TAU NEUTRINO APPEARANCE IN THE FLUX OF ATMOSPHERIC NEUTRINOS

AN UNAMBIGUOUS CONFIRMATION OF

NEUTRINO OSCILLATIONS -

DIRECT DETECTION OF OSCILLATION INDUCED

TAU NEUTRINOS IN THE ATMOSPHERE

IN THE ENERGY RANGE OF \sim 3.5 GeV TO 100 GeV.

92%

neutrinos,

Inputs

to the

MC studies

shown here

are with 500

years of

each

signal and

weighed to

a live-time of

the detector

of ~9 years

events have

been scaled

background,

otherwise.

(SK-IV run

period).

The tau

to the

unless

stated

background,

AT SUPER-KAMIOKANDE

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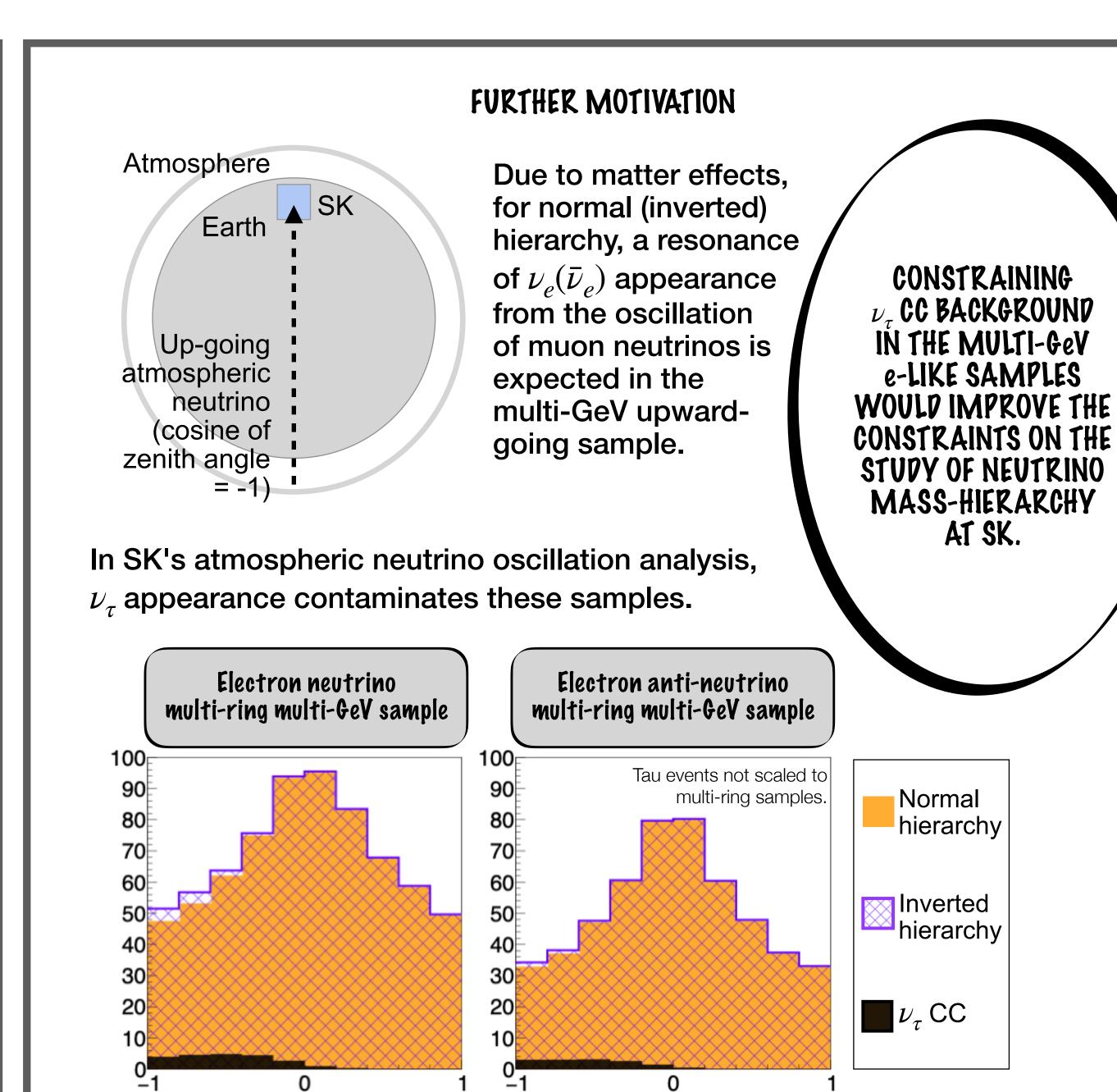




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FOR NUCLEAR

CENTRE



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Signal

 $(\nu_{\tau} CC)$

interactions)

~800 events

Secondary cosmic rays decay to produce atmospheric neutrinos which are mainly of μ and e flavors.

Atmospheric neutrinos

Super-Kamiokande (SK)

- Water Cherenkov detector.
- Cyclindrical tank filled with ultra-pure water until 2019.
- Lined inside with ~11000 PMTs.

In 2019, **SK** water was doped with gadolinium, increasing its capacity for detecting neutron captures.

Charge (pe)

MUON NEUTRINOS IN THE FLUX OF ATMOSPHERIC NEUTRINOS ARE EXPECTED TO OSCILLATE PREDOMINANTLY TO TAU NEUTRINOS.

The background to the ν_{τ} CC signal is made up of the following interactions of the atmospheric

Background

~9000 events

 ν_{μ} CC interactions,

 ν_{e} CC interactions, and NC interactions of all flavors.

THE MANY CHARGED PARTICLES PRODUCED IN THESE EVENTS ARE SEEN AS MULTIPLE RINGS AT SK.

Tau produced in ν_{τ} CC interactions, decay to produce many particles, which makes the reconstruction of the tau difficult.

> SK USES A NEURAL NETWORK (NN) TO PICK THE TAU SIGNAL.

A typical tau neutrino CC PIS interaction at SK 2 mu-e 2 decay 500 1000 1500 200 Times (ns) A typical background to the tau signal, muon neutrino CC PIS interaction at SK

> 2 mu-e 2 decay 500 1000 1500 2000 Pi+ e+ Times (ns)

LARGER FIDUCIAL VOLUME (FV) IS EXPECTED WHAT'S NEW? TO ADD 12% MORE TAU EVENTS.

SK NN performs well in the expanded FV, selecting tau neutrinos with an efficiency of 74% (signal purity of 75%).

Standard FV, $d_{\text{wall}} > 200$ cm.

neutrino',

interaction

vertex

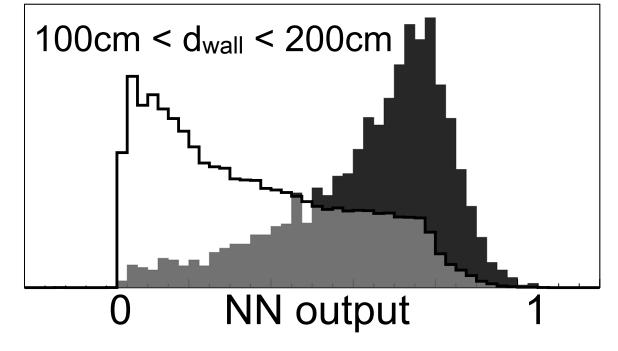
SK inner

detector

dwall

wal

Expanded FV, $d_{\text{wall}} > 100 \text{cm}$.



0.0 0.2 0.4 0.6 0.8 1.0 Fraction of energy carried by the most energetic ring PID of the most energetic ring

O.O O.2 O.4 O.6 O.8 1.0 Clustered sphericity 3.5 4 4.5 5 5.5 6 log₁₀(visible energy

in MeV) log₁₀ (distance to the decay electron in cm)

Decay electrons

5 10 15 20 25 30 Cherenkov ring fragments

METHOD AND RESULTS NN output Signal (tau neutrino CC inter THE SK NN IDENTIFIES -action) TAU NEUTRINOS WITH Back-72% EFFICIENCY ground AND A SIGNAL PURITY OF 77%. 0.5 NN output

Cosine of zenith angle

Cosine of zenith angle

* Events with NN ouput > 0.5 are classified as tau-like. Signal $(\nu_{ au}\, {\sf CC}$ 80% interactions) Background ~2000 events ~500 events

 $Bkg(\nu_{\mu}CC)$ $Bkg(\nu_eCC)$ — BkgNCSignal Tau signal not scaled to background. 0.5 NN output ν_{μ} CC ν_{ρ} CC NC % of misclassified background 13% 37% 70%

On average, more neutrons are expected to be produced in the NC interactions than in the tau signal. NN inputs related to neutron captures could declassify the misidentified NC background.

WHAT'S NEXT?

OF ALL THE

BACKGROUND

22% 18

MISIDENTIFIED

AS TAU-LIKE.

- Data from the latest SK periods, combined with the results of higher neutron tagging efficiency from the SK-Gd upgrade.
- Software upgrades to the SK NN.

NEW TAU CC CROSS-SECTION AND NORMALISATION MEASUREMENTS, AND IMPROVED MASS-HIERARCHY RESULTS.

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Previous analysis: Z. Li et al. (SK Collaboration) Phys. Rev. D 98, 052006