3-flavour results with NOvA

Juan Miguel Carceller



University College London



February 25, 2021

NOvA

- Long baseline neutrino accelerator experiment
- A beam of neutrinos is produced in Fermilab
- The Near Detector is 1 km away from the source. 100 m underground, shielded against cosmic rays
- The Far Detector is 810 km away from the source and 14 mrad off axis, in Minnesota. On the surface, high rate of cosmic rays





Near Detector



Far Detector

NOvA

- Calorimetric, liquid scintillator detector
- The cells have horizontal and vertical orientation in each plane allowing for 3D reconstruction
- There is a fibre inside each cell that carries the light to the readout



Identifying events and particles

- For 3-flavour analyses, a Convolutional Neural Network (CNN) is used to identify particles
- Calibrated hits (with cosmic muons) are provided as pixel maps
- The network classifies these events in several signal and background categories
- Neutrino and antineutrino modes are trained independently



Neutrino energy estimation



 Energy estimated from muon length and hadronic shower

1

 Energy estimated from the electromagnetic shower and hadronic component

$$E_{v} = f (E_{EM}, E_{had})$$
Hadronic
EM Shower

The sample: POT

- Exposure is measured in Protons On Target (POT), the number of protons that collided with the target
- $13.6 \cdot 10^{20}$ POT in neutrino mode
- $12.5 \cdot 10^{20}$ POT in antineutrino mode



Making predictions

- To make predictions first we have to
 - Detect the neutrino event
 - Identify the particles
 - Reconstruct its energy

• The predictions for the FD are obtained from the ND: **extrapolation**



The ν_{μ} sample: Hadronic energy quartiles

Neutrino Mode

- Oscillation sensitivity depends on the shape of the spectrum
- By separating in fraction of hadronic energy, high and low resolution events are separated increasing the sensitivity to the oscillation dip NOvA Preliminary NOvA Preliminary



Neutrino Mode

The ν_{μ} sample: p_t extrapolation

- Detector acceptance is different in the transverse direction for the ND and FD because of their different size
- Bins of *p_t* are made to better match samples between the ND and FD and increase oscillation sensitivity



The u_{μ} sample



211 events with 8.2 background



105 events with 2.1 background

The ν_e sample



Bayes Theorem



- Credible intervals or regions (regions where the posterior is maximized)
- Posterior depends on the choice of prior, uninformative (i.e. constant) prior typically chosen
- Likelihood: binned Poisson likelihood (same as in the frequentist analysis)

$$\ln P(\vec{x}|\vec{\theta}) = -2\sum_{i=1}^{N} \left[E_i(\vec{\theta}) - O_i + O_i \ln \frac{O_i}{E_i(\vec{\theta})} \right]$$

Nuisance parameters (due to systematics) are marginalized (integrated)

Markov Chain Monte Carlo (MCMC)

- One way of aproximating the posterior by using a Monte Carlo sampling method
- Iterative process: at each iteration parameters are deviated and the new values can be accepted or be rejected (duplicating the previous value in the chain)
- The result is a series of values whose distribution approximates the posterior



- Two methods implemented in the CAFAna fitting framework in NOvA:
 - Metropolis-Hastings
 - Hamiltonian MCMC

Markov Chain Monte Carlo (MCMC)

Metropolis-Hastings

• Proposal function: gaussian, Σ is the step size, needs tuning

$$g(\vec{\theta}'|\vec{\theta}) = (2\pi)^{-\frac{N}{2}} (\det \Sigma)^{-\frac{1}{2}} \exp\left(-\frac{1}{2} \left(\vec{\theta}' - \vec{\theta}\right)^T \Sigma^{-1} \left(\vec{\theta}' - \vec{\theta}\right)\right)$$
(1)

Acceptance function, this is where the likelihood is included

T(

$$A(ec{ heta},ec{ heta'}) = \min\left(1,rac{P(ec{ heta'}|ec{ heta})}{P(ec{ heta}|ec{ heta'})}rac{g(ec{ heta}|ec{ heta'})}{g(ec{ heta'}|ec{ heta})}
ight)$$

- Draw next sample from (1) and then decide whether to accept or reject based on (2)
- Hamiltonian MCMC: Add momentum to the minimization $(H = -\log P(\vec{\theta}|\vec{x}))$

$$\frac{d\vec{q}}{dt} = \frac{\partial H}{\partial \vec{p}} = \frac{\partial T}{\partial \vec{p}}$$
$$\frac{d\vec{p}}{dt} = -\frac{\partial H}{\partial \vec{q}} = -\frac{\partial T}{\partial \vec{q}} - \frac{\partial V}{\partial \vec{q}}$$
$$\vec{q}, \vec{p}) = \frac{1}{2}\vec{p}^T M^{-1} \vec{p} + \log|M| + \text{cons}$$

(2)

Bayesian Results: δ_{CP} and $\sin^2 \theta_{23}$



-----3 σ

 $\frac{3\pi}{2}$

Bayesian Results: Δm_{32}^2



Slightly prefer upper octant (bayes factor 2.1)

• Slightly prefer normal mass ordering (bayes factor 1.7)

	N. Ordering	I. Ordering	
U. Octant	41.7%	20.9%	62.6%
L. Octant	25.8%	11.5%	37.4%
	67.5%	32.5%	



Bayesian Results: $\sin^2 2\theta_{13}$

- Typically θ_{13} is used to constrain the fit. In this case it is measured by NOvA
- Larger θ_{13} favour the lower octant and small θ_{13} favour lower octant
- Results agree with reactor experiments
- $\sin^2 2\theta_{13} = 0.085^{+0.020}_{-0.016}$



Jarlskog Invariant

 $J = c_{12}c_{13}^2c_{23}s_{12}s_{13}s_{23}\sin\delta_{\rm CP}$

- Measure of CP violation independent of the parameterization
- Prior flat in $\sin \delta_{\rm CP}$ at the top and flat in $\delta_{\rm CP}$ at the bottom



Comparison: Bayesian and Frequentist





Frequentist Results: $\sin^2 \theta_{23}$ results

Best fit $\sin^2 \theta_{23} = 0.57^{+0.03}_{-0.04}$ $\Delta m_{32}^2 = (+2.41 \pm 0.07) \cdot 10^{-3} \text{ eV}^2$

- Maximal mixing disfavoured at 1.1σ
- Upper Octant is preferred at 1.2σ





Frequentists Results: δ_{CP} results

- Excludes $\delta_{CP} = \frac{\pi}{2}$ at $> 3\sigma$ in IH and $\delta_{CP} = \frac{3\pi}{2}$ at $\sim 2\sigma$ in NH
- Normal hierarchy preferred at 1σ





Summary

- New Bayesian fit to study NOvA data, deeper look into the data
- Results agree with the frequentist fit: $\delta_{CP} = \frac{\pi}{2}$ rejected at 3 sigma, slight preference for the upper octant and normal mass ordering
- First measurement of θ_{13} by NOvA
- New results later this year (NOvA T2K fit)

Backup

The Beam



- 120 GeV protons collide with a graphite target
- Magnetic horns are used to focus charged pions
- Pions decay in the decay pipe, producing muon neutrinos or antineutrinos
- Muon monitors allow us to study the beam, for example beam alignment

Uncertainties



• Systematic uncertainties dominated by calibration and energy scale

Jarlskog invariant



Comparison: NOvA and T2K

