

# Gravitational-wave event validation by Advanced LIGO and Advanced Virgo detectors

Procedures and challenges for the  
upcoming observing runs

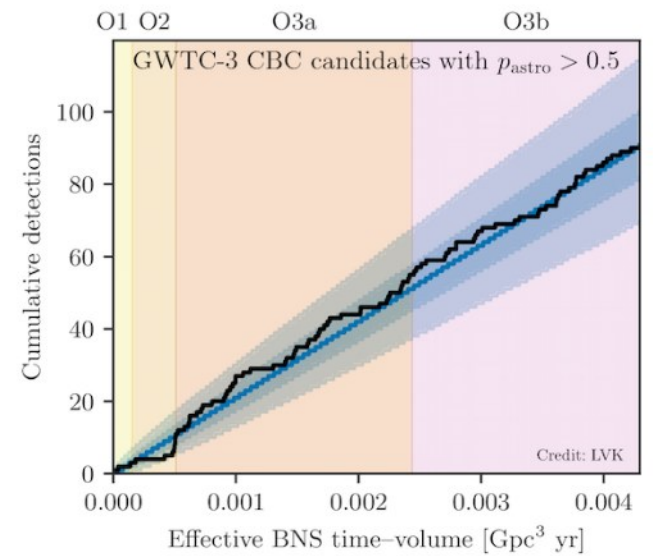
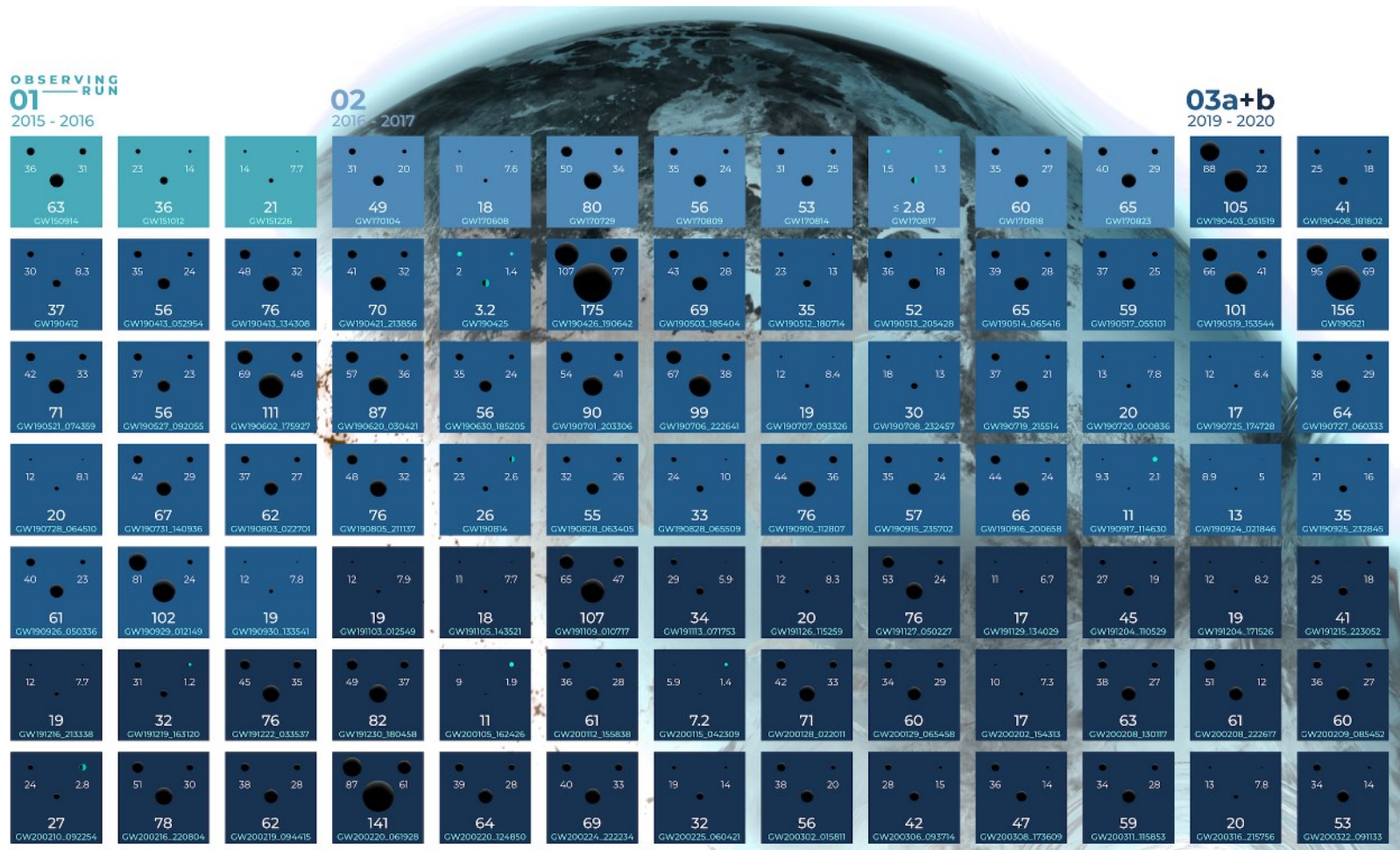


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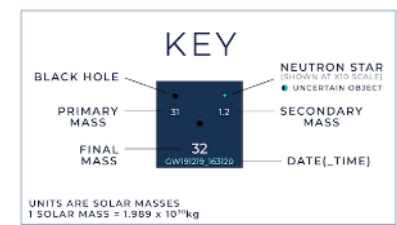
Francesco Di Renzo  
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# Compact binary observations by Advanced LIGO and Advanced Virgo detectors

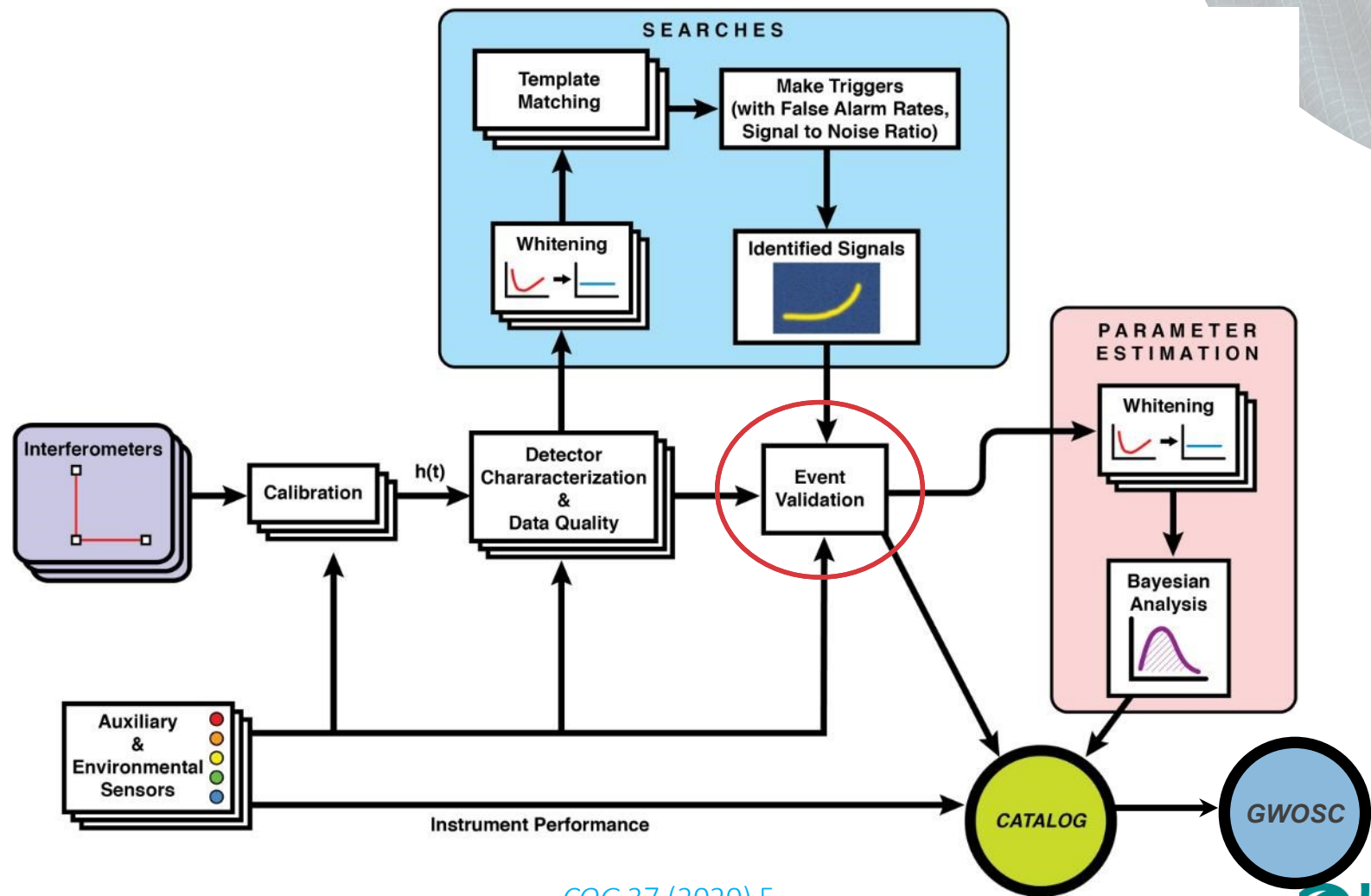


[arXiv:2111.03606](https://arxiv.org/abs/2111.03606)

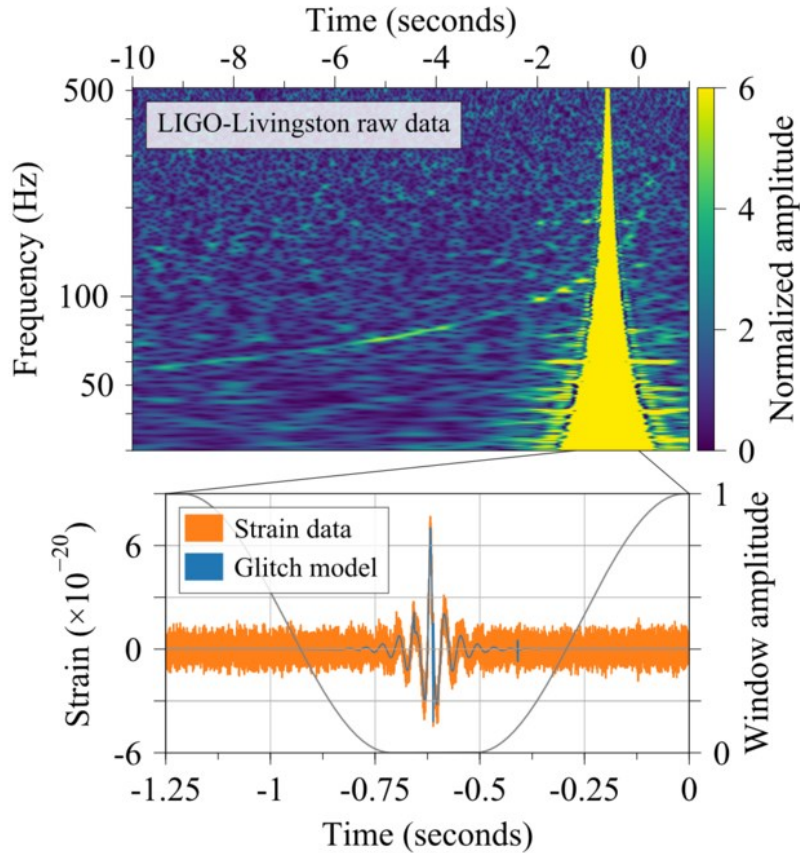




# Data processing overview: from detectors to publications



# The validation of gravitational-wave events



[PRD 98, 084016 \(2018\)](#)

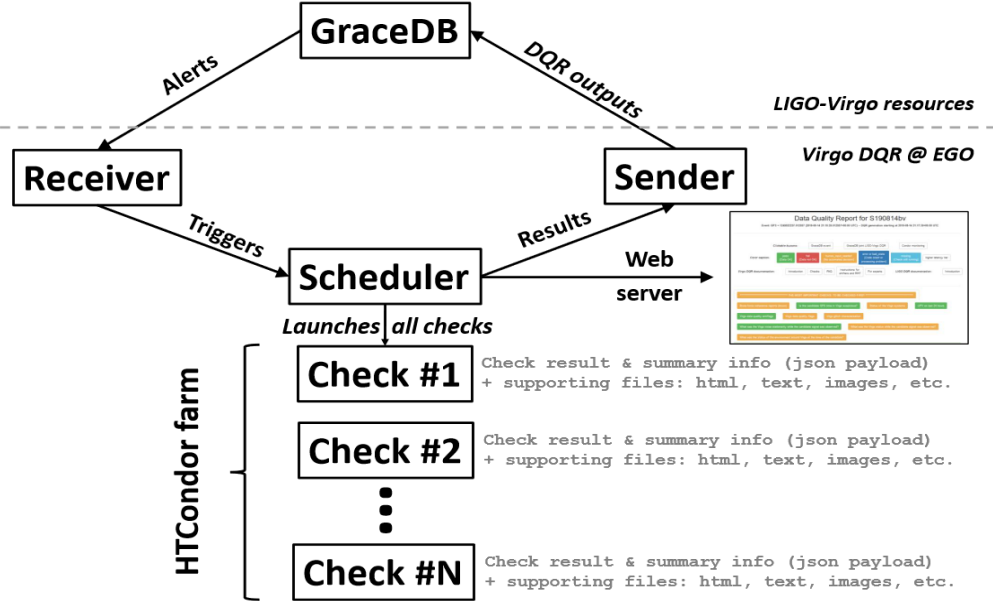
- Event validation consists of a set of procedures to verify if **data quality (DQ) issues**, such as instrumental artifacts, environmental disturbances, or anomalies in the search pipelines, can impact the analysis results and **decrease the confidence of a detection**;
- It is applied to all gravitational-wave transient **candidate events** found by both *online* and *offline* search pipelines;
- Typically, candidate events undergo **two stages of validation**:
  - **Prompt validation (online triggers only):**

Accompanies every public alerts and is typically completed within  $\mathcal{O}(10 \text{ min})$  from the data acquisition. It has the role to **vet** an event trigger if there is evidence of terrestrial origin or other severe DQ issues;
  - **Final validation (all):**

Completed as a final check before publication for all events found by online and/or offline pipelines. The typical timescale is days or even months after the time of the event.

# The Data Quality Report framework

Schematics of the Virgo O3 DQR architecture, from [arXiv:2205.01555](#)



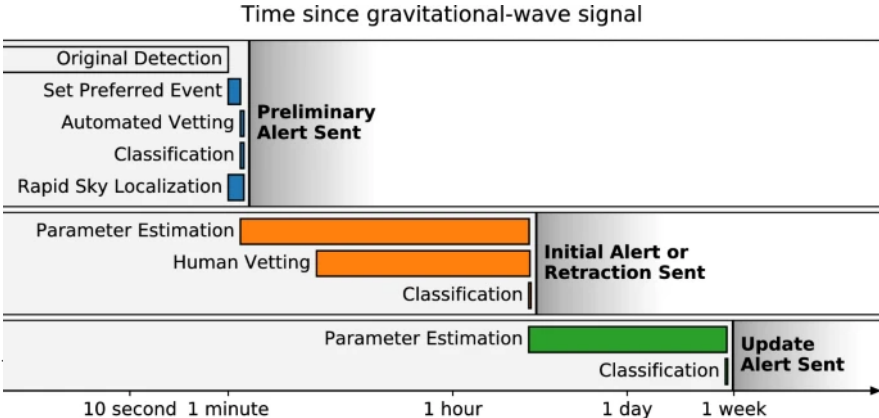
- A Data Quality Report (DQR) is a [framework developed by LIGO and Virgo](#) consisting in a set of DQ checks;
- It is automatically prompted after each gravitational-wave candidate trigger with false alarm rate (FAR) of 1/day is being generated on [GraceDB](#);
- The results are uploaded back to [GraceDB](#) and used by the **Rapid Response Team** to validate or vet the associated event, and afterwards for the final event validation.

Table: Performance of Virgo DQR during O3b, from [arXiv:2205.01555](#)

Operation	Time taken [s]		
	Median	Mean	95 <sup>th</sup> percentile
Data acquired → Candidate on GraceDB	52	166	331
Candidate on GraceDB → LVAalert trigger	4	4	11
LVAalert trigger → Virgo DQR configured	331	339	383
Virgo DQR configured → Virgo DQR started	8	10	21

Operation	Time from start [s]		
	Median	Mean	95 <sup>th</sup> percentile
Quick key checks	374	383	619
Adding Omicron trigger distributions	868	816	935
Adding full Omicron scans	1740	2159	4690
End	5185	4954	6330

[Living Rev.Rel.](#)  
[23 \(2020\) 1, 3](#)

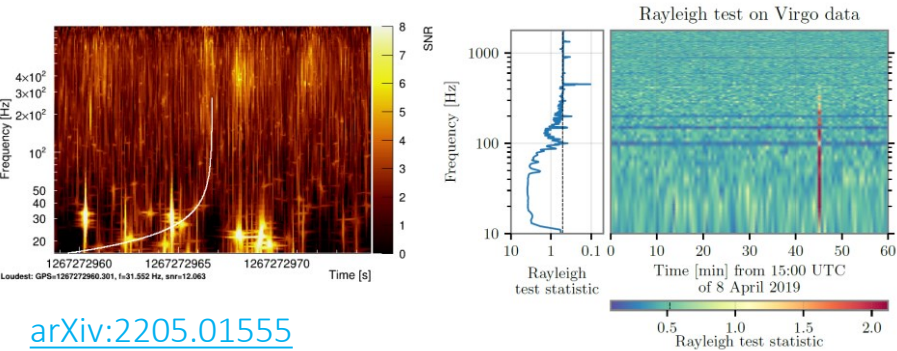




# Prompt event validation of low-latency alerts

DMS		ITF Mode: Commissioning (2019-04-08 00:00)										ITF State: LOCKED_ARMED (2019-04-08 00:00)										UTC: 2022-07-08 08:00	
Injection	SIB1_IP		SIB1_BENCH		SIB1_BR		SIB1_Vert		SIB1_Temp		MC_TE		SIB1_Guard		SIB1_Electr								
	MC_IP		MC_PAY		MC_BR						MC_TE												
	Laser		LaserAmp		LaserChiller		SL_TempController		RFC				LNFS		PC								
	SLC_Ba_MC_Temp		MC_Power		PSTAB		MC_AA		MC_AA_GALVO		MC_F0_2				BPC		BPC_Electr						
Detection	PD		PD_BR		QPD_Bip		QPD_B2		QPD_B4		QPD_B5		OMC		PhotoDiode		Shutter						
	SDB1_IP		SDB1_LC		SDB1_BR				SDB1_Vert				SDB1_TE		SDB1_Guard		SDB1_Electr						
ISC		PR_parking										SR_parking											
ALS		NE_ALS_Laser				NE_ALS_ARM				WE_ALS_Laser				WE_ALS_ARM				CEB_ALS_Laser					
Suspensions	BS_IP		BS_F7		BS_PAY		BS_BR		BS_Vert		BS_TE		BS_Guard		BS_Electr								
	NI_IP		NI_F7		NI_PAY		NI_BR		NI_Vert		NI_TE		NI_Guard		NI_Electr								
	NE_IP		NE_F7		NE_PAY		NE_BR		NE_Vert		NE_TE		NE_Guard		NE_Electr								
	PR_IP		PR_F7		PR_PAY		PR_BR		PR_Vert		PR_TE		PR_Guard		PR_Electr								
	SR_IP		SR_F7		SR_PAY		SR_BR		SR_Vert		SR_TE		SR_Guard		SR_Electr								
	WI_IP		WI_F7		WI_PAY		WI_BR		WI_Vert		WI_TE		WI_Guard		WI_Electr								
	WE_IP		WE_F7		WE_PAY		WE_BR		WE_Vert		WE_TE		WE_Guard		WE_Electr								
Environment	CD_Hall		MC_Hall		TCS_zones		NE_Hall		WE_Hall		WindActivity		Seismon		BRMSon								
	InJ_A_Room		DET_A_Room		EE_Room		DAQ_Room		External		DeadChannel		FlatChannel_ENV		Lights		SnkActivity						

Example of Virgo DMS. From [Virgo logbook entry #56363](#)  
(NOT a candidate event) [VIR-0191A-12](#)



[arXiv:2205.01555](#)

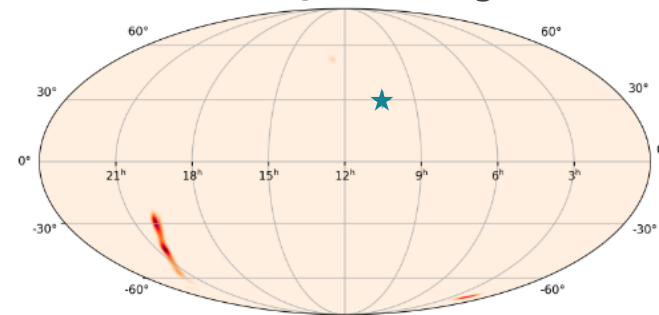
166.88	SDB1_LC_TZ_fb (0.24)	SDB1_LC_TZ_corr (0.24)	SDB1_LC_TZ_err (0.24)	SDB1_LC_TZ (0.24)	SDB1_LC_COIL_FL_V (0.24)
167.00	SDB1_LC_TZ_fb (0.32)	SDB1_LC_TZ_corr (0.32)	SDB1_LC_COIL_FL_V (0.32)	SDB1_LC_COIL_BR_V (0.32)	SDB1_LC_COIL_BR_V (0.32)
167.12	SDB1_LC_COIL_FR_V (0.45)	SDB1_LC_COIL_BR_V (0.45)	SDB1_LC_COIL_FL_V (0.45)	SDB1_LC_COIL_BR_V (0.45)	SDB1_LC_TZ_err (0.45)

- This stage has the role to **vet those event triggers with severe noise contamination**, for which an astrophysical origin should be excluded;
- Otherwise, it serves to enforce the confidence in the event type and **sky-localization** to support **multimessenger follow-up**.
- The **main DQ checks** based on the DQR are:
  - Operational **status of the detector** and its subsystems at the time of the trigger and around it;
  - Scan of the **main DQ flags**:  $h_{\text{rec}}$  correctly computed, detector observational intent and working condition, injections of spurious signals, etc.
  - **Noise characterization**: stationarity and Gaussianity, including the presence of glitches and their distribution; correlation with auxiliary channels; status of the environment, etc.

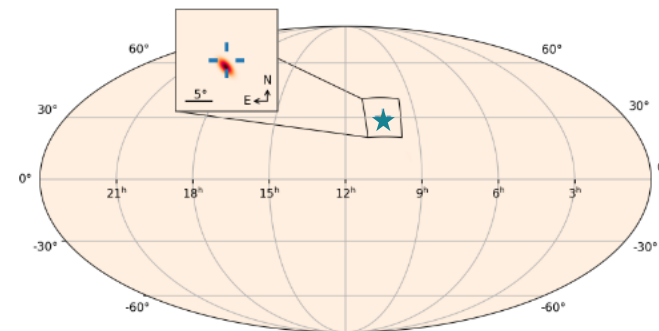
# Final validation before publications

- **Every LVK publications** (catalogs and exceptional events) undergo a final, comprehensive validation procedure before data analysis reruns;
- This includes all the events found online and pre-validated and those found by offline pipelines;
- An event validation team is in charge of this procedure. Each event requires  $\mathcal{O}(1 \text{ hour})$  per person involved if no DQ issue is found;
- The goal is to assess whether the parameter estimation of the astrophysical source can be affected by noise artifacts;  
[CQG 35 \(2018\) 15, 155017](#)
- If no DQ issue is found, the candidate event is considered validated;
- For those events where noise artifacts are found in the vicinity of the putative GW signal, or even overlapping with it, a procedure of **noise mitigation** is implemented. This requires additional time and person power.

Glitch: 90% credible region 137 deg<sup>2</sup>

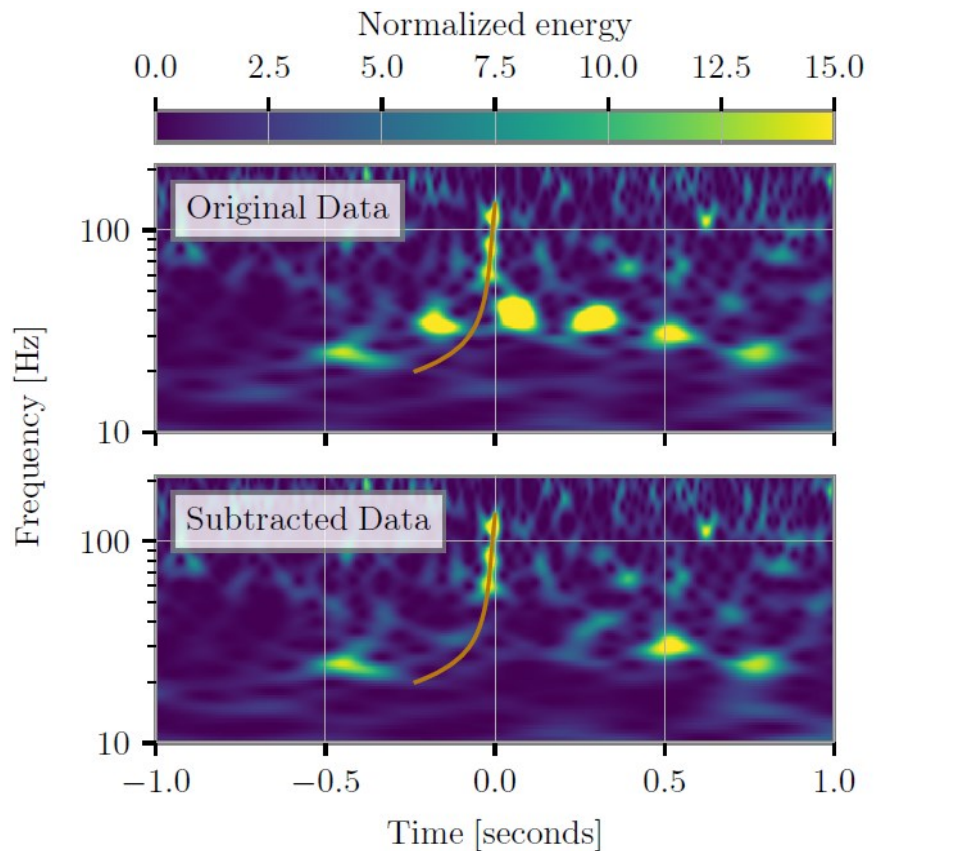


No glitch: 90% credible region 8 deg<sup>2</sup>



Effect on sky-localization of a blip glitch 30 ms after a GW150914-like event. [PRD 105 \(2022\) 103021](#)

# Noise artifacts mitigation of gravitational-wave detector data



[PRX 11, 021053 \(2021\)](#)

- Applied to those events flagged to have DQ issues: transient noise, namely **glitches**, superimposing the putative astrophysical signal (orange curve);
- Metric based on the **PSD variation** to assess the extent of each non-stationary region identified [[CQG 37 \(2020\) 21](#)];
- Deglitched frames mostly produced with BayesWave pipeline [[CQG 32 \(2015\) 13](#)];
- Assessment of subtraction by means of the previous stationarity metric. Parameter Estimation comparison tests to check for bias and systematics;
- **16 events** ( $\approx 20\%$ ) required glitch subtraction during O3. This process involves lots of human input and slows down downstream analyses..



# Lessons learned and preparation for O4

Run	Total	Online	Retracted	Offline	DQ issues
O3a	44	27-3	8+3	12+8	4
O3b	35	18	16	17	12
O3	79	42	27	37	16

O3 catalog events ([GWTC-2.1](#) + [GWTC-3](#)) with  $p_{\text{astro}} > 0.5$

## Expectation for O4 events

- $2 \div 4$  times larger rate of candidates than O3;
- If the glitch rate remains the same, expect 20% of them with DQ issues;
- Consider  $1/4 \div 1/2$  of the time to be dedicated to event validation and glitch subtraction with respect to O3.

## O4 preparation tasks

- Expect much **more data (and DQ issues)** and the same person power;
- During O3, typical event validation required  $\mathcal{O}(1\text{h})$  per person, and a few days for noise mitigation:
  - Try to reduce the requirement of **human inputs**;
  - Improve **training** of validation rota members;
  - Invest in **automatization**: “vanilla events” with no DQ issues should be automatically validated with no or just minimal further human inputs;
  - Identify a set of **common scenarios** and prepare clearer guidelines to speed up the analyses.
- Update the DQR infrastructure (in progress);
- **Share tools and techniques and plan collaborative trainings.**

# Conclusions

- Event validation is an **integrating part of gravitational-wave data analysis** with the role of enforcing the confidence in the astrophysical origin of a transient signal detected by search pipelines, and the reliability of the source parameter estimation results;
- Preliminary event validation, jointly with low-latency alerts, should be completed in the quickest time to **support multimessenger follow-up** searches by the astronomical community;
- The presence of **DQ issues** delays the completion of the validation tasks and requires additional noise mitigation procedures;
- For the expected rate increase in O4 of transient gravitational-wave events, all **the above tasks should be sped up**. LVK is working on the automatization of part of the validation procedure, reducing the requirement of human input;
- **Cooperation inside LVK** is paramount to implementing common validation procedures, included in the DQR framework, and the training and support of an event validation team before the next observing run.

*Thanks for the attention!*

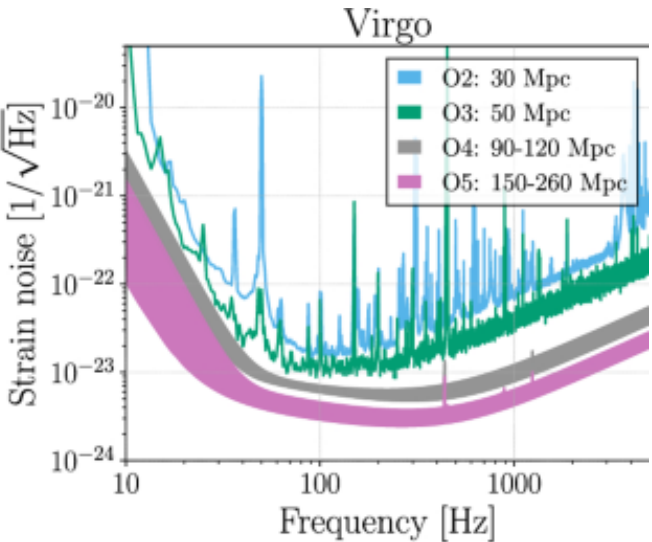
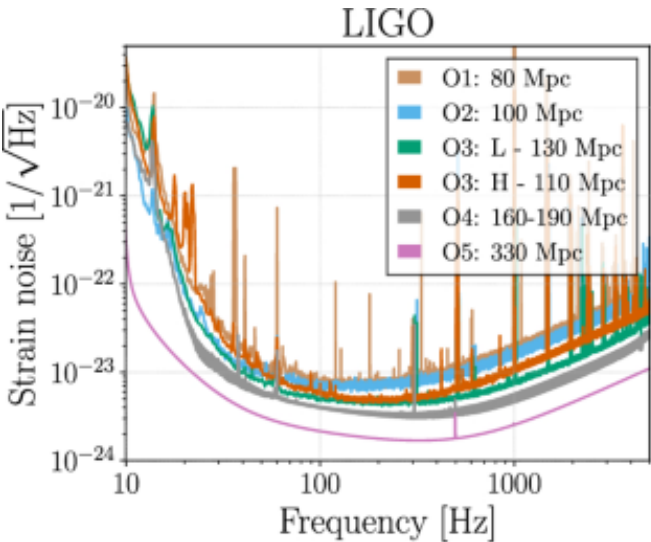


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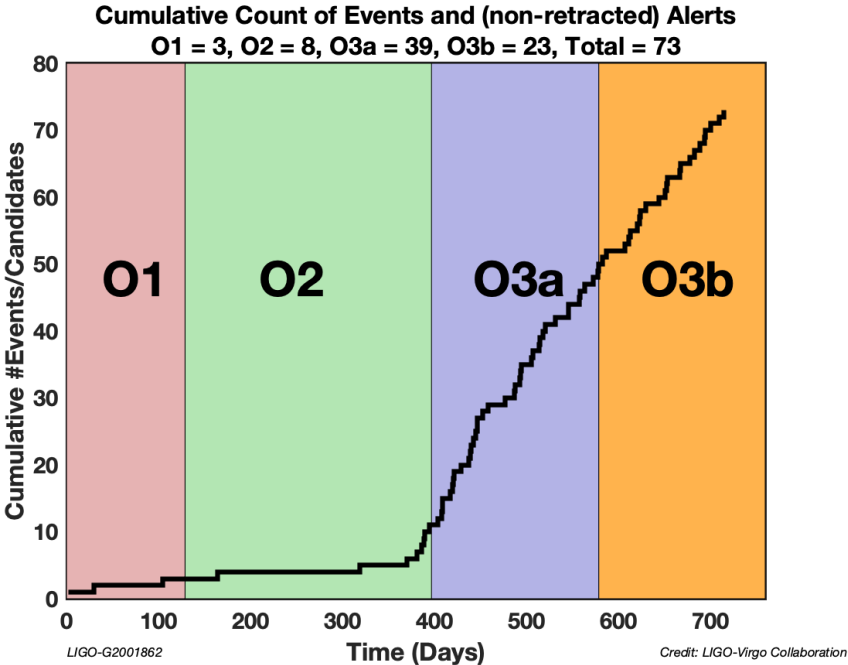


# Sensitivity improvements and event rate vs time

A sensitivity range improvement of  $50 \div 200\%$  implies an increase in observed volume of  $2 \div 4$  times, and an equal increase in detected events.



[Living Rev.Rel. 23 \(2020\) 1, 3](#)



[LIGO-G2001862](#)