

Markarian 501
Markarian 421
Geming
Milky Way
Messages from very

Libra
high energy γ rays

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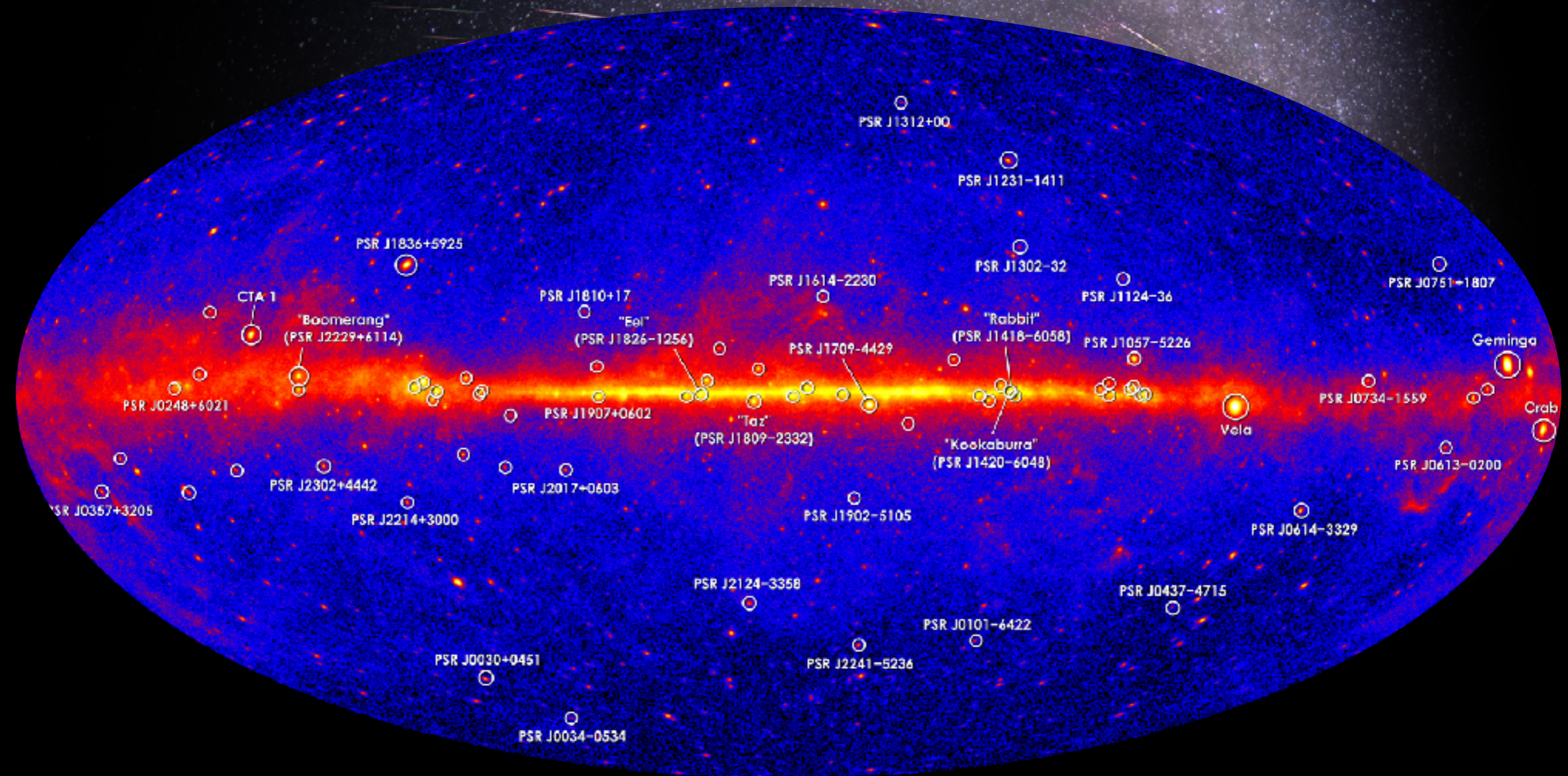
Outline

- ▶ Introduction and Scientific Motivation
- ▶ Biased selection of recent results:
 - ▶ Galactic Pevatrons
 - ▶ New TeV sources
 - ▶ Jets of a microquasar
 - ▶ Gamma-ray bursts
 - ▶ Flaring quasars
- ▶ Conclusions & Outlook

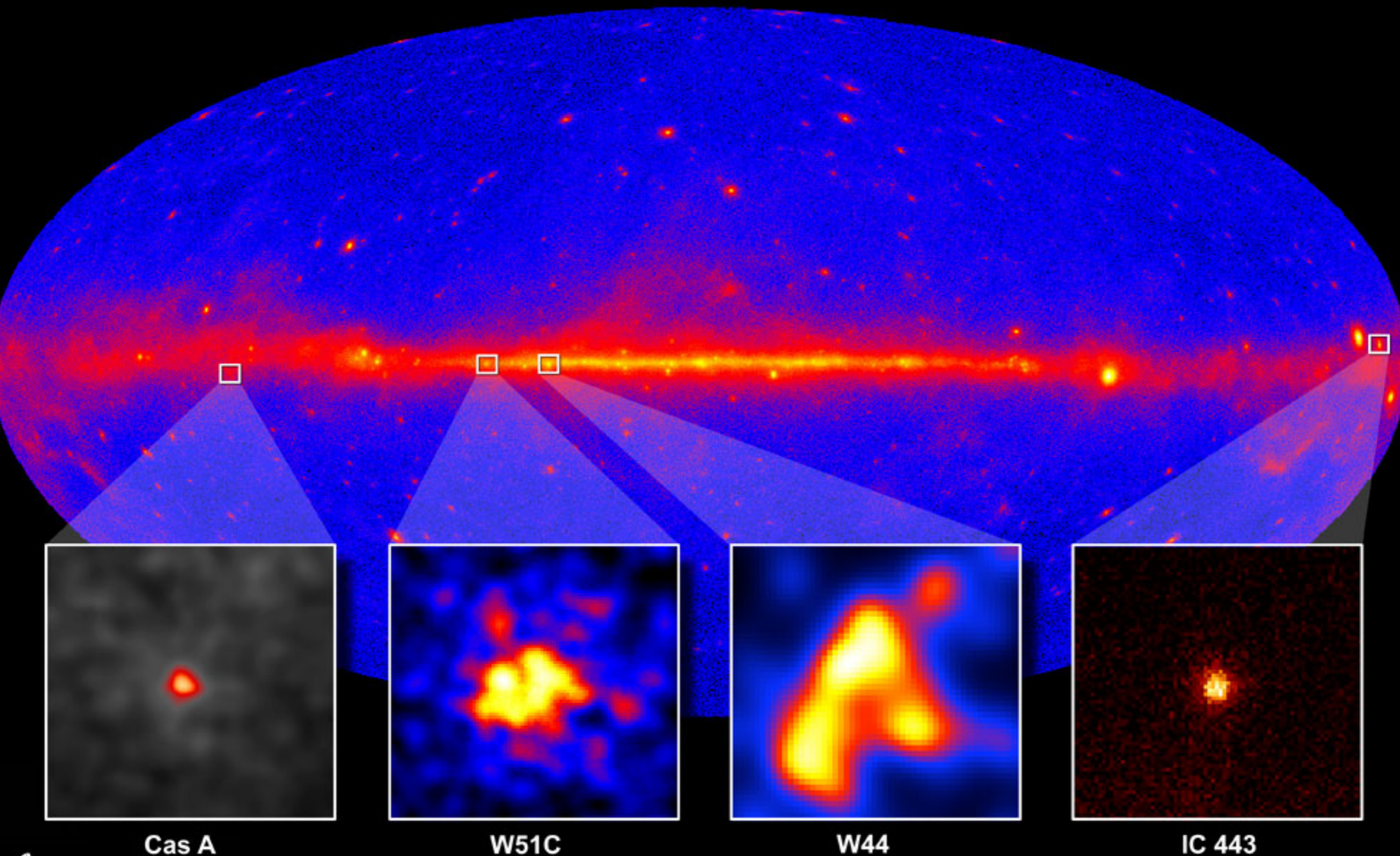
Scientific Motivation

What are the sources of the highest energy particles in the Universe?

Gamma rays



NASA's Fermi telescope resolves supernova remnants at GeV energies



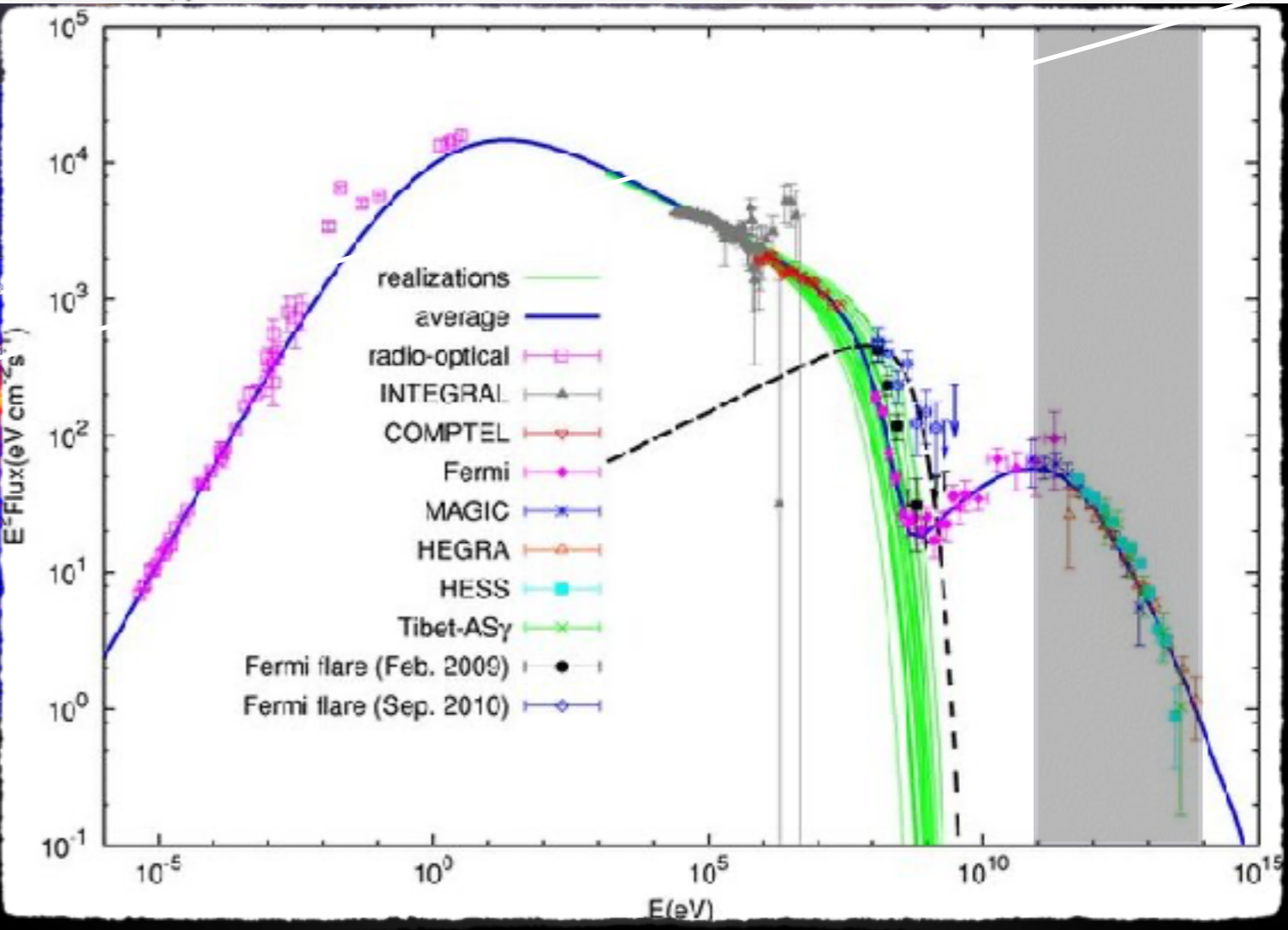
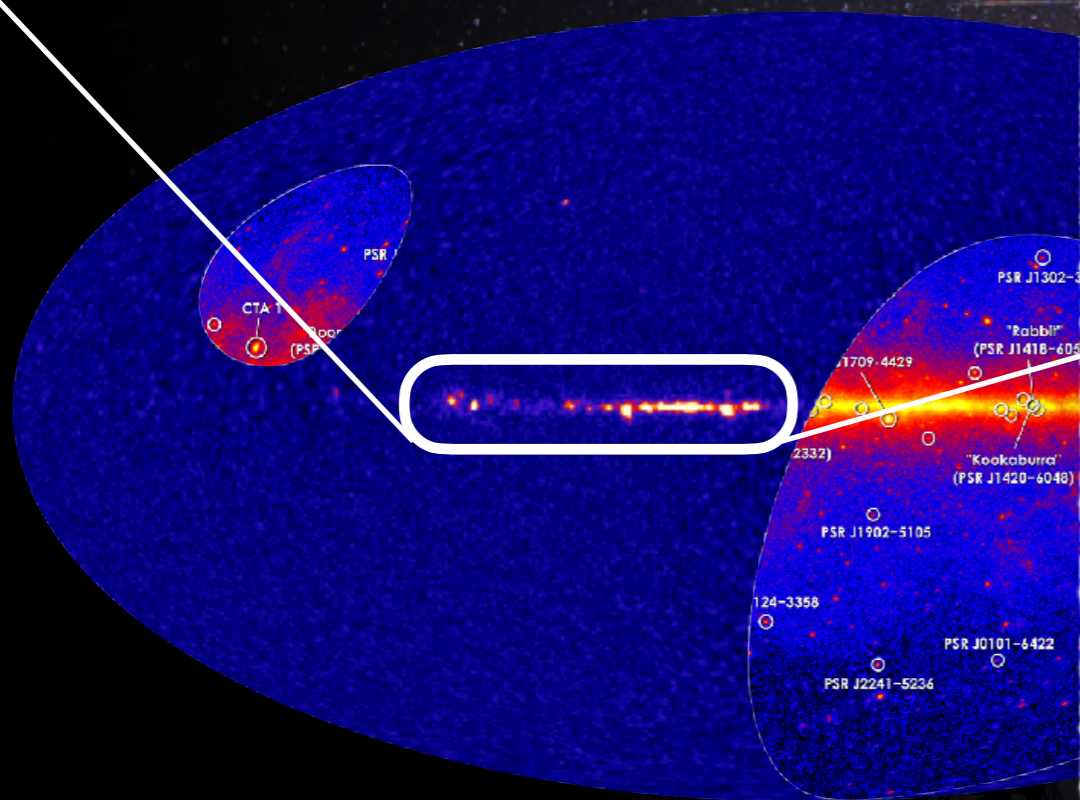
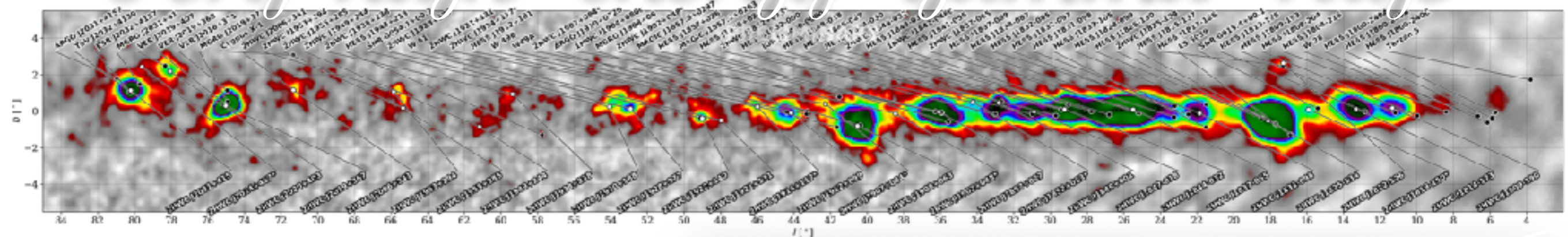
Cas A

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Very-high energy gamma rays

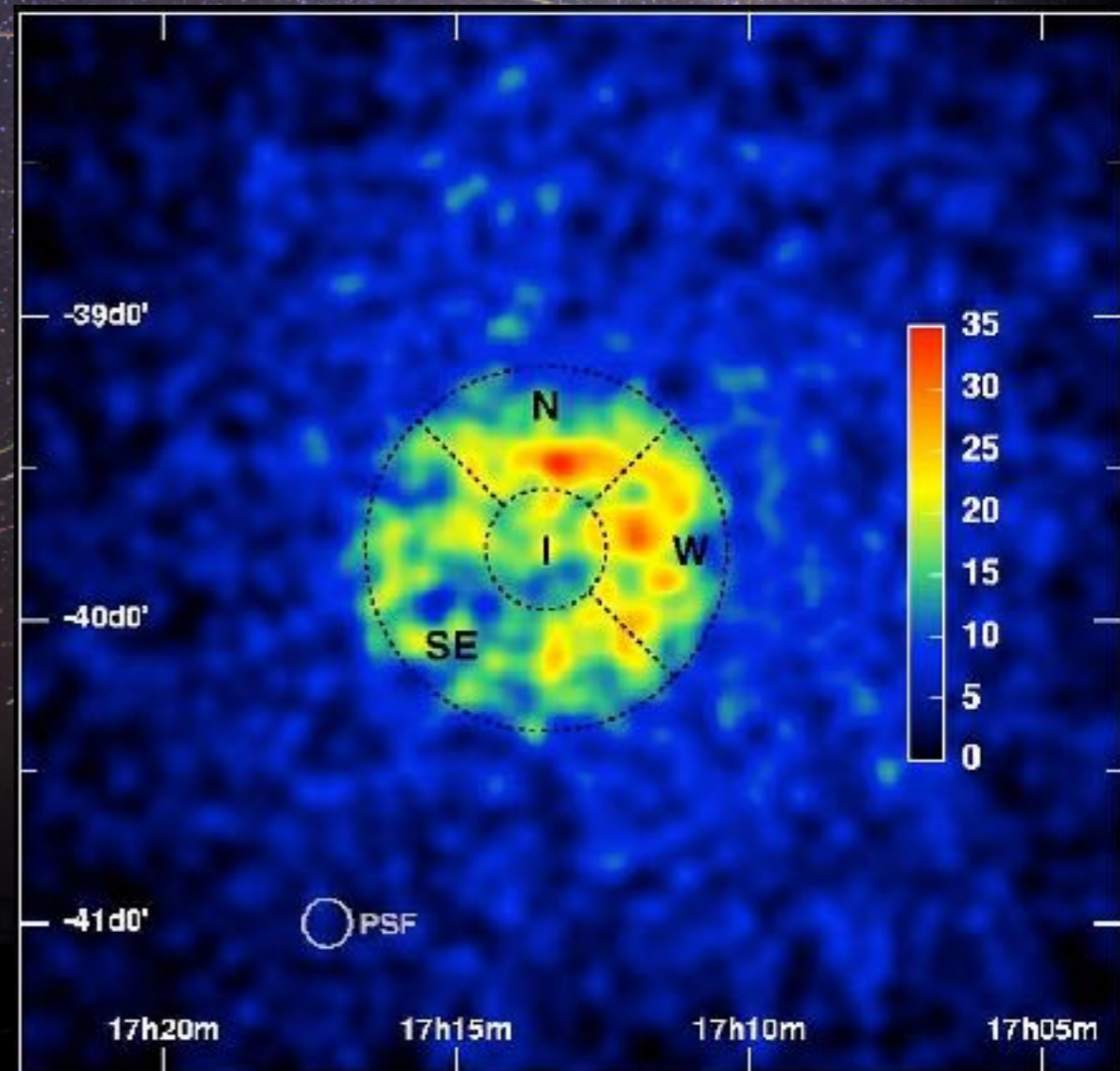


TeV energy range

Accelerating photons

First resolved **TeV γ -ray** image
of a Shell type SNR
(Resolution ~ 10 arcmin)

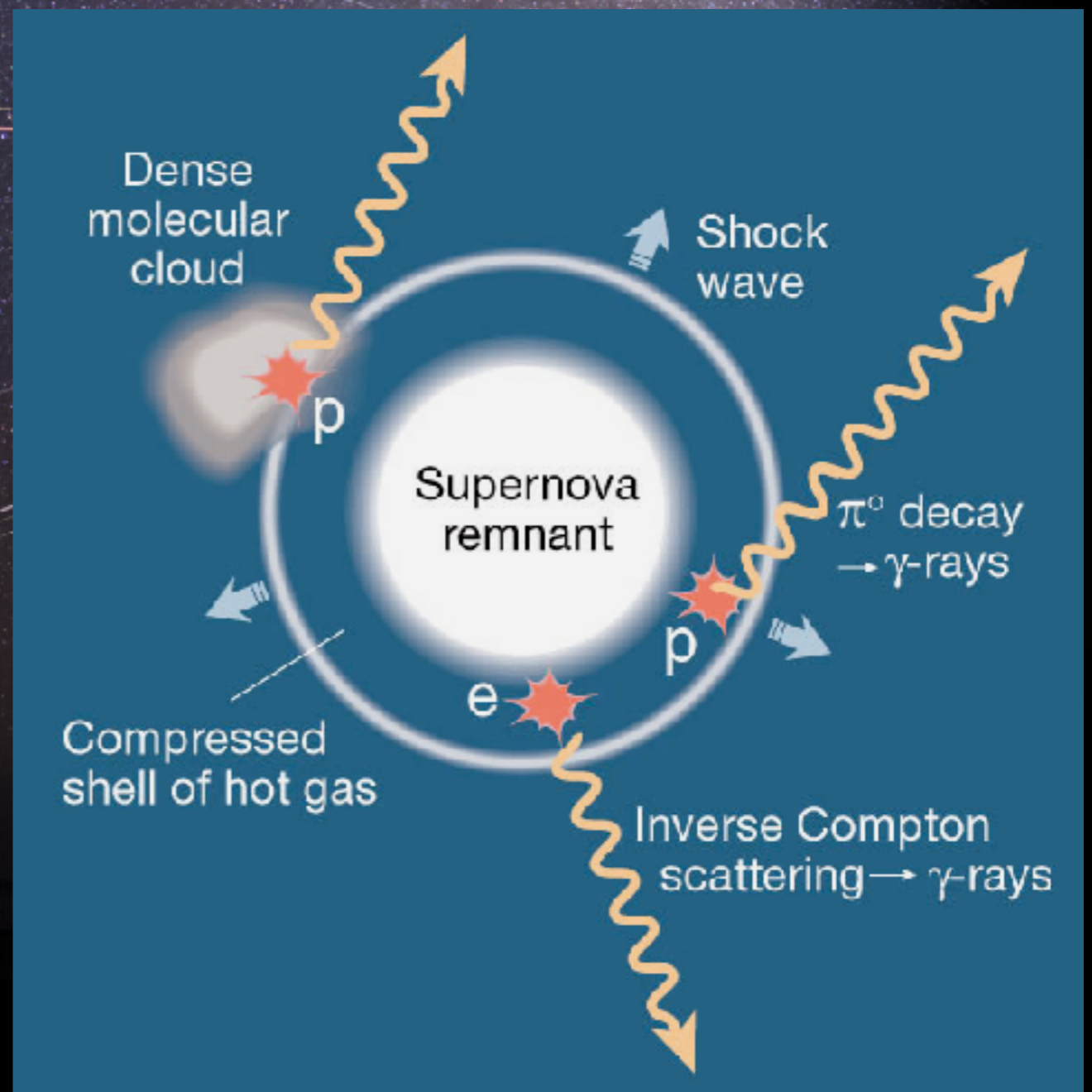
Acceleration source of cosmic
rays, but is it evidence of
protons?



Accelerating photons

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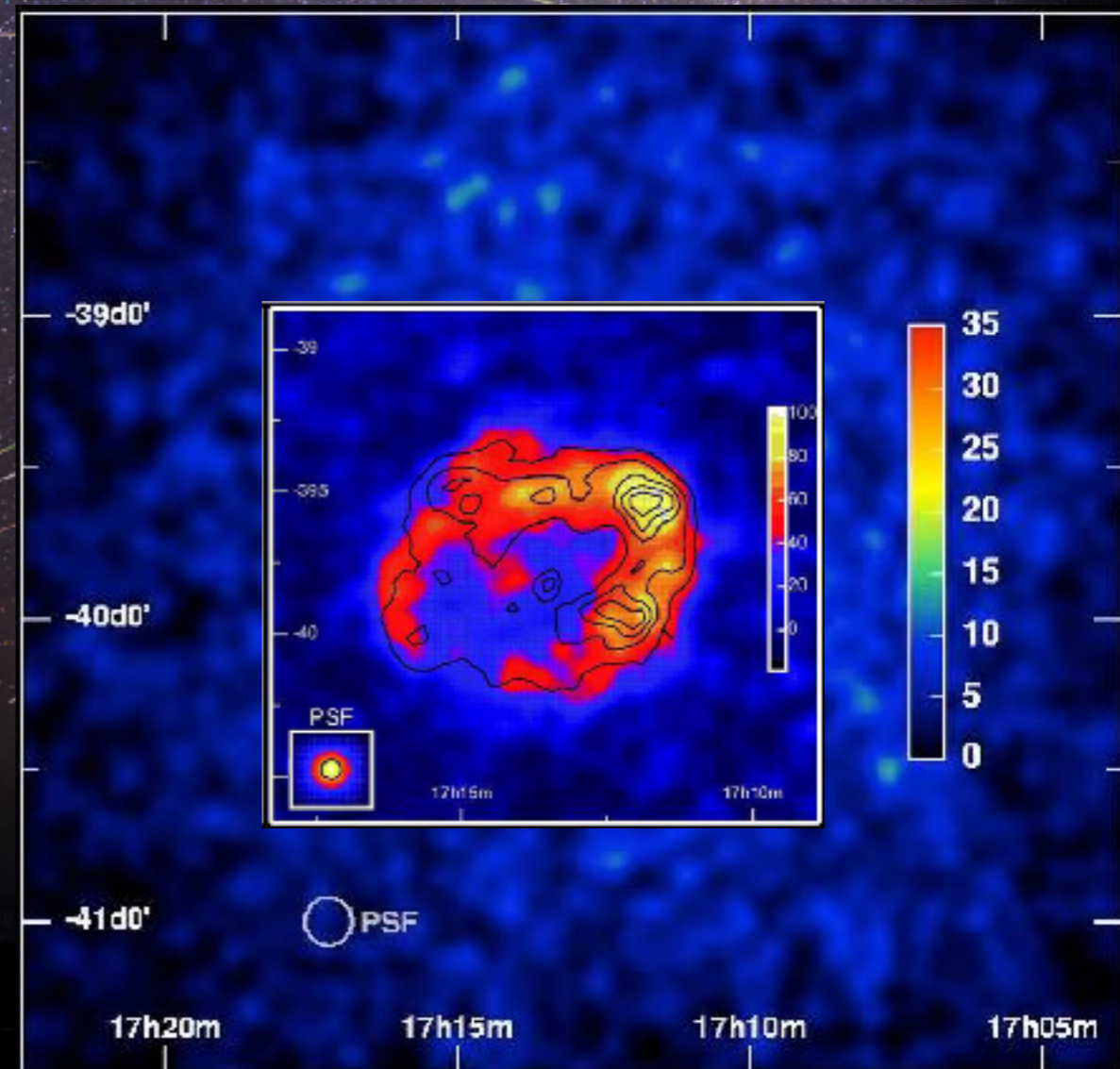
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Accelerating photons

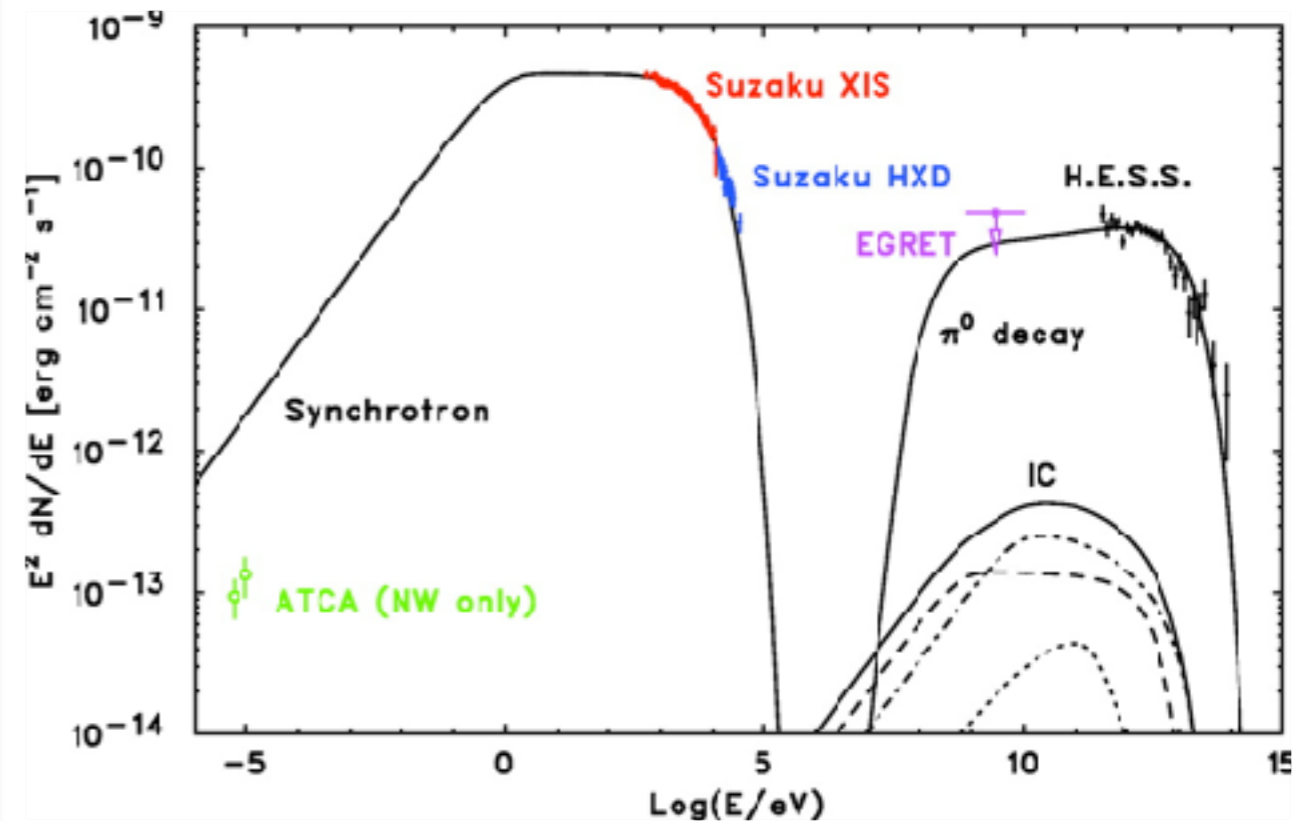
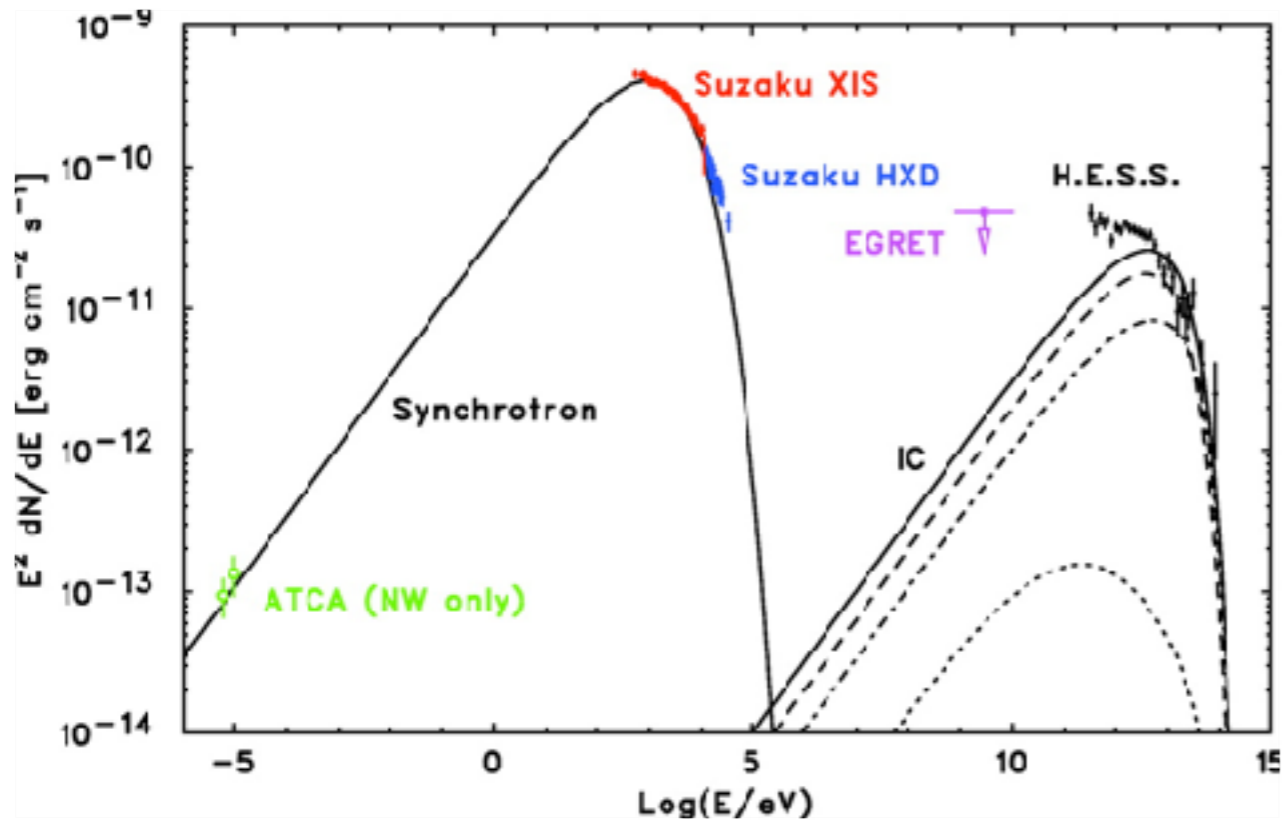
First resolved **TeV γ -ray** image
of a Shell type SNR
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Acceleration source of cosmic
rays, but is it evidence of
protons?



Accelerating photons

Tanaka et al., The Astrophysical Journal 685 (2008) 988



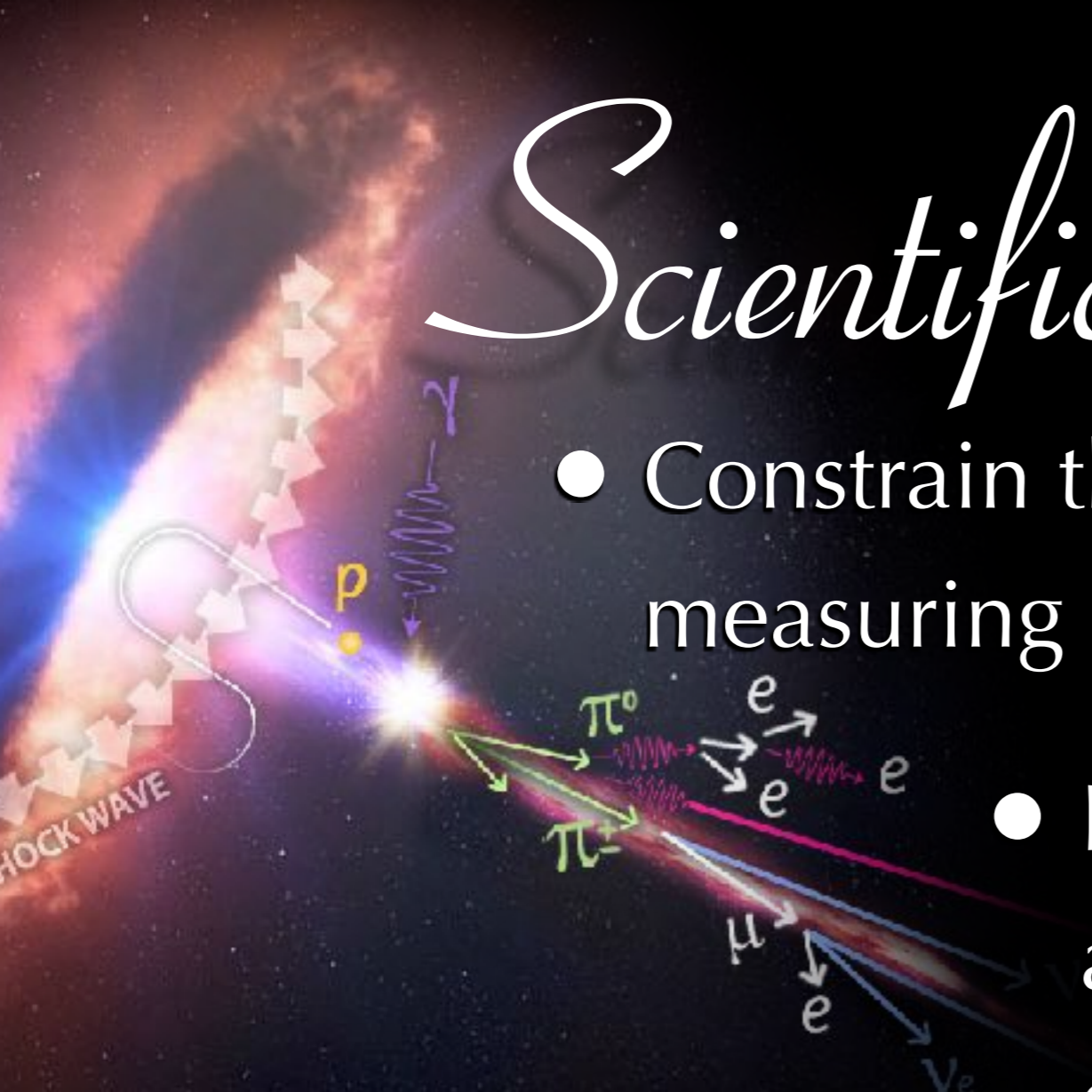
Leptonic origin (i.e., electrons) *vs.* Hadronic origin (i.e., protons)

Scientific motivation

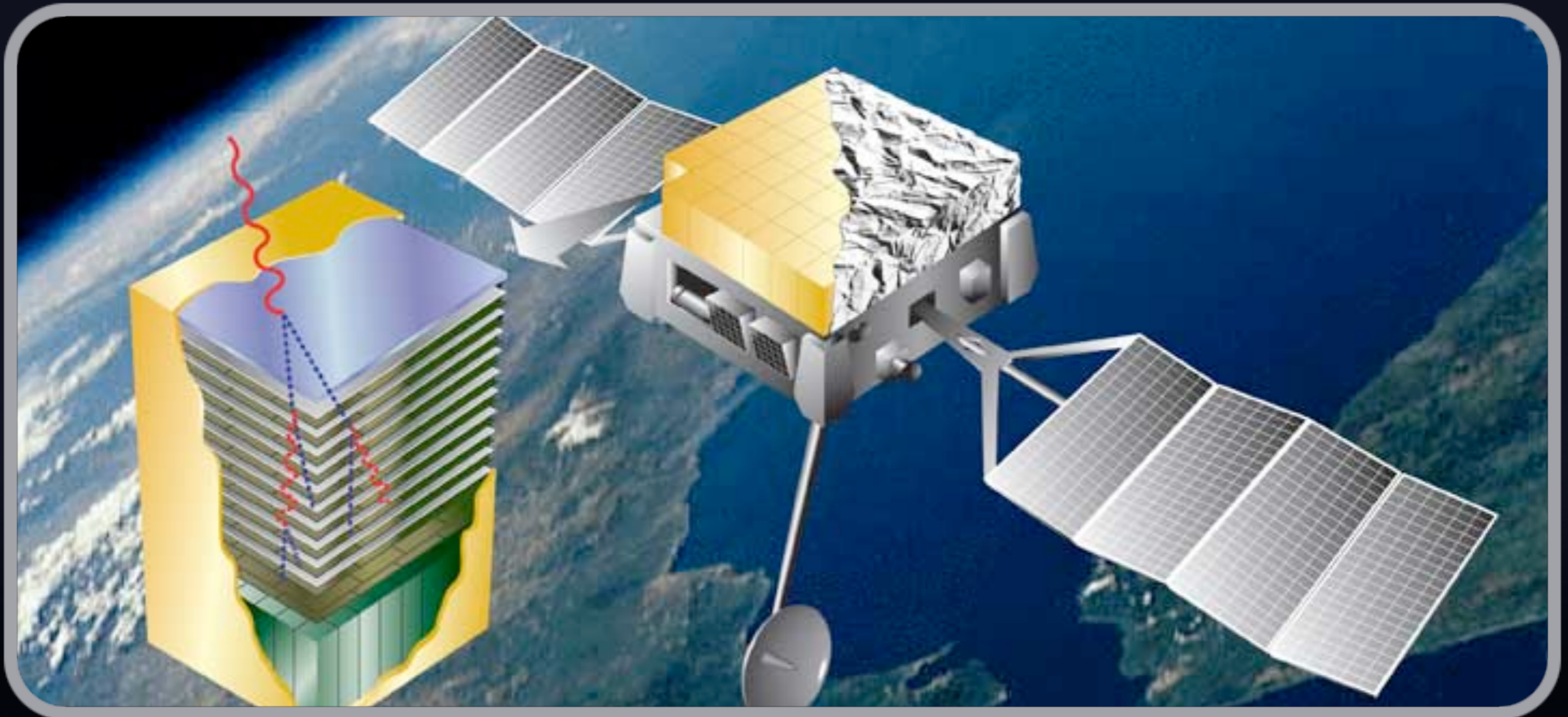
- Constrain the **origin of cosmic rays** by measuring gamma-ray spectra to 100 TeV.

- Probe **particle acceleration** in astrophysical objects with a complimentary set of instruments.

- Explore **new physics** in the TeV energy range.



Experimental Techniques



- **Space-based detectors**
 - Low energy threshold**
 - EGRET, Fermi-LAT

Experimental Techniques

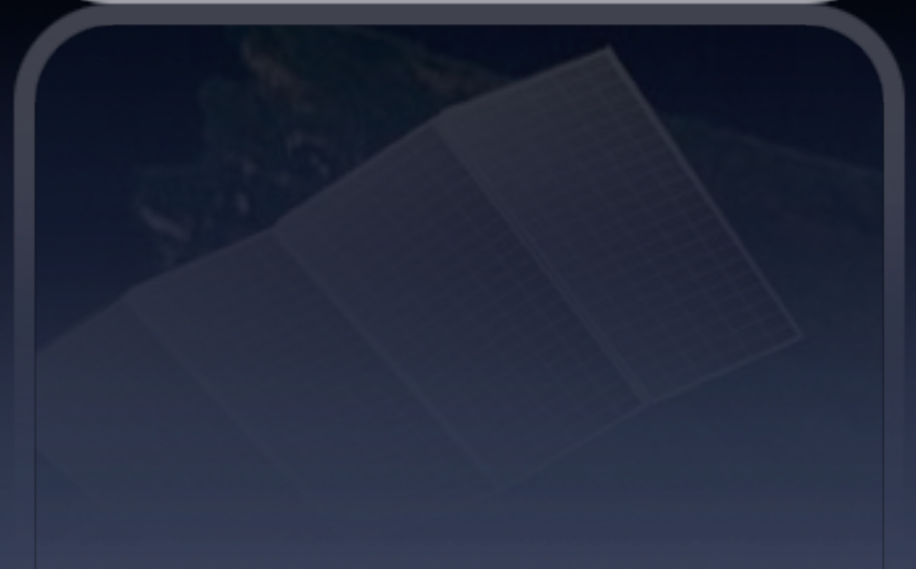
- ✓ Background free
- ✓ Large duty cycle
- ✓ Large aperture

- Small area

● Space-based detectors

Low energy threshold

EGRET, Fermi-LAT



Experimental Techniques



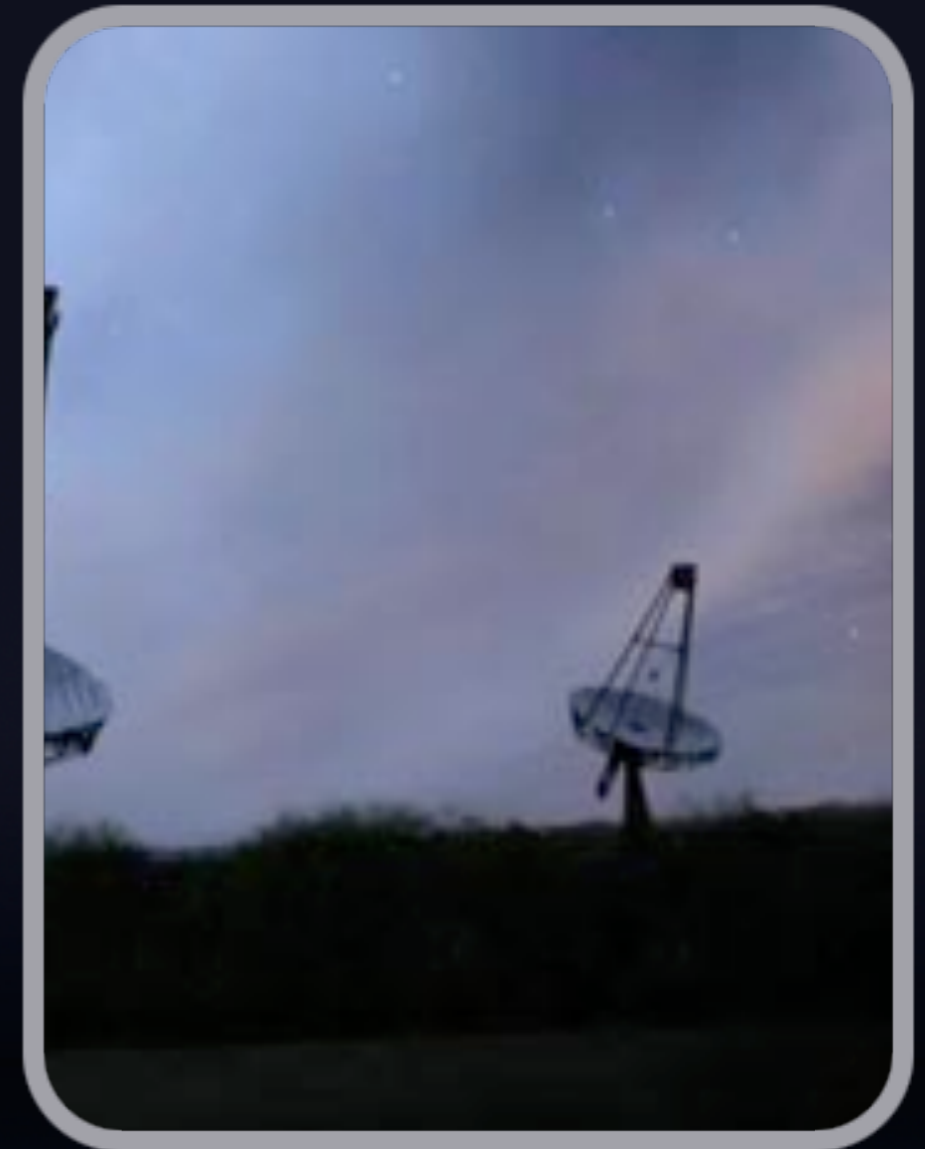
● Imaging Atmospheric Cherenkov Telescopes

High sensitivity

HESS, MAGIC, VERITAS, CTA

Experimental Techniques

- ✓ Large effective area
- ✓ Excellent background rejection
- Small aperture
- Low duty cycle



● Imaging Atmospheric Cherenkov Telescopes

High sensitivity

HESS, MAGIC, VERITAS, CTA

Experimental Techniques



- **Ground array of air-shower particle detectors**
Large aperture + High duty cycle
Milagro, Tibet, ARGO, HAWC, LHAASO

Experimental Techniques

- ✓ Large aperture
- ✓ Excellent background rejection
- ✓ Large duty cycle

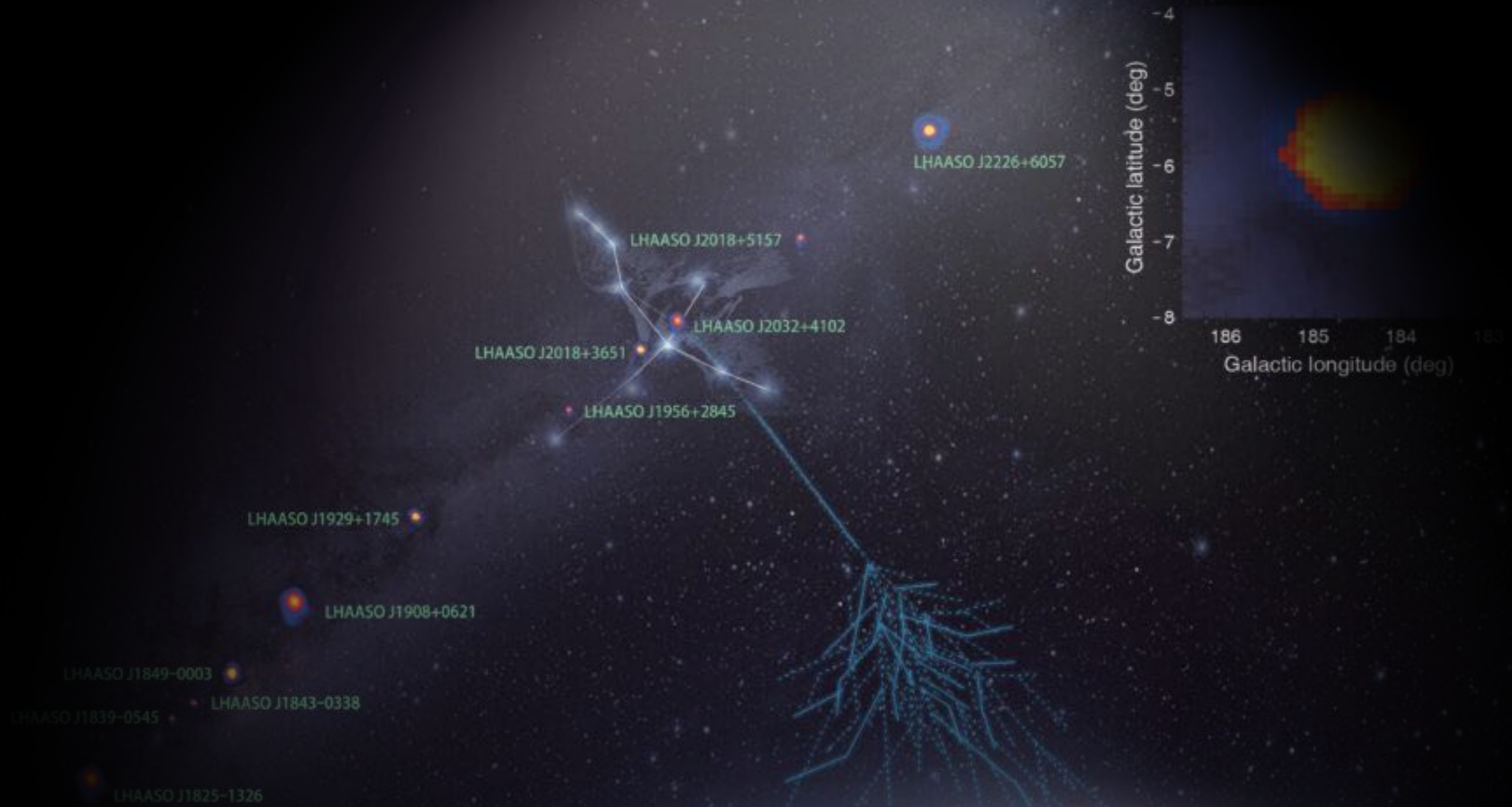
- Moderate area

● Ground array of air-shower particle detectors

Large aperture + High duty cycle

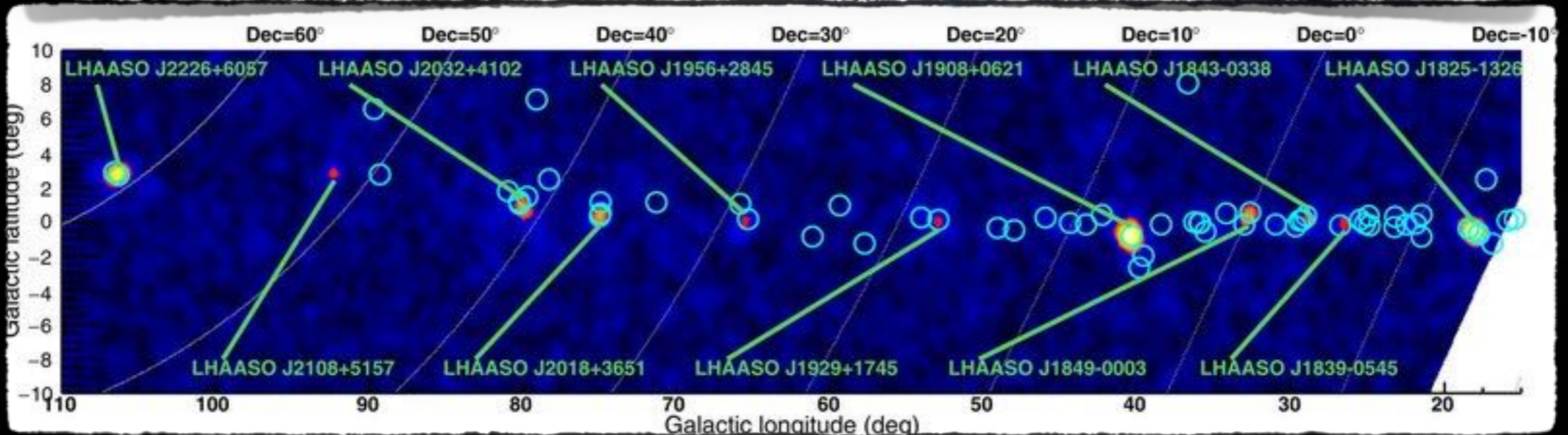
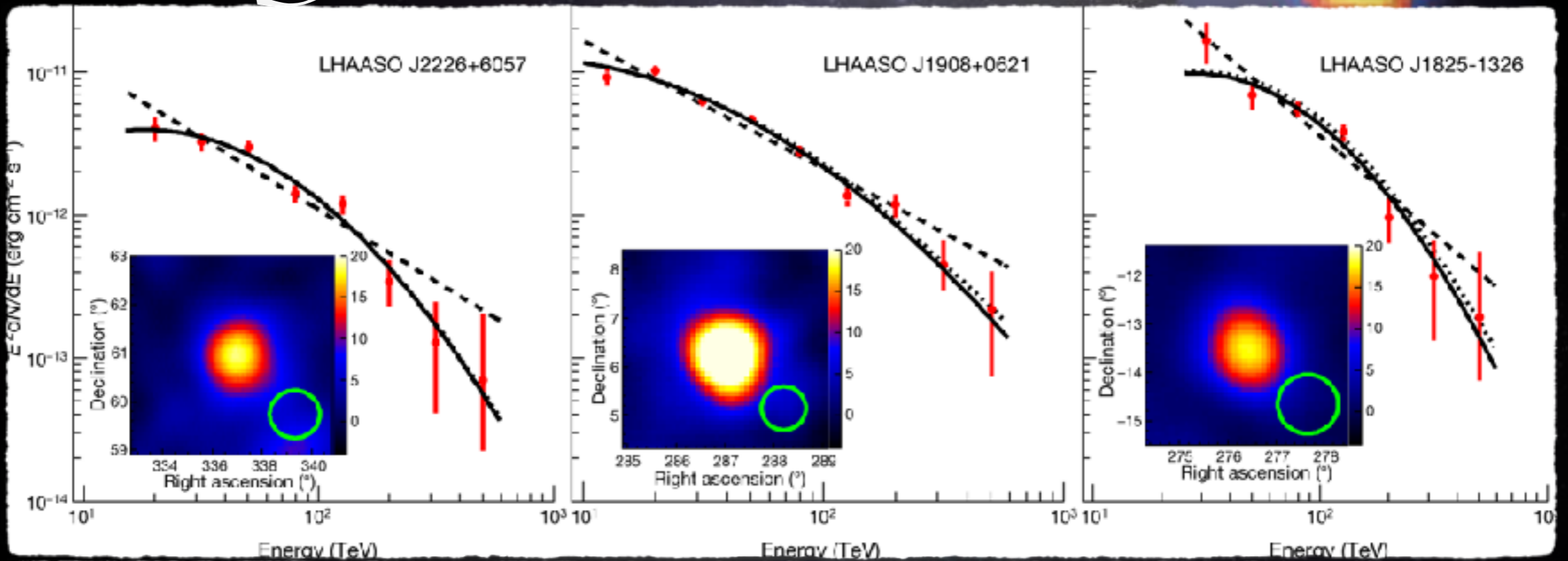
Milagro, Tibet, ARGO, HAWC, LHAASO



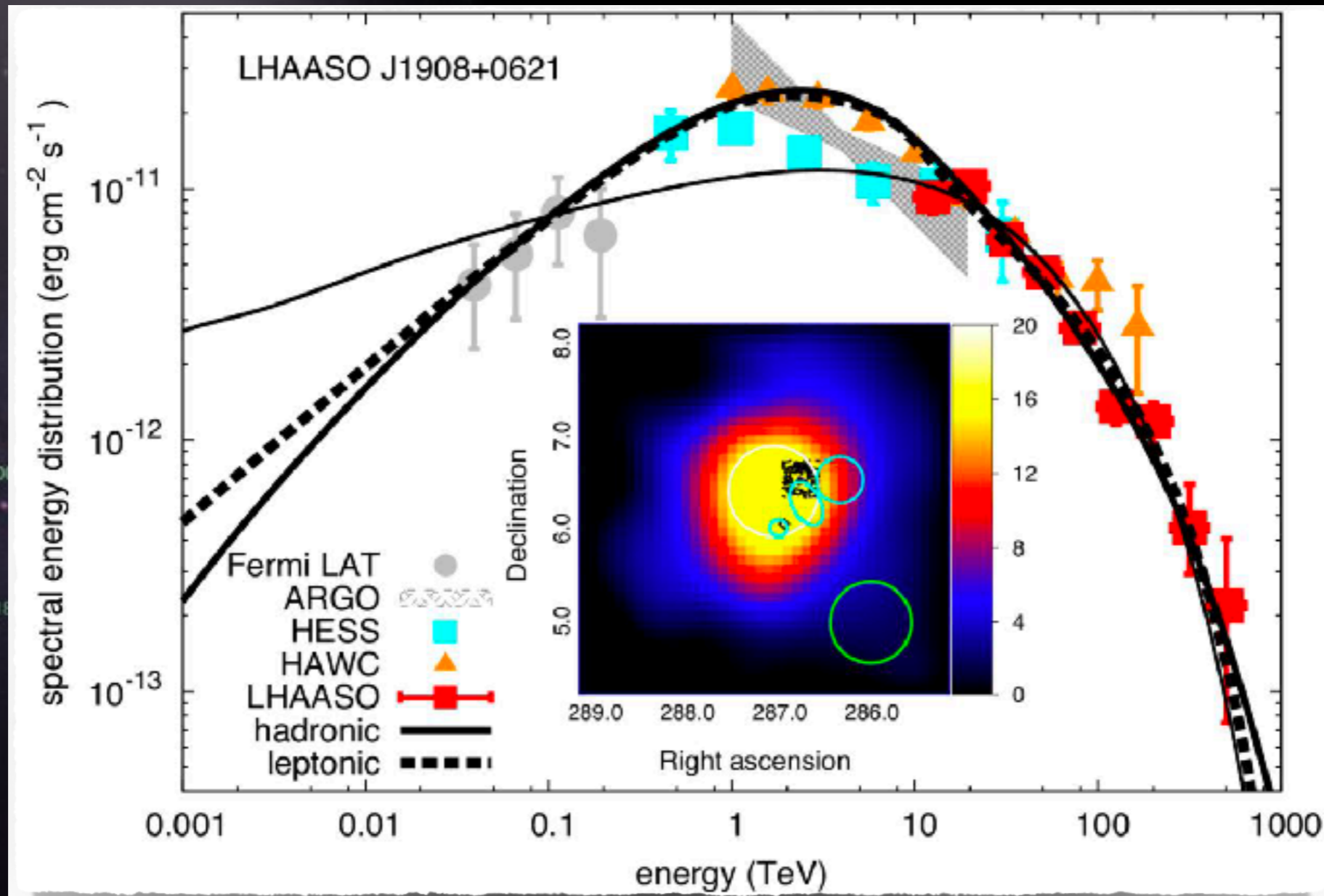


Selected Results

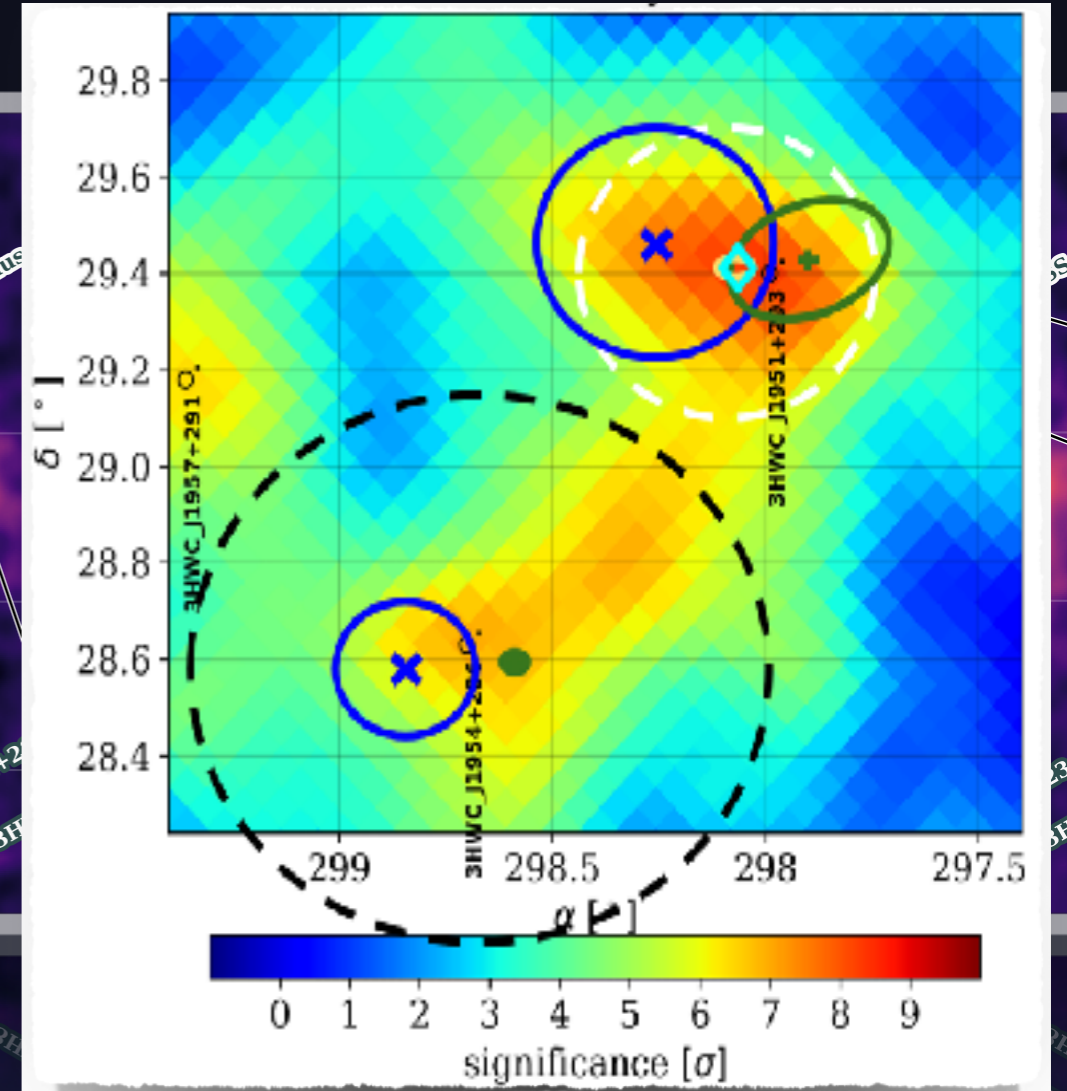
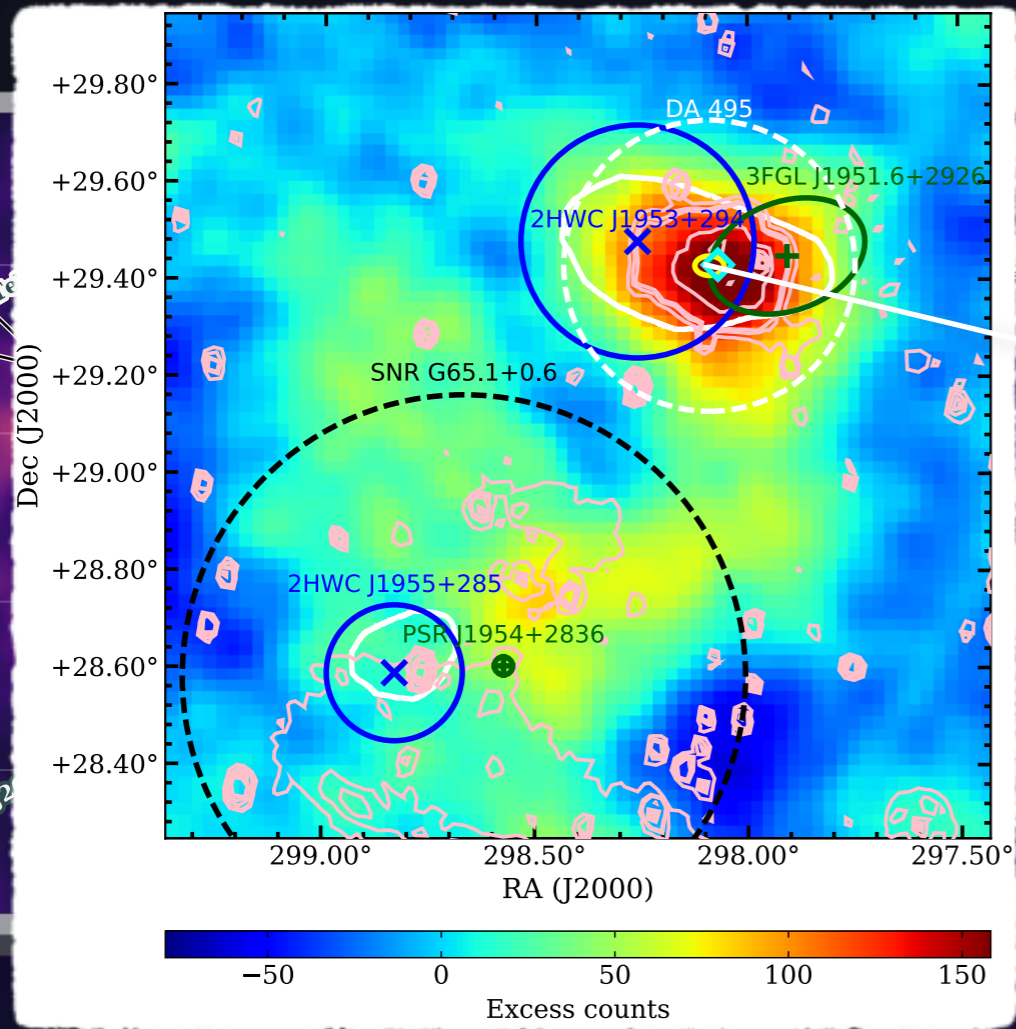
Galactic Pevatrons



Galactic Pevatrons



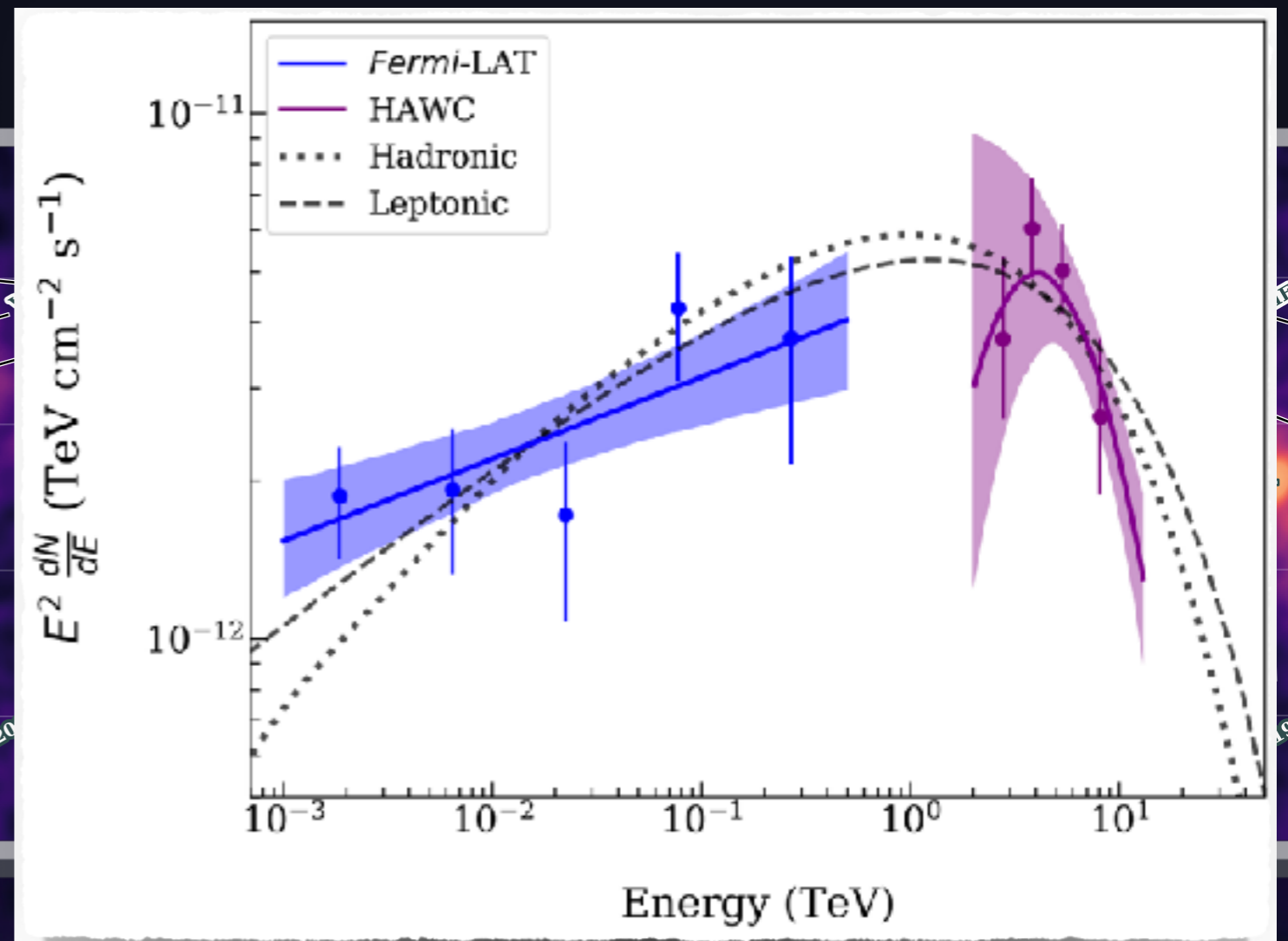
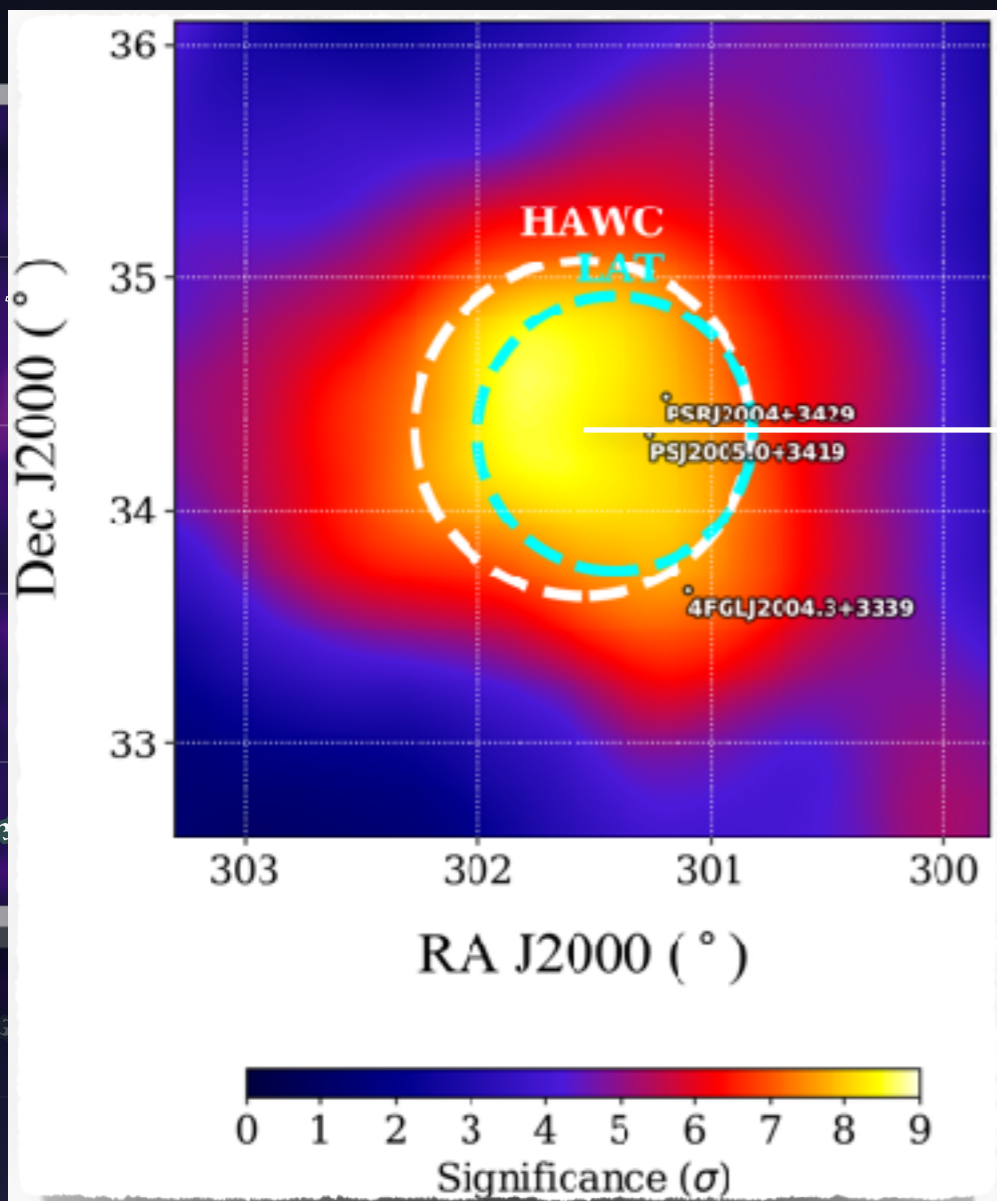
New TeV sources



VERITAS counts map

3HWC significance map

New TeV sources

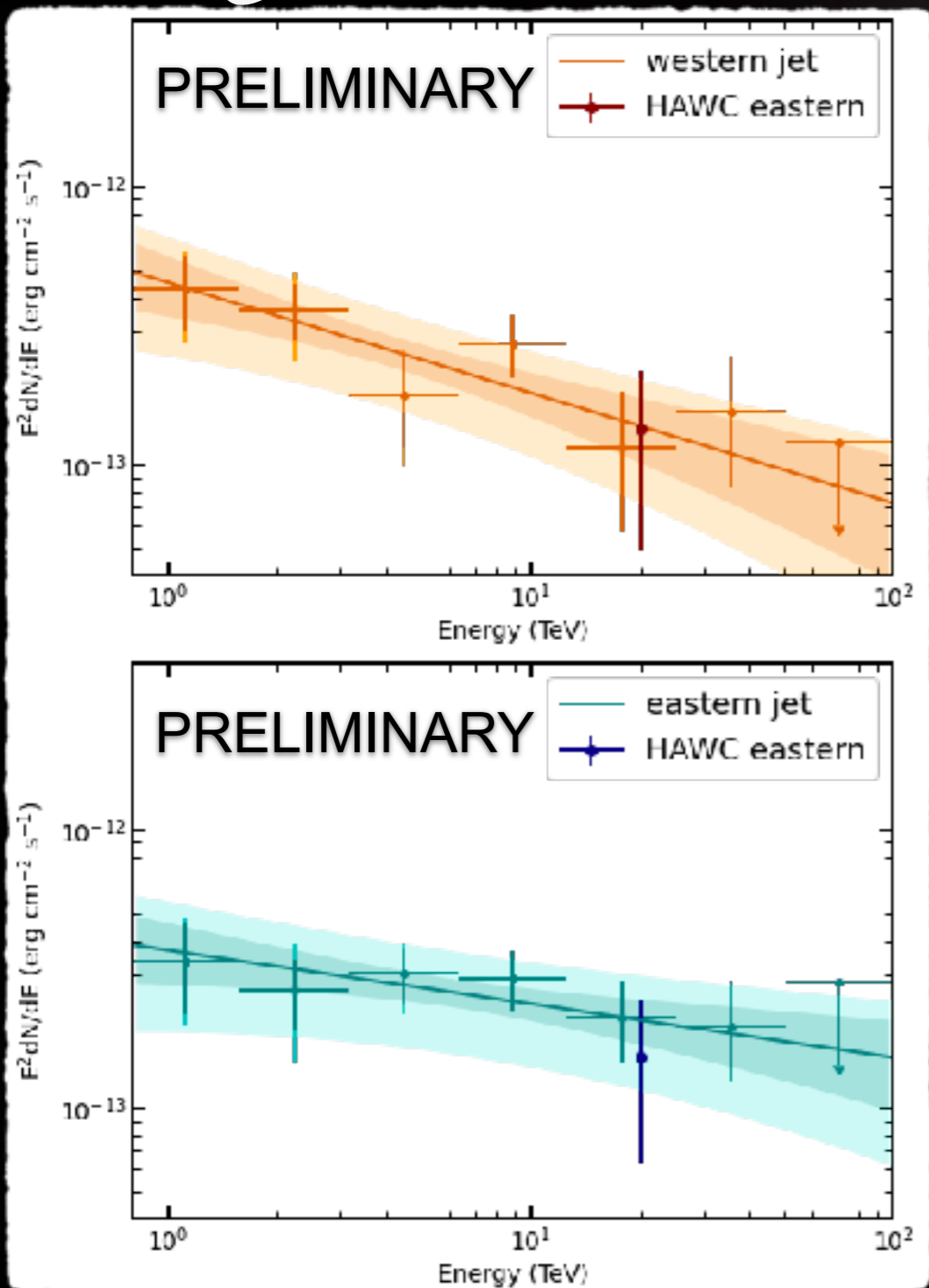


SED from HAWC and LAT data

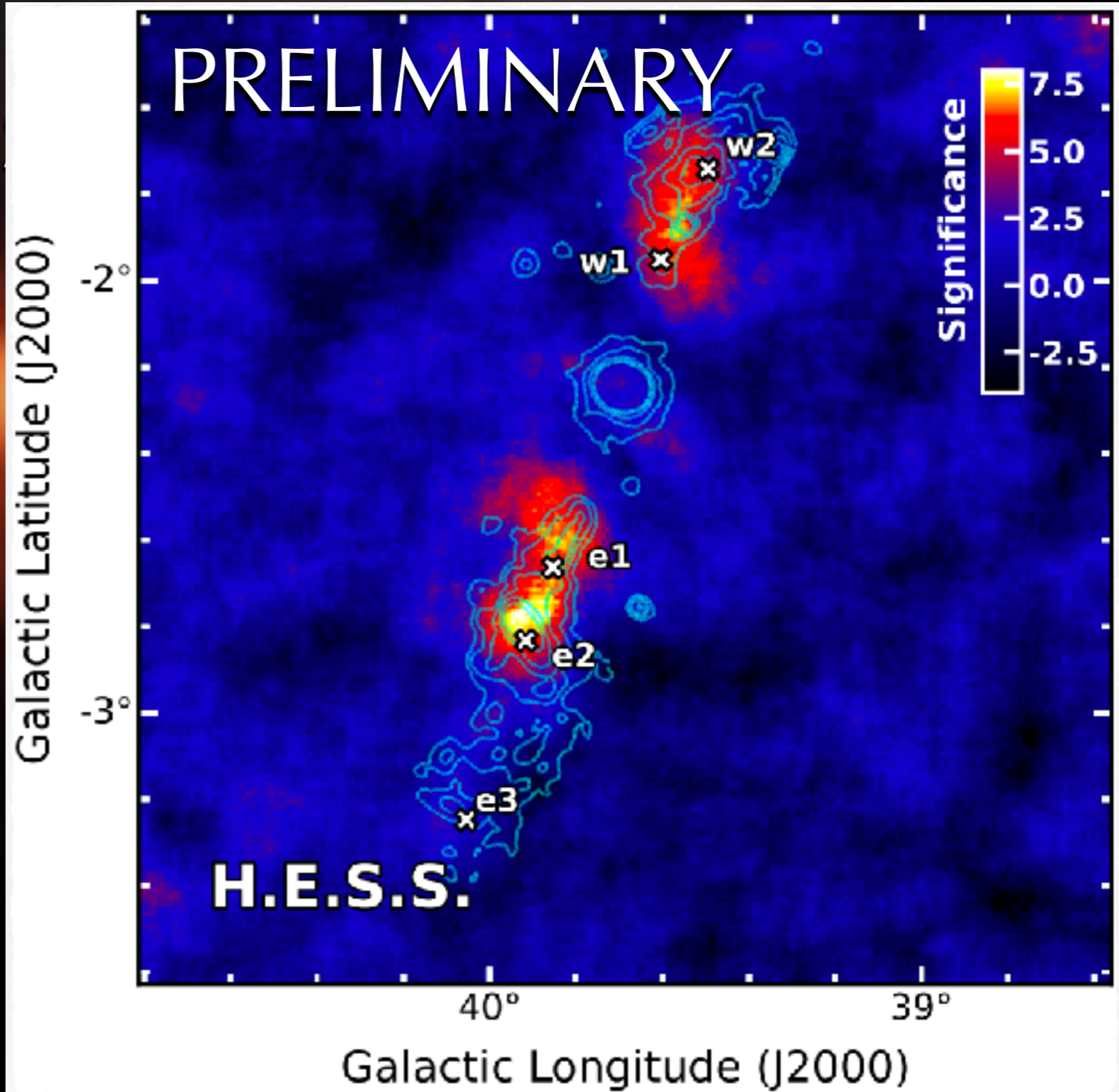
3HWC significance map

[HAWC+Fermi-LAT] ApJL 903 (2020) L14

Jets of a Microquasar



S



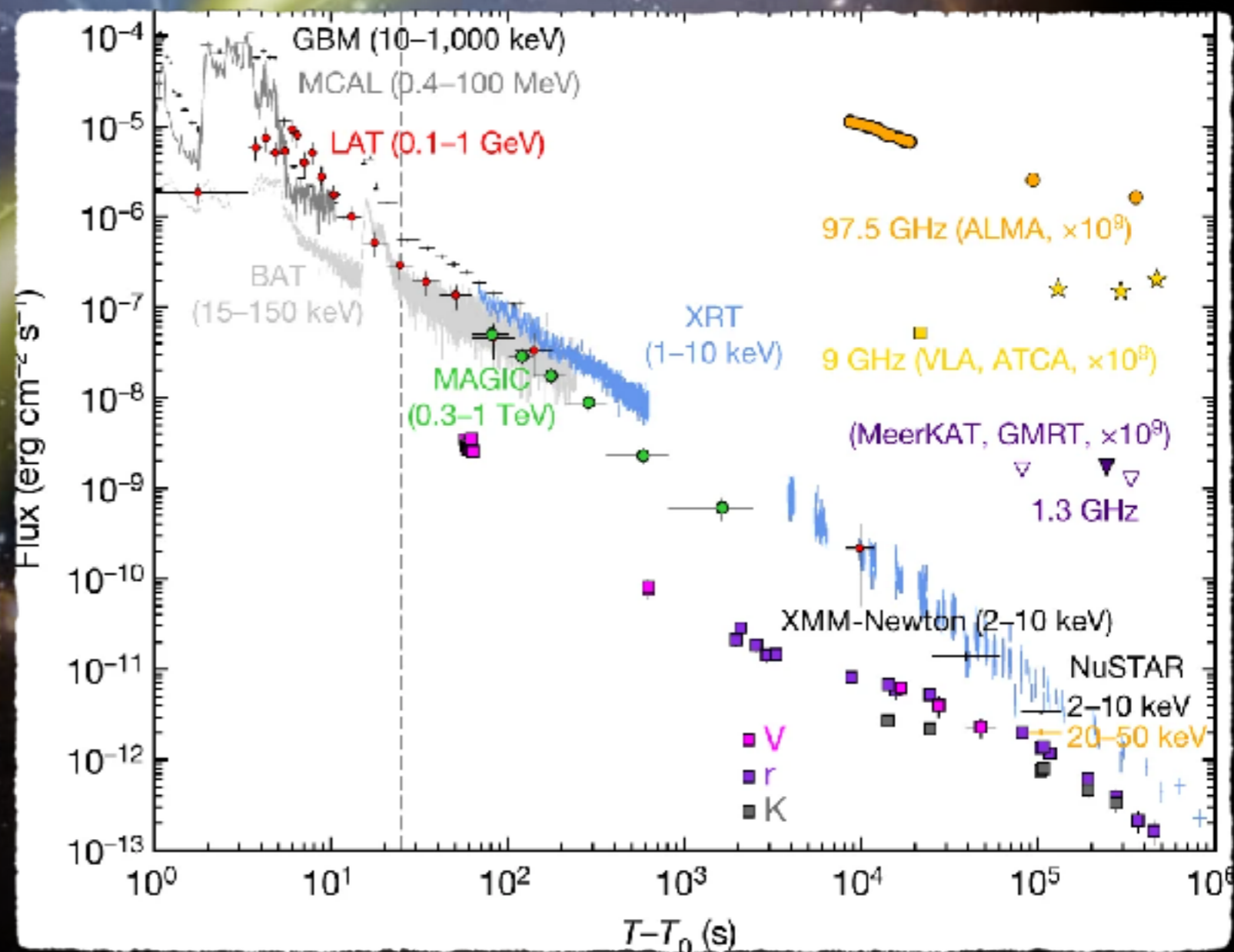
Gamma Ray Bursts

GRB190114C: 1st detection of VHE emission!

Brightest VHE γ -ray source!

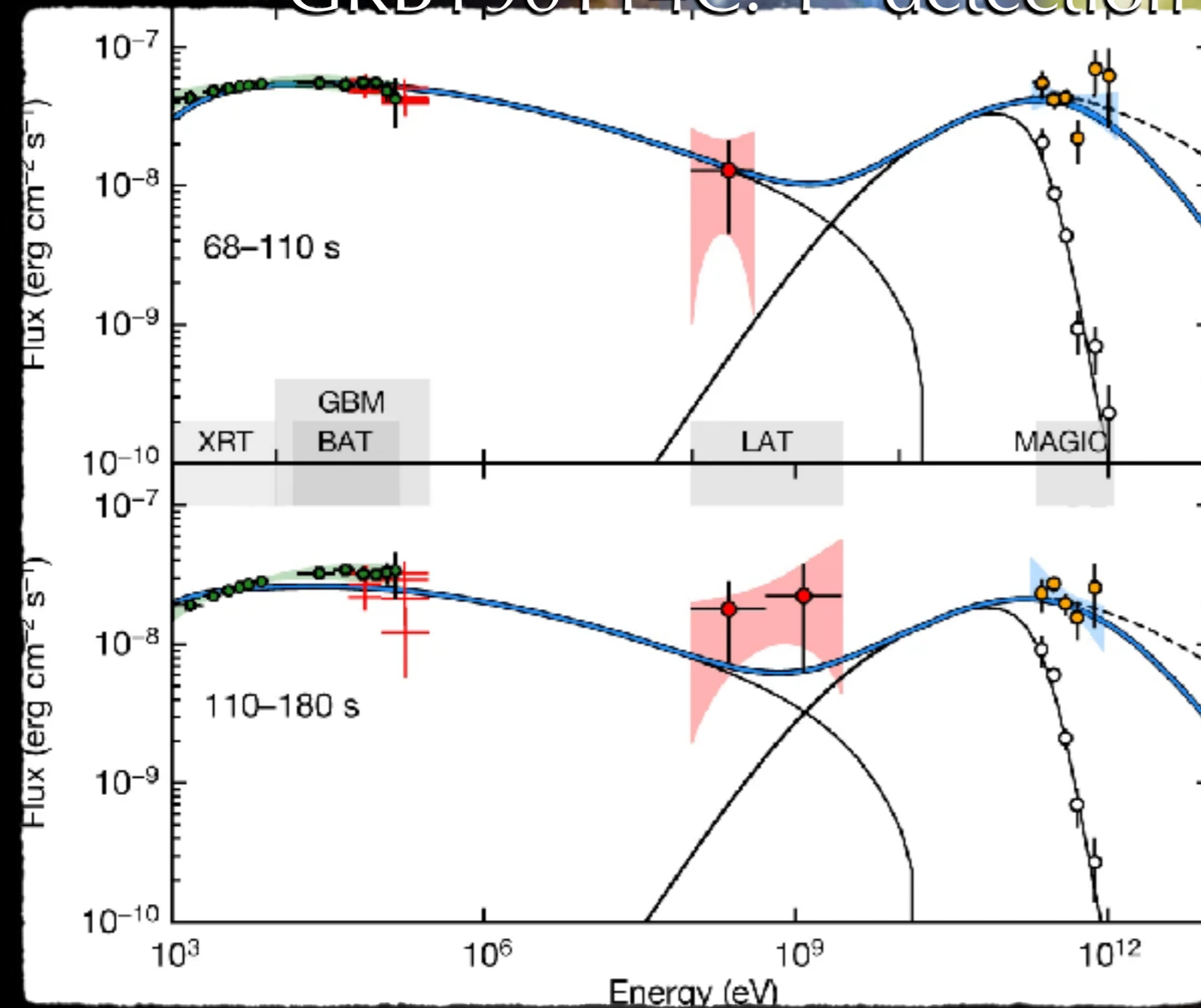
Long GRB
($T_{90} \sim 6$ min)

Observations started < 1 min after BAT trigger



Gamma Ray Bursts

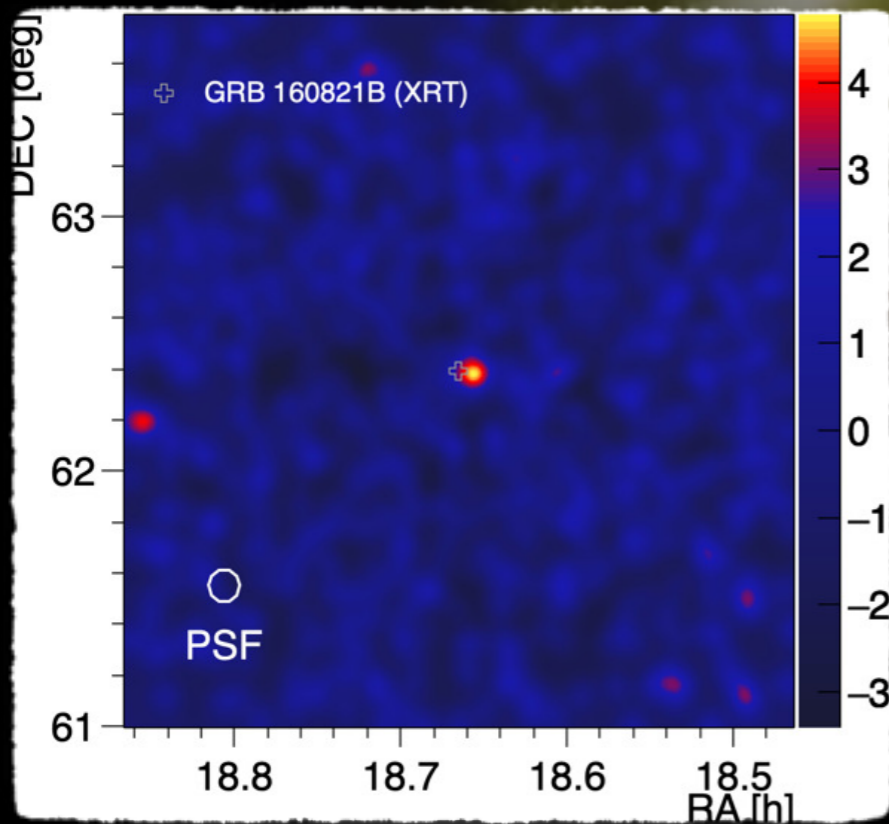
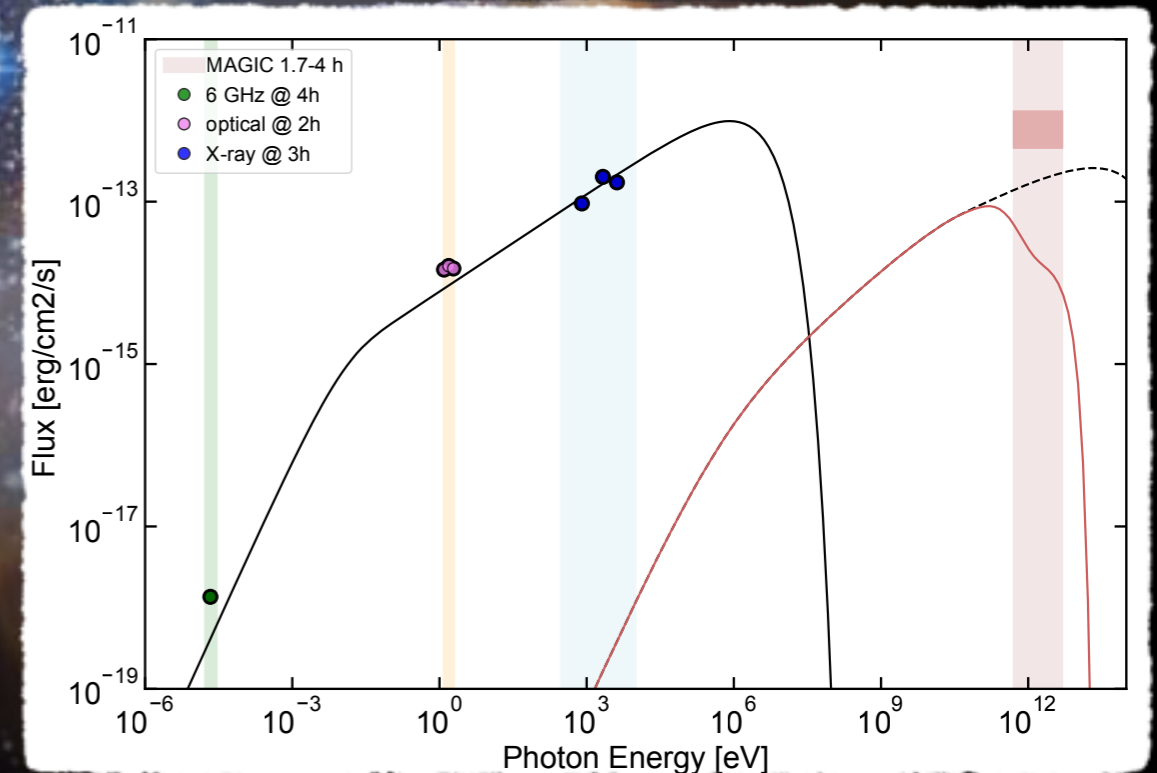
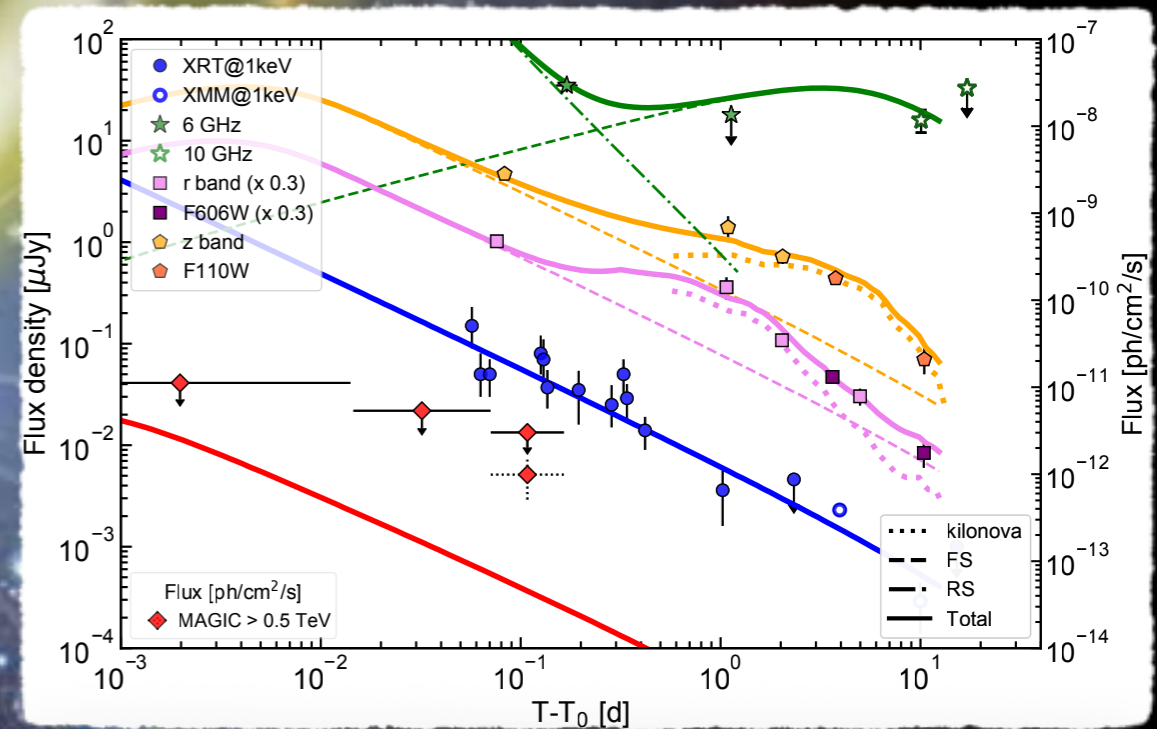
GRB190114C: 1st detection of VHE emission!



Highest energy emission strongly absorbed by EBL
VHE γ -ray emission can be modeled in SSC scenario

Gamma Ray Bursts

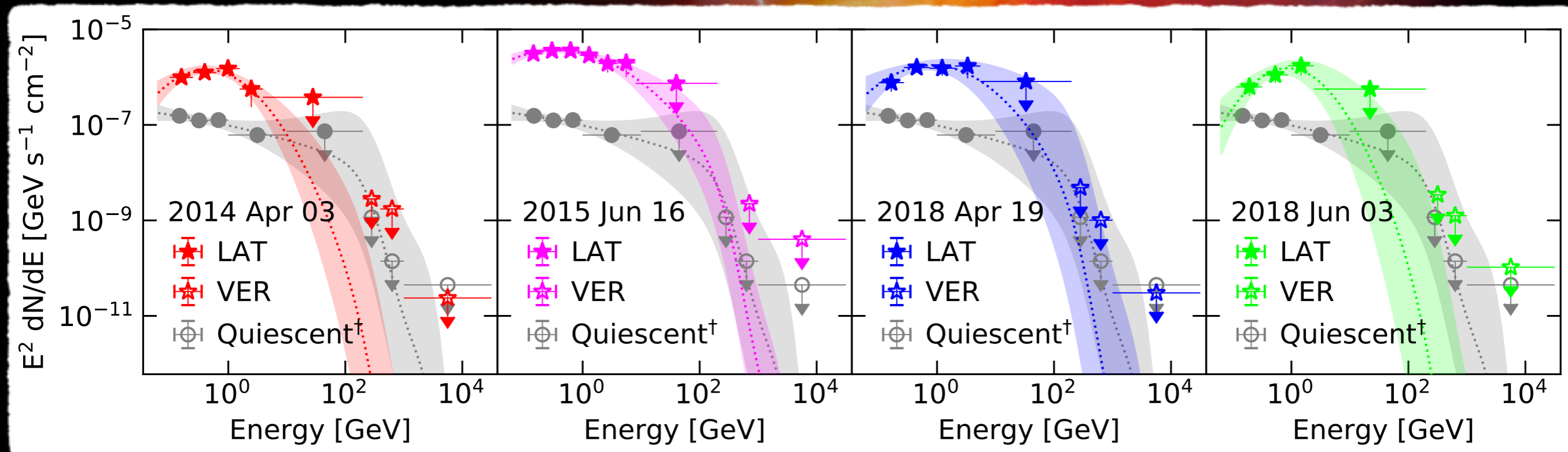
GRB160821B: $\sim 3\sigma$ evidence of γ -ray emission above ~ 0.5 TeV
One-zone models of SSC emission have a hard time explaining the putative TeV flux



Flaring quasars

γ -ray variability and spectral characteristics
of FSRQs during bright GeV flares

Flux upper limits on 3C 279



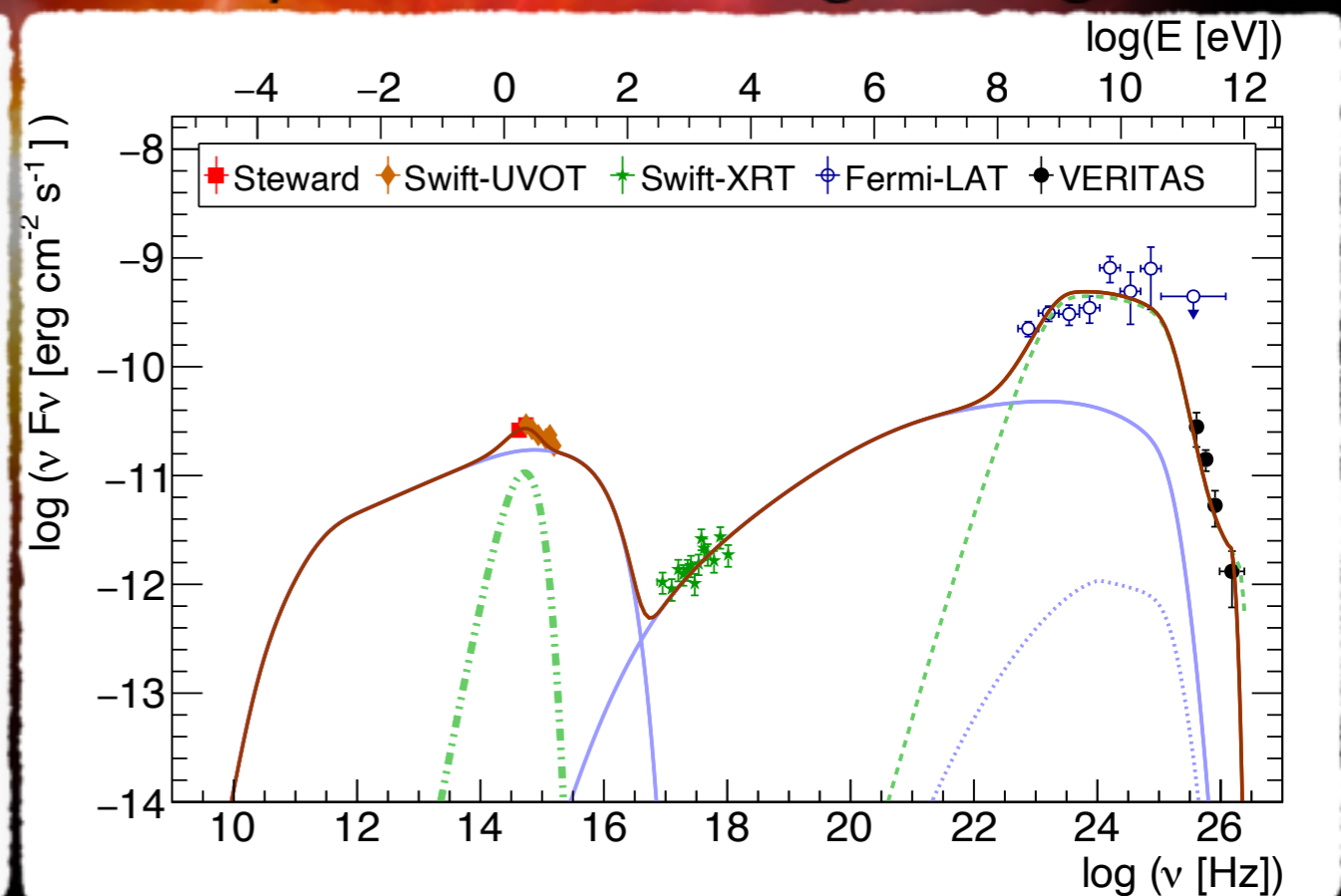
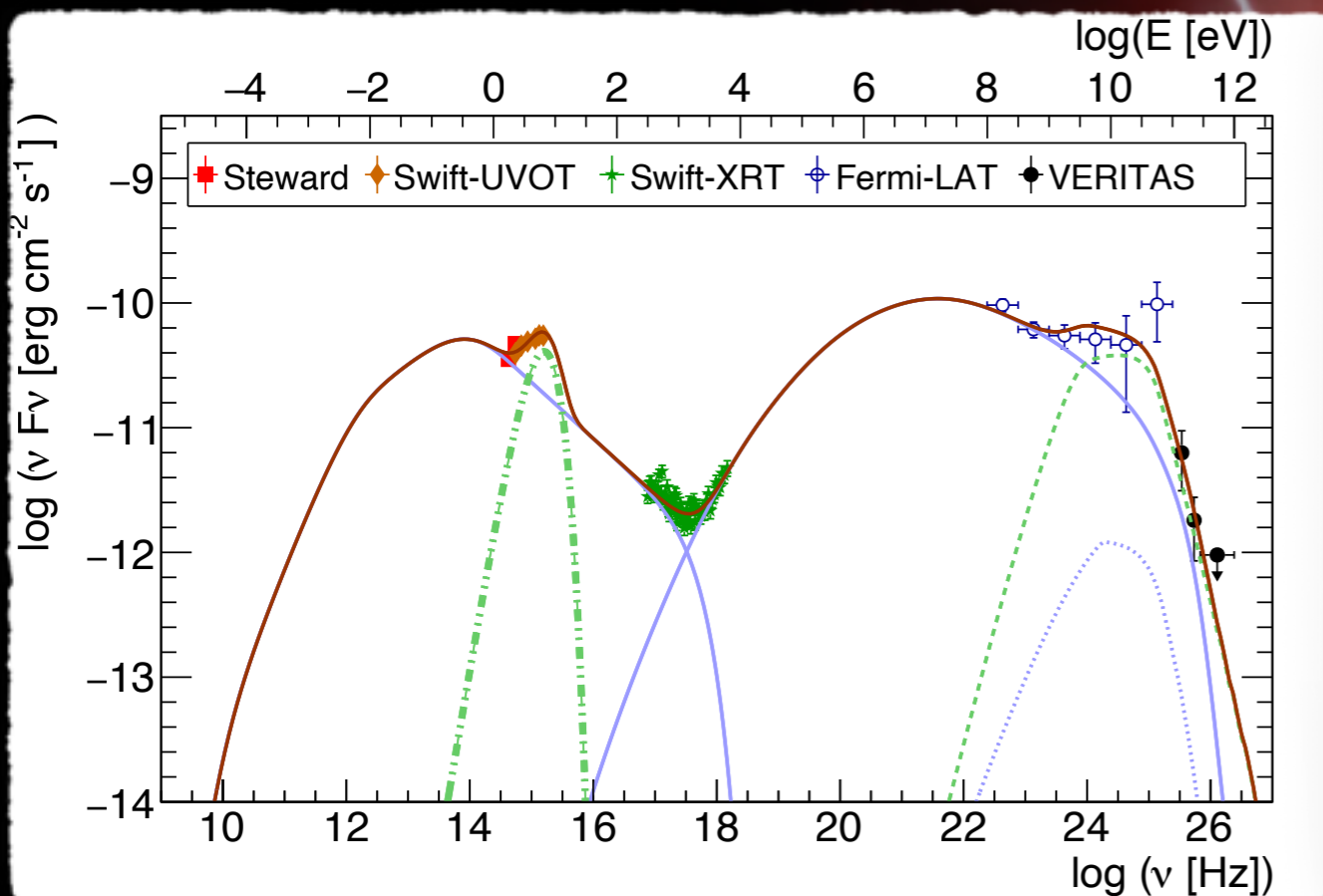
100 h of VERITAS observations over 10 y + LAT data

Adams *et al* [VERITAS+Fermi-LAT] *ApJ* 924 (2022) 95

Flaring quasars

γ -ray variability and spectral characteristics
of FSRQs during bright GeV flares

Both PKS 1222+216 and TON 599 detected by VERITAS during flaring states!



100 h of VERITAS observations over 10 y + LAT data

Adams *et al* [VERITAS+Fermi-LAT] *ApJ* 924 (2022) 95

Multimessenger Astrophysics

Extensive follow-up/MM alert programs

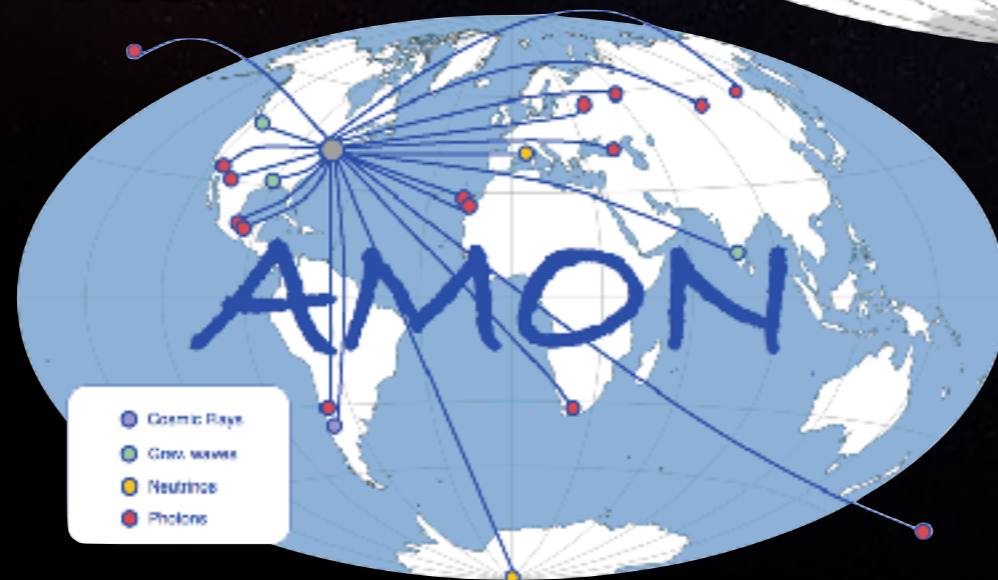
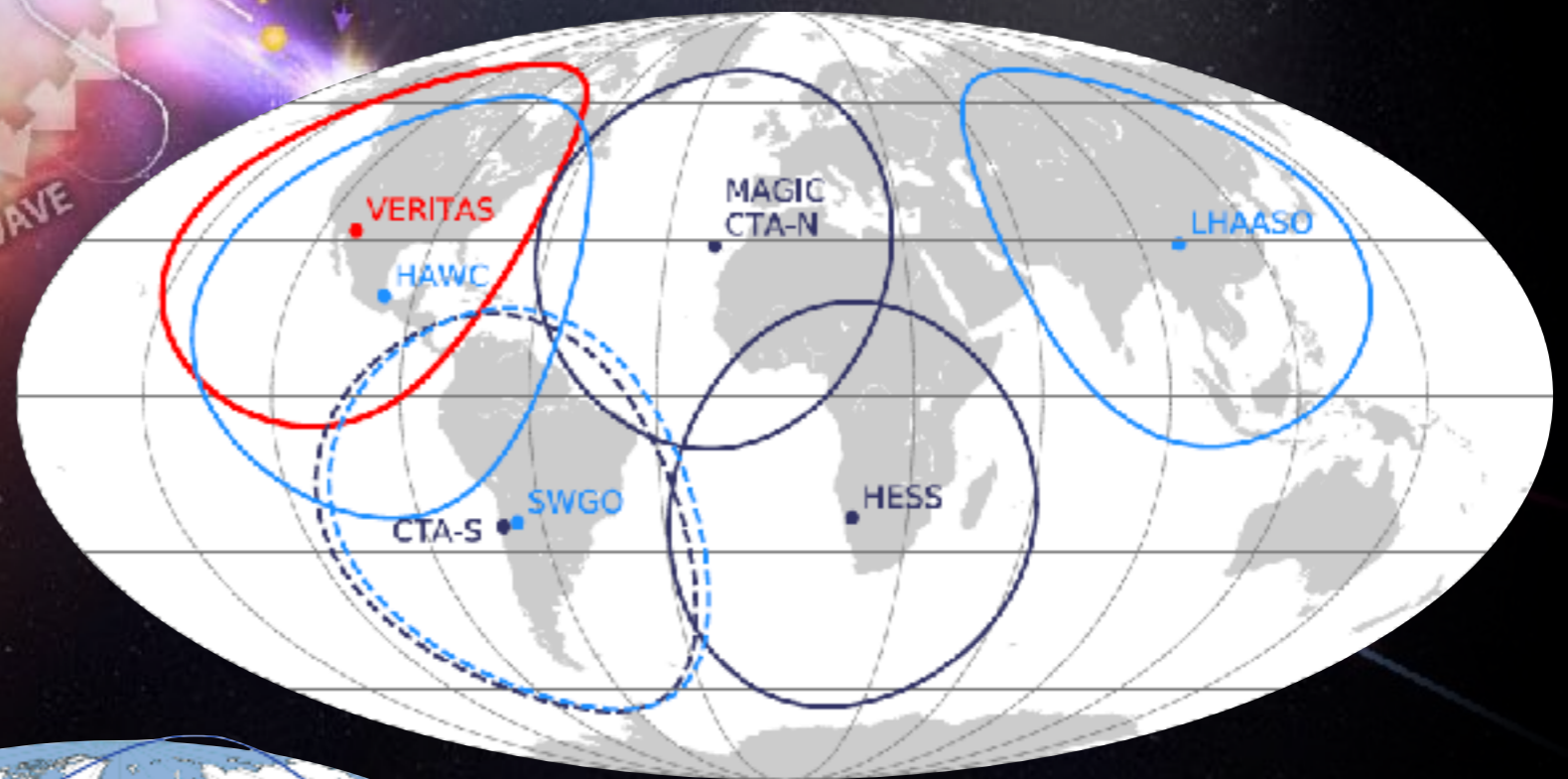
Neutrinos

GWs

FRBs

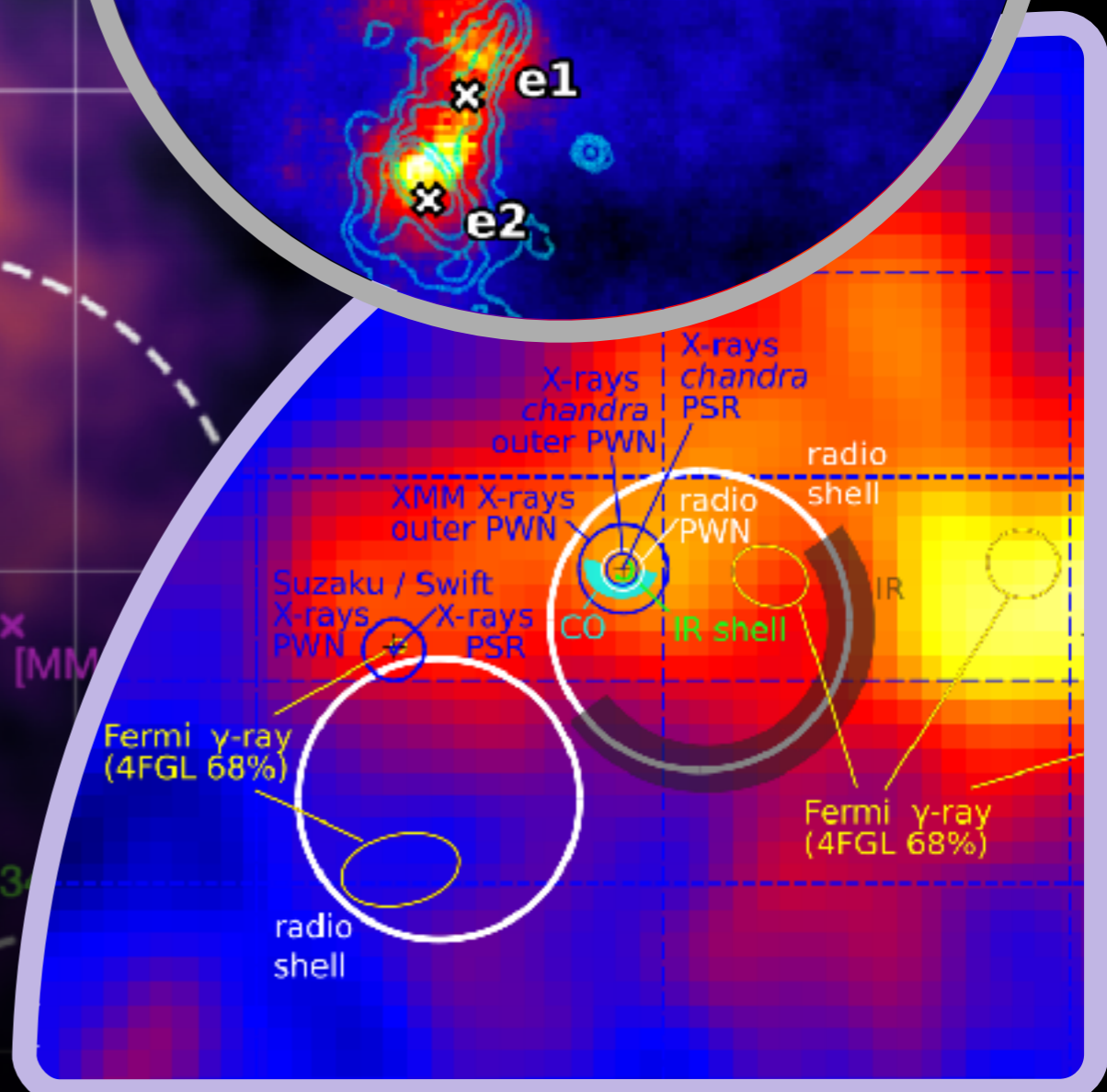
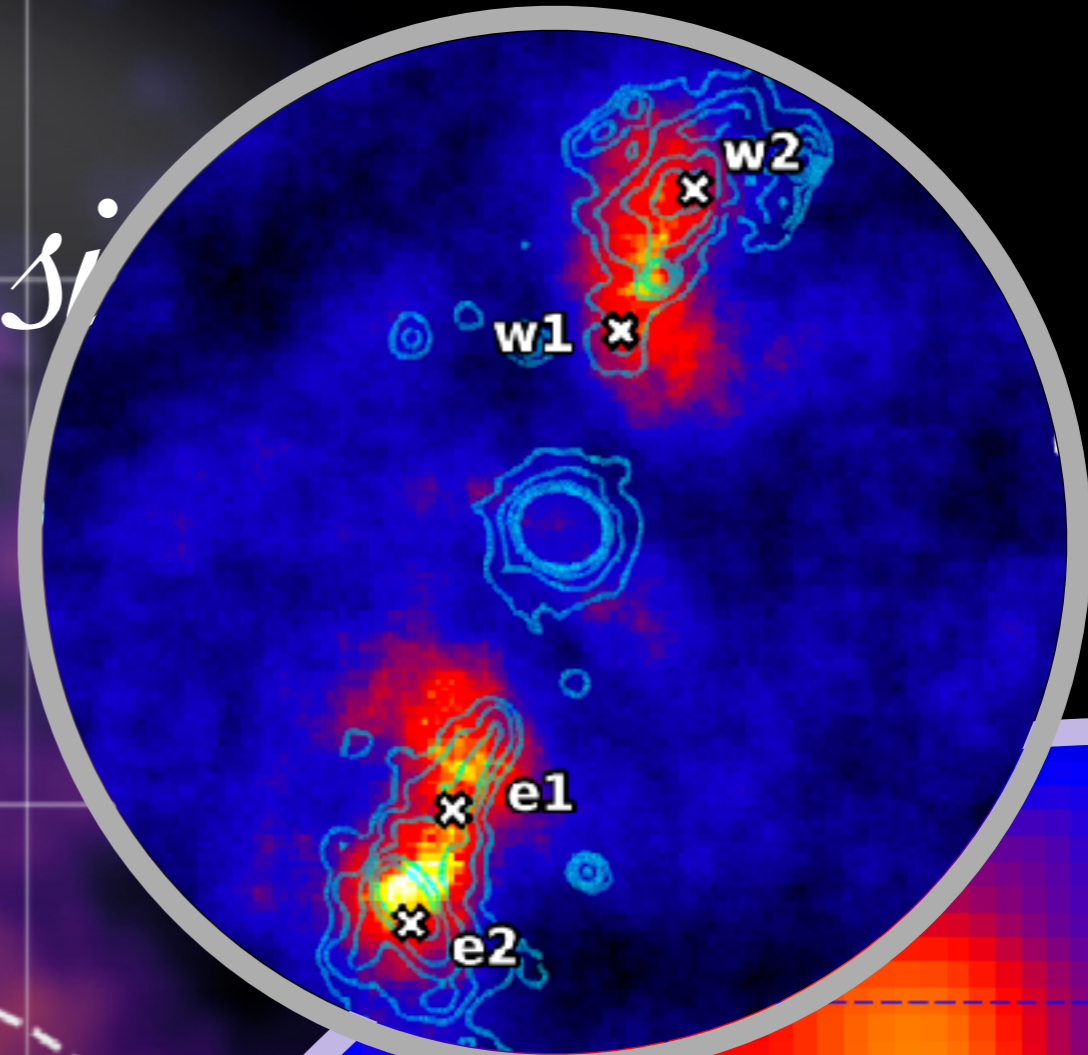
GRBs

etc., etc.



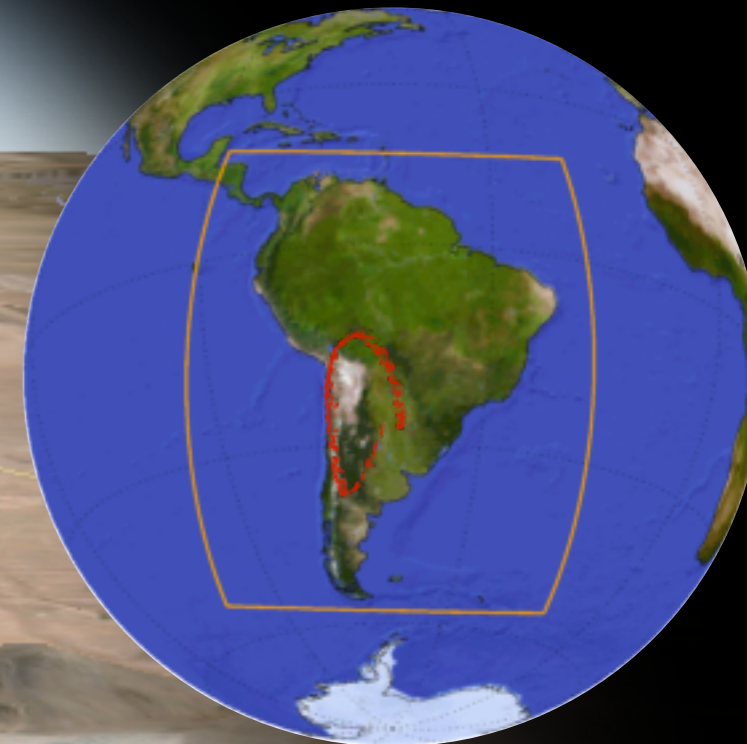
Conclusions

Discoveries from
HAWC, H.E.S.S.,
MAGIC, LHAASO,
and VERITAS

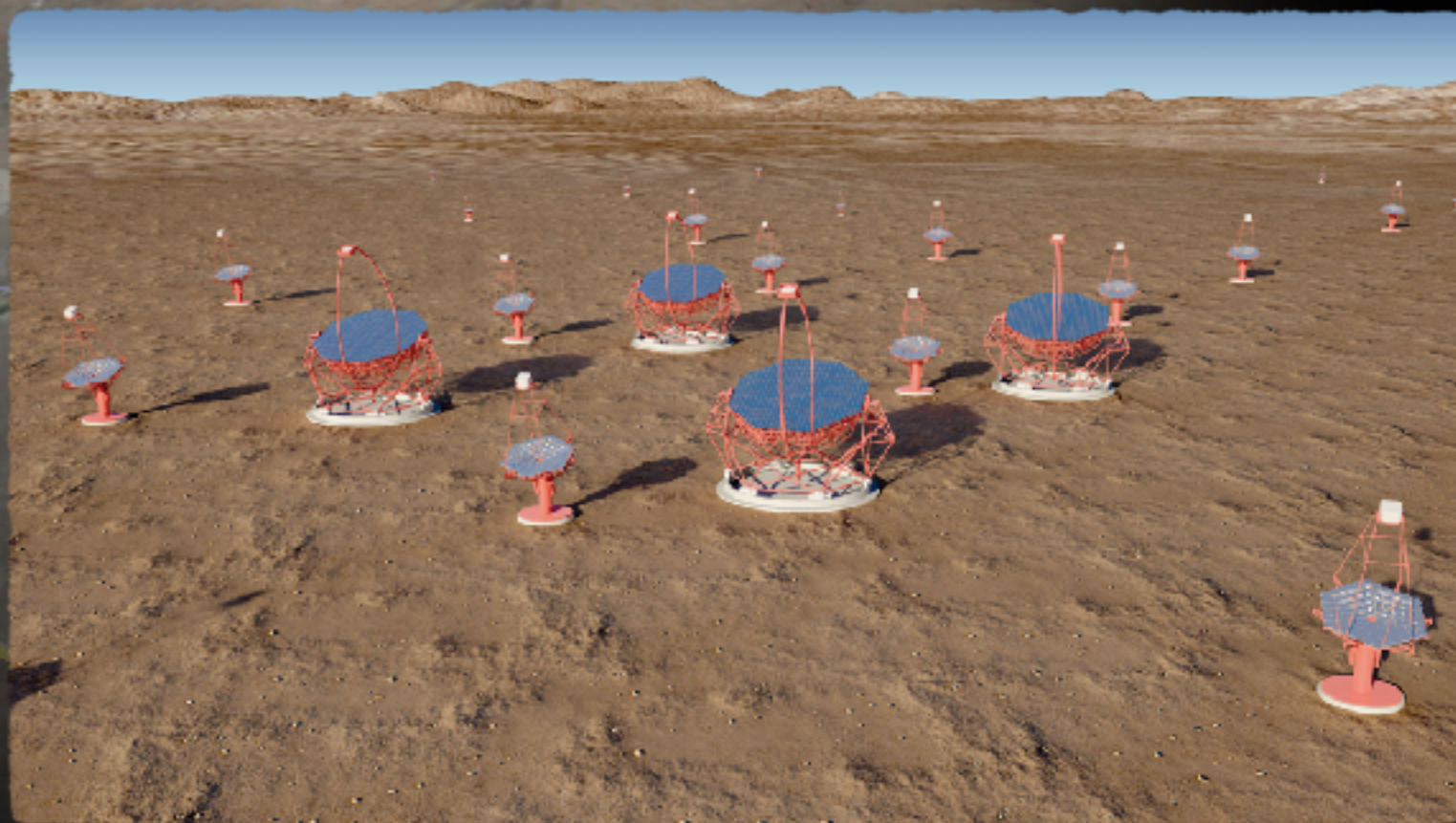


Conclusions

Discoveries from
HAWC, H.E.S.S.,
MAGIC, LHAASO,
and VERITAS
Next-generation
experiments!



HAWC-S QUBIC



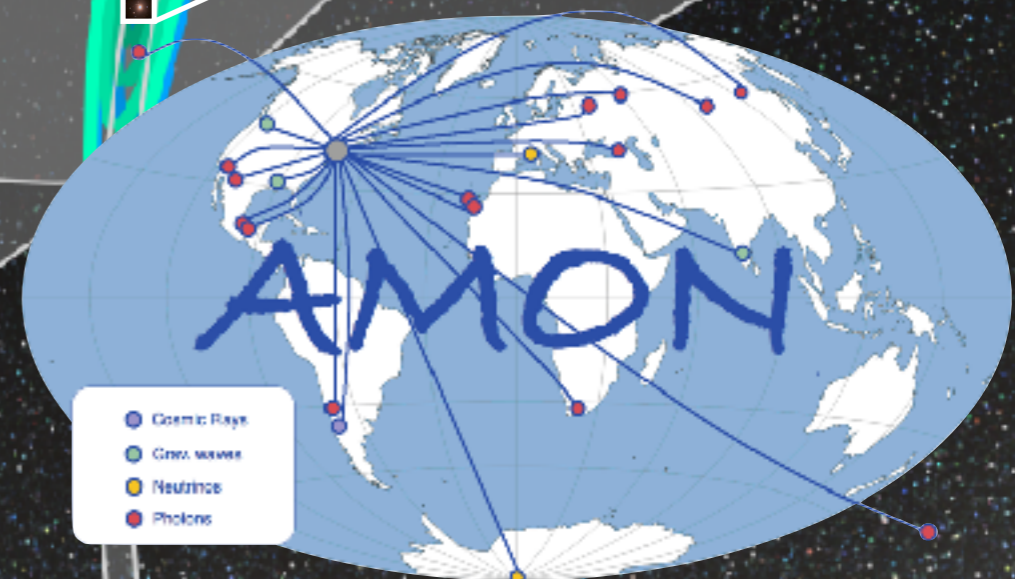
Conclusions

Discoveries from
HAWC, H.E.S.S.,
MAGIC, LHAASO,
and VERITAS

Next-generation
experiments!

New Era of
Multimessenger
Astrophysics

GW170817
DECam observation
(0.5–1.5 days post merger)



Thank you very much!