



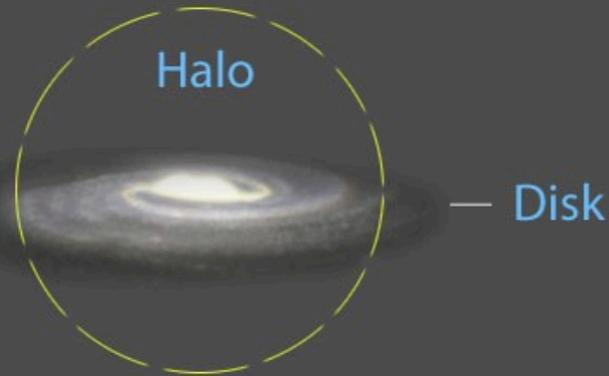
# ***Direct Detection of Dark Matter with DarkSide-20k***

**Paolo Agnes, GSSI**  
*RICAP 2022*  
*Rome, 7th Sept 2022*



# Expected rate in direct detection

Dark matter halo



## Dark Matter particles:

- are massive
- are stable
- non relativistic
- interact gravitationally
- electromagnetically neutral

**WIMP:**  
Weakly Interacting  
Massive Particle

Cross section

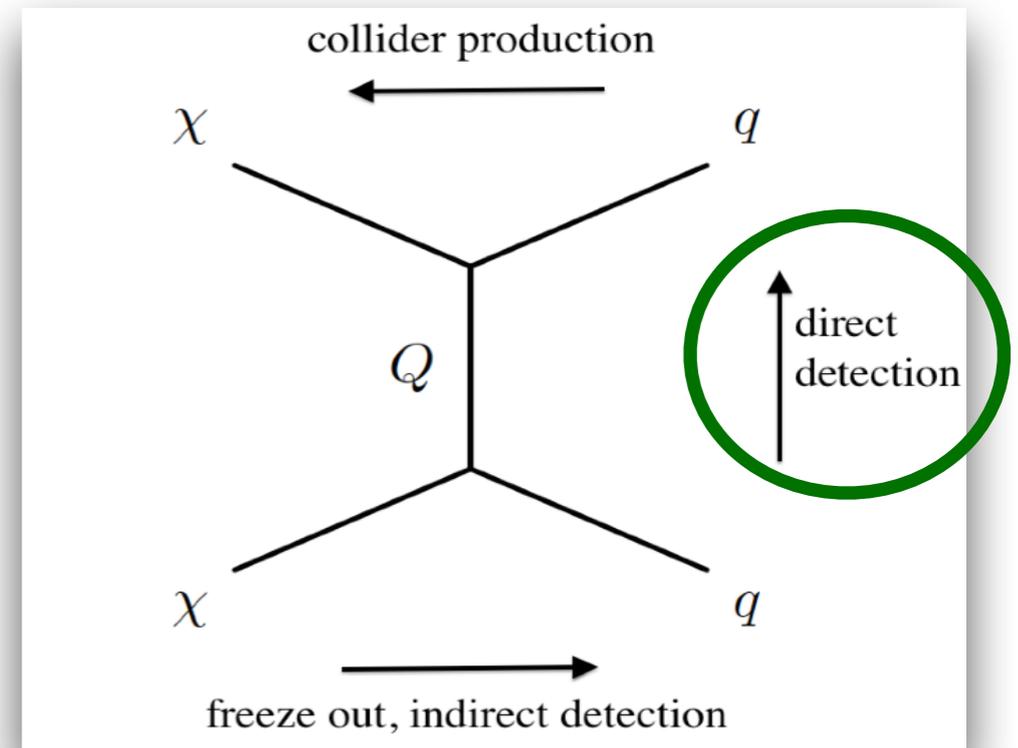
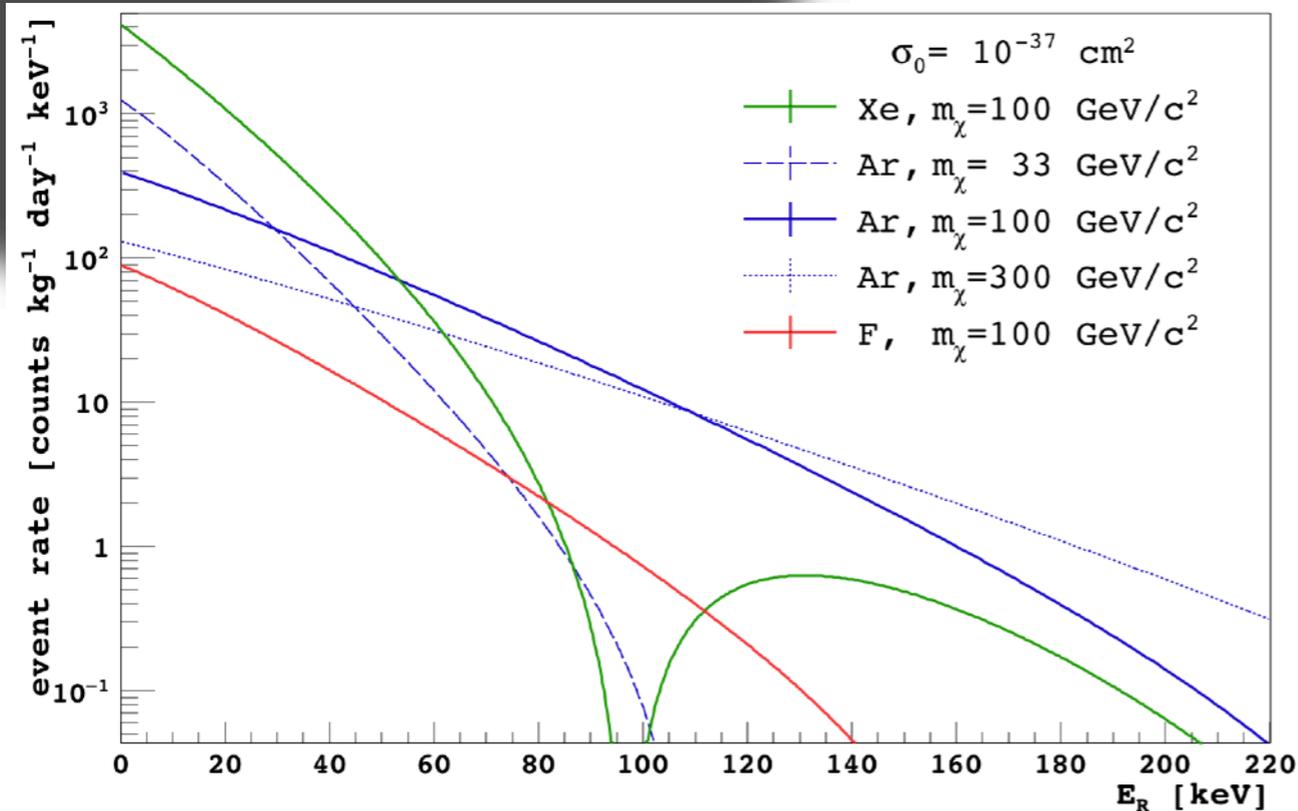
Nuclear Form Factor

Astrophysics

$$\frac{dR^{NR}}{dE_R} = \underbrace{\sigma_n}_{\text{Dark Matter Properties}} \underbrace{\frac{\rho_\chi}{M_\chi} \frac{m_A}{2\mu_{n\chi}^2} A^2}_{\text{Target}} \underbrace{F_A(q)^2}_{\text{Nuclear Form Factor}} \underbrace{\int_{v_{min}(E_R)}^{v_{esc}} d^3v \frac{f(v, v_E)}{v}}_{\text{Astrophysics}}$$

Dark Matter Properties

Target



# Noble liquids

- **Low energy threshold**

( $E < 100$  keV)

- **Large mass**

( $\sim 1$  event/tonne/yr @  $10^{-47}$  cm<sup>2</sup> in noble liquids)

- **Background suppression**

Deep underground

Passive/active shielding

Low intrinsic radioactivity

ER background discrimination

...

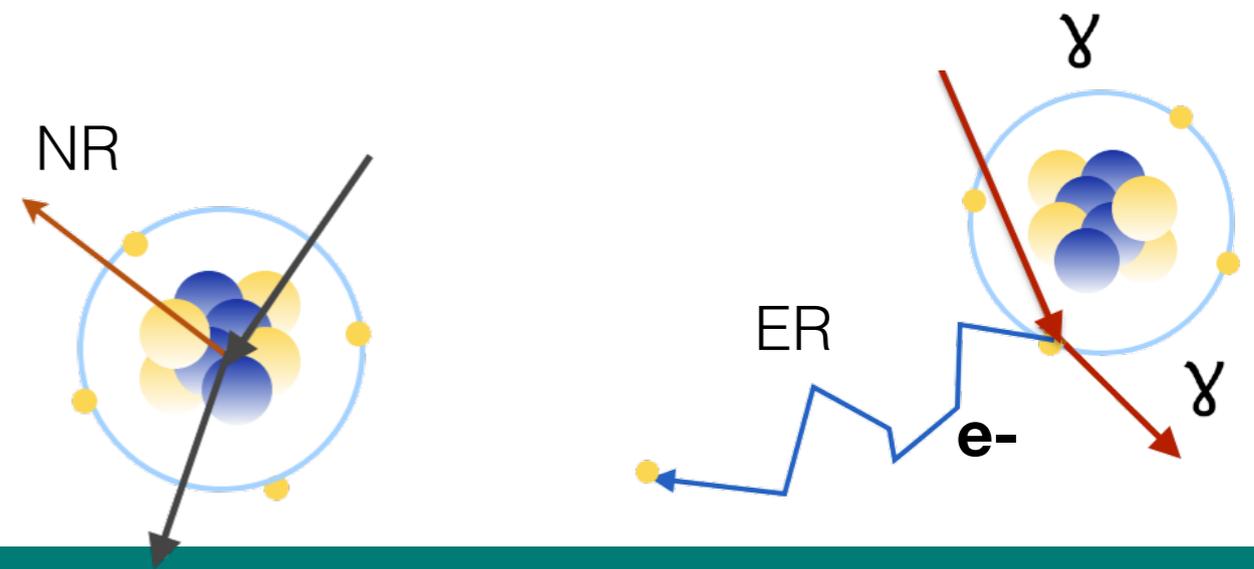
**Noble liquids are suitable targets:**

✓ scalable, easy to purify

✓ large ionization/scintillation yields ( $W \sim 10$  eV)

✓ ER recoil background discrimination

**Complementarity: great value in case of an excess**

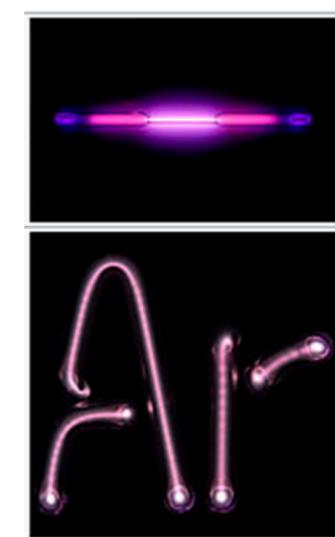


**NR backgrounds**

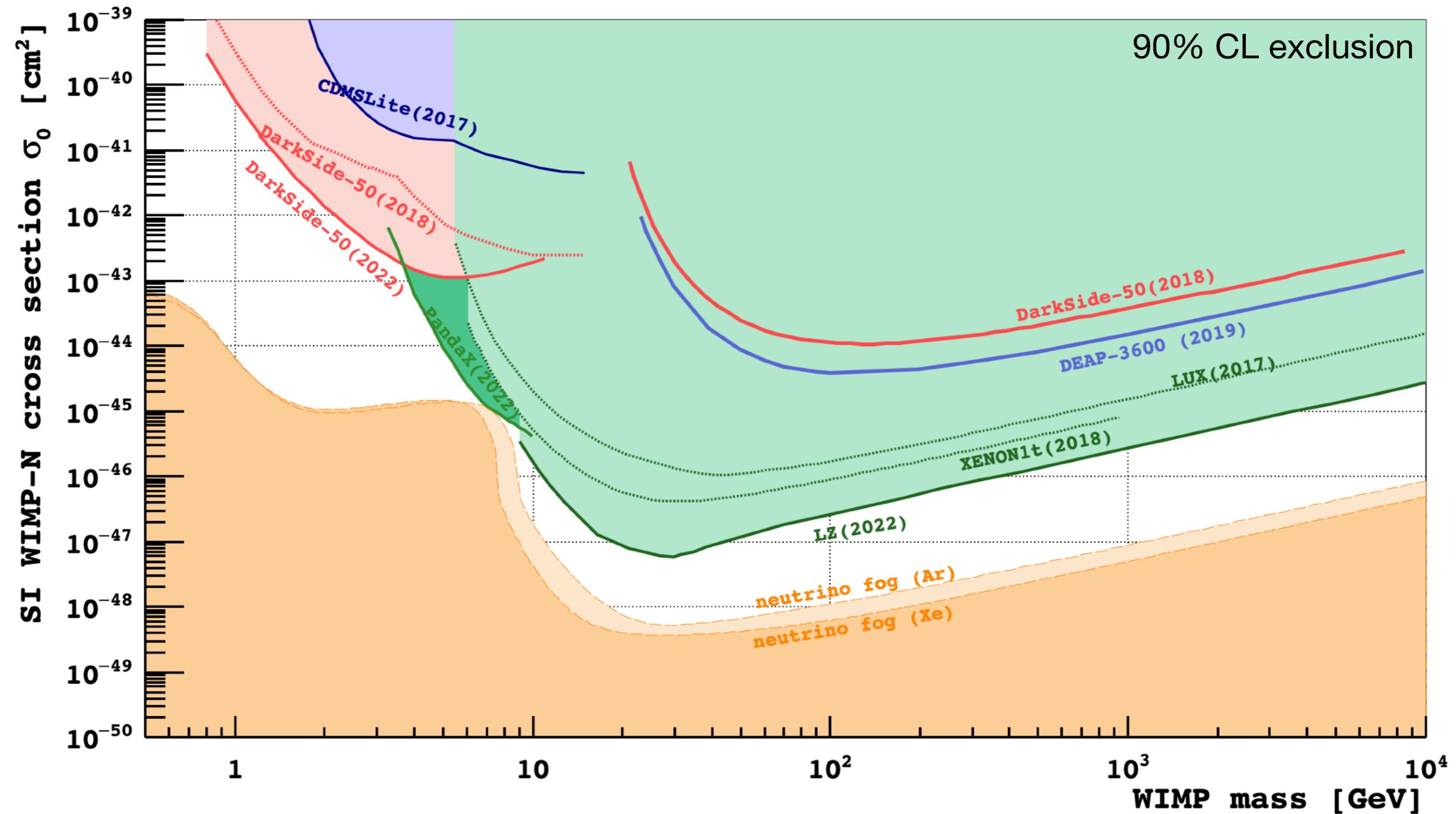
radiogenic and cosmogenic  
neutrons  
 $\alpha$ 's,  $\nu$ 's

**ER backgrounds**

internal and  
external  $\beta$ 's,  $\gamma$ 's



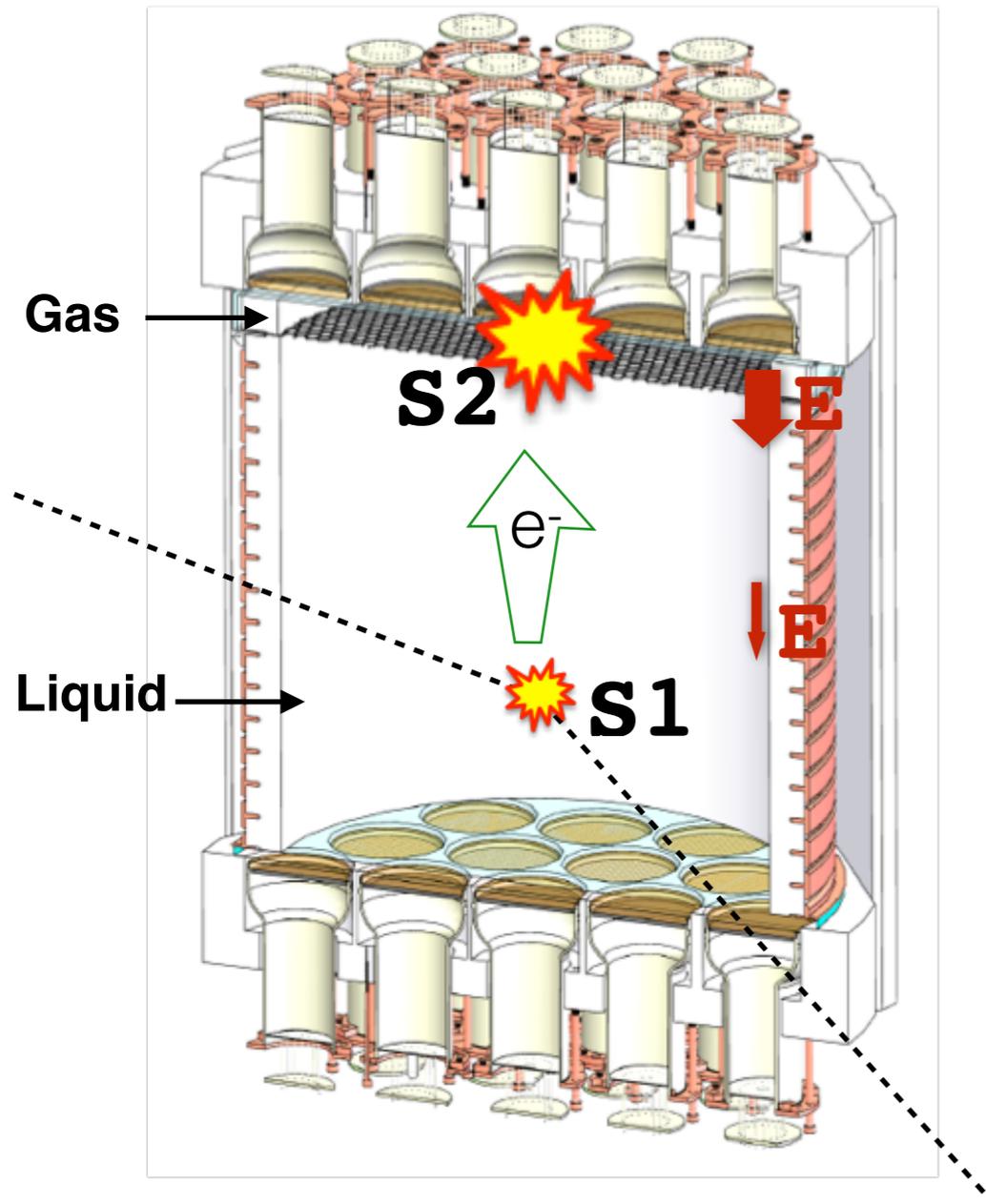
# Direct detection of dark matter (SI)



Noble liquids leading searches both at high- and low-mass

Potential of LAr in low-mass WIMP region

# Dual-phase TPCs — DarkSide-50



==> 3D vertex reconstruction  
(surface events, multi-sited events) !

Energy deposition

~~Heat~~

Excitation

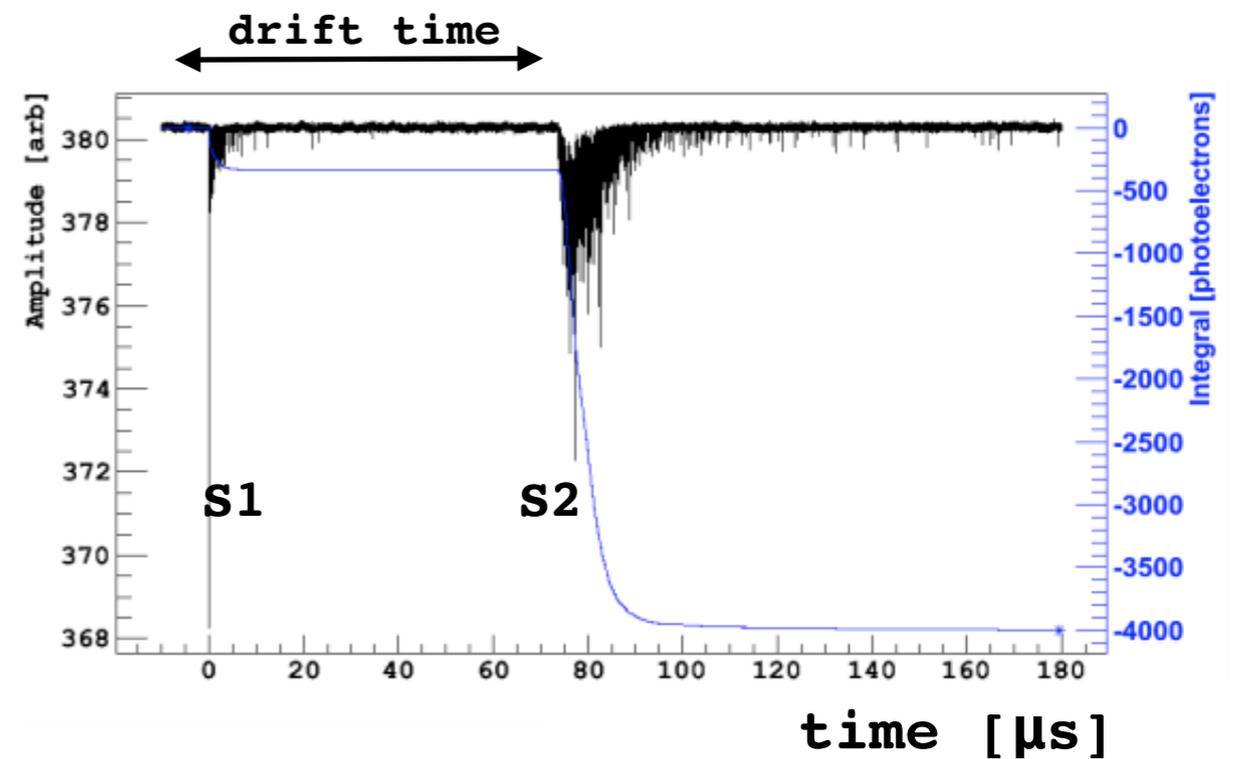
Ionization

## S1 and S2 Yields:

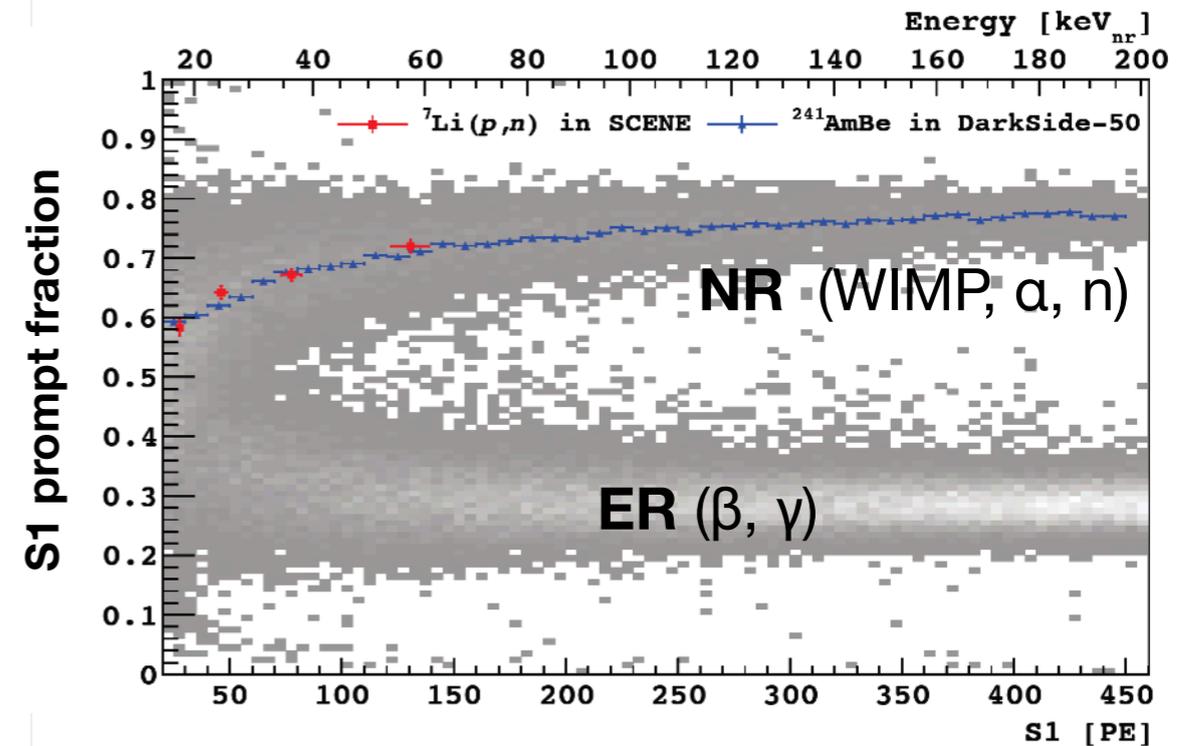
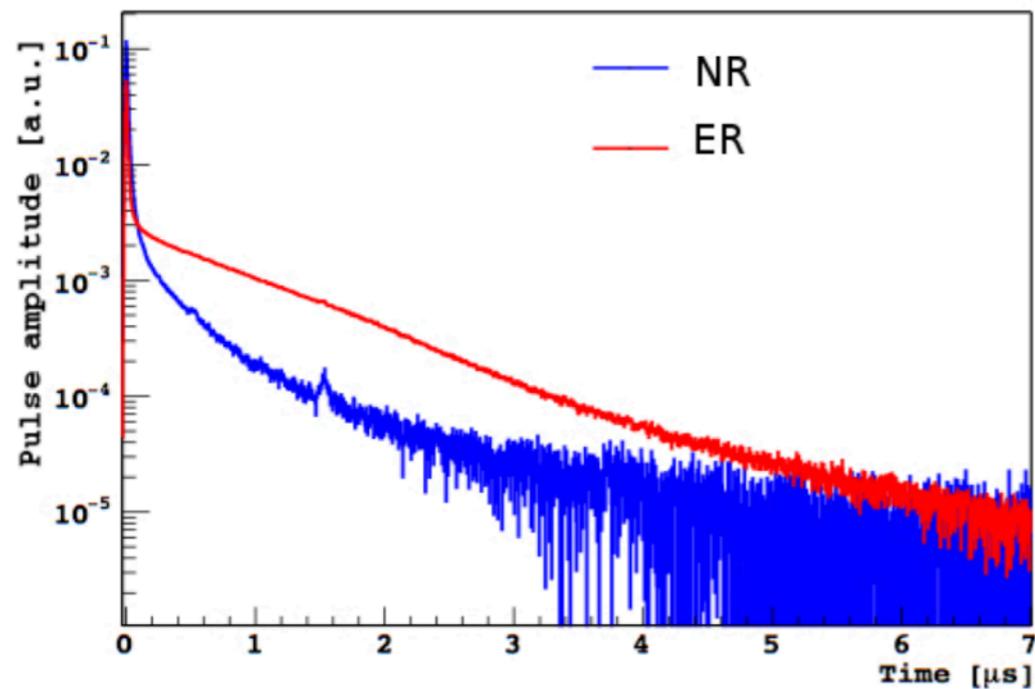
- S1 Yield  $\sim 7.9$  pe/keV at null field
- S1 Yield  $\sim 7.0$  pe/keV at 200 V/cm
- S2 yield  $\sim 23$  pe / e<sup>-</sup>

**Electron lifetime > 5 ms**

Maximum drift time: 376  $\mu$ s



# Pulse Shape Discrimination in Liquid Argon



## Liquid argon scintillation:

~ 40-50  $\gamma$  (128 nm) per keV deposited

DEAP: 6.5 PE / keV<sub>ee</sub>

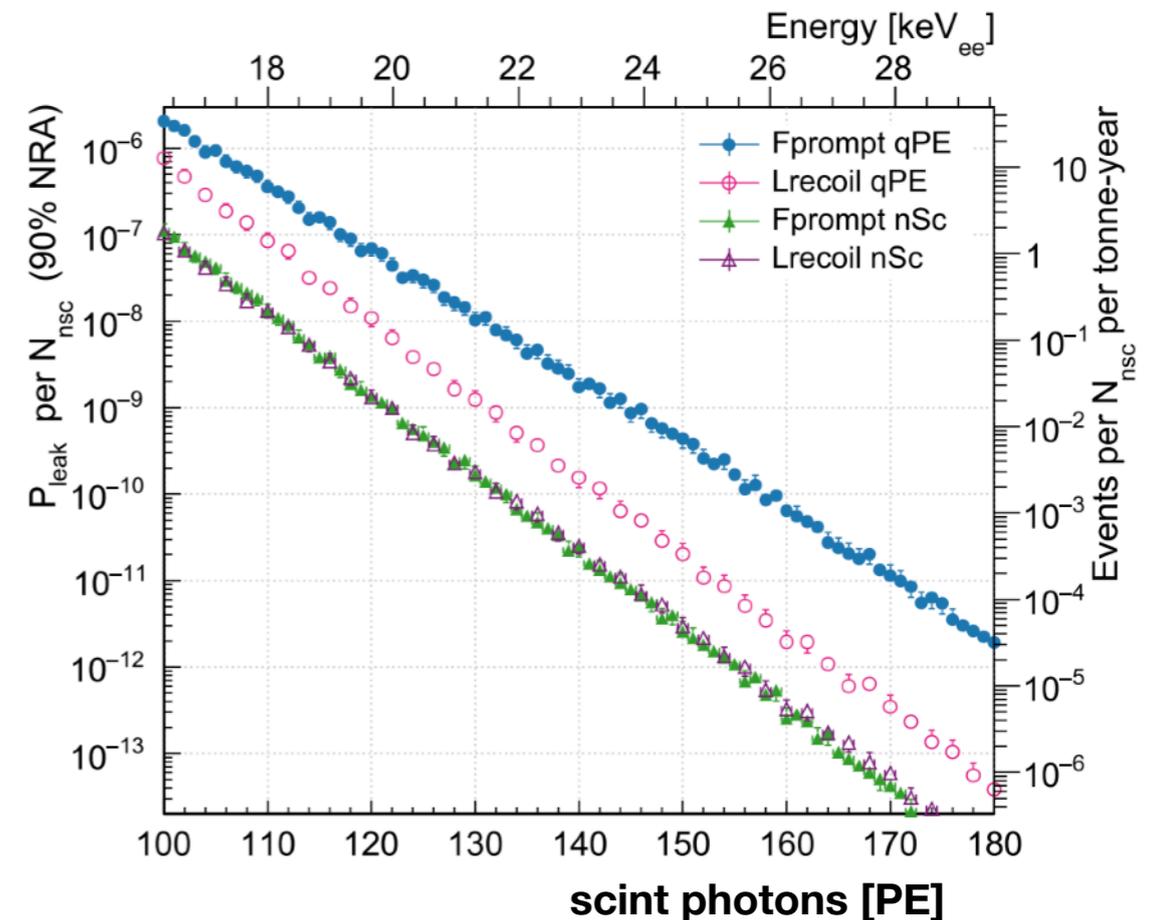
DarkSide-50: 7.9 PE / keV<sub>ee</sub>

Excited dimers decay with two decay constants,

$\tau_{\text{fast}} \sim 6 \text{ ns}$  and  $\tau_{\text{slow}} \sim 1.6 \mu\text{s}$

**ER and NR excite fast and slow in different proportion**

**==> ER Rejection factor: ~ 10<sup>8</sup> in LAr**



DEAP-3600: *Eur. Phys. J. C* 81,823 (2021)

# Argon from underground

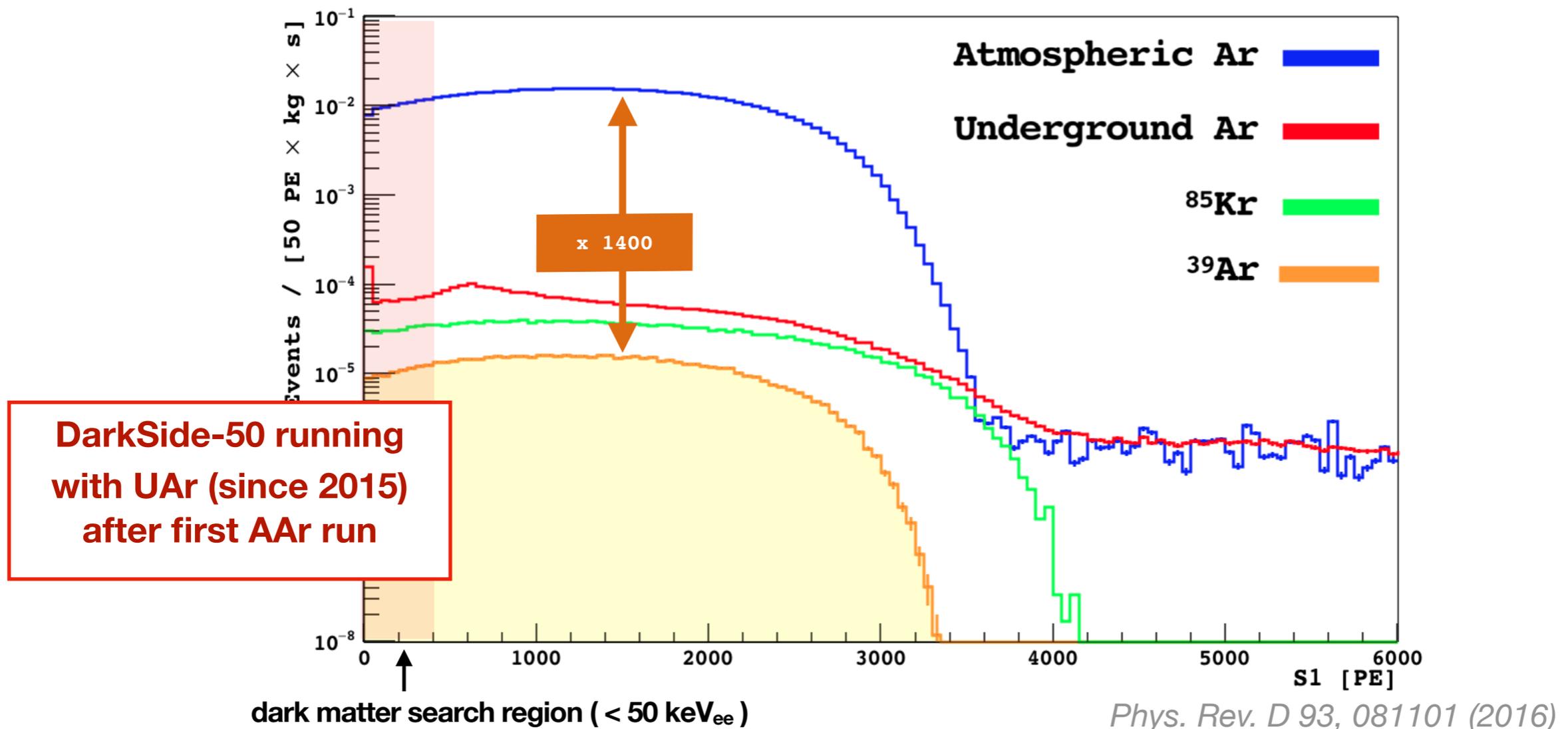
**$^{39}\text{Ar}$**  is produced by cosmic rays in the atmosphere.  **$\beta$ -decay** with  **$Q = 565 \text{ keV}$** ;  $\tau_{1/2} = 269 \text{ yr}$

►  **$^{39}\text{Ar}$  activity in atmospheric argon** ( $\sim 1 \text{ Bq/kg}$ ): limiting dual-phase target mass

**==> extract argon from underground** ( $\text{CO}_2$  well in Colorado) !

►  **$^{39}\text{Ar}$  activity in underground argon** ( $0.73 \pm 0.10 \text{ mBq/kg}$ )

► Possibly smaller: presence of  $^{85}\text{Kr}$  indicate possible presence of atmospheric air



# The DarkSide Program

## Background suppression

- Ultra-low background materials
- **Depleted Liquid Argon**
- Low background photo-detectors
- Low background material components

## Active Shielding

- **Active Neutron Veto**
- Water Cherenkov against muons (WCD)

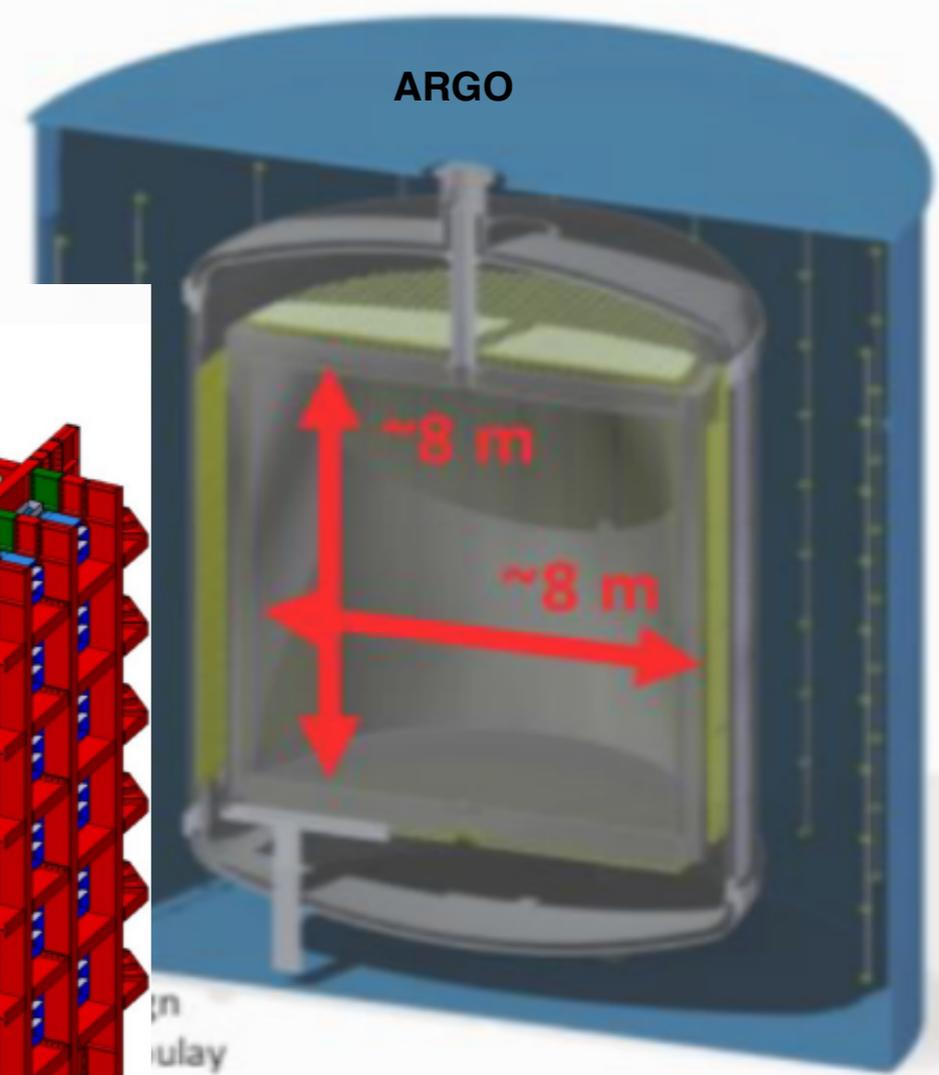
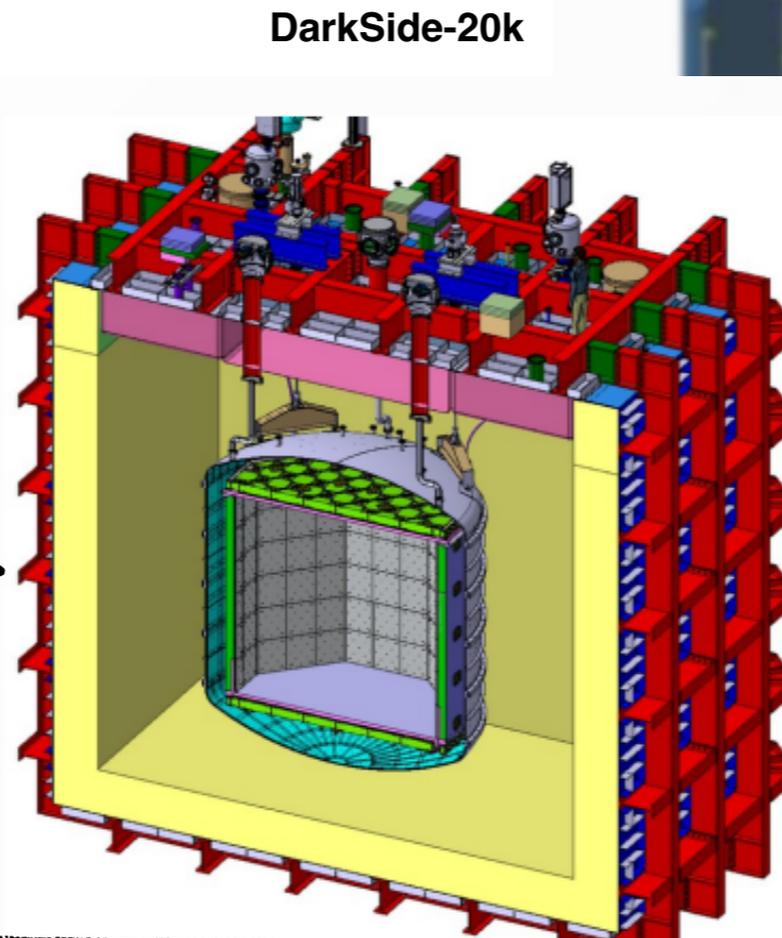
## Main goal: bg-free exposure

## Background identification

- **Pulse Shape Discrimination (PSD)**
- Ionization/scintillation ratio
- Position reconstruction (surface events)
- Multiple scatters within the TPC



x 1000



2011      2013-2018 at LNGS      2025+      2030+

0.03 t.yr exposure      200 t.yr target exposure      3000 t.yr target exposure

# DarkSide-50 — High-mass WIMP results

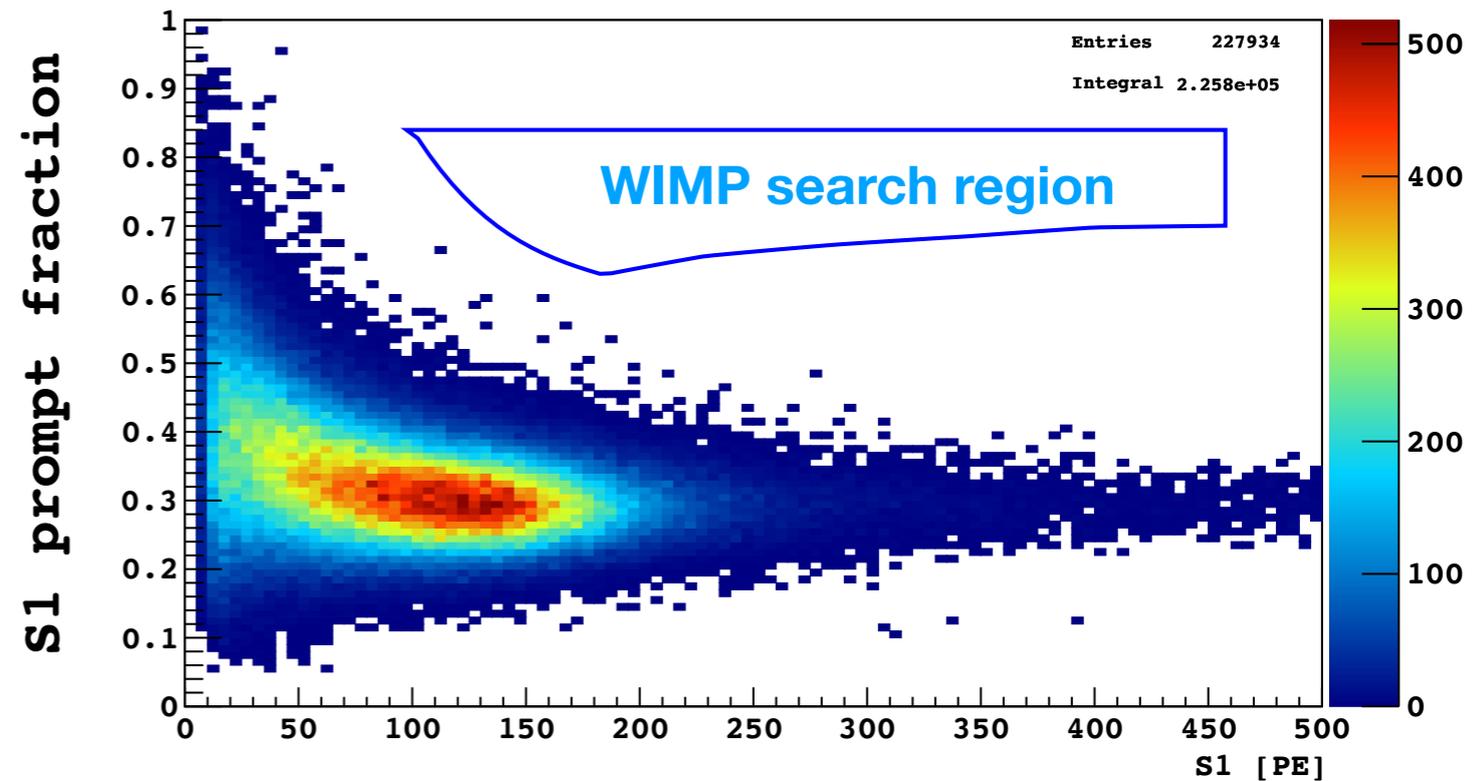
## Blind analysis published in 2018

- 524 live-days
- Use first 70 days of UAr dataset to tune cuts
- **Minimise backgrounds** while maximising acceptance to NR

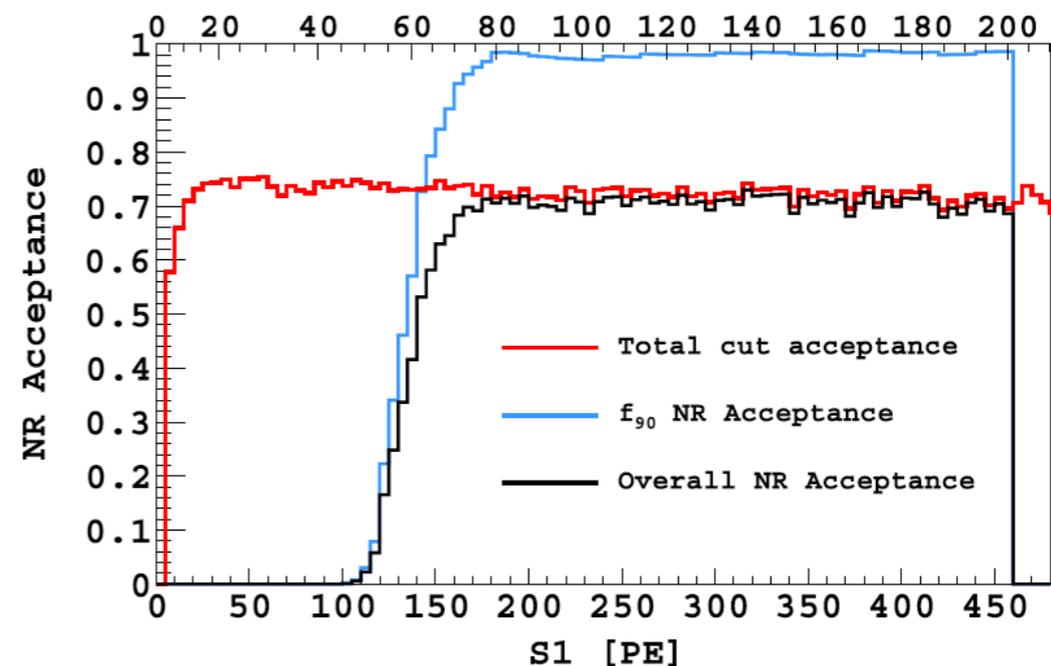
**$17.10^3$  kg.d: background free!**

*Expected backgrounds in ROI, before opening box*

NR	surface alphas	0.001
NR	cosmogenic neutrons	<0.00035
NR	radiogenic neutrons	<0.005
ER	electron recoil	0.08
		$0.09 \pm 0.04$



*NR acceptance after all cuts. Threshold driven by PSD:*



# DarkSide-50 low mass

Look at the **ionization only** spectrum  
 ( $W_{\text{ion}} = 23.5 \text{ eV}$ , gain in the gas: 23 PE/e<sup>-</sup>)

**Below 3 keV<sub>ee</sub>**: give up the scintillation signal  
 (too small to trigger the detector), and thus

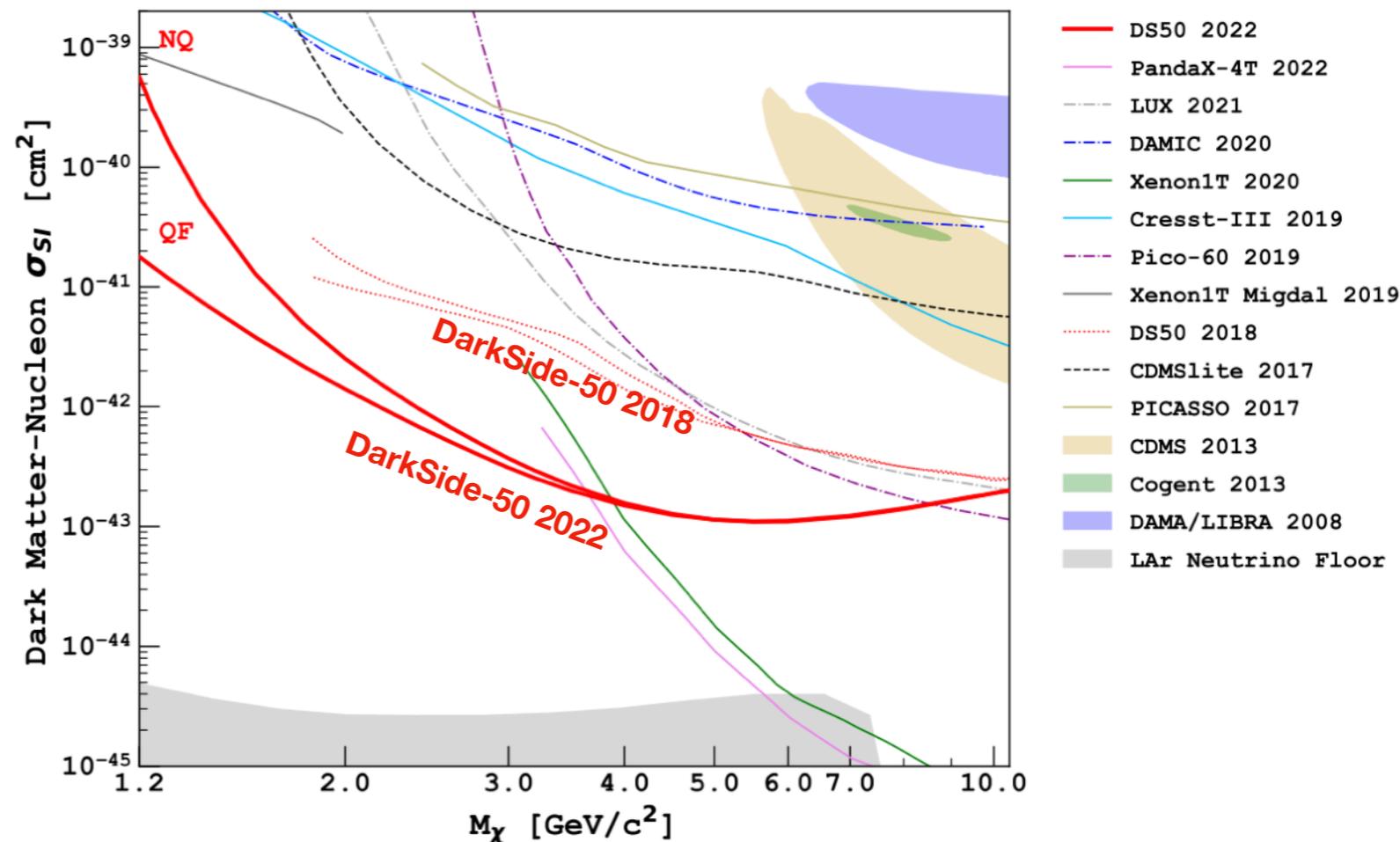
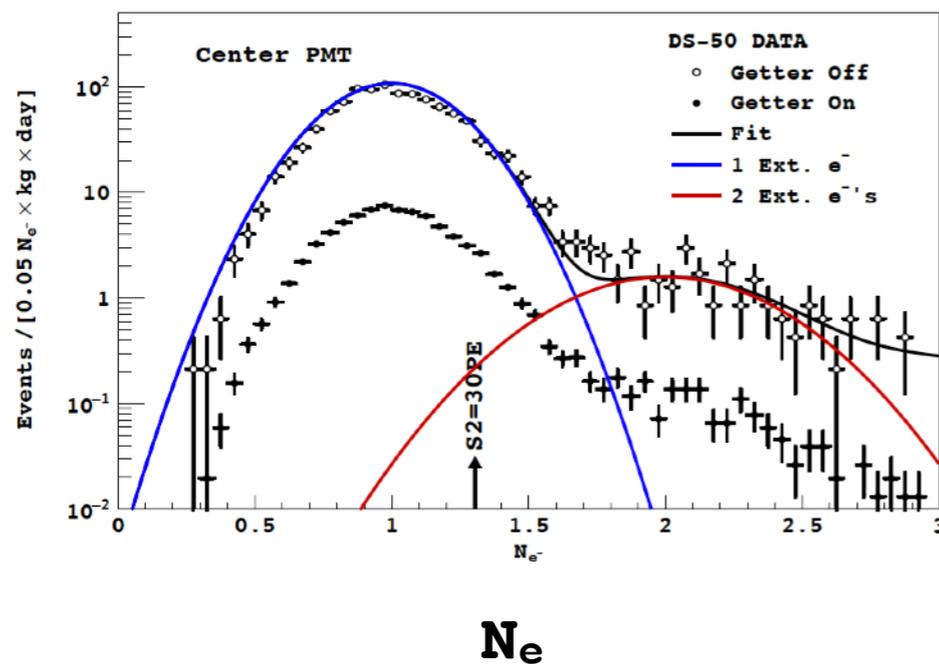
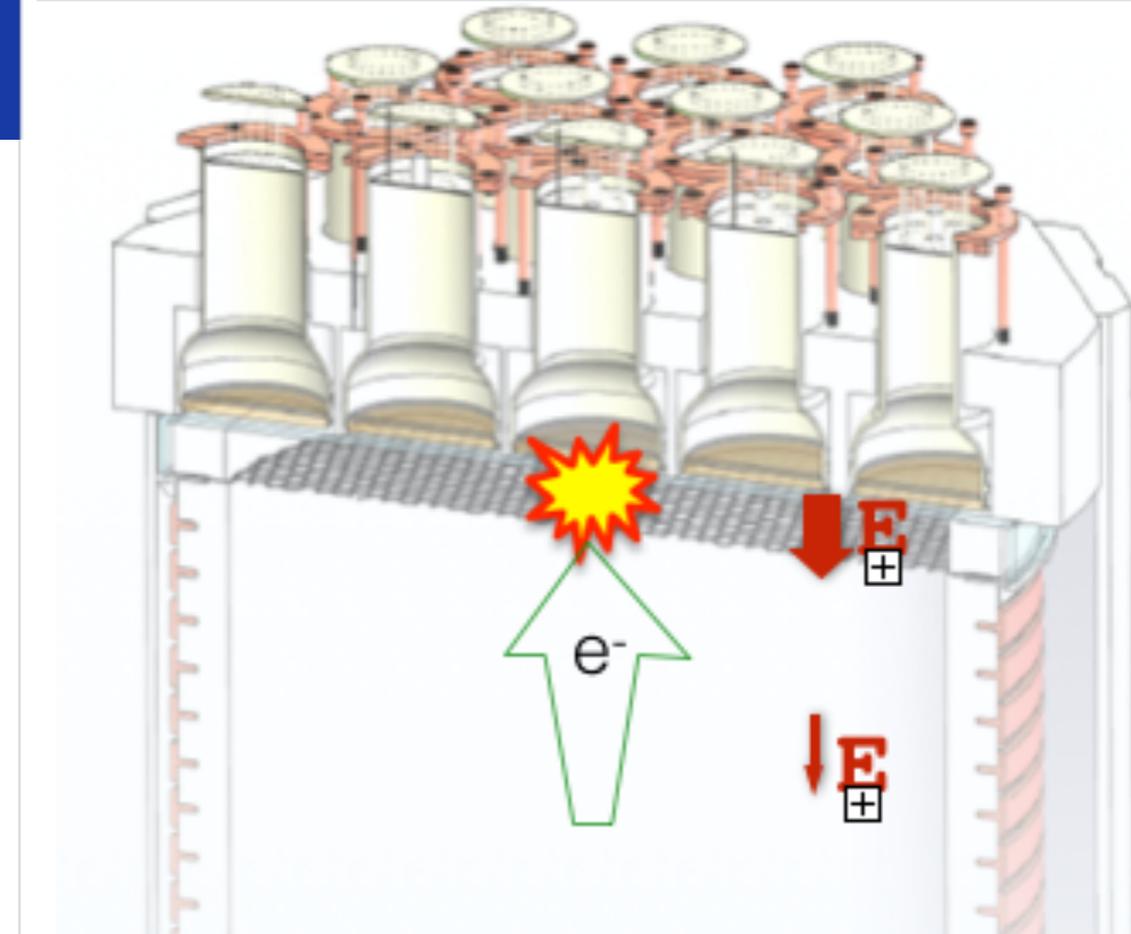
- **minimal fiducialization** (only radial)
- **no PSD**

**No more background free**

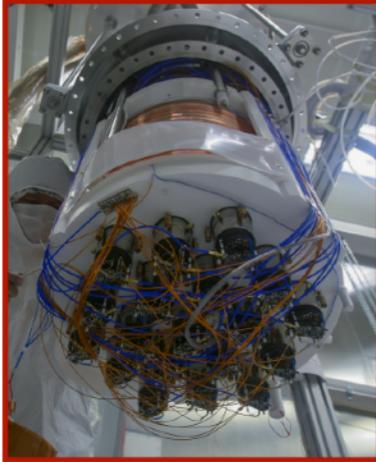
==> **Background model** for DarkSide-50

First analysis in 2018, recently updated !

- > WIMP-N 2207.11966
- > Migdal effect 2207.11967
- > WIMP-electron 2207.11968



DarkSide-50 @ LNGS



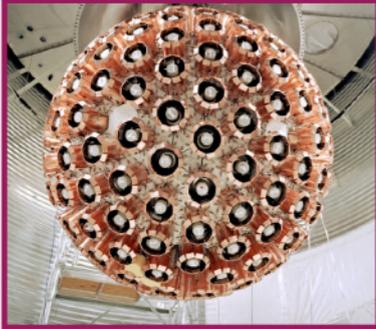
MiniClean @ Snolab



ArDM @ Canfranc

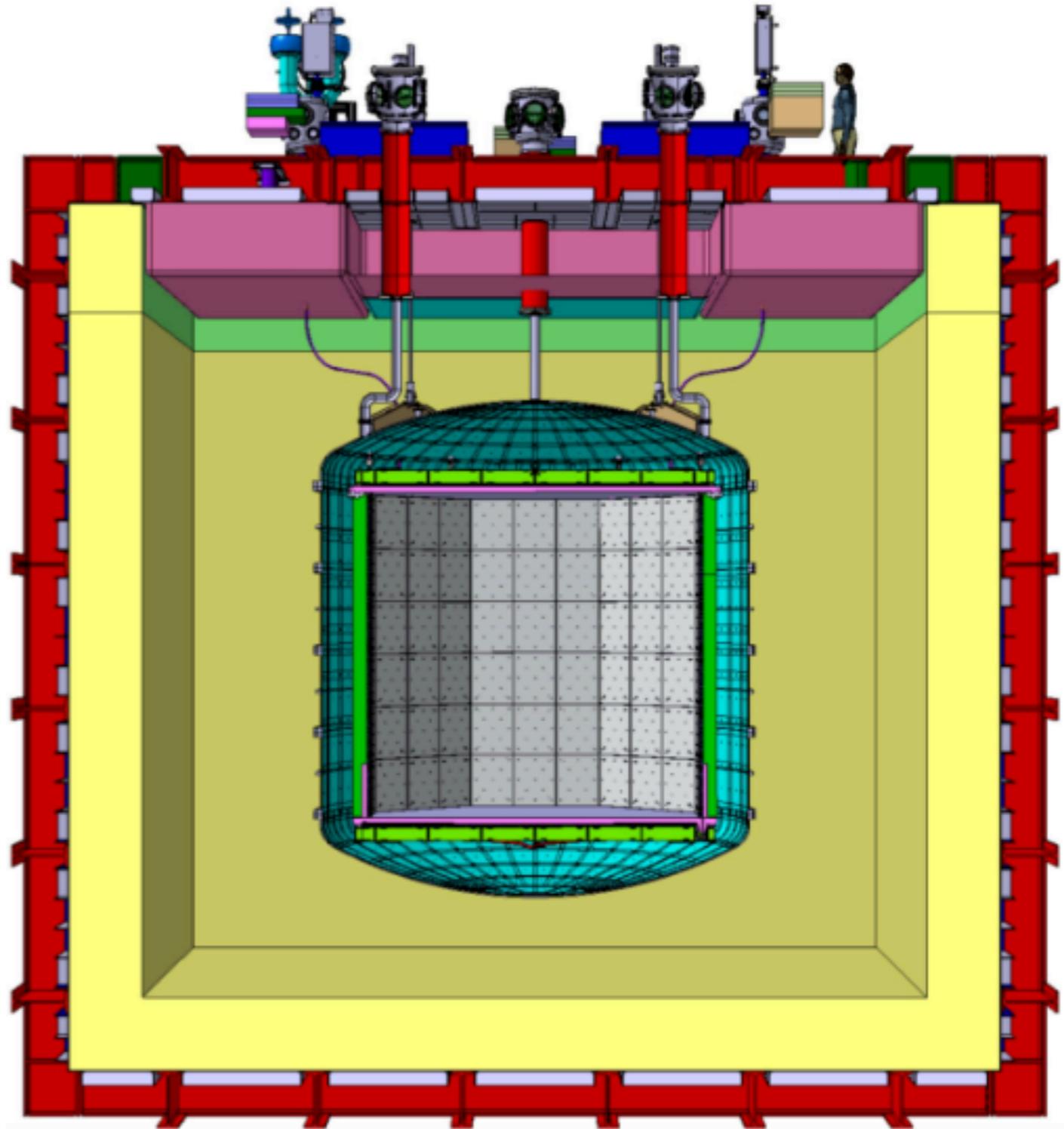


DEAP @ Snolab

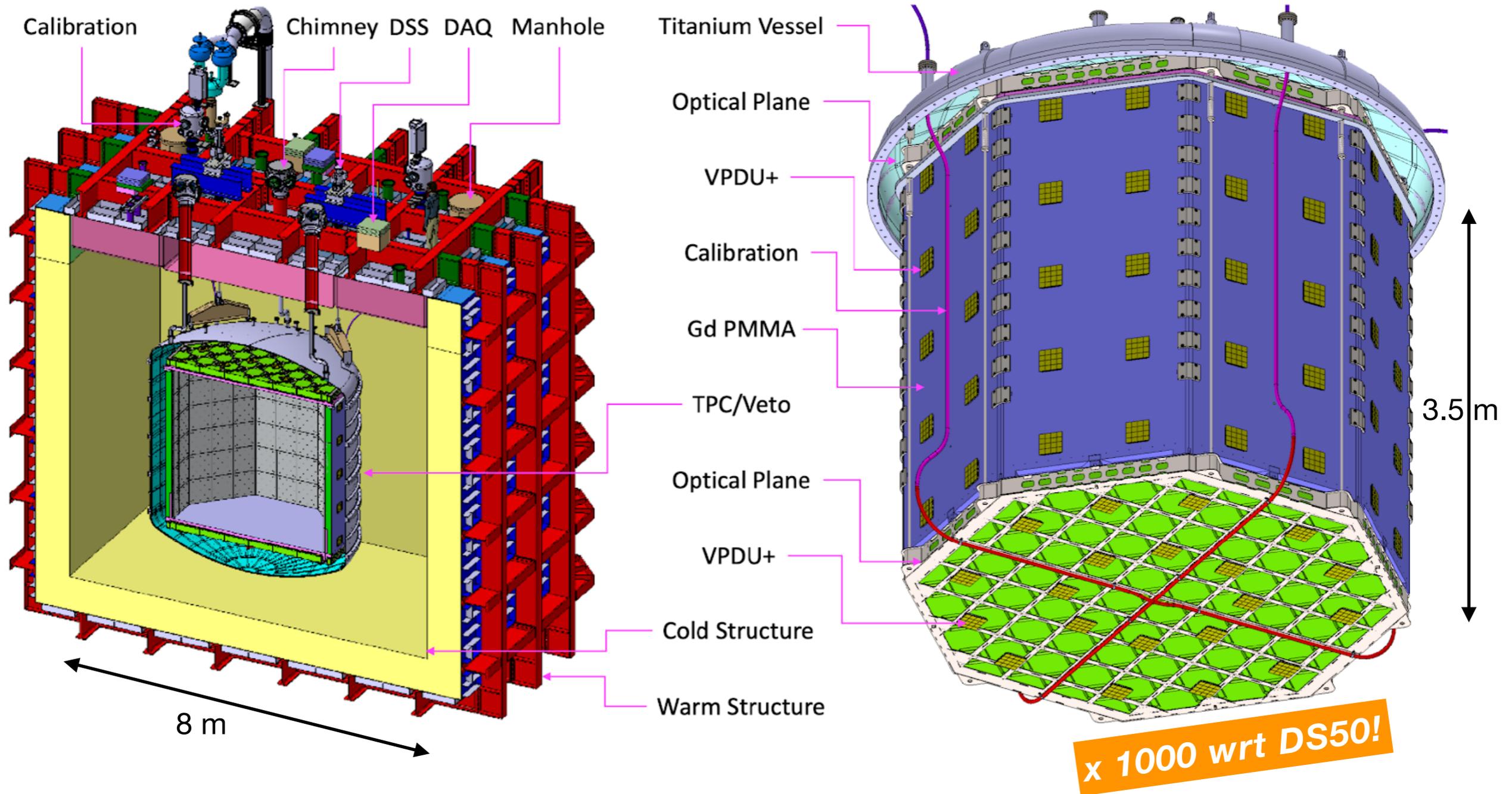


# DarkSide-20k

by the Global Dark Matter Argon Collaboration

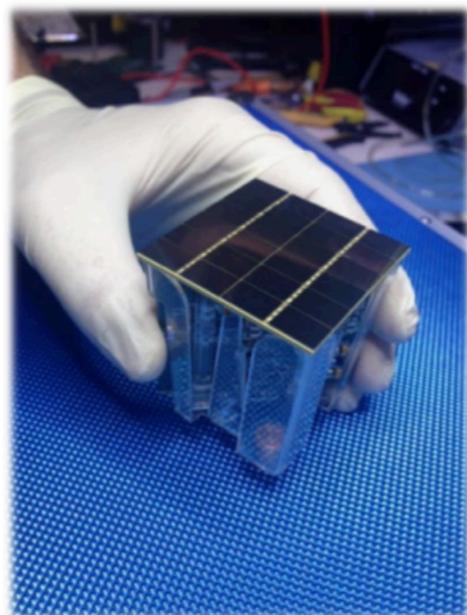


# DarkSide-20k



- To be installed in Hall C at LNGS
- Hosted inside a 700 t AAr LAr bath, in a cryostat à la ProtoDUNE
- **Target: 50 t UAr as WIMP target**
- **Veto: 35 t UAr + custom developed Gd-PMMA**, optimized for radiogenic neutrons
- **Novel readout system for the scintillation light**, based on **grouped SiPM arrays (> 25 m<sup>2</sup>)**

# From PMTs to SiPM arrays

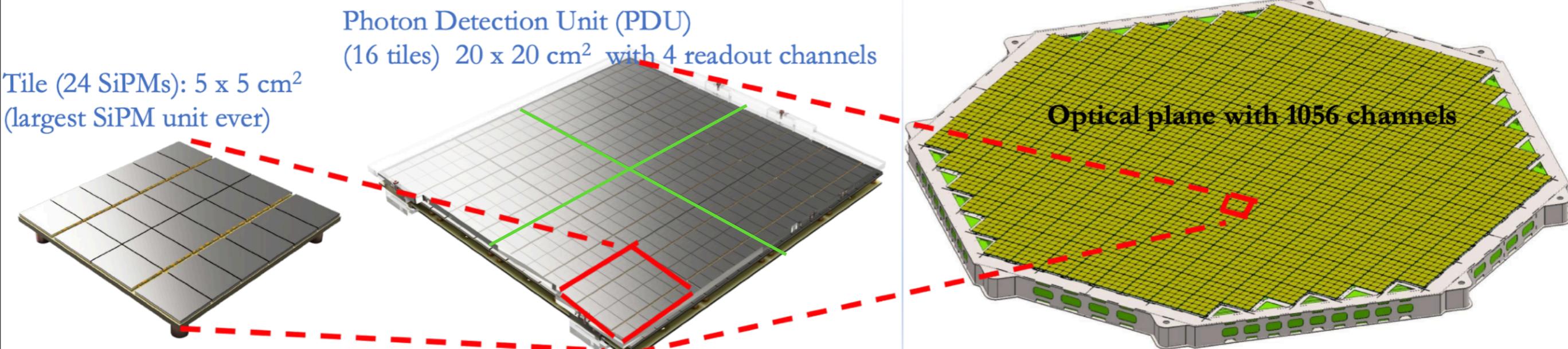


## PROS

- Cryogenic temp stability
- Better single photon resolution
- **Higher photo-detection efficiency**
- Low voltage operation
- **Lower background** (Si intrinsically radiopure)
- Lower cost

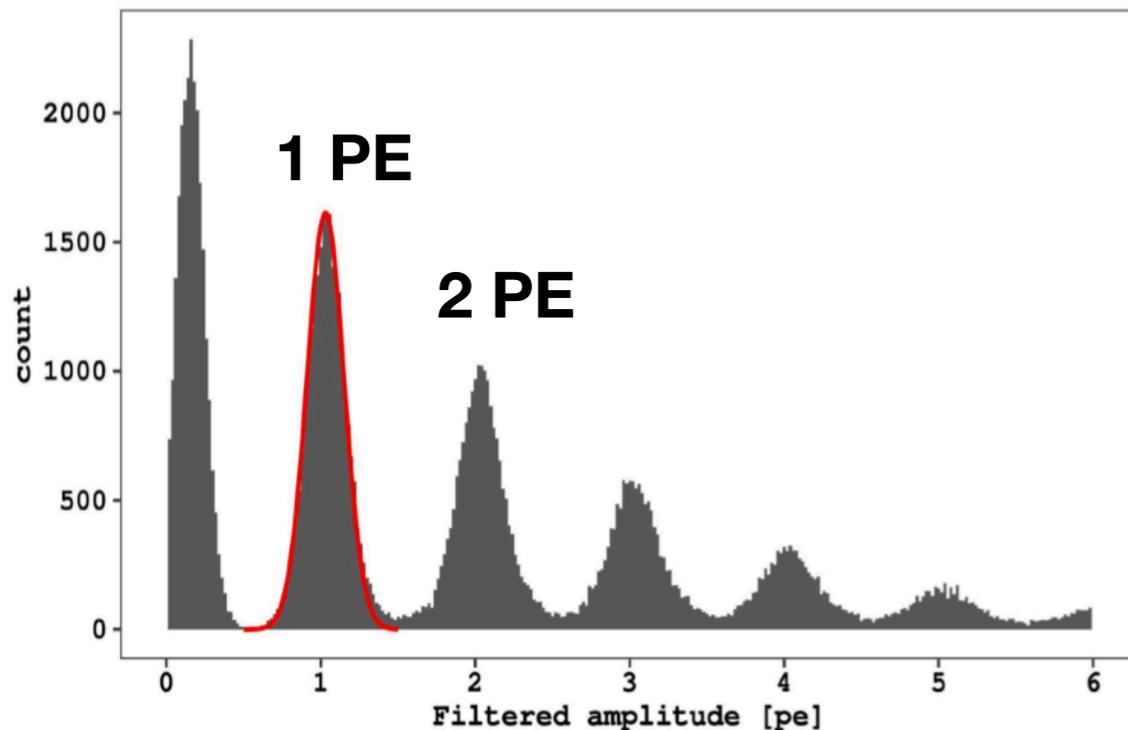
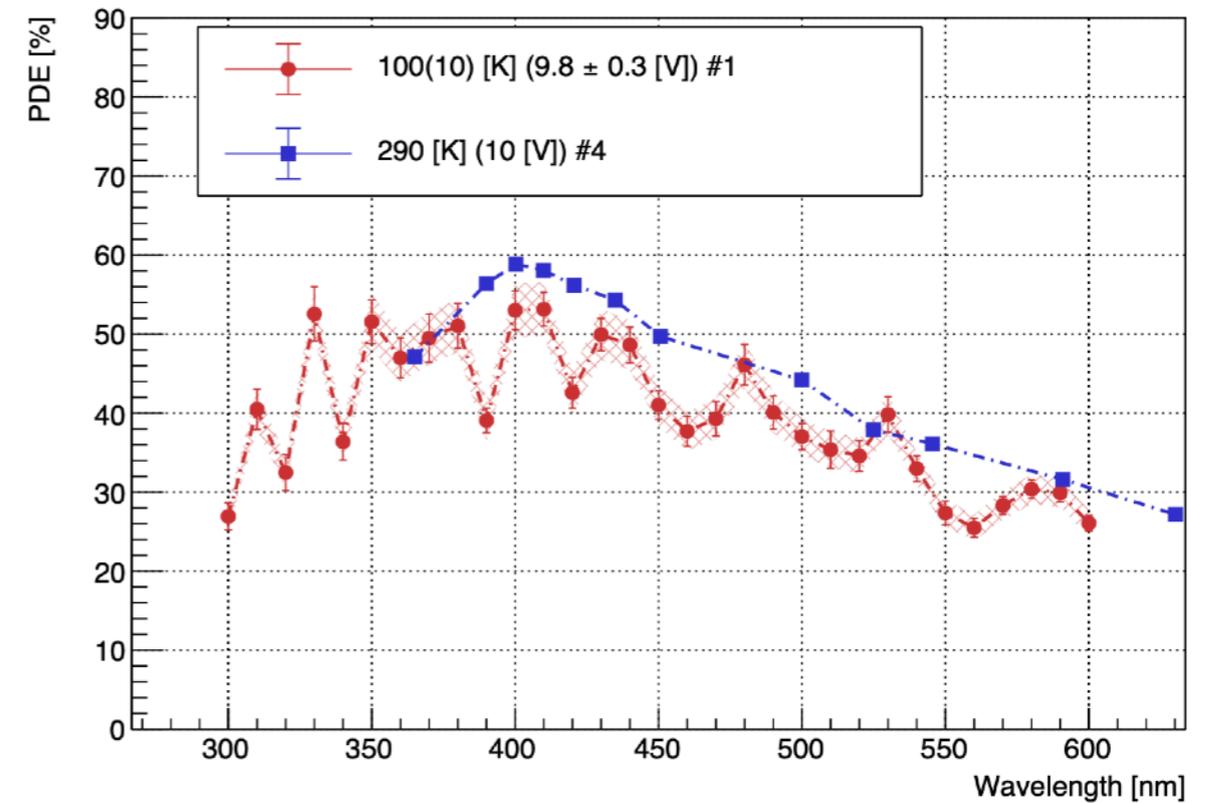
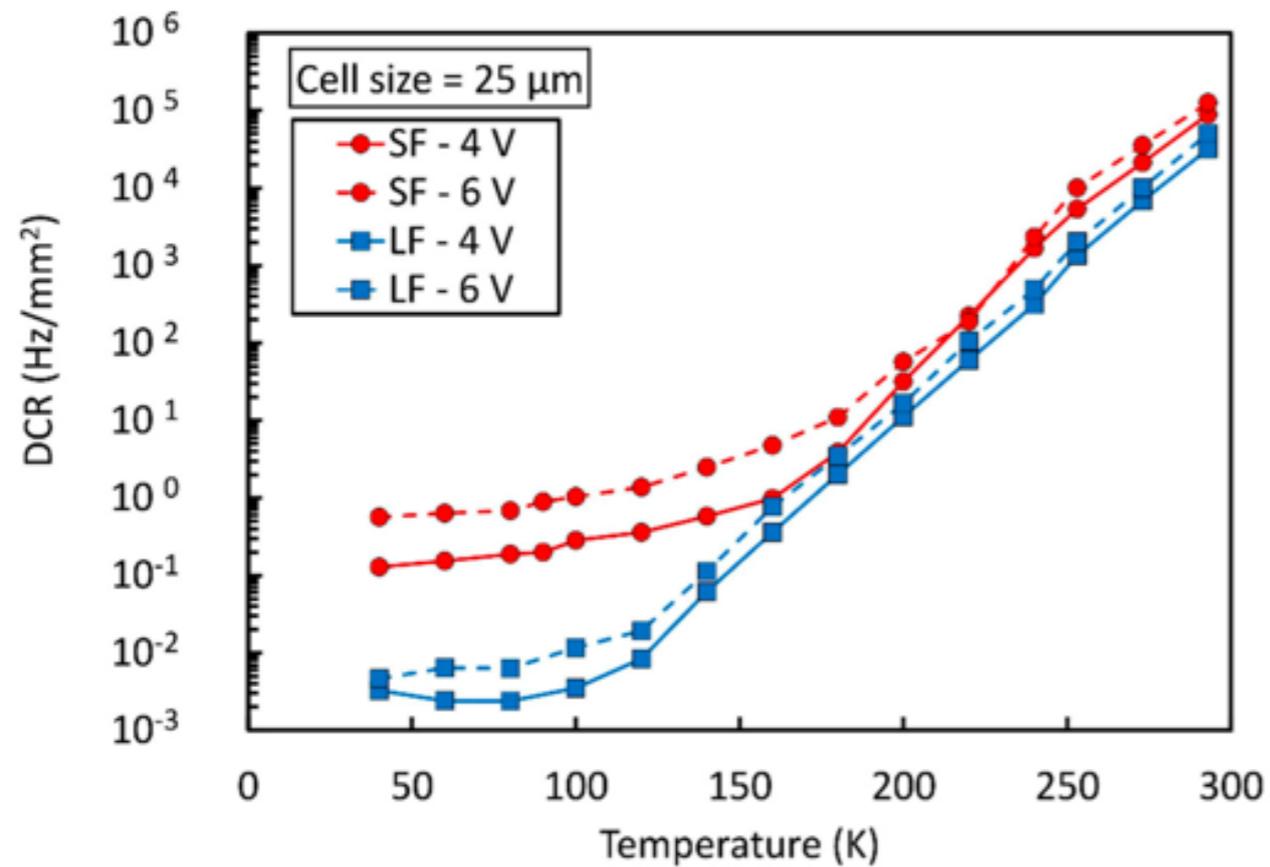
## CONS

- Small area  $\approx \text{cm}^2$  (group them)
- High dark rate (solved, + operated at 87K)
- High output capacitance for large devices ( $\sim 0.5 \text{ us}$  recharge)



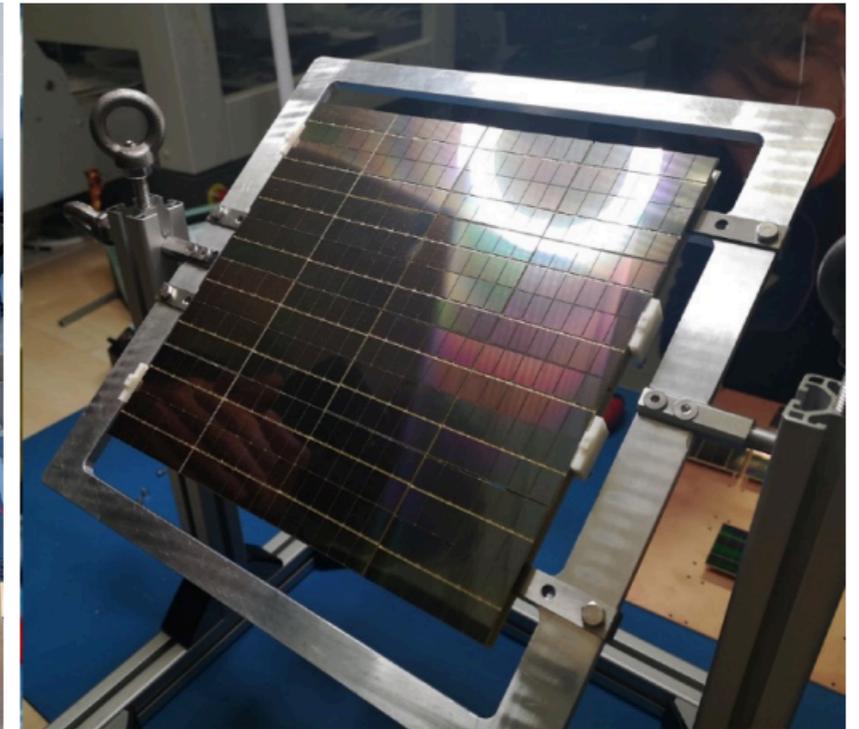
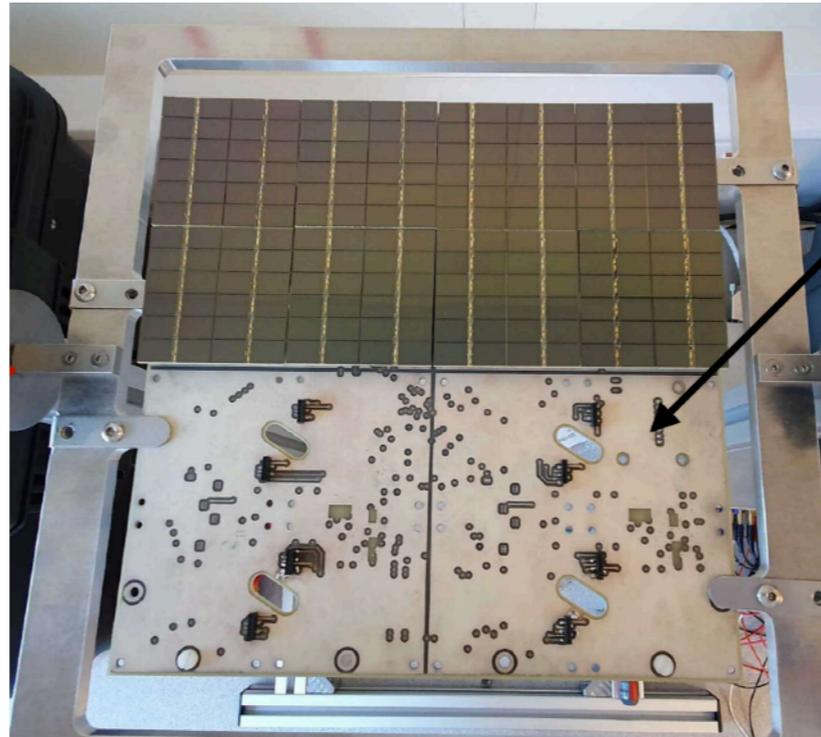
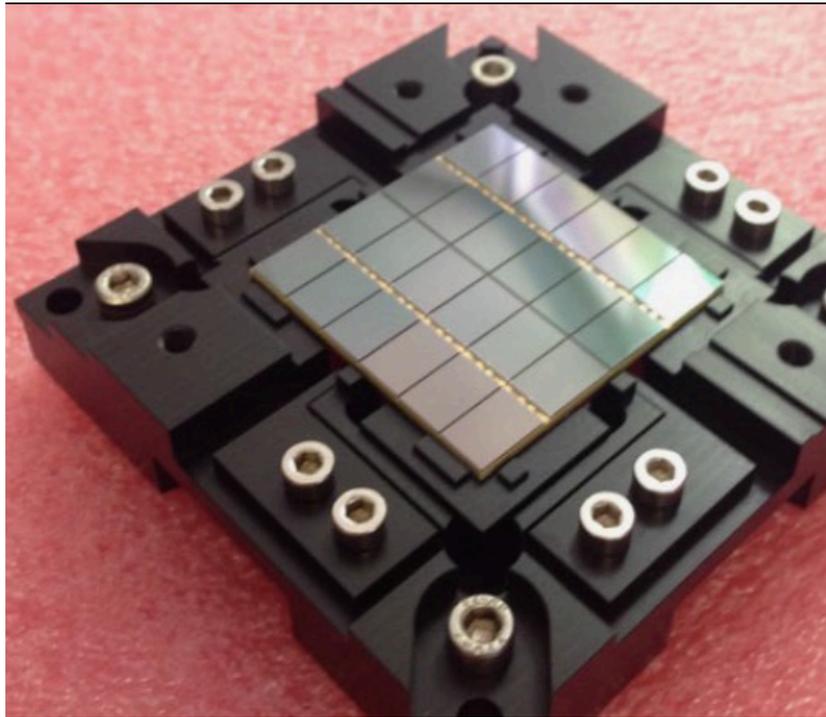
2 optical planes for the TPC  
+ 480 channels to instrument the UAr veto

# PDU design, results of years of R&D



parameter	spec required	spec achieved
PDE @ 420 nm	> 40%	> 42%
DCR (87 K)	250 Hz / tile	~ 20 Hz / tile
correlated noise probabilities (afterpulses, cross talk)	< 50% + 50%	< 10% + 35%
SiPM gain	> 1E6	> 1E6
SNR after ARMA filter	> 8	> 15
time resolution	~ 10 ns	~ 15 ns

# PDU mass production (and testing) is starting



- Wafer delivery from LFoundry started in 2022.
- Packaging and assembly for TPC sensors: **Nuova Officina Assergi (NOA)**, about to start operations
- Packaging and assembly for Veto sensors: RAL and Liverpool, UK
- Several test facilities to qualify production: Naples, Liverpool, Edinburgh, AstroCent...



# The UAr target: URANIA & ARIA

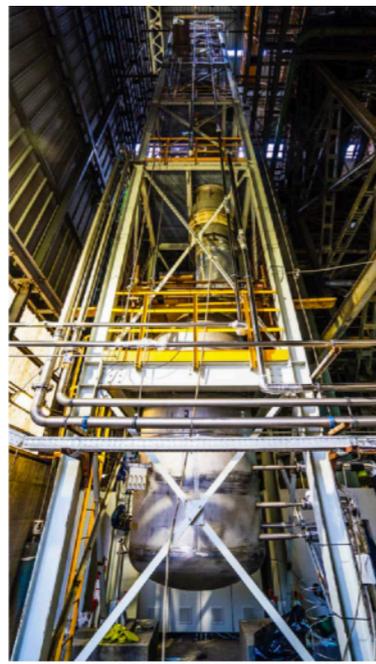
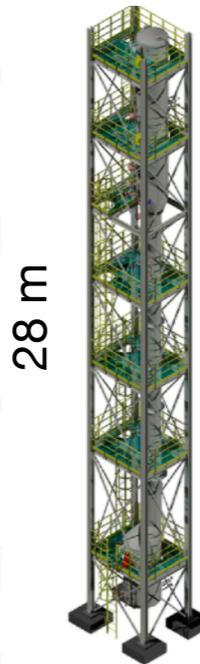
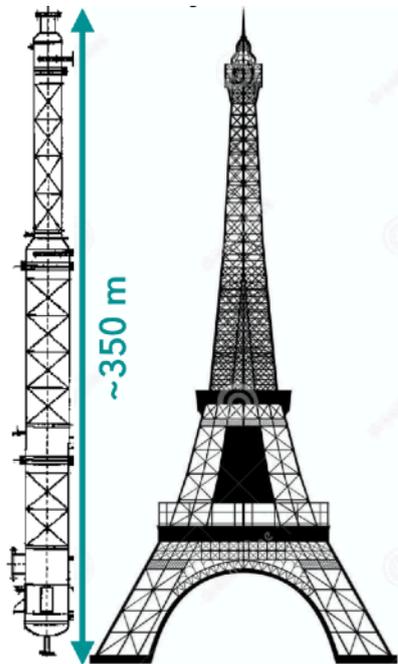
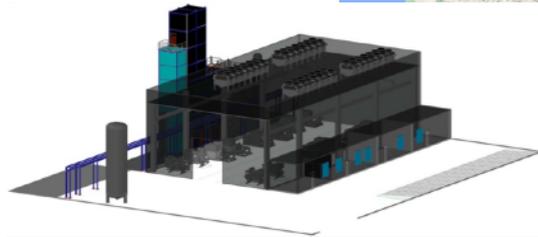
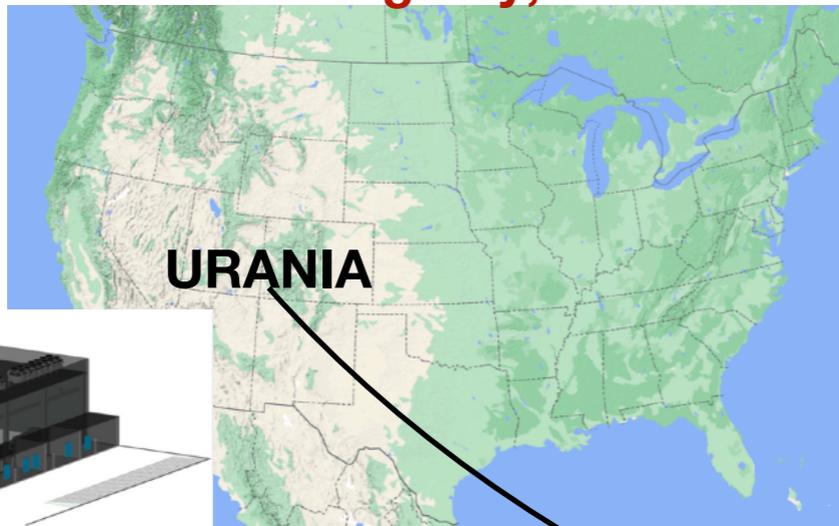
## 1) UAr extraction at the URANIA plant.

Industrial CO<sub>2</sub> extraction plant in Cortez, Colorado;

Plant ready to be shipped;

Expected argon purity at outlet: 99.99%;

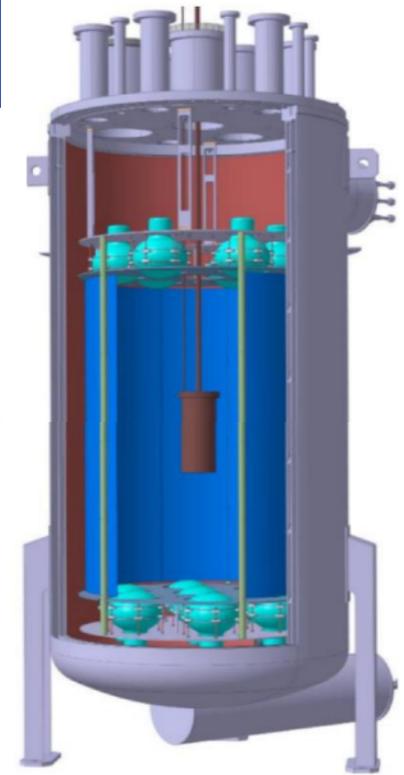
**UAr extraction rate: 250-330 kg/day;**



(*Eur. Phys. J. C* (2021) 81:359)

## 3) Qualification at Canfranc, DArT in ArDM

A single-phase LAr detector with active volume ~1L, capable of measuring UAr to AAr <sup>39</sup>Ar depletion factors of the order of 1000 with 10% precision in weeks



## 2) Cryogenic distillation at the ARIA facility

Installed in the shaft of a coal mine

**Chemical purification rate: 1 t/day**

First module operated according to specs with nitrogen

Run completed with Ar at the end of 2020: results to be published soon.

Full assembly about to start

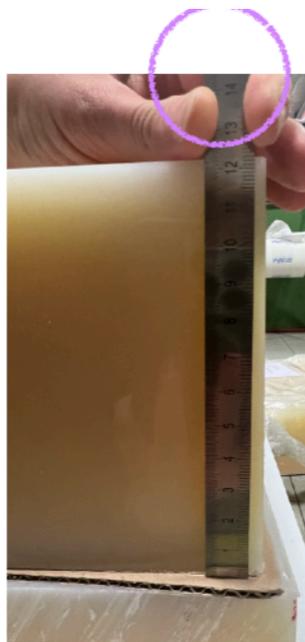
# Radiogenic neutron veto

- TPC surrounded by a single phase (S1 only) detector in UAr (35 tonne)
- Neutrons are captured by the Gd-loaded acrylic
- From the capture gamma ray shower up to 8 MeV
- Scintillation light is shifted by PEN wls and detected by SiPMs (400 channels) in both buffer and TPC
- **~90% tagging efficiency from simulation**, acceptable accidental lifetime loss

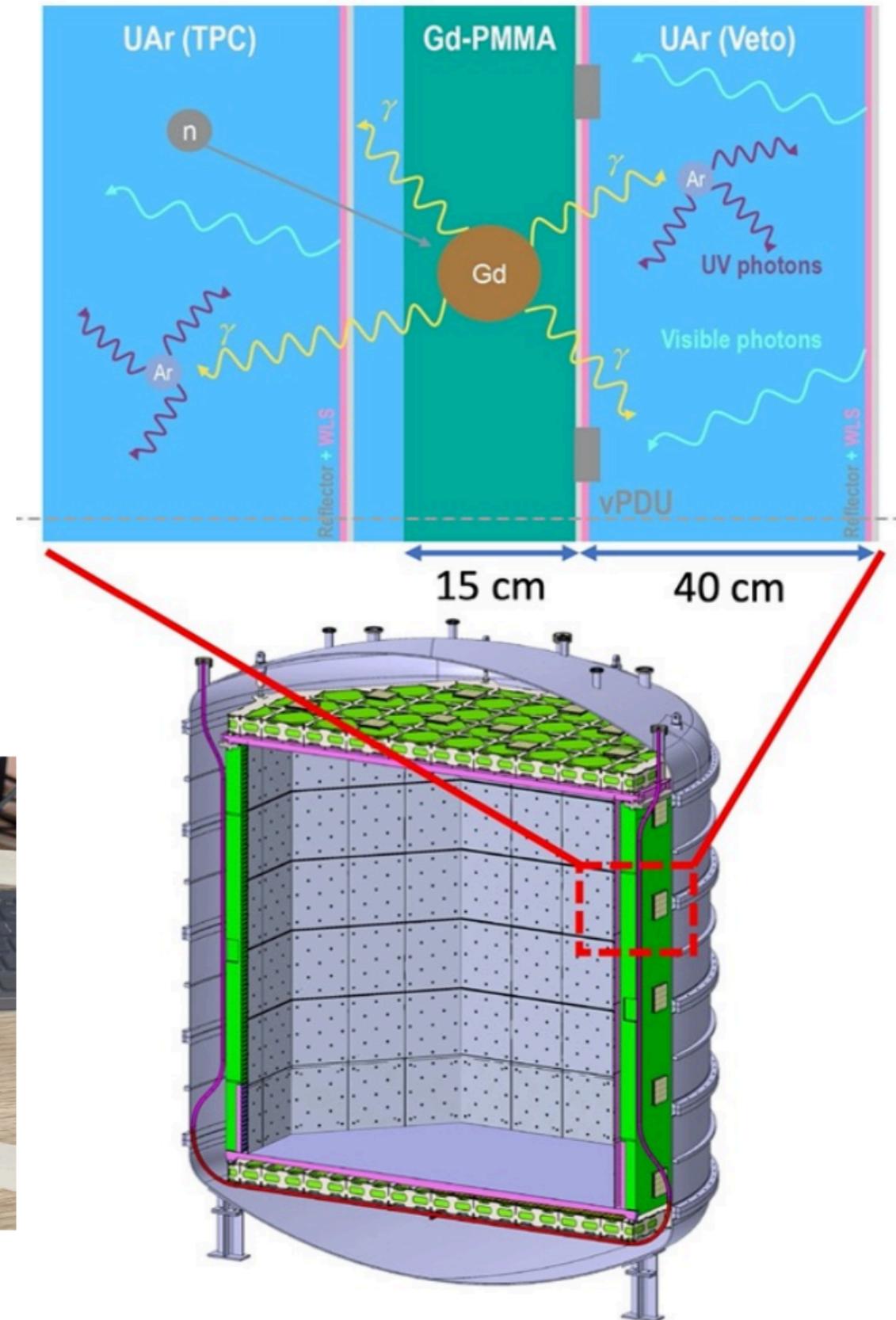
R&D(s) for Gd-loaded acrylic concluded. Radioactivity assay satisfactory. Moving to production



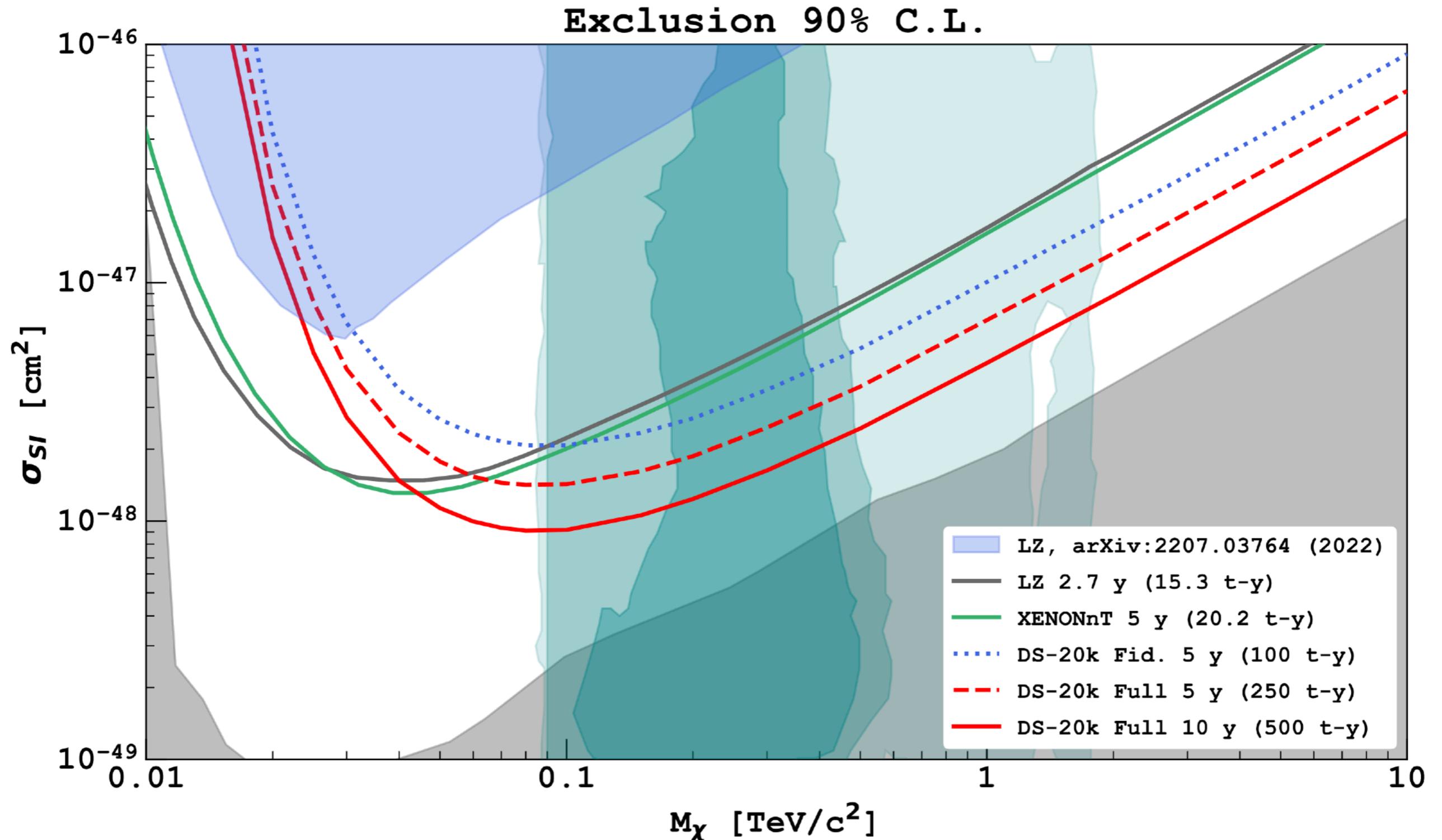
Gadolinium oxide nanograins ( $Gd_2O_3$ ) mechanical dispersion in MMA.



$Gd(MAA)_3$  doped acrylic sheet (5 cm thick)



# Projected sensitivity



Projected sensitivity based on bg-free exposure (<0.1 events in 200 t.yr) using the FV (innermost 20 tonnes of UAr) AND using **the full active volume** (PLR approach, bg pdfs known)

**Main backgrounds** (in 200 t.yr): **CEvNS (3.2 ev)**, radiogenic and cosmogenic neutrons, ER+Cherenkov, S1+S2 accidental coincidence...

# Conclusions

- R&D phase for the DS-20k detector completed, **construction is starting!**
- Production of SiPMs started and PDU mass test facility is ready
- Underground Ar procurement and characterisation project is ongoing (URANIA & ARIA & DArT-ArDM)
- Underground Argon cryogenic system has been successfully tested at CERN and is being relocated at INFN-LNGS
- Construction of the DS-20k cryostat will start soon in Hall C
- Mechanical mockup testing scheduled to take place at INFN-LNGS starting this year
- Data taking expected in 2026
- Wide range of physics:
  - **most sensitive WIMP search of the next decade**
  - potential at **lower WIMP masses** being assessed (dedicated detector?)
  - sensitivity to **supernova  $\nu$  bursts**
  - ...



# Expected backgrounds

**Full background budget** obtained thanks to:

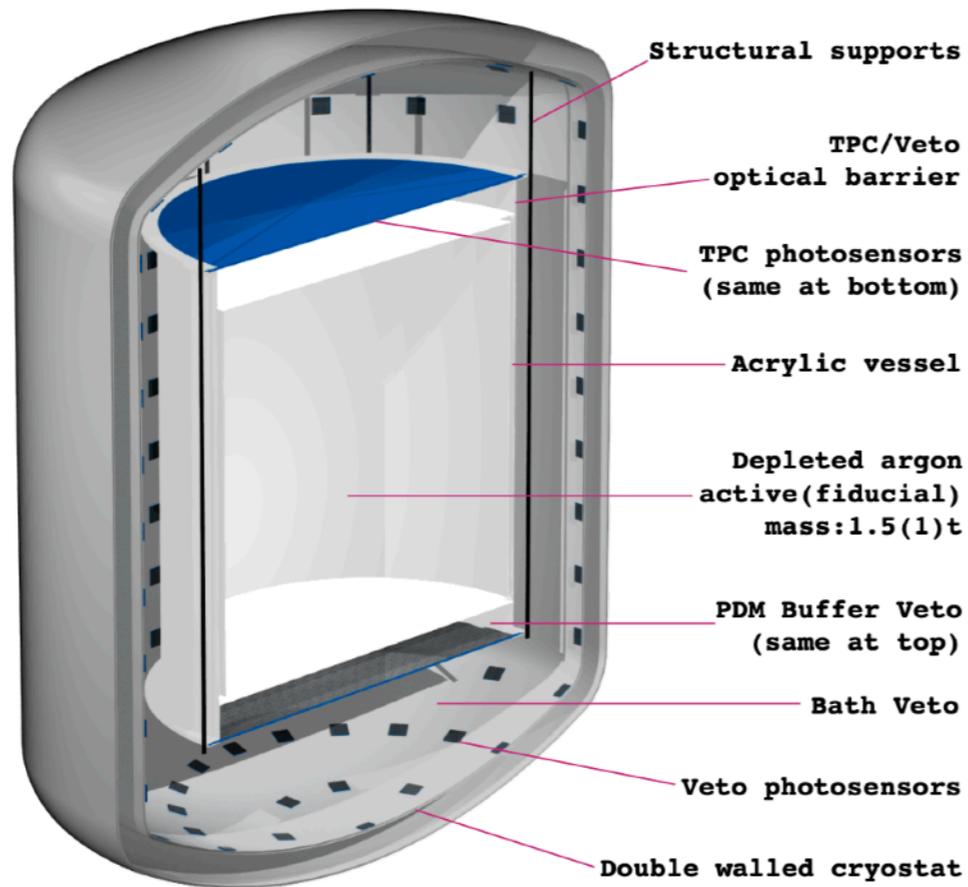
- Massive material screening (Jagiellonian University, CIEMAT, SNO, Canfranc, LNGS, Boulby) and strict material selection (+ account for broken equilibrium)
- Independent tools for ( $\alpha$ , n) emission probability (SaG4n: <http://win.ciemat.es/SaG4n/>)
- Full simulation of detector geometry, particle tracking, UAr scintillation, detector response and analysis strategy

Irreducible CEvNS: 3.2 evts in ROI in 200 t.yr

Background type	Bg events in ROI [200 t yr] <sup>-1</sup>	Material selection	Cleaning and purification	Event Topology	PSD	AAr instrumentation	Neutron veto
( $\alpha, n$ ) neutrons from U and Th	$9.5 \times 10^{-2}$	✓✓		✓✓			✓✓
Fission neutrons from U-238	$<2.3 \times 10^{-3}$	✓✓	✓	✓✓			✓✓
Neutrons from Rn-222 diffusion and surface plate-out	$<1.4 \times 10^{-2}$	✓✓	✓	✓✓			
Cosmogenic neutrons (stat limited)	$<6.0 \times 10^{-1}$					✓	
Neutrons from the lab rock	$1.5 \times 10^{-2}$	✓✓		✓✓			✓
Random surface $\alpha$ decay + S2 coincidence	$<5.0 \times 10^{-2}$	✓✓	✓	✓✓			
Correlated ER + Cherenkov	$<1.8 \times 10^{-2}$	✓✓		✓✓			
Uncorrelated ER + Cherenkov	$<3.0 \times 10^{-2}$	✓✓		✓✓			
ER	$<1.0 \times 10^{-1}$				✓		

TABLE 26. Nuclear recoil (NR) backgrounds expected during the full DS-20k exposure, based on current data and Monte Carlo simulations. The right column is the total number of events surviving the veto cut, fiducial volume cut, and PSD.

# DarkSide-LowMass



## Design optimization:

### – to reduce gamma background:

- material choice
- geometry
- active veto
- fiducialization

### – to increase single electron sensitivity:

- high extraction field

R&D needed to understand/reduce the single electron bg that dominates  $< 4$  ne

