# Neutrinos - second lecture

importance of "*neutrinoless double beta decay*" process; global numbers; significance of neutrinos masses in particle physics; introduction to Majorana's formalism; extension of the standard model; implications

Francesco VISSANI - Laboratori Nazionali del Gran Sasso - July 21, 2022

### a crucial process for particle physics

the importance of the search for process

### $(A, Z) \rightarrow (A, Z + 2) + 2 e^{-1}$

receives consistent support from *several* argumentations, of different character

**OBSERVATIONAL:** the need to verify the global symmetries of SM

**PHENOMENOLOGICAL:** implications of evidences of physics beyond the SM

**AXIOMATIC:** the hypothesis posed in 1937 by Majorana for neutrino mass

- **EVOLUTIONARY:** the development of gauge theories and search for their extensions
- We expose these argumentations today, highlighting the many connections with *neutrino* physics

# history of global symmetries in particle physics

**Jean Perrin** 

Les hypothèses moléculaires

La Revue Scientifique 13 Avril 1901

PHYSIQUE. – *Les hypothèses moléculaires*. **Jean Perrin**, La Revue Scientifique

- 4e série Tome XV:N°15 13 Avril 1901
- Conférence faite aux étudiants et aux amis de l'université de Paris, le 16 février 1901. par M. Jean Perrin, chargé du cours de Chimie à la Sorbonne.

In the section titled 'Division de l'atome en corpuscles' we read statements such as:

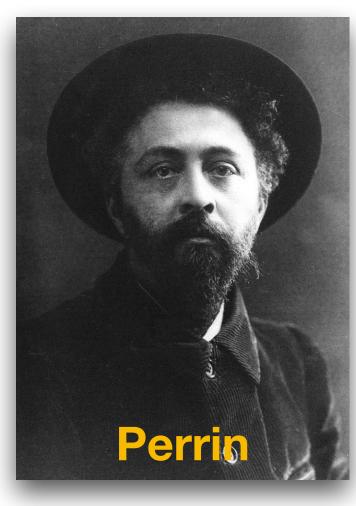
Every atom will consist, on the one hand, of various masses strongly charged with positive electricity, a kind of **positive suns** (soleils) whose charge will be much greater than that of a corpuscle, and on the other hand, of a multitude of corpuscles, a kind of **small negative planets** (planètes), the whole of whose masses gravitate under the action of electrical forces, and the total negative charge exactly equals the total positive charge, so that the atom is electrically neutral.

And also, equally impressive and eloquent,

- Negative corpuscles are all equal to each other, whatever the chemical nature of the atom [...]
- the mass [...] is only the thousandth part of the hydrogen atom.



**Télécharger** 



*ver the chemical nature of the atom […] on atom.* 



### from radioactivity to the nucleus

### discovery of radioactivity

$$lpha, eta, \gamma$$
 rays

 $\checkmark$  Th $\rightarrow$ Ra transmutation [Soddy+Rutherford 1901]

<sup>×</sup>energy mass equivalence

discovery of nucleus [Geiger-Marsden+Rutherford 1911]

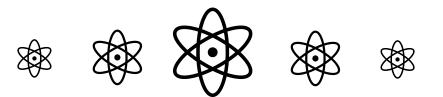


[Rutherford 1899, Villard 1900]

[Einstein 1905]



### electrons. A first understanding of light quanta begins to emerge.



At this stage of the discussion, matter is thought to consist of nuclei and

### electrons. A first understanding of light quanta begins to emerge.



NOW: if - based on  $E = mc^2$  - someone had dared to raise the question, whether such a reaction could ever happen

the answer could have been that "the mass-number is conserved," which is a corollary of *radioactive displacement law* (Russell, Fajans, Soddy)

At this stage of the discussion, matter is thought to consist of nuclei and



 $\alpha \rightarrow p + p + \gamma \qquad \alpha \rightarrow p + p + p + \beta^{-}$ 

### which are the components of the nucleus? (first attempts to devise a model)



- van den Broek (1911) suggested that the nuclei consist of  $\alpha$  and  $\beta$  particles
- Harkins (1915) instead hypothesised that the basic components were H-nuclei and  $\beta$  particles
- both are elegant models, that explain the main facts and <u>exclude previous processes</u> from occurring



### which are the components of the nucleus? (first attempts to devise a model)



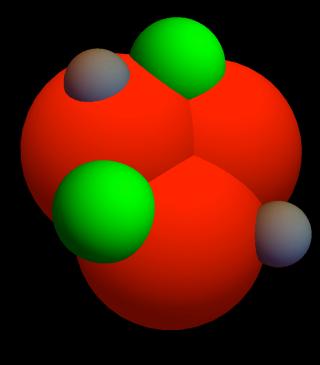
# but these models contradict $\beta$ rays observations!

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- both are elegant models, that explain the main facts and exclude previous processes from occurring



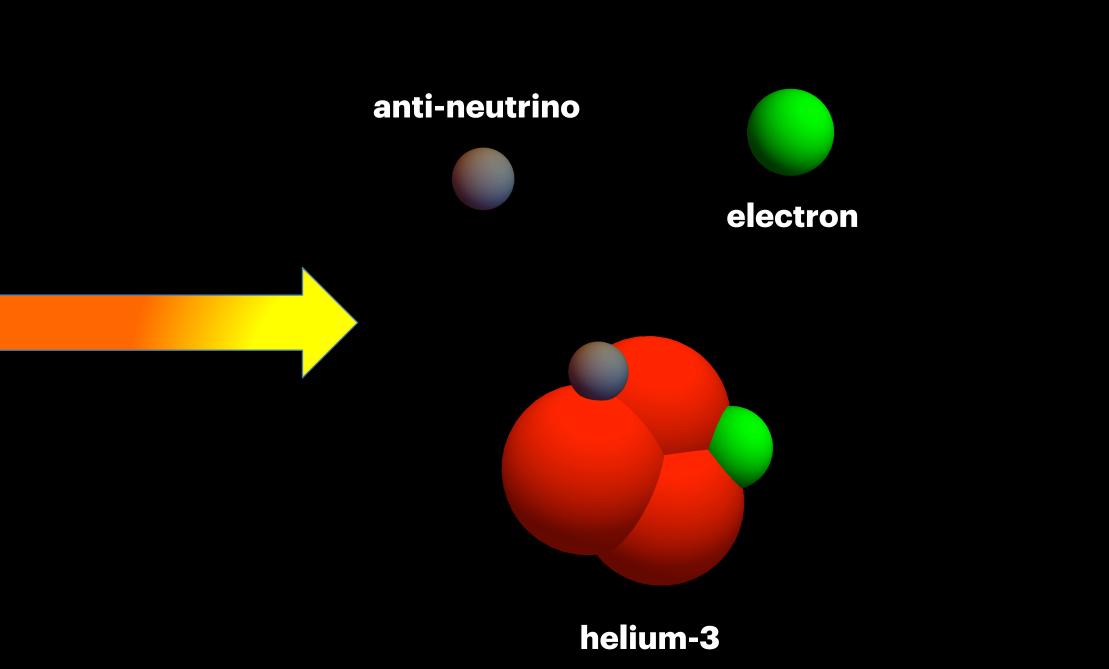


## a new model, a new particle (Pauli, 1930)

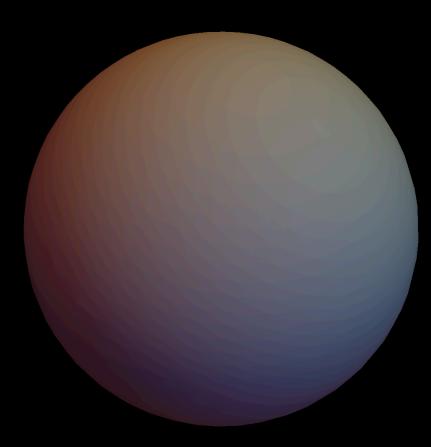


tritium

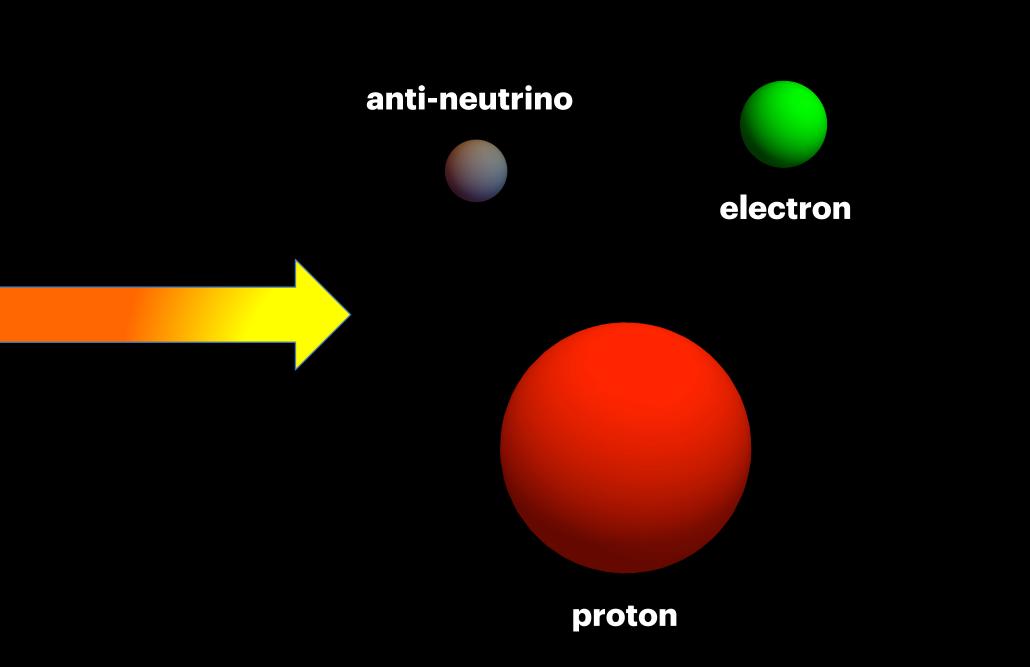
nuclei contain also neutrinos, that steal energy and spin (in fact they have spin 1/2, as all other matter particles)



## a relativistic theory of $\beta$ -rays (Fermi, 1933)

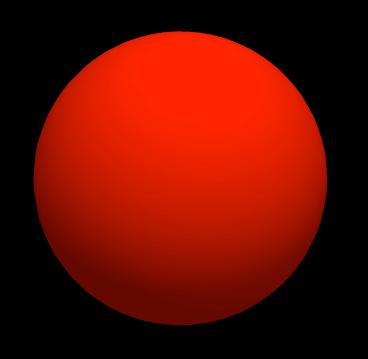


neutron

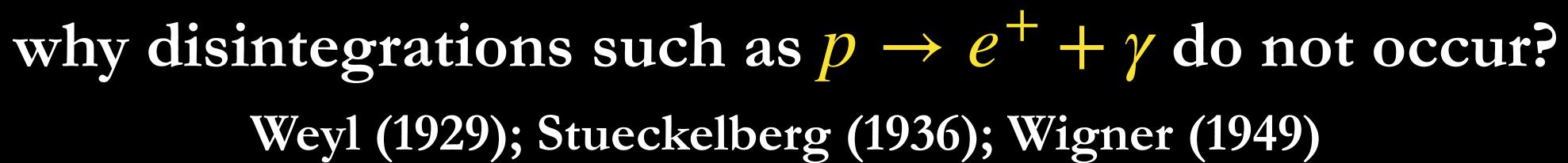


some particles of matter disappear and others appear, just as Einstein's "lichtquanten" do

### this brings us back to matter stability dilemma

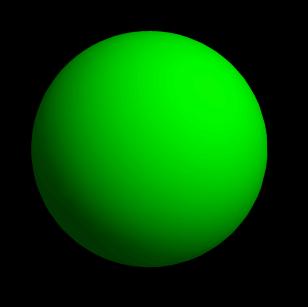


proton



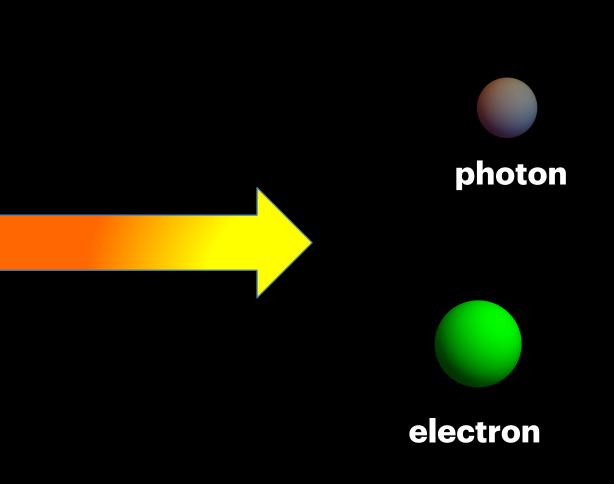


### and similar dilemma for leptons



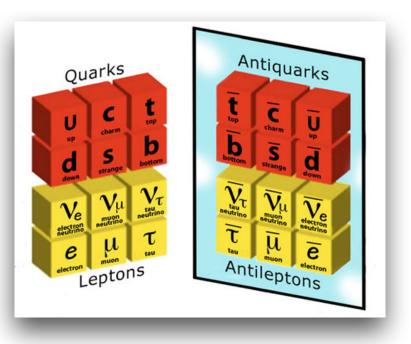
muon

### why disintegrations such as $\mu \rightarrow e + \gamma$ do not occur? Marx; Zel'dovich; Konopinsky Mahmoud 52-53 - but also Pontecorvo 47, Puppi 48)



### foundations of the standard model

- $\approx$  number of baryons **B** (1929-1949) and leptons **L** (1947-1953) are postulated to be conserved, to reconcile observations and theory
- after the discovery of parity violation (1956), a further hypothesis is invoked to proceed: that neutrinos are massless (1957)
- $\overleftrightarrow$  this paved the way to **V-A** theory of weak interaction (1957-1958) a cornerstone of the SM (1960-1967)



# status of global numbers

#### neutrino appearance experiments proved that there is only one basic type of lepton

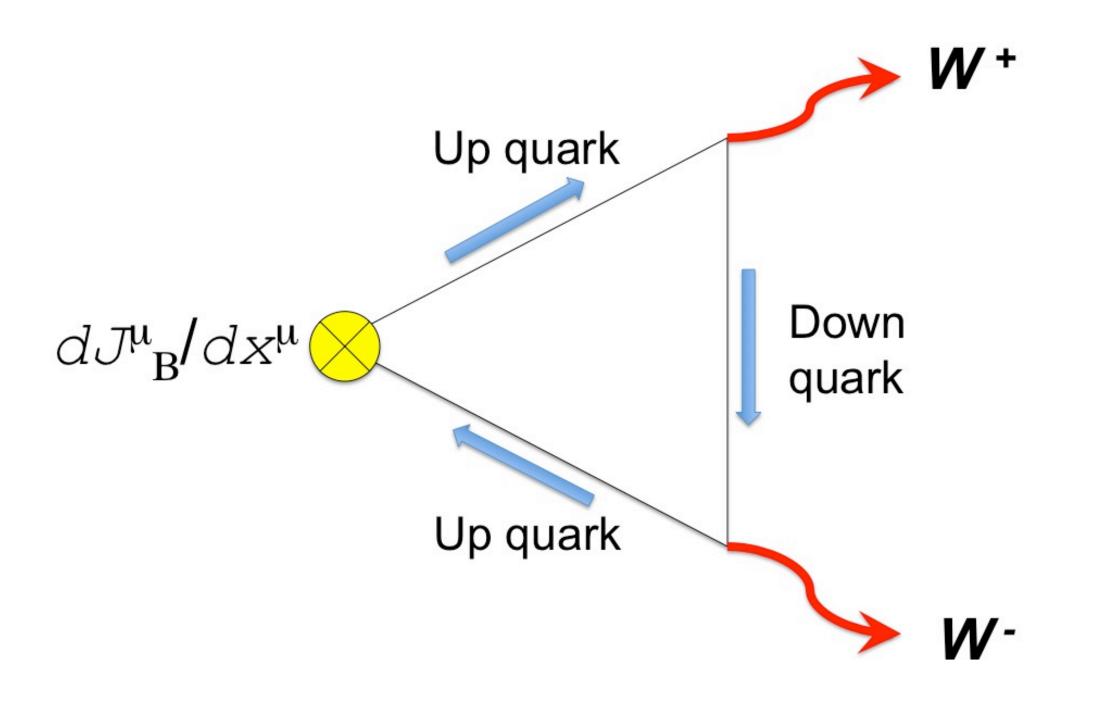
(=at the scrutiny of T2K, NOvA, OPERA, SK, DeepCore, only total lepton number L survived )

	ΔL <sub>e</sub>	ΔL <sub>μ</sub>	ΔL <sub>τ</sub>	ΔL
v <sub>µ</sub> →v <sub>e</sub>	+1	-1	0	0
ν <sub>μ</sub> →ν <sub>τ</sub>	0	-1	+1	0

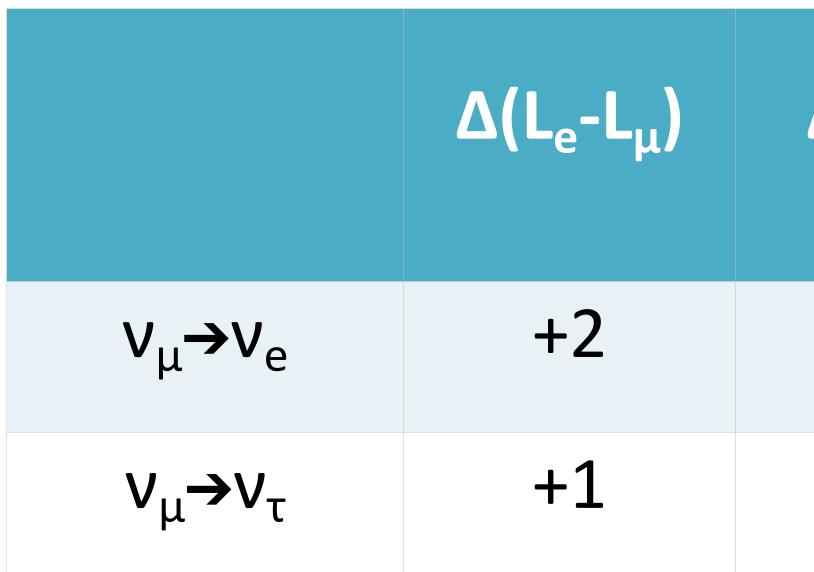
We have tested that all global symmetries of SM are violated, except **L** and **B**. Conversion among families is possible, we have only two fundamental types of matter particles: leptons and quarks

#### but in the SM, B and L are not separately conserved: B-L is conserved exactly; instead, B, L, B+L are not.

#### thus, in SM L and B are intimately connected

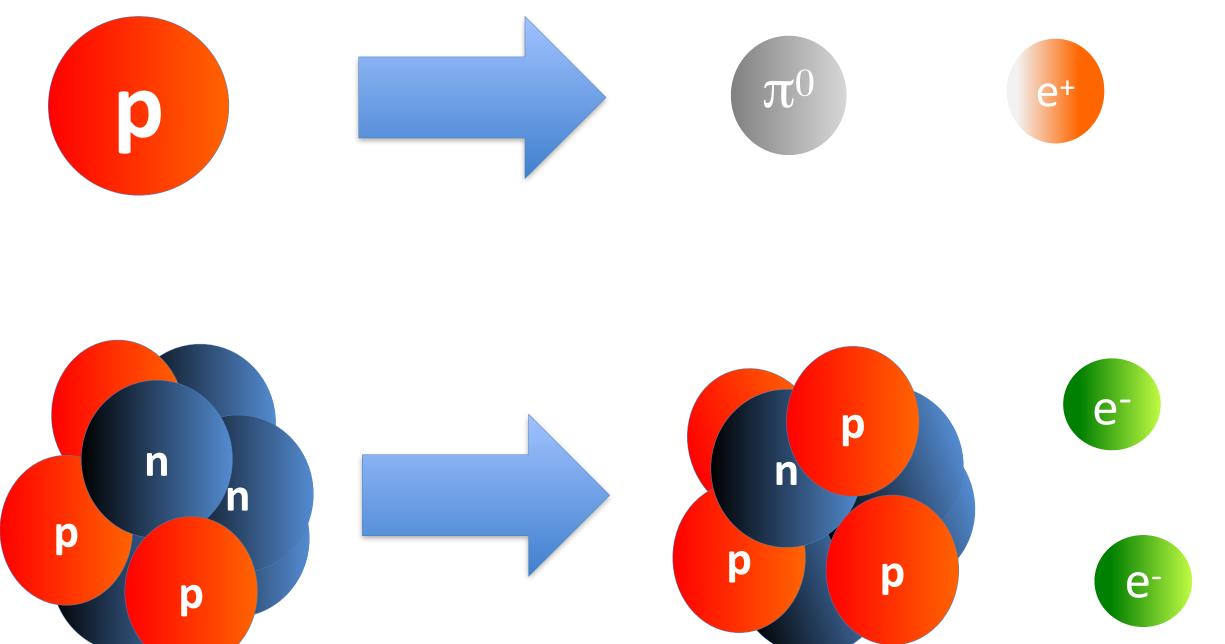


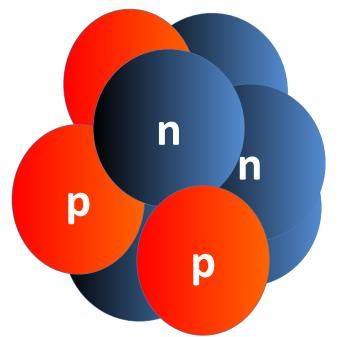
#### neutrino appearance experiments + SM imply that the only potentially exact symmetry is B-L

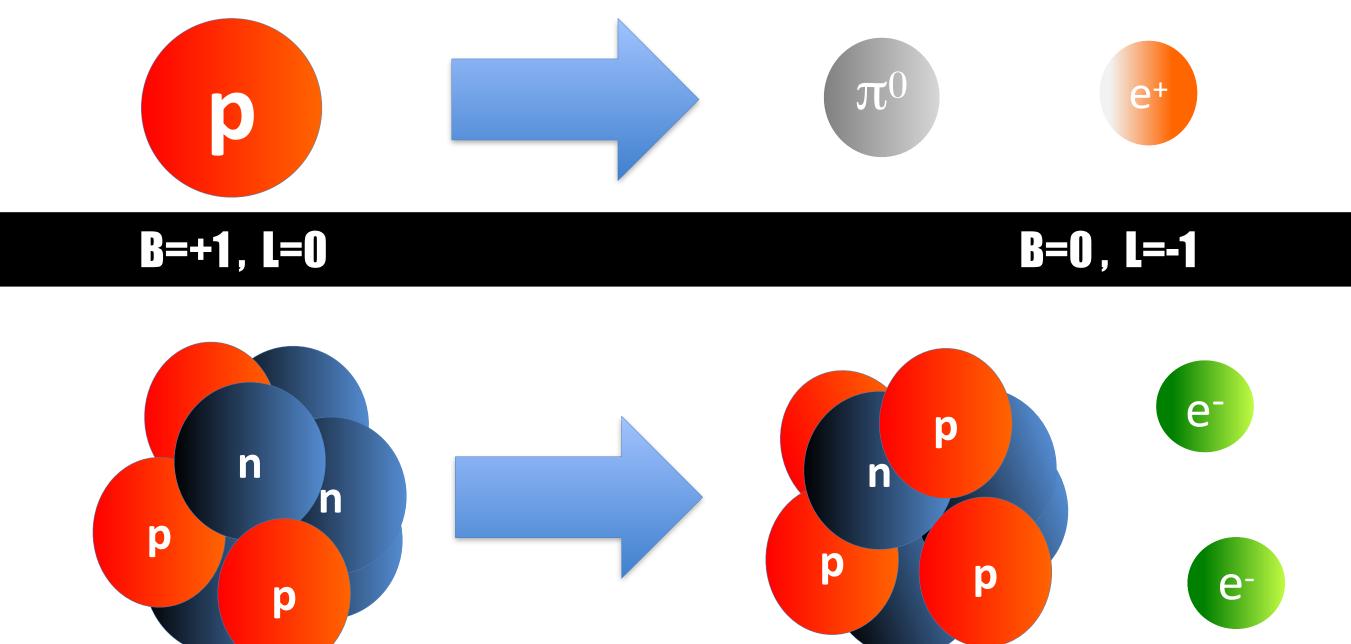


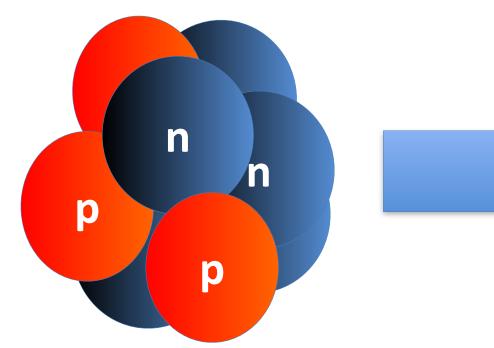
 $\Rightarrow$  there is an intimate connection between leptons and quarks. One question that immediately arises is what is the degree of violation of **B**, **L**, etc

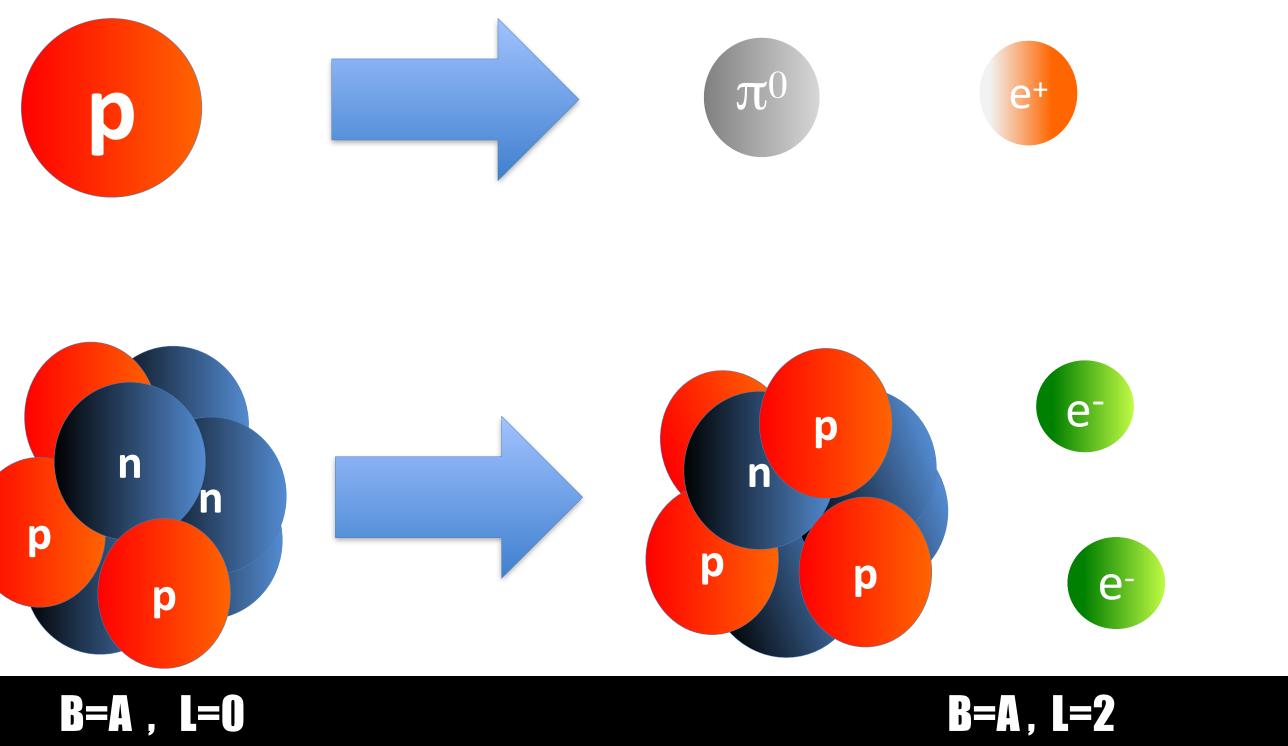
$\Delta(L_{\mu}-L_{\tau})$	Δ(L <sub>τ</sub> -L <sub>e</sub> )	<b>Δ(B-L)</b>
-1	-1	0
-2	+1	0

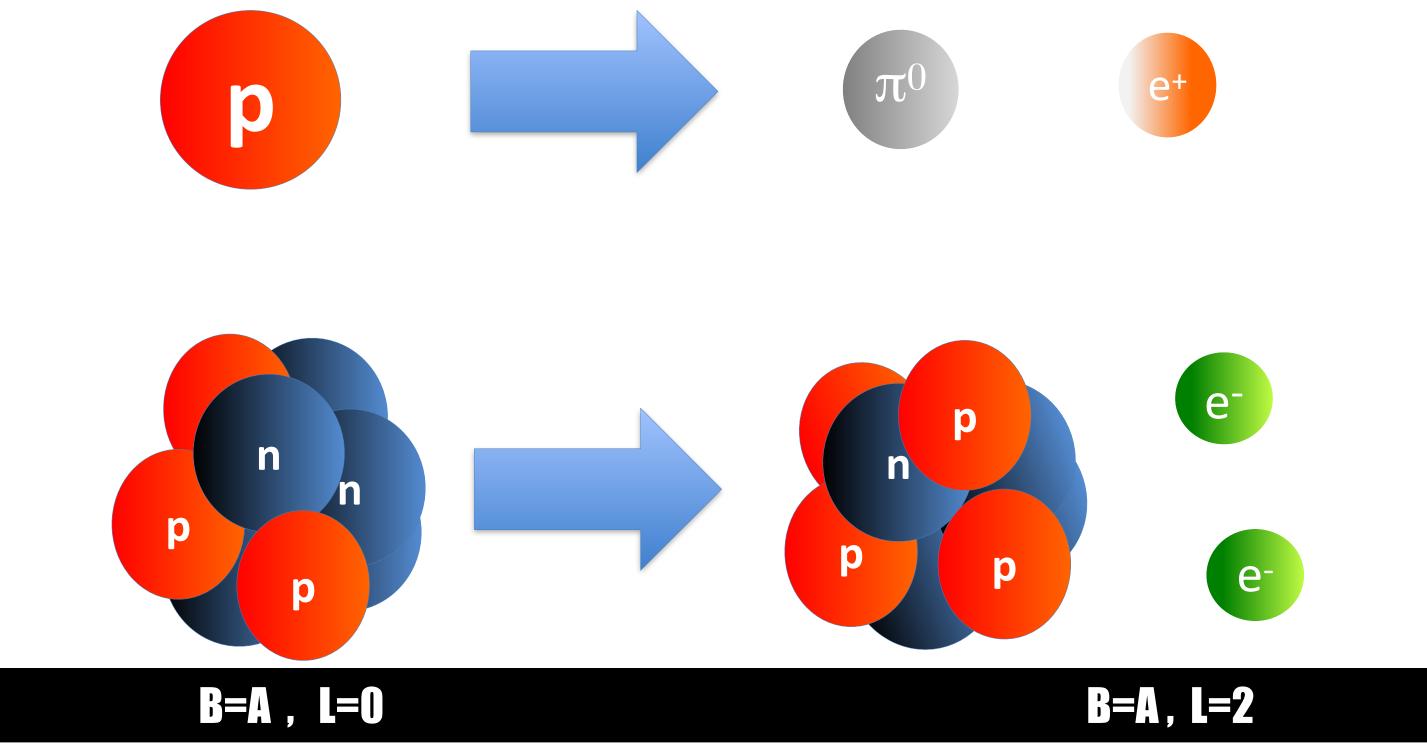




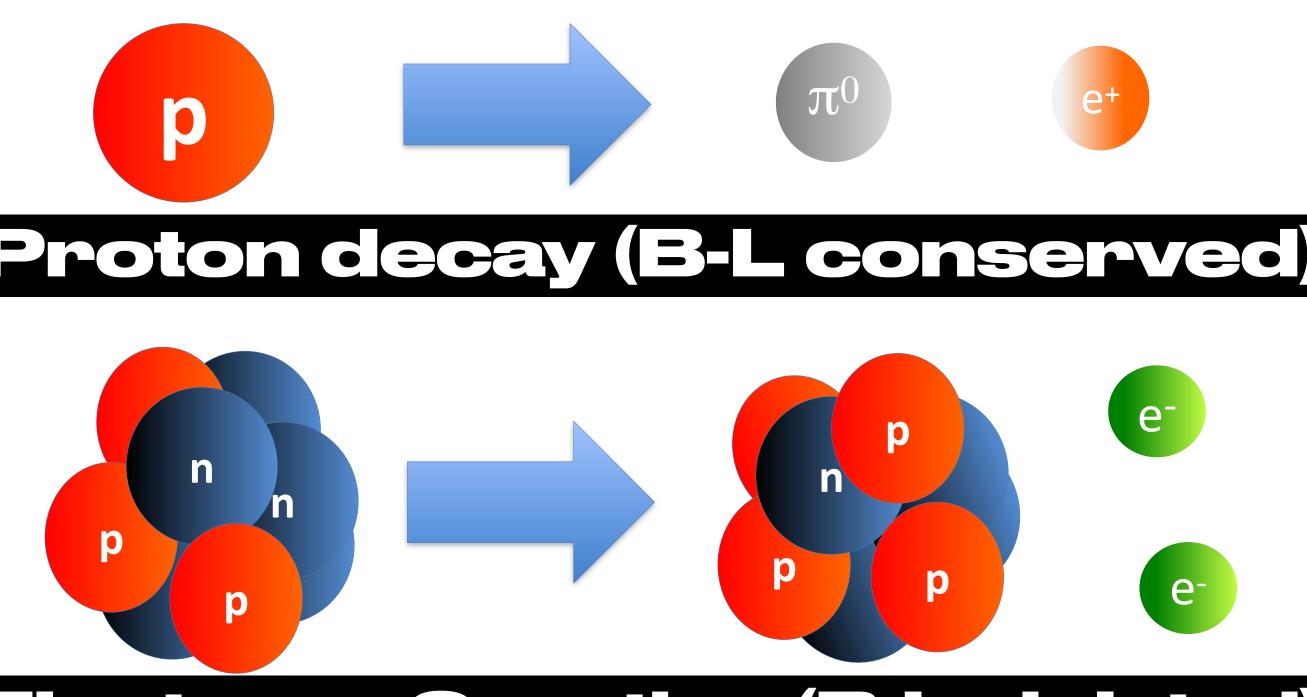


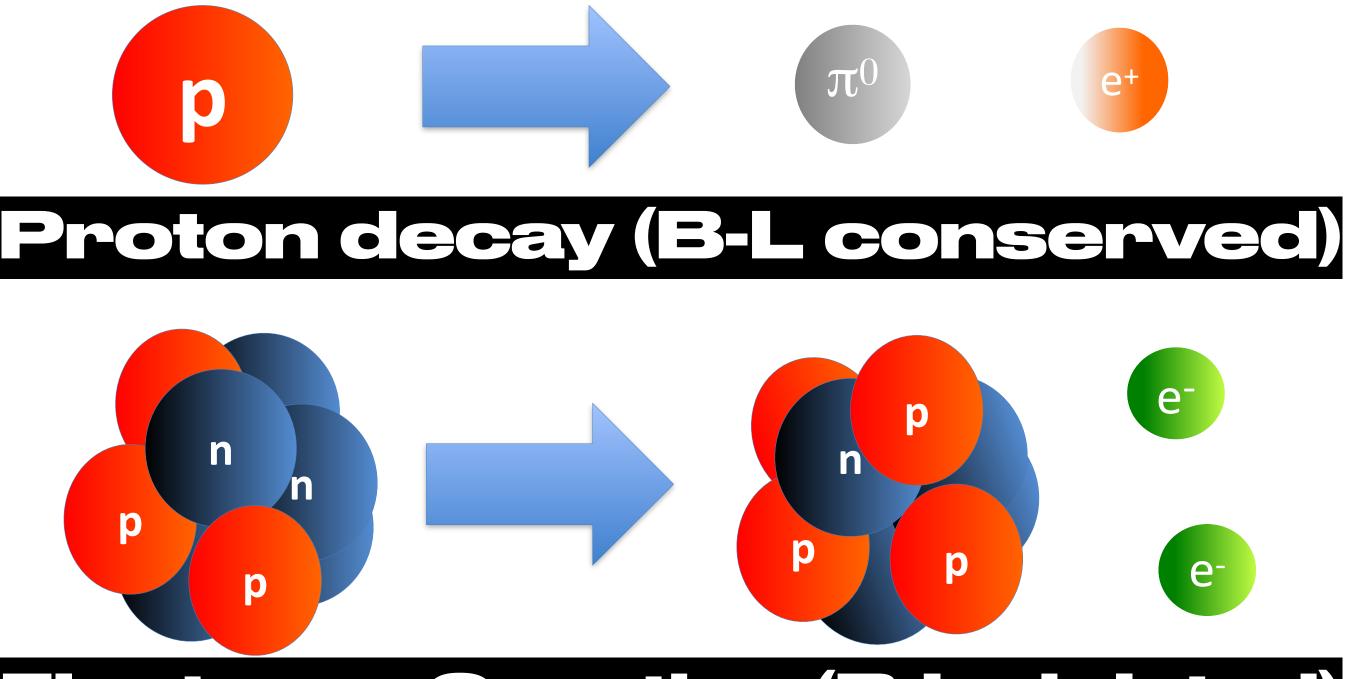






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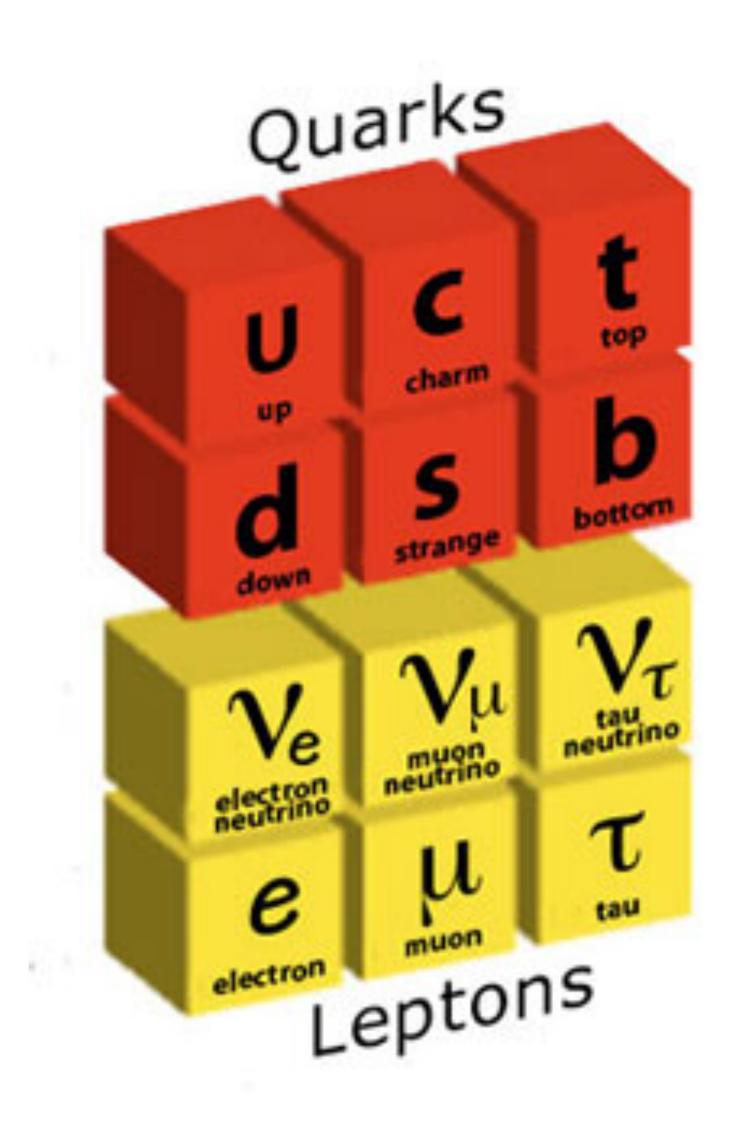


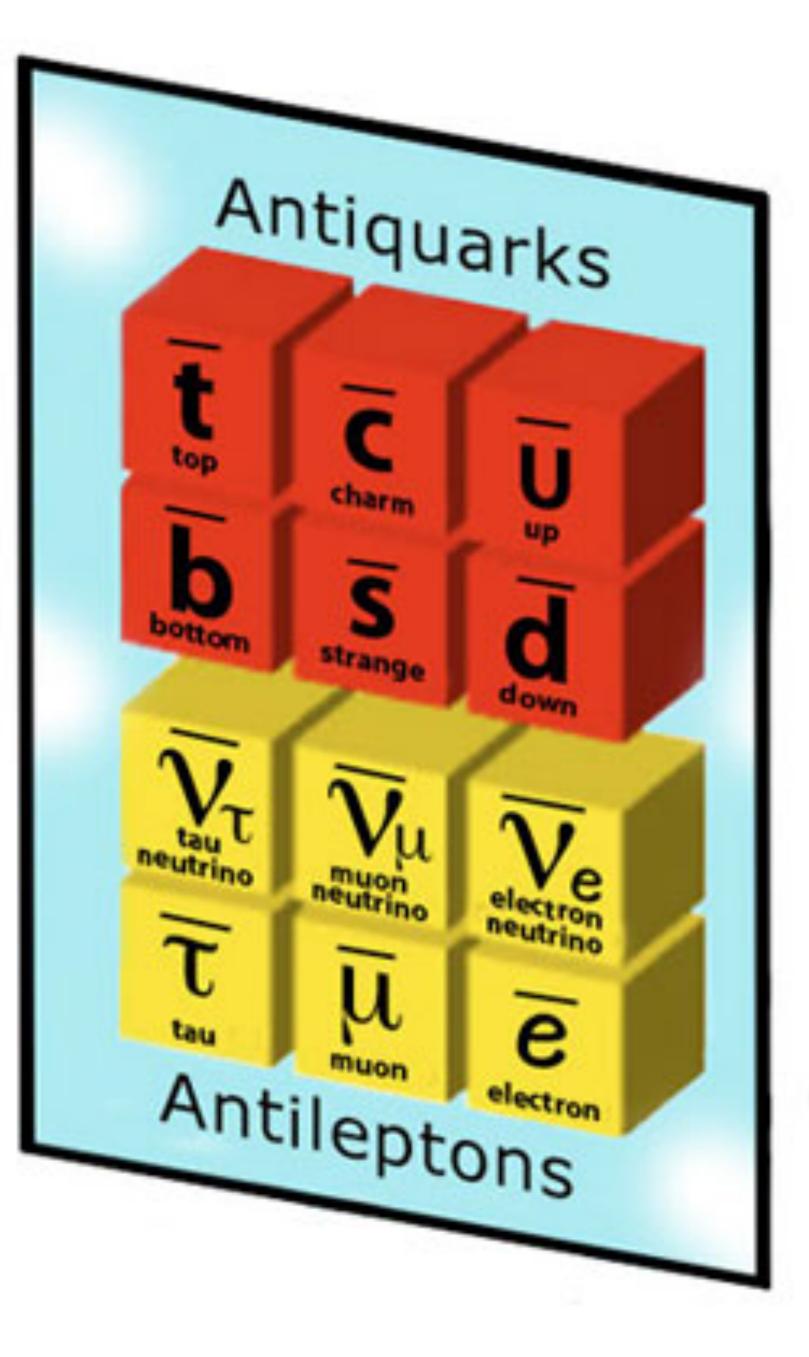


#### Electrons Creation (B-L violated)

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# features of the standard model and proof that it has to be extended





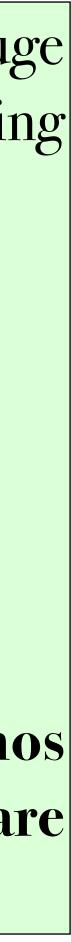
Matter and antimatter particles Credit: Fermilab

This useful picture conveys a huge amount of information, evoking the concepts of:

**\*** particles/antiparticles *\* quarks/leptons \* family replication* 

But it raises a question: what distinguishes neutrinos and antineutrinos, as they are both chargeless?





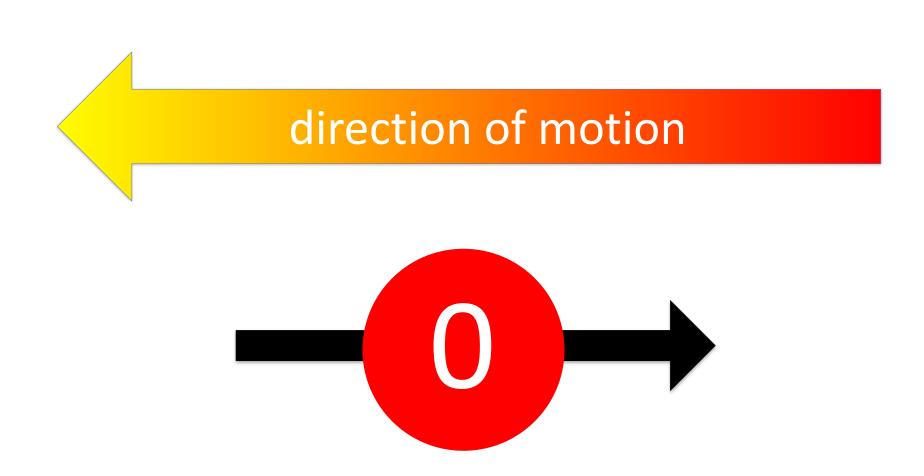
## on the structure of the standard model

in perturbation theory. Their differences Li-Li and B-L are exact

the standard model predicts that the 3 lepton numbers are all conserved

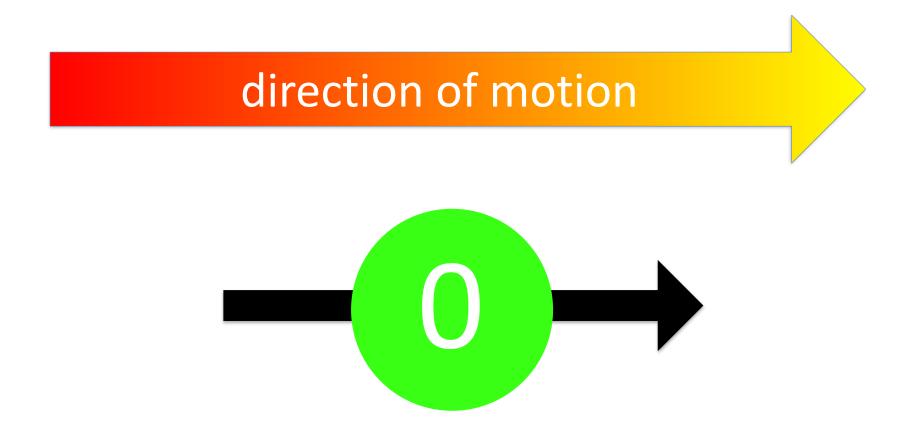
### on the structure of the standard model

in perturbation theory. Their differences Li-Li and B-L are exact



lacksquarebased on the masslessness of neutrinos.

the standard model <u>predicts</u> that the 3 lepton numbers are all conserved



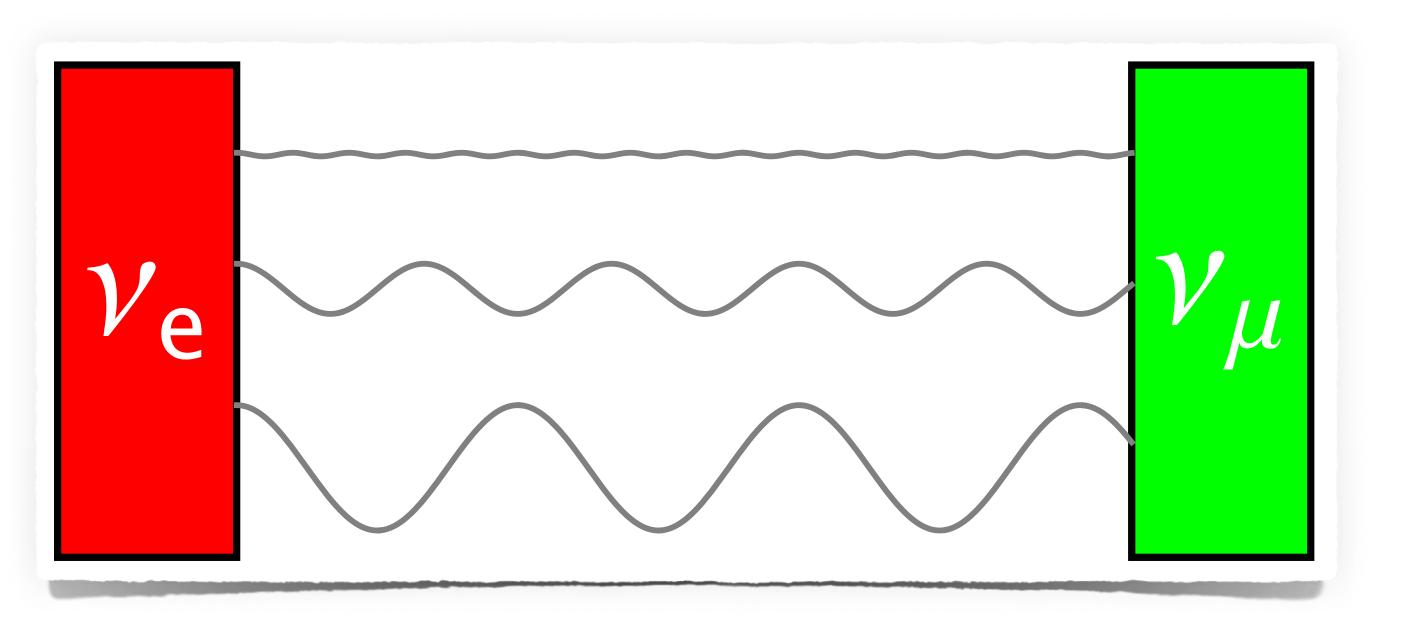
Helicity distinguishes neutrinos from antineutrinos - a feature of SM,

#### however, neutrinos do have mass.

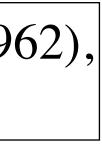
a quantum phenomenon, neutrino oscillations (1957-1967), indicates this beyond any doubt.

the proof, achieved with great efforts lasting more than 30 years, was recognized by the Nobel Prize awarded to Kajita and McDonald (2015)

oscillations, B. Pontecorvo (1957-1967); neutrino mixing, Y. Katayama, K. Matumoto, S. Tanaka, E. Yamada (1962), Z. Maki, M. Nakagawa, S. Sakata (1962) M. Nakagawa, H. Okonogi, S. Sakata, A. Toyoda (1963)







the evidence for neutrino mass implies physics beyond SM, and it shakes the foundations of the difference between neutrinos and antineutrinos (inherent in the SM)



an argument for neutrino-antineutrino identity or, introducing Majorana fermions



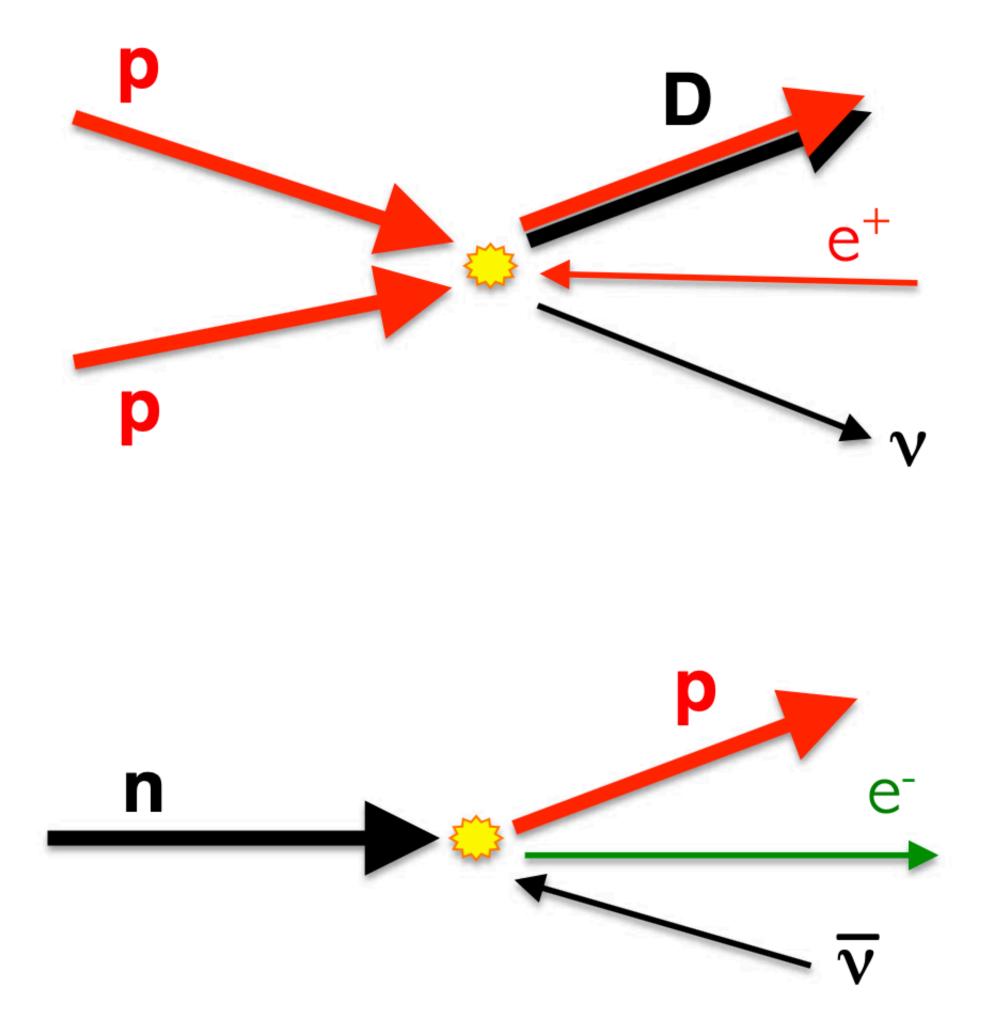


#### $p+p \rightarrow D + e^+ + v$

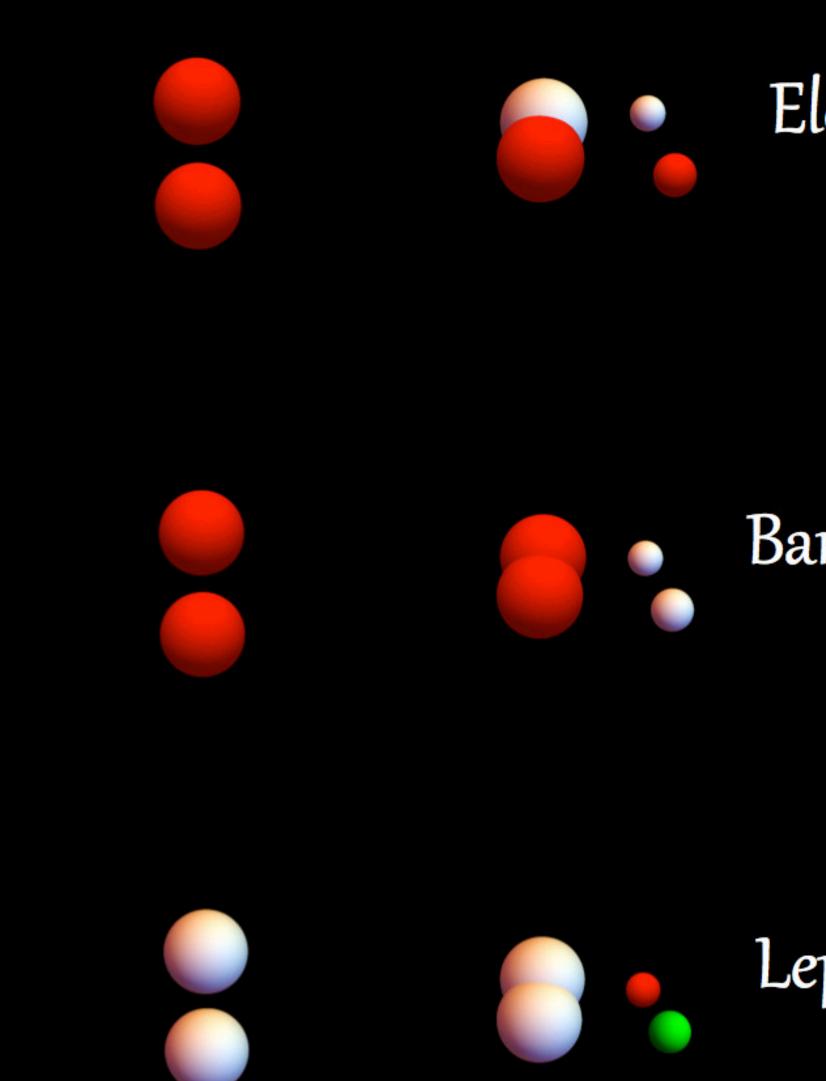
The neutrino is the matter particle that accompanies the positron

### $n \rightarrow p + e + \overline{v}$

The antineutrino is the antimatter particle that accompanies the electron



### proton fusion: $p+p \rightarrow D+e^++V$

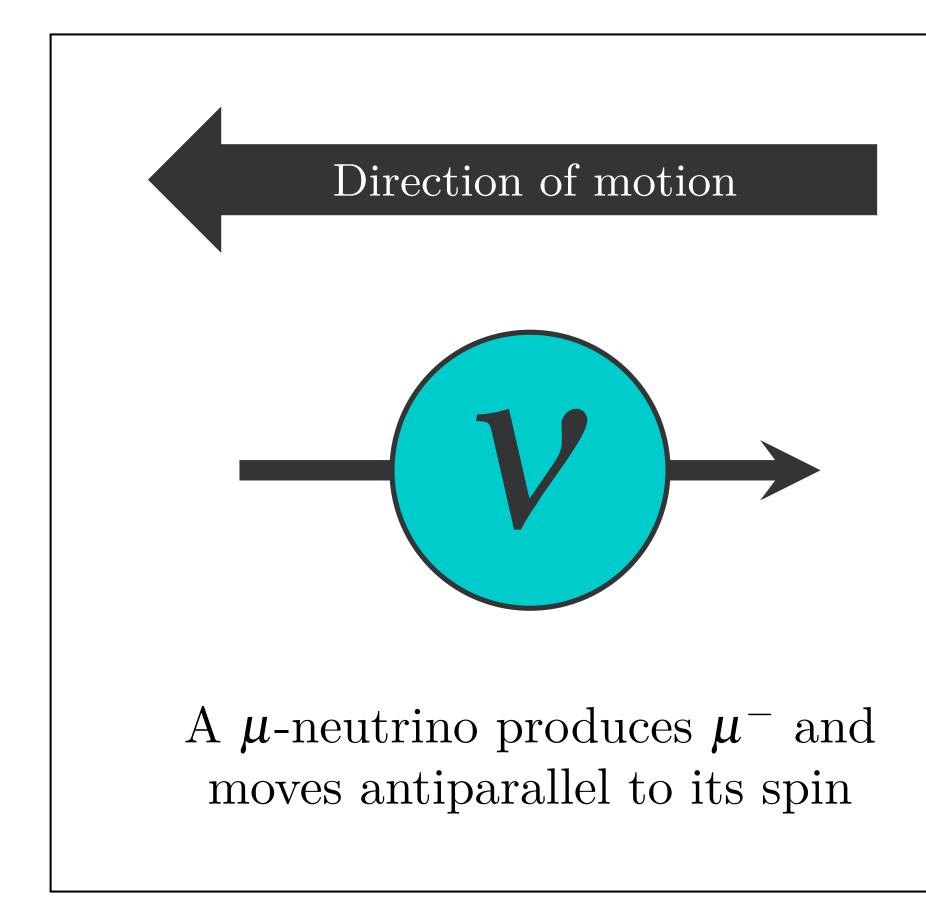


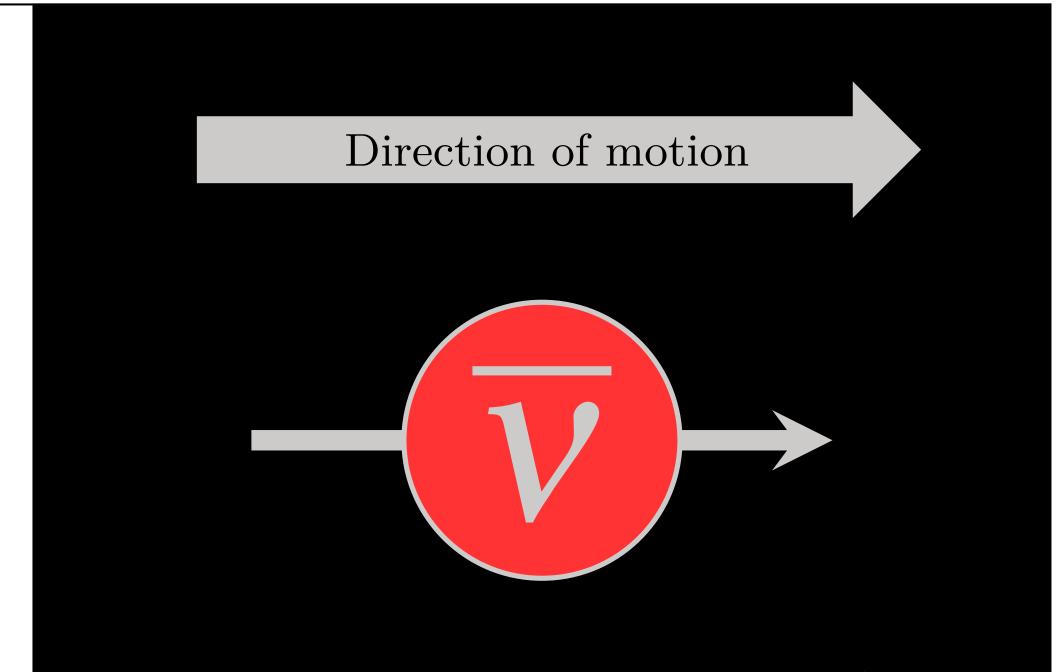
### Electric charge is conserved: 1+1=1+1+0

### Baryon number is conserved : 1+1=2+0+0

Lepton number is conserved: 0+0=0-1+1

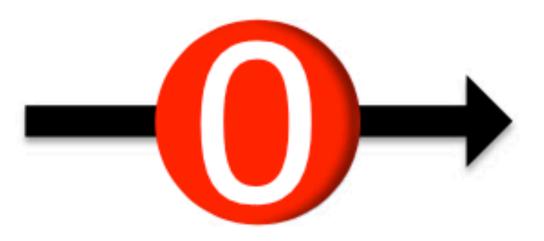
# Therefore neutrinos should be classified among matter particles, aren't them?



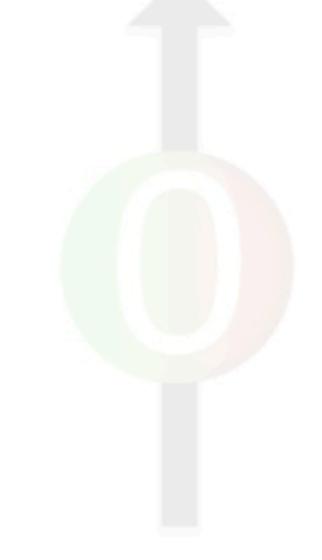


#### A $\mu$ -antineutrino produces $\mu^+$ and moves parallel to its spin

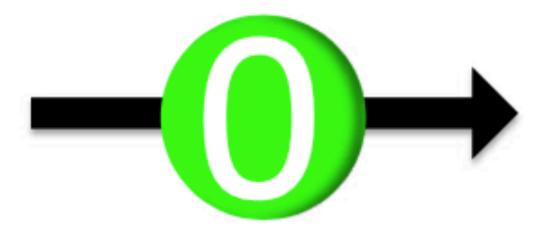
direction of motion



### neutrinos and antineutrinos distinguished by their helicities

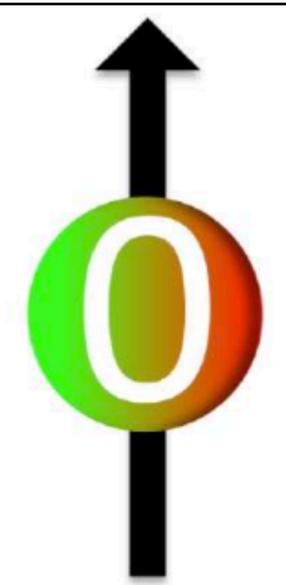


direction of motion



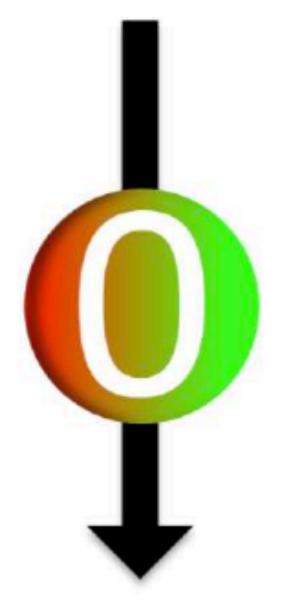
direction of motion

### and in their rest frame? they could be indistinguishable...

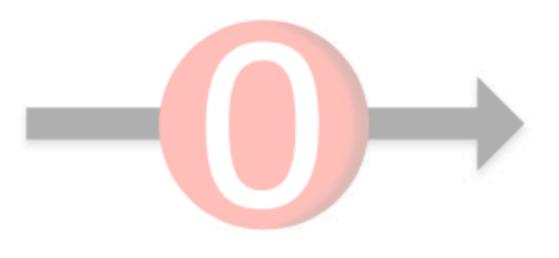


direction of motion

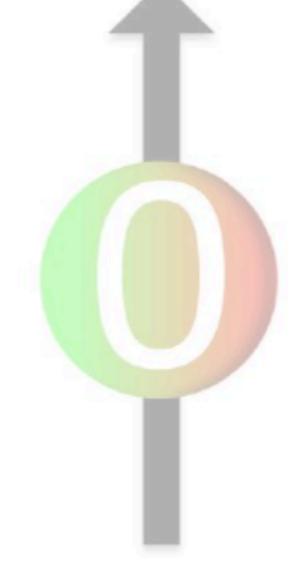




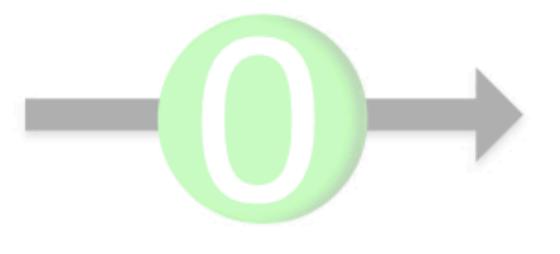
direction of motion

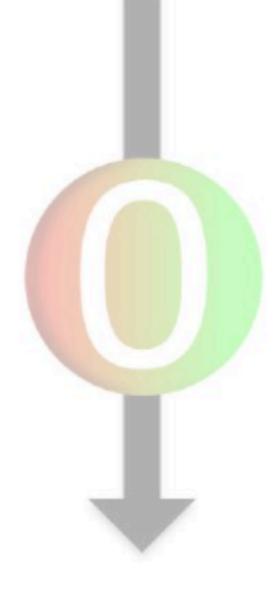


### Majorana theory: they're matter & antimatter at once



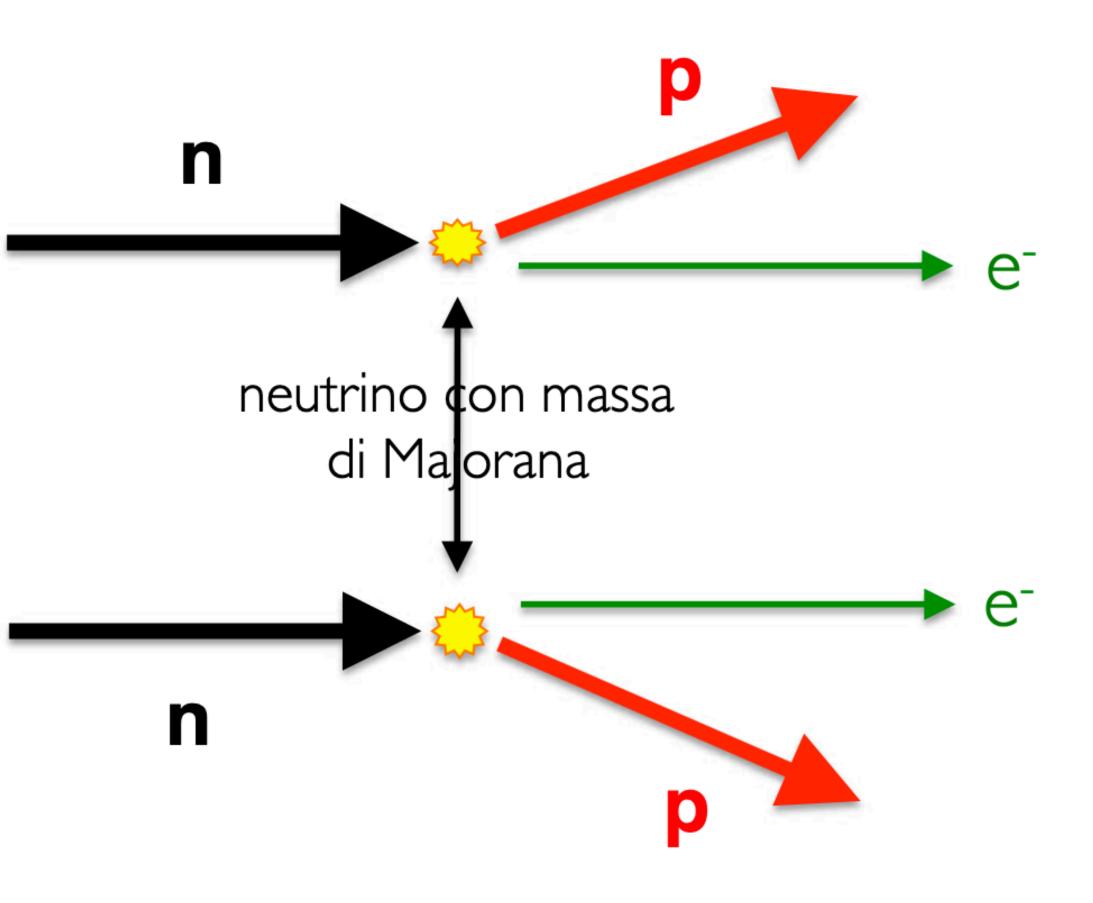
direction of motion

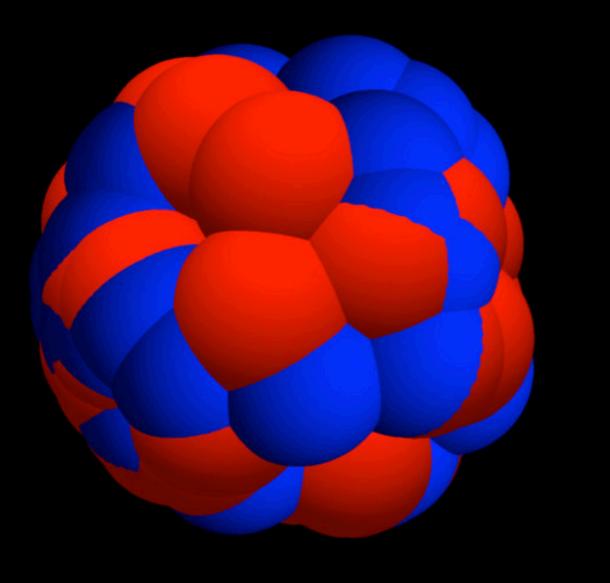


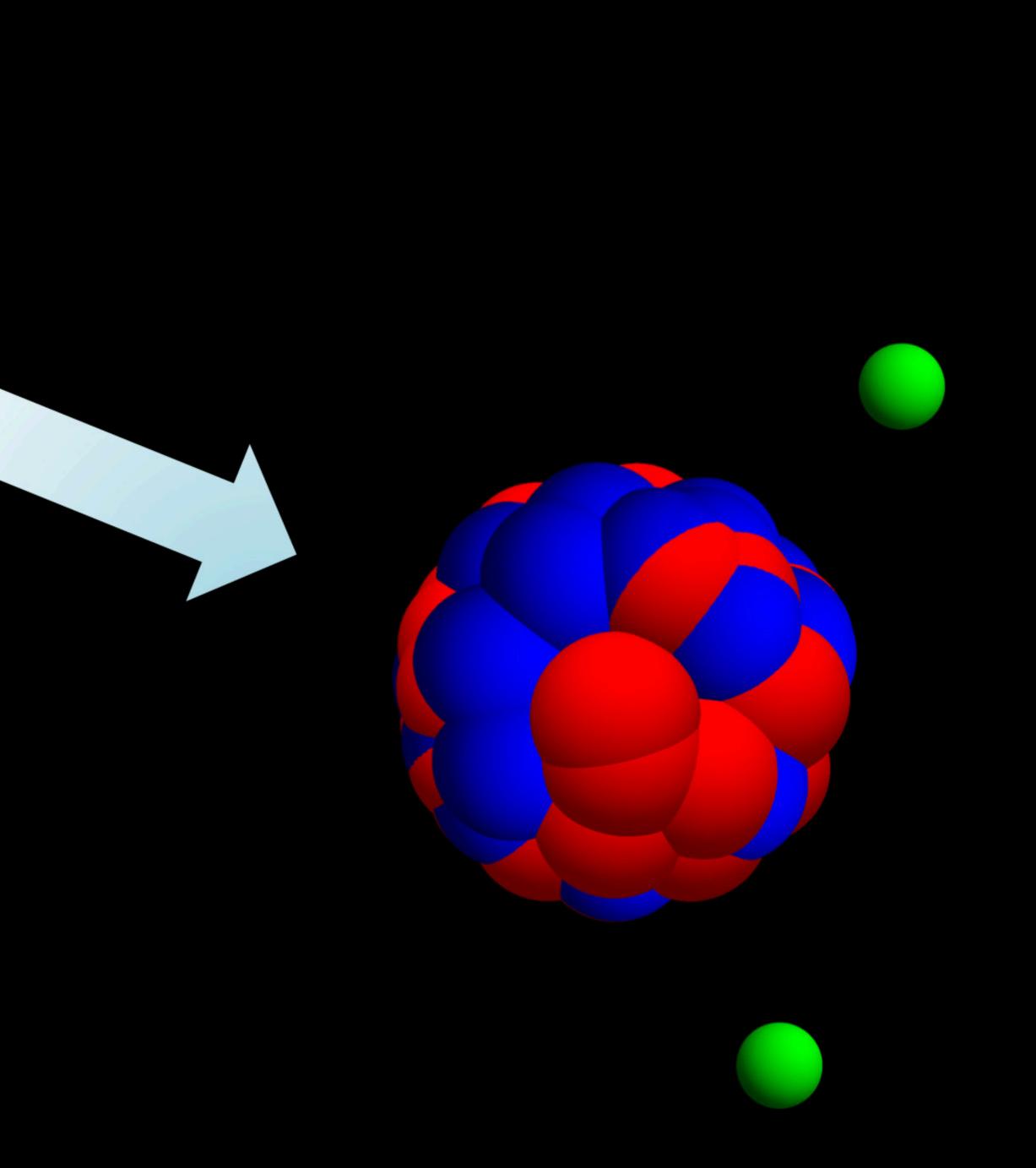


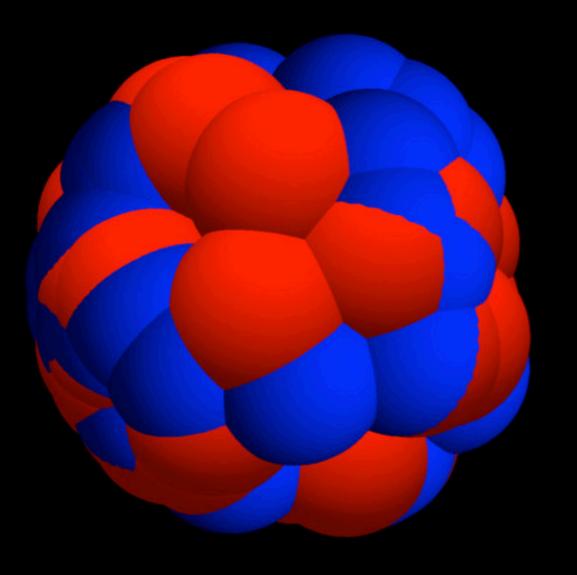
#### 2 n → 2 p + 2 e

The neutrino of Majorana is matter and antimatter at the same time: it leads to the creation of a couple of electrons



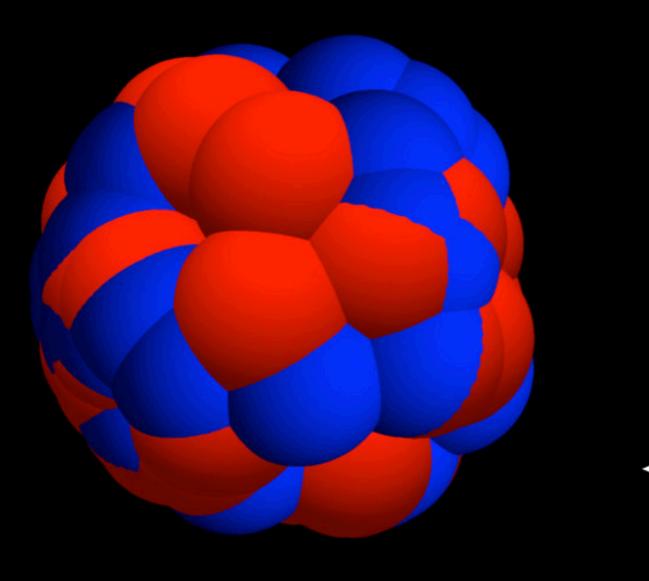




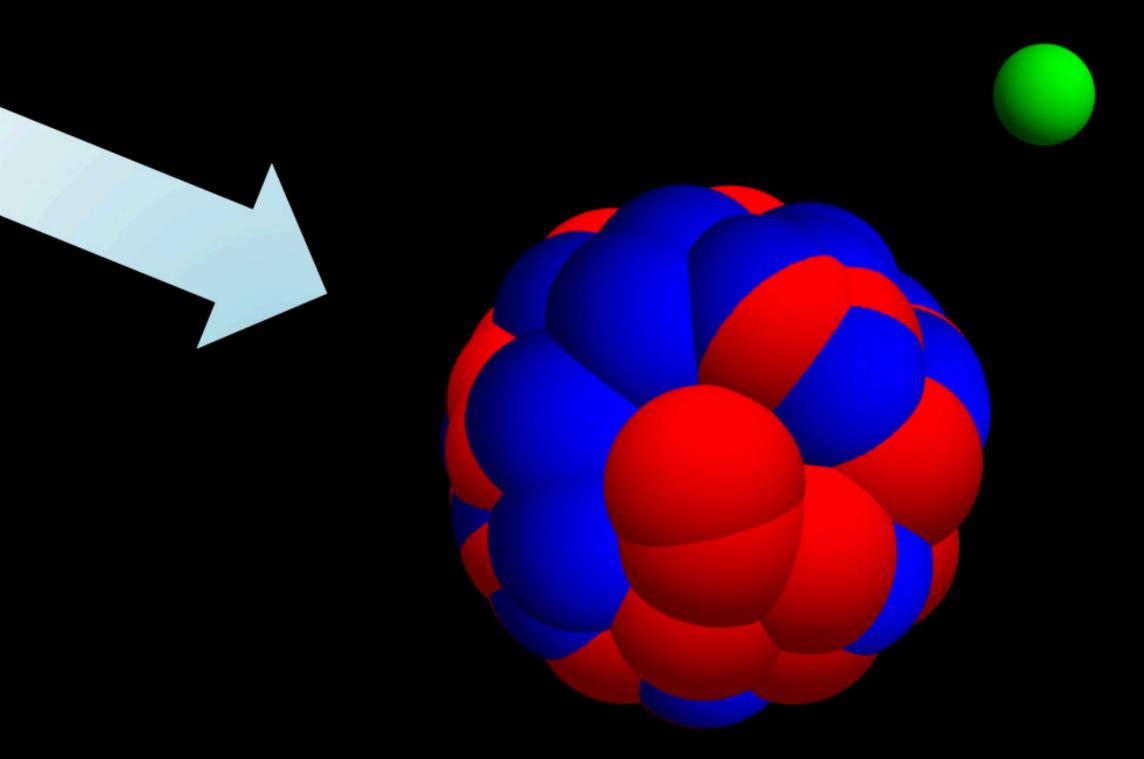


# electron creation





#### aka, neutrinoless double beta decay





# Several experiments at the Gran Sasso labs and elsewhere are looking for evidence...





#### Toward the Discovery of Lepton Creation with Neutrinoless Double- $\beta$ Decay

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Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, UK

Giovanni Benato

INFN, Laboratori Nazionali del Gran Sasso, 67100 Assergi, L'Aquila, Italy

Jason A. Detwiler<sup>‡</sup>

Center for Experimental Nuclear Physics and Astrophysics, and Department of Physics, University of Washington, Seattle, WA 98115 - USA

Javier Menéndez<sup>§</sup>

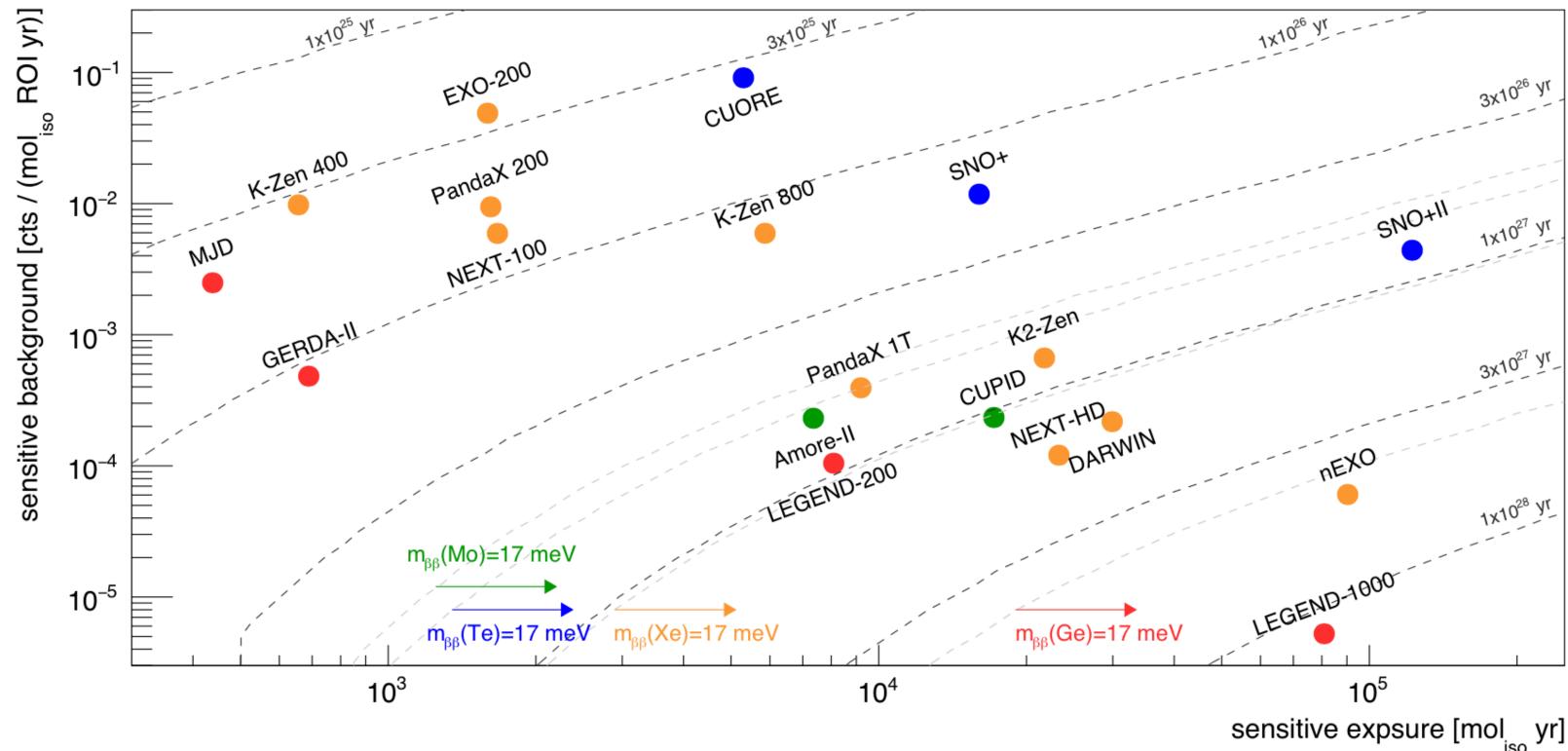
Department of Quantum Physics and Astrophysics and Institute of Cosmos Sciences, University of Barcelona, 08028 Barcelona, Spain

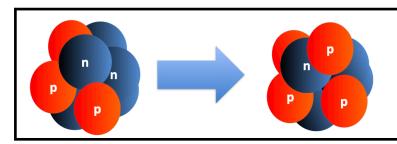
Francesco Vissani

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(Dated: September 24, 2021)

The discovery of neutrinoless double- $\beta$  decay could soon be within reach. This hypothetical ultra-rare nuclear decay is a portal to new physics beyond the Standard Model. Its observation would constitute the discovery of a matter-creating process, corroborating leading theories of why the universe contains more matter than antimatter. It would also prove that neutrinos and anti-neutrinos are not two distinct particles, but can transform into each other, generating their own mass in the process. The recognition that neutrinos are not massless necessitates an explanation and has boosted interest in neutrinoless double- $\beta$  decay. The field is now at a turning point. A new round of experiments is currently being proposed for the next decade to cover an important region of the parameter space. Advancements in nuclear theory are laying the groundwork to connect the nuclear decay with its underlying mechanisms. Meanwhile, the particle theory landscape continues to find new motivations for neutrinos to be their own antiparticle. This review brings together the experimental, nuclear theory, and particle theory aspects connected to neutrinoless double- $\beta$  decay, with the goal of exploring the path toward - and beyond - its discovery.









# insights on Dirac equation and formal introduction to Majorana ideas



### **insights into the Dirac equation** to reason about neutrinos instead of electrons

Dirac's equation is essential for dealing with fermions, including neutrinos.

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But it was developed for atomic physics, where the motions are non-

- Dirac's equation is essential for dealing with fermions, including neutrinos.
- relativistic; initially Dirac stressed the theory of the spin and Landé factor (!)

By P. A. M. DIRAC, St. John's College, Cambridge.

(Communicated by R. H. Fowler, F.R.S.-Received January 2, 1928.)

The new quantum mechanics, when applied to the problem of the structure of the atom with point-charge electrons, does not give results in agreement with experiment. The discrepancies consist of "duplexity" phenomena, the observed number of stationary states for an electron in an atom being twice the number given by the theory. To meet the difficulty, Goudsmit and Uhlenbeck have introduced the idea of an electron with a spin angular momentum of half a quantum and a magnetic moment of one Bohr magneton. This model for the electron has been fitted into the new mechanics by Pauli,\* and Darwin,†

The Quantum Theory of the Electron.



### **insights into the Dirac equation** to reason about neutrinos instead of electrons

- Dirac's equation is essential for dealing with fermions, including neutrinos.
- But it was developed for atomic physics, where the motions are non-
- relativistic; initially Dirac stressed the theory of the spin and Landé factor (!)
- To understand neutrinos well, it is useful to deepen at least these aspects
  - its Lorentz structure
  - Ite formal origin of antiparticles
    Stueckelberg 1941
  - It the treatment of neutral particles
    Majorana 1937

#### **Weyl 1929**

### spinors and relativity let us begin with rotations, part of Lorentz group

As it is well-known, spinors transform with half-angles

With them, we can define one scalar and axial vector (=polarization)

$$S = \psi^{\dagger} \psi$$
 a

that transform as expected, e.g.

 $\delta \psi = -\frac{i \vec{\theta} \vec{\sigma}}{2} \psi$ 

- and  $\overrightarrow{V} = \psi^{\dagger} \overrightarrow{\sigma} \psi$

$$= \overrightarrow{\theta} \wedge \overrightarrow{V}$$

[check it]

#### spinors and relativity let us continue with velocity transformations

Under **boosts**, spinors transform as

 $\delta \psi =$ 

One uses them with the quantity already defined

$$S = \psi^{\dagger}\psi$$
 and  $\overrightarrow{V} = +\psi^{\dagger}\overrightarrow{\sigma}\psi$ 

and we find

$$= + \frac{\overrightarrow{\eta} \overrightarrow{\sigma}}{2} \psi$$

 $\delta S = \overrightarrow{\eta} \overrightarrow{V}$  and  $\delta \overrightarrow{V} = \overrightarrow{\eta} S$ 

[check it]

### spinors and relativity let us continue with velocity transformations

Under **boosts**, spinors transform (alternatively) as

$$S = \psi^{\dagger}\psi$$
 and  $\overrightarrow{V} = -\psi^{\dagger}\overrightarrow{\sigma}\psi$ 

and we find

 $\delta S = \overrightarrow{n} V$ 

 $\delta \psi = -\frac{\eta' \,\overline{\sigma}}{2} \,\psi$ 

One uses them with the quantity already defined (in an alternative manner)

and 
$$\delta \overrightarrow{V} = \overrightarrow{\eta} S$$

[check it]

### spinors and relativity summary

There are **two types** of bi-spinors in relativity (Weyl)  $\delta \psi_{\pm} = -\frac{i \overrightarrow{\theta} \overrightarrow{\sigma}}{2} \psi_{\pm}$ 

$$\pm , \quad \delta \psi_{\pm} = \pm \frac{\overrightarrow{\eta} \overrightarrow{\sigma}}{2} \psi_{\pm}$$

### spinors and relativity summary

There are **two types** of bi-spinors in relativity (Weyl)  $\delta \psi_{\pm} = -\frac{i \overrightarrow{\theta} \overrightarrow{\sigma}}{2} \psi_{\pm}$ 

we better think in terms of 4-vectors

$$\pm , \quad \delta \psi_{\pm} = \pm \frac{\overrightarrow{\eta} \overrightarrow{\sigma}}{2} \psi_{\pm}$$

- $V^0 \equiv S = \psi_+^{\dagger} \psi_{\pm}$  and  $\overrightarrow{V} = \pm \psi_+^{\dagger} \overrightarrow{\sigma} \psi_{\pm}$

### spinors and relativity summary

There are **two types** of bi-spinors in relativity (Weyl)  $\delta \psi_{\pm} = -\frac{i \overrightarrow{\theta} \overrightarrow{\sigma}}{2} \psi_{\pm}$ 

we better think in terms of 4-vectors  $V^0 \equiv S = \psi_+^\dagger \psi_+$ 

 $\delta V^{\mu}$ 

[obtain it by yourself]

$$\pm , \quad \delta \psi_{\pm} = \pm \frac{\overrightarrow{\eta} \overrightarrow{\sigma}}{2} \psi_{\pm}$$

and 
$$\overrightarrow{V} = \pm \psi_{\pm}^{\dagger} \overrightarrow{\sigma} \psi_{\pm}$$

and, with a suitable definition of the 6 parameters  $\omega^{\mu\nu} = -\omega^{\nu\mu}$ , we have

$$=\omega^{\mu}_{\nu}V^{\nu}$$

### spinors and relativity one revision exercise

$$\mathscr{L} = i \ \psi^{\dagger}_{+} \ \sigma^{\mu} \ \partial_{\mu} \psi$$

- consider the lagrangian density
  - with  $\sigma^{\mu} = (\sigma^0, \overrightarrow{\sigma})$
- and the associated wave-equation

$$i \ \sigma^{\mu} \partial_{\mu} \ \psi_{+} = 0$$

show that the positive energy solution represents a massless particle with positive helicity

### formal origin of antiparticles generalities

The coupling to electromagnetic interactions is described replacing

$$P_{\mu} \rightarrow \mathscr{P}_{\mu} = P_{\mu} - qA_{\mu}$$
 where  $P_{\mu} = i\hbar\partial_{\mu}$ 

### formal origin of antiparticles generalities

The coupling to electromagnetic interactions is described replacing

$$P_{\mu} \rightarrow \mathscr{P}_{\mu} = P_{\mu} - qA_{\mu}$$
 where  $P_{\mu} = i\hbar\partial_{\mu}$ 

Let us discuss the cases of scalar particles (described by Klein-Gordon equation) and the one of fermions (described by Dirac's equation)

### formal origin of antiparticles the case of scalar particles

Consider the simplest wave equation for an incoming wave

$$(\mathcal{P}_{\mu}\mathcal{P}^{\mu}-m^2) \ \varphi = ((P_{\mu}-qA_{\mu})^2-m^2) \ \varphi = 0$$

### formal origin of antiparticles the case of scalar particles

 $(\mathcal{P}_{\mu}\mathcal{P}^{\mu}-m^{2}) \varphi = ($ 

Consider the simplest wave equation for an incoming wave

$$((P_{\mu} - qA_{\mu})^2 - m^2) \varphi = 0$$

- the outgoing wave  $\phi^*$  satisfies automatically
- $((\mathcal{P}_{\mu}\mathcal{P}^{\mu})^* m^2) \varphi^* = ((P_{\mu} + qA_{\mu})^2 m^2) \varphi^* = 0$

### formal origin of antiparticles the case of scalar particles

 $(\mathcal{P}_{\mu}\mathcal{P}^{\mu}-m^{2}) \varphi = ($ 

Consider the simplest wave equation for an incoming wave

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- $((\mathcal{P}_{\mu}\mathcal{P}^{\mu})^{*} m^{2}) \varphi^{*} = ((P_{\mu} + qA_{\mu})^{2} m^{2}) \varphi^{*} = 0$

waves with "negative energies" can be thought as regular waves with  $q \rightarrow -q$ 



### formal origin of antiparticles the case of spinorial particles

 $(\mathcal{P}_{\mu}\gamma^{\mu}-m)\psi = (($ 

Dirac' equation

$$(P_{\mu}-qA_{\mu})\gamma^{\mu}-m)\psi=0$$

use Majorana representation, such as  $\gamma_{\mu}^* = -\gamma_{\mu}$ , finding

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 $(\mathcal{P}_{\mu}\gamma^{\mu}-m)^{*}\psi^{*}=(($ 

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use Majorana representation, such as  $\gamma_{\mu}^* = -\gamma_{\mu}$ , finding

$$(P_{\mu} + qA_{\mu})\gamma^{\mu} - m)\psi^* = 0$$

### formal origin of antiparticles the case of spinorial particles

$$(\mathcal{P}_{\mu}\gamma^{\mu}-m)\psi = ((P_{\mu}-qA_{\mu})\gamma^{\mu}-m)\psi = 0$$

$$(\mathcal{P}_{\mu}\gamma^{\mu} - m)^{*} \psi^{*} = ((P_{\mu} + qA_{\mu})\gamma^{\mu} - m) \psi^{*} = 0$$

Dirac' equation

use Majorana representation, such as  $\gamma_{\mu}^* = -\gamma_{\mu}$ , finding

waves with "negative energies" can be thought as regular waves with  $q \rightarrow -q$ 





there are particles such as the photon or the  $\pi^0$  which coincide with their own antiparticles, other instead such as the neutron or the  $K^0$  that instead don't

## neutral particles



#### we do not know in which class neutrino is

## neutral particles

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Majorana (1937) proposed that they are in the first class (completely neutral, or real) In this case their quantised field would obey,

[Majorana representation of gamma matrices]

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 $= \Psi *$ 



## neutral particles

#### we do not know in which class neutrino is

Majorana (1937) proposed that they are in the first class (completely neutral, or real) In this case their quantised field would obey,

## $\Psi = C \Psi^t$

[generic representation of gamma matrices]

there are particles such as the photon or the  $\pi^0$  which coincide with their own antiparticles, other instead such as the neutron or the  $K^0$  that instead don't



## consolidation exercises

- ★ study the three most common <u>representations</u> of the gammamatrices: Dirac's, Weyl's, Majorana and in particular:
- 1.obtain the expressions of the charge conjugation matrix  ${m C}$
- 2.0btain the expressions of the **chirality** matrix  $\gamma_5$
- write down the <u>equation of motion</u> of free 4-fermions in the Weyl representation and compare it with that of 2-fermions
- show that a <u>Majorana fermion</u> does not couple to the photon, neither by charge, nor by magnetic moment

# how we use this...

#### Given a set of masses $m_{a}$



E.g., with the 3 known neutrinos  $\boldsymbol{\nu}_\ell$  define

 $\boldsymbol{\chi}_i = U_{\ell i}^* \, \boldsymbol{\nu}_\ell + U_\ell$ 

then the Lagrangian density reads,

 $\mathcal{L}_{ extsf{Majorana}} = -rac{1}{2} m_{
u,\ell\ell'} \, ar{m{\iota}}$ 

with the complex symmetric mass matrix

 $m_{
u} =$ 

$$_i$$
 and of real fields,  $oldsymbol{\chi}_i = C \overline{oldsymbol{\chi}_i}^t$ ,

$$=-rac{1}{2}m_i \,\overline{oldsymbol{\chi}}_i oldsymbol{\chi}_i$$

$$\int_{\ell i} C \overline{\boldsymbol{\nu}_{\ell}}^t \left\{ \begin{array}{l} \ell = \mathbf{e}, \mu, \tau \\ i = 1, 2, 3 \end{array} \right.$$

$$\overline{\boldsymbol{\nu}_{\ell}} C \overline{\boldsymbol{\nu}_{\ell'}}^t + \frac{1}{2} m^*_{\boldsymbol{\nu},\ell\ell'} \, \boldsymbol{\nu}_{\ell}^t C^{\dagger} \boldsymbol{\nu}_{\ell'}$$

$$oldsymbol{U}$$
diag $(m)oldsymbol{U}^t$ 

# Given a set of masses $m_i$ and of real fields, $\chi_i = C \overline{\chi_i}^t$ , $\mathcal{L}_{Majorana} =$ E.g., with the 3 known neutrinos $u_{\ell}$ define $\boldsymbol{\chi}_i = U_{\ell i}^* \, \boldsymbol{ u}_\ell + U_\ell$ then the Lagrangian density reads, $\mathcal{L}_{ extsf{Majorana}} = -rac{1}{2}m_{ u,\ell\ell'}~ar{m{ u}}$ with the complex symmetric mass matrix $m_{ u} =$

$$=-rac{1}{2}m_i\,\overline{oldsymbol{\chi}}_i{oldsymbol{\chi}}_i$$

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$$oldsymbol{U}$$
diag $(m)oldsymbol{U}^t$ 

$$\mathcal{L}_{\mathsf{Majorana}}$$
 =

$$\boldsymbol{\chi}_{i} = U_{\ell i}^{*} \, \boldsymbol{\nu}_{\ell} + U_{\ell i} \, C \overline{\boldsymbol{\nu}_{\ell}}^{t} \quad \begin{cases} \ell = \mathbf{e}, \mu, \tau \\ i = 1, 2, 3 \end{cases}$$

$$\mathcal{L}_{\text{Majorana}} = -\frac{1}{2} m_{\nu,\ell\ell'} \ \overline{\boldsymbol{\nu}_{\ell}} C \overline{\boldsymbol{\nu}_{\ell'}}^t + \frac{1}{2} m_{\nu,\ell\ell'}^* \ \boldsymbol{\nu}_{\ell}^t C^{\dagger} \boldsymbol{\nu}_{\ell'}$$

 $m_
u =$ 

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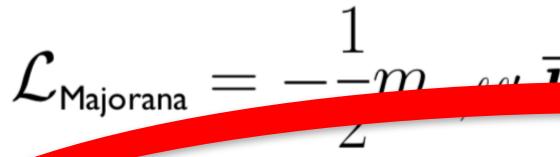
with the complex symmetric mass matrix

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$$\mathcal{L}_{\mathsf{Majorana}}$$
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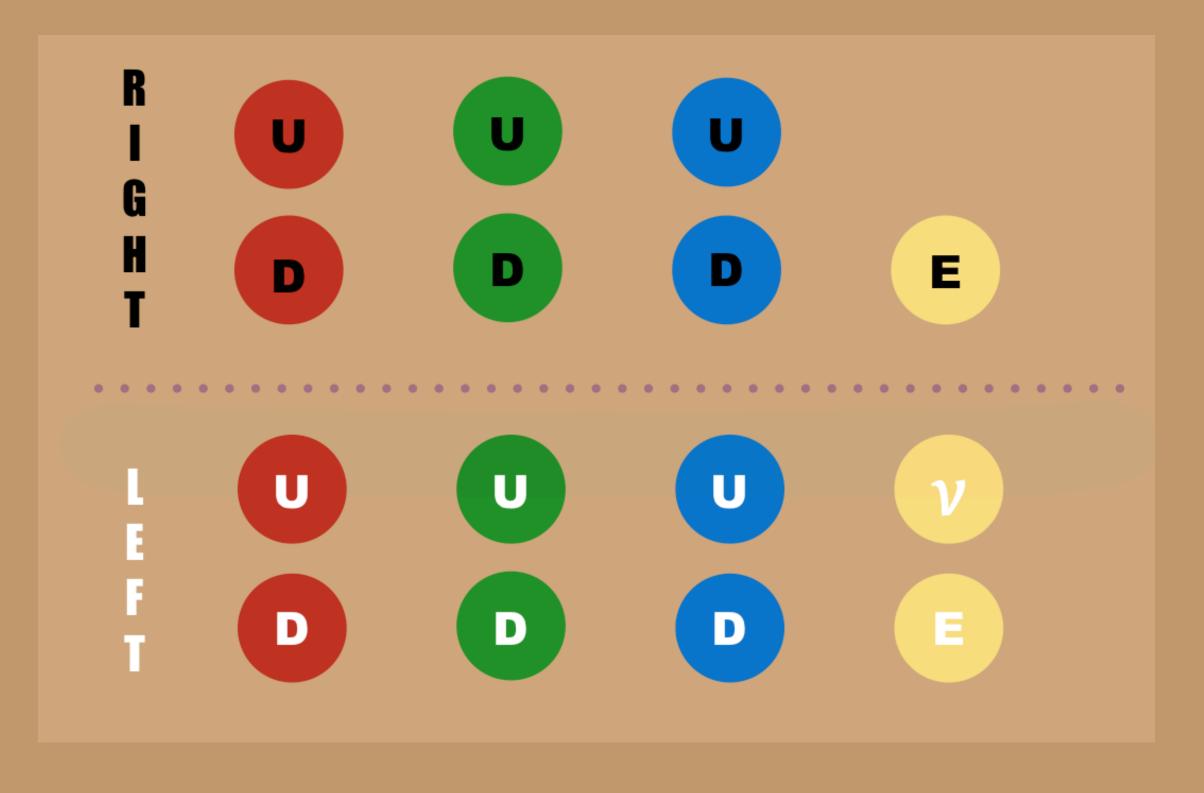
$$\int_{\ell i} C \overline{\boldsymbol{\nu}_{\ell}}^t \left\{ \begin{array}{l} \ell = \mathbf{e}, \mu, \tau \\ i = 1, 2, 3 \end{array} \right.$$

then the Lagrangian density reads,

$$\overline{\boldsymbol{\nu}}_{\ell}^{T} \overline{\boldsymbol{\nu}}_{\ell}^{t} + \frac{1}{2} m_{\ell}^{*} \ell^{\ell} \boldsymbol{\nu}_{\ell}^{t} \mathcal{\nu}_{\ell}^{\dagger} \mathcal{\nu}_{\ell'}$$

$$oldsymbol{U}{\mathsf{d}}{\mathsf{iag}}(m)oldsymbol{U}^t$$

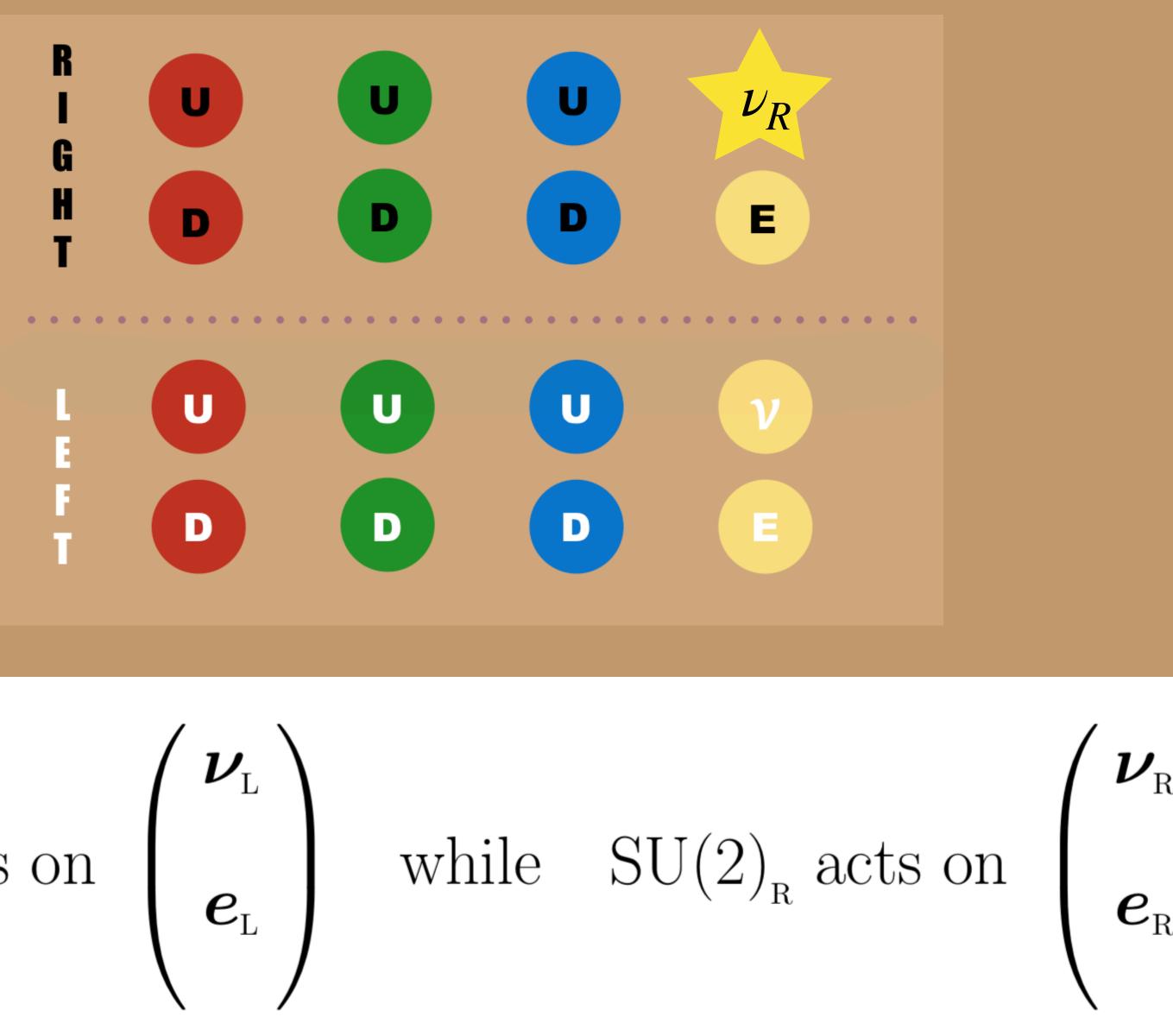
evolution of gauge models

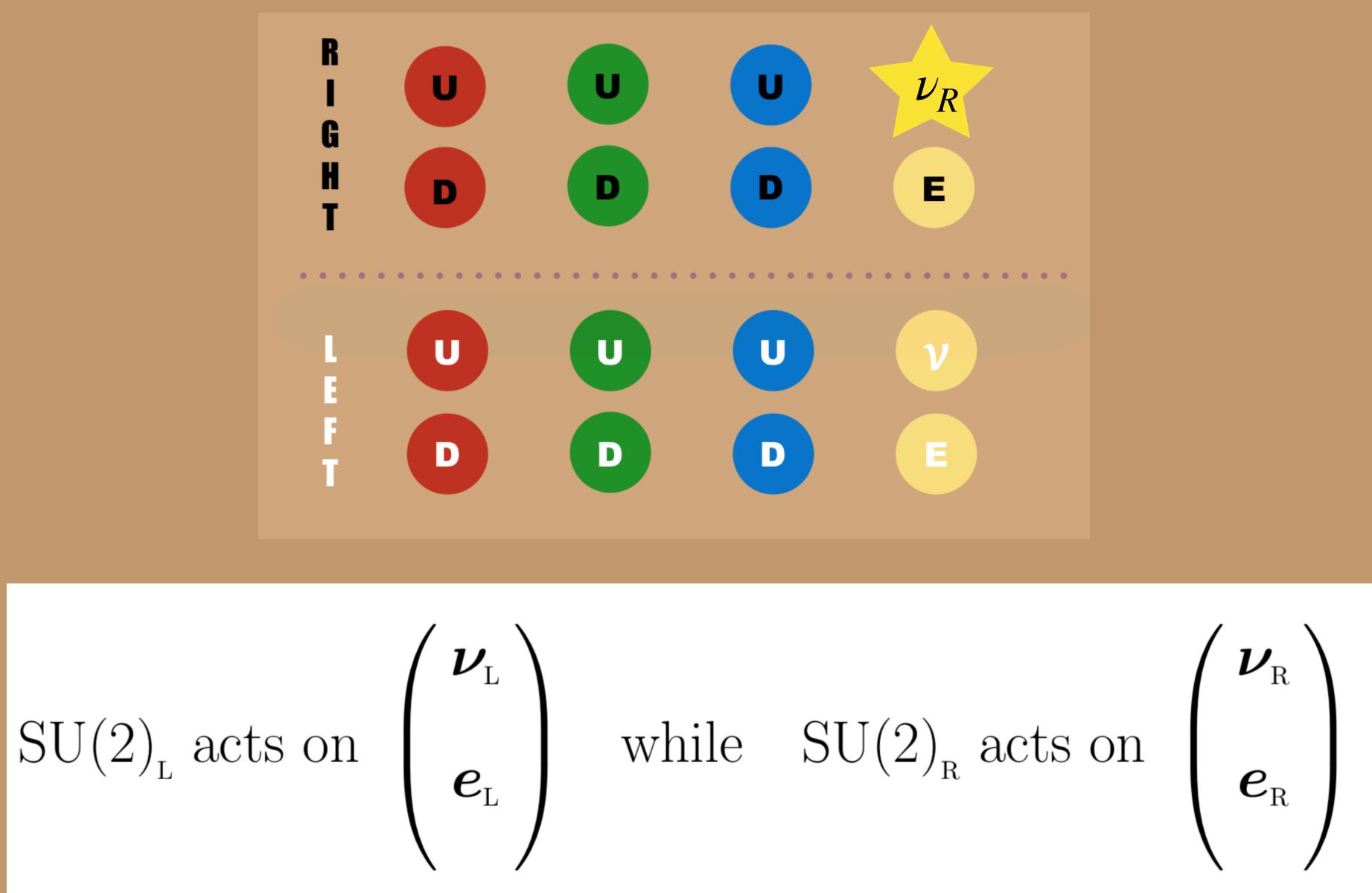


this diagram depicts more accurately which are the particles of the standard model in each family

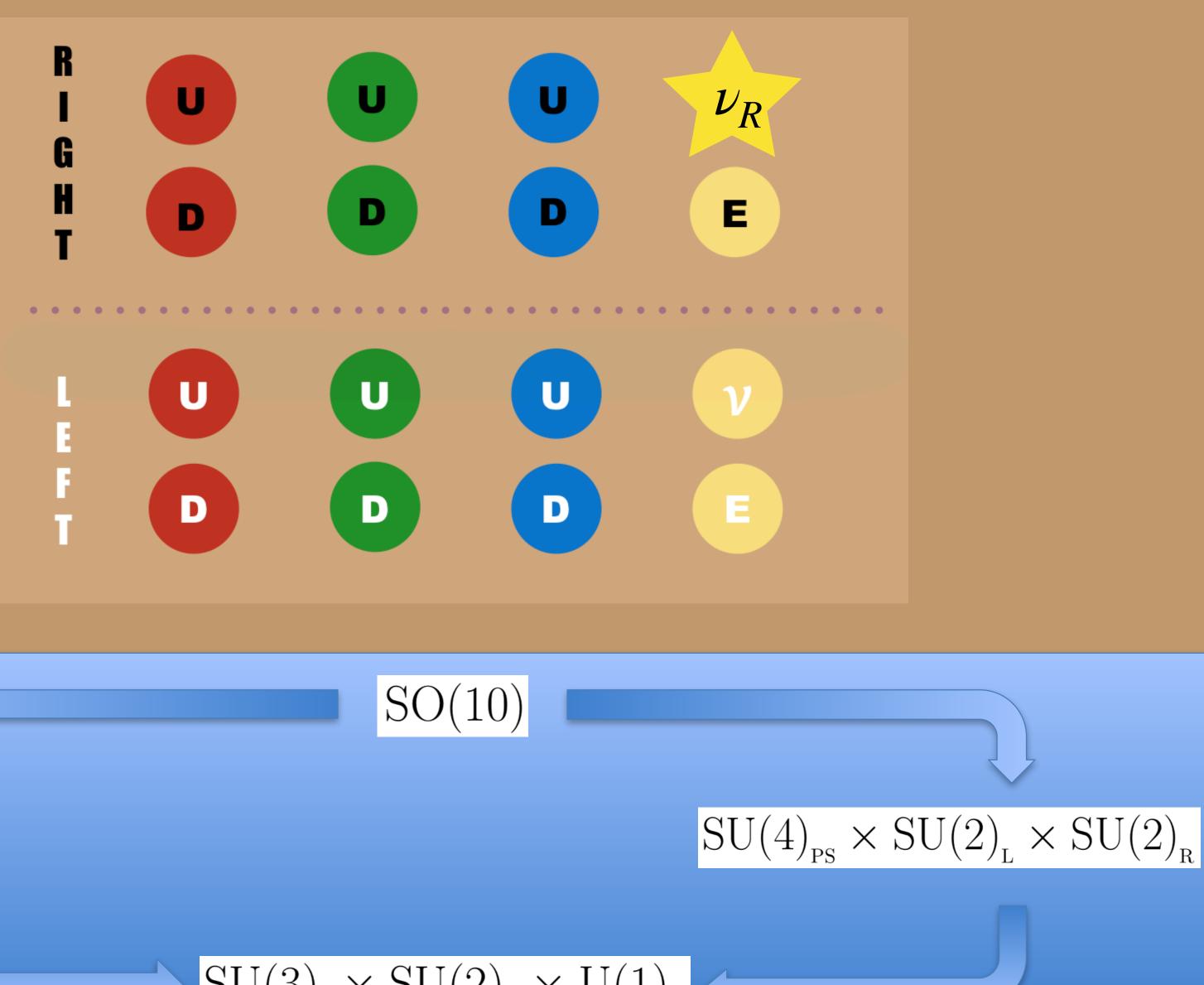
this new representation highlights a significant asymmetry concerning neutrinos

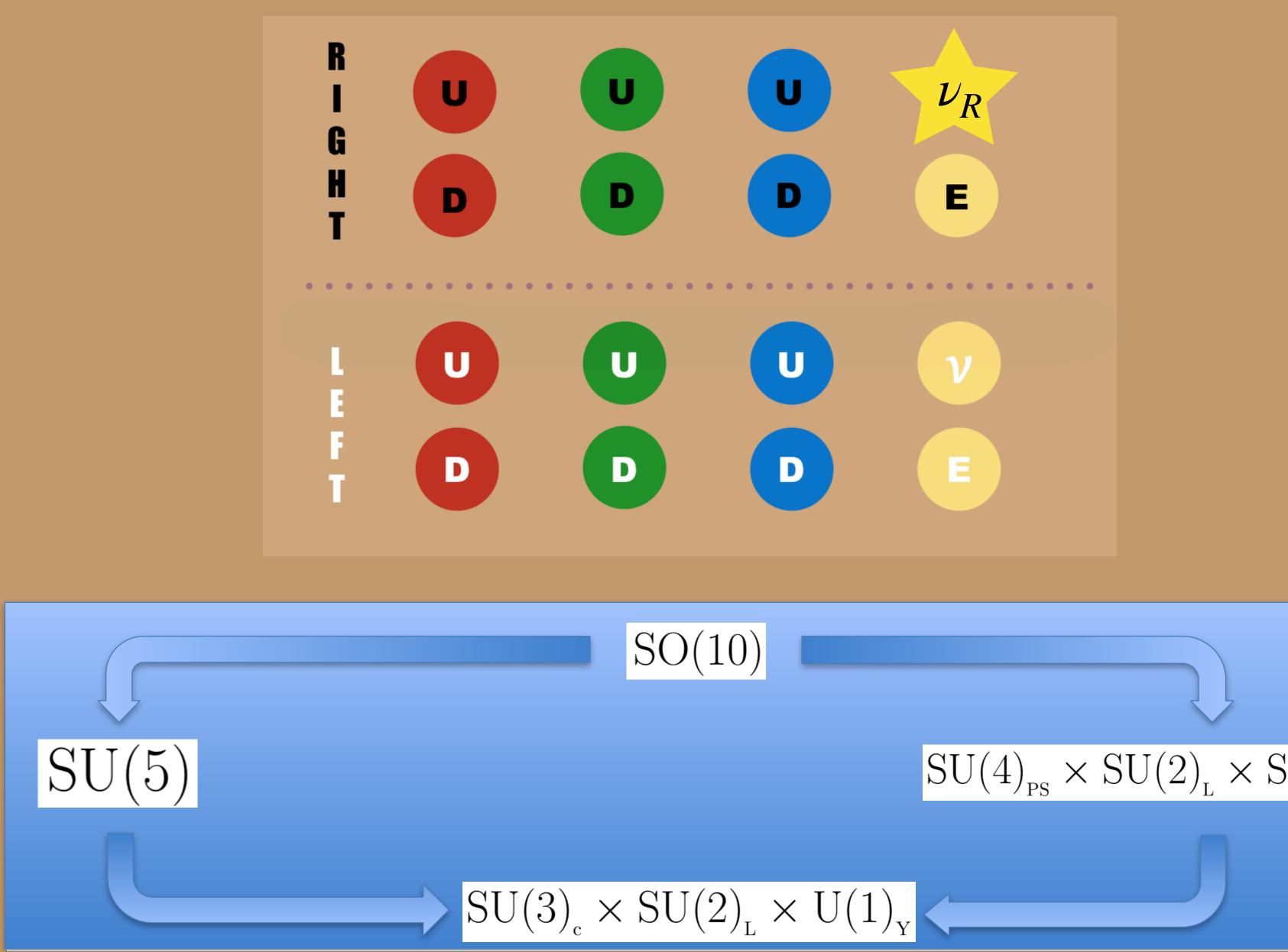
Glashow 1961; Weinberg 1967; Salam 1967; Gross & Wilczek 1973; Politzer 1973





Mohapatra & Pati 1975; Senjanovic & Mohapatra 1975





Georgi & Glashow 1974; Pati & Salam 1974; Georgi 1975; Fritzsch & Minkowski 1975

## on the mass scale of heavy neutrinos

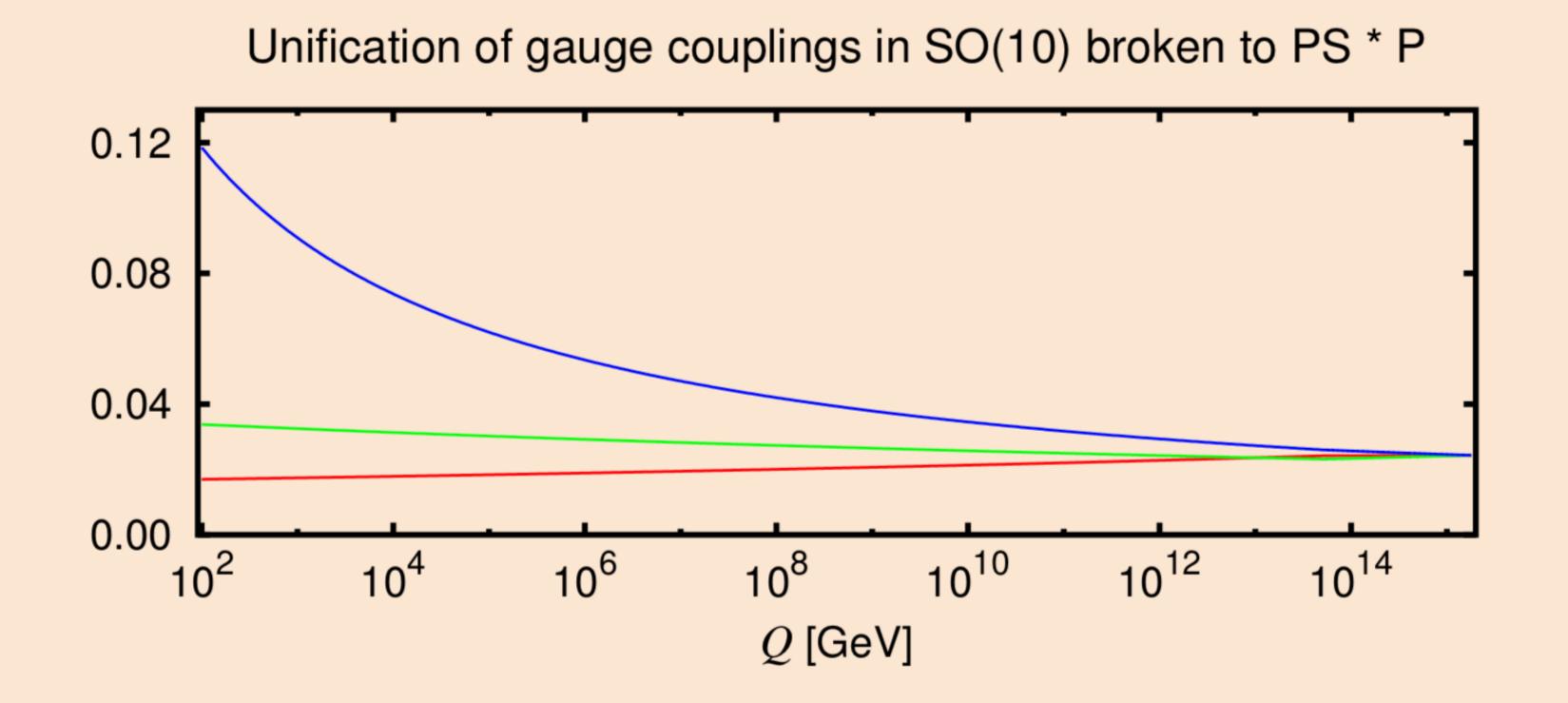
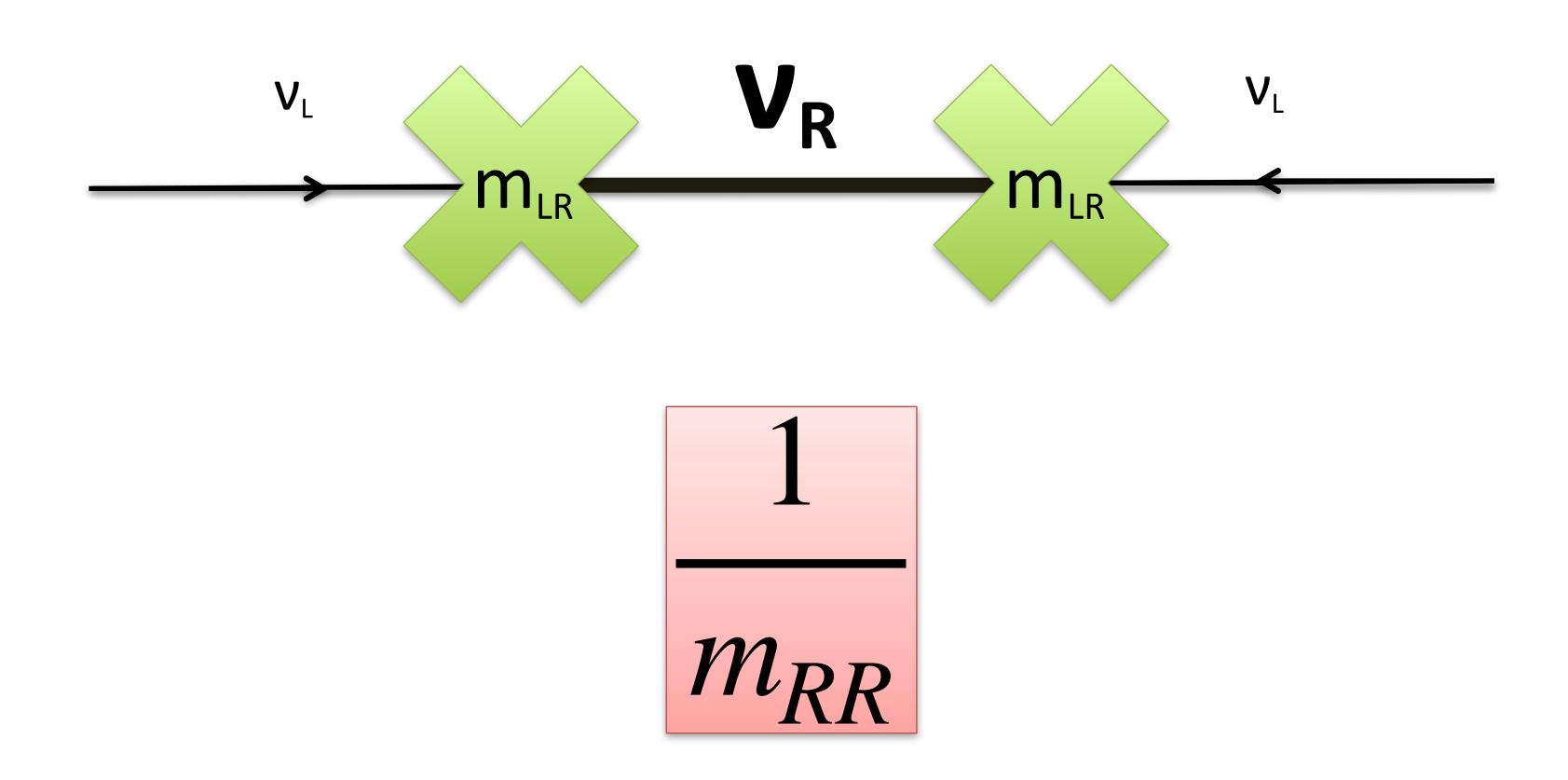


Figure 2: Evolution of the gauge coupling constants in a GUT model with intermediate scale. Here,  $M_{\rm interm.} \approx 5 \times 10^{13}$  GeV.

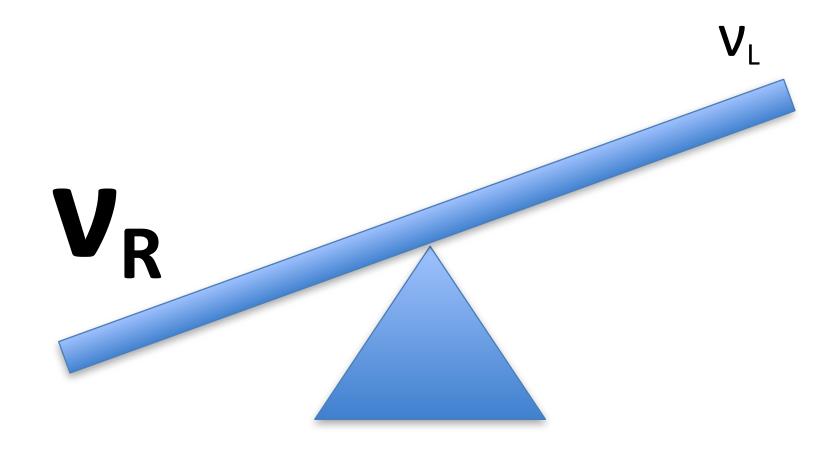
FV at the Cryodet meeting at Gran Sasso lab (2006)



Minkowski 1977; Yanagida 1979; Gell-Mann, Ramond, Slansky 1979; Mohapatra Senjanovic 1980

if the new neutrinos  $\nu_{\rm R}$  are heavy enough, the ordinary ones take on a small mass, just as we observe

# this is called "seesaw"



Minkowski 1977; Yanagida 1979; Gell-Mann, Ramond, Slansky 1979; Mohapatra Senjanovic 1980



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 $oldsymbol{H} = \left(egin{array}{c} H^+ \ H^0 \end{array}
ight)$ 



# dim5 operator...

June 07, 2019

$$oldsymbol{L}_{oldsymbol{\ell}} = \left(egin{array}{c} 
u_\ell \ \ell \end{array}
ight) \;\; ext{with} \; \ell = extbf{e}, \mu, au$$

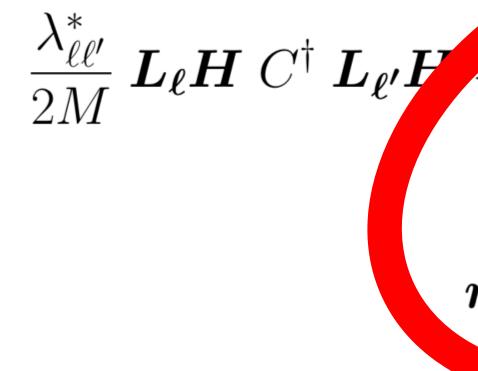
 $oldsymbol{L}oldsymbol{H}=oldsymbol{L}^t\,i\sigma_2\,oldsymbol{H}=oldsymbol{
u}\langle H^0
angle+2$  field

$$= \lambda_{\ell\ell'} \frac{\langle H^0 \rangle^2}{2M} \nu_{\ell} C^{\dagger} \nu_{\ell'} + \text{interactions}$$

$$\boldsymbol{m}_{\boldsymbol{\nu}} = \frac{\left\langle H^0 \right\rangle^2}{2M} \, \boldsymbol{\lambda}$$

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 $oldsymbol{H} = \left( egin{array}{c} H^+ \ H^0 \end{array} 
ight)$ 



# ... yields Majorana mass in SM!

June 07, 2019

$$oldsymbol{L}_{oldsymbol{\ell}} = \left(egin{array}{c} 
u_\ell \ \ell \end{array}
ight) \;\; {
m with} \; \ell = {
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 $\boldsymbol{L}\boldsymbol{H} = \boldsymbol{L}^t \, i\sigma_2 \, \boldsymbol{H} = \boldsymbol{
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## a plausible scenario for baryogenesis (Fukugita-Yanagida's implementation of Sakharov's program)

(1) During big-bang, the decay of heavy (right-handed) neutrinos create  $\Delta L$ 

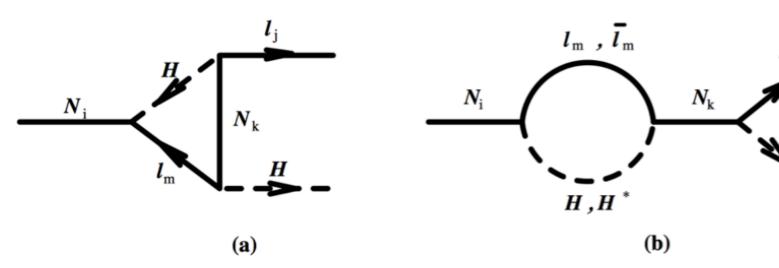


Figure 1: Diagrams contributing to the vertex (Fig. 1a) and wave function (Fig. 1b) CP violation in the heavy singlet neutrino decay.





Covi et al. '96

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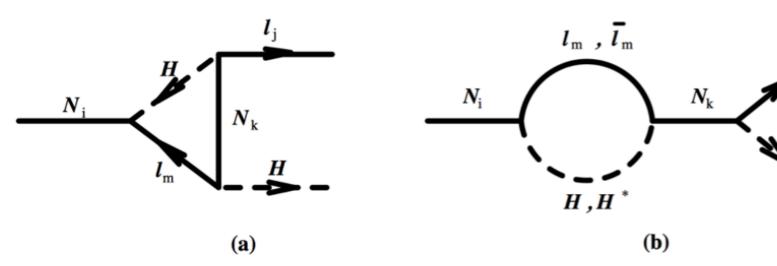
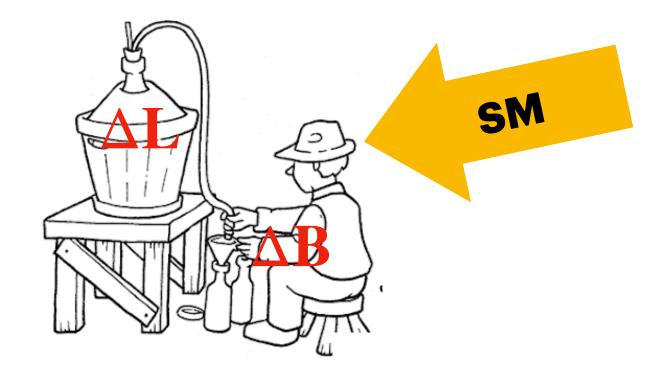


Figure 1: Diagrams contributing to the vertex (Fig. 1a) and wave function (Fig. 1b) CP violation in the heavy singlet neutrino decay.

(2) Subsequently,  $\mathbf{B} + \mathbf{L}$  violating effects convert it into  $\Delta \mathbf{B}$ 

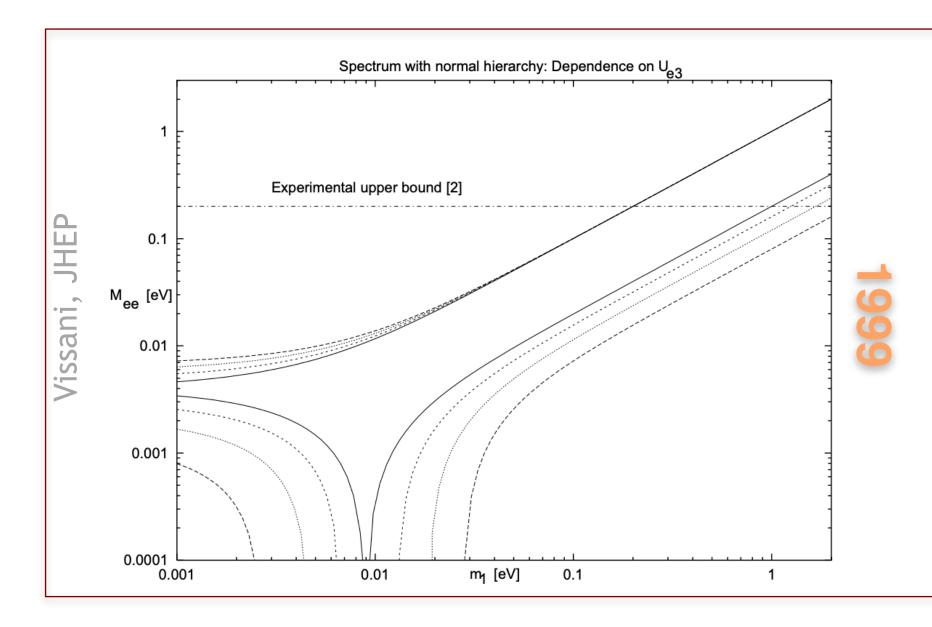


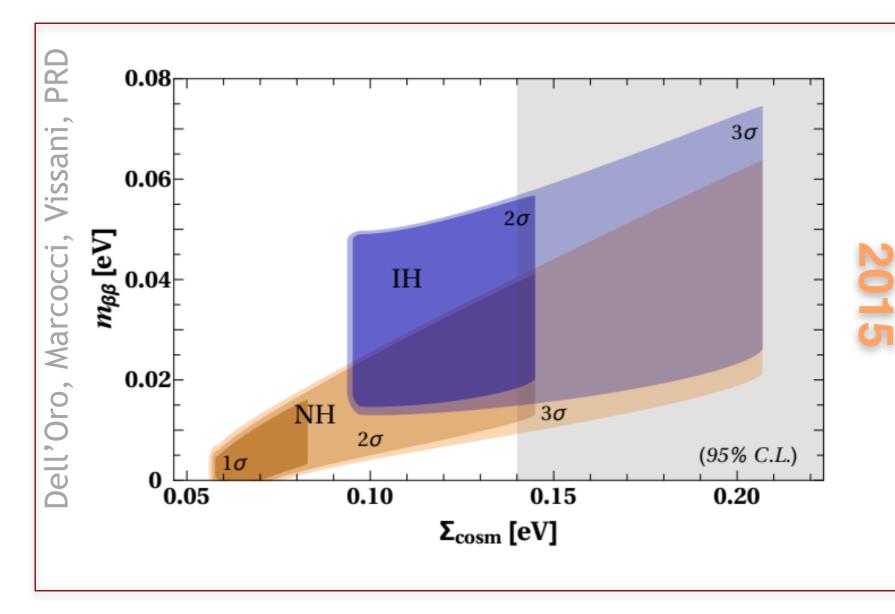
Covi et al. '96

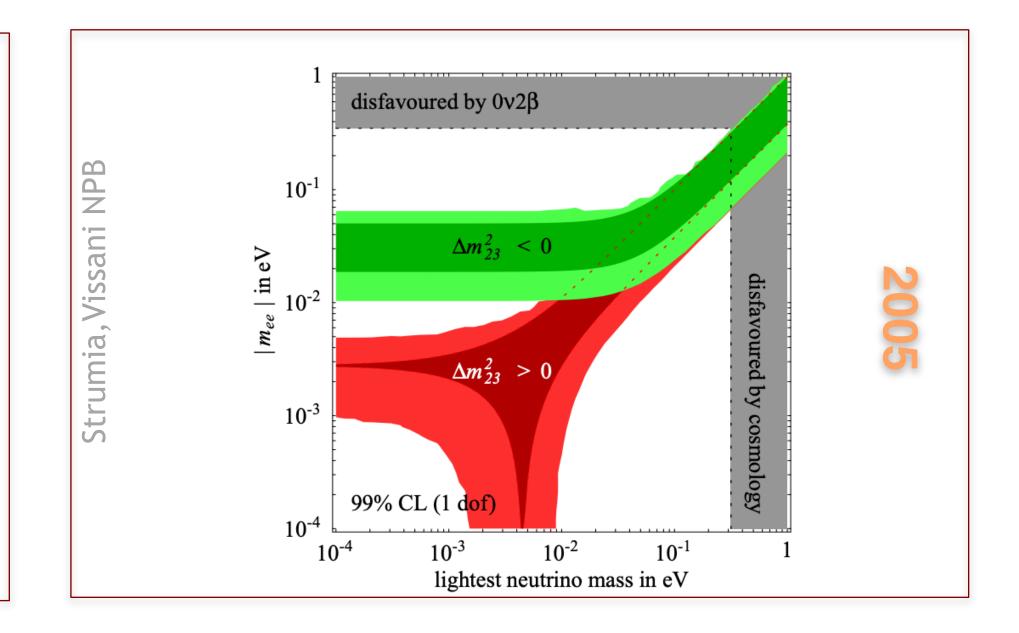


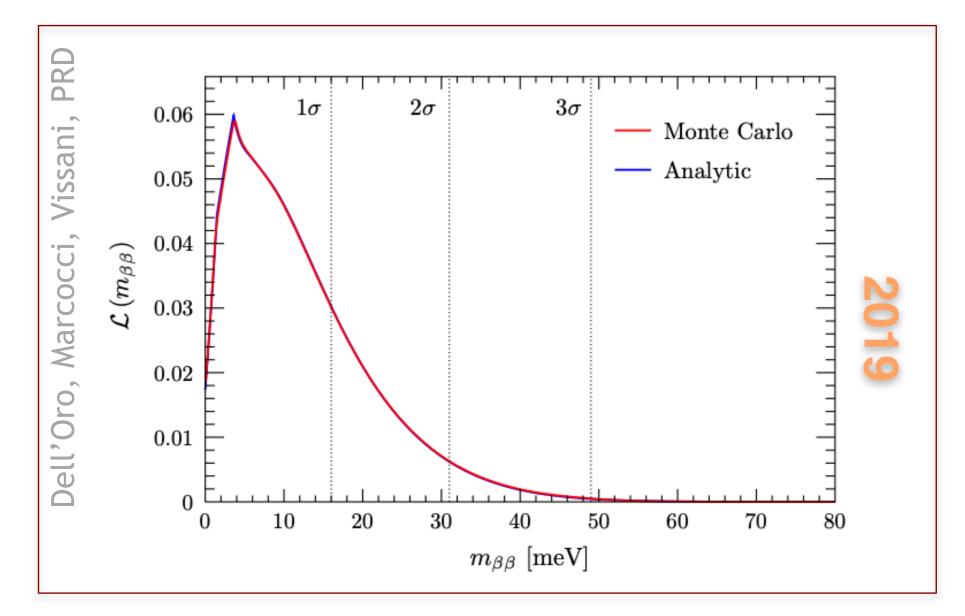
implications

## constraints on the Majorana mass relevant to $2n \rightarrow 2p+2e$









#### Testing the Inverted Neutrino Mass Ordering with Neutrinoless Double-Beta Decay

Matteo Agostini,<sup>1,\*</sup> Giovanni Benato,<sup>2,†</sup> Jason A. Detwiler,<sup>3,‡</sup> Javier Menéndez,<sup>4,§</sup> and Francesco Vissani<sup>2,5,¶</sup>

<sup>1</sup>Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, UK <sup>2</sup>INFN, Laboratori Nazionali del Gran Sasso, 67100 Assergi, L'Aquila, Italy <sup>3</sup>Center for Experimental Nuclear Physics and Astrophysics, and Department of Physics, University of Washington, Seattle, WA 98115 - USA <sup>4</sup>Department of Quantum Physics and Astrophysics and Institute of Cosmos Sciences, University of Barcelona, 08028 Barcelona, Spain <sup>5</sup>Gran Sasso Science Institute, 67100 L'Aquila, Italy

We quantify the extent to which future experiments will test the existence of neutrinoless doublebeta decay mediated by light neutrinos with inverted-ordered masses. While it remains difficult to compare measurements performed with different isotopes, we find that future searches will fully test the inverted ordering scenario, as a global, multi-isotope endeavor. They will also test other possible mechanisms driving the decay, including a large uncharted region of the allowed parameter space assuming that neutrino masses follow the normal ordering.

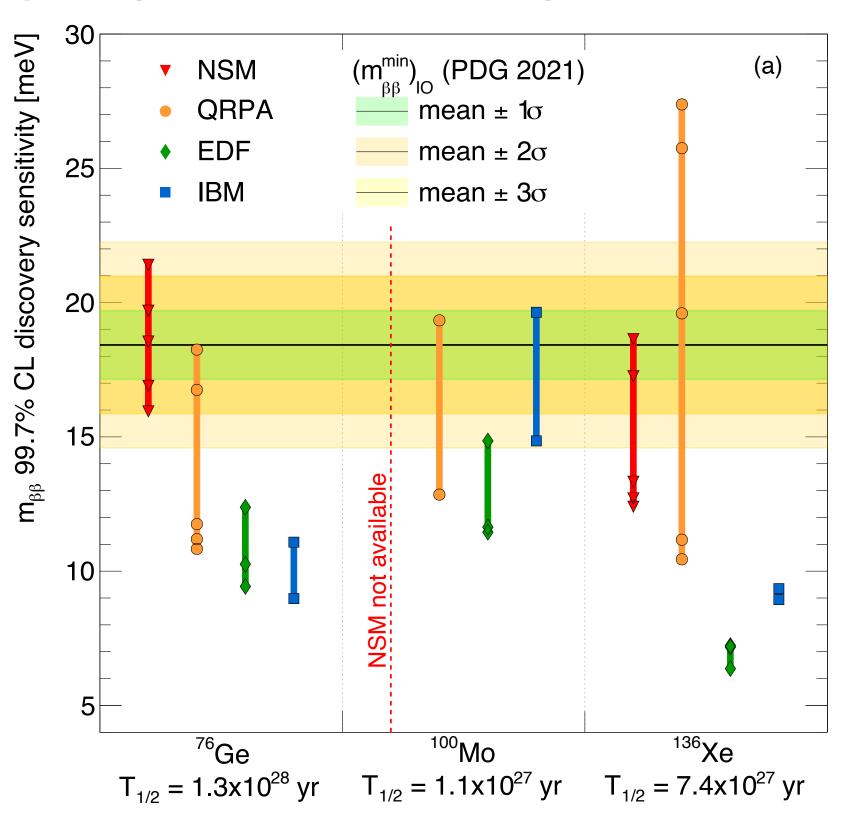


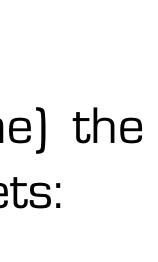
FIG. 1. Comparison of  $m_{\beta\beta}$  99.7%-CL discovery and 90%-CL median exclusion sensitivities for different isotopes at stated halflife sensitivities [30–32], grouped by nuclear many-body frameworks with matrix element ranges from Table I. The horizontal bands show the variation on  $(m_{\beta\beta}^{min})_{IO}$  under variation of the neutrino oscillation parameters.

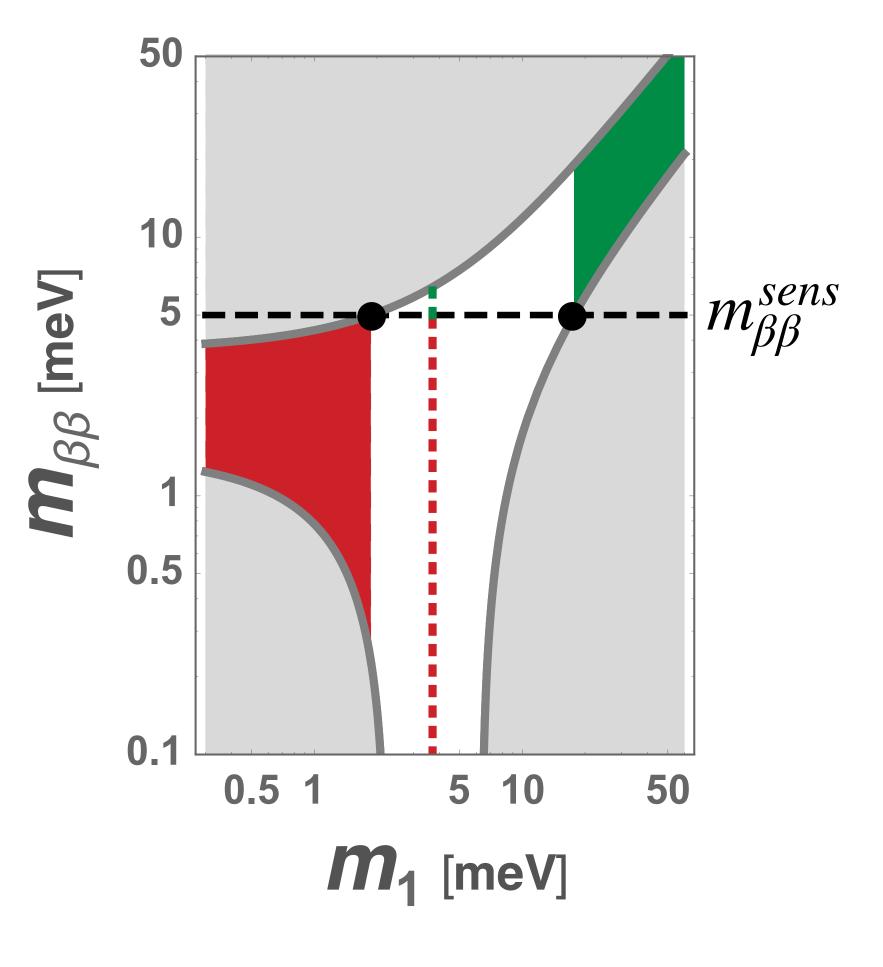
(Dated: July 21, 2021)

# Majorana neutríno masses

Given a sensitivity  $m_{\beta\beta}^{sens}$  (horizontal line) the possible true values of  $m_1$  fall in 3 subsets:

- 1.
- the true value  $m_{\beta\beta}(m_1) < m_{\beta\beta}^{sens}$  [red] it is more,  $m_{\beta\beta}(m_1) > m_{\beta\beta}^{sens}$  [green], 2.
- it depends upon Majorana phases (white) З.



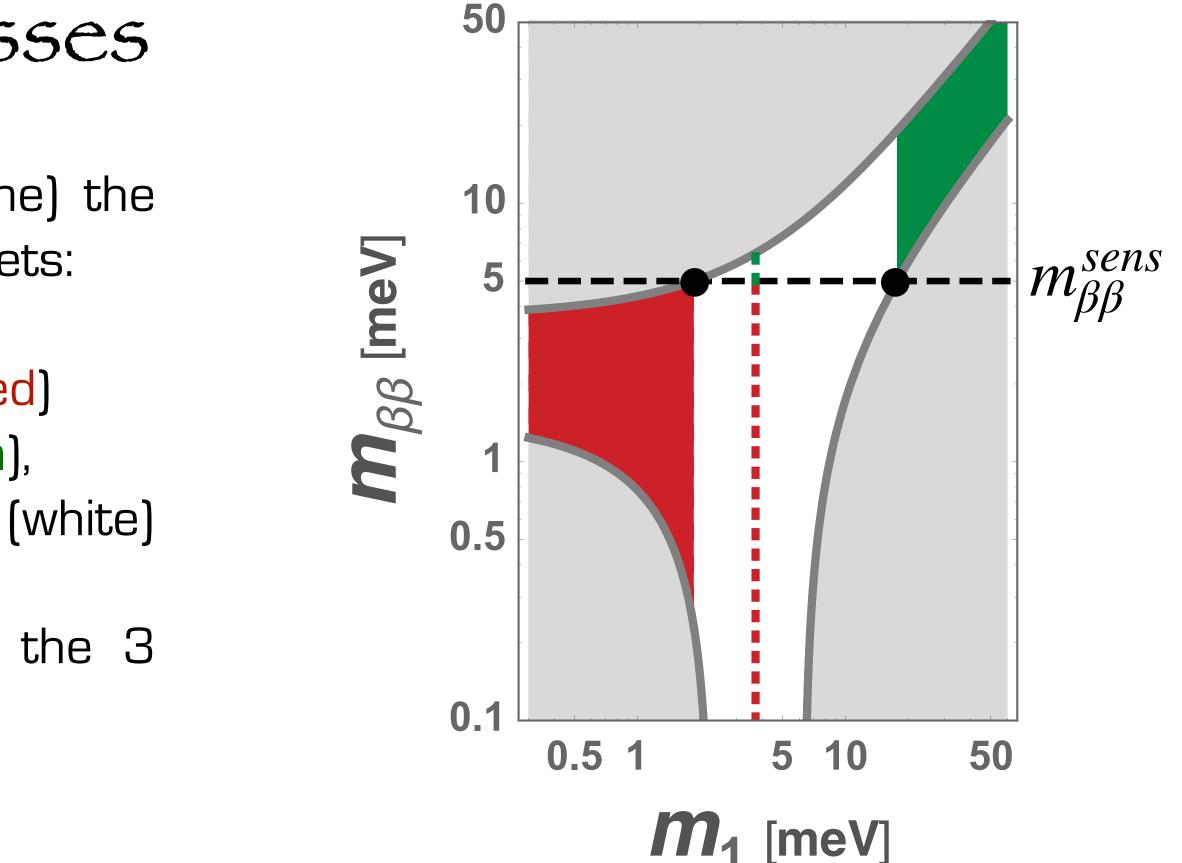


# Majorana neutríno masses

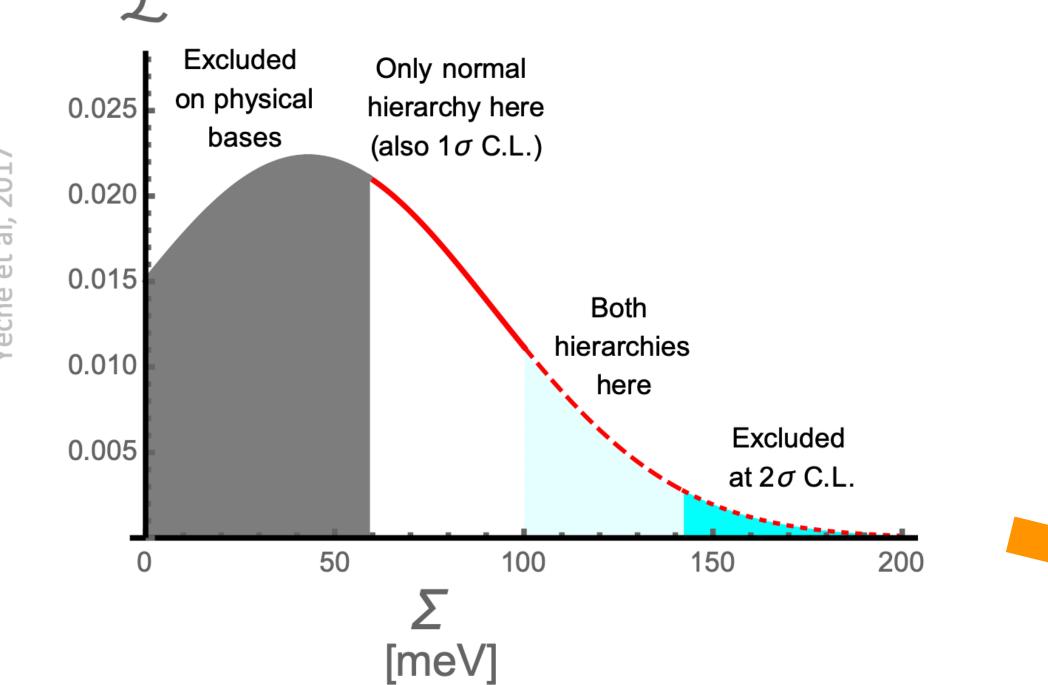
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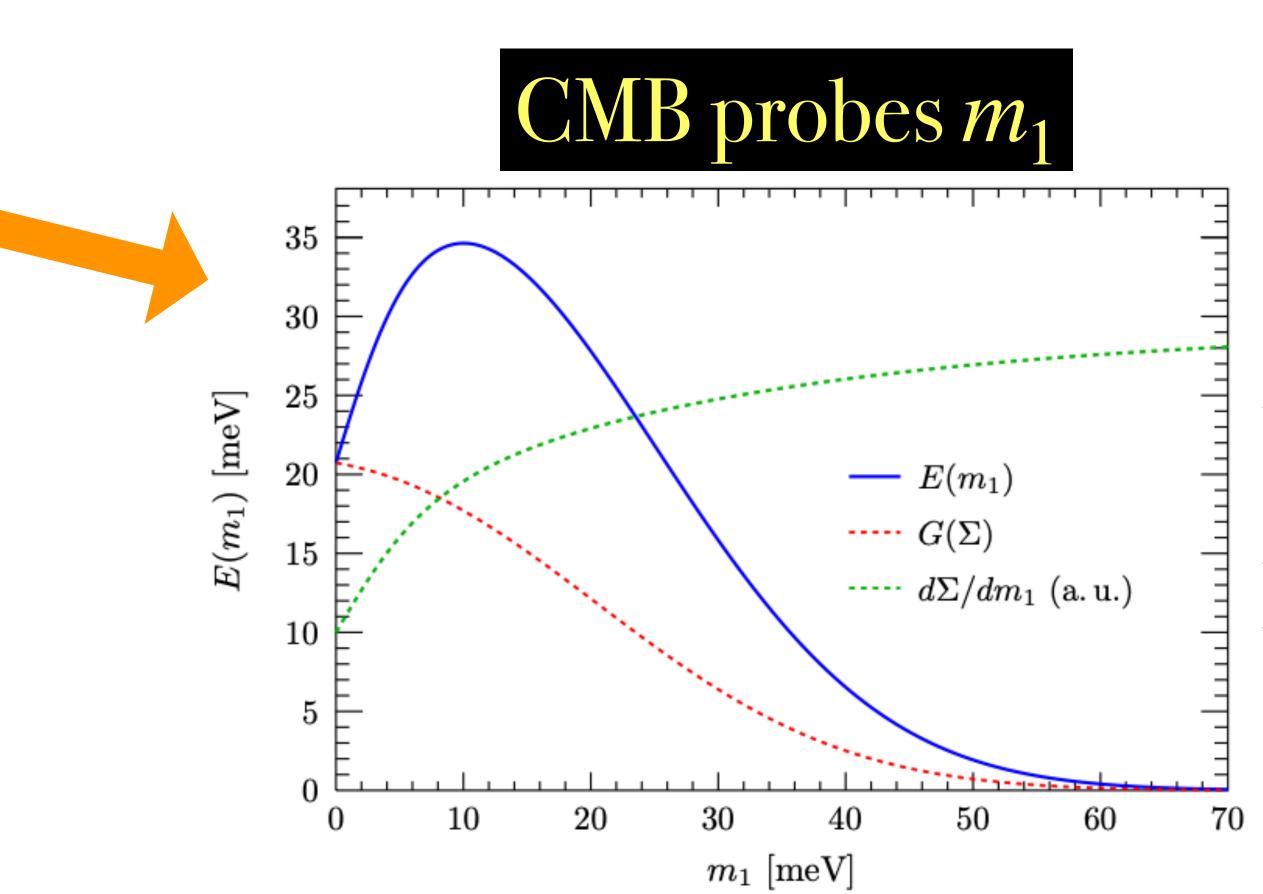
By measuring  $m_1$  the probabilities of the 3 subsets can be quantified



CMB probes  $\Sigma = m_1 + m_2 + m_3$ 



Yeche et al, 2017

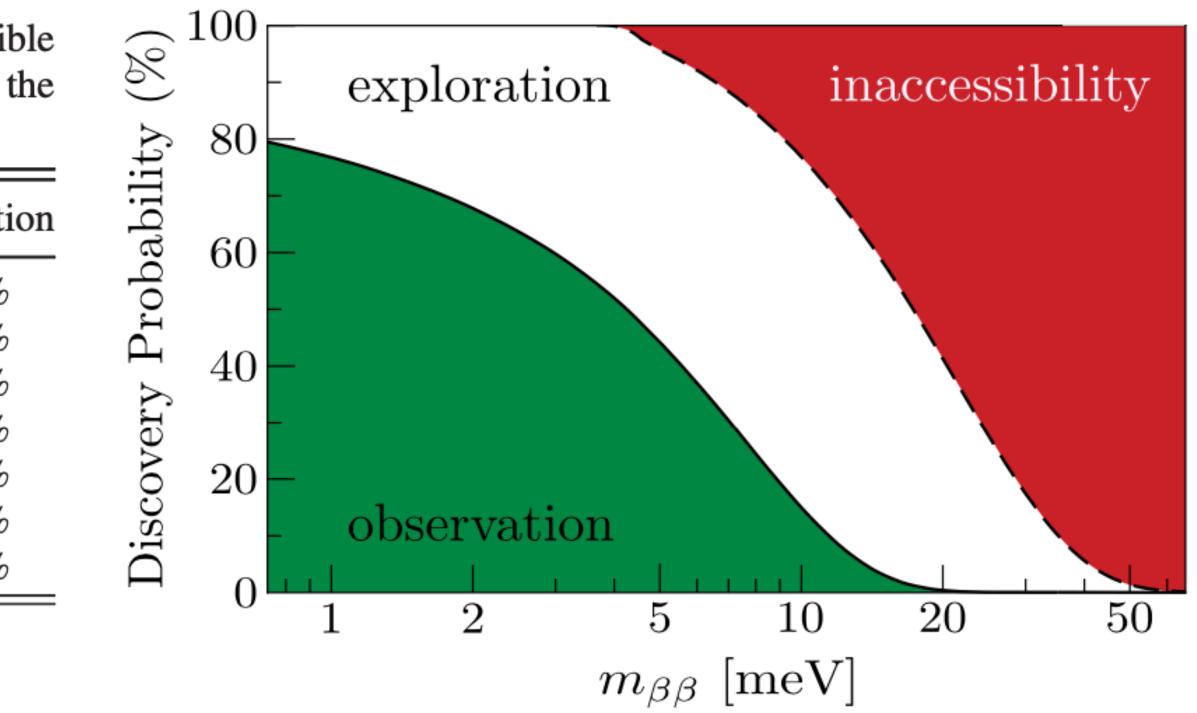




# the three probabilities & "flag viewgraph"

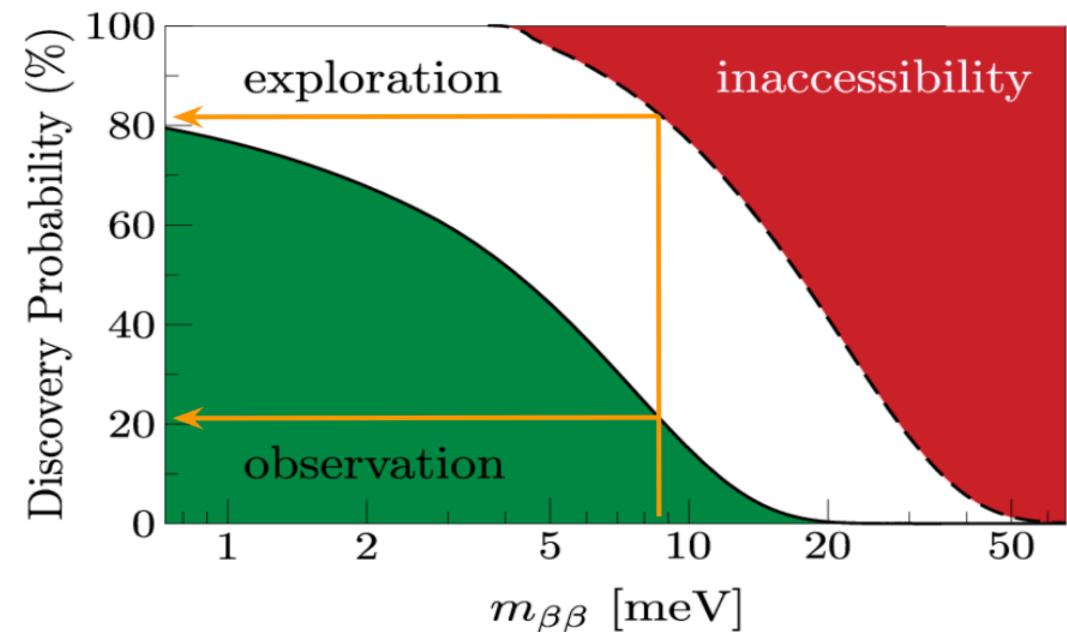
TABLE II. Probability (expressed in percent) of the possible outcomes of a search for  $0\nu\beta\beta$  given as a function of the experimental sensitivity  $m_{\beta\beta}^*$ .

$m^*_{\beta\beta}$ [meV]	Inaccess.	Exploration	Observati
50	98.7%	1.3%	0.0%
20	58.6%	41.1%	0.3%
15	41.9%	55.1%	3.0%
10	23.1%	62.0%	14.9%
5	4.4%	51.4%	44.2%
2	0.0%	32.3%	67.7%
0	0.0%	12.4%	87.6%



100% for inverted ordering;

 $m_{\beta\beta} = \sqrt{\Delta m_{12}^2} = 8.6 \text{ meV} \text{ is achieved}$ 

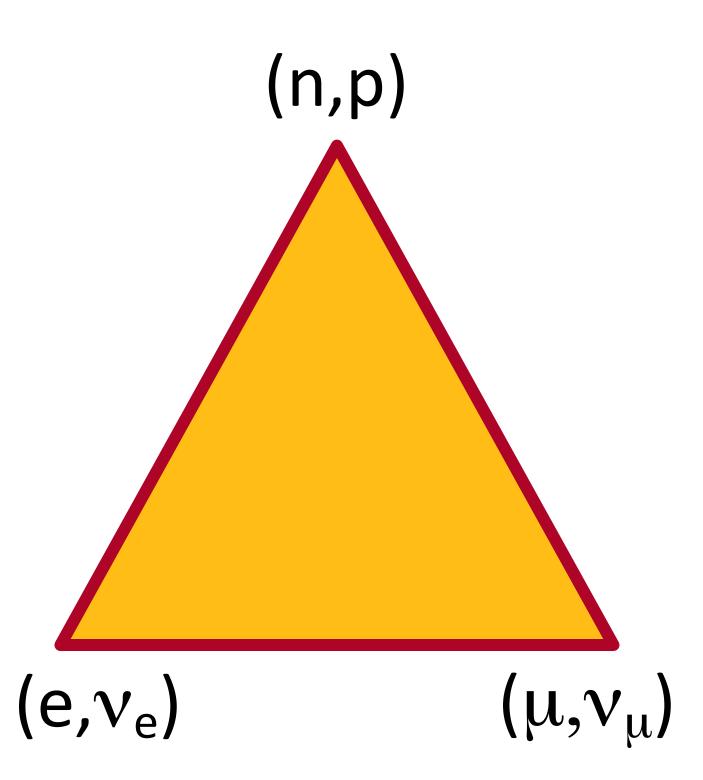


- current estimation of discovery probability:
- between 20% an 80% for normal ordering, if

# end of the second lecture

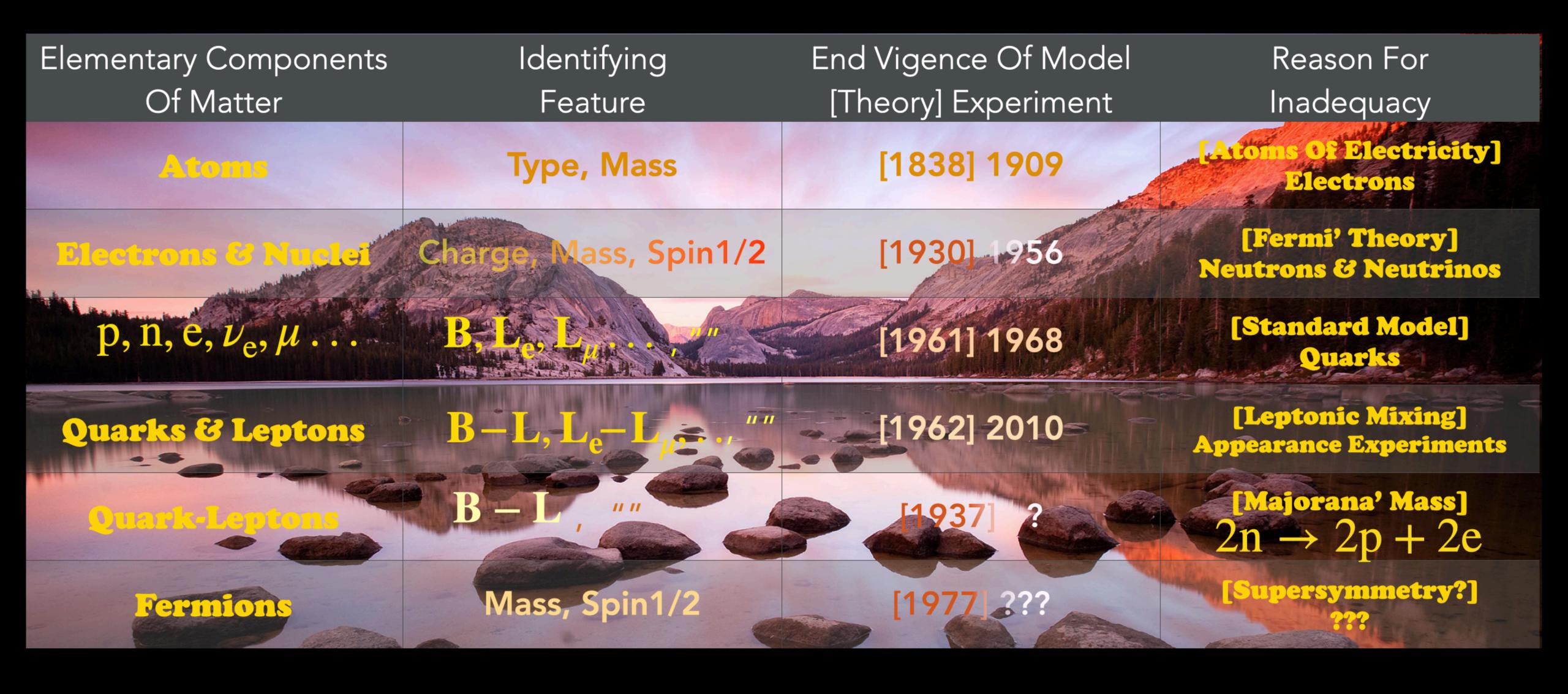
# Other types of neutrinos

- Pontecorvo studied e-capture and  $\mu$ -capture and suggested that the **coupling is same** (1947)
- Puppi argued that the muon is associated to a new type of neutrino, as depicted in the triangle (1948)
- This indicates **universality**: weak interactions treat in the same way the various pairs of particles
- (For the pair (n,p), this is almost true)



Puppi's triangle

# WHAT IS "MATTER" MADE OF?



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# SM effective operators (Weinberg; Wilczek & Zee 79)

accept in the SM Lagrangian density also the operators with canonical dimension >4 that conserve gauge symmetry e.g.

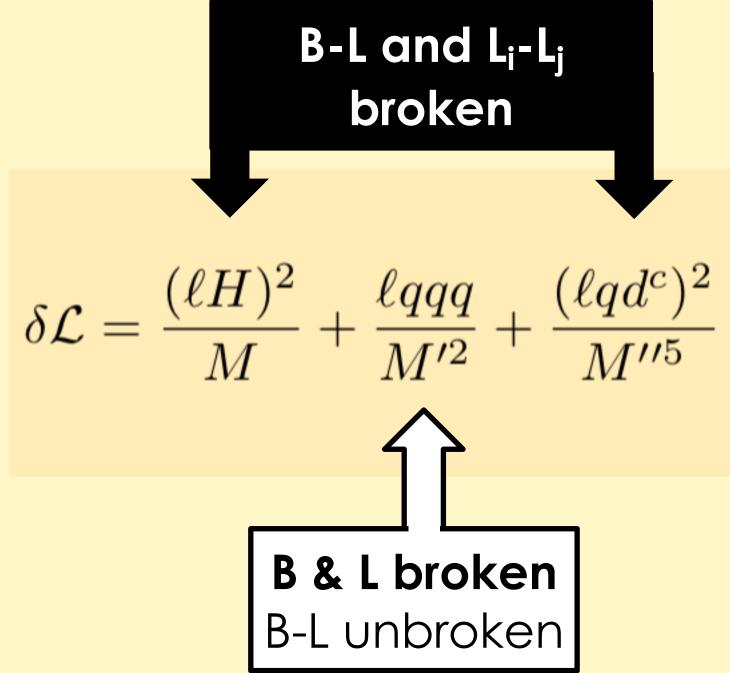
 $\delta \mathcal{L} = \frac{(\ell H)}{M}$ 

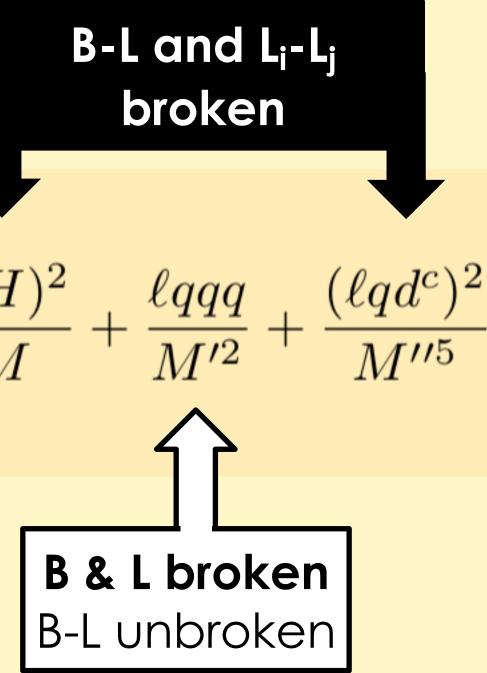
June 07, 2019

$$\frac{(\ell q q q)^2}{M'^2} + \frac{(\ell q q q)^2}{M''^5}$$

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# SM effective operators (Weinberg; Wilczek & Zee 79)

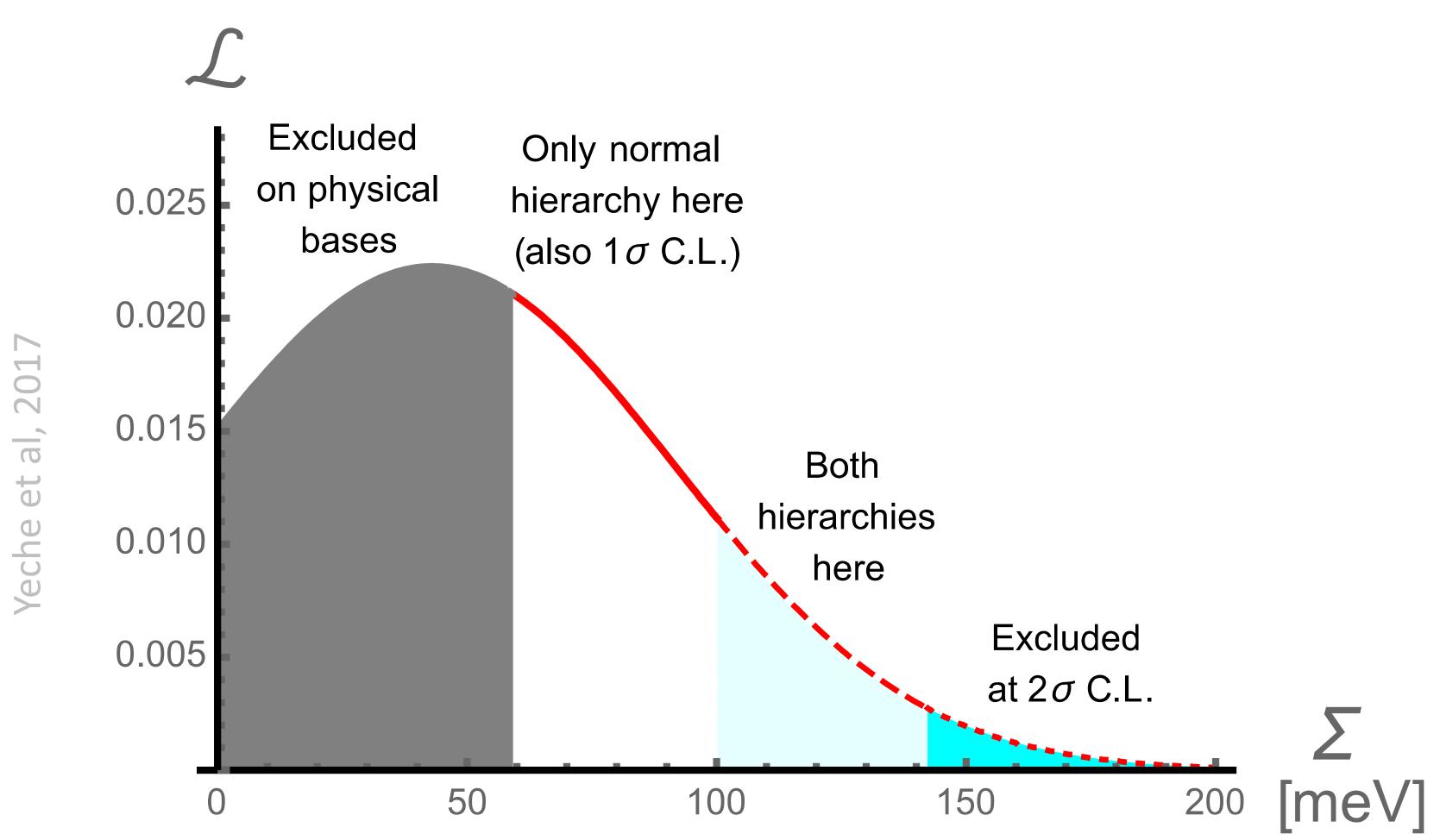




June 07, 2019

## accept in the SM Lagrangian density also the operators with canonical dimension >4 that conserve gauge symmetry e.g.

## CMB is sensitive to $\Sigma = m_1 + m_2 + m_3$



### Discovery probabilities of Majorana neutrinos based o cosmological data

M. Agostini<sup>(D)</sup>,<sup>1,2,\*</sup> G. Benato<sup>(D)</sup>,<sup>3,†</sup> S. Dell'Oro<sup>(D)</sup>,<sup>4,5,‡</sup> S. Pirro<sup>(D)</sup>,<sup>6,§</sup> and T. VISSAR <sup>1</sup>Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, United Kingdom since <sup>2</sup>Physik-Department, Technische Universität München, 85748 Garching, G <sup>3</sup>INFN, Laboratori Nazionali del Gran Sasso, 67100 Assergi, L'Ac **Planck 2015** <sup>4</sup>INFN Sezione di Milano–Bicocca, 20126 Milano, <sup>5</sup>University of Milano–Bicocca, 20126 Milano findings, this is the most <sup>6</sup>INFN, Laboratori Nazionali del Gran Sasso, 67100 As 'Gran Sasso Science Institute, 67100 L'A sensitive probe of (Received 5 January 2021; accepted 5 February 2021 absolute neutrino masses. and the best chance of measuring them in the future

We discuss the impact of the cosmological measurements on the p neutrinos, the parameter probed by neutrinoless double-beta decay assumptions, we quantify the probabilities of discovering neutrinoless new graphical representation that could be of interest for the community DOI: 10.1103/PhysRevD.103.033008

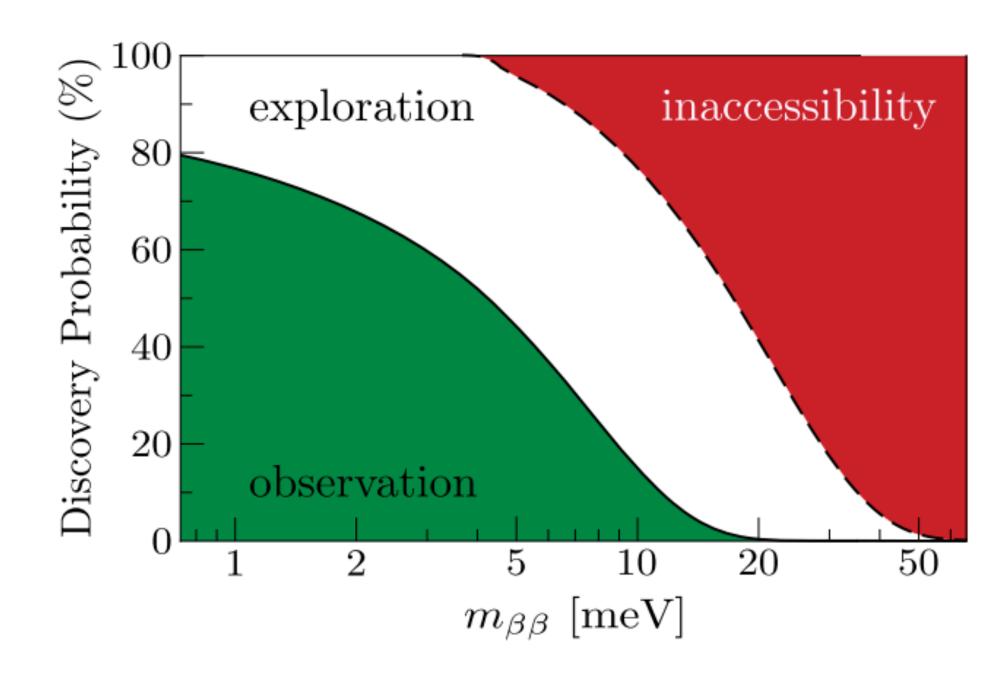


FIG. 2. Discovery probability as a function of the experimental sensitivities to  $m_{\beta\beta}$  for the most unfavorable scenario (black solid line,  $m_{\beta\beta}^{\min}$ ) and the most favorable one (black dashed line,  $m_{\beta\beta}^{\max}$ ). The colored areas express the probability for the three possible outcomes of an experiment: observing a signal even in the worst case scenario (green, observation), not observing a signal even in the best case scenario (red, inaccessibility), and when observing a signal depends on the value of the Majorana phases (white, exploration).



# references

## Neutrino masses and mixings and...

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## Abstract

We review experimental and theoretical results related to neutrino physics with emphasis on neutrino masses and mixings, and outline possible lines of development.

<u>hep-ph/0606054</u> [hep-ph]

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Open Access Review

## What Is Matter According to Particle Physics, and Why Try to Observe Its Creation in a Lab?

by <mark>용 Francesco Vissani</mark> <sup>1,2</sup> 🖂 🝺

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# Special Issue Neutrino Masses and Oscillations 2015 View this Special Issue Review Article | Open Access Volume 2016 | Article ID 2162659 | https://doi.org/10.1155/2016/2162659 Show citation Neutrinoless Double Beta Decay: 2015 Review

Stefano Dell'Oro,<sup>1</sup> Simone Marcocci,<sup>1</sup> Matteo Viel,<sup>2,3</sup> and **Francesco Vissani**<sup>1,4</sup> Show more

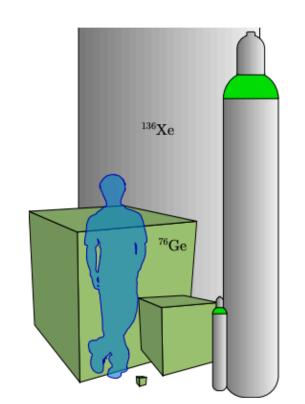


FIG. 18. Masses corresponding to present, mega and ultimate exposures, assuming zero background condition and 5 years of data acquisition. The cubes represent the amount of  $^{76}$ Ge, the (150 bar) bottles the one of  $^{136}$ Xe. The smallest masses depict the present exposure, while the biggest bottle is out of scale.

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### Toward the discovery of matter creation with neutrinoless double-beta decay

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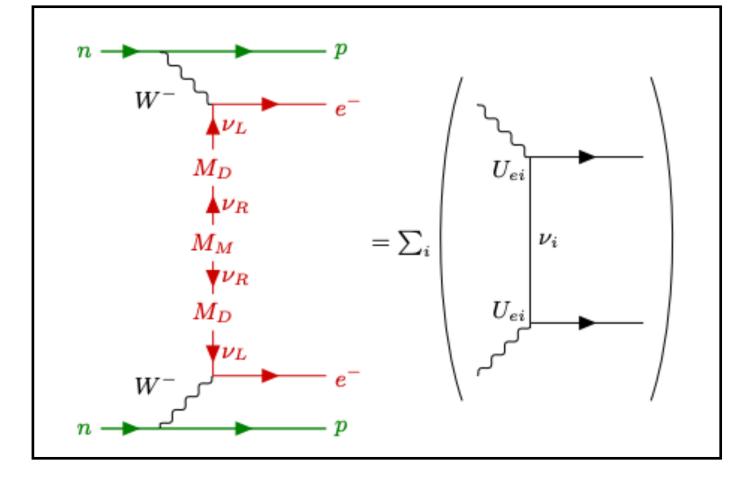
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#### I. INTRODUCTION

In all physical processes observed so far the creation or destruction of matter is compensated by that of antimatter. However, our universe contains an abundance of matter, a fact to which we owe our very existence. In various theories the balance of matter and antimatter can be broken, accounting for this asymmetry of our universe. At present, the most promising avenues for the detection in the laboratory of processes that alter the abundance of matter are proton decay, altering the number of baryons, and electron creation, altering the number of leptons.

The quest to observe electron creation is being pursued vigorously in the form of searches for a nuclear decay where the atomic number Z increases by two units while the nucleon number A remains constant:  $(A, Z) \rightarrow$ (A, Z+2)+2e. This is commonly known as "neutrinoless"  $\beta\beta$  decay" ( $0\nu\beta\beta$  decay). Here, the creation of electrons can be enabled by the "transmutation" of neutrinos into antineutrinos, which is only possible if the neutrino has a peculiar type of mass, named after Majorana. Thus the matter-antimatter imbalance and neutrino masses could have a common origin.

A symmetry between neutrinos and antineutrinos was postulated by Majorana and further discussed by Racah in 1937. This led Furry to propose the existence of  $0\nu\beta\beta$  decay in 1939, building on Goeppert Mayer's ideas on  $\beta\beta$  transitions. Pioneering searches for  $0\nu\beta\beta$  decay started in the 40s using time-coincidence counting techniques or visual detection of tracks in cloud chambers and photographic emulsions. Since then, experiments have



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