

# Neutrino upscattering to HNLs as an explanation of MB LEE

## ICHEP '22

Jaime Hoefken Zink

In collaboration with A. Abdullahi, M. Hostert, D. Massaro, S. Pascoli

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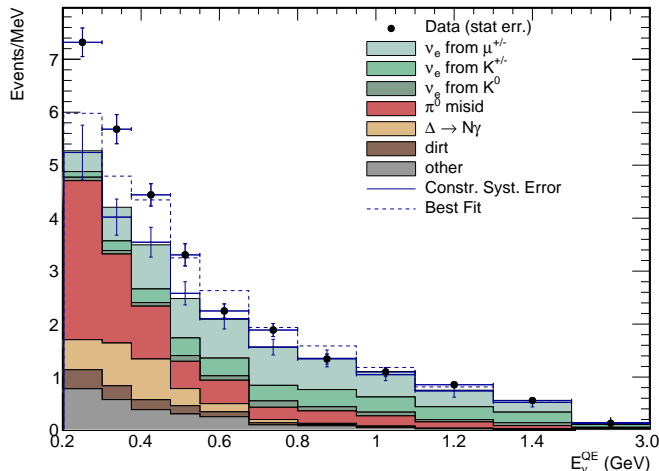
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# Anomaly: MiniBooNE Low Energy Excess

Neutrino  
upscattering to  
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explanation of MB  
LEE

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$560.6 \pm 119.6(77.4 \pm 28.5)$  in  $\nu(\bar{\nu})$  mode:  **$4.8\sigma$**

Problem: Anomaly

Dark Photon  
Model

Simulation

Results

# Anomaly: possible solutions

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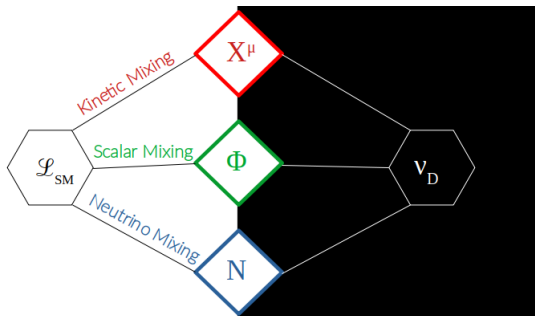
Simulation

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Topology	Model	Diagram	Signal
single $\gamma$	Transition magnetic moment		$N \rightarrow \nu \gamma$
	neutrino-induced inverse-Primakoff scattering		$a^* + A \rightarrow \gamma A$
$e^+ e^-$	upscattering to $N$		$N \rightarrow \psi_{inv} e^+ e^-$
	bremsstrahlung of light $Z'$		$N^* \rightarrow \nu (Z' \rightarrow e^+ e^-)$
	neutrino-induced $Z'$ fusion		$Z'^* Z'^* \rightarrow (S \rightarrow e^+ e^-)$
	neutrino-induced inverse-Primakoff scattering to $Z'$		$S^* \gamma^* \rightarrow (Z' \rightarrow e^+ e^-)$
$\gamma \gamma$	neutrino-induced $Z'$ fusion		$Z'^* Z'^* \rightarrow (S \rightarrow e^+ e^-)$

# Dark Photon Model

## Portals



Dark photon model:

1.  $U(1)_{\text{dark}}$  extension
2. dark Higgs singlet
3.  $\nu_S$  and  $\nu_D$

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# Dark Photon Model

## Lagrangian

$$\begin{aligned}
 \mathcal{L} = & \mathcal{L}_{\text{SM}} + \bar{\nu}_D i \not{D}^x \nu_D \\
 & + (D_\mu^x \Phi)^\dagger (D^{x\mu} \Phi) - V(\Phi, H) \\
 & - \frac{1}{4} X_{\mu\nu} X^{\mu\nu} - \frac{\sin \chi}{2} B_{\mu\nu} X^{\mu\nu} \\
 & + \bar{N} i \not{\partial} N - [y_\nu^\alpha (\bar{L}_\alpha \cdot \tilde{H}) N^c + \frac{\mu'}{2} \bar{N} N^c + y_N \bar{N} \nu_D^c \Phi + \text{h.c.}]
 \end{aligned}$$

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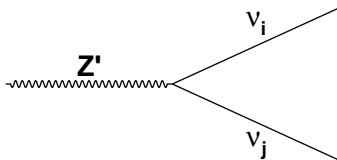
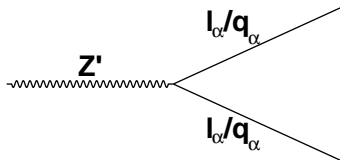
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$$\mathcal{L} \subset \epsilon e Z'_\mu J_{\text{EM}}^\mu + g_D U_{Di} U_{Dj}^* Z'_\mu (i \bar{\nu}_i \not{\partial} \nu_j) \quad (1)$$



# Dark Photon Model

## Neutrino Mixing

$$\mathcal{L}_{\text{mass}}^{\nu} = -\frac{1}{2} \begin{pmatrix} \overline{\nu}_{\alpha} & \overline{N} & \overline{\nu}_D \end{pmatrix} \begin{pmatrix} 0_{3 \times 3} & m_D^T & 0 \\ m_D & \mu' & \Lambda^T \\ 0 & \Lambda & 0 \end{pmatrix} \begin{pmatrix} \nu_{\alpha} \\ N^C \\ \nu_D^C \end{pmatrix} \quad (2)$$

$$\begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \\ N^C \\ \nu_D^C \end{pmatrix} \xrightarrow[\text{mass basis}]{\text{From flavor to}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ N_4 \\ N_5 \end{pmatrix}$$

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# Dark Photon Model

## MiniBooNE detector

- ▶ **Material:** pure mineral oil ( $CH_2$ ) in spherical shape (radius: 6.1 m)
- ▶ **Baseline:** 541 m
- ▶ **Mean** energy: 800 MeV
- ▶ **Beam:**  $\nu_\mu$ ,  $\bar{\nu}_\mu$  (mainly), from BNB
- ▶ **Peaks:** 600 MeV ( $\nu_\mu$ ), 400 MeV ( $\bar{\nu}_\mu$ )
- ▶ **Detection:** Cherenkov, PMTs

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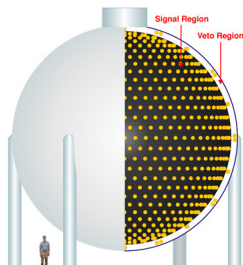
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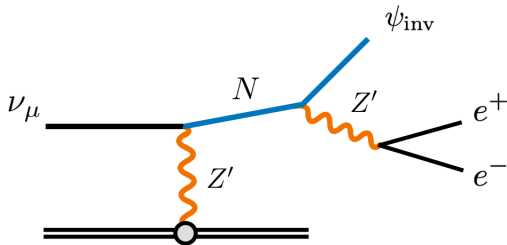
Results

MiniBooNE Detector

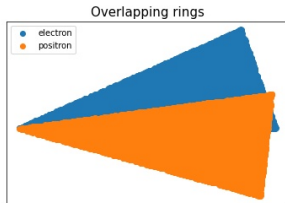
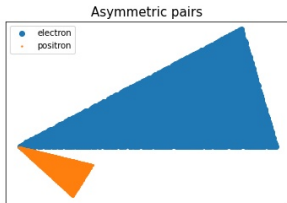


# Dark Photon Model

Misreconstruction at MiniBooNE



Condition:  $< 13^\circ$  or  $E_\nu < 0.03$  GeV



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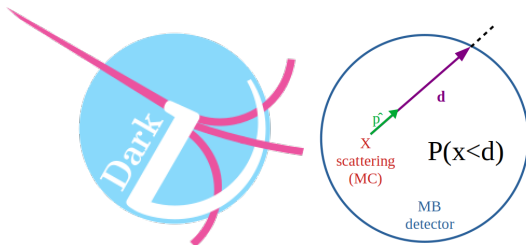
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# Simulation and Fitting

DarkNews



$$-2 \ln L =$$

$$\sum_{i,j} (D_i - P_i) M_{ij}^{-1} (D_j - P_j)$$

$$+ \ln \det M$$

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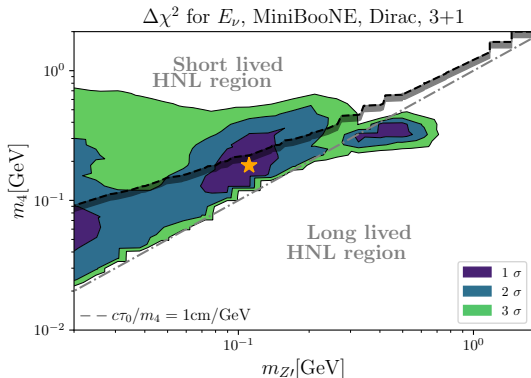
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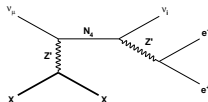
# Results

3+1



**Best Fit:**  $m_{Zl}=111.7\text{MeV}$ ,  $m_4=186\text{MeV}$ ,

$$\chi_{\text{red}}^2=1.255, |U_{\mu 4}^{\text{max}}|=10^{-3}, \epsilon=8 \times 10^{-4}$$



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## 3+1 best fit

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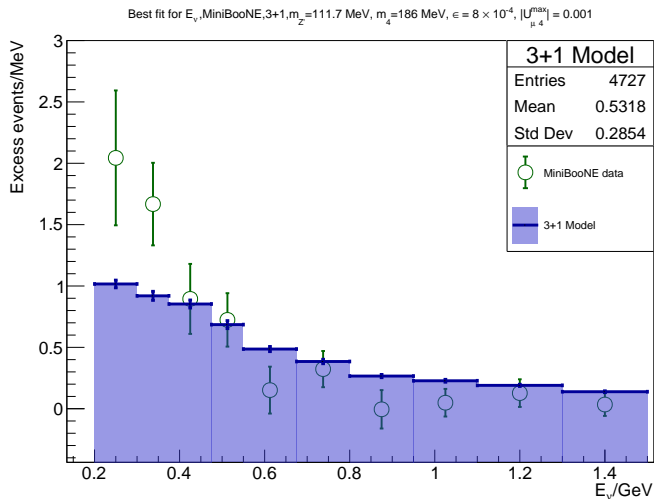
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# Results

## Signal efficiency

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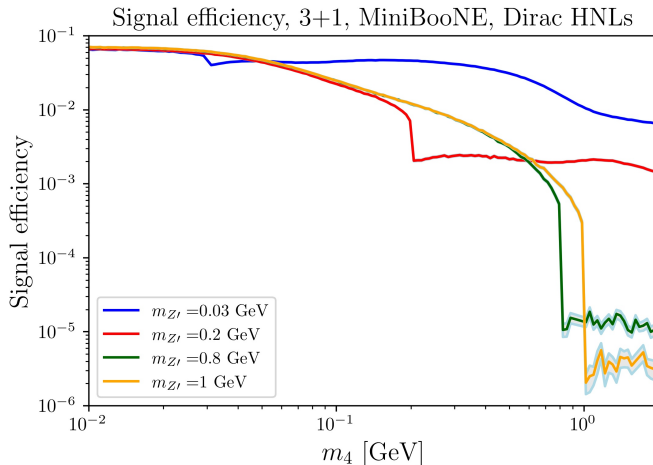
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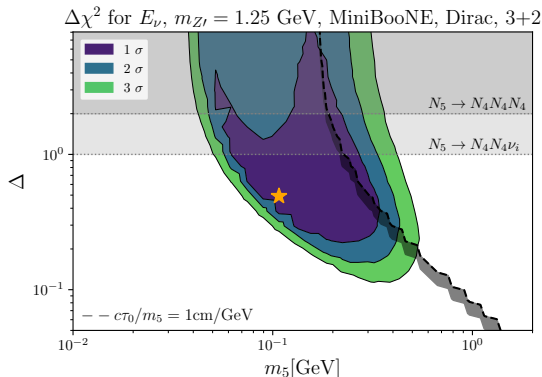
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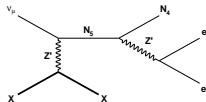
# Results

3+2



**Best Fit:**  $m_{Z'}=1.25\text{GeV}$ ,  $m_5=107.5\text{MeV}$ ,  
 $m_4=72.07\text{MeV}$ ,

$$\chi_{\text{red}}^2=0.91, |U_{\mu 5(4)}^{\text{max}}|=10^{-3}, \epsilon=10^{-2}$$



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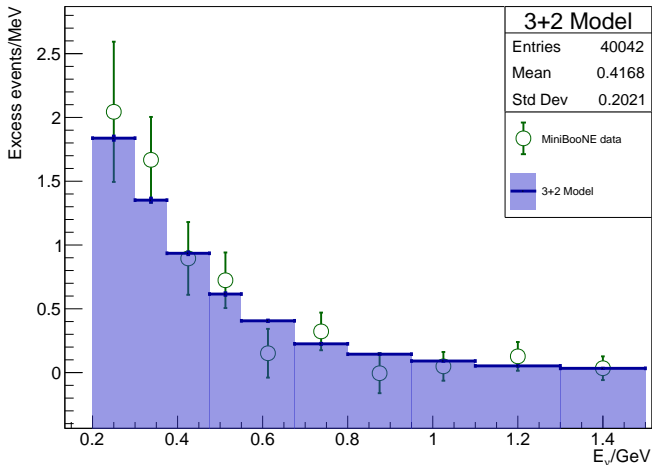
Simulation

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# Results

## 3+2 best fit

Best fit for  $E_\nu$  MiniBooNE, 3+2,  $m_z = 1.25$  GeV,  $m_5 = 107.5$  MeV,  $m_4 = 72.07$  MeV,  $\epsilon = 10^{-2}$ ,  $|U_{\mu 4}^{\text{max}}| = 0.001$



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## 3+2 benchmark point

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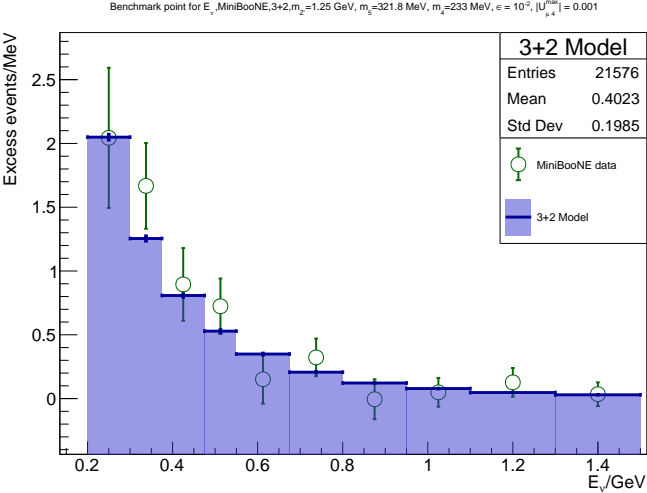
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## Couplings vs mass, 3+2

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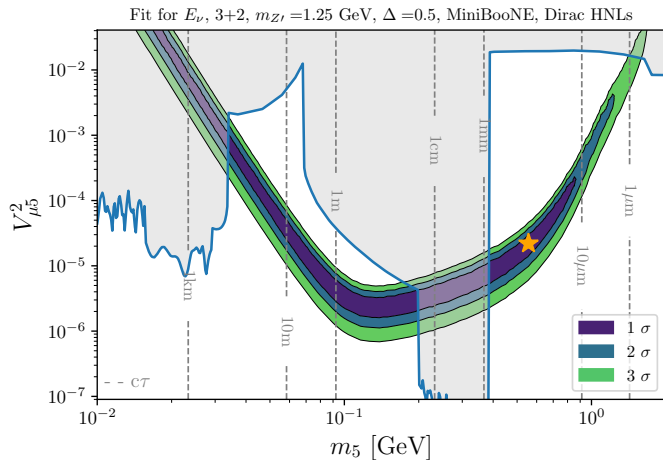
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# Summary

1. The  $3+2$  dark sector model can explain MB LEE in the off-shell decay regime of the HNL
2. It is a simple model with small number of parameters
3. The SBN program will be able to test it
4. It is not so constrained as other models, such as  $3+1$  DS

# THANK YOU!

# Limits for $U_{\mu N}$

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