

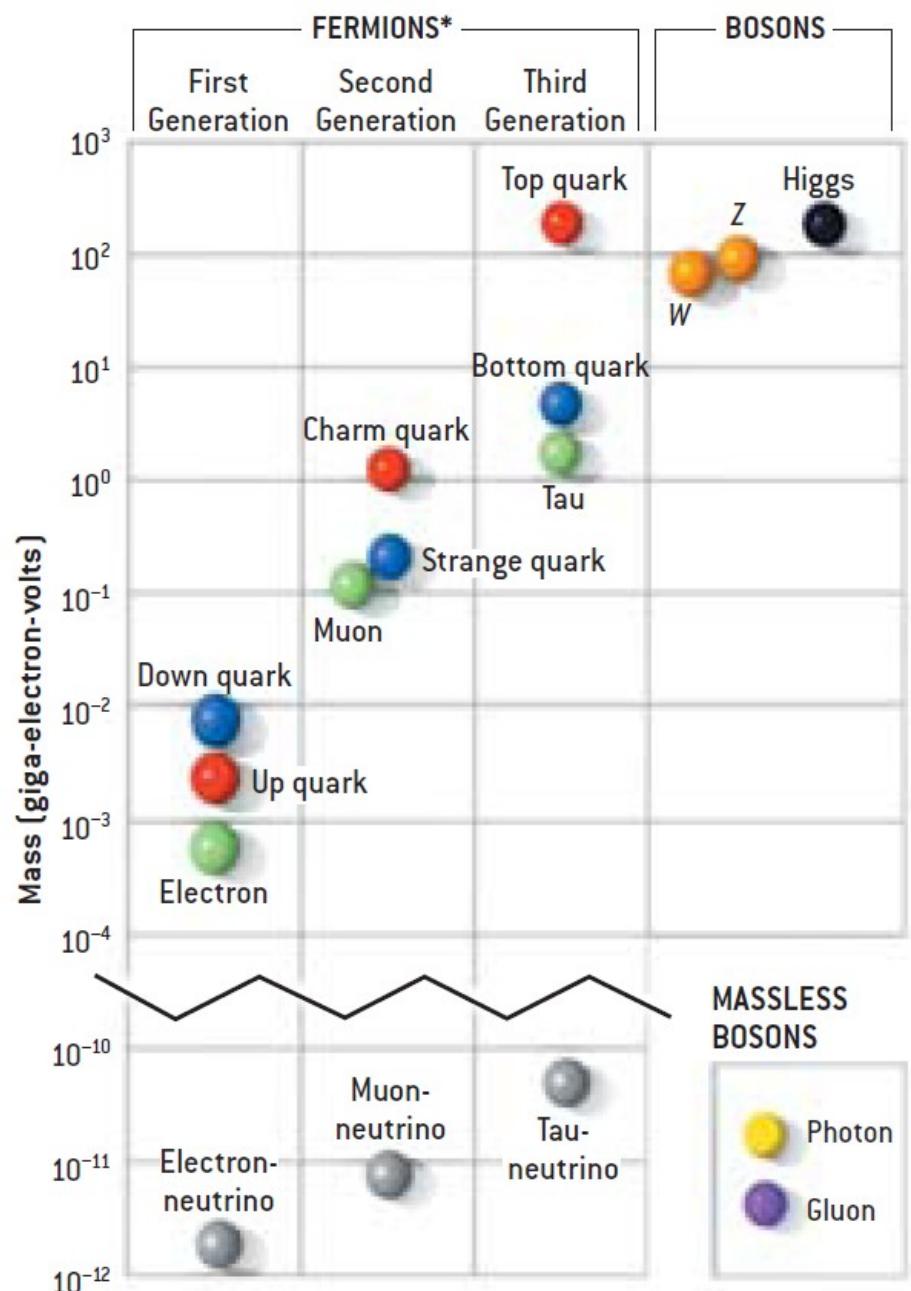
First Sub-eV Neutrino Mass Limit from the KATRIN Experiment

V.M. Hannen for the KATRIN collaboration

Institut für Kernphysik,
Westfälische Wilhelms-Universität Münster

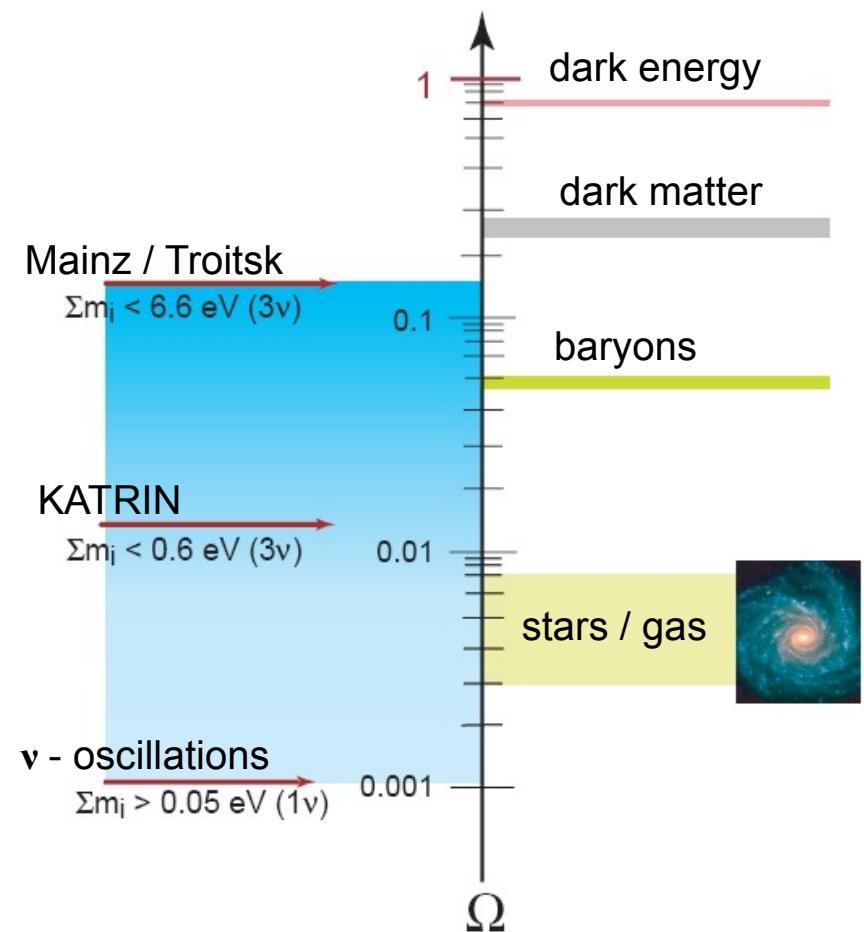
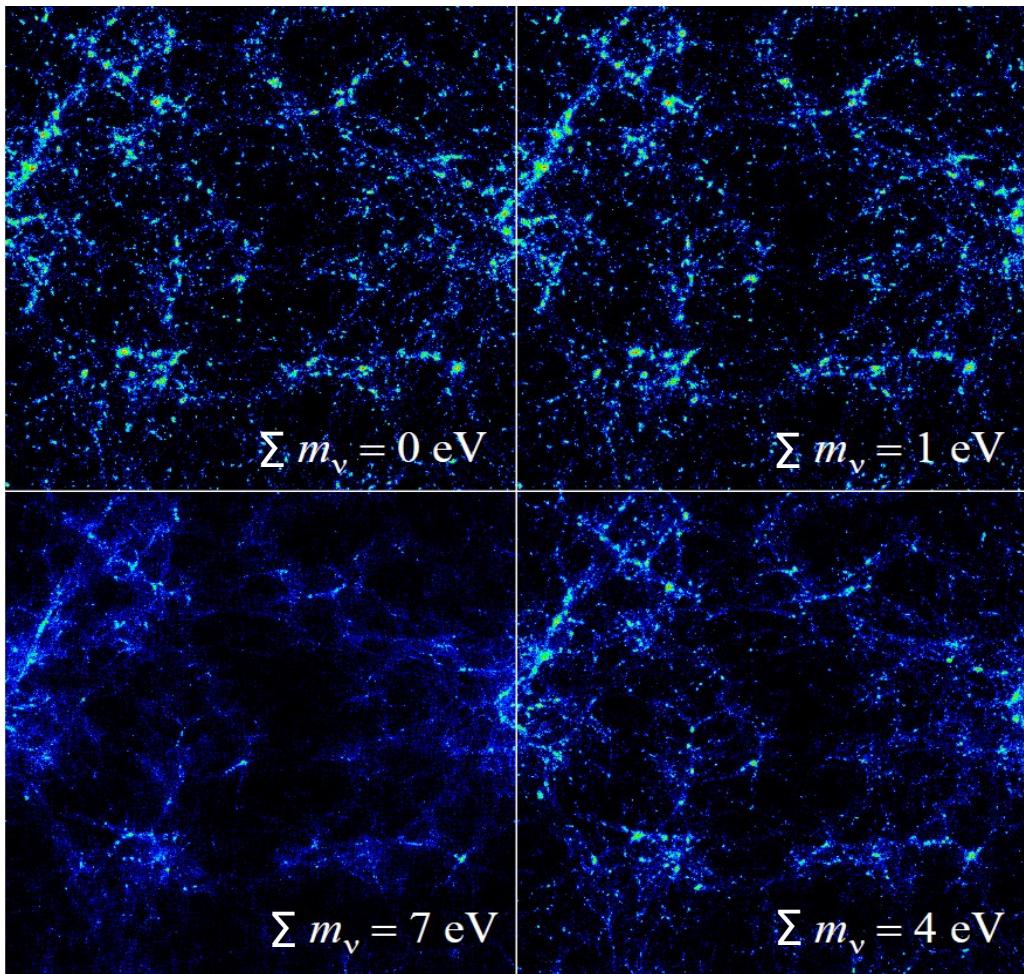
Neutrino mass in particle physics

- Nature of the neutrino: Majorana or Dirac particle, i.e. is the neutrino it's own anti-particle ?
- How to explain the many orders of magnitude difference between neutrino mass limits and masses of the charged fermions of the standard model
→ sea-saw type I and type II mechanisms
- Possible connection to the generation of the observed matter - antimatter asymmetry in the universe
→ leptogenesis



Neutrino mass in cosmology

- Neutrinos are (after γ 's) the second most abundant particle species in the universe
- As part of the hot dark matter, neutrinos have a significant influence on structure formation



- For large $\sum m_\nu$ values fine grained structures are washed out by the free streaming neutrinos

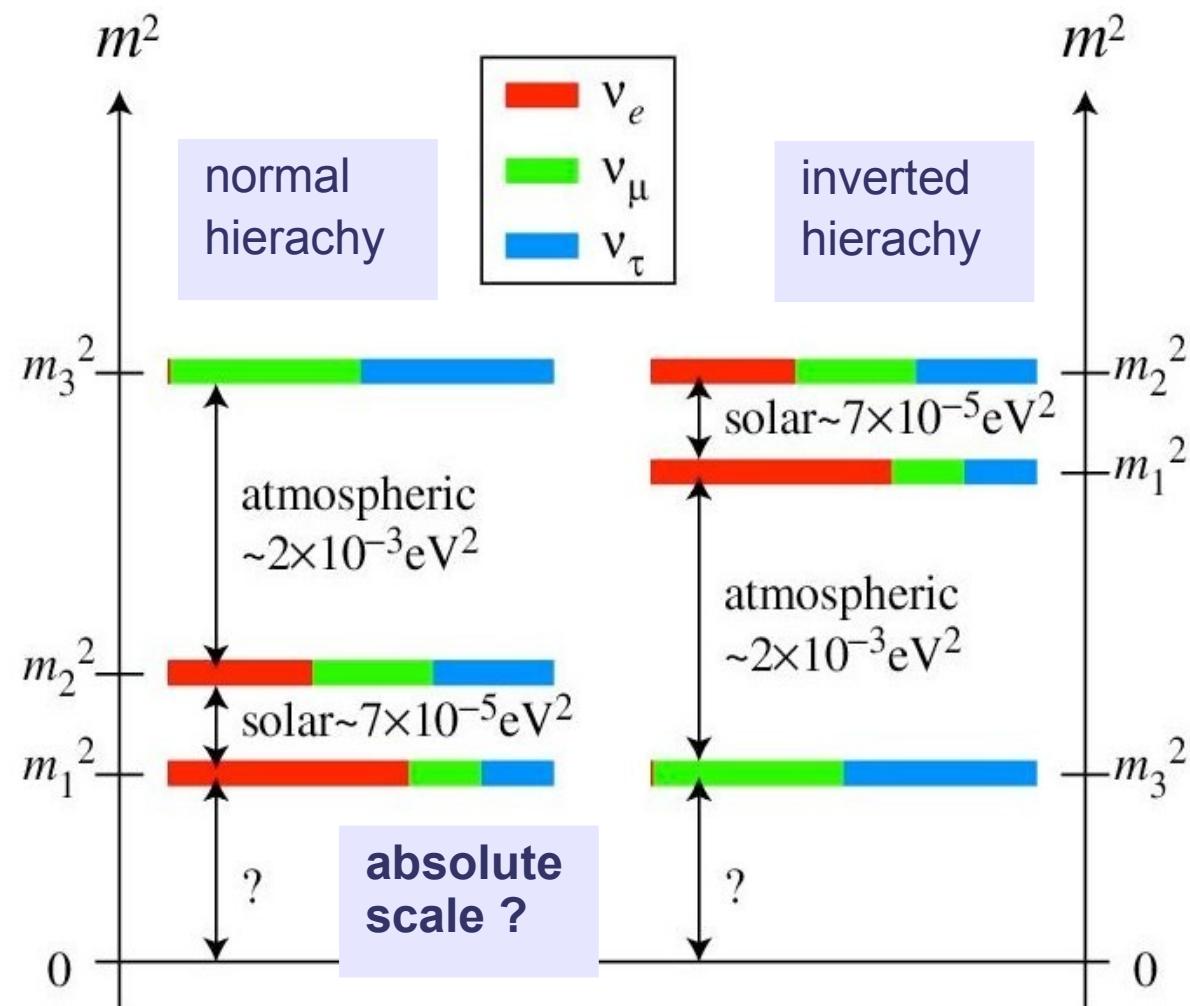
Chung-Pei Ma 1996

What we know (from ν oscillations):

- Neutrino flavour eigenstates differ from their mass eigenstates
- Neutrinos oscillate, hence they must have mass
- Mixing angles and Δm^2 values known (with varying accuracies)

What we don't know :

- Normal or inverted hierarchy ?
- Dirac or Majorana particle ?
- CP violating phases in mixing matrix ?
- **No information about absolute mass scale !** (only upper limits)
- Existence of sterile neutrinos ?



Search for neutrino mass

β -decay: absolute ν -mass

model independent, kinematics

status: $m_\nu < 0.8$ eV

potential: $m_\nu \approx 0.2$ eV

e.g.: KATRIN, Project-8, ECHO,
HOLMES, NuMECS

$0\nu\beta\beta$ -decay: eff. Majorana mass

model-dependent (CP-phases)

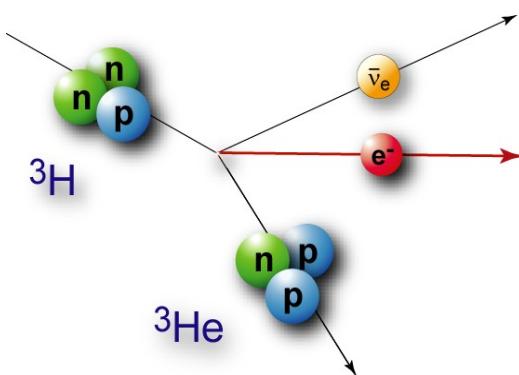
status: $m_{\beta\beta} < 61$ meV to 165 meV

(^{136}Xe , PRL 117 (2016) 082503)

potential: $m_{\beta\beta} \approx 20\text{-}50$ meV

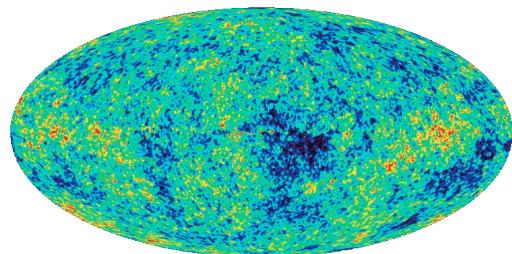
e.g.: KamLAND-Zen, GERDA, CUORE, EXO,
SNO+, Majorana, Nemo 3, COBRA

$$m_\nu^2 = \sum |U_{ei}|^2 m_i^2$$



neutrino mass measurements

$$m_{\beta\beta} = \left| \sum U_{ei}^2 m_i \right|$$

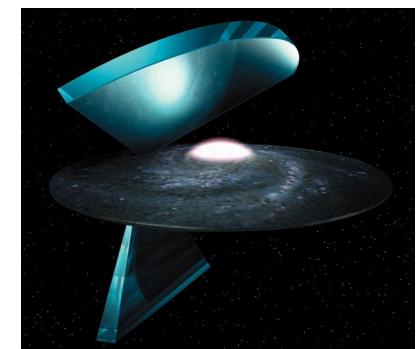


cosmology: ν hot dark matter Ω_ν

model dependent, analysis of CMB and structure formation data

status: $\sum m_i < 0.12$ eV

(Planck 2018 results,
Astron. & Astrophys. 641 (2020) A6)



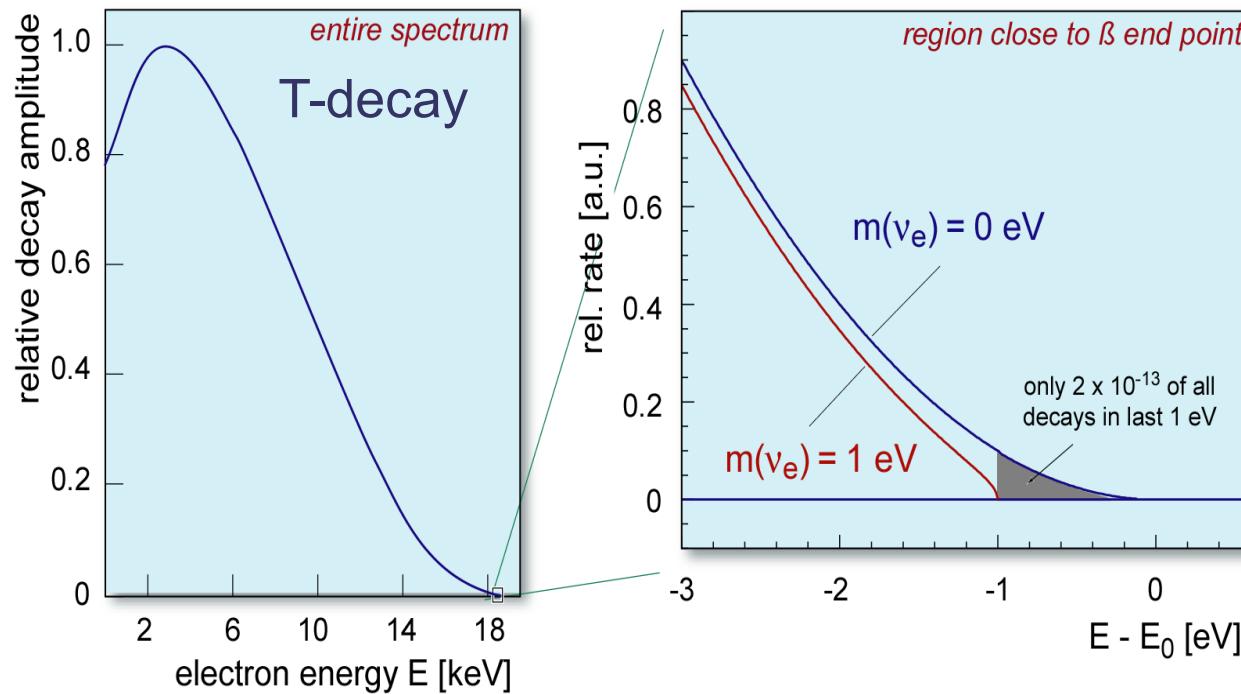
Kinematic determination of $m(\nu_e)$

$$\frac{d\Gamma}{dE} = C p(E+m_e)(E_0-E)\sqrt{(E_0-E)^2 - m_\nu^2} F(Z+1, E) \Theta(E_0-E-m_\nu) S(E)$$

$$C = \frac{G_F^2}{2\pi^3} \cos^2 \theta_C |M|^2$$

(modified by final state distribution, recoil corrections, radiative corrections, ...)

$$m_\nu^2 = \sum_{i=1}^3 |U_{ei}|^2 m_i^2$$



Suitable Isotopes:

Tritium

- $E_0 = 18.6 \text{ keV}$, $T_{1/2} = 12.3 \text{ a}$
- $S(E) = 1$ (super-allowed)

Rhenium

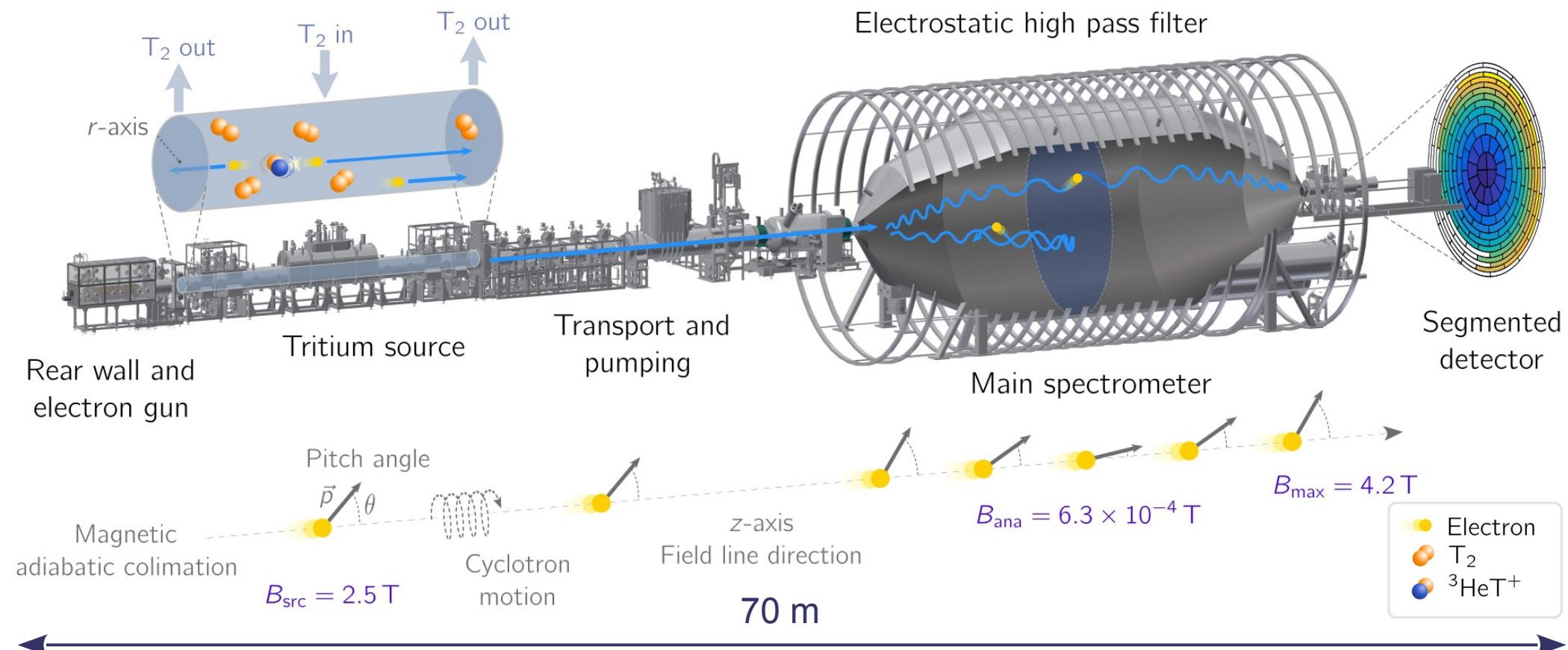
- $E_0 = 2.47 \text{ keV}$, $T_{1/2} = 43.2 \text{ Gy}$

alternative approach:

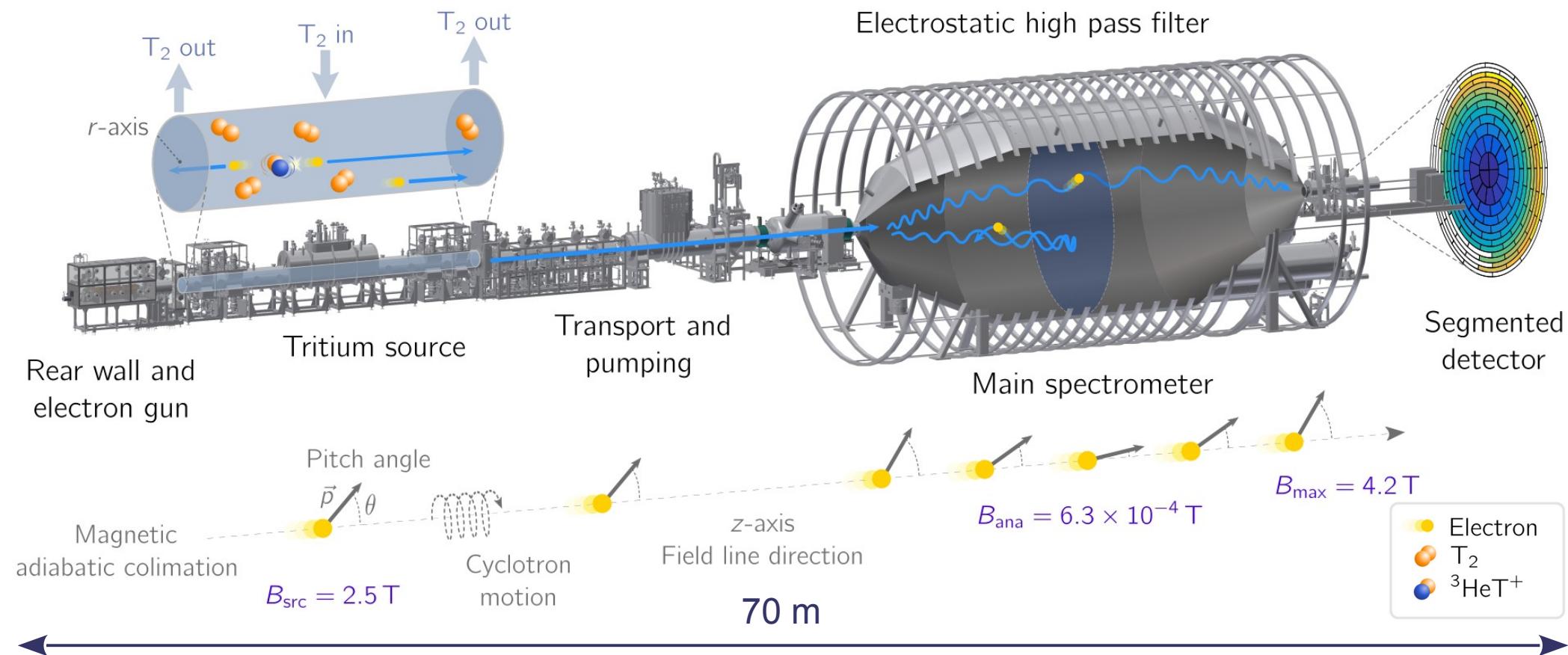
Holmium (EC decay)

- $Q_{EC} \approx 2.5 \text{ keV}$, $T_{1/2} = 4570 \text{ y}$

KATRIN experiment at KIT



KATRIN experiment at KIT



KATRIN design goals:
collect **1000 days** of data

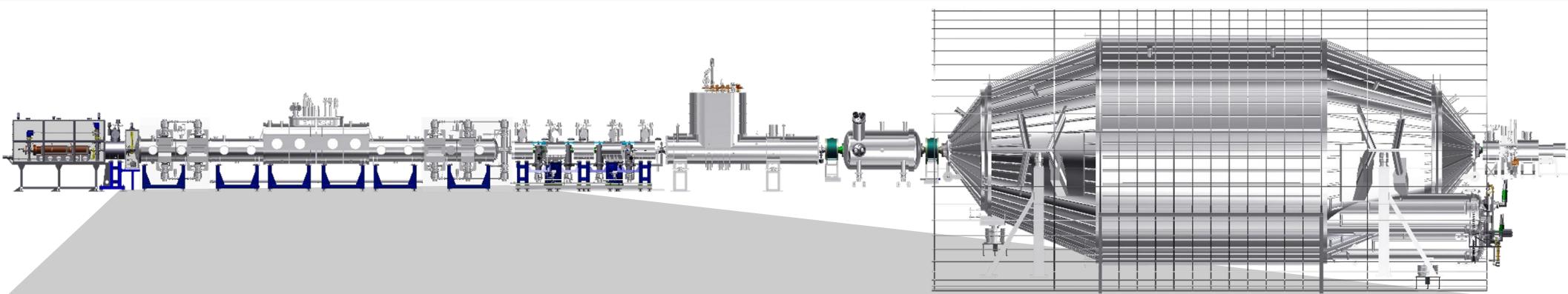
statistical uncertainty
systematic uncertainty

$$\sigma_{\text{stat}} \approx 0.018 \text{ eV}^2$$

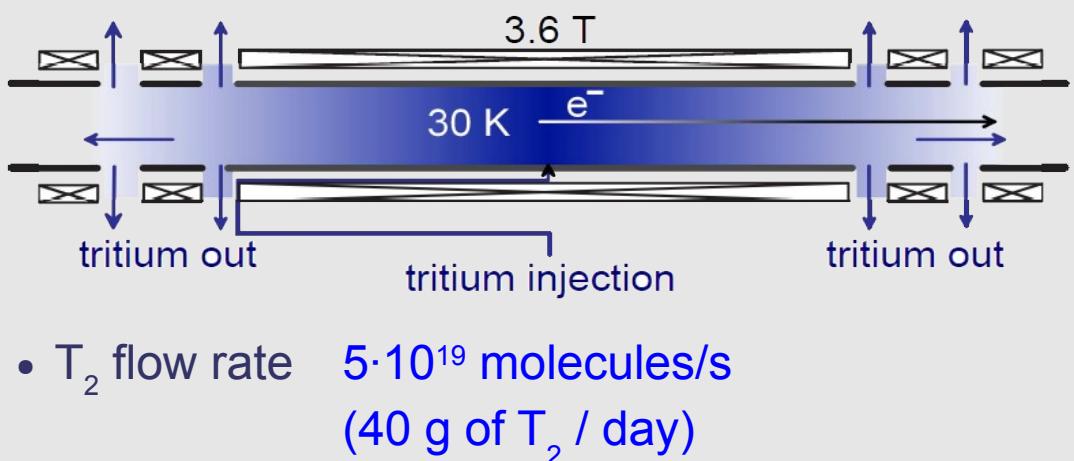
$$\sigma_{\text{sys,tot}} \approx 0.017 \text{ eV}^2$$

→ sensitivity for upper limit: **0.2 eV** (90% C.L.)
 $m_\nu = 0.35 \text{ eV}$ observable with 5σ

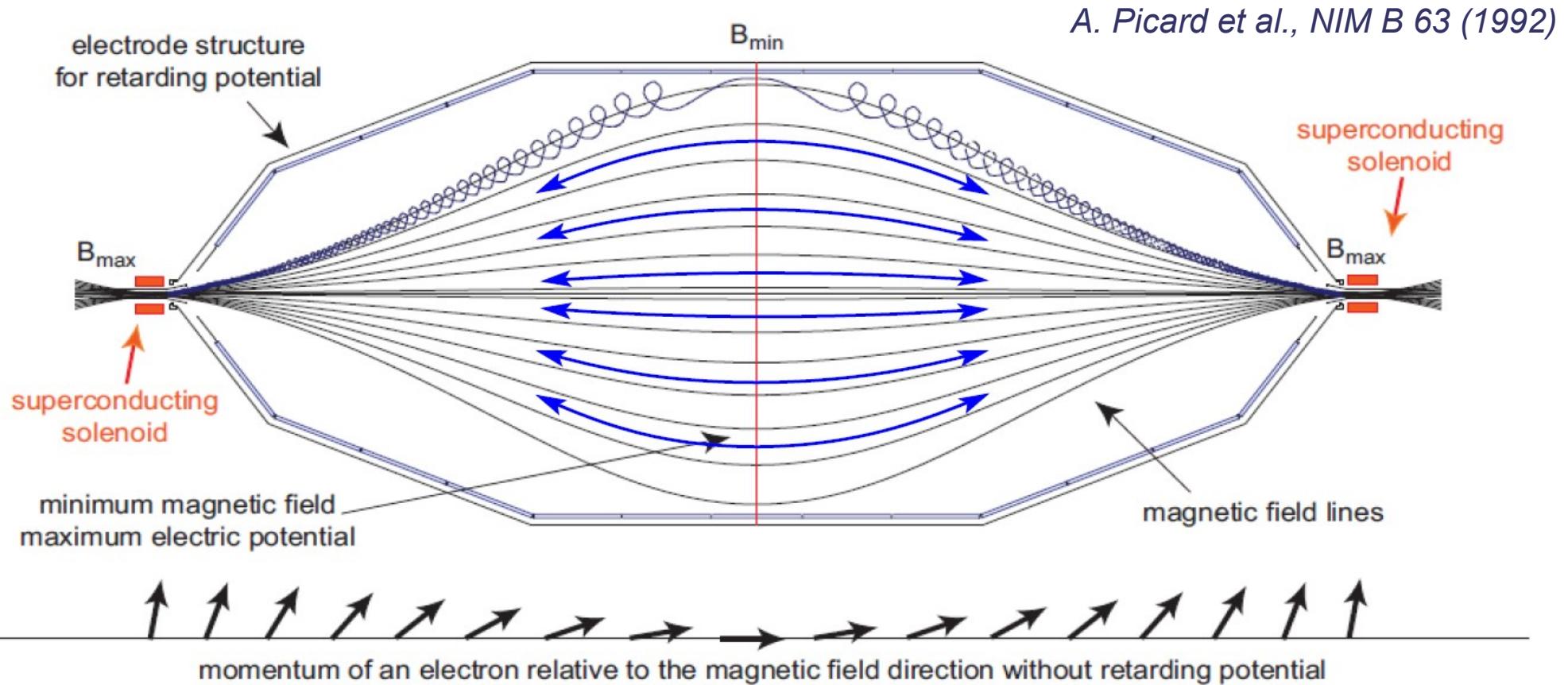
Windowless Gaseous Tritium Source



- beam tube $\varnothing = 9 \text{ cm}$, $L = 10 \text{ m}$
- guiding field 3.6 T
- temperature $T = 30 \text{ K} / 80 \text{ K} \pm 30 \text{ mK}$,
- T_2 purity $95\% \pm 0.1 \%$
- column density $5 \cdot 10^{17} \text{ T}_2/\text{cm}^2$
- luminosity $1.7 \cdot 10^{11} \text{ Bq}$

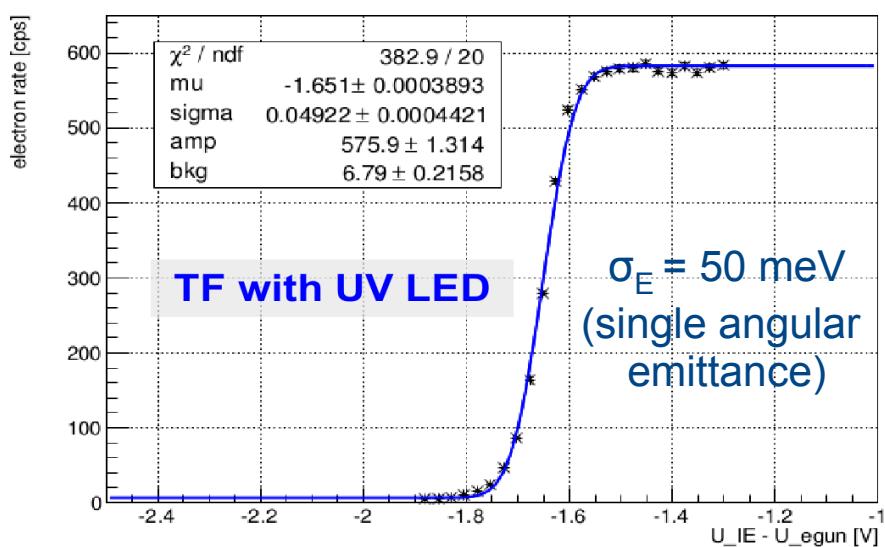
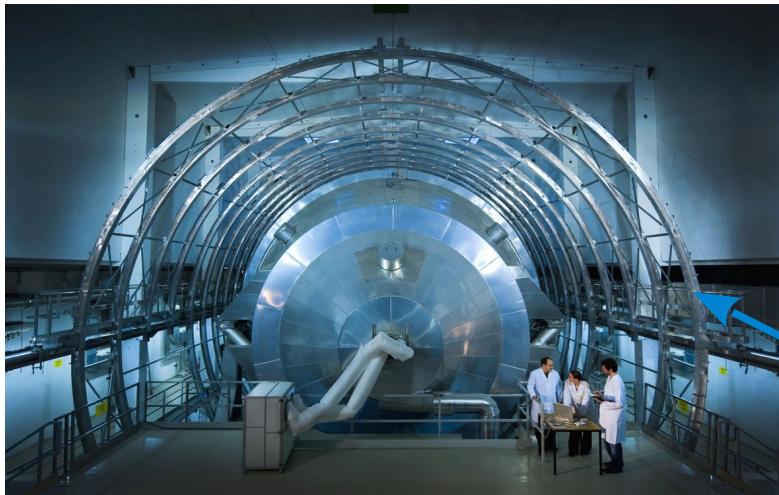
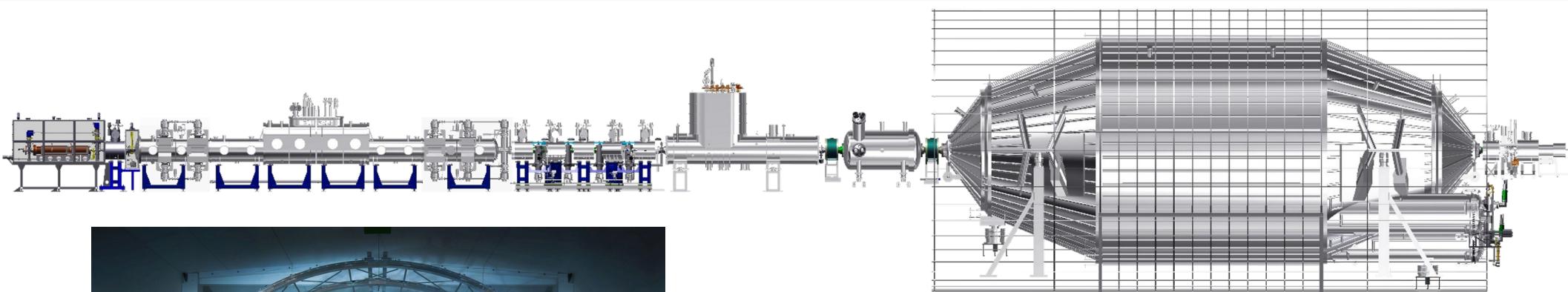


Magnetic Adiabatic Collimation with Electrostatic Filter



- adiabatic transport $\rightarrow \mu = E_{\perp} / B = \text{const.}$
- B drops by $2 \cdot 10^4$ from solenoid to analyzing plane $\rightarrow E_{\perp} \rightarrow E_{||}$
- only electrons with $E_{||} > eU_0$ can pass the retardation potential
- Energy resolution $\Delta E = E_{\perp, \text{max, start}} \cdot B_{\min} / B_{\max} < 1 \text{ eV}$

Main-Spectrometer

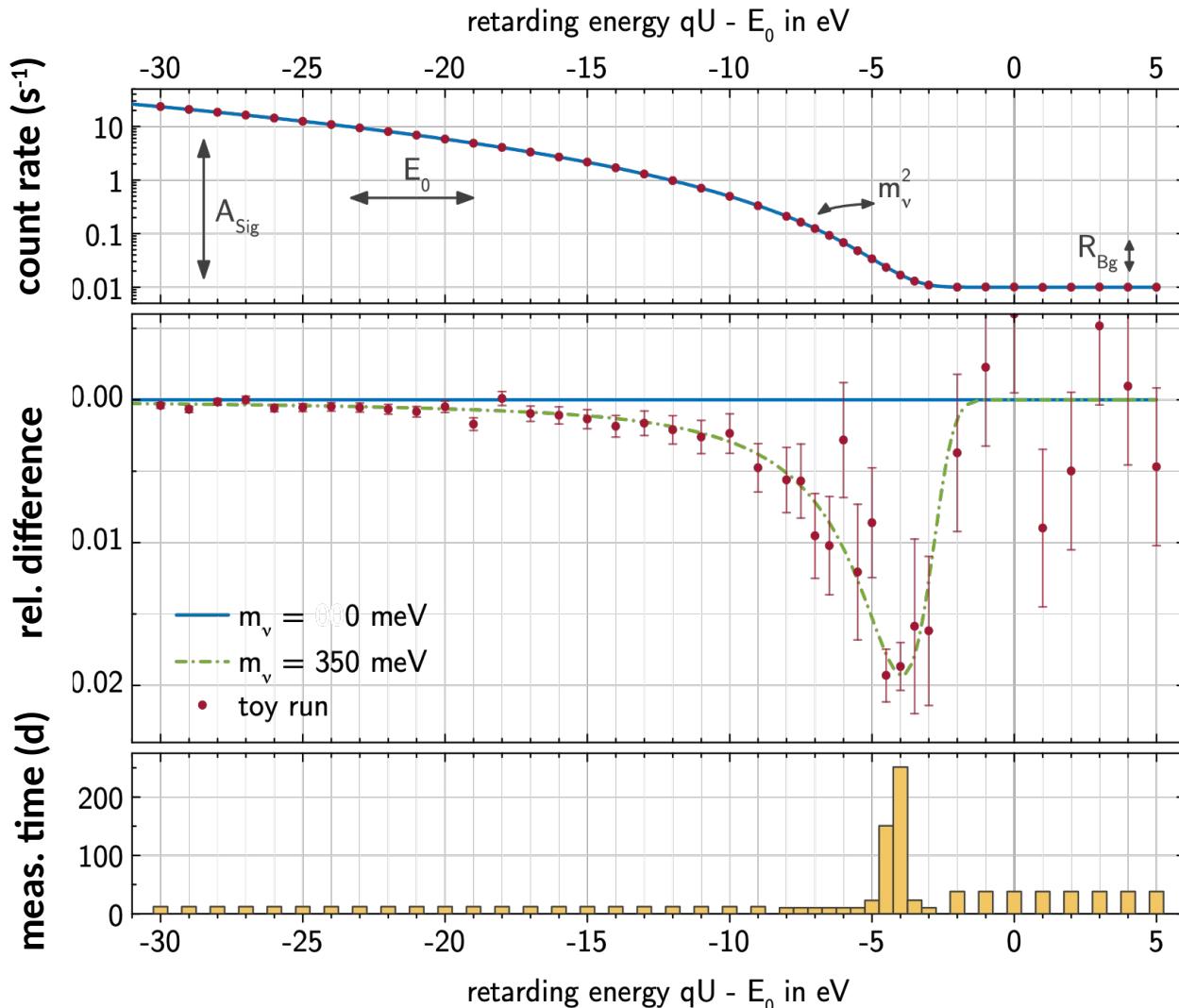


- 18.6 kV retardation voltage, $\sigma < 60 \text{ meV}$
- 0.93 eV resolution
- pressure $< 10^{-11} \text{ mbar}$
- Air coils for earth magnetic field compensation
- Double layer wire electrode for background reduction and field shaping



Measurement principle

- Direct **shape** measurement of **integrated β spectrum**



Four fit parameters:

spectrum
ampl. A_{sig}

spectrum
endpoint E_0

background
rate R_{bg}

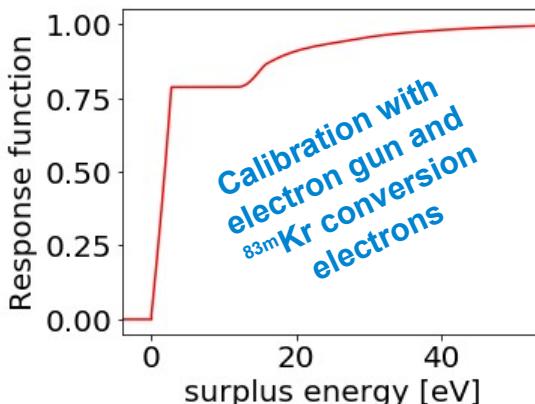
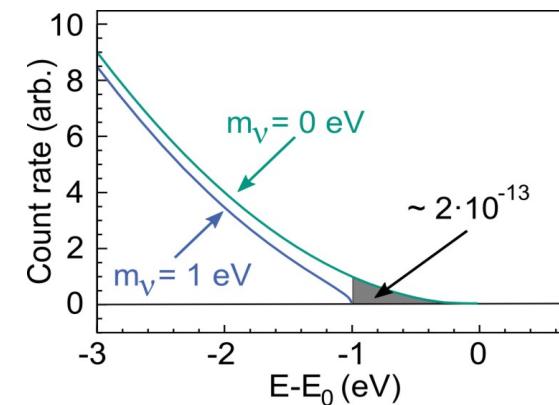
squared
mass m_ν^2

~10⁻⁸ of all β -decays in scan
region ~40 eV below endpoint

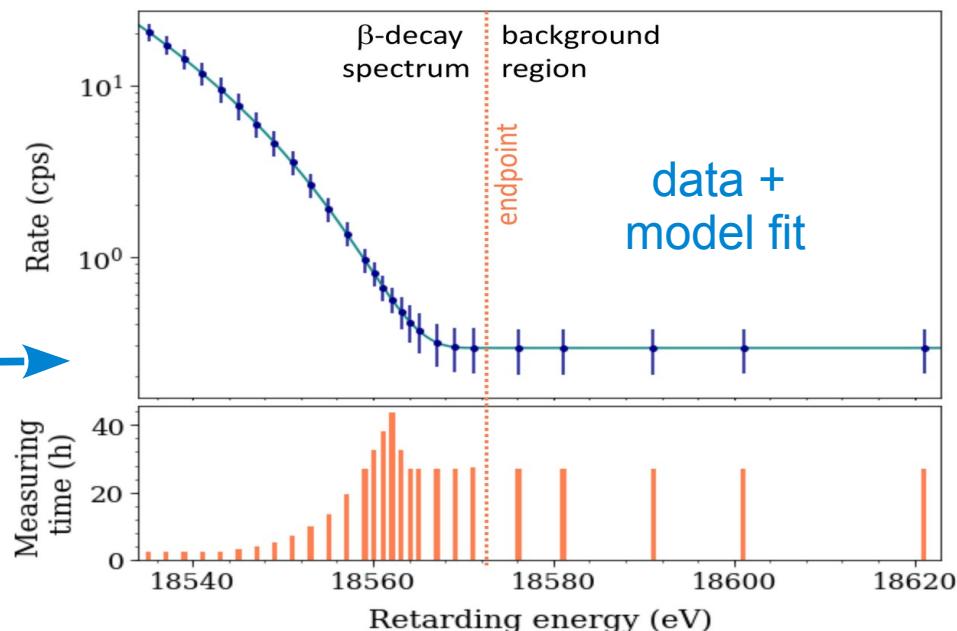
Eur. Phys. J. C 79 (2019) 204

Model of the experimental spectrum

Beta spectrum: $R_\beta(E, m^2(\nu_e))$



KATRIN β -scan



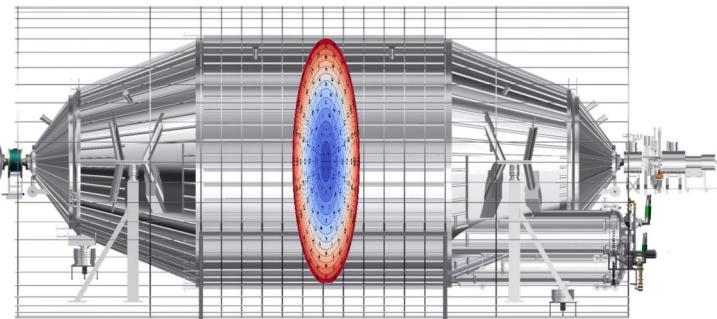
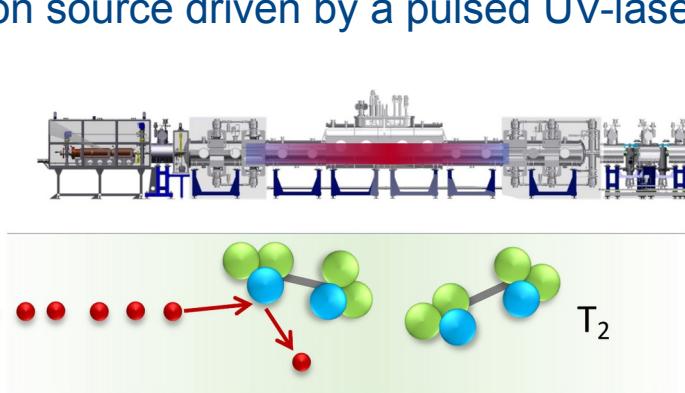
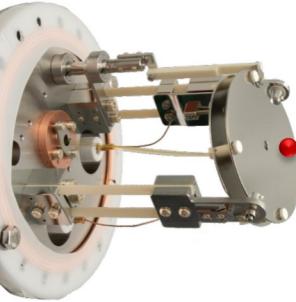
$$R(qU) = A_s \cdot N_T \int_{qU}^{E_0} R_\beta(E, m^2(\nu_e)) \cdot f(E - qU) dE + R_{bg}$$

Experimental response: $f(E - qU)$

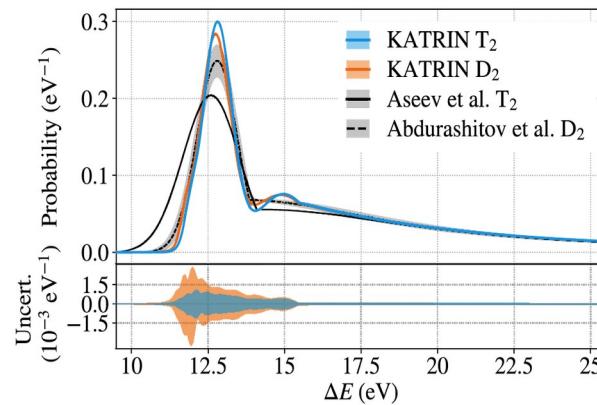
*PRL 123 (2019) 221802, EPJ C 79 (2019) 204
+ detailed analysis PRD 104 (2021) 012005
+ energy loss measurement EPJ C 81 (2021) 579*

Response function measurement

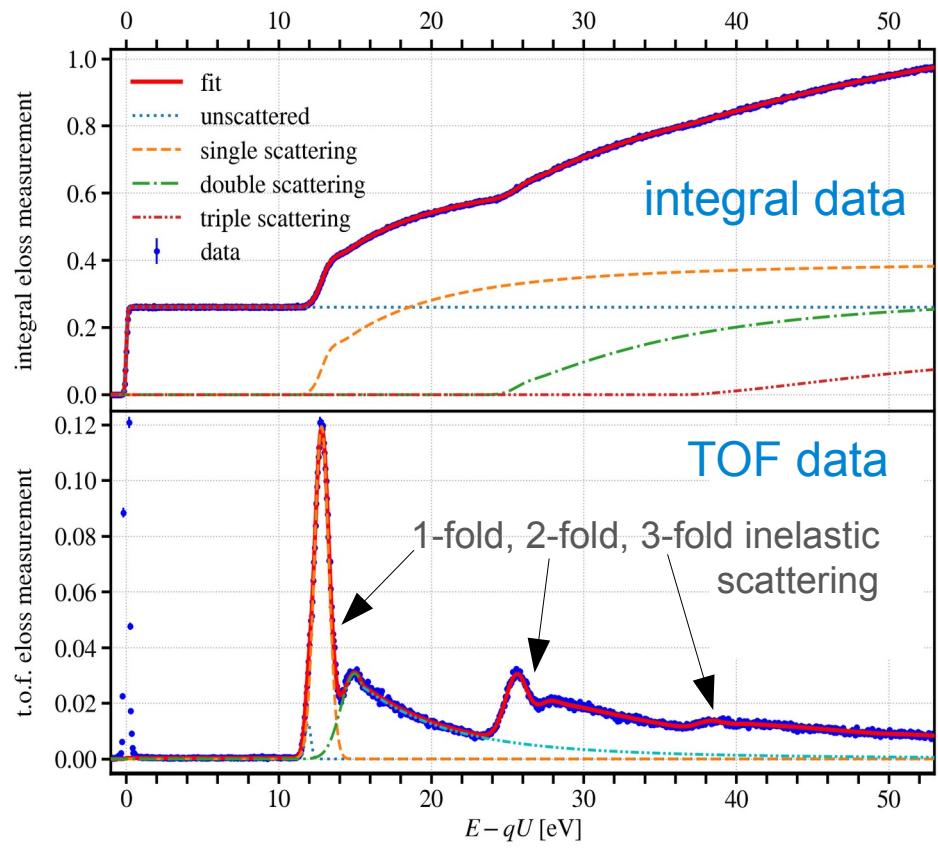
- Determination of the experimental response function using a mono-energetic, angular selective photo-electron source driven by a pulsed UV-laser



- Measurement of integral and differential (using TOF method) spectra at different column densities
→ Extraction of spectrometer transmission function and energy losses due to scattering (*EPJ C 81 (2021) 579*)
- Additional measurements with a gaseous ^{83m}Kr source to investigate plasma potential (*J. Phys. G 47 (2020) 065002*)



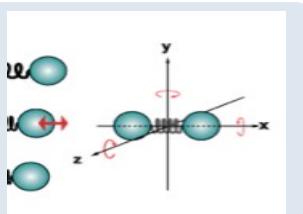
Extract differential energy loss function



Main systematic uncertainties

Molecular final states

- Quantum-chemical computations



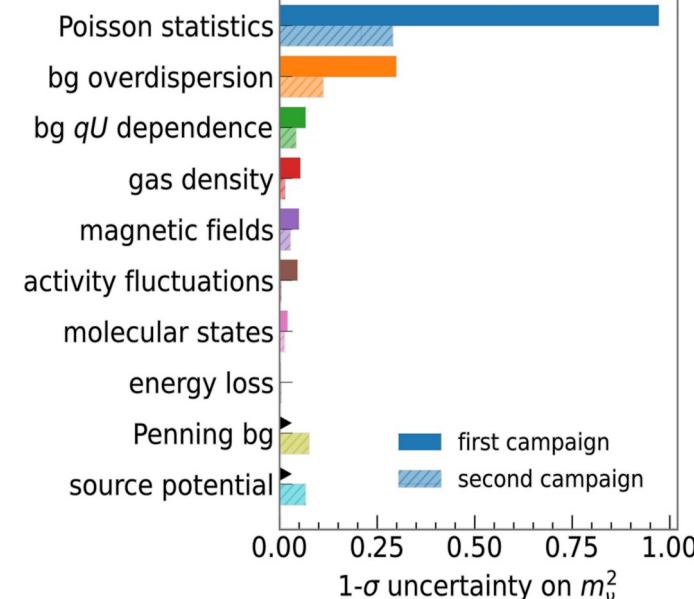
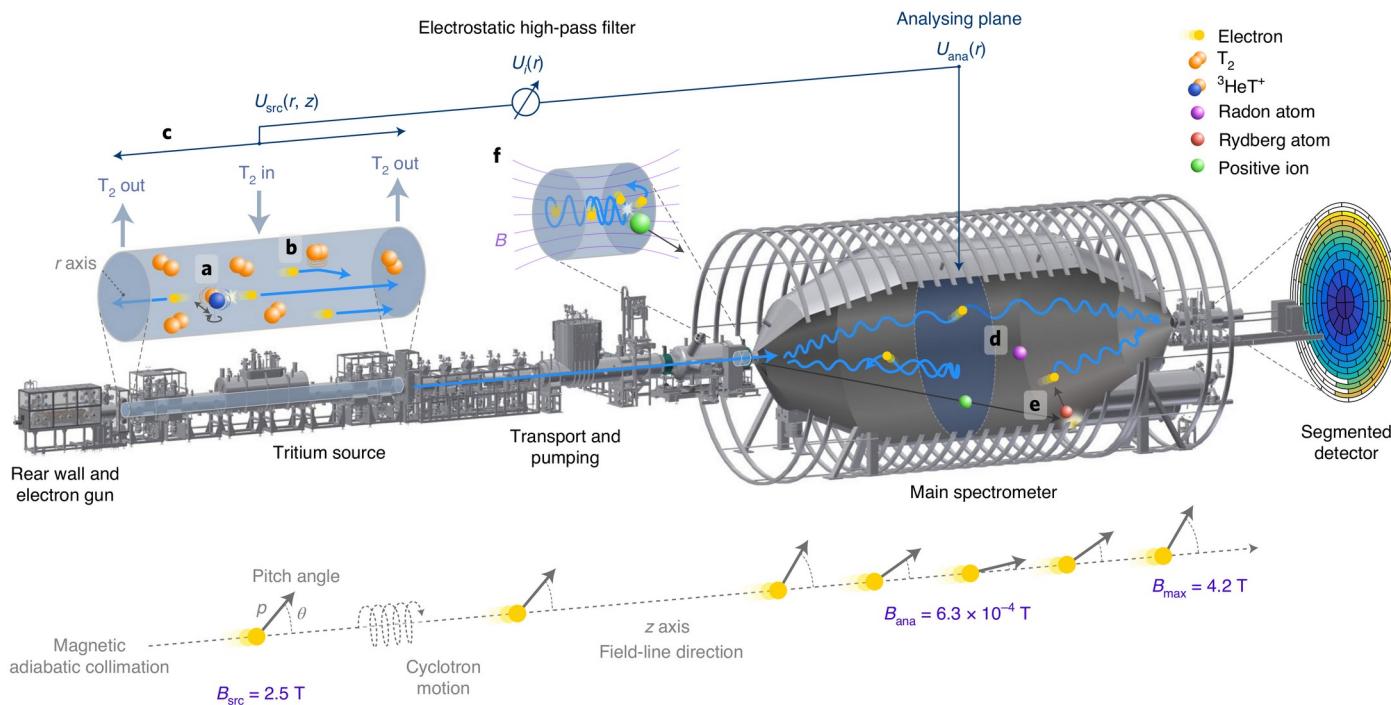
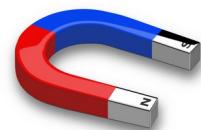
Source electric potential

- plasma properties
- surface conditions

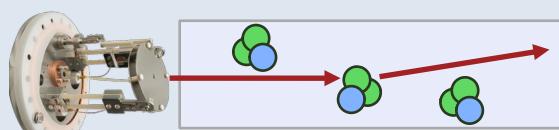


Magnetic fields

- source
- spectrometer
- detector



Energy loss by scattering

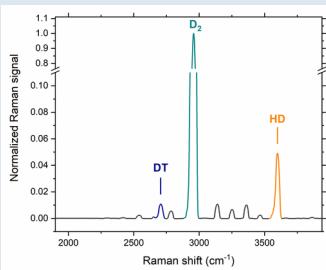


EPJ C 79 (2019) 204
EPJ C 81 (2021) 579

Activity fluctuations

- column density
- Tritium purity (T₂, DT, HT)

Sensors 20 (2020) 4827

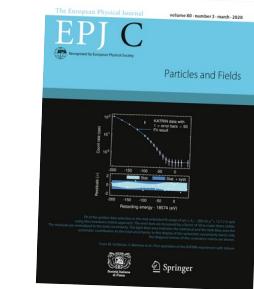


Background

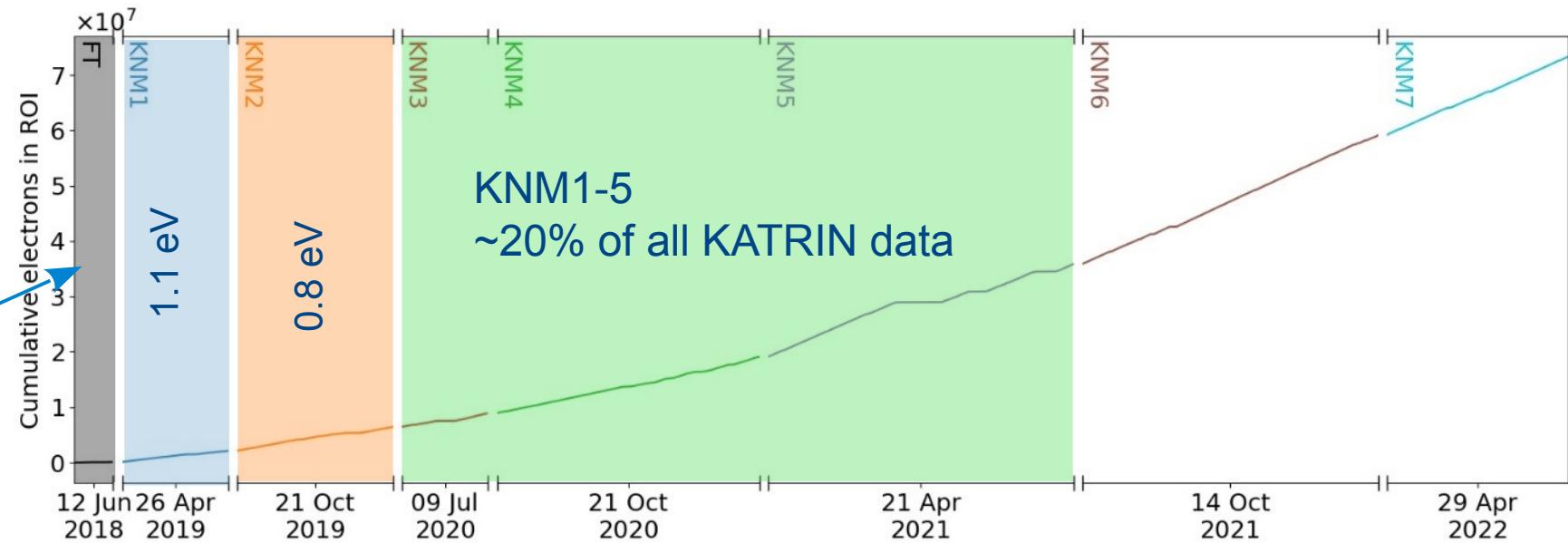
- Volume dependent “Rydberg” background
- Voltage dependent background
- Background time structure due to trapped electrons

arXiv:2011.05107v1 (2020)

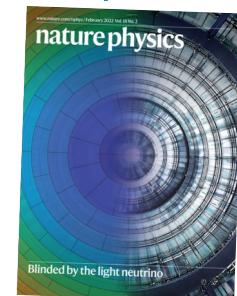
KATRIN data taking



EPJC 80 (2020) 264



PRL 123 (2019) 221802



Nat. Phys. 18 (2022) 160



Light sterile neutrinos



relic neutrinos



PRL 129 (2022) 011806

- analysis of runs KNM3 to KNM5 ongoing, expected statistical sensitivity ~ 0.5 eV (90% CL)
- data unblinding will happen soon

Extracted neutrino mass limits

1st campaign (spring 2019)

- total statistics: 2 million events
- best fit result: $m_\nu^2 = -1.0^{+0.9}_{-1.1} \text{ eV}^2$
- mass limit: $m_\nu < 1.1 \text{ eV} (90\% CL)$

2nd campaign (autumn 2019)

- total statistics: 4.3 million events
- best fit result: $m_\nu^2 = 0.26^{+0.34}_{-0.34} \text{ eV}^2$
- mass limit: $m_\nu < 0.9 \text{ eV} (90\% CL)$

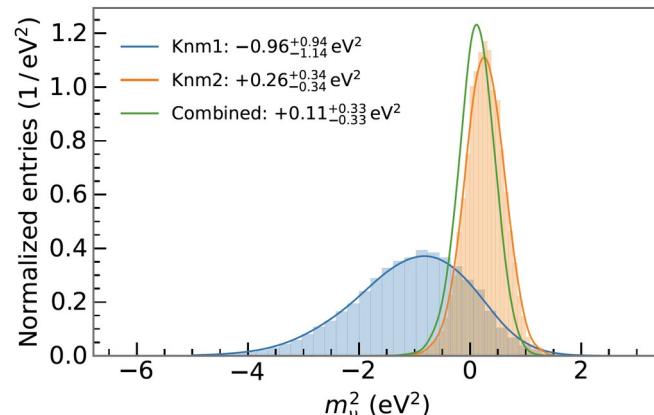
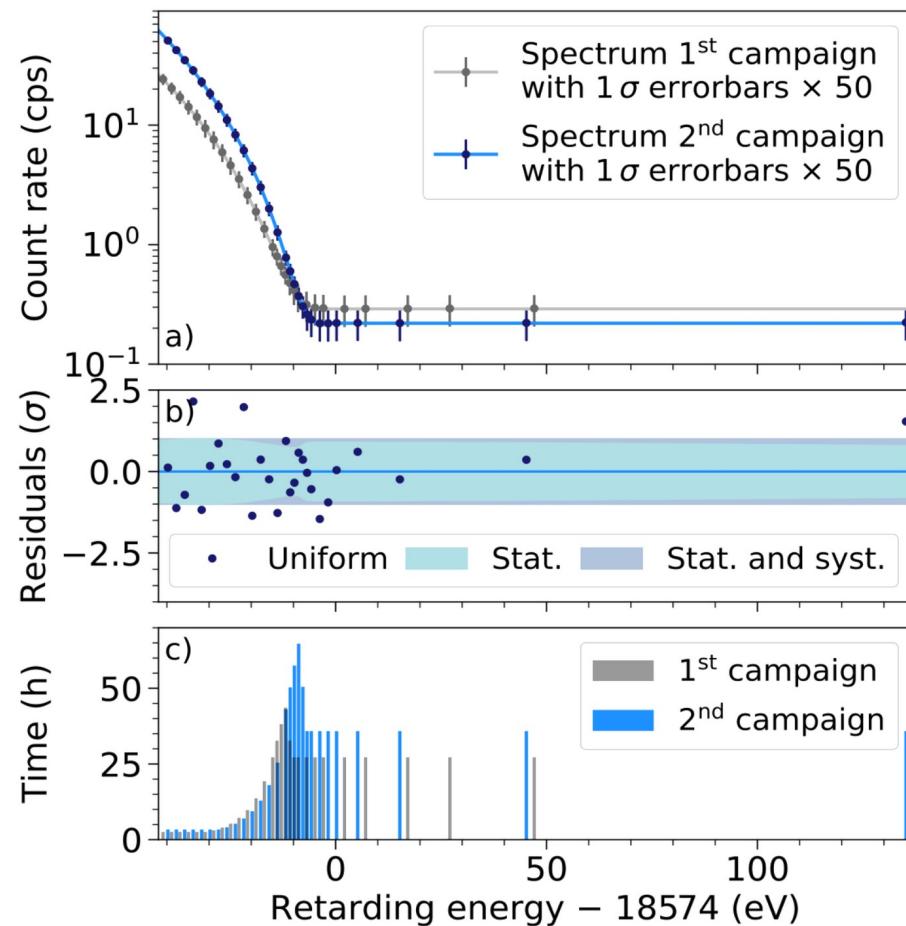
Combine 1st and 2nd campaign:

- mass limit: $m_\nu < 0.8 \text{ eV} (90\% CL)$

Cross-check: endpoint energy

$$E_0 = 18573.69 \pm 0.03 \text{ eV} \rightarrow \text{Q-value: } 18575.2 \pm 0.5 \text{ eV}$$

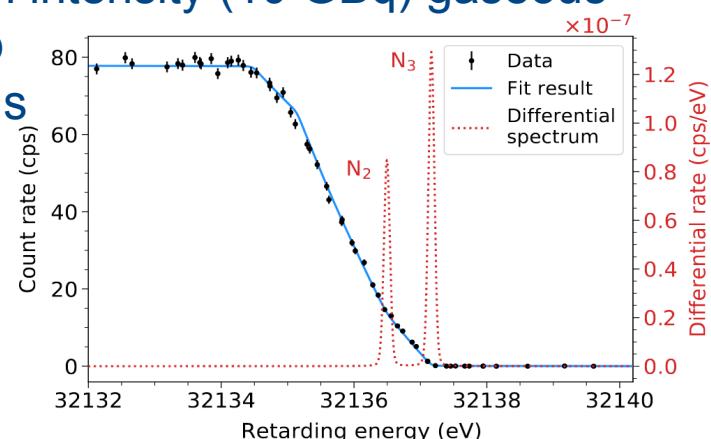
→ good agreement with Penning trap experiments:
 $Q = 18575.72 \pm 0.07 \text{ eV}$ (*PRL 114 (2015) 013003*)



KATRIN improvements

- New spectrometer field configuration (shifted analysis plane) reduced background by factor 2 and removes non-poissonian backgrounds
- Calibration with high intensity (10 GBq) gaseous ^{83m}Kr source to map out source potentials

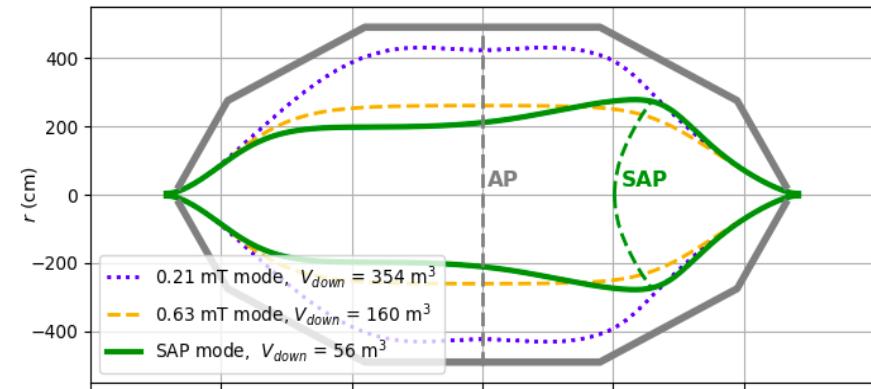
→ co-circulation with T_2 for simultaneous monitoring



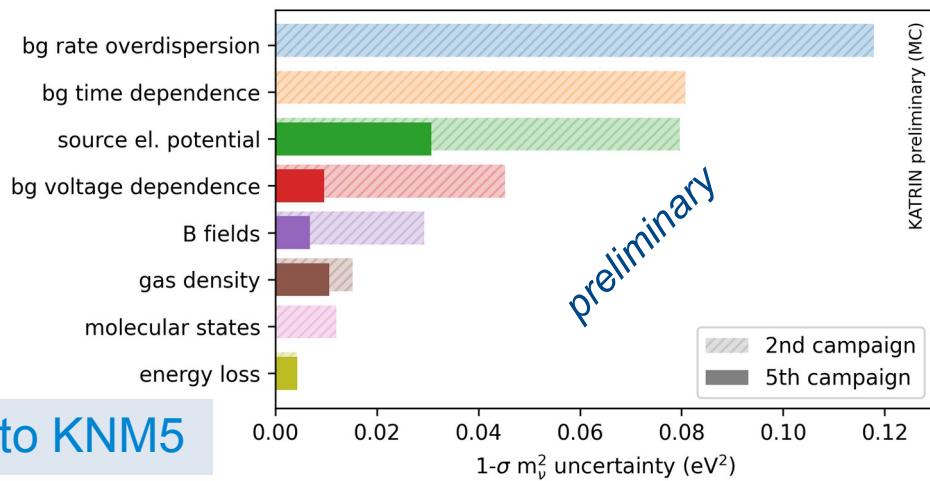
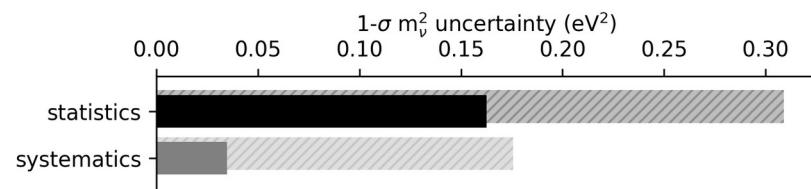
J. Phys. G 47 (2020) 065002

- Removal of inter-spectrometer Penning trap to eliminate background time dependence
- New calculations of molecular final states in progress
- Many additional accuracy improvements ...

Expected uncertainty budget for KNM3 to KNM5

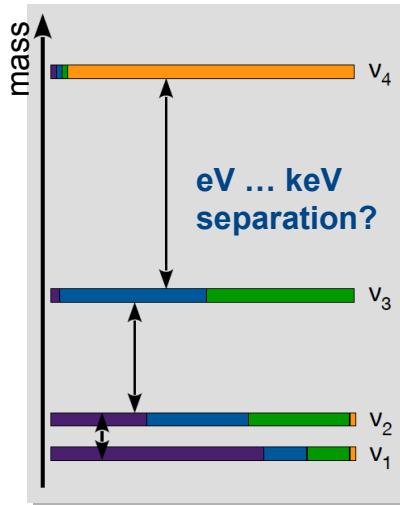


Eur. Phys. J. C 82 (2022) 258

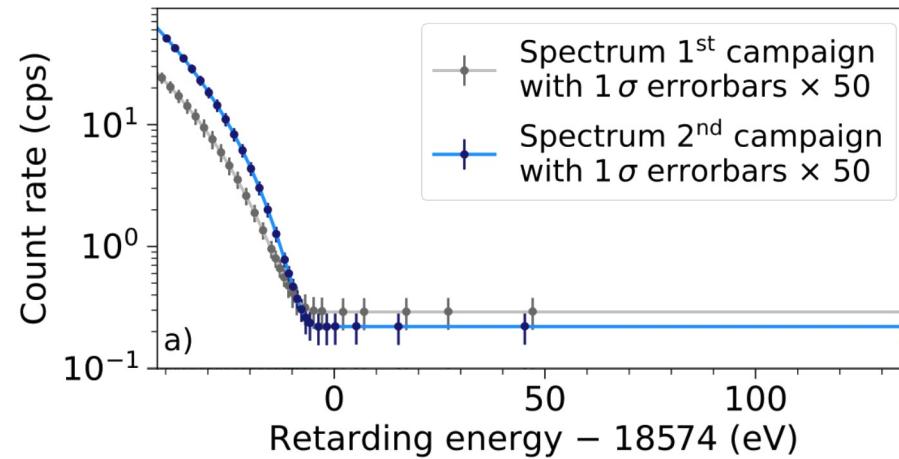


Beyond neutrino mass searches

Is there a fourth (sterile) neutrino?



Neutrino mixing: “Kink” in normal β -spectrum (eV scale) or deep β -spectrum (keV scale)

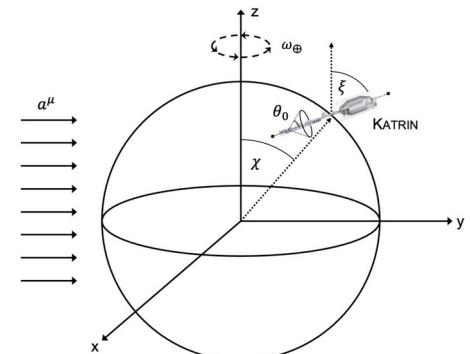


High statistics, high precision β spectrum

Constrain local over-density of cosmic relic neutrinos (peak search)

Search for exotic weak interactions (spectrum shape)

Search for Lorentz invariance violation (sidereal modulation)



Phys. Rev. D 105 (2022) 072004
arXiv:2207.06337v1 (2022)

PRL 129 (2022) 011806

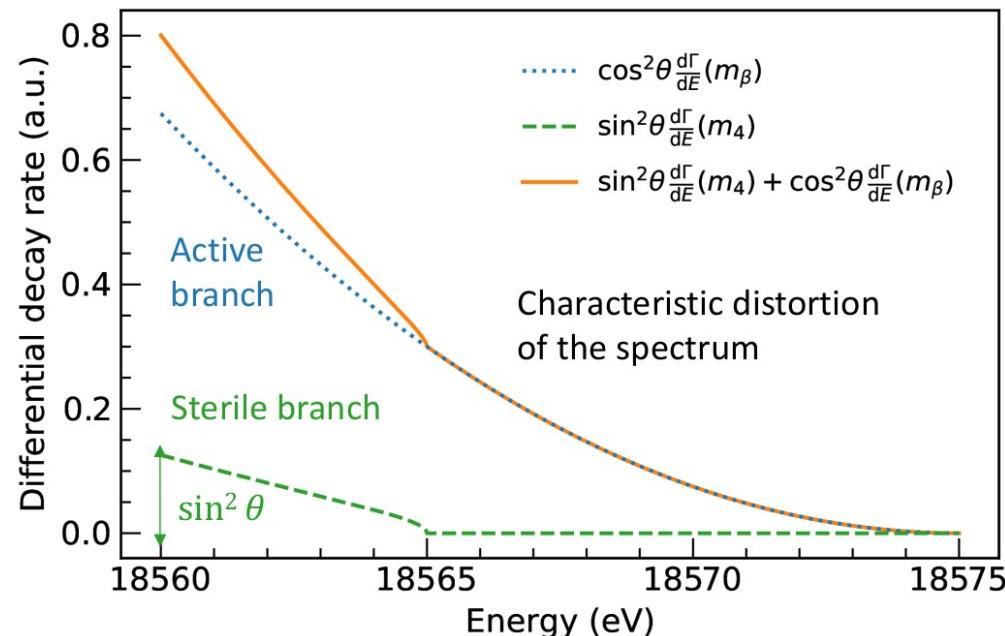
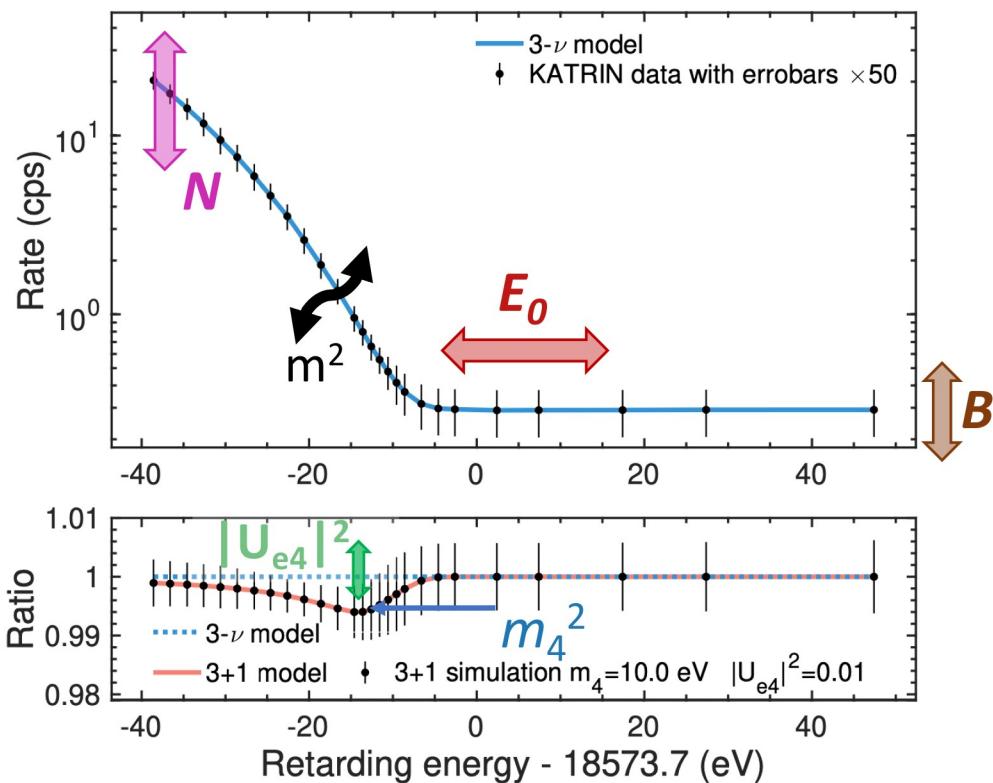
arXiv:2207.06326v1 (2022)

(Light) sterile neutrino signature

- 3+1 sterile neutrino model
- Same data-set as for the neutrino mass
- Grid search in m_4 , $|U_{e4}|^2$ plane

$$\frac{d\Gamma}{dE} = \left(1 - |U_{e4}|^2\right) \frac{d\Gamma}{dE}(m_\beta^2) + |U_{e4}|^2 \frac{d\Gamma}{dE}(m_4^2)$$

light neutrino **heavy neutrino**



6 Fit Parameters:

- | | |
|--------------|------------------------------------|
| m^2 | - neutrino mass (free/constrained) |
| E_0 | - endpoint |
| N | - amplitude |
| B | - background rate |
| m_4^2 | - 4 th neutrino mass |
| $ U_{e4} ^2$ | - 4 th neutrino mixing |

Limits on light sterile neutrinos

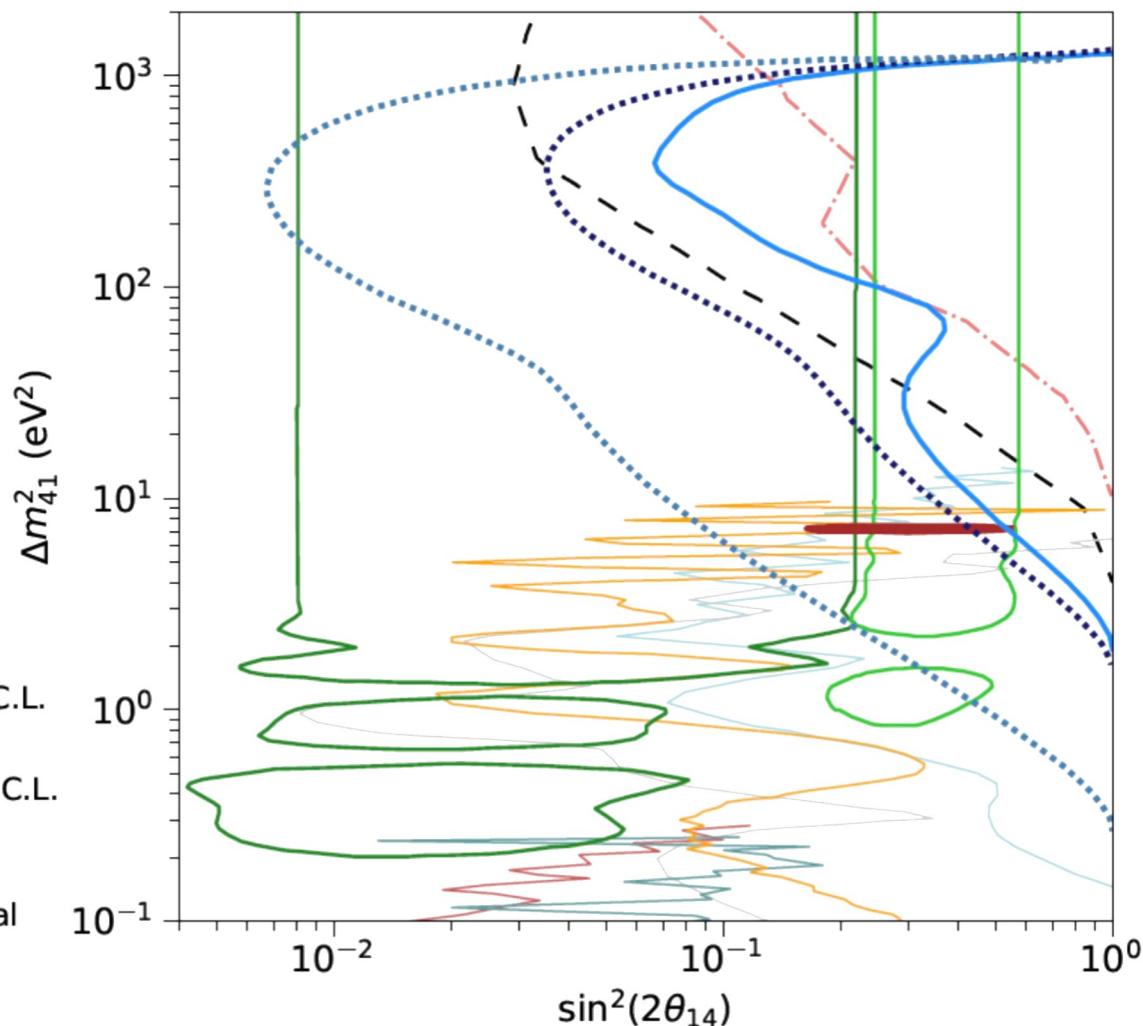
Current dataset

- KATRIN starts to probe very interesting parameter space, complementary to oscillation searches
- approaching the BEST allowed regions with $\Delta m_{41}^2 > 10 \text{ eV}^2$

Final dataset

- Probing large portion of RAA, BEST and Neutrino-4 allowed regions
- comparable sensitivity to neutrinoless double β -decay

- | | |
|--------------------------|--|
| — Mainz 95 % C.L. | — RAA 95 % C.L. |
| — Troitsk 95 % C.L. | — BEST + GA 95.45 % C.L. |
| — Prospect 95 % C.L. | — Neutrino-4 2 σ |
| — DANSS 95 % C.L. | — KATRIN (KSN1) 95 % C.L. |
| — Daya Bay 90 % C.L. | — KATRIN (KSN1+2) sensitivity 95 % C.L. |
| — Double Chooz 95 % C.L. | — Projected KATRIN final sensitivity 95 % C.L. |
| — Stereo 95 % C.L. | |



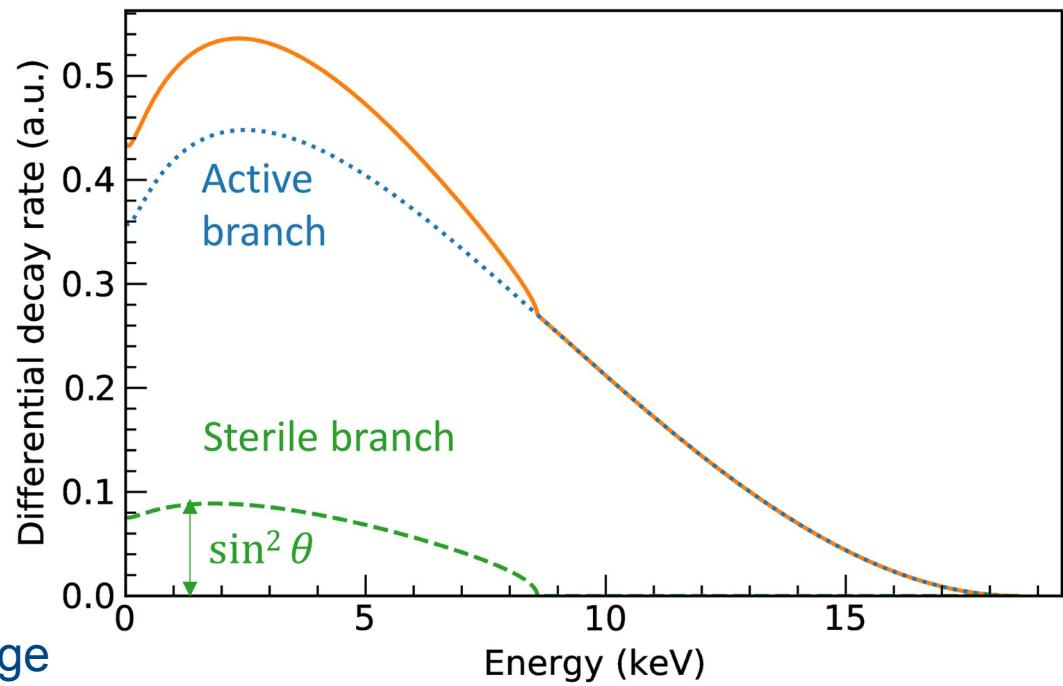
PR D 105 (2022) 072004

$$\sin^2(2\theta_{ee}) = 4|U_{e4}|^2(1 - |U_{e4}|^2)$$

Future plans: keV scale sterile neutrinos

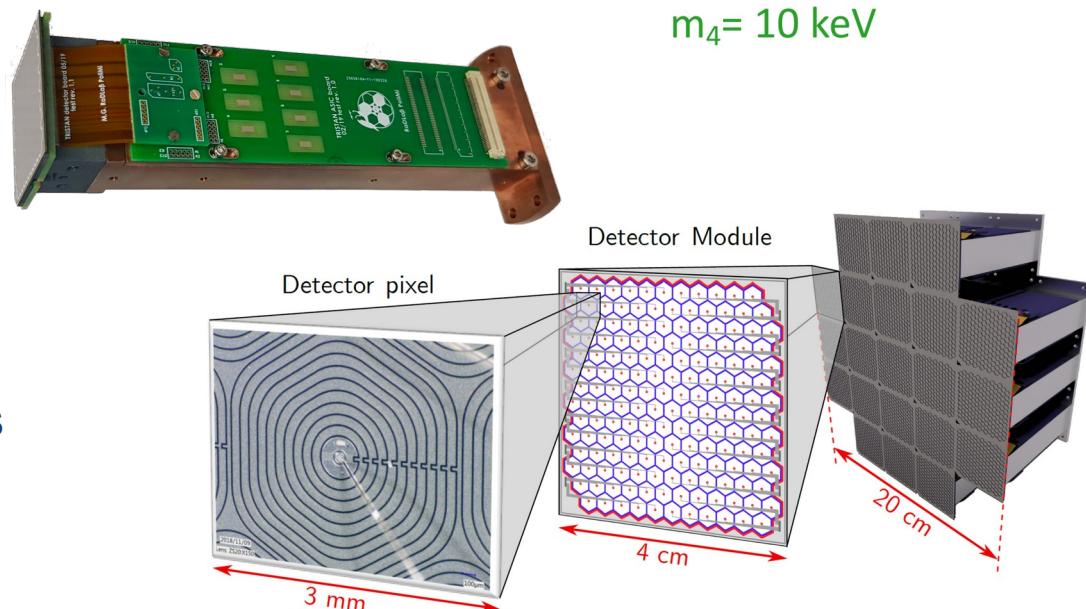
keV-scale sterile neutrino search:

- sterile neutrinos on keV-scale are a viable candidate for dark matter in the universe (WDM)
- First scans deep into the β -spectrum during FT campaign at 0.5% c.d.
[arXiv:2207.06337v1 \(2022\)](https://arxiv.org/abs/2207.06337v1)
- high-sensitivity search requires new high-rate detector system (TRISTAN) to handle huge electron rates from WGTS over large spectral range



TRISTAN project in KATRIN:

- novel multi-pixel Silicon Drift Detector array
- large count rates
- excellent energy resolution
- prototypes installed as monitoring devices @ KATRIN
- target sensitivity: $\sin^2\theta < 10^{-6}$



- Studies of β -decay kinematics offer a model-independent way to determine the neutrino mass, complementary to cosmology and $0\nu\beta\beta$ searches
- The KATRIN experiment has finalized the analysis of the first two science runs and published the first sub-eV neutrino mass limit with $m_\nu < 0.8 \text{ eV}$
- Several improvements allowed to strongly reduce experimental background and systematic uncertainties
- Analysis of KNM3 to KNM5 science runs ongoing, unblinding will happen soon
- KATRIN has the capability to study several physics topics beyond neutrino mass:
 - eV-scale sterile neutrinos (first upper limits published)
 - keV scale sterile neutrinos (future project with new focal plane detector TRISTAN)
 - upper limit on local relic neutrino overdensity
 - investigations of Lorentz invariance
 - search for exotic weak interactions