



Discovering neutrinoless double-beta decay in the era of precision cosmology

Manuel Ettengruber¹, Matteo Agostini², Allen Caldwell¹, Philipp Eller³, Oliver Schulz¹

¹Max-Planck-Institut für Physik, ²Department of Physics and Astronomy, University College London, ³Technical University Munich (TUM)

Introduction

- Neutrinoless Double Beta Decay ($0\nu\beta\beta$) is a nuclear beta Decay which emits two left-handed electrons but no neutrinos.
- This could be possible due to the possible Majorana nature of neutrinos.
- The next-Gen $0\nu\beta\beta$ experiments are funded and designed to investigate the whole parameter space of Inverted Ordering.
- Questions addressed in this work are:
 - How is the Discovery Probability (P_D) of these experiments in case of Normal Ordering?
 - And how is the $0\nu\beta\beta$ hunt influenced by cosmological measurements of the neutrino mass sum?

Discovery Probability of $0\nu\beta\beta$ -experiments

- The half life for $0\nu\beta\beta$ can be calculated via

$$\frac{1}{T_{1/2}} = G g_A^4 \mathcal{M}^2 \left(\frac{m_{\beta\beta}}{m_e} \right)^2, \quad (1)$$

where G is the kinematically allowed phase space factor, $g_A^4 \simeq 1.276$ is the axial-vector coupling, \mathcal{M} the nuclear matrix element (NME).

- The crucial parameter is the effective majorana mass $m_{\beta\beta}$ which is calculated the following way

$$m_{\beta\beta} = |c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 e^{i\alpha_1} + s_{13}^2 m_3 e^{i\alpha_2}|, \quad (2)$$

where c_{ij} and s_{ij} are the cosines and sines of the oscillation angles, m_i the eigenvalues of the neutrino mass eigenstates and α_i are the so-called Majorana phases.

- We calculate a combined posterior of data from neutrino oscillation experiments, cosmology and previous $0\nu\beta\beta$ experiments.

- By sampling from the posterior, we generate pseudo-data sets for the future $0\nu\beta\beta$ -experiments and calculate the Posterior Odds.

- We define our **Discovery Criteria** as follows

$$\mathcal{O}_1 = \frac{P(D|H_1)P(H_1)}{P(D|H_0)P(H_0)} > 10, \quad (3)$$

where $P(D|H)$ are the probabilities of the data given the hypothesis that $0\nu\beta\beta$ exists (H_1) or not (H_0), and $P(H_{1/0})$ their corresponding priors which are assumed to be equal.

Results

- We perform a scan over m_1 for different NME models.
- We find that the P_D starts rising for values of $m_1 > 1\text{meV}$.

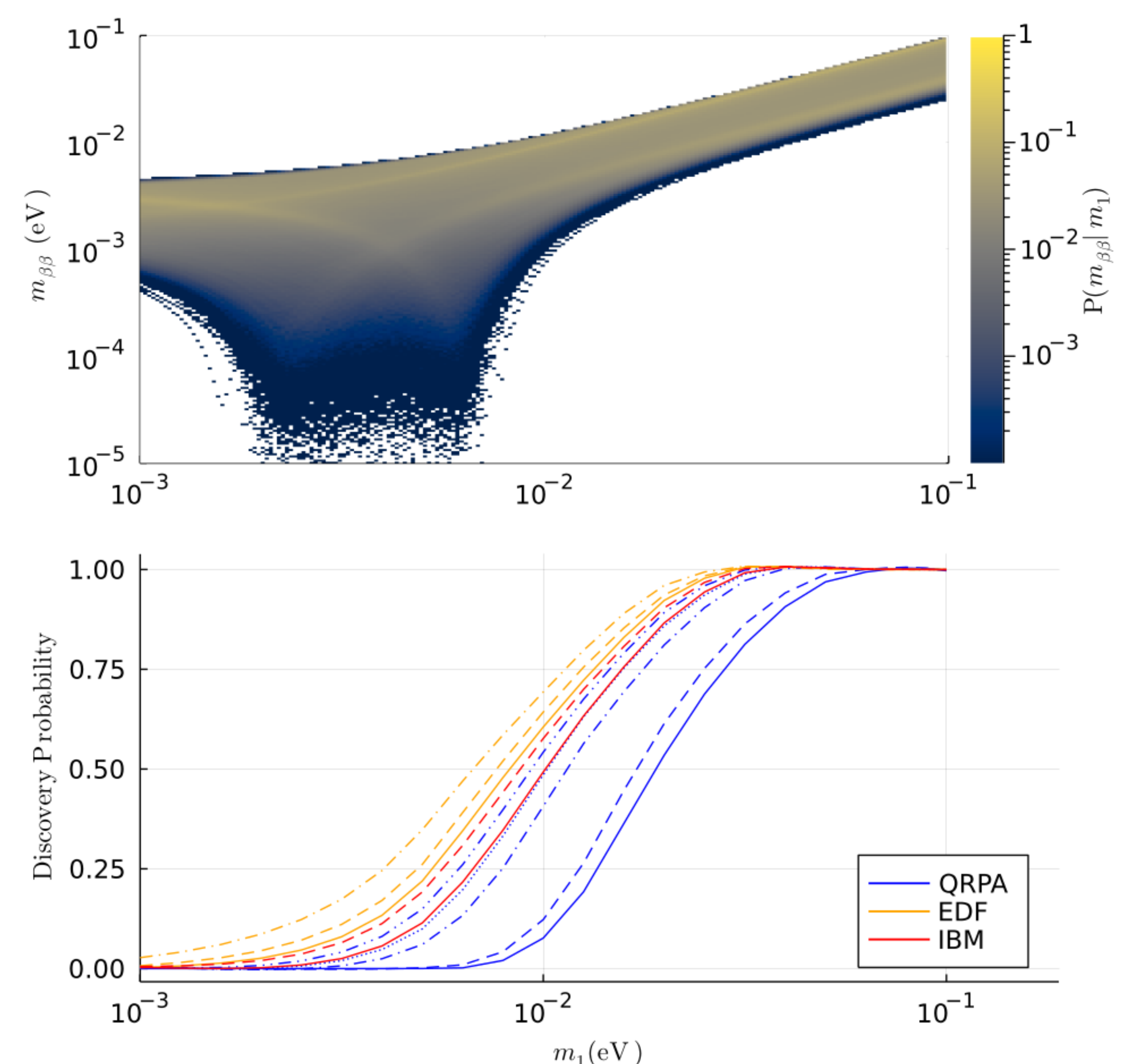


Figure 1: (Top) Conditional one-dimensional posterior probabilities for $m_{\beta\beta}$ computed for a scan of fixed m_1 values. (Bottom) Discovery probability, P_D , for a combined analysis of CUPID, LEGEND and nEXO as a function of the true value of m_1 .

- In the second analysis we investigated the influence of current cosmological bounds and possible future detections or bounds on the neutrino mass sum.

- The most optimistic scenarios reach a P_D up to 80 – 90%.
- The most pessimistic scenario can still reach a P_D up to 50%.
- All calculations are heavily influenced by the chosen NME model (horizontal ticks in FIG.2).

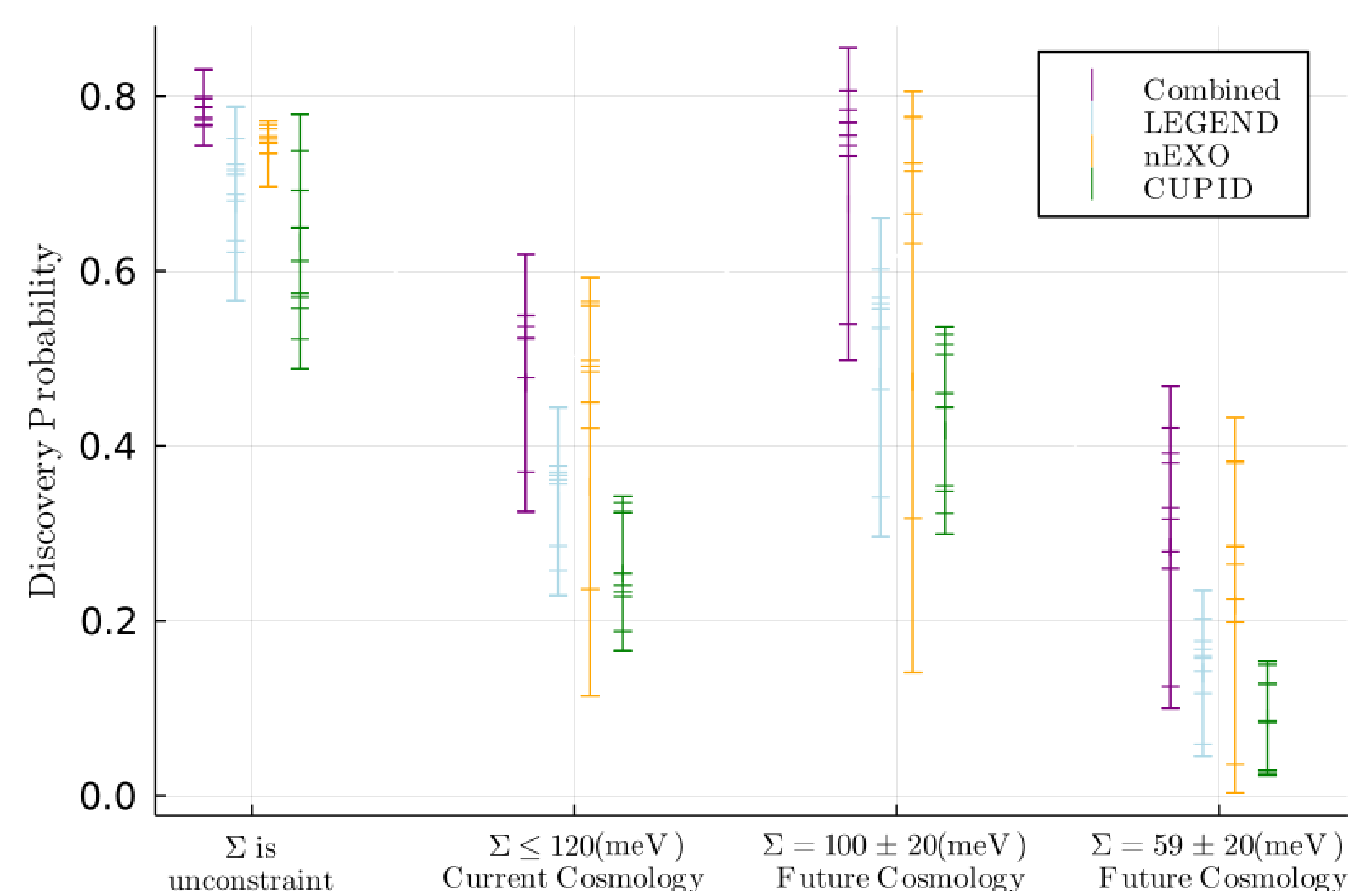


Figure 2: Discovery probabilities for a selection of proposed experiments and their combination under different scenarios and set of NME values. The calculation has been performed using fixed sets of NME values.