5-SEP-2022 @NOW 2022

RECENT OSCILLATION RESULTS AND FUTURE PROSPECTS OF SUPER-KAMIOKANDE

Yasuo TAKEUCHI (Kobe University) for SK collaboration





Inside of SK detector during refurbishment work (July 15, 2018)

The Super-Kamiokande collaboration





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~230 collaborators from 51 institutes in 11 countries (as of May 2022)

Outline

- SK detector & summary of recent results
- Atmospheric neutrino oscillation results
- Solar neutrino oscillation results
- Status and prospects of SK-Gd
- Summary

Super-Kamiokande detector

https://www-sk.icrr.u-tokyo.ac.jp/sk/



50 kton water Cherenkov detector

- ~2 m OD viewed by 8-inch PMTs
- 32 kt ID viewed by 20-inch PMTs
- 22.5 kt fid. vol. (conventional)
- SK-I: April 1996~

SK-VII is running

Physics targets:
Nucleon decay search
Neutrino oscillation study
Astrophysical neutrino search

Inner Detector (ID) PMT: ~11100 (SK-I,III~VII), ~5200 (SK-II) Outer Detector (OD) PMT: 1885



Summary of recent results



Nucleon decay search

- PRD 101, 052011 (2020), PRD 102, 112011 (2020), PRD 103, 012008 (2021)
- Current result: SK 450 kt·yr (27.2 kt enlarged fid. vol., SK-I~IV)
 τ/B(p→e⁺ π⁰) > 2.4x10³⁴ years (90%CL)
- Atmospheric v oscillation analysis
 - PRD97, 072001 (2018): SK 5326 days, 328 kt·yr
 - Current preliminary analysis: SK-I~V 6511 days, 484.2 kt·yr
- Solar v oscillation analysis
 - PRD94, 052010 (2016): (SK-IV: until Feb. 2014, 1664 days)
 - Current data set: SK 5805 days (until end of SK-IV) Updated in May 2022

Astrophysics

- Pre-supernova alert system: ApJ. 935, 40 (2022)
- Solar anti v search: Astropart. Phys. 139, 102702 (2022)
- DSNB search: PRD 104, 122002 (2021)
- GRB v search: PTEP 2021, 103F01 (2021)
- GW v search: ApJ. 918, 78 (2021)

May 2022

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<u>Atmospheric v</u>

- Produced by cosmic-rays in Earth's atmosphere
- Measure v flavor, energy, zenith angle, etc. in SK
 - Sensitive to θ_{23} and Δm_{32}^2 parameters.
- Matter effect: enhancement of v_e in several GeV energy region and in Earth core
 - θ_{13} and mass hierarchy could be studied.
- Solar term: enhancement of v_e in sub-GeV region
 - θ_{23} octant degeneracy could be studied.
- Interference: CP phase could be studied.
- In SK-Gd: purity of v_e-like will be improved.



https://www-sk.icrr.u-tokyo.ac.jp/en/sk/about/research/



Neutrino Energy [GeV] (d) $P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$

8

Atmospheric v categories in



Zenith angle & lepton momentum distributions

UPER



Data files (PRD97, 072001 (2018) version): https://www-sk.icrr.u-tokyo.ac.jp/sk/publications/result-e.html#atmosci2018

Recent progresses in atm. v analysis

- PRD97, 072001 (2018): SK 5326 days, 328 kt·yr
- Updates in May 2020
 - 364.8 kt·yr
 - Incorporated neutron tagging for ν / ν̄ separation
 - New multi-ring event selection (likelihood \rightarrow BDT)
 - Updated MC simulation (NEUT 5.4.0), improved systematic errors
 - Changed zenith binning
 - Applied constrain with sin²θ₁₃
 - Updates in May 2022
 - Added SK-V: SK-I ~ V, 6511 live days, 484.2 kt·yr
 - Enlarged fid. vol.: 27.2 kt (conventional: 22.5 kt)
 - Applied improved T2K model constrain
 - Run 1-9, including recent v data
 - with sin²θ₁₃=0.0220+/-0.0007
 - Cf. joint fit: SK-IV only + T2K Run 1-10

In this afternoon:

Junjie Xia (ICRR), "T2K-SK joint nu oscillation sensitivity"

Zenith angle distributions using neutrons (SK-IV)

UPER



Enlarged fiducial volume

- Vertex distance from nearest ID surface = "wall"
- "wall > 2 m" is conventional in SK \rightarrow enlarged to 1 m
- No significant bias in reconstruction
- No significant increase of external background





(preliminary)

300



Atm. v oscillation analysis with T2K model



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Solar neutrinos

- Produced by nuclear fusion in the sun (initial = v_e)
- High statistics measurement of flux, energy spectrum, and time variations of ⁸B solar neutrinos in SK.
- Energy spectrum: sensitive to θ_{12} and Δm_{21}^2 parameters. Global analysis is more powerful.
- Matter effect in the sun: energy dependence of v_e survival probability is observed.
 - SK is trying to measure "Upturn"
- Time variation: matter effect in the earth could be studied from day-night flux asymmetry.
 - Sensitive to Δm_{21}^2 parameter.



Recent progresses in solar v analysis



PRD94, 052010 (2016): (SK-IV: until Feb. 2014, 1664 days)

Updates in May 2020

- Improved analysis tools: energy reconstruction, spallation cut
 - Spallation cut: consider cosmogenic neutrons (arXiv:2112.00092)
- Did oscillation analysis (spectrum and day-night) with full SK-IV (2970 days)

Δm² difference between Solar and KamLAND reduced.

Confirmed low-energy data quality in SK-V

Updates in May 2022

- Re-estimated systematic errors, then re-do oscillation analysis (spectrum and day-night) up to SK+SNO (using SK-I~IV data).
- Global analysis with other solar experiments is ongoing.
- First solar signal in 2.49-3.49 MeV (kinetic) in SK was obtained.
- Confirmed energy scale in low-energy region in SK-V and SK-VI

SK-IV solar neutrino signal



(preliminary) May 2020

- According to improved spallation cut, signal efficiency in solar analysis was increased by 12.6% in 3.49-19.49 MeV (kinetic).
- Signal (in SK-IV): 63890 +381 -379(stat.) +/-907 (syst.) events.
 - Total: > 10⁵ solar v events (during SK-I~IV, 5805 days)



Solar v oscillation results

- Re-estimated systematic errors, then re-do oscillation analysis
- No big change from the preliminary results in 2020
- ~1.5 σ level tension in Δm^2_{21} between SK+SNO analysis and KamLAND is still remaining.





Preliminary

SK 5805 days

Spectral fitting (quadratic)

The same parametrization with SNO (PRC 88, 025501 (2013))

A combined fitting can be done.



FIG. 29. Allowed survival probability 1σ band from the combined data of SK and SNO (red). Also shown are predictions based on the oscillation parameters of a fit to all solar data (green) and a fit to all solar + KamLAND data (blue). The pastel colored bands are the separate SK (green) and SNO (blue) fits.

JPFR

May 2022

Solar v signal with lowest energy in

May 2022

Preliminary

Use Wideband Intelligent Trigger (WIT) data in SK-IV

- SK-IV WIT data, 2.49 MeV 3.49 MeV electron kinetic energy, 858 live days good data during Oct. 2015 - May 2018
- Applied a boosted decision tree (BDT) selection and the standard spallation cut for the 858 days WIT sample.
 - Inputs: standard reconstructed variables in the traditional solar analysis
 - ~6 times better background rejection comparing to the traditional solar analysis
- A clear peak is observed.
 - First time in SK
- References:
 - <u>10.5281/zenodo.6759244</u>
 - J. Phys. Conf. Ser. 888, 012189 (2017)

Solar angle distribution (SK-IV WIT 858 days, 2.49-3.49 MeV)



<u>SK-V solar v data</u>



Cleaning of inside detector was done during the refurbishment work before SK-V.

- Better water flow control was achieved in SK-V using additional water piping.
- Low-energy background level in SK-V is lower than SK-IV.
 - Probably lower radon-222 contamination



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SK-Gd Phase



SK-Gd Phase:

Add gadolinium (Gd) to enhance neutron tagging efficiency of the SK detector.

Physics targets:

- Detect the world's first Supernova Relic Neutrinos (SRN) (or Diffuse Supernova Neutrino Background, DSNB)
- Improve pointing accuracy for supernova
- Early warning of nearby supernova from pre-burst signal (silicon burning)
- Enhance v or v discrimination in atmospheric v & T2K analysis
- Reduce backgrounds in proton decay search

SK refurbishment was done before SK-V

- Fix water leakage, clean inside the detector
- Replace dead PMTs
- Improve water piping in the SK detector



Capture efficiencies in water

- 0.01% Gd [Gd₂(SO₄)₃ 10t] : ~50%
- 0.03% Gd [Gd₂(SO₄)₃ 30t] : ~75%
- 0.1% Gd [Gd₂(SO₄)₃ 100t] : ~90%

Gd concentration SK-VI: 0.011% (18 Aug. 2020-) SK-VII: 0.03% (5 Jul. 2022-)

Calibration with neutrons in SK-VI

NIM A 1027 (2022) 166248

- Using Am/Be + BGO source, response of SK-Gd detector to neutrons was studied.
 - Trigger: scintillation light of BGO by 4.4 MeV gamma-ray from Am/Be
 - Neutron capture time was measured
- Geant4.9.6p04 and G4NDL4.2 are used to obtain the relation between Gd concentration and capture time constant.
- Observed neutron capture time = $115 + -1 \mu s \leftarrow \rightarrow 111 + -2 ppm$ Gd concentration.
 - Consistent with other estimations



Neutron tagging tools in SK-Gd

- Several tools are under development to detect neutron events in SK-Gd
 - Neural network, cut-based method, ...
- They are almost ready, but need more fine tunings
 - Neutron multiplicity, neutron tag efficiency, ...





Atmospheric v in SK-V & SK-VI

N Preliminary C May 2022

F٧

SK5

SK6

SK5

SK6

0.0

 $\cos\theta_{zenith}$

Multi-GeV 1-ring μ -like (FCFV)

0.0

 $\cos\theta_{zenith}$

0.5

0.5

10

20

1.0

1.0

0

Vertex z [m]

Multi-GeV 1-ring e-like (FCFV)

-20

-1.0

-1.0

-10

-0.5

SK5

-0.5

🕂 SK6

 Compare basic distributions and event rates in SK-V and SK-VI data.

Data set:

SK-V 461 days
 Feb. 2019 – Jul. 2020

SK-VI 577 days
 Jul. 2020 – Apr. 2022

They are consistent.



<u>Neutron signals in SK-VI atmospheric v</u>

BK

- Applied the neural network neutrino tagging to the SK-VI 577 days atmospheric v sample.
- The peak position of the neutron travel distance looks to be energy dependent.
- The capture time of neutron looks similar.
- Expected improvement in atmospheric v analysis in SK-Gd:
 - Purity of v_e-like will be improved. It improves δ_{CP} and MH sensitivities.
 - Reference: <u>10.5281/zenodo.6781493</u>



Solar neutrino signal in SK-VI



Preliminary

- Applied conventional solar selections to SK-V and SK-VI data above 6.49 MeV.
- No significant difference among SK-IV, SK-V, and SK-VI.
- Spallation cut in solar neutrino analysis would be more efficient in SK-VI & VII, thanks to improved cosmogenic neutron tagging (under study)
- Below 6.49 MeV: under study (need careful treatment of trigger efficiency, water convection, PMT dark rate, etc.)



Neutron events in SK-VII

Preliminary

- Cosmic-ray muons produce neutrons by spallation.
- Theses cosmogenic neutrons are observed in SK-IV and in SK-VI
 - Reference: <u>arXiv:2112.00092</u> (SK-IV) and <u>M. Shinoki@UGAP2022</u> (SK-VI)
- At the beginning of SK-VII, the cosmogenic neutron candidates are monitored when the additional Gd is being loaded from the bottom of the SK detector.
- Am/Be + BGO calibrations are also carried out.
- Clear increase of the neutron candidates and shorter capture time is observed.



<u>Summary</u>

- We are improving atmospheric and solar neutrino oscillation analyses at SK, continuously.
 - Current preliminary results use:
 - 6511 live days (SK-I~V, for atmospheric)
 - 5805 live days (SK-I~IV, for solar)
- The new phase (SK-Gd) is started in 2020.
 - SK-VI: 0.011% Gd concentration (18 Aug. 2020-)
 - SK-VII: 0.03% (5 Jul. 2022-)
- Clear neutron signals are observed in SK-VI and SK-VII.
- Improvements in the oscillation analyses are expected.
 - Atmospheric v: better v / \overline{v} separation \rightarrow better v_e-like purity \rightarrow improve δ_{CP} and mass hierarchy sensitivities
 - Solar v: better efficiency of cosmogenic neutron tagging → better spallation cut → less statistical uncertainties of solar v measurements

Supplements





Typical high-energy events

Atmospheric v: Partially contained (PC)

Atmospheric v: Fully contained (FC)

Typical low-energy event

Resolutions (for 10 MeV electrons)Energy: 14%Vertex: 87cmDireEnergy: 14%Vertex: 55cmDire

(software improvement) Direction: 26° SK-I Direction: 23° SK-III, IV

SK-Gd sensitivity on DSNB search

Preliminary May 2022

