# Footprints of Majorana Neutrinos in Rare Meson Decays

Julia Harz, Technical University of Munich

in collaboration with **F. Deppisch** and **K. Fridell** (hep-ph/2009.04494)

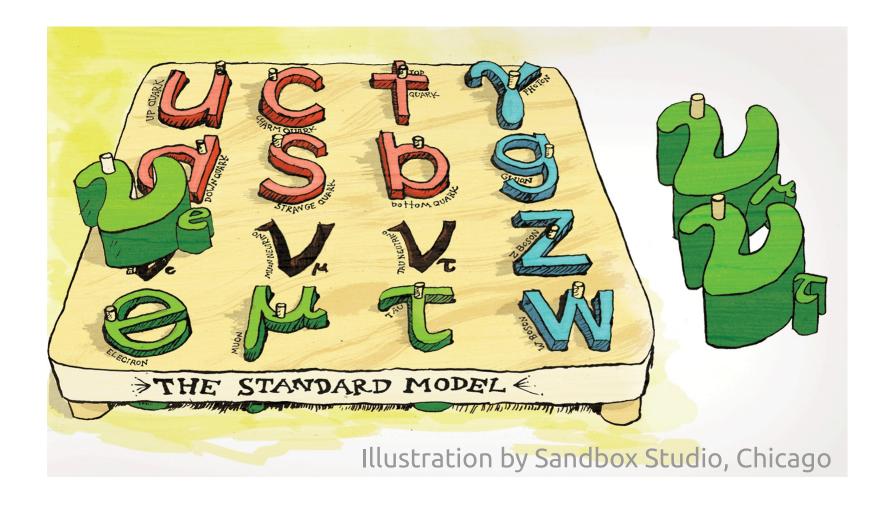
and A. J. Buras (upcoming work)

Bologna, July 9th 2022 ICHEP conference 2022





## Neutrinos – the Standard Model misfits



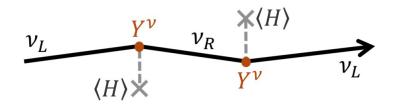


# Neutrinos - Dirac or Majorana?

#### Dirac mass

$$y_{\nu}L\epsilon H\overline{\nu}_R\supset m_D\nu_L\overline{\nu}_R$$

→ lepton number no accidental symmetry anymore



## Majorana mass

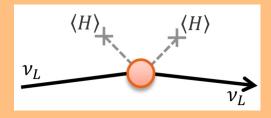
$$m_M \overline{
u}_R 
u_R^c$$

→ higher dimensional operator

$$m_M \overline{
u}_L 
u_L^c$$

not at tree-level within the SM possible

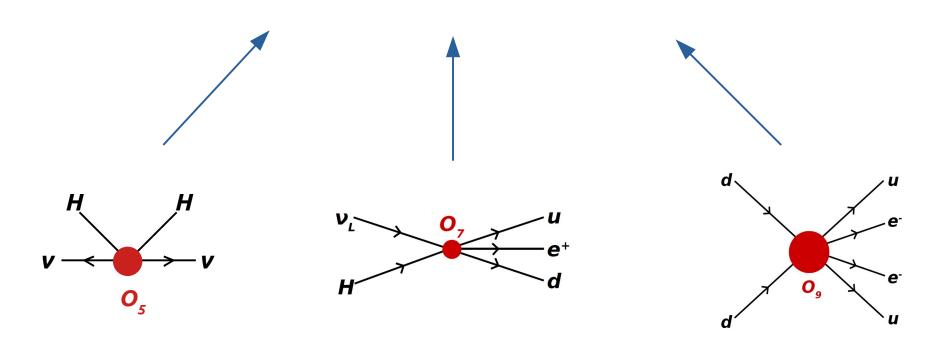
→ Lepton number violation (LNV)



# **Lepton Number Violation**

## LNV occurs only at odd mass dimension beyond dim-4:

$$\mathcal{L} = \mathcal{L}_{SM} + \frac{1}{\Lambda_1} \mathcal{O}_1^{(5)} + \sum_i \frac{1}{\Lambda_i^3} \mathcal{O}_i^{(7)} + \sum_i \frac{1}{\Lambda_i^5} \mathcal{O}_i^{(9)} + \cdots$$



See surveys of all LNV operators up to dim-11 e.g. in

Babu, Leung (2001), Gouvea, Jenkins (2008), Graf, JH, Deppisch, Huang (2018)



# **Lepton Number Violation**

$$\mathcal{L} = \mathcal{L}_{SM} + \frac{1}{\Lambda_1} \mathcal{O}_1^{(5)} + \sum_i \frac{1}{\Lambda_i^3} \mathcal{O}_i^{(7)} + \sum_i \frac{1}{\Lambda_i^5} \mathcal{O}_i^{(9)} + \cdots$$







$$\mathcal{O}_1^{(5)} = L^{\alpha} L^{\beta} H^{\rho} H^{\sigma} \epsilon_{\alpha\rho} \epsilon_{\beta\sigma}$$

$$\mathcal{O}_{3b}^{(7)} = L^{\alpha}L^{\beta}Q^{\rho}d^{c}H^{\sigma}\epsilon_{\alpha\rho}\epsilon_{\beta\sigma}$$

$$\mathcal{O}_{16}^{(9)} = L^{\alpha} L^{\beta} e^c d^c \bar{e}^c \bar{u}^c \epsilon_{\alpha\beta}$$

See surveys of all LNV operators up to dim-11 e.g. in

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# **Lepton Number Violation**

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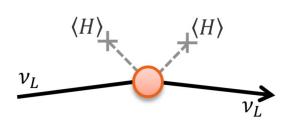


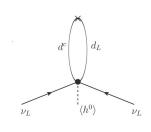


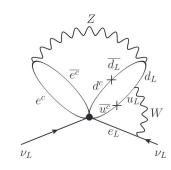
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$$m_{\nu} pprox rac{v^2}{\Lambda_1}$$

$$m_{\nu} pprox rac{y_d}{16\pi^2} rac{v^2}{\Lambda_{3b}}$$

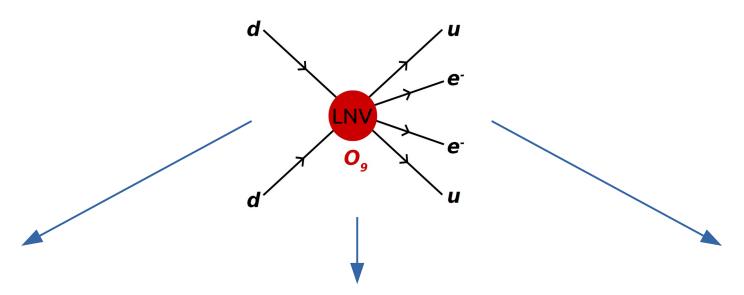
$$m_{\nu} \approx \frac{y_d y_u g^4}{(16\pi^2)^4} \frac{v^2}{\Lambda_{16}}$$

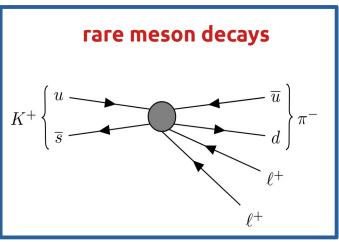
→ the discovery of a higher dimensional operator will point towards a Majorana contribution to neutrino masses

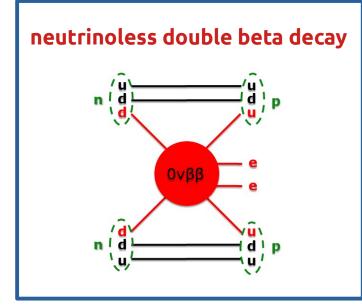
Gouvea, Jenkins (2008), Graf, JH, Deppisch, Huang (2018)

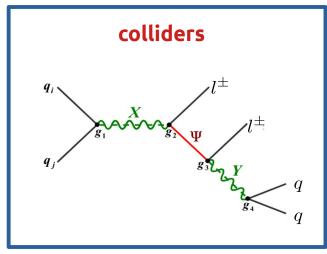


# **Probing LNV interactions**

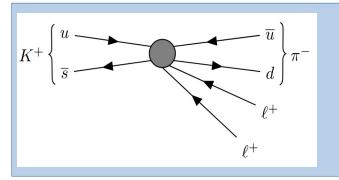








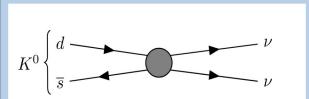
# Probing LNV interactions with Kaon decays



## Same-sign leptonic final state

- LNV is directly tested
- dim-9 SMEFT only
- for first generation, 0vββ stronger
- constraints very weak

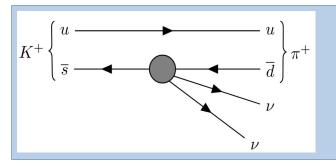
Liu, Zhang, Zhou (2016) Quintero (2017) Chun, Das, Mandal, Mitra, Sinha (2019)



## Decay into neutrino final state

- No experimental searches?
- dim-7 SMEFT

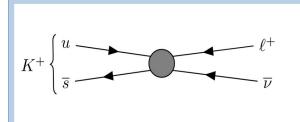
Gninenko (2014)



### Neutrino final state

- LNV needs to be independently confirmed
- dim-7 SMEFT

Li, Ma, Schmidt (2019) Deppisch, Fridell, JH (2020) Buras, JH (2022)



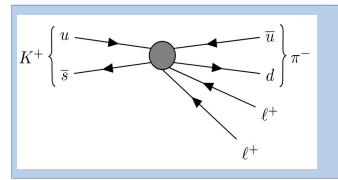
## Charged lepton + neutrino final state

- Neutrino needs to be detected (Cooper et al. 1982)
- dim-7 SMEFT

Deppisch, Fridell, JH (2020)



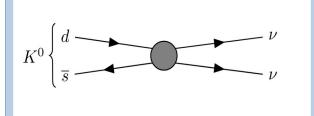
# Probing LNV interactions with Kaon decays



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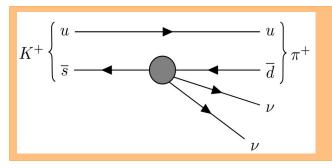
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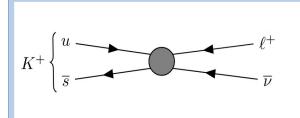
Gninenko (2014)



#### Neutrino final state

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Li, Ma, Schmidt (2019) Deppisch, Fridell, JH (2020) Buras, JH (2022)



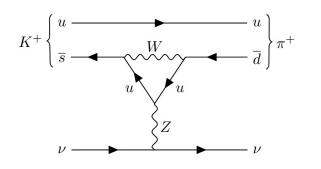
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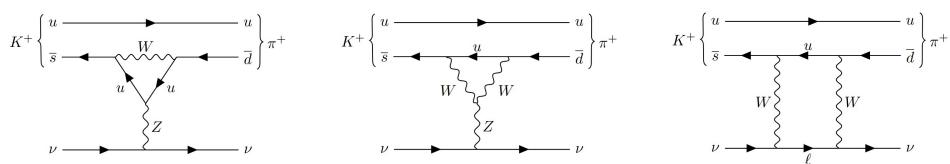
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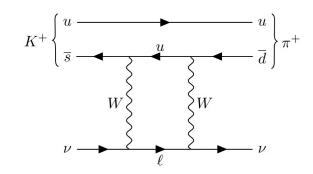
Deppisch, Fridell, JH (2020)



# $K^+ \rightarrow \pi^+ \nu \nu$ and $K_{\scriptscriptstyle I} \rightarrow \pi^0 \nu \nu$ in the Standard Model







## **Branching Ratios:**

$$BR(K^+ \to \pi^+ \nu \bar{\nu}) = \kappa^+ (1 + \Delta_{EM}) \left[ \left( \frac{Im(V_{ts}^* V_{td} X_t)}{\lambda^5} \right)^2 + \left( \frac{Re(V_{cs}^* V_{cd})}{\lambda} P_c + \frac{Re(V_{ts}^* V_{td} X_t)}{\lambda^5} \right)^2 \right]$$

$$\mathrm{BR}(K_L o \pi^0 
u ar{
u}) = \kappa_L \left( \frac{\mathrm{Im}(V_{ts}^* V_{td} X_t)}{\lambda^5} \right)^2$$
 GIM suppressed in the SM!

Buchalla, Buras (1999)

## Small hadronic uncertainties due to relation to well measured $K o \pi \ell^+ u_\ell$ :

$$\kappa^{+} = (0.5173 \pm 0.025) \times 10^{-10} (|V_{us}|/0.0225)^{8}$$

$$\kappa_L = (2.231 \pm 0.013) \times 10^{-10} (|V_{us}|/0.0225)^8$$

Mescia, Smith (2007)



# Theoretical and experimental status

## Theoretical prediction

$$BR(K^+ \to \pi^+ \nu \bar{\nu})_{SM} = (8.60 \pm 0.42) \times 10^{-11}$$

$$BR(K_L \to \pi^0 \nu \bar{\nu})_{SM} = (2.94 \pm 0.15) \times 10^{-11}$$

Buras, Buttazzo, Girrbach-Noe, Knegjens (2015) updated by Buras, Venturini (2021)

**Golden Channel!** 

## **Experimental measurements**

$${\rm BR}(K^+ \to \pi^+ \nu \bar{\nu})_{\rm E949} = \left(1.73^{+1.15}_{-1.05}\right) \times 10^{-10}$$
 E949 collaboration (2009)

$$BR(K^+ \to \pi^+ \nu \bar{\nu})_{NA62} = (10.6^{+4.0}_{-3.5} \pm 0.9) \times 10^{-11}$$

NA62 collaboration (2021)

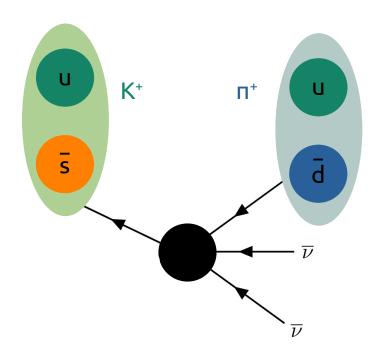
BR
$$(K_L \to \pi^0 \nu \bar{\nu})_{KOTO} = \left(2.1^{+2.0(+4.1)}_{-1.1(-1.7)}\right) \times 10^{-9}$$

KOTO collaboration (2019)

## → NA62 aims to reach SM sensitivity!

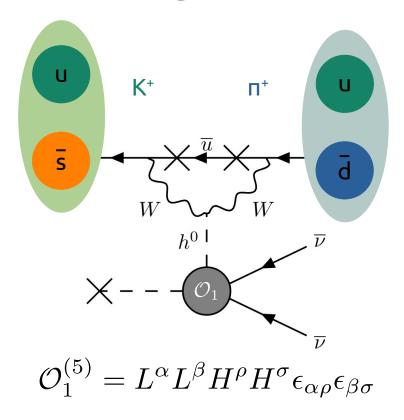
What would a deviation from the SM expectation imply for new physics?

# Constraining new physics in rare kaon decays

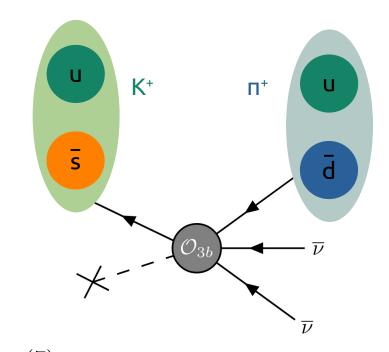


As neutrinos are not explicitly measured, a new physics contribution could also be lepton number violating!

# Constraining LNV interactions with rare kaon decays



- GIM suppressed
- Majorana neutrino mass



$$\mathcal{O}_{3b}^{(7)} = L^{\alpha}L^{\beta}Q^{\rho}d^{c}H^{\sigma}\epsilon_{\alpha\rho}\epsilon_{\beta\sigma}$$

- No GIM suppression
- Majorana contribution to neutrino masses
- → Footprints of Majorana neutrinos in rare meson decays?
- → BUT: no unambiguous signal of LNV is there a way to disentangle LNV from LNC?

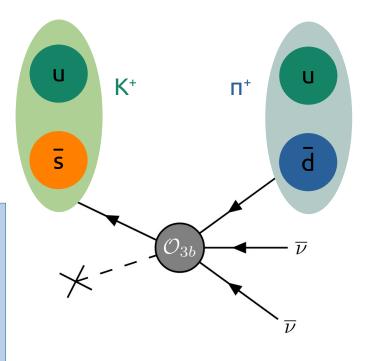


# Lepton number violating vs conserving current

$$\frac{\Gamma\left(K \to \pi \nu_i \nu_j\right)}{ds \, dt} = \frac{1}{1 + \delta_{ij}} \frac{1}{(2\pi)^3} \frac{1}{32m_K^3} |\overline{\mathcal{M}}|^2$$

• **SM**, lepton number **conserving vector** current

$$\mathcal{L}_{\mathrm{SM}}^{K \to \pi \nu \bar{\nu}} = \frac{1}{\Lambda_{\mathrm{SM}}^2} \left( \bar{\nu}_i \gamma^{\mu} \nu_i \right) \left( \bar{d} \gamma_{\mu} s \right)$$
$$|\mathcal{M}|^2 = \frac{6}{\Lambda_{\mathrm{SM}}^4} \left[ m_K^2 \left( t - m_{\pi}^2 \right) - t \left( s + t - m_{\pi}^2 \right) \right] f_+^K(s)^2$$



BSM, lepton number violating scalar current

$$\mathcal{L}_{\text{BSM}}^{K \to \pi \nu \nu} = \frac{v}{\Lambda_{\text{BSM}}^3} \left( \nu_i \nu_j \right) \left( \bar{d}s \right)$$
$$|\mathcal{M}|^2 = \frac{v^2}{\Lambda_{\text{BSM}}^6} \left( \frac{m_K^2 - m_\pi^2}{m_s - m_d} f_0^K(s) \right)^2 s$$

Li, Ma, Schmidt (2019) Deppisch, JH, Fridell (2020) Buras, JH (2022)



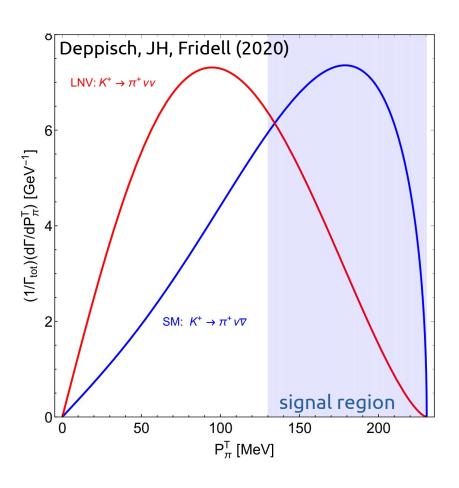
## LNV vs LNC current – KOTO

$$\langle \pi^0 | \, \bar{d}s \, | \bar{K}^0 \rangle = \langle \pi^0 | \, \bar{s}d \, | K^0 \rangle$$
$$\langle \pi^0 | \, \bar{d}\gamma^\mu s \, | \bar{K}^0 \rangle = -\langle \pi^0 | \, \bar{s}\gamma^\mu d \, | K^0 \rangle$$

$$i\mathcal{M}\left(K_L \to \pi^0 \nu \nu\right) = \frac{1}{\sqrt{2+2|\epsilon|^2}} \left(F(1+\epsilon) \langle \pi^0 | C | K^0 \rangle + F^*(1-\epsilon) \langle \pi^0 | C | \bar{K}^0 \rangle\right) \nu \nu$$

LNV mode → scalar current → real part

LNC mode → vector current → imaginary part

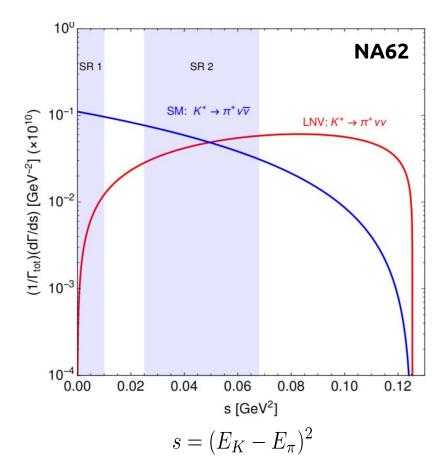


- → no CP phase needed in the LNV case
- → different phase space distribution

Deppisch, Fridell, JH (2020)



## LNV vs LNC current – NA62



$$BR_{LNV}(K^+ \to \pi^+ \nu_i \nu_j) = 10^{-10} \left(\frac{19.2 \text{ TeV}}{\Lambda_{ijsd}}\right)^6$$

$$BR_{LNV}(K_L \to \pi^0 \nu_i \nu_j) = 10^{-10} \left(\frac{24.9 \text{ TeV}}{\Lambda_{ijsd}}\right)^6$$

Process	Experimental limit	0	$\Lambda_{ijkn}^{\mathrm{NP}}$ [TeV]
$K^+ \to \pi^+ \nu \nu$	$BR_{future}^{NA62} < 1.11 \times 10^{-10}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 19.6$
$K^+  o \pi^+ \nu \nu$	$BR_{current}^{NA62} < 1.78 \times 10^{-10} [67]$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 17.2$
$K_L \to \pi^0 \nu \nu$	$BR_{\text{current}}^{\text{KOTO}} < 3.0 \times 10^{-9} [71]$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 12.3$

- → scalar and vector current lead to different phase space distribution
- → competitive limits on the new scale of physics

Deppisch, Fridell, JH (2020)



# Complementarity with other LNV observables

Process	Experimental limit	O	$\Lambda_{ijkn}^{\mathrm{NP}}$ [TeV]
$K^+ \to \pi^+ \nu \nu$	$BR_{\text{future}}^{\text{NA62}} < 1.11 \times 10^{-10}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 19.6$
$K^+  o \pi^+ \nu \nu$	$BR_{current}^{NA62} < 1.78 \times 10^{-10}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 17.2$
$K_L \to \pi^0 \nu \nu$	$BR_{current}^{KOTO} < 3.0 \times 10^{-9}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 12.3$
$B^+ \to \pi^+ \nu \nu$	$BR < 1.4 \times 10^{-5}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iibd} > 1.4$
$B^+  o K^+  u  u$	$BR < 1.6 \times 10^{-5}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iibs} > 1.4$
$B^0 \to \pi^0 \nu \nu$	$BR < 9 \times 10^{-6}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iibd} > 1.5$
$B^0  o K^0  u  u$	$BR < 2.6 \times 10^{-5}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iibs} > 1.3$
$K^+ \to \mu^+ \bar{\nu}_e$	$BR < 3.3 \times 10^{-3}$	$\mathcal{O}_{3a}$	$\Lambda_{\mu esu} > 2.4$
$\pi^+ \to \mu^+ \bar{\nu}_e$	$BR < 1.5 \times 10^{-3}$	$\mathcal{O}_{3a}$	$\Lambda_{\mu eud} > 1.9$
$\pi^0 \to \nu \nu$	$BR < 2.9 \times 10^{-13}$	$\mathcal{O}_{3b}$	$\Lambda_{\nu\nu ud} > 3.4$
0 uetaeta	$T_{1/2}^{136\text{Xe}} \ge 1.07 \times 10^{26} \text{ yrs}$	$\mathcal{O}_{3b}$	$\Lambda_{eeud} > 330$
$\mu^- \to e^+$	$R_{\mu^-e^+}^{\mathrm{Ti}} < 1.7 \times 10^{-12}$	$\mathcal{O}_{14b}$	$\Lambda_{\mu eud} > 0.01$

- $\rightarrow$  While limits weaker than from  $0v\beta\beta$  decay, different flavours are probed
- → B-meson constraints still in LHC reach



# General NP contribution in $K^+ \to \pi^+ \nu \nu$ or $K_L \to \pi^0 \nu \nu$

NOW: allow for most generic NP contribution from dim-6 operators.

#### **Dim-6 LEFT operators** (lepton number conserving)

$$\mathcal{O}_{uL}^{V} = (\overline{u_L}\gamma^{\mu}u_L)(\overline{\nu}\gamma^{\mu}\nu) , \qquad \mathcal{O}_{dL}^{V} = (\overline{d_L}\gamma^{\mu}d_L)(\overline{\nu}\gamma^{\mu}\nu) , 
\mathcal{O}_{uR}^{V} = (\overline{u_R}\gamma^{\mu}u_R)(\overline{\nu}\gamma^{\mu}\nu) , \qquad \mathcal{O}_{dR}^{V} = (\overline{d_R}\gamma^{\mu}d_R)(\overline{\nu}\gamma^{\mu}\nu) ,$$

#### **Dim-6 LEFT operators** (lepton number violating)

$$\mathcal{O}_{uRL}^{S} = (\overline{u_R}u_L)(\overline{\nu^C}\nu) , \qquad \mathcal{O}_{dRL}^{S} = (\overline{d_R}d_L)(\overline{\nu^C}\nu) , 
\mathcal{O}_{uLR}^{S} = (\overline{u_L}u_R)(\overline{\nu^C}\nu) , \qquad \mathcal{O}_{dLR}^{S} = (\overline{d_L}d_R)(\overline{\nu^C}\nu) , 
\mathcal{O}_{uLR}^{T} = (\overline{u_R}\sigma^{\mu\nu}u_L)(\overline{\nu^C}\sigma_{\mu\nu}\nu) , \qquad \mathcal{O}_{d}^{T} = (\overline{d_R}\sigma^{\mu\nu}d_L)(\overline{\nu^C}\sigma_{\mu\nu}\nu) .$$

For flavour diagonal contributions tensor contribution vanishes.

Li, Ma, Schmidt (2019)



# General NP contribution in $K^+ \rightarrow \pi^+ \nu \nu$ or $K_{\mu} \rightarrow \pi^0 \nu \nu$

#### Most generic Branching ratios:

$$\mathcal{B}(K_L \to \pi^0 \nu \widehat{\nu}) = 36.27 G_F^{-2} \sum_{\alpha \le \beta} \left( 1 - \frac{1}{2} \delta_{\alpha\beta} \right) \left| C_{dRL}^{S,sd\alpha\beta} + C_{dLR}^{S,sd\alpha\beta} + C_{dRL}^{S,ds\alpha\beta} + C_{dLR}^{S,ds\alpha\beta} \right|^2$$

$$+ 0.236 G_F^{-2} \sum_{\alpha,\beta} \left| C_{dL}^{V,sd\alpha\beta} + C_{dR}^{V,sd\alpha\beta} - C_{dL}^{V,ds\alpha\beta} - C_{dR}^{V,ds\alpha\beta} \right|^2$$

$$\mathcal{B}(K^+ \to \pi^+ \nu \widehat{\nu}) = 17.05 G_F^{-2} \sum_{\alpha \le \beta} \left( 1 - \frac{1}{2} \delta_{\alpha\beta} \right) \left( \left| C_{dRL}^{S,sd\alpha\beta} + C_{dLR}^{S,sd\alpha\beta} \right|^2 + \left| C_{dRL}^{S,ds\alpha\beta} + C_{dLR}^{S,ds\alpha\beta} \right|^2 \right)$$

$$0.219 G_F^{-2} \sum_{\alpha,\beta} \left| C_{dL}^{V,sd\alpha\beta} + C_{dR}^{V,sd\alpha\beta} \right|^2,$$

#### Parameterization:

$$C_{dL}^{V,sd\alpha\beta} + C_{dR}^{V,sd\alpha\beta} = (|C_{\rm SM}|e^{i\phi_{\rm SM}} + |C_{V}|e^{i\phi_{V}})\delta_{\alpha\beta} \qquad C_{dRL}^{S,sd\alpha\beta} + C_{dLR}^{S,sd\alpha\beta} = |C_{S}|e^{i\phi_{S}}\delta_{\alpha\beta},$$

$$C_{dL}^{V,ds\alpha\beta} + C_{dR}^{V,ds\alpha\beta} = (|C_{\rm SM}|e^{-i\phi_{\rm SM}} + |C_{V}|e^{-i\phi_{V}})\delta_{\alpha\beta}, \qquad C_{dRL}^{S,sd\alpha\beta} + C_{dLR}^{S,sd\alpha\beta} = |C_{S}|e^{-i\phi_{S}}\delta_{\alpha\beta}.$$

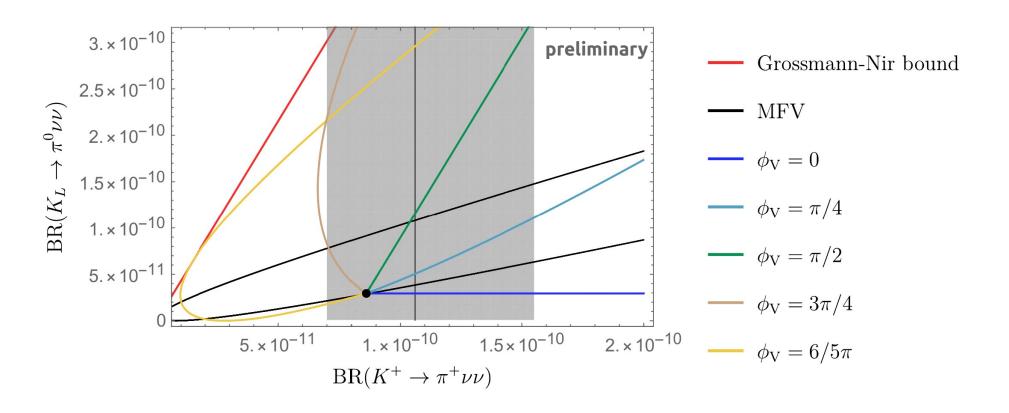
#### **Identify SM parameters:**

$$\mathcal{B}_{\text{SM}}(K_L \to \pi^0 \nu \nu) = 12 J_2^{K_L} |C_{\text{SM}}|^2 \sin^2 \phi_{\text{SM}}, \qquad \mathcal{B}_{\text{SM}}(K^+ \to \pi^+ \nu \nu) = 3 J_2^{K^+} |C_{\text{SM}}|^2$$

$$|C_{\text{SM}}| = 1.33 \times 10^{-10} \text{ GeV}^{-2}, \qquad \phi_{\text{SM}} = 0.09 \pi.$$



# New vector contribution in $K^+ \rightarrow \pi^+ \nu \nu$ and $K_+ \rightarrow \pi^0 \nu \nu$

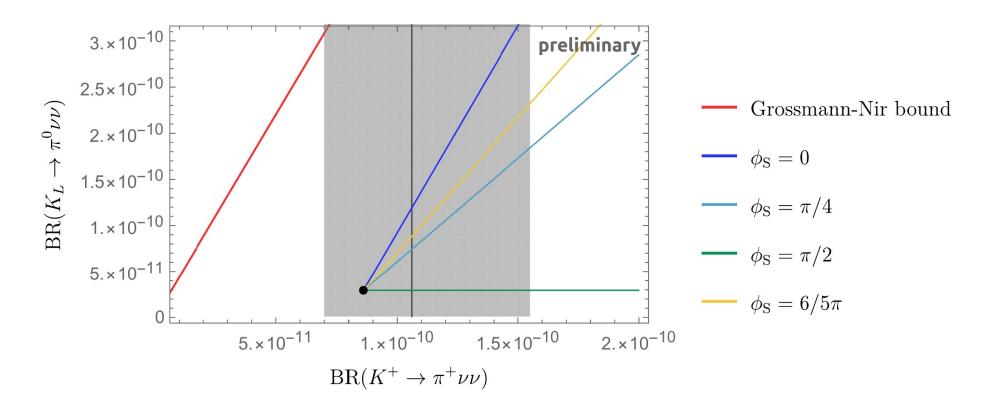


- → Vector contribution can lead to larger and smaller BRs than in the SM
- → Vector contribution can lie everywhere below Grossman-Nir bound

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# New scalar contribution in $K^+ \to \pi^+ vv$ and $K_L \to \pi^0 vv$



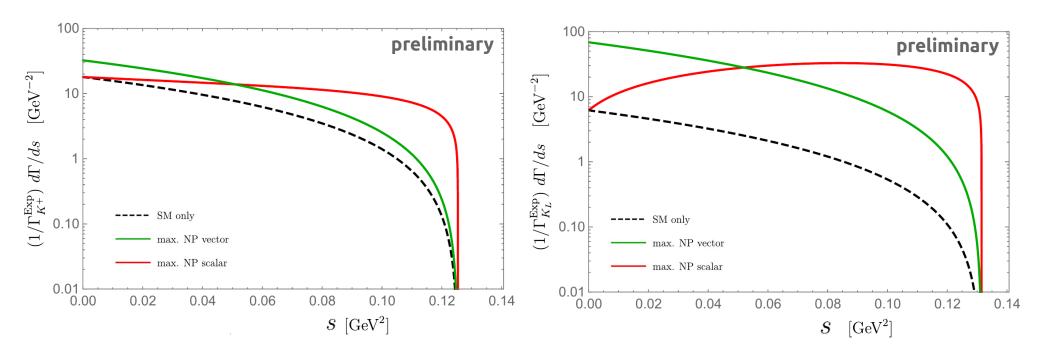
- → Scalar contribution can only lead to larger BRs than in the SM
- → Scalar contribution confined between blue and green line
- → Measuring a lower BR than in the SM implies (also) a vector contribution.



# Disentangling the NP contribution

Allow for a NP scalar or vector contribution additionally to the SM such that the experimental upper bound  $\mathcal{B}(K^+ \to \pi^+ \nu \bar{\nu}) = 1.55 \times 10^{-10}$  is saturated.

We fixed  $\phi_V=\pi/2$  and  $\phi_S=0$  .



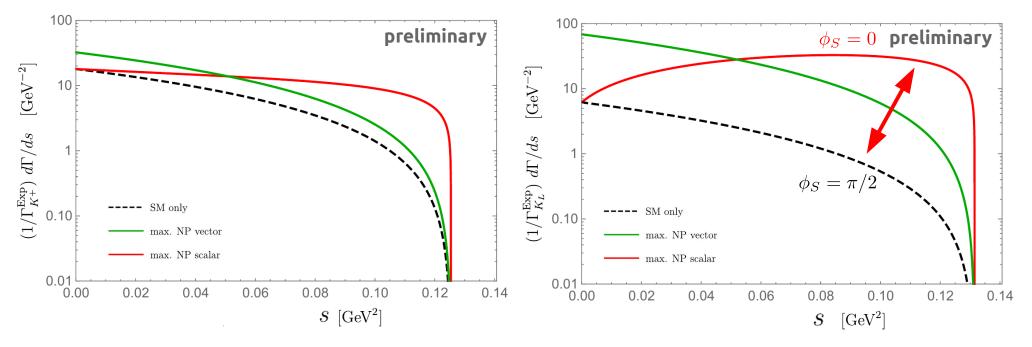
→ A NP scalar contribution additionally to the SM leads to a striking difference in the distribution when comparing to a vector contribution only.



# Disentangling the NP contribution

Allow for a NP scalar or vector contribution additionally to the SM such that the experimental upper bound  $\mathcal{B}(K^+ \to \pi^+ \nu \bar{\nu}) = 1.55 \times 10^{-10}$  is saturated.

We fixed  $\phi_V=\pi/2$  and  $\phi_S=0$  .



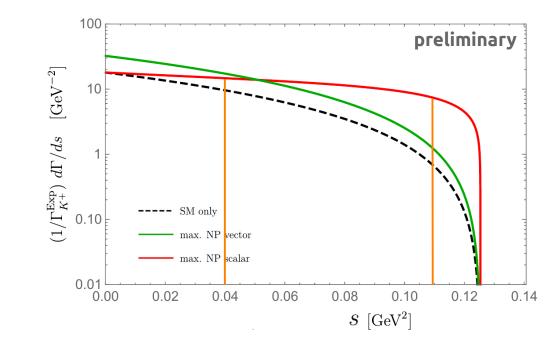
independent of scalar phase

dependent on scalar phase

→ additional insight by comparing K<sub>1</sub> and K<sup>+</sup> distributions!



# Proposal to disentangle a possible NP contribution



## Measuring distribution at two different values $s_1$ and $s_2$ :

$$\mathcal{D}_{+}^{\exp}(s) \equiv \frac{d\Gamma(K^{+} \to \pi^{+}\nu\widehat{\nu})}{ds} = C_{S}^{+}f_{S}^{+}(s) + C_{V}^{+}f_{V}^{+}(s)$$

$$C_V^+ = \frac{f_S^+(s_2)\mathcal{D}_+^{\text{exp}}(s_1) - f_S^+(s_1)\mathcal{D}_+^{\text{exp}}(s_2)}{f_V^+(s_1)f_S^+(s_2) - f_V^+(s_2)f_S^+(s_1)} \qquad C_S^+ = \frac{f_V^+(s_2)\mathcal{D}_+^{\text{exp}}(s_1) - f_V^+(s_1)\mathcal{D}_+^{\text{exp}}(s_2)}{f_S^+(s_1)f_V^+(s_2) - f_S^+(s_2)f_V^+(s_1)},$$

- $\rightarrow$  measuring non-zero C<sub>s</sub> implies the existence of a scalar current
- $\rightarrow$  measuring non-zero C<sub>v</sub> not in agreement with SM, implies new vector currents



## **Conclusions**

- Observation of LNV interactions would imply Majorana contribution to neutrino masses
- A deviation from the SM expectation in the golden channel K<sup>+</sup> → π<sup>+</sup>vv would point towards new physics (either LNV or LNC)
- While not a clear proof of LNV, pion energy distribution provides possibility to disentangle scalar LNV from vector LNC contribution
- Interplay between  $K^+ \to \pi^+ vv$  and  $K_L \to \pi^0 vv$  can give further insights
- In case of a deviation from SM consistent with LNV contribution, interplay with other experiments is crucial, such as collider searches and 0vββ decay
- → For experiments: dedicated analyses and limits for different currents highly interesting!

# Thank you for your attention!

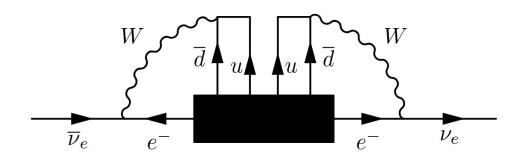






# Does LNV directly imply Majorana neutrinos?

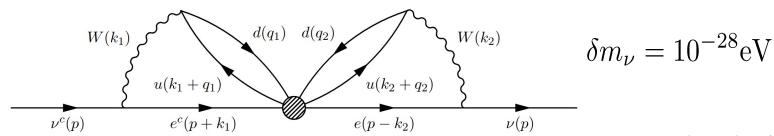
Schechter-Valle Theorem ("Black box" theorem)



Schechter, Valle (1982)

Any  $\Delta L = 2$  operator that leads to 0vbb will induce a Majorana mass contribution via loops.

Caveat



Dürr, Merle, Lindner (2011)

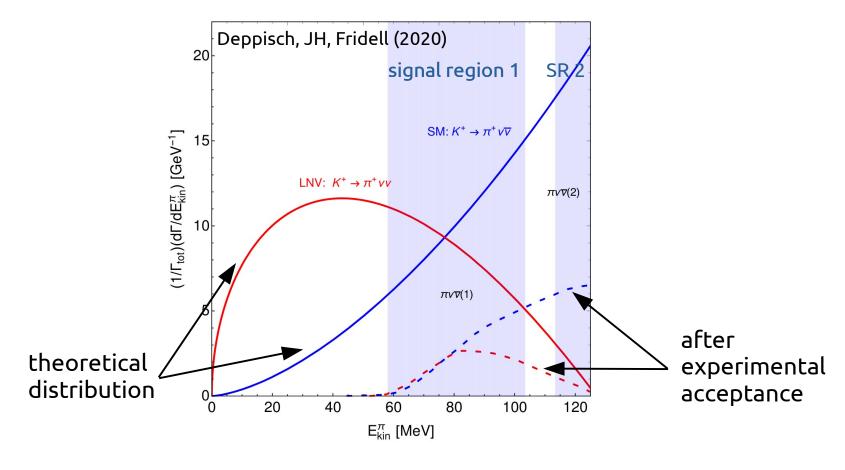
**E.g.** 9-dim  $\Delta L = 2$  operator will lead to 0vbb while contributing only little to the neutrino mass.

Observation of LNV implies some Majorana nature of neutrinos, but not necessarily the dominant contribution.



## Limits from E949

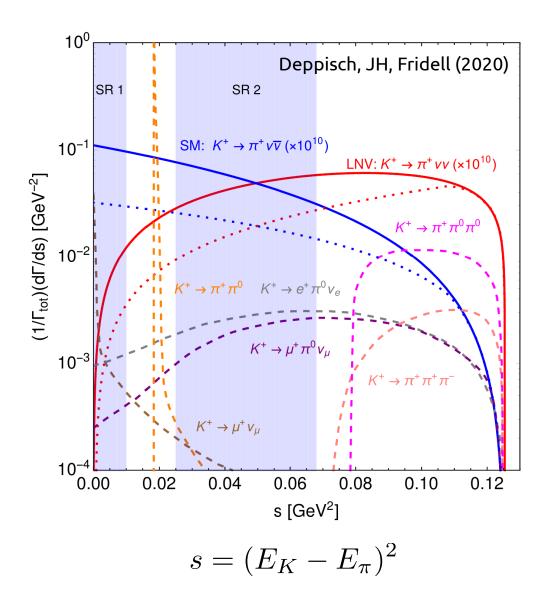
$$\frac{\Gamma(K \to \pi \nu_i \nu_j)}{ds \, dt} = \frac{1}{1 + \delta_{ij}} \frac{1}{(2\pi)^3} \frac{1}{32m_K^3} |\overline{\mathcal{M}}|^2$$



→ LNV and LNC current lead to a different phase space distribution

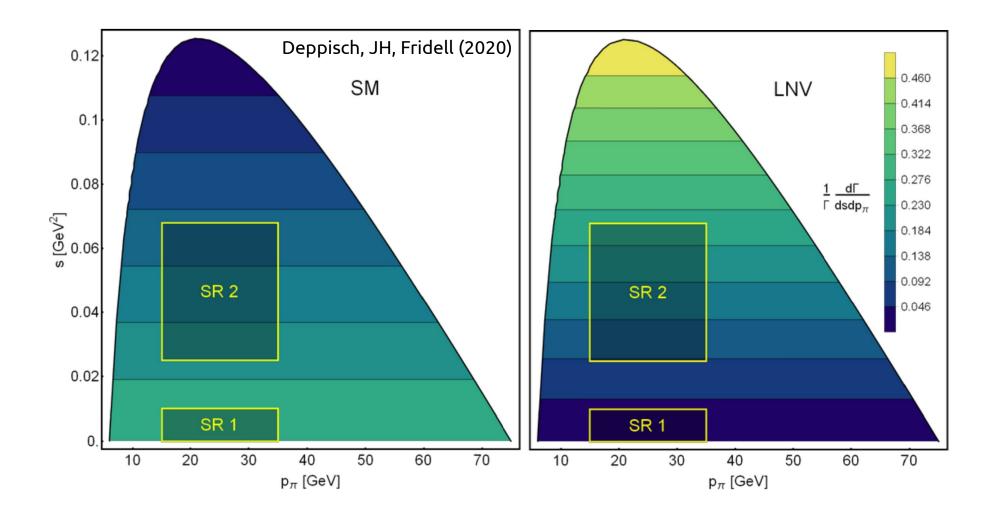


# Limits and prospects of NA62





# Disentangeling LNV and LNC currents at NA62





# Disentangeling LNV and LNC currents at NA62

Summary of sensitivities for a scalar current (based on kinematics only):

Experiment	SM (vector)	LNV (scalar)
NA62 SR 1	6%	0.3%
NA62 SR 2	17%	15%
$E949 \pi \nu \overline{\nu}(1)$	29%	2%
E949 $\pi\nu\overline{\nu}(2)$	45%	38%
КОТО	64%	30%

→ Possibility to disentangle a LNV scalar vs LNC vector current by improving on experimental sensitivity and strategy?



# Analysis for charged final states

O	$1/\Lambda_{M^+ \to \ell_i^+ \bar{\nu}_j}^2$	$\Lambda_{\mu eus}$ [TeV]	$\Lambda_{\mu eud}$ [TeV]	$\mid m_{ u} \mid$	$\Lambda^{m_{\nu}}$ [TeV]
3a	$\frac{v}{\Lambda^3}$	2.2	1.7	$\frac{y_d g^2}{(16\pi^2)^2} \frac{v^2}{\Lambda}$	69
$3a^{H^2}$	$f(\Lambda)\frac{v}{\Lambda^3}$	1.3	1.1	$\frac{y_d g^2}{(16\pi^2)^2} \frac{v^2}{\Lambda} f(\Lambda)$	0.4
4a	$\frac{v}{\Lambda^3}$	2.2	1.7	$\frac{y_u}{16\pi^2} \frac{v^2}{\Lambda}$	$2.4 \times 10^4$
$4a^{H^2}$	$f(\Lambda)\frac{v}{\Lambda^3}$	1.3	1.1	$\frac{y_u}{16\pi^2} \frac{v^2}{\Lambda} f(\Lambda)$	150
$4b^{\dagger}$	$\frac{v}{\Lambda^3}$	2.2	1.7	$\begin{array}{ c c c c c c }\hline 10\pi^{2} & 1 & 1 & 1 \\ \hline \frac{y_{u}g^{2}}{(16\pi^{2})^{2}} & \frac{v^{2}}{\Lambda} & 1 & 1 \\ \hline \end{array}$	33
$4b^{\dagger H^2}$	$f(\Lambda)\frac{v}{\Lambda^3}$	1.3	1.1	$ \frac{y_u g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda} $	0.2
6	$f(\Lambda)\frac{v}{\Lambda^3}$	1.3	1.1	$\frac{y_u}{\sqrt{1-y_u^2}} \frac{v^2}{\sqrt{1-y_u^2}}$	150
7	$\frac{v^3}{\Lambda^5}$	0.8	0.7	$\frac{y_e g^2}{(16\pi^2)^2} \frac{v^2}{\Lambda} f(\Lambda)$	0.6
8	$\frac{v}{\Lambda^3}$	2.2	1.7	$\frac{y_e y_d y_u g^2}{(10^{\circ})^2} \frac{v^4}{\Lambda 3}$	$4.3 \times 10^{-4}$
$8^{H^2}$	$f(\Lambda)\frac{v}{\Lambda^3}$	1.3	1.1	$\frac{y_e y_d y_u g^2}{(16\pi^2)^2} \frac{v^4}{\Lambda^3} f(\Lambda)$	$7.9 \times 10^{-5}$
11a	$\frac{1}{16\pi^2} \frac{y_d v}{\Lambda^3}$	0.2	0.1	$ \frac{y_e y_d y_u g^2}{(16\pi^2)^2} \frac{v^4}{\Lambda^3} f(\Lambda)  \frac{y_d^2 g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda}  \frac{y_u^2}{(16\pi^2)^2} \frac{v^2}{\Lambda}  \frac{y_u^2 g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda} $ $ \frac{y_u^2 g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda} $ $ \frac{y_u^2 g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda} $	$1.2 \times 10^{-5}$
12a	$\frac{1}{16\pi^2} \frac{y_u v}{\Lambda^3}$	0.6	0.5	$\frac{y_u^2}{(16\pi^2)^2} \frac{v^2}{\Lambda}$	$1.9 \times 10^{-3}$
$12b^*$	$\frac{1}{16\pi^2} \frac{y_u v}{\Lambda^3}$	0.7	0.6	$\frac{y_u^2 g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda}$	$2.6 \times 10^{-6}$
13	$\frac{1}{16\pi^2} \frac{y_e v}{\Lambda^3}$	0.2	0.2	$\frac{gegu}{(16\pi^2)^2}\frac{c}{\Lambda}$	$4.5 \times 10^{-4}$
14a	$\frac{1}{16\pi^2} \frac{(y_u + y_d)v}{\Lambda^3}$	0.6	0.5	$\frac{y_u y_d g^2}{(16\pi^2)^3} \frac{v^2}{\Lambda}$	$5.6 \times 10^{-6}$
16	$\frac{1}{16\pi^2} \frac{y_e v}{\Lambda^3}$	0.1	0.1	$\frac{y_d y_u g^4}{(16\pi^2)^4} \frac{v^2}{\Lambda}$	$7.4 \times 10^{-9}$
19	$\frac{1}{16\pi^2} \frac{y_d v}{\Lambda^3}$	0.1	0.1	$\frac{y_e y_u y_d^2 g^2}{(16\pi^2)^3} \frac{v^4}{\Lambda^3} \\ \underline{y_e y_u^2 y_d g^2}_{AB} \frac{v^4}{\Lambda^3}$	$2.4 \times 10^{-6}$
20	$\frac{1}{16\pi^2} \frac{y_u v}{\Lambda^3}$	0.5	0.4	$\frac{y_e y_u^2 y_d g^2}{\left(16\pi^2\right)^3} \frac{v^4}{\Lambda^3}$	$1.8 \times 10^{-6}$





# Constraints on the scale of New Physics

Process	Experimental limit	O	$\Lambda_{ijkn}^{\mathrm{NP}} [\mathrm{TeV}]$
$K^+ \to \pi^+ \nu \nu$	$BR_{\text{future}}^{\text{NA62}} < 1.11 \times 10^{-10}$	$\mathcal{O}_{3b}$	$\sum_{i} \Lambda_{iisd} > 19.6$
$K^+  o \pi^+ \nu \nu$	$BR_{\text{current}}^{\text{NA62}} < 1.78 \times 10^{-10}$	$\mid \mathcal{O}_{3b} \mid$	$\sum_{i} \Lambda_{iisd} > 17.2$
$K_L \to \pi^0 \nu \nu$	$BR_{current}^{KOTO} < 3.0 \times 10^{-9}$	$\mathcal{O}_{3b}$	$\sum_{i=1}^{n} \Lambda_{iisd} > 12.3$

## Sensitivity to different flavors than most constraining $0v\beta\beta$ !

O	$1/\Lambda_{K \to \pi \nu \nu}^2$	$\sum_{i} \Lambda_{iisd}^{\text{E949}} \text{ [TeV]}$	$m_ u$	$\Lambda^{m_{\nu}} [\text{TeV}]$
$1^{y_d}$	$\frac{v^3}{\Lambda^5}$	2.4	$\frac{y_d}{16\pi^2} \frac{v^4}{\Lambda^3}$	11.6
3b	$\frac{v}{\Lambda^3}$	11.5	10% - 11	$5.2 \times 10^4$
$3b^{H^2}$	$f(\Lambda)\frac{v}{\Lambda^3}$	5.7	$\frac{y_d}{16\pi^2} \frac{v^2}{\Lambda} f(\Lambda)$	330
5	$\frac{1}{16\pi^2} \frac{v}{\Lambda^3}$	2.6	$\frac{y_d}{(16\pi^2)^2} \frac{v^2}{\Lambda}$	330
10	$\frac{1}{16\pi^2} \frac{y_e v}{\Lambda^3}$	0.8	$\frac{y_e y_d}{(16\pi^2)^2} \frac{v^2}{\Lambda}$	$9.6 \times 10^{-4}$
11b	$\frac{1}{16\pi^2} \frac{y_d v}{\Lambda^3}$	0.8	$\frac{y_d^2}{(16\pi^2)^2} \frac{v^2}{\Lambda}$	$330$ $9.6 \times 10^{-4}$ $8.9 \times 10^{-3}$ $4.1 \times 10^{-3}$ $330$ 1st generation couplings
14b	$\frac{1}{16\pi^2} \frac{y_u v}{\Lambda^3}$	2.9	$\frac{y_d y_u}{(16\pi^2)^2} \frac{v^2}{\Lambda}$	$4.1 \times 10^{-3}$
66	$f(\Lambda)\frac{v}{\Lambda^3}$	5.1	$\frac{y_d}{16\pi^2} \frac{\dot{v}^2}{\Lambda} f(\Lambda)$	330 1st generation couplings



