

Thermal production of sexaquarks in heavy ion collisions

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on behalf of

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Outline

- I Introduction
- II uuddss recent searches
- III Thermal model results on uuddss
- IV Summary, Conclusions and Outlook

H-uuddss

The uuddss state can be maximally bound due to its symmetry.
 Due to being a flavor singlet it does not bind to pions resulting in a compact configuration.
 Assuming it can bind to lightest flavor singlet mesons like the f0 a radius of 0.1-0.3 fm is estimated.
 The different size of the S and baryons means that amplitudes involving the S and 2 baryons are strongly suppressed.
 Lacking coupling via pions its interaction with matter is lower than that of ordinary hadrons supporting the hypothesis that it can be a DM candidate.

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and
 G. R. Farrar, (2018), arXiv:1805.03723 [hep-ph]

The uuddss diquark state has been initially proposed by Robert Jaffe (MIT) who made a bag model based calculations and predicted the H-dibaryon with a mass below 2m(Lambda) of about 2150 MeV and unstable.

R. L. Jaffe, Phys. Rev. Lett. **38**, 195 (1977)

The H dibaryon has been searched by several experiments without finding it.

B. H. Kim et al. (Belle), Phys. Rev. Lett. **108**, 222002 (2012),
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R. L. Jaffe, Phys. Rev. Lett. **38**, 195 (1977)

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S-uuddss

Recently a new possibility for the uudds multiquark state has been proposed by Glennys Farrar (NYU) in which the uudds is a state with small radius (0.1-0.4 fm) and mass below 2 GeV (so called Sexaquark).

This state has been proposed to be stable and could be produced in the experimental searches till now since (depending on its mass) it can be absolutely stable or have a lifetime of the order of the age of the universe.

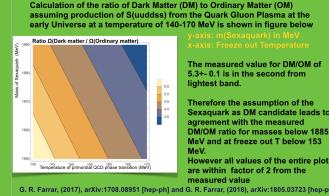
Such a stable state could be a candidate for Dark Matter; if produced out of the primordial Quark Gluon Plasma at the QCD phase transition from partons to hadrons.

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and
 G. R. Farrar, (2018), arXiv:1805.03723 [hep-ph]

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S-uuddss

Calculation of the ratio of Dark Matter (DM) to Ordinary Matter (OM) assuming production of S-uuddss from the Quark Gluon Plasma at the early Universe at a temperature of 140-170 MeV is shown in figure below



G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and G. R. Farrar, (2018), arXiv:1805.03723 [hep-ph]

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S-uuddss

The previous calculation of the ratio of Omega Dark Matter to Omega Ordinary Matter assuming a Sexaquark as DM candidate is valid for all forms of DM including equal u,d,s pairs (like stable sexaquarks, quark nuggets or primordial black holes) or a combination of them.

As a result the observed value of Omega(DM)/Omega(Matter) has been reproduced by assuming that DM is composed of equal number of u,d quarks

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and G. R. Farrar, (2018), arXiv:1805.03723 [hep-ph]

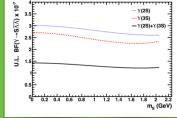
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Experimental searches

Even though not everyone agrees, its possible cosmological implications as DM candidate cannot be excluded and it has been recently searched in the BaBar experiment that set upper limits in

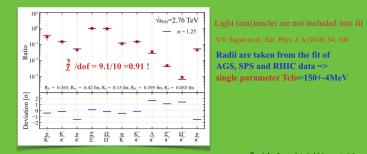
the BaBar experiment that set upper limits in

BABAR Coll. J. P. Lees et al., Phys.Rev.Lett. **122** (2019) no.7, 072002



90% confidence level upper limits on the branching fraction $\Upsilon(3S,3S) \rightarrow S + \bar{\Lambda} + \bar{\Lambda}$ as well as the combined sample

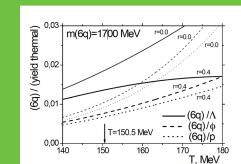
Hadron Resonance Gas Model (HRGM) with Induced Surface Tension EOS Results for LHC energy



K. A. Bugaev et al., Nucl. Phys. **A970** (2018) 133-155 and references therein,
 K. A. Bugaev et al., Universe **5** (2019) 63

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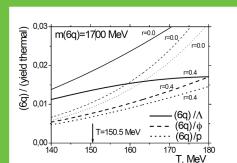
PRELIMINARY



D. Blaschke et al., Int.J.Mod.Phys.A **36** (2021) 25, 2141005

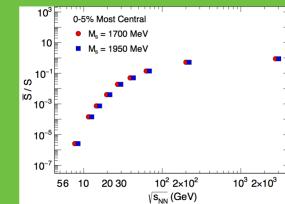
Thermal production of S-uuddss in heavy ion collisions

PRELIMINARY



D. Blaschke et al., Int.J.Mod.Phys.A **36** (2021) 25, 2141005

Antisexaquark to Sexaquark ratio vs collision energy in Au+Au collisions for two assumed masses



Thermal model calculation based on
 S. Kabana, P. Minkowski, New J.Phys. **3** (2001) 4

* In the framework of the thermal model by K. Bugaev et al., and of another thermal model by S. Kabana and P. Minkowski, we estimated ratios of Sexaquarks to hadrons.

* Sexaquarks are produced at relatively high rates, for vanishing and finite radius of 0.4 fm, and for masses of 1700 and 1950 MeV at the LHC.

* At $T=170$ MeV at LHC the ratio of thermal Sexaquark with mass 1950 GeV to thermal deuteron is about 0.45

* Antisexaquark ratios are shown here for the first time and are maximal at top RHIC energy and similar to LHC.

Summary, Conclusions and Outlook