



Weak gravity, supergravity and de Sitter vacua

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Motivation

- ▶ Our universe is expanding due to positive background energy density of order $\Lambda \sim 10^{-122} M_P^4$.
- ▶ A lot of effort in finding de Sitter vacua within flux compactifications of supersymmetric string theory.
- ▶ Turns out, it is extremely hard to find **stable de Sitter** in string theory, if possible at all.

See e.g. Danielsson, Van Riet '18

Can we uncover such problem directly within 4D supergravity?

→ For 4D EFTs derived from string theory

$$V = f^2 - 3m_{3/2}^2 > 0 \implies \underline{\text{Supersymmetry Breaking.}}$$

→ Supersymmetry breaking in 4D supergravity typically from:

- ▶ Superpotentials. (*fluxes*)
- ▶ Gaugings. (*fluxes*)
- ▶ Non-linear realizations. (*anti-branes*)

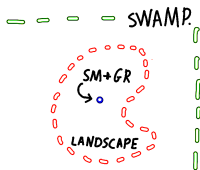
Can we restrict these ingredients? - what lessons do we learn?

Plan:

- Swampland and gaugings
- Non-linear supersymmetry and condensates
- Outlook

Swampland and gaugings

- ✗ Swampland: EFTs that **do not** arise from S.T.
- ✓ Landscape: EFTs that **do** arise from S.T.



We distinguish swamp from landscape with so-called **conjectures**: *General properties that we guess and test on compactifications, black holes, quantum gravity etc., e.g.:*

Weak gravity conjecture (WGC) postulates that in EFTs with electromagnetism there is a UV cut-off

$$\Lambda_{UV} < qg M_P .$$

Arkani-Hamed, Motl, Nicolis, Vafa '06

The 4D N=1 Lagrangian of supergravity + FI term (*Freedman '77*) is

$$e^{-1} \mathcal{L} = -\frac{1}{2} M_P^2 R + \frac{1}{2} \epsilon^{\kappa\lambda\mu\nu} (\bar{\psi}_\kappa \bar{\sigma}_\lambda D_\mu \psi_\nu - \psi_\kappa \sigma_\lambda D_\mu \bar{\psi}_\nu) \\ - \frac{1}{4g^2} F_{\mu\nu} F^{\mu\nu} + i\xi \epsilon^{\kappa\lambda\mu\nu} \bar{\psi}_\kappa \bar{\sigma}_\lambda \psi_\mu A_\nu - 4g^2 \xi^2 M_P^4.$$

- ▶ Because ξ is the gravitino charge, the WGC gives

$$\Lambda_{UV} \sim g \xi M_P.$$

- ▶ The vacuum energy is $4g^2 \xi^2 M_P^2$ and so the Hubble is

$$H \sim g \xi M_P \sim \Lambda_{UV},$$

therefore such genuine 4D de Sitter vacuum cannot arise from string theory. *Cribiori, FF, Tringas '20*

Why is $H \sim \Lambda_{UV}$ problematic? See e.g. Cribiori, Dall'Agata, FF '20

- ▶ From an EFT perspective we have

$$e^{-1} \mathcal{L}_{grav.} = M_P^2 \left(\frac{1}{2} R + \frac{\mathcal{O}(1)}{\Lambda_{UV}^2} R^2 + \dots \right).$$

- ▶ On a spatially flat de Sitter background one has

$$R = 12H^2,$$

therefore to be able to trust the 2-derivative EFT we need

$$H \ll \Lambda_{UV}.$$

- ▶ Supported by further arguments related to scalar fluctuations in de Sitter but also by the interpretation of H as an IR cut-off.

Further results:

- ▶ We have shown that **all known stable de Sitter vacua of $N=2$** (e.g. Fre, Trigiante, Van Proeyen '02) have the same issue and massless gravitini. *Cribiori, Dall'Agata, FF '20, Dall'Agata, Emelin, FF, Moritsu '21*
- ▶ $N=2$ dS by Catino, Scrucca, Smyth '13: no explicit model.
- ▶ We have deemed it as a **conjecture** that:
Supergravity de Sitter vacua with vanishing gravitino mass belong to the swampland.
Cribiori, Dall'Agata, FF '20
- ▶ All known stable de Sitter $N=2$ vacua have $m_{3/2} = 0$, $q_{3/2} \neq 0$, so violate also the festina lente bound. *Montero, Van Riet, Venken '19, Montero, Vafa, Van Riet, Venken '21*

Non-linear supersymmetry and condensates

Dall'Agata, Emelin, FF, Moritsu '22

- ▶ The Volkov–Akulov model is

$$\mathcal{L}_{VA} = -f^2 + i\partial_m \bar{G} \bar{\sigma}^m G + \frac{1}{4f^2} \bar{G}^2 \partial^2 G^2 - \frac{1}{16f^6} G^2 \bar{G}^2 \partial^2 G^2 \partial^2 \bar{G}^2.$$

- ▶ This EFT is defined with a cut-off $\Lambda \leq \sqrt{f}$, and we want to lower it to uncover the existence of composite states à la Nambu–Jona-Lasinio.
- ▶ In terms of *linear SUSY* it is described by

$$K = X\bar{X}, \quad W = fX + \frac{1}{2}TX^2,$$

where the variation of T gives $X^2 = 0$, which gives

$$X = G^2/2F \implies \text{NL SUSY.}$$

Check M. Moritsu's poster for details!

- ▶ Using the Functional RG Flow (with some approx.) we find

$$K = \alpha|X|^2 + \beta|T|^2 + g|T|^2|X|^2 + \frac{1}{4}q|X|^4, \quad W = fX + \frac{1}{2}TX^2.$$

- ▶ We find for $t_{RG} = \log \Lambda/\Lambda' \ll 1$ that

$$\alpha \simeq 1, \quad \beta \simeq \frac{1}{16\pi^2}t^2, \quad g \simeq \frac{2t}{\Lambda'^2}, \quad q \simeq \frac{2t}{\Lambda'^2}, \quad f \gtrsim \Lambda^2.$$

- ▶ Around the “V–A” point $T = X = 0$ we find tachyons

$$V = f^2, \quad V'' < 0.$$

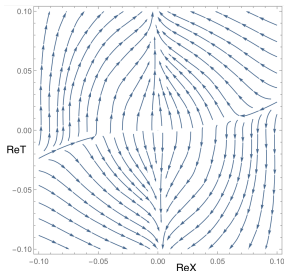
- ▶ You can imagine the composite states to have the form

$$X \sim G^2/f, \quad T \sim \partial^2 \overline{G}^2/f^2.$$

What happens with de Sitter?

- ▶ The V–A model is easily coupled to 4D $N=1$ supergravity to get de Sitter. *See e.g. Lindstrom, Rocek '79, Bergshoeff, Freedman, Kallosh, Van Proeyen '15*
- ▶ Doing the ERG within supergravity is actually beyond the state-of-the-art.
- ▶ We simply **directly couple the effective theory at Λ' to supergravity.**

1. Tachyons persist in SG.
2. Similarly due to NL SUSY of D3, also in KKLT.
3. Agreement with gravitino condensates bibliography. *E.g. Alexandre, Houston, Mavromatos '13-'15*



Outlook

4D supergravity can give us a very strong handle on de Sitter:

- ▶ Break SUSY and uplift with **gaugings** → Clash with WGC or festina lente bound.
- ▶ Anti-brane uplift with **NL SUSY** → New tachyonic instability towards goldstino condensates.
 - ▶ New stable vacua / Matter couplings?
 - ▶ We need to understand ERG in supergravity.
 - ▶ Revisit anti-brane uplifts.
- ▶ What about the **superpotential**?
(Maybe using p-forms? See e.g. Martucci, Sorokin, et al.)

Thank you