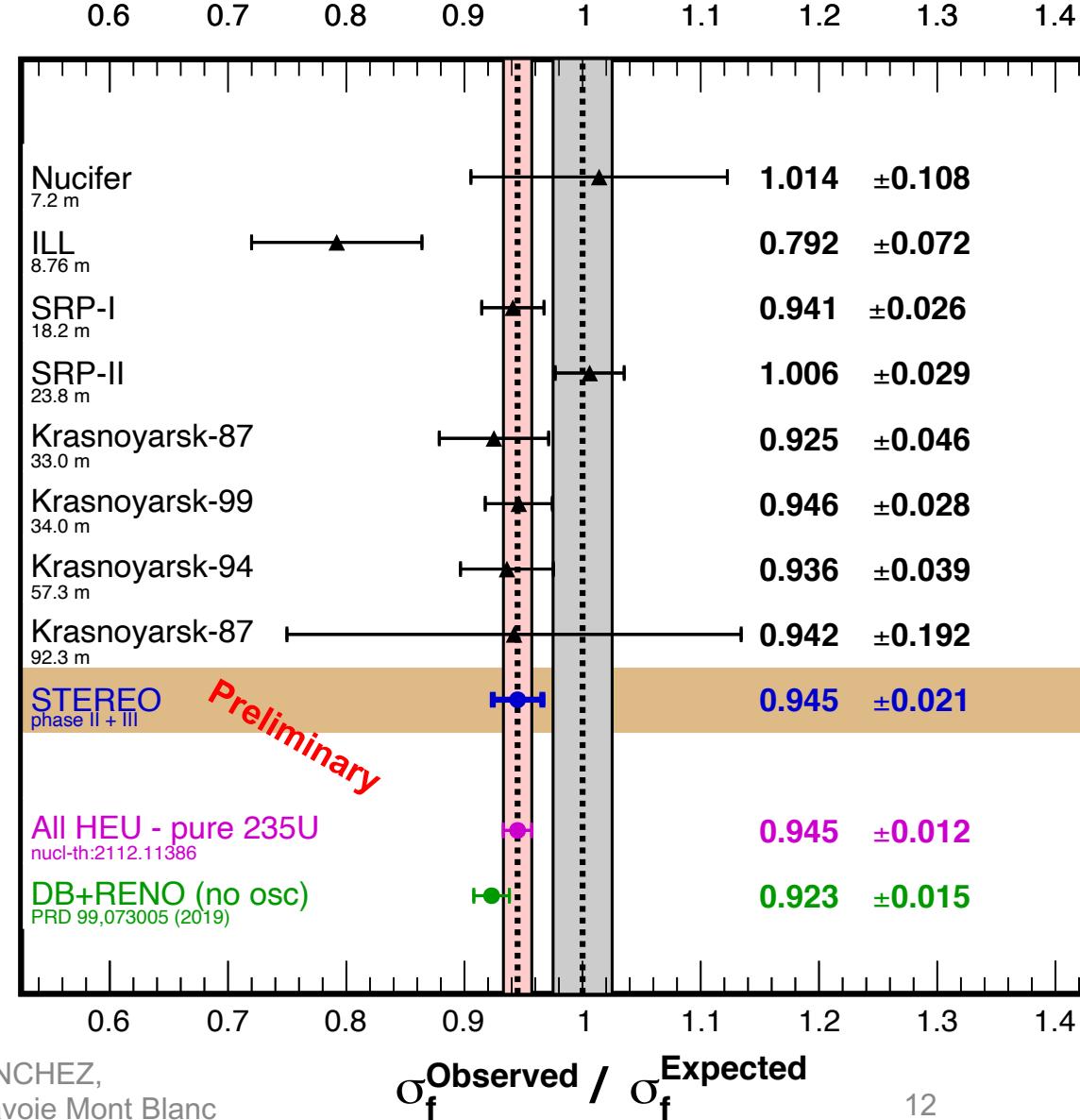
# Reactor flux

- Better neutron efficiency treatment (0.7% syst)
- Dominant exp error: reactor  $P_{th}$  (1.4% syst)
- No aging of diaphragm in primary circuit, as checked by ILL staff ✓

Normalization	Reactor thermal power Proton number Neutron efficiency Neutrino extraction method Solid angle Others [5] → Total	1.4% 1.0% 0.7% 0.7% 0.5% 0.3% <b>2.1%</b>
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Overall ~5% deficit confirmed

Most precise  $^{235}\text{U}$  flux measurement



# Reactor spectrum measurement

- Joint Phases II + III unfolding ( $E_{\text{rec}} \rightarrow E_{\nu}$ )

$$\chi^2(\lambda_l, \alpha_k) = \sum_p^{\text{phases}} \sum_j^{N_{\text{Ebins}}} \left( \frac{D_{j,p} - \varphi_p \times R_{ij,p}(\alpha_k) \times \phi_j(\lambda_l)}{\sigma_{j,p}} \right)^2 + \sum_k \left( \frac{\alpha_k}{\sigma_k} \right)^2 + \text{corr pull } (\varphi_p) + r \times \sum_l (\lambda_l - \lambda_{l+1})^2$$

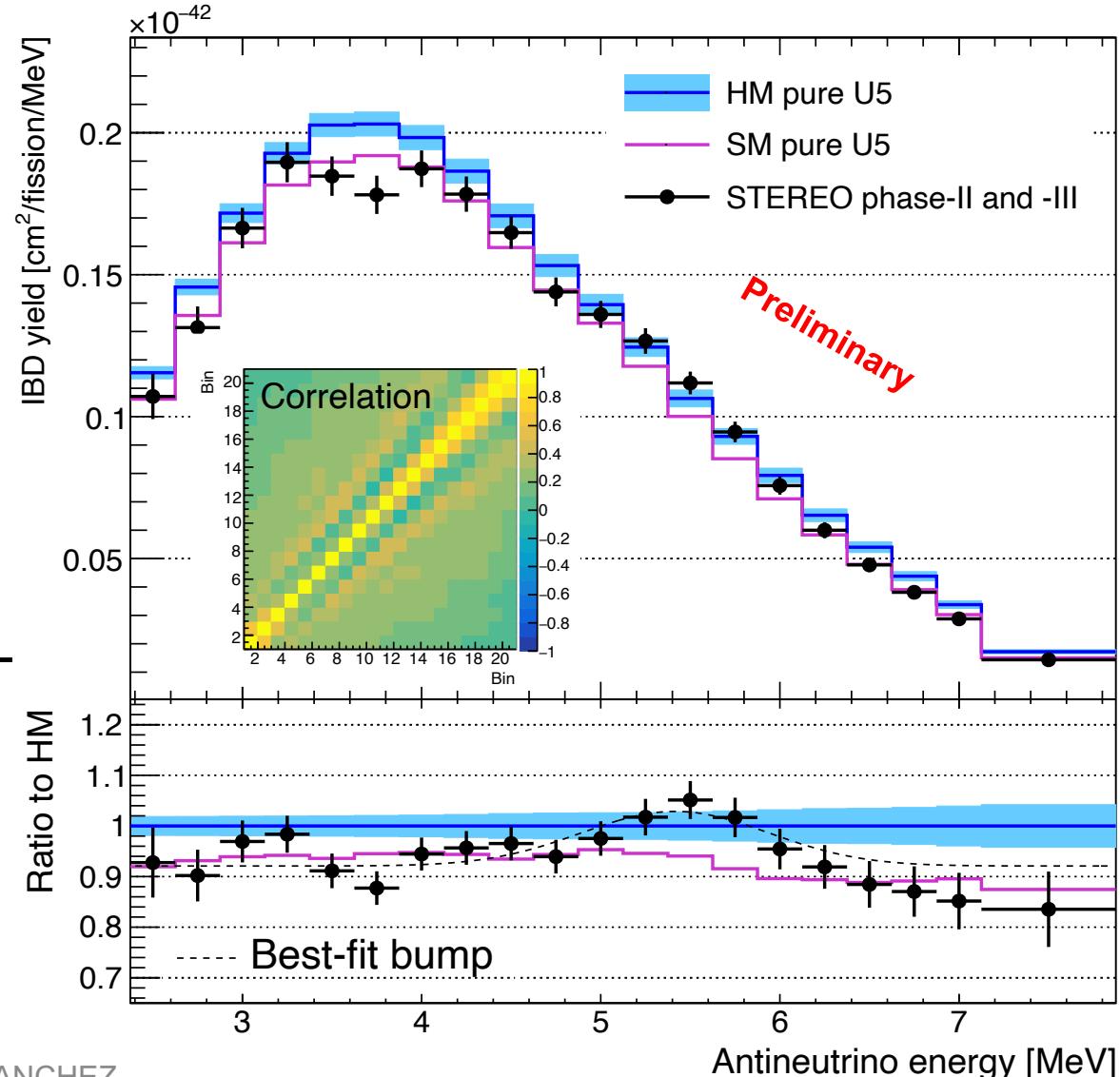
regularisation term

where:  $D_j$  = data energy spectrum bin

$\lambda_l$  = ratio of spectrum to ref spectrum (p.ex. HM)

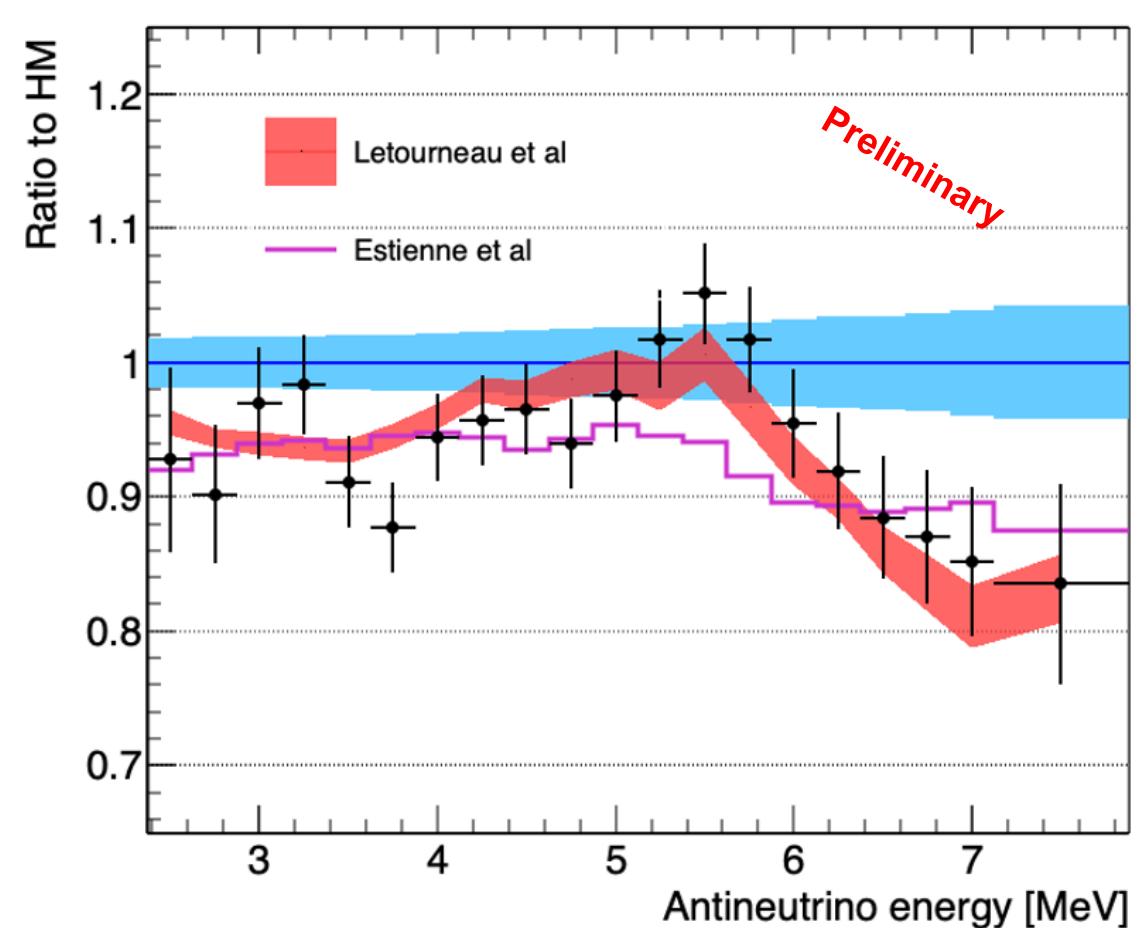
$R_{ij}$  = detector response matrix

- Tikhonov regularisation, strength  $r$  tuned with pseudo-experiments to minimise prior dependence. Covariance-matrix formalism gives ~identical results
- Bump observed ( $4.6 \sigma$ ) at  $\sim 5.5$  MeV
- Reference  $^{235}\text{U} \nu$  spectrum will be provided to community

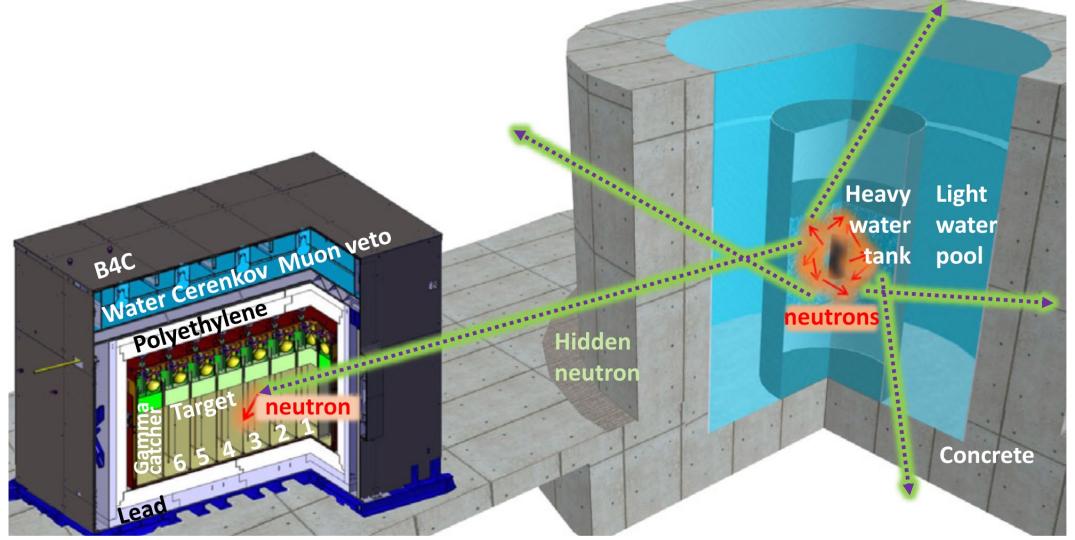


# Reactor neutrino spectra: tool for nuclear data validation

- RAA: if not a sterile neutrino, a nuclear data problem?
- Test nuclear data with reactor neutrinos!
- A. Letourneau *et al* ( arXiv:2205.14594, sub. to PRL ):
  - Observe bias when comparing ENSDF nuclear database data with isotopes measured with TAGS
  - Model of bias
  - Correct all isotopes (not TAGS measured) for modelled bias
  - Predict spectrum
  - Test predicted spectrum against STEREO's measured reactor spectrum
  - Good qualitative agreement!
- Towards reactor neutrino data as reference nuclear data?



# BONUS: neutrons and hidden branes



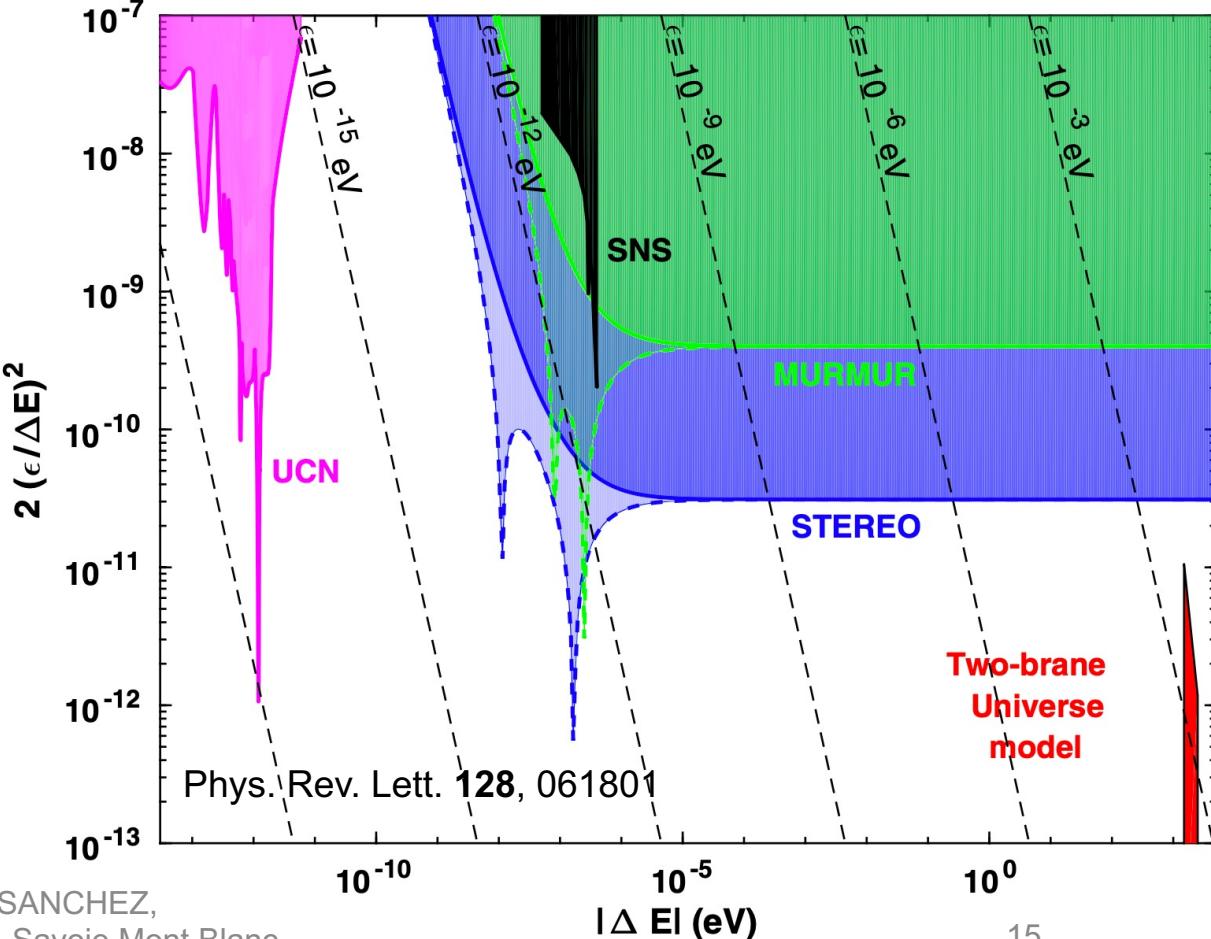
In some models:

- Neutron collision has  $\neq 0$  prob of sending neutron to hidden brane and viceversa
  - Hidden neutrons can escape reactor and reappear in STEREO
- constrain swapping probability parameters

$$\epsilon \text{ and } \Delta E$$

$\epsilon$ , coupling between hidden and visible states

$\Delta E$ , energy difference between hidden and visible states



# Summary & outlook

- We present here the analysis of STEREO's final dataset
- STEREO confirms ~5% deficit in reactor antineutrino flux
- Exclude most of the RAA allowed parameter space at low masses (below ~4 eV<sup>2</sup>)
- Provide most precise measurement of pure <sup>235</sup>U reactor  $\bar{\nu}$  spectrum  
→ Towards reactor neutrino data as reference nuclear data?
- Ongoing reactor  $\bar{\nu}$  spectrum joint unfolding with other experiments' data



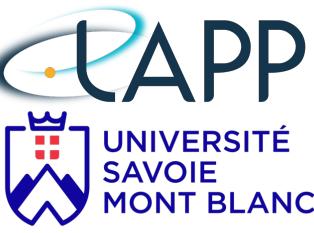


Photo: S. Schoppmann

### The STEREO Collaboration

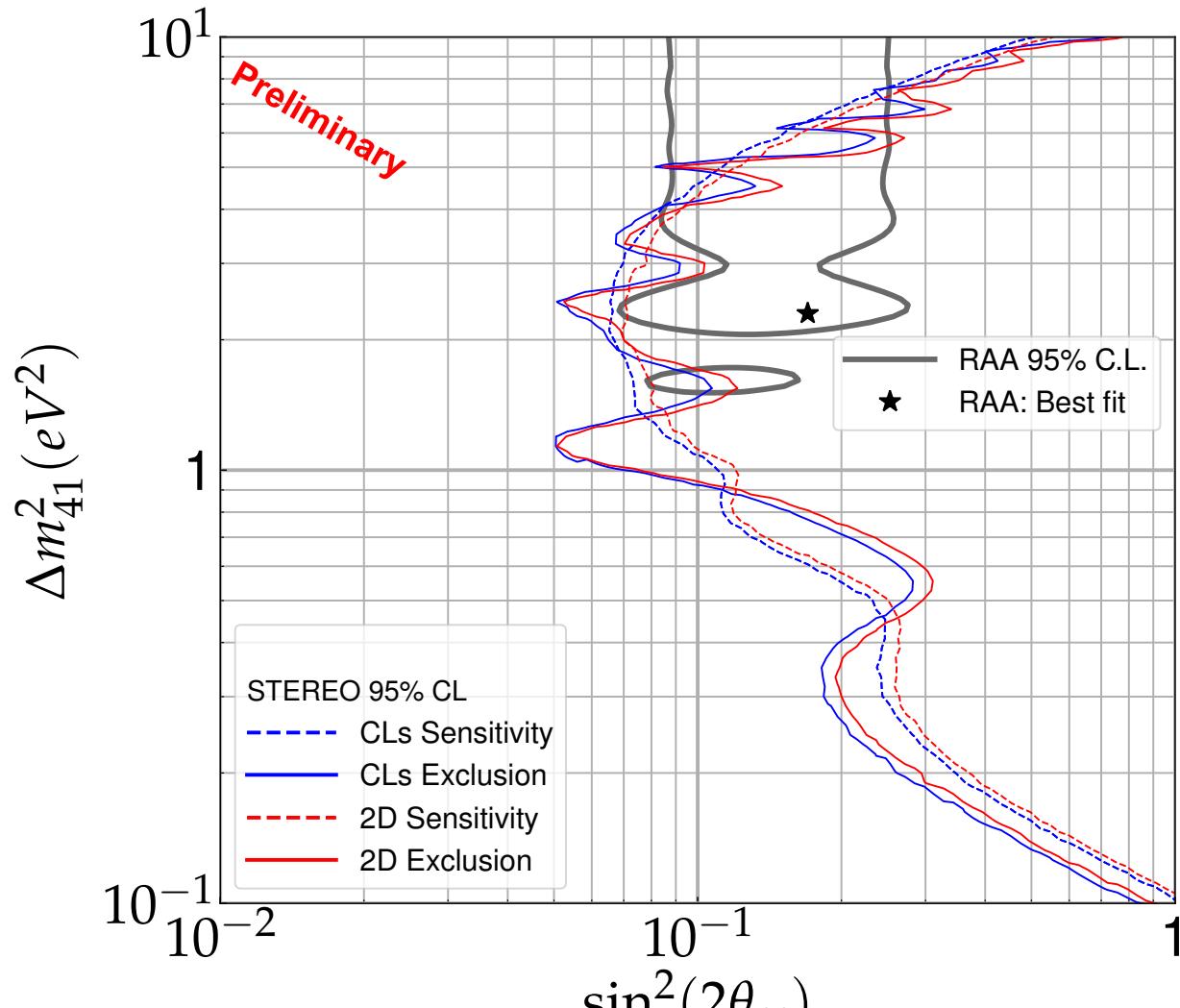
07/09/2022





# BACK UP SLIDES

# Statistics crosscheck: CLs VS 2D

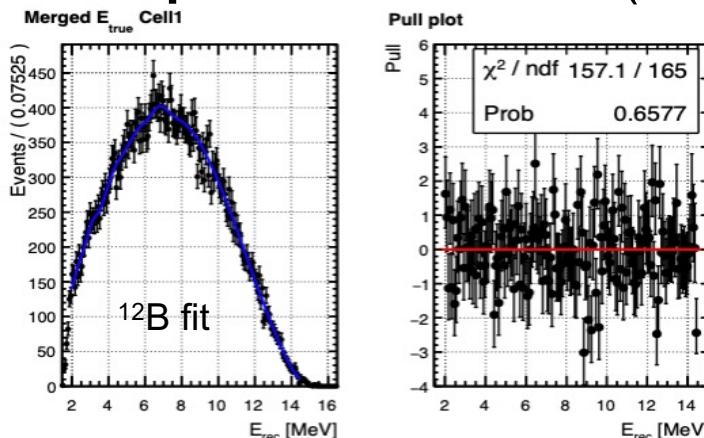


Gaussian CLs method:  
*Qian et al,*  
*Nucl.Instrum.Meth.A*  
827 (2016) 63-78

# Oscillation analysis crosscheck

Approach: instead of fixed response model  $M_{l,i}$  from Geant4 MC with allowed linear variation on each cell's energy and norm, use analytical parameterisation in order to leave free in fit ~all response model params.

- Parametrise and fit analytical response model: 15 params/cell (3 for  $E_{\text{rec}} = E_{\text{rec}}(E_{\text{true}})$ , 3 for  $\sigma = \sigma(E_{\text{rec}})$ , rest for shoulder and non-gaussian tail)
- PDF in  $E_{\text{rec}}$  = convolution of  $E_{\text{true}}$  PDF (free spectrum x acceptance x oscillation probability) with analytical response model
- Derive systematics from fit to  $^{12}\text{B}$  data
- Finer binning (250 keV VS 500 keV), only 118 reactor ON days
- **Independent, radically different oscillation search compatible with published results (Phys.Rev.D 102 (2020) 5, 052002)**



Data VS no  
oscillation  
model after  
nuisance  
params fit:

