

Latest results from CUPID-0

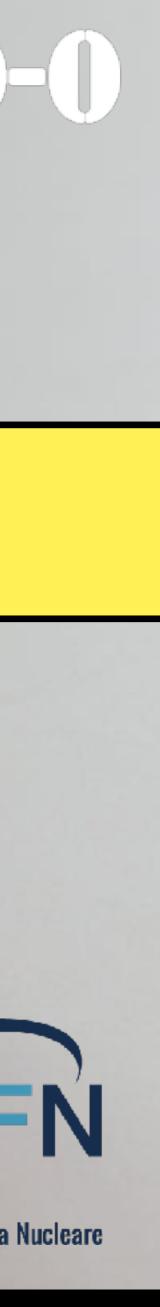




Guido Fantini on behalf of the CUPID-0 Collaboration







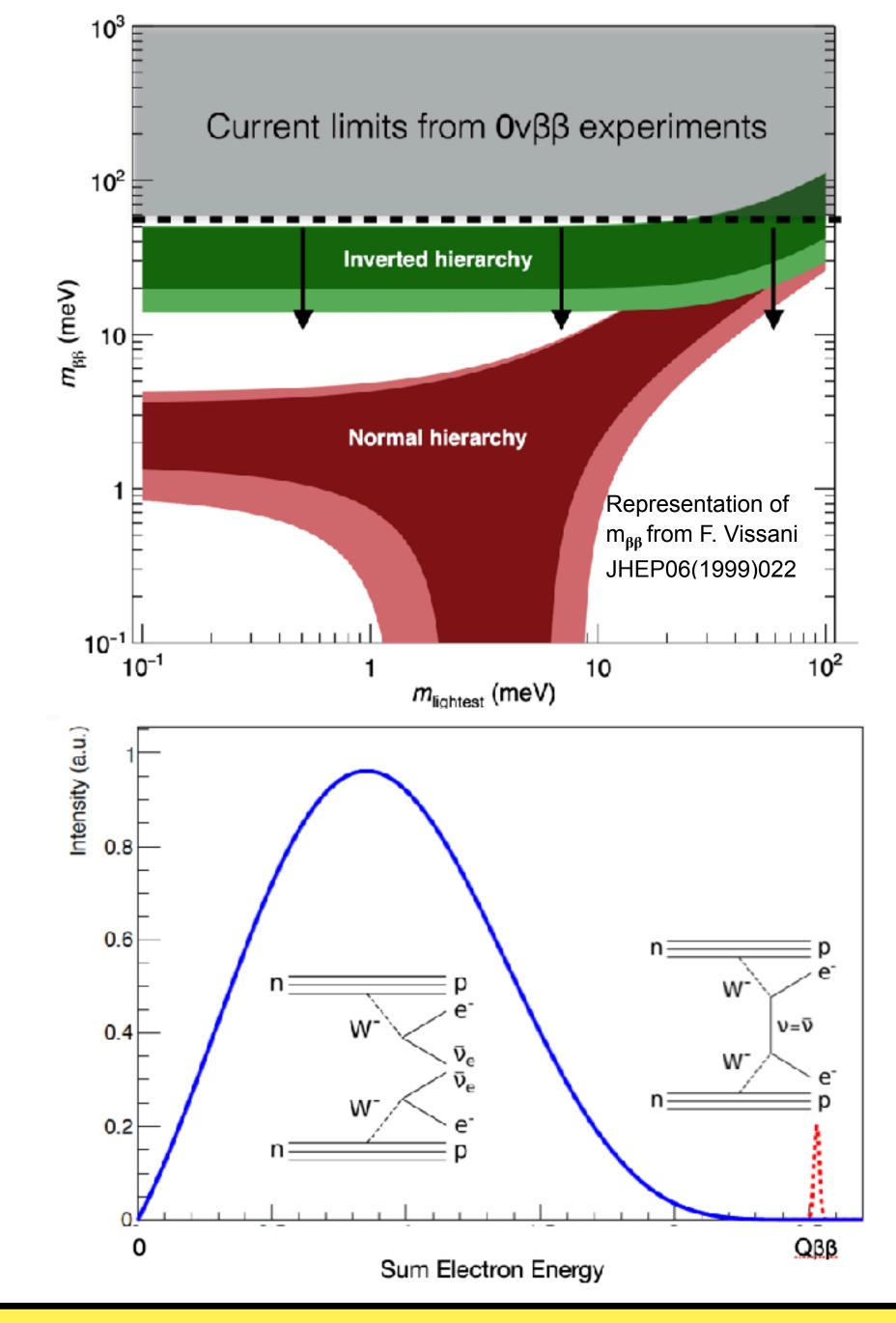
Physics case

- lepton number violation ($\Delta L = +2$)
- neutrino mass hierarchy, absolute scale, Majorana mass component
- matter anti-matter asymmetry in the Universe (matter creation in the lab)
- peak at 2vßß spectral endpoint

$$\Gamma_{0\nu} = G_{0\nu} |M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

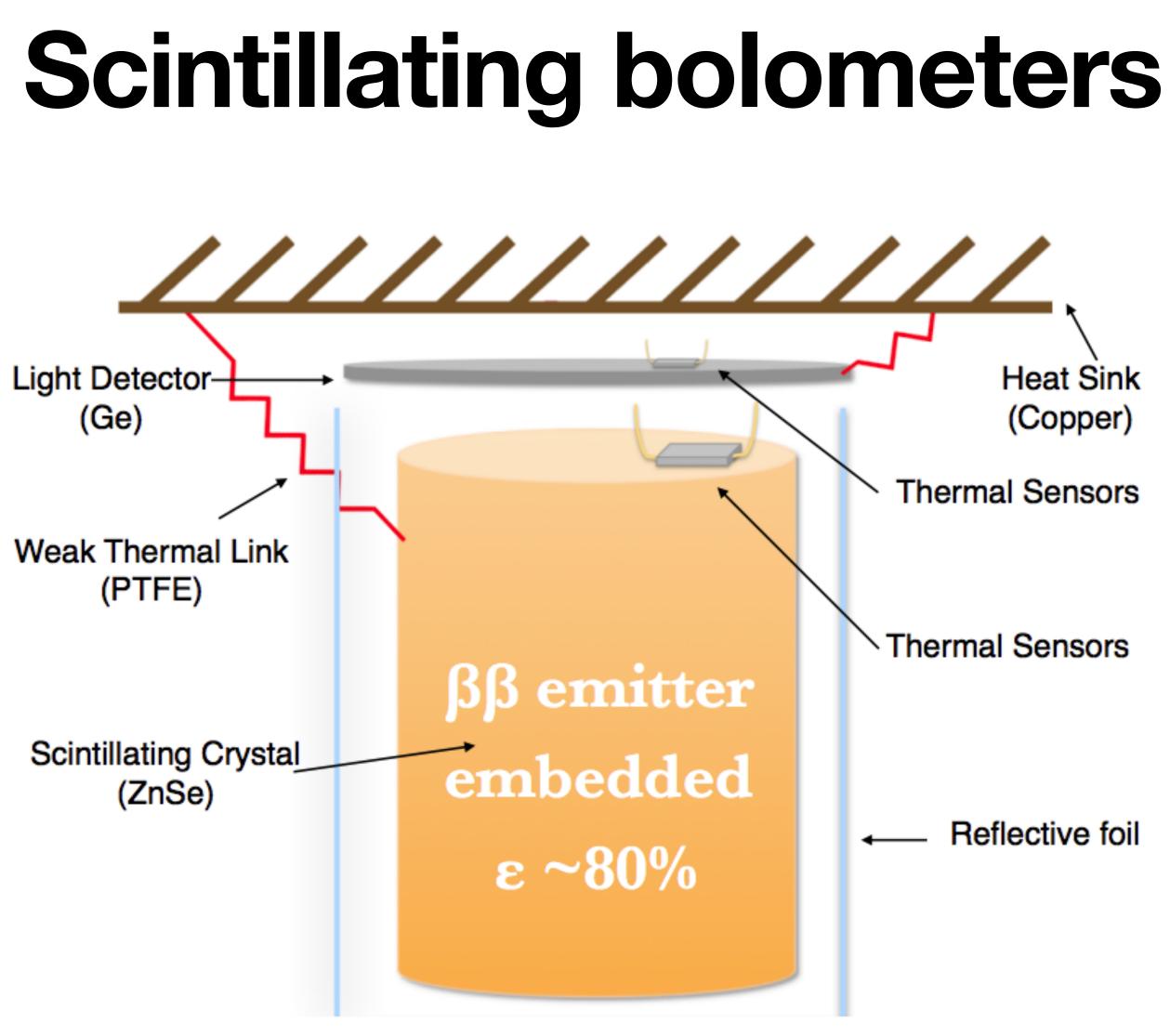






G. Fantini on behalf of the CUPID-0 Collaboration





Eur. Phys. J. C (2018) 78:428 10.1140/epjc/s10052-018-5896-8

Latest results from CUPID-0

- \
leat Sink
(Copper)

Reflective foil

- T~10 mK temperature makes heat capacity very low
- ββ decays heat the crystal
- temperature pulse readout via NTD thermistor as voltage pulse
- scintillation light heats Ge light detector (same readout)
- weak thermal link to heat bath brings detector back to base T

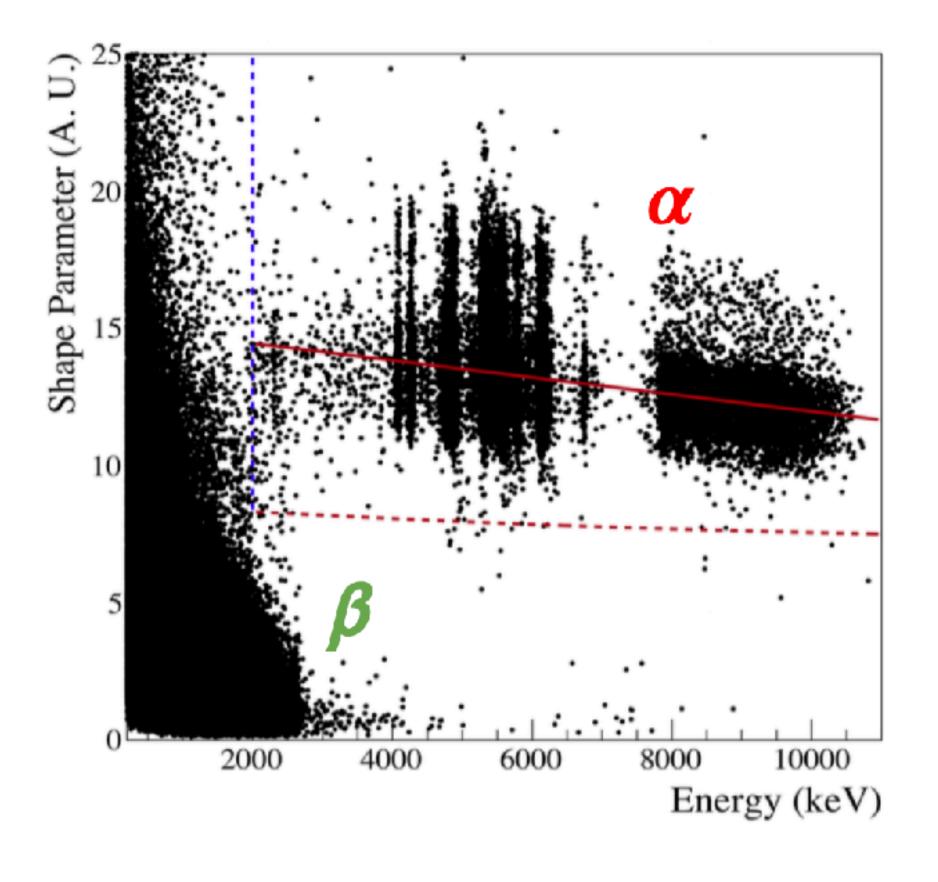








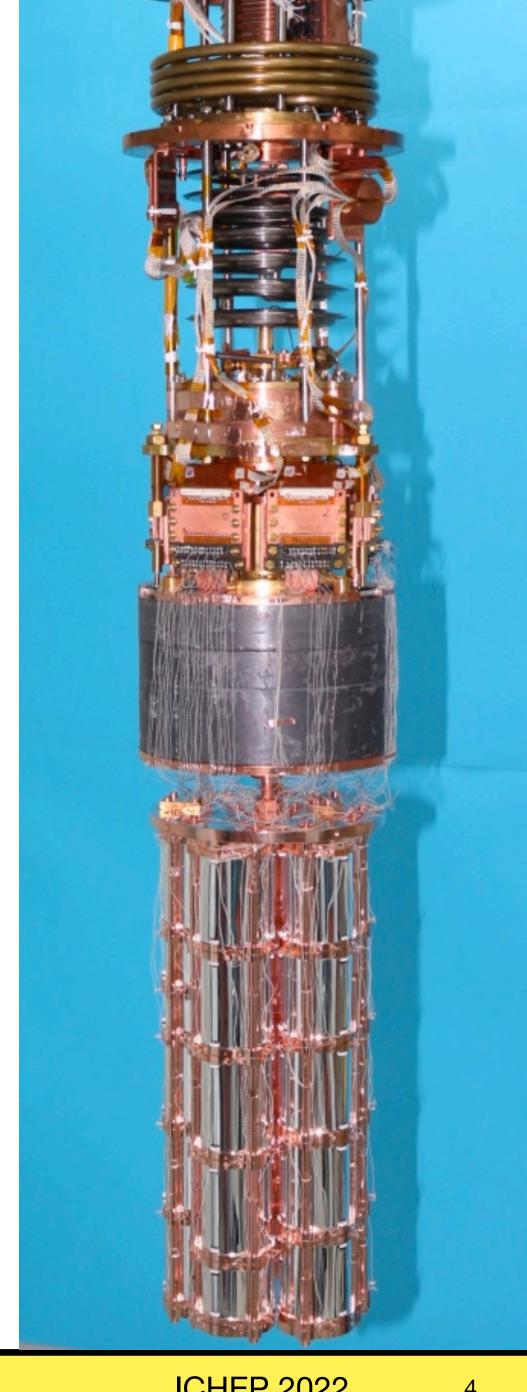
The CUPID-0 experiment



- 26 ZnSe cylindric scintillating bolometer crystals, 2 natural + 24 95% enriched in ⁸⁶Se
- surrounded by VIKUITI (3M) reflective foil
- installed underground in the INFN LNGS lab to shield from cosmic radiation
- achieve full $\alpha/\beta\gamma$ discrimination with dual heat/light readout
- operation phase I (3/2017-12/2018) with reflective foil
- operation phase II (6/2019-2/2020) w/o reflective foil



Eur. Phys. J. C (2018) 78:428 0.1140/epjc/s10052-018-5896-8

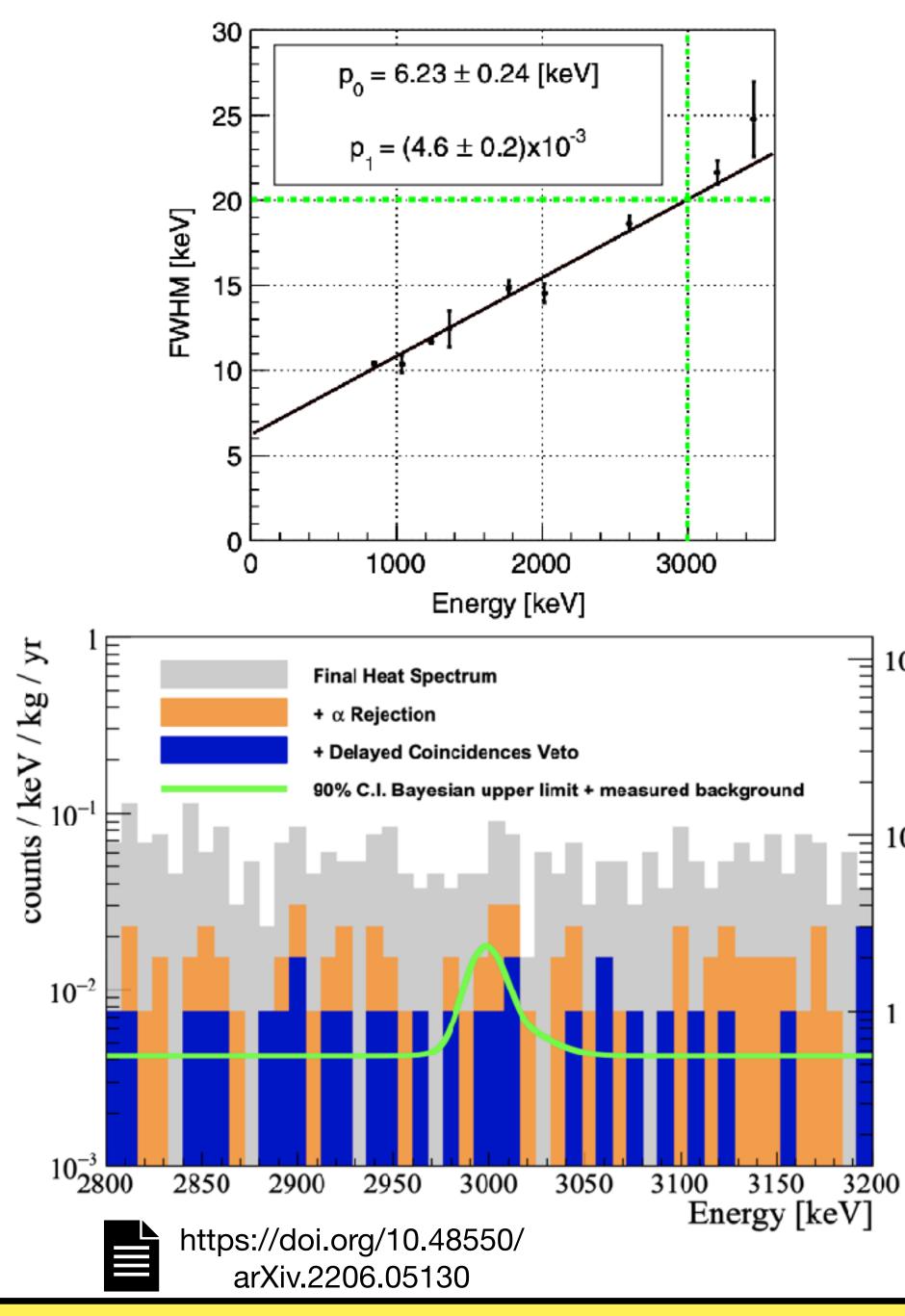


Ovββ results

- 8.82 kg.yr ⁸²Se exposure (phase I + II)
- energy resolution (20.05 ± 0.34) keV at ~3 MeV
- background index $(3.5 \pm 1.0) \ 10^{-3} \ cts/(keV kg yr) \ phase$ $(5.5 \pm 1.5) \ 10^{-3} \ cts/(keV kg yr) \ phase II$
- no evidence found, Bayesian 90% C.I. set $T_{1/2}^{0v} > 4.6 \times 10^{23} \text{ yr}$ $m_{\beta\beta} < (263 - 545) \text{ meV}$

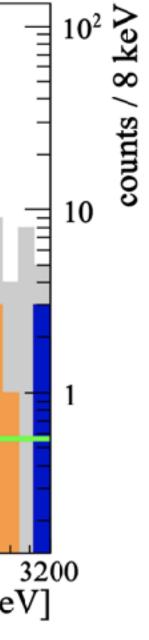


Phys. Rev. Lett 123, 032501 (2019) 10.1103/PhysRevLett.123.032501





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Background Model

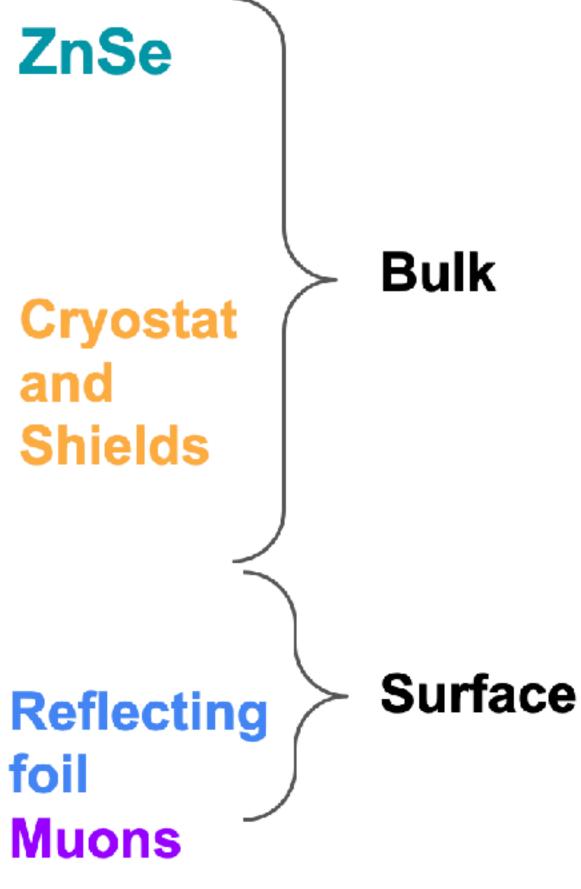
Component	Mass (kg)	Source	Index	Activity (Bq/kg)
Crystals	10.5	$2\nu\beta\beta$	1	$(9.96 \pm 0.03) \times 10^{-4}$
		⁶⁵ Zn	2	$(3.52 \pm 0.06) \times 10^{-4}$
		⁴⁰ K	3	$(8.5 \pm 0.4) \times 10^{-5}$
		⁶⁰ Co	4	$(1.4 \pm 0.3) \times 10^{-5}$
		¹⁴⁷ Sm	5	$(1.6 \pm 0.3) \times 10^{-7}$
		$^{238}U^{-226}Ra$	6	$(5.51 \pm 0.10) \times 10^{-6}$
		²²⁶ Ra- ²¹⁰ Pb	7	$(1.54 \pm 0.02) \times 10^{-5}$
		²¹⁰ Pb- ²⁰⁶ Pb	8	$(7.05 \pm 0.16) \times 10^{-6}$
		$^{232}\text{Th}-^{228}\text{Ra}$	9	$(2.74 \pm 0.10) \times 10^{-6}$
		$^{228}Ra^{-208}Pb$	10	$(1.20 \pm 0.03) \times 10^{-5}$
		$^{235}U^{-231}Pa$	11	$(5.3 \pm 0.7) \times 10^{-7}$
		²³¹ Pa- ²⁰⁷ Pb	12	$(7.8 \pm 0.4) \times 10^{-7}$
Holder	3.10	⁵⁴ Mn	13	$(2.2 \pm 0.3) \times 10^{-4}$
CryoInt (^a)	36.9	²³² Th	14	$< 4.5 \times 10^{-5}$
		^{238}U	15	$(7 \pm 3) \times 10^{-5}$
		⁴⁰ K	16	$(3.0 \pm 0.6) \times 10^{-3}$
		⁶⁰ Co	17	$(6.8 \pm 1.3) \times 10^{-5}$
IntPb	202	²³² Th	18	$< 6.3 \times 10^{-5}$
		$^{238}\mathrm{U}$	19	$<7.3\times10^{-5}$
CryoExt	832	⁶⁰ Co	20	$(2.6 \pm 0.9) \times 10^{-5}$
ExtPb (^b)	24694	²³² Th	21	$(4.3 \pm 0.6) \times 10^{-4}$
		²³⁸ U	22	$(2.5 \pm 1.2) \times 10^{-4}$
		⁴⁰ K	23	$(2.8 \pm 0.8) \times 10^{-3}$
		²¹⁰ Pb	24	7.8 ± 0.3
Component	Surface (cm^2)	Source	Index	Activity (Bq/cm^2)
Crystals	2574	226 Ra- 210 Pb-0.01 μ m	25	$(2.63 \pm 0.15) \times 10^{-8}$
		$^{228}Ra^{-208}Pb^{-0.01}\mu m$	26	$(6.5 \pm 1.1) \times 10^{-9}$
		$^{226}Ra^{-210}Pb^{-10}\mu m$	27	$< 2.3 \times 10^{-9}$
		$^{228}Ra^{-208}Pb^{-10}\mu m$	28	$(4.2 \pm 1.6) \times 10^{-9}$
Reflectors (^c)	2100	²³² Th–10µm	29	$< 7.3 \times 10^{-10}$
		$^{226}Ra^{-210}Pb^{-10}\mu m$	30	$(8.7 \pm 1.3) \times 10^{-9}$
		²¹⁰ Pb- ²⁰⁶ Pb-10µm	31	$(1.0 \pm 0.5) \times 10^{-8}$
		$^{210}\text{Pb}-^{206}\text{Pb}-0.01\mu\text{m}$	32	$(1.43 \pm 0.02) \times 10^{-7}$
Muons	Flux in units of $\mu/(\text{cm}^2\text{s})$		33	$(3.7 \pm 0.2) \times 10^{-8}$



foil

Latest results from CUPID-0





Aim at reproducing the energy spectrum with a weighted sum of different sources

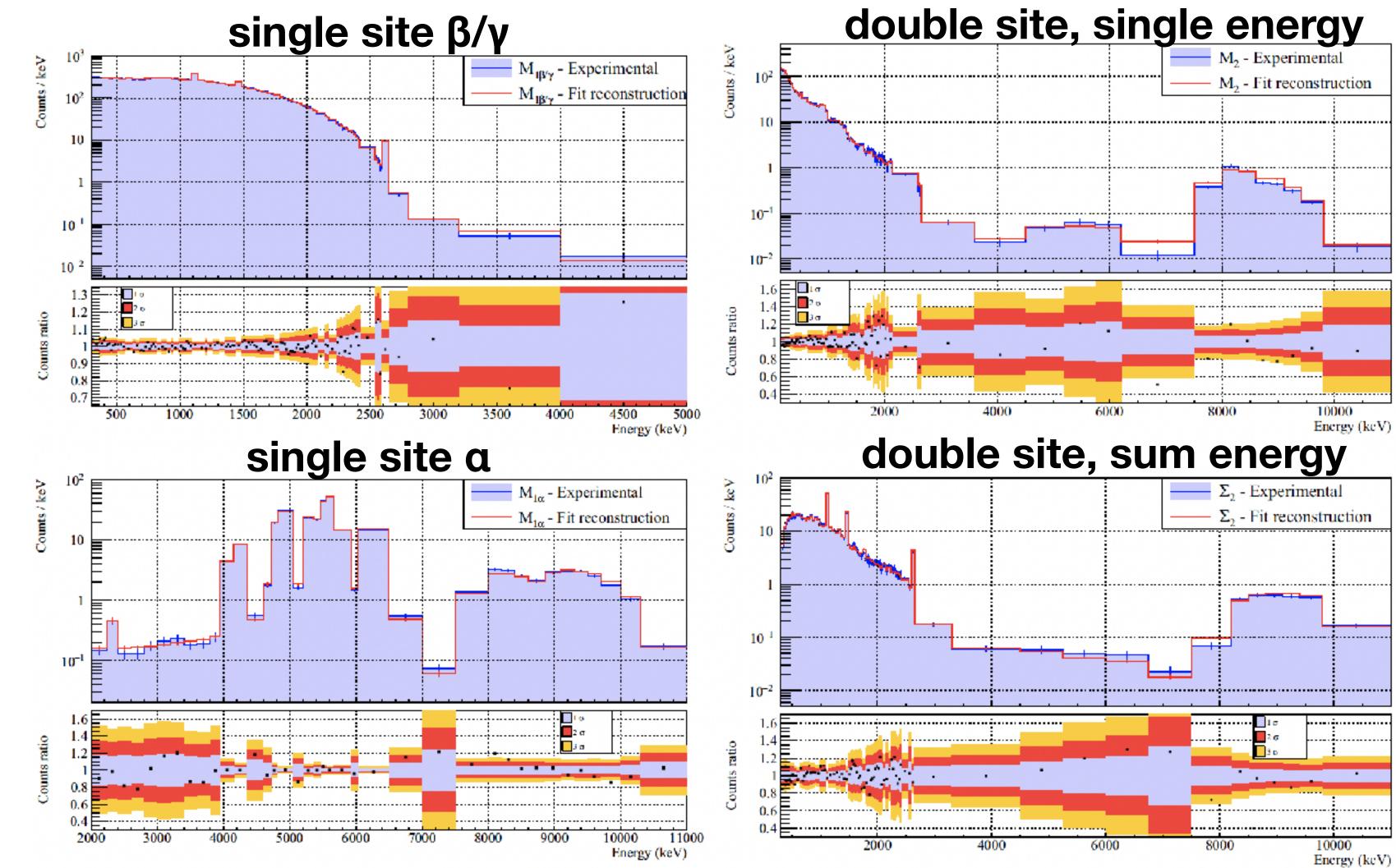
Source selection based on data (prominent lines)

Each source is a MC simulation, the activity is fit to the data





Background Model



Eur. Phys. J. C (2019) 79:583 10.1140/epjc/s10052-019-7078-8

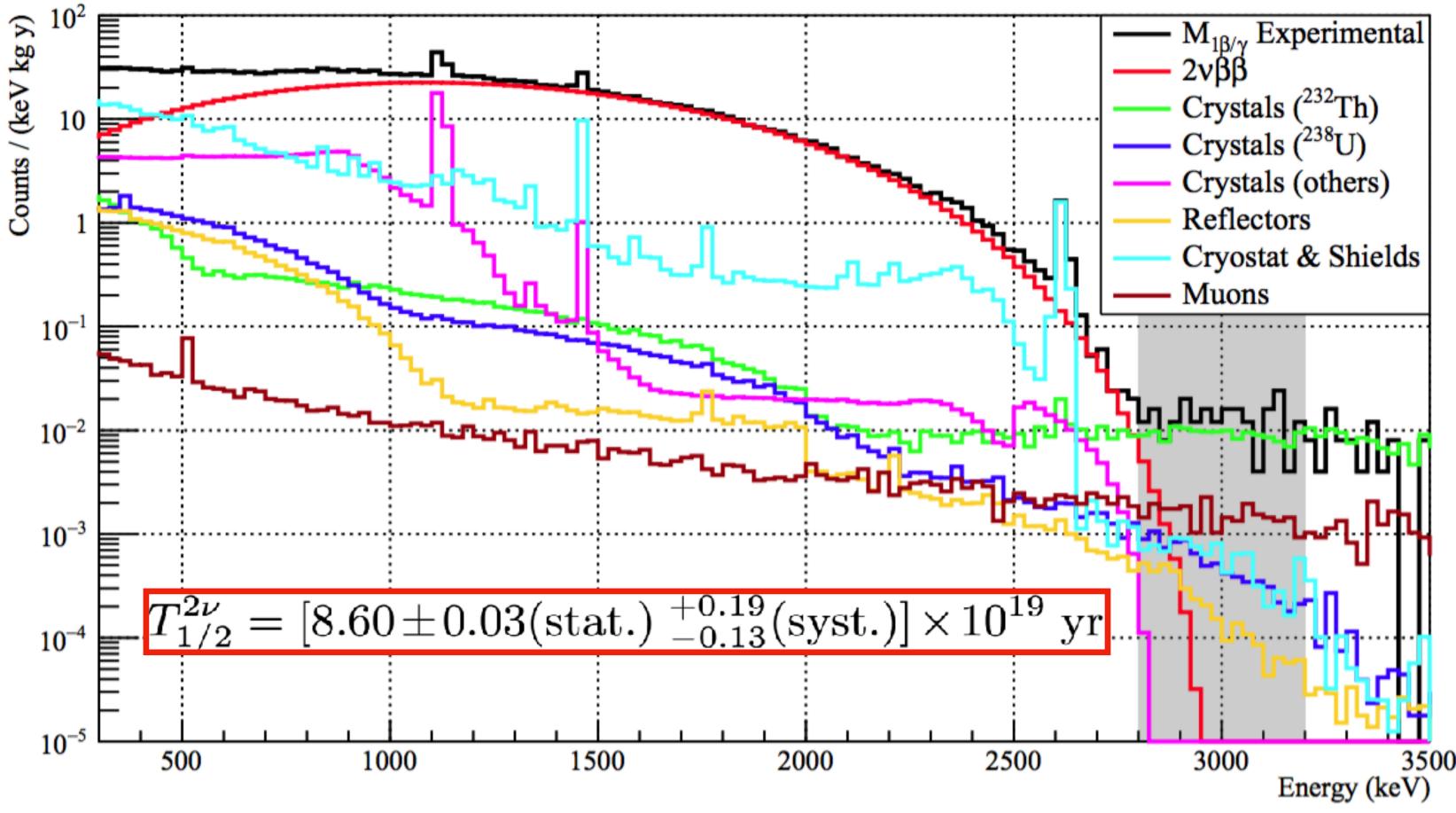


Latest results from CUPID-0

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2vββ results



(the time veto for ²⁰⁸Tl events is not included in the plot)

The dominant contribution < 2.5 MeV is 2vββ of ⁸²Se



Eur. Phys. J. C (2019) 79:583 10.1140/epjc/s10052-019-7078-8



Phys. Rev. Lett 123, 262501 (2019) 10.1103/PhysRevLett.123.262501

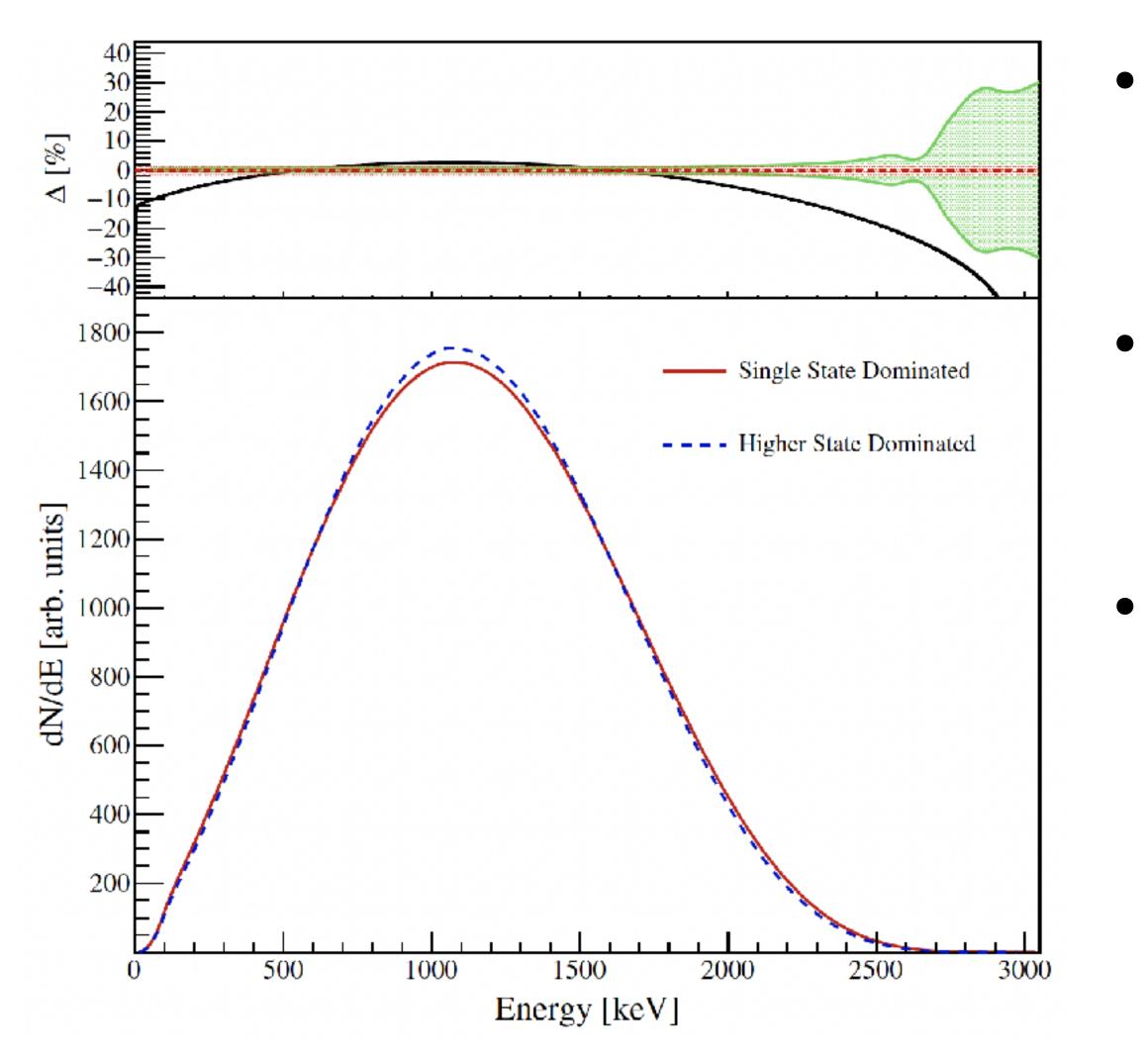
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2vββ results - spectral shape sensitivity



Latest results from CUPID-0

- High statistics and signal purity can be used to disentangle tiny spectral shape differences in 2vßß events
- CUPID-0 provided evidence for Single State Dominance of 2vββ transition in ⁸²Se over Higher State Dominance
 - Many exotic processes have their signature as a distortion of the $2\nu\beta\beta$ energy spectrum



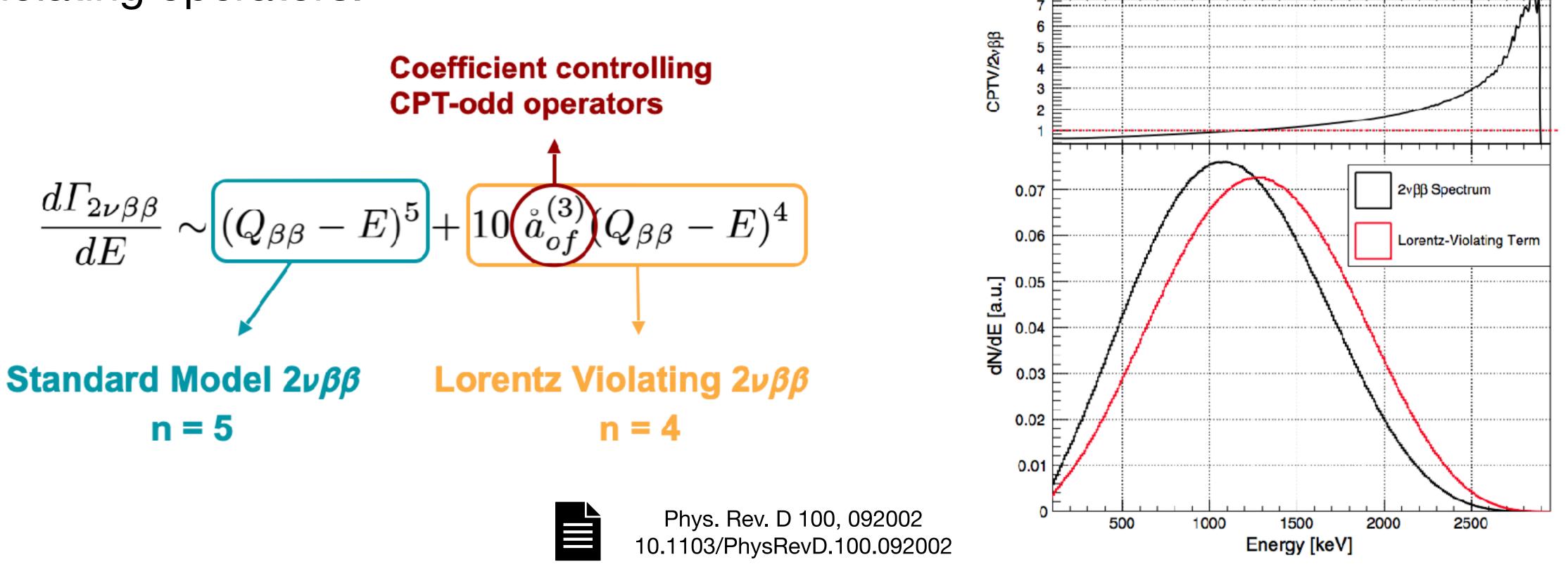
Phys. Rev. Lett 123, 262501 (2019) 10.1103/PhysRevLett.123.262501





Lorentz violation

violating operators.





The introduction of a consistent quantum gravity theory predicts new physics at the Planck scale, that has a low energy counterpart in the SME as Lorentz-

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Lorentz violation

Added a new component to the background model

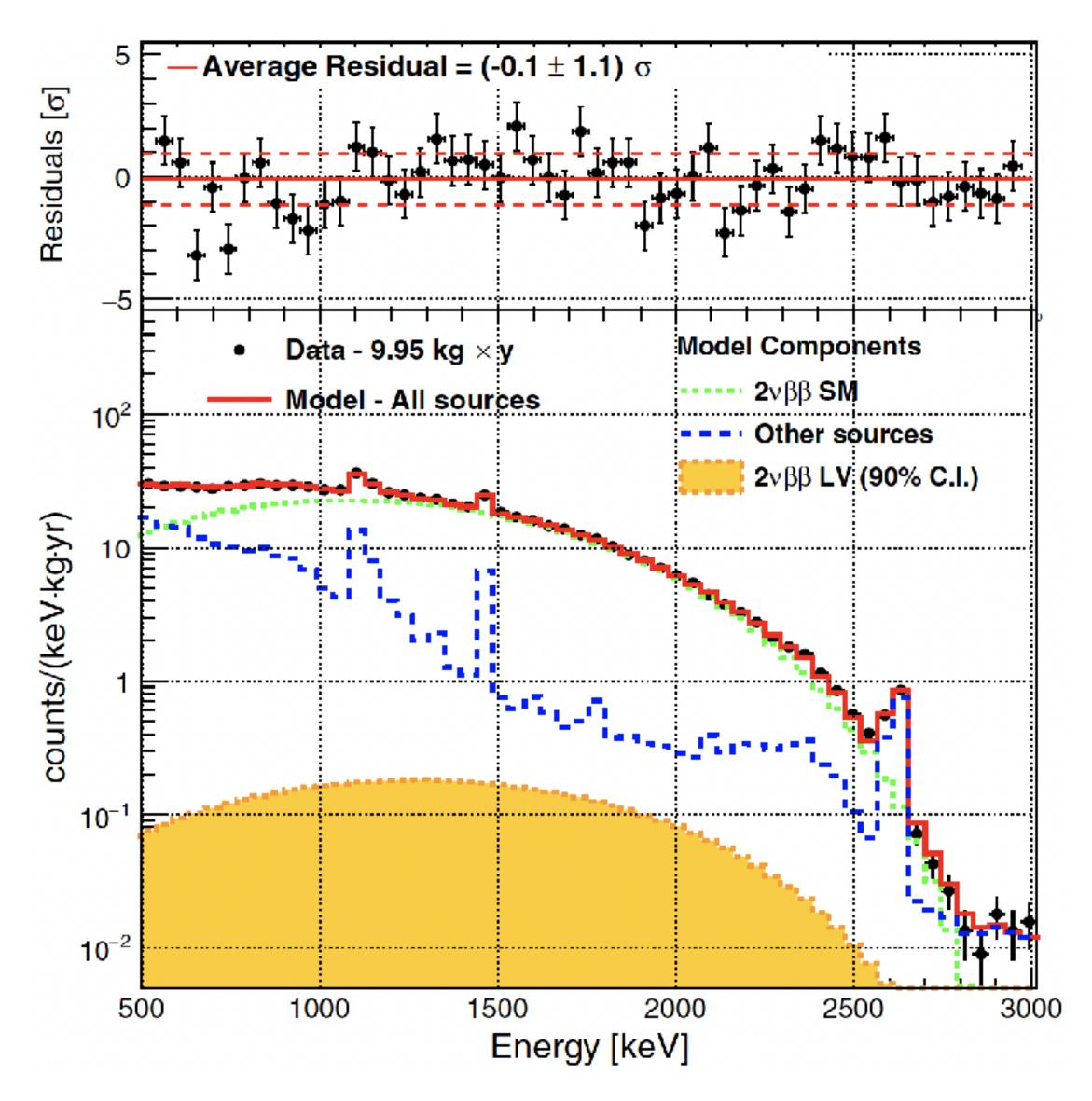
Systematic effects from energy scale, binning, source choice, threshold accounted for

No evidence for Lorentz violation, can set Bayesian 90% C.I. limit

$$a_{of}^{o(3)} < 4.1 \times 10^{-6} \text{ GeV}$$



Phys. Rev. D 100, 092002 10.1103/PhysRevD.100.092002



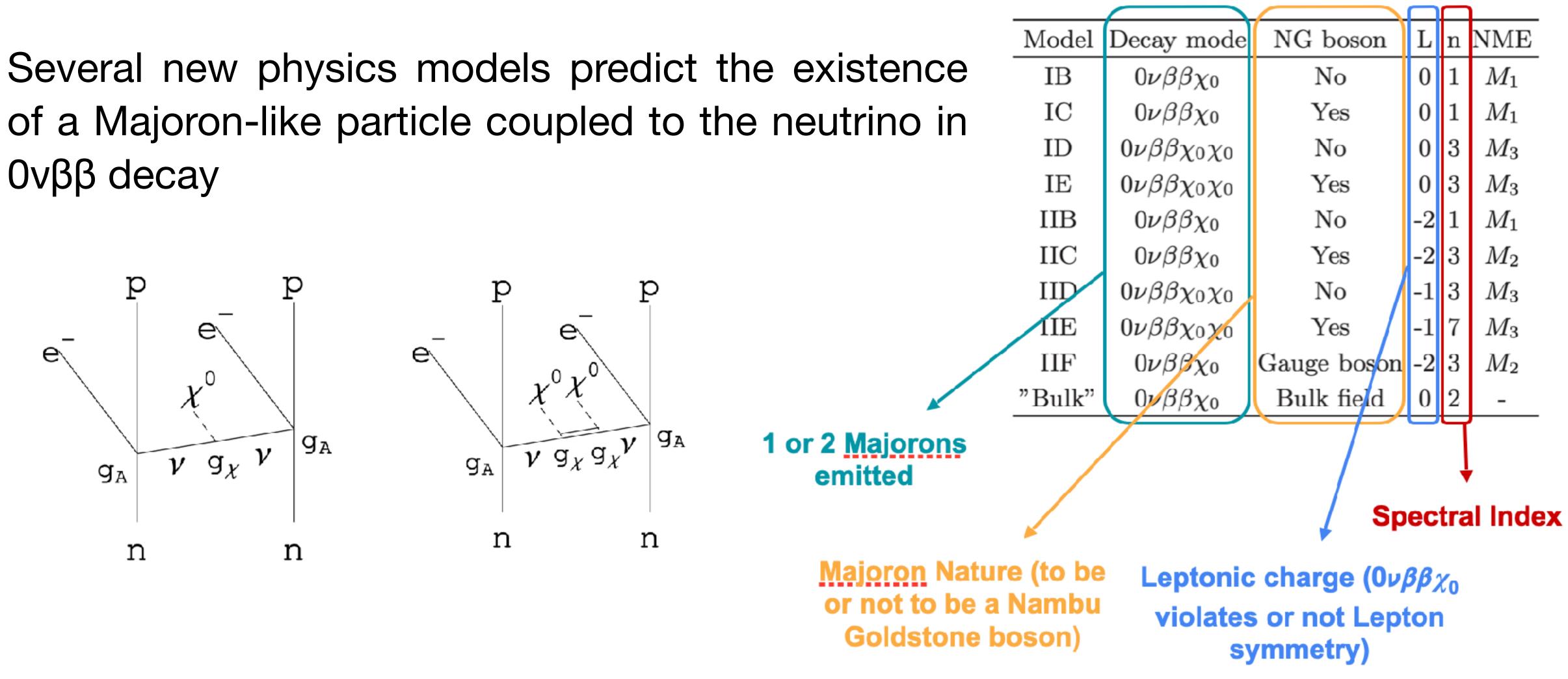
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ICHEP 2022

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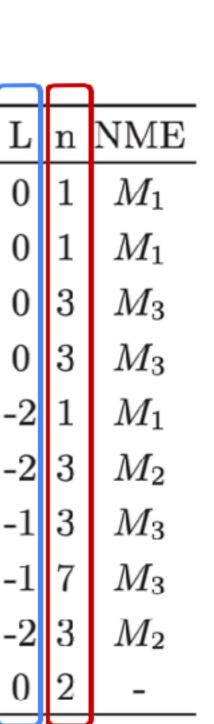
Ovββ with Majoron emission

 $0\nu\beta\beta$ decay





Phys. Rev. C 103, 044302 10.1103/PhysRevC.103.044302





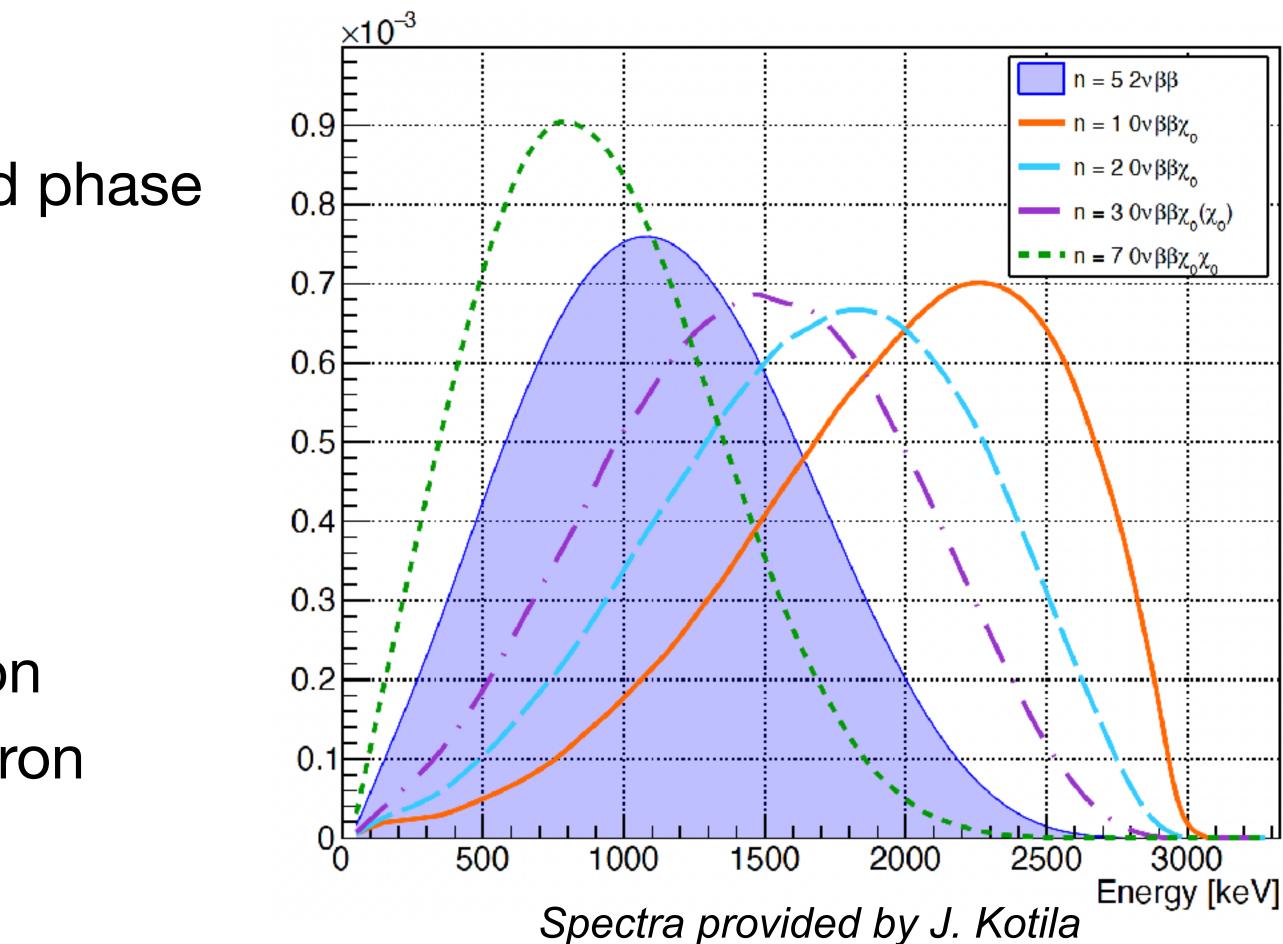


Ovββ with Majoron emission

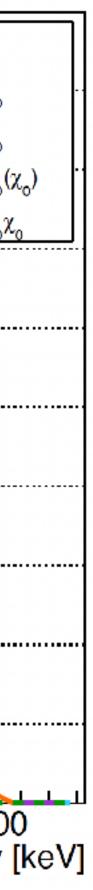
The experimental signature of different Majoron emitting channels is a modified phase space as

$$\frac{d\Gamma_{0\nu\beta\beta\chi_0}}{dE} \sim (Q_{\beta\beta} - E)^n$$

We investigated single Majoron emission modes $0\nu\beta\beta\chi$ n=1,2,3 and double Majoron emission modes 0vββχχ n=3,7



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Οvββ with Majoron emission

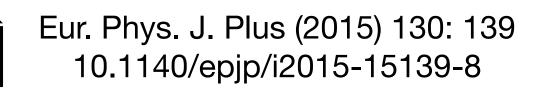
- No evidence for Majoron emitting modes
- Upper limits at 90% Bayesian C.I.
- Interpretation as upper limits on Majoronneutrino coupling constant

$$[t_{1/2}]^{-1} = |\langle g_{\chi_0} \rangle|^{2m} G^{(0)}_{(m,n)} |M_{(m,n)}|^{2m} G^{(0)}_{(m,n)}|^{2m} G^{(0)}_{(m,n)}|^$$

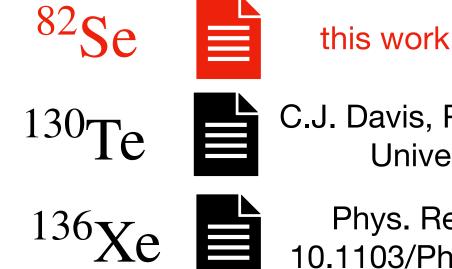
where m = 1,2 number of final-state Majorons ⁸²Se



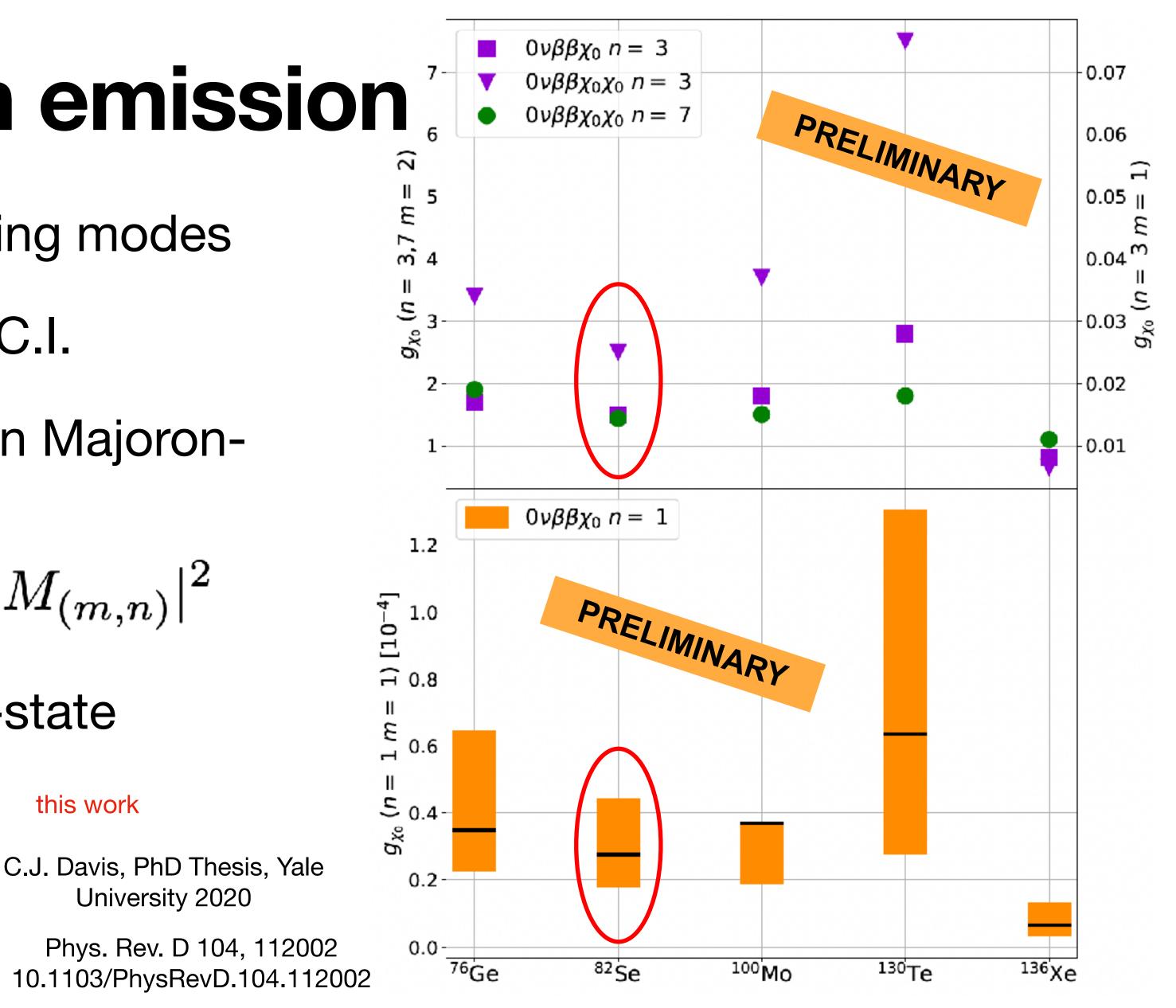
¹⁰⁰Mo



Eur. Phys. J. C (2019) 79:440 10.1140/epjc/s10052-019-6948-4



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CUPID baseline design

45 x 45 x 45 mm³ $Li_2^{100}MoO_4$ crystals

Crystal mass: 280 g

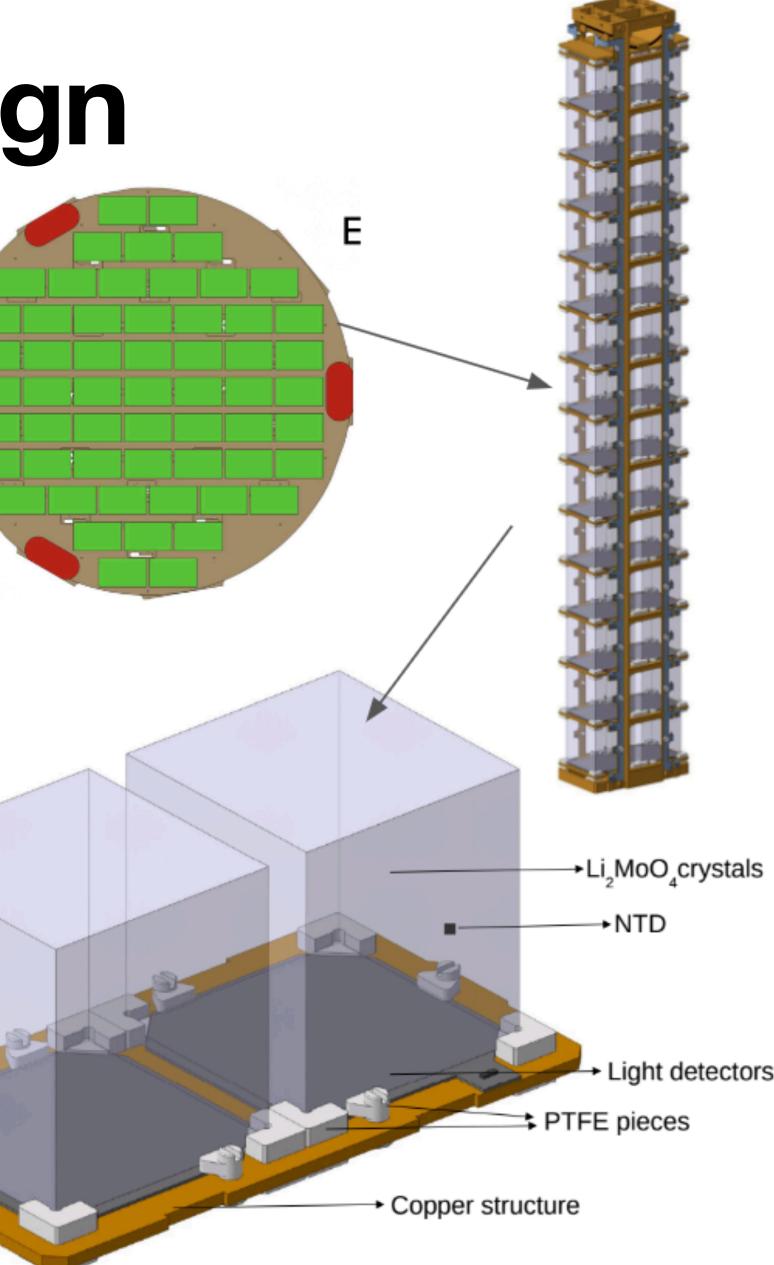
1596 total crystals

- 450 kg of Li₂¹⁰⁰MoO 95% enrichment in ¹⁰⁰Mo: 240 kg of ¹⁰⁰Mo •
- 57 towers of 28 crystals. 14-floors of 2x1 • crystal pairs. Gravity-assisted design

Ge light detectors with SiO anti-reflective coating

- Each crystal has top and bottom light detectors
- No reflective foils •

Muon veto for muon-induced background suppression





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Summary

- CUPID-0 successfully concluded operation in 2020
- The full CUPID-0 statistics can constrain 0vββ to
- Just phase-I data can produce multiple results on the $\beta\beta$ continuous energy spectrum:

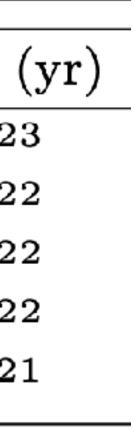
 $T_{1/2}^{2\nu} = [8.60 \pm 0.03 (\text{stat.})]_{-0.13}^{+0.19} (\text{syst})$

 CUPID will use dual readout technol Li₂¹⁰⁰MoO₄ to fully explore the inverte of neutrino masses

$$T_{1/2}^{0v} > 4.6 \times 10^{23} \text{ yr}$$

$$a_{of}^{o(3)} < 4.1 \times 10^{-6} \text{ GeV}$$





Backup

Backup – References for Majoron emission

- n = 1 (black line), 3 and 7: <u>https://doi.org/10.1103/physrevc.103.044302</u>
- Orange Bands for n = 1:
 - <u>https://doi.org/10.1016/j.nuclphysa.2008.12.005</u>
 - <u>https://doi.org/10.1088/1361-6471/aa9bd4</u>
 - <u>https://doi.org/10.1103/PhysRevC.93.024308</u>
 - <u>https://doi.org/10.1103/physrevc.101.044315</u>
 - o https://doi.org/10.1103/PhysRevC.91.034304
 - <u>https://doi.org/10.1103/PhysRevD.102.095016</u>
 - o <u>https://doi.org/10.1103/physrevc.103.044302</u>
 - o <u>https://doi.org/10.1103/PhysRevC.87.064302</u>
 - <u>https://doi.org/10.1103/PhysRevC.87.045501</u>
 - o https://doi.org/10.1103/PhysRevC.98.064325
 - <u>https://doi.org/10.1103/PhysRevC.91.024613</u>
 - <u>https://doi.org/10.1103/PhysRevC.97.045503</u>
 - <u>https://doi.org/10.1103/physrevc.102.044303</u>
 - <u>https://doi.org/10.1103/PhysRevLett.105.252503</u>
 - <u>https://doi.org/10.1103/PhysRevLett.111.142501</u>
 - <u>https://doi.org/10.1103/PhysRevC.95.024305</u>
 - <u>https://doi.org/10.1103/PhysRevC.88.064322</u>
 - o <u>https://doi.org/10.1103/PhysRevC.91.024316</u>
- Phase Spaces: https://doi.org/10.1103/physrevc.91.064310

