

LIME: a gaseous TPC with optical readout



Stefano Piacentini - Sapienza Università di Roma & INFN Roma 1
on behalf of:

F.. D. Amaro, E. Baracchini, L. Benussi, S. Bianco, C. Capoccia, M. Caponero, D. S. Cardoso, G. Cavoto, A. Cortez, R. J. de Cruz Roque, I. A. Costa, E. Dané, E. Di Marco, G. Grilli di Cortona, G. D'Imperio, G. Dho, F. Di Giambattista, R. R. M. Gregorio, F. Iacoangeli, H. P. Lima Júnior, G. Maccarrone, R. D. P. Mano, M. Marafini, G. Mazzitelli, A. G. Mc Lean, A. Messina, M. L. Migliorini, C.M.B. Monteiro, R. A. Nóbrega, A. Orlandi, I. F. Pains, E. Paoletti, L. Passamonti, F. Petrucci, S. Pelosi, S. Piacentini, D. Piccolo, D. Pierluigi, D. Pinci, A. Prajapati, F. Renga, F. Rosatelli, A. Russo, J.M.F. dos Santos, G. Saviano, A. da Silva Lopes Júnior, N. Spooner, R. Tesauro, S. Tomassini, S. Torelli

Introduction

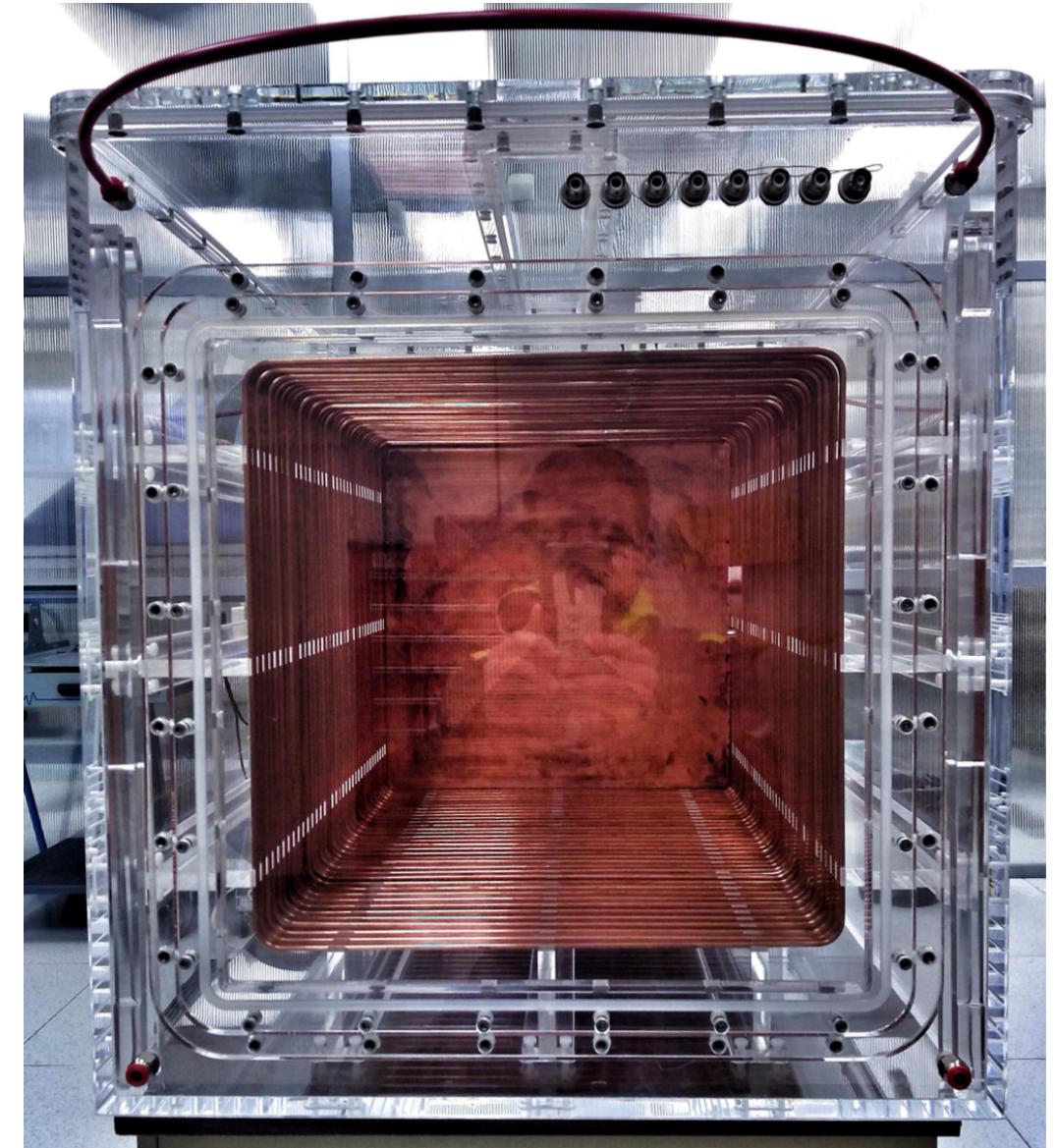


- In this contribution we present **LIME**, the latest 50 L prototype of the **CXGNO** project.
- The **CXGNO** experimental approach is based on a **gaseous** Time Projection Chamber (**TPC**) **with optical readout**.
- The **goal** of the collaboration is to contribute to the **direct search** for **dark matter** candidates with a detector with **directional capabilities**
- **LIME** has been **recently** installed **underground** at the Laboratori Nazionali del Gran Sasso (**LNGS**):
 - to assess the **performance** of the **CXGNO** experimental approach in a low background environment
 - to validate the **background model**, measure the **neutron flux** and determine the best **shielding configuration**

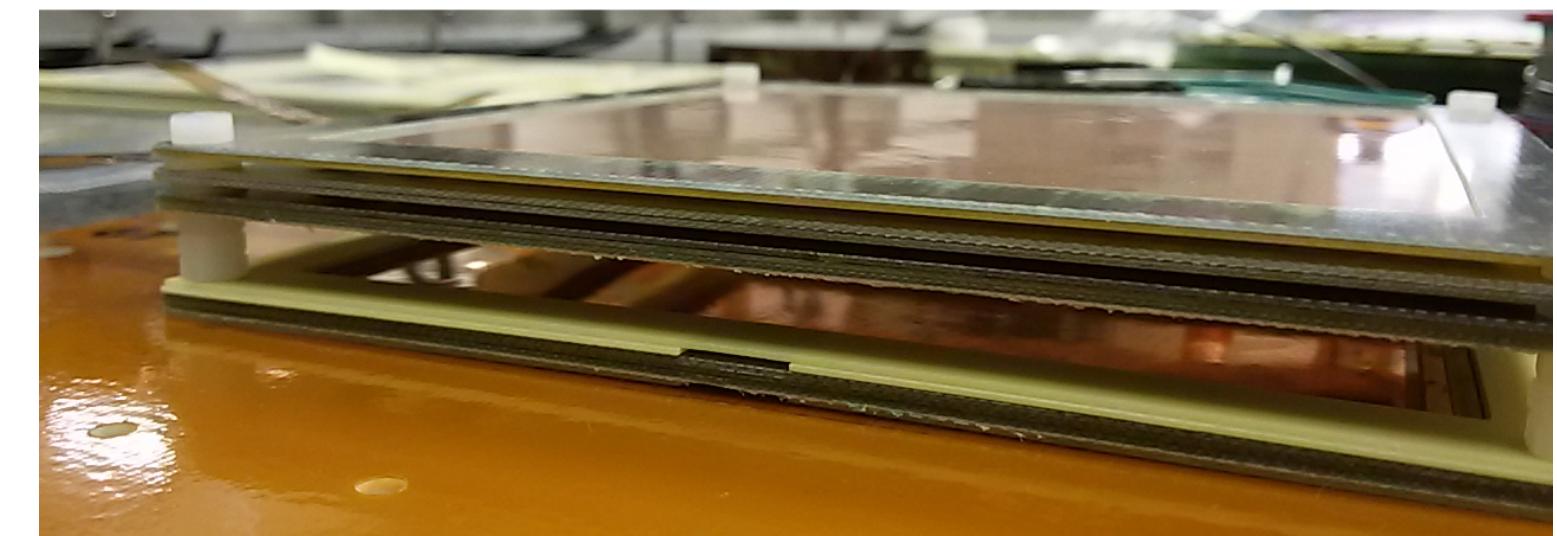
The CXENO project



- **Aiming at** a large detector for high precision **3D tracking of rare low energy nuclear recoils (keV)** as for example WIMPs.
- **Experimental challenges:** rate $O(\text{evt}/\text{kg}/\text{day})$, background rejection, and energy threshold (keV)
- **Strategy: photograph nuclear recoils** in a (1 atm) He:CF₄ TPC with a GEM amplification stage
 - low energy events in 1 atm gas \Rightarrow visible tracks
 - Access to position, direction, total released energy, dE/dx (head/tail), PID
 - **optical sensors:** high granularity, very low noise, and high sensitivity
 - **optical coupling:** sensors outside the sensitive volume, acquire large surfaces with small sensors



TPC of the LIME prototype @ LNF



Amplification with GEMS

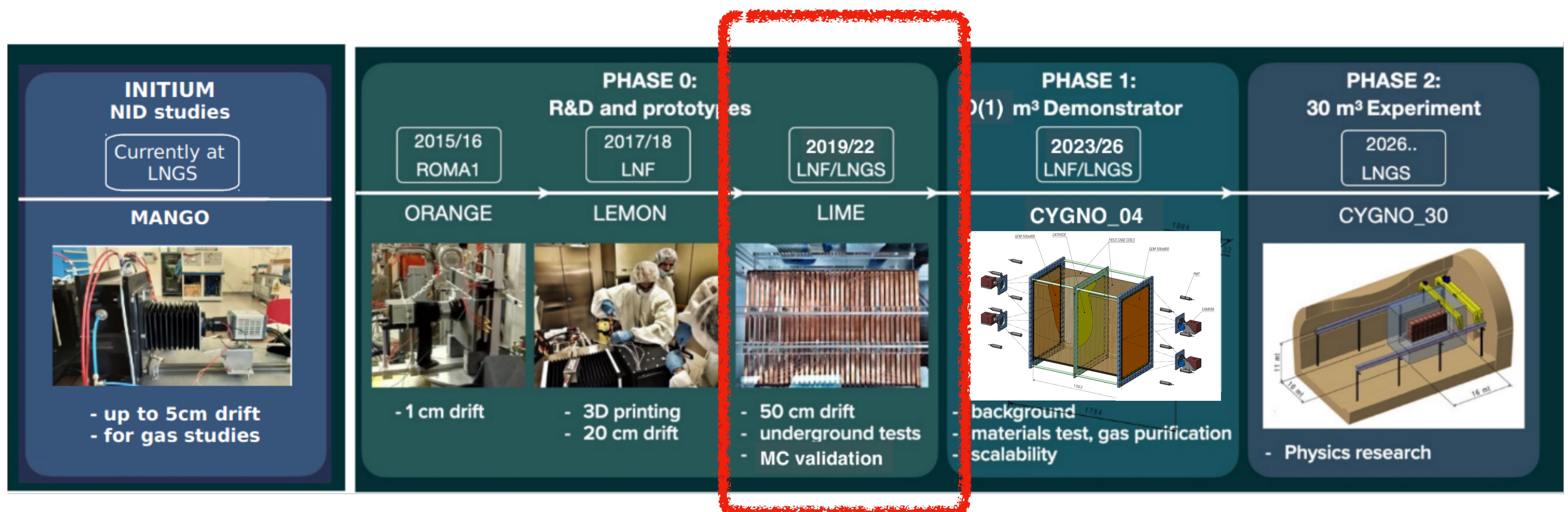


SCMOS camera [Hamamatsu Orca-Fusion]

4 PMTs [Hamamatsu R7378]



The CXGNO timeline



Instruments 6 (2022) 1, 6

JINST 15 (2020) 12, T12003

JINST 15 (2020) P08018

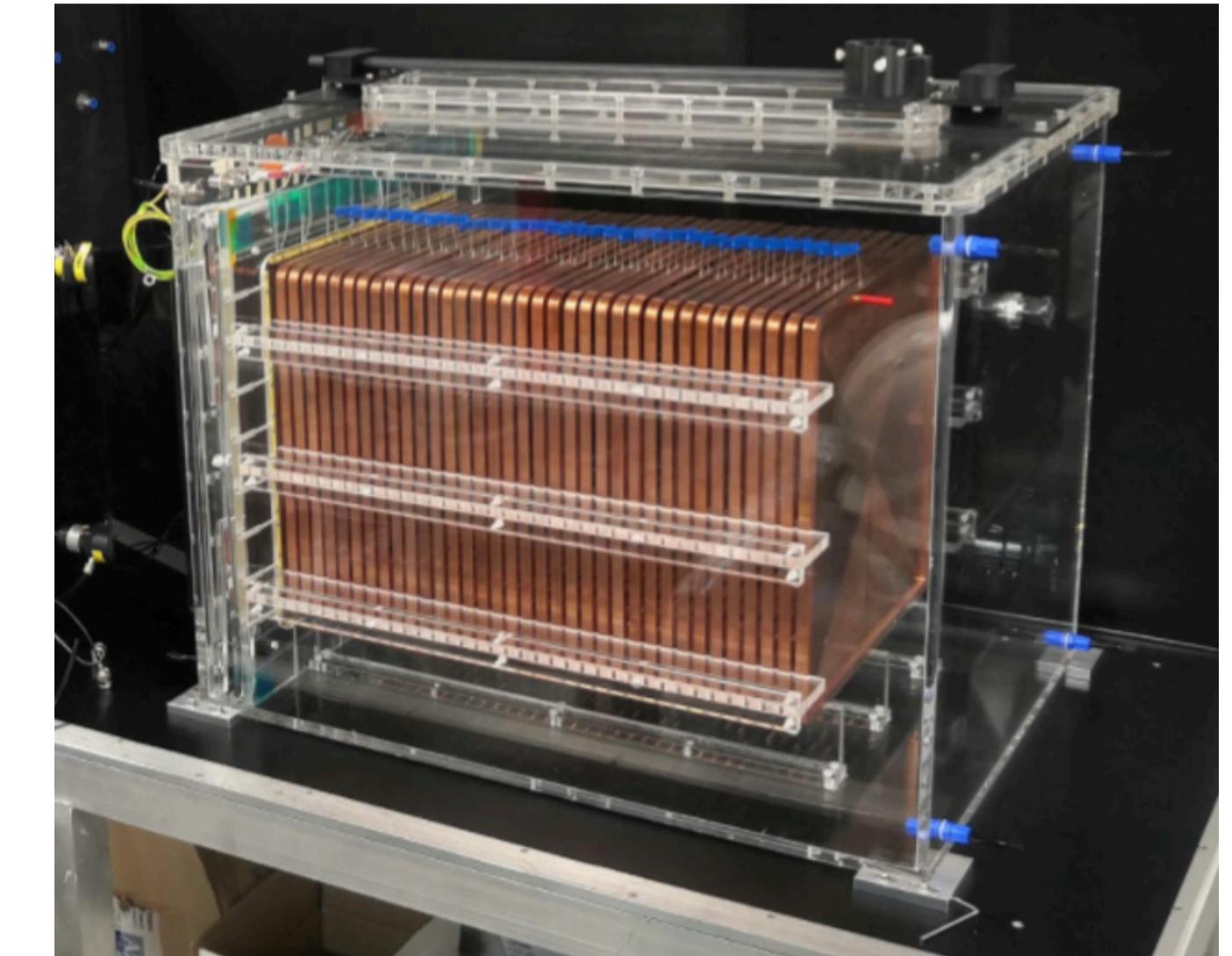
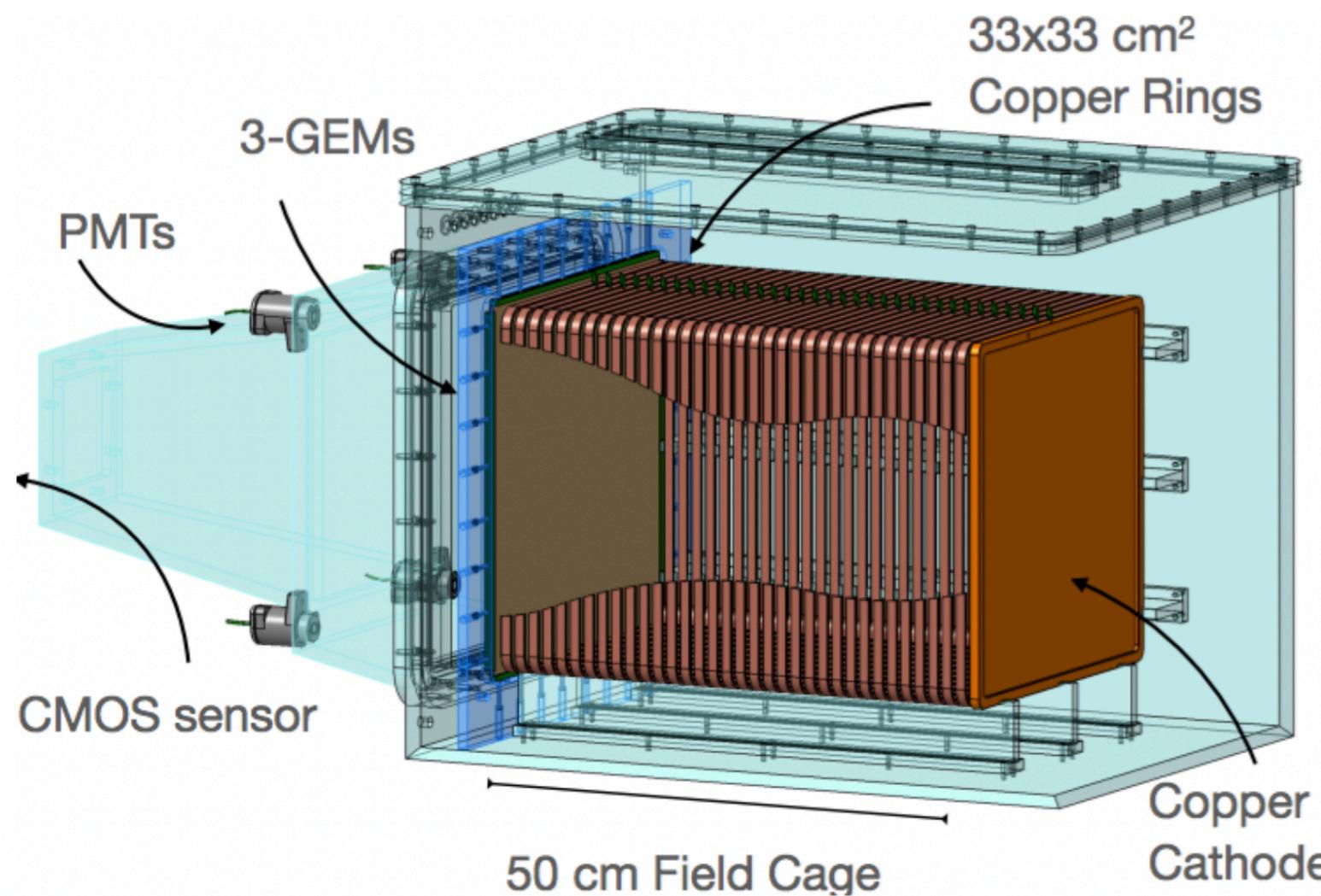
Measur.Sci.Tech. 32 (2021) 2, 025902

JINST 15 (2020) P10001

2019 JINST 14 P07011

NIM A 999 (2021) 165209

CXGNO PHASE 0: the LIME prototype



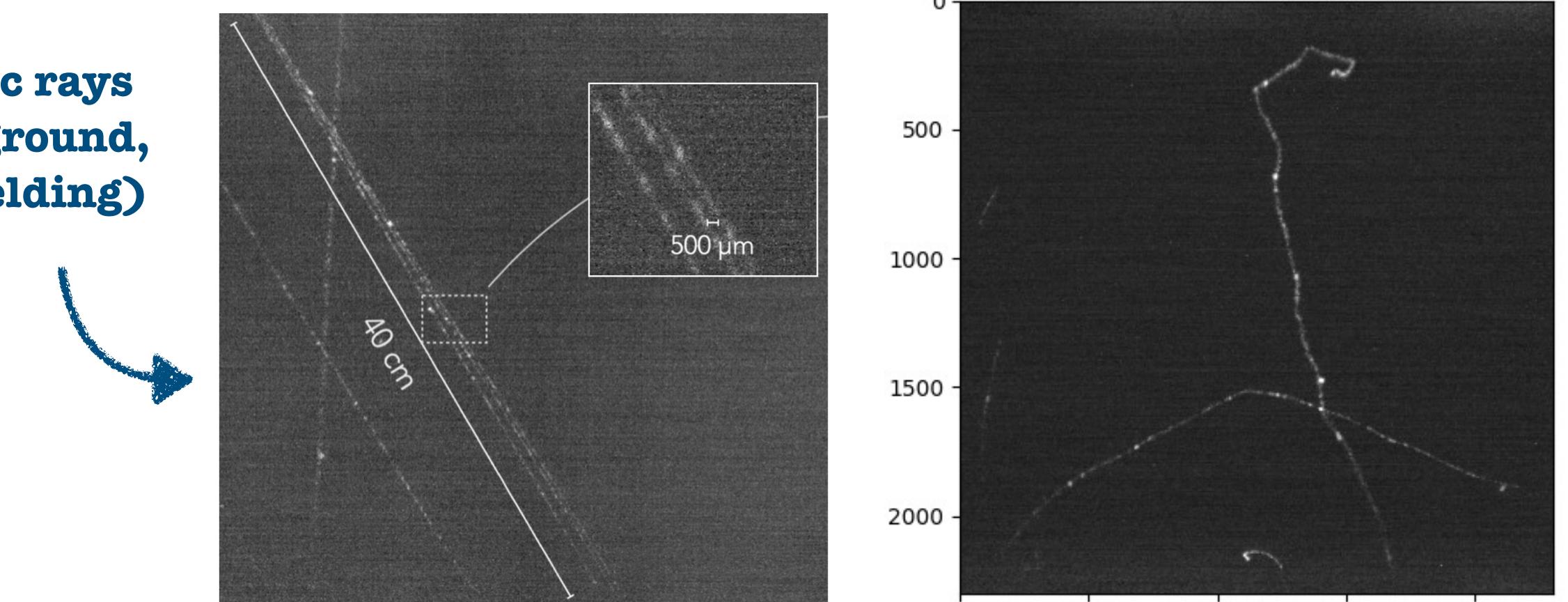
ORCA-Fusion

HIGH RESOLUTION
 2304×2304
5.3 Megapixels

READOUT NOISE
0.7 electrons rms
Ultra-quiet Scan



Natural radioactivity
(underground,
no shielding)

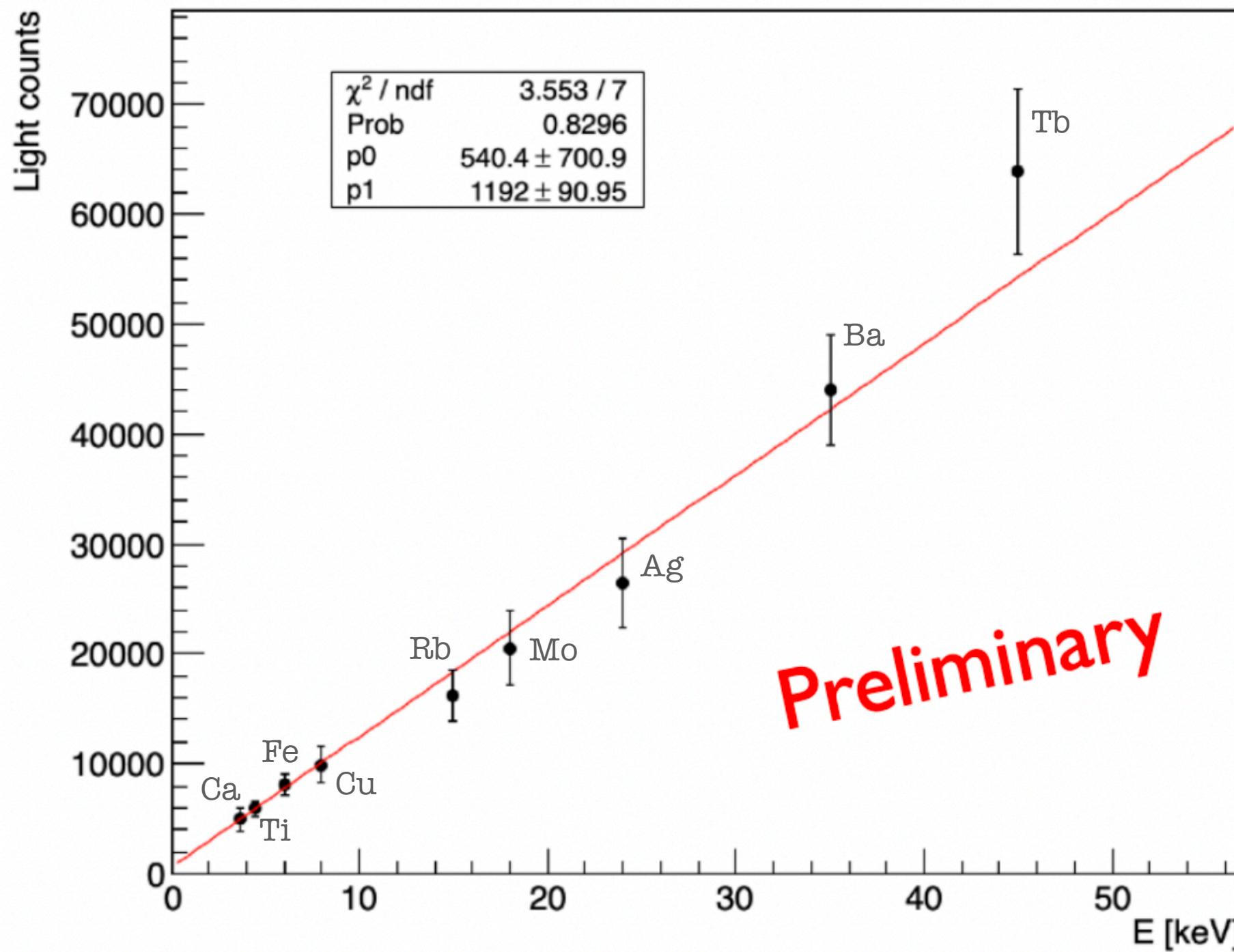


- He:CF₄ @ 1 atm
- Copper ring field cage, 50 cm drift
- 1 sCMOS sensor + 4 PMT
- 3 GEMs for a 33 x 33 cm² sensitive area

Cosmic rays
(overground,
no shielding)

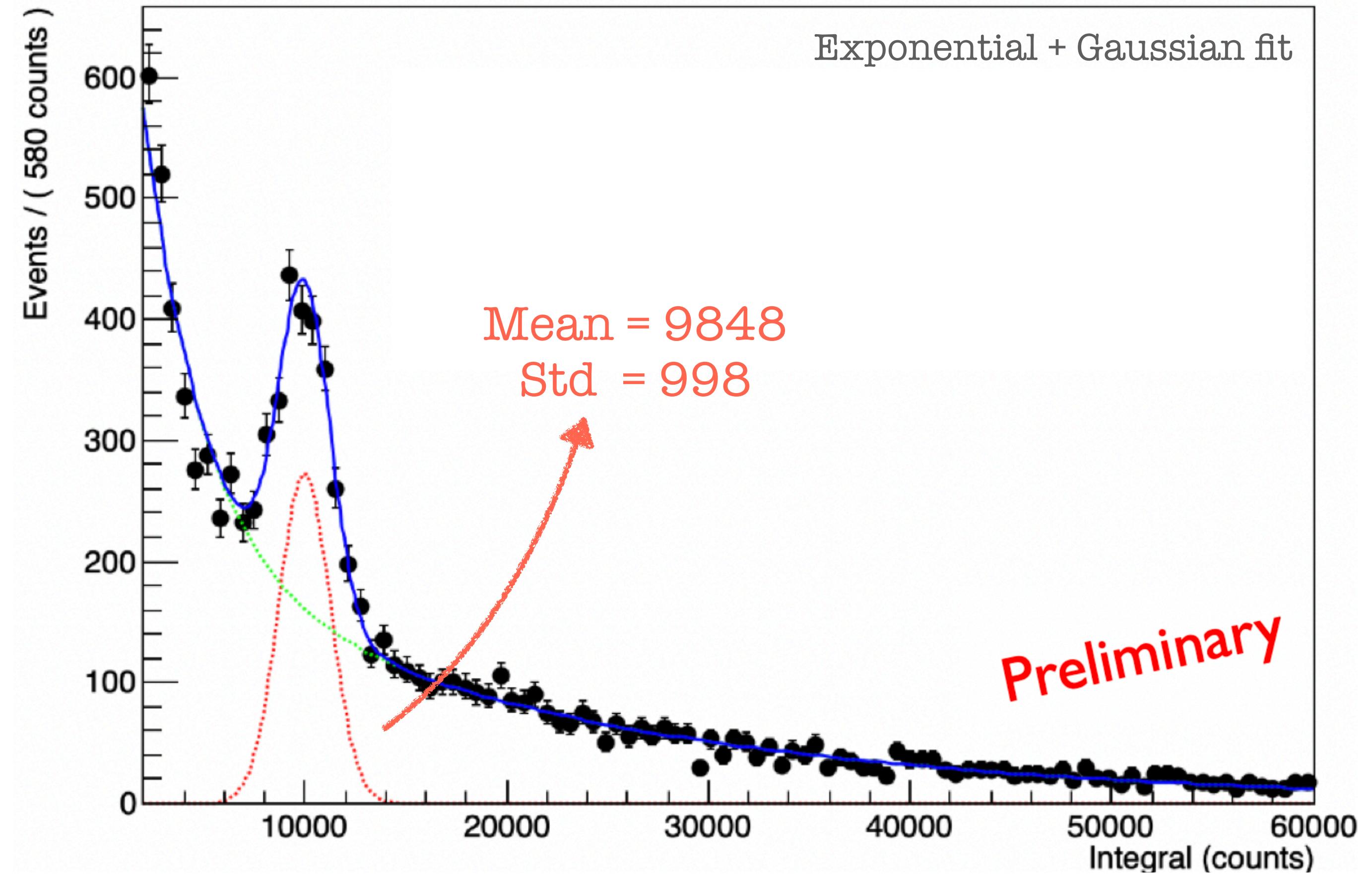
Overground tests: ER energy response

Electronic recoil (ER) calibration:



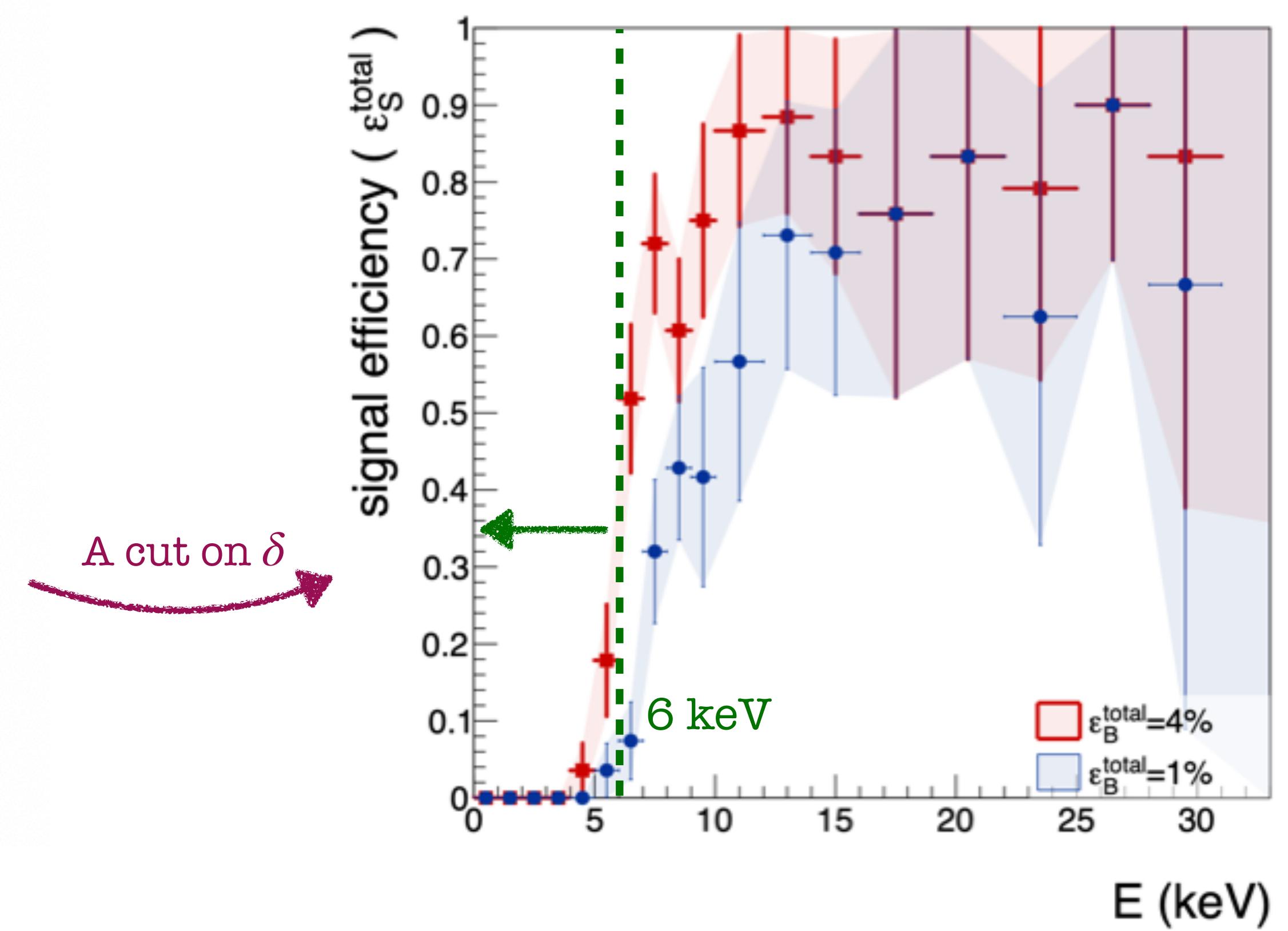
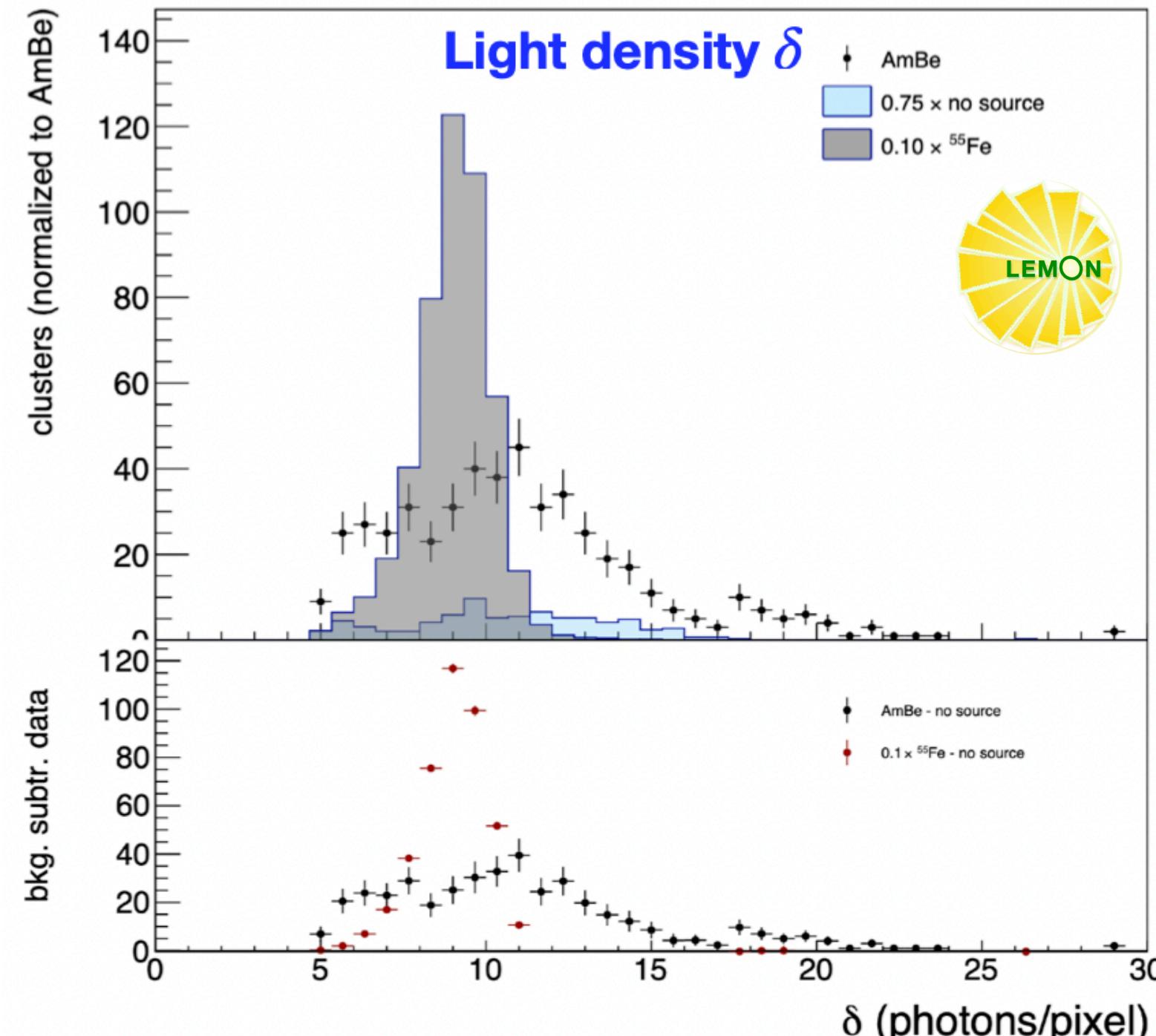
- **~ 13% energy resolution**
- **good linearity** in the response

Example: calibration with the 8 keV Cu X-rays



NR vs ER discrimination

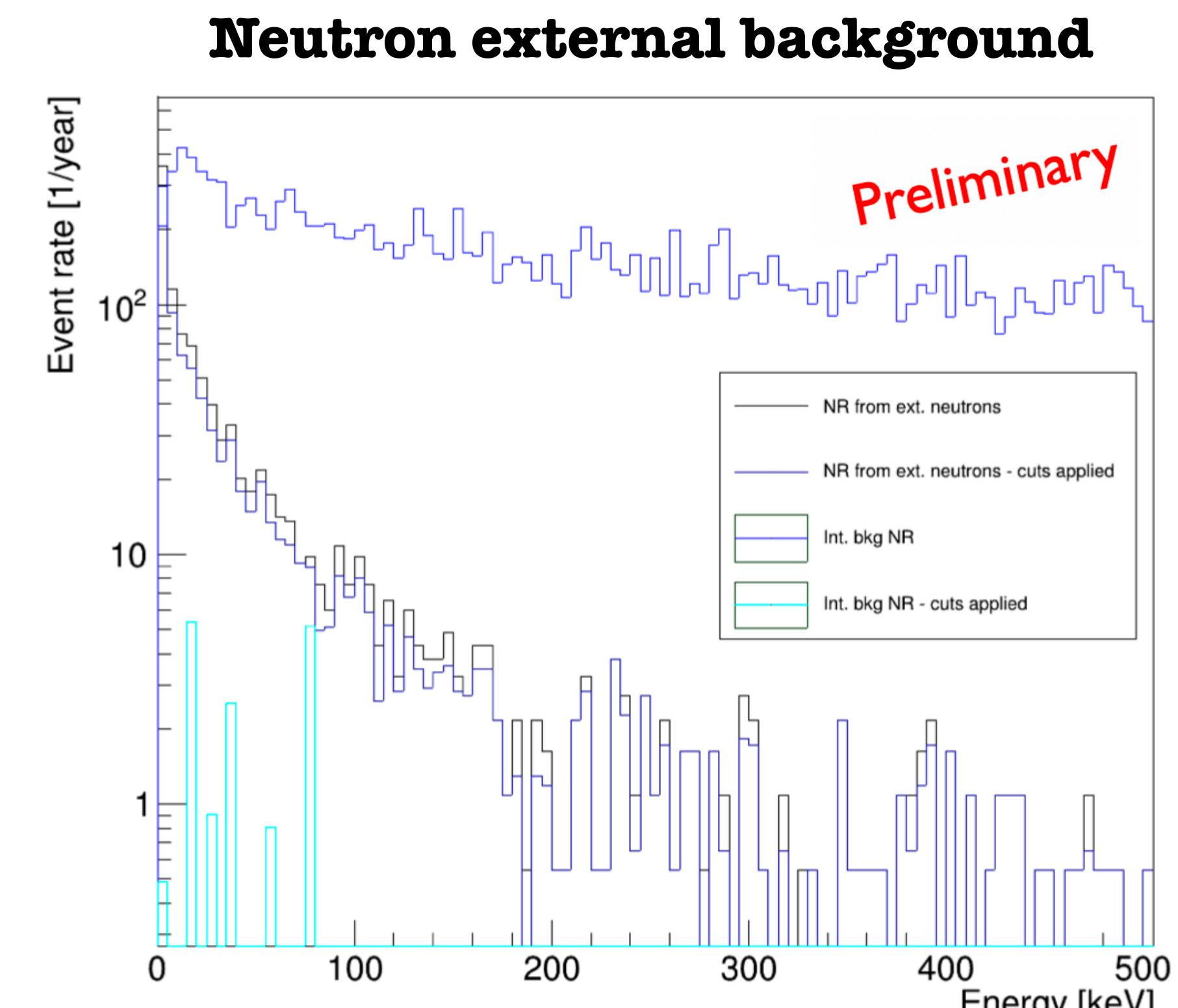
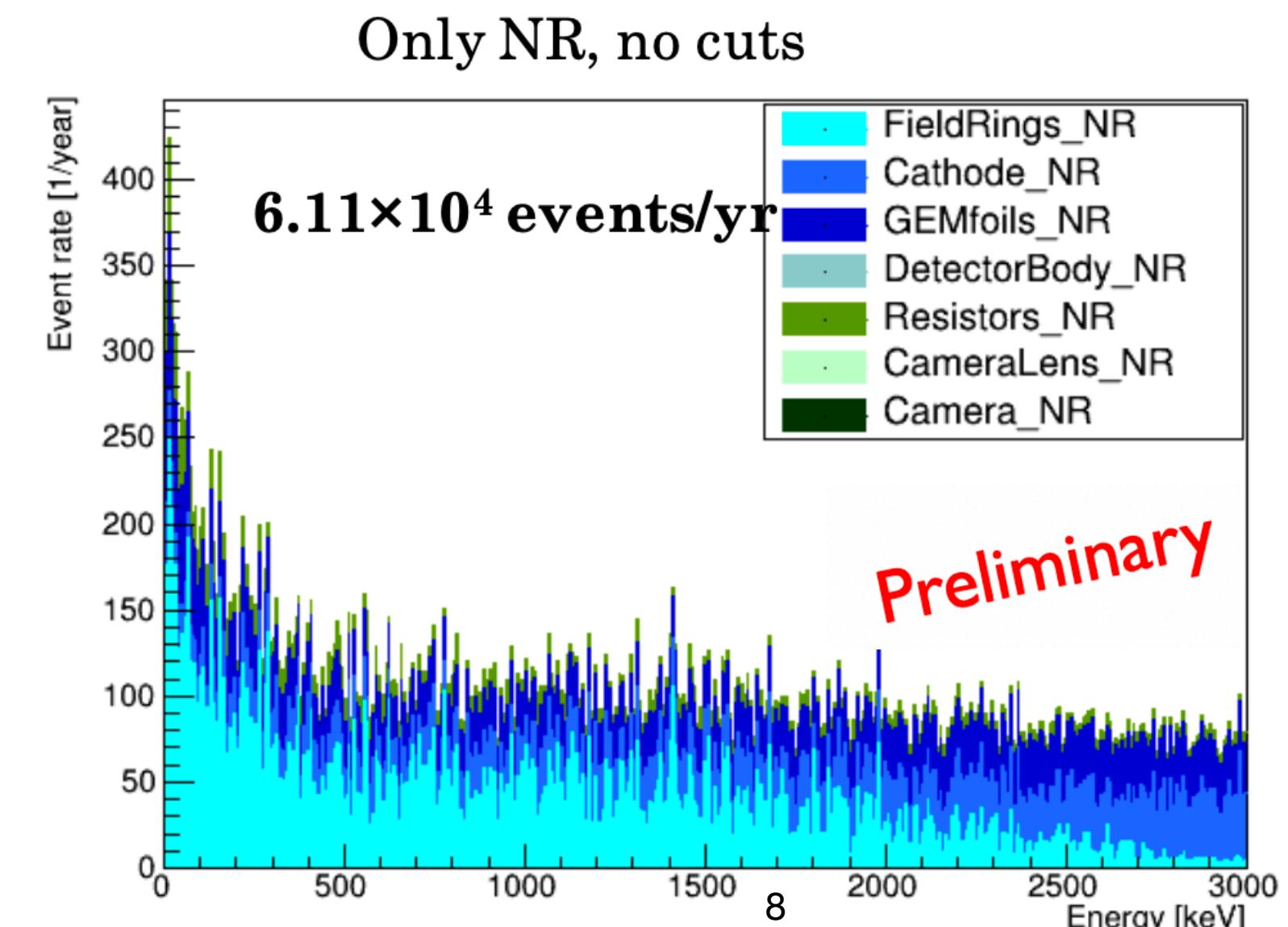
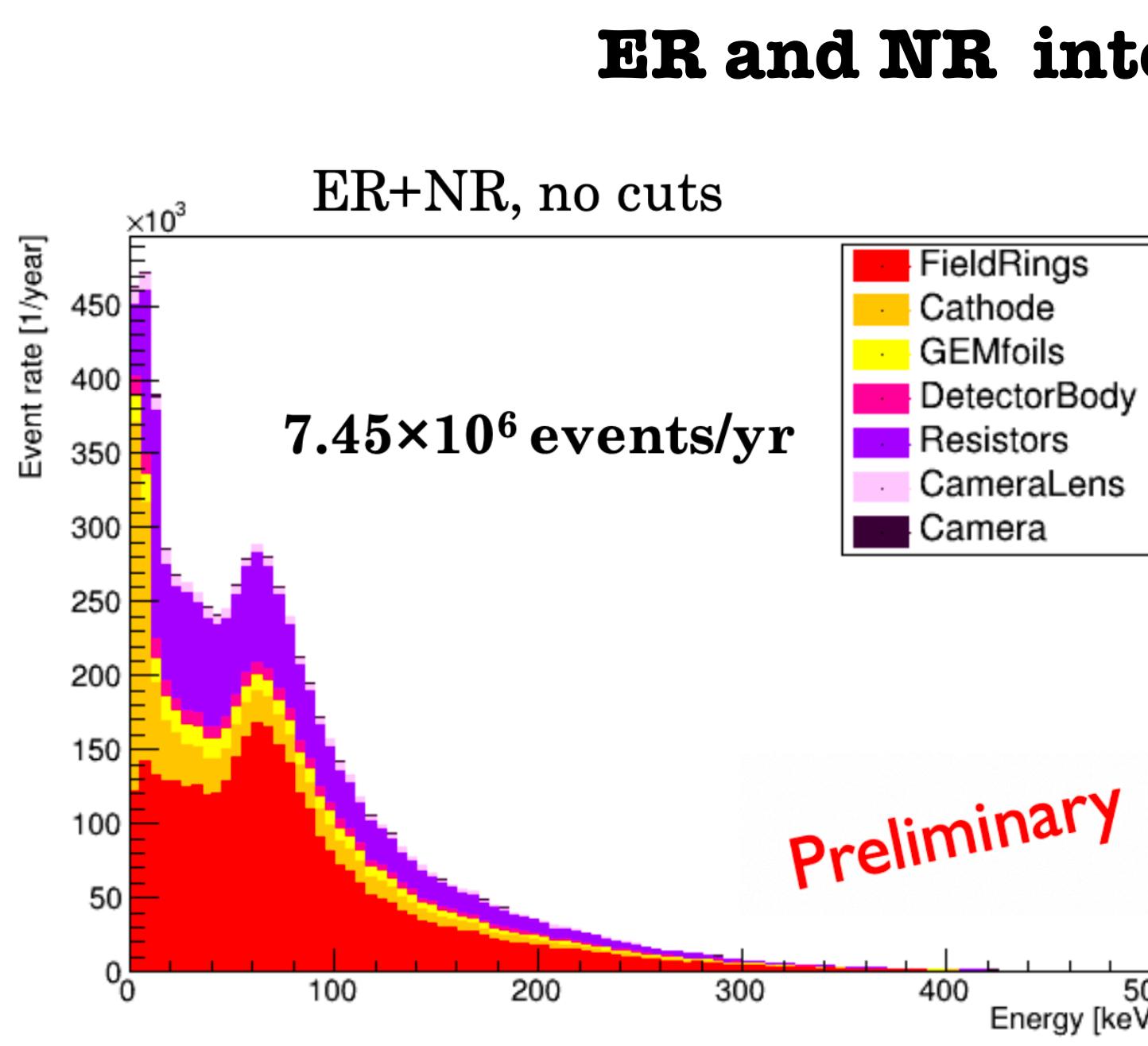
- NR vs ER discrimination power: preliminarily assessed with the AmBe LEMOn campaign



**40% NR efficiency
at 6 keV_{ee}**
obtained with 96%
rejection
against ^{55}Fe

Background studies for LIME

- **Activity** of the **components**: measured @ LNGS
- Main contribution to the **internal backgrounds**: copper rings, resistors, GEM/cathode
- 96% (99%) reduction of ER (NR) events with **fiducial cuts**
- **Studies** about the **shielding** are **ongoing** (10 cm copper + 40 cm water for the underground campaign)



Underground installation

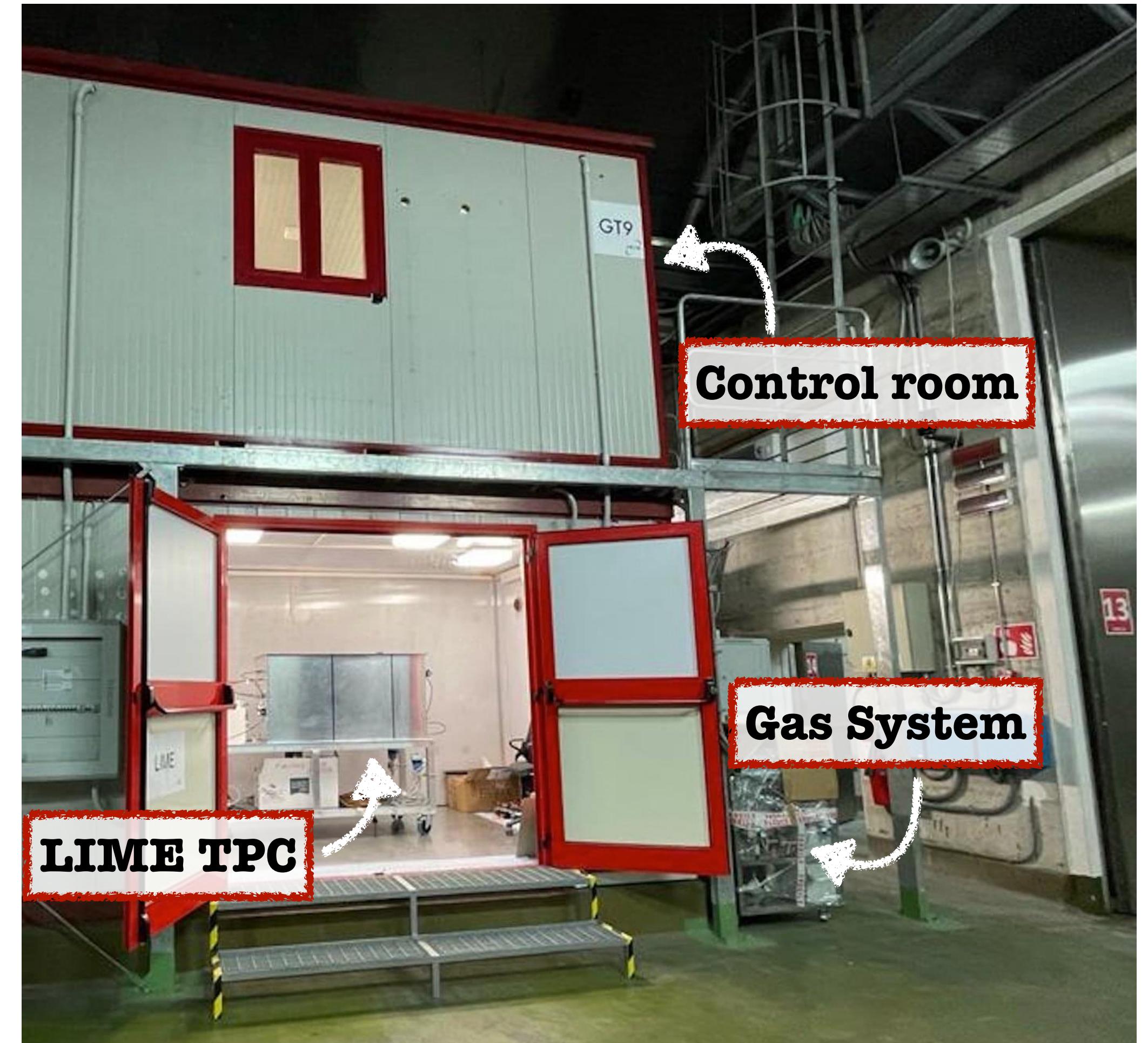
- The LIME prototype has been preliminarily assembled and **tested overground** at Laboratori Nazionali di Frascati (**LNF**)
- Moved underground** at Laboratori Nazionali del Gran Sasso (**LNGS**) the beginning of 2022



The TPC inside the Faraday cage



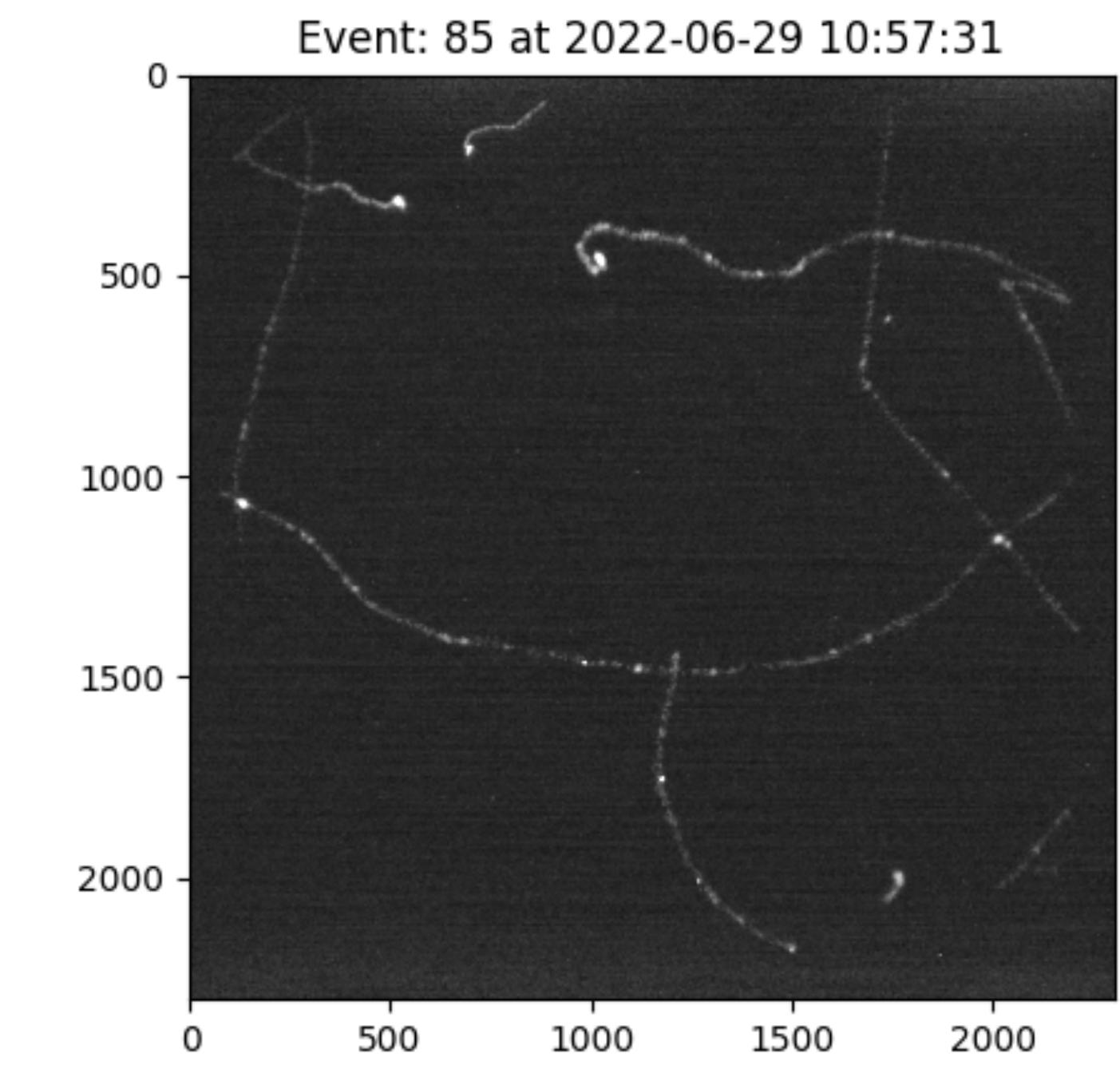
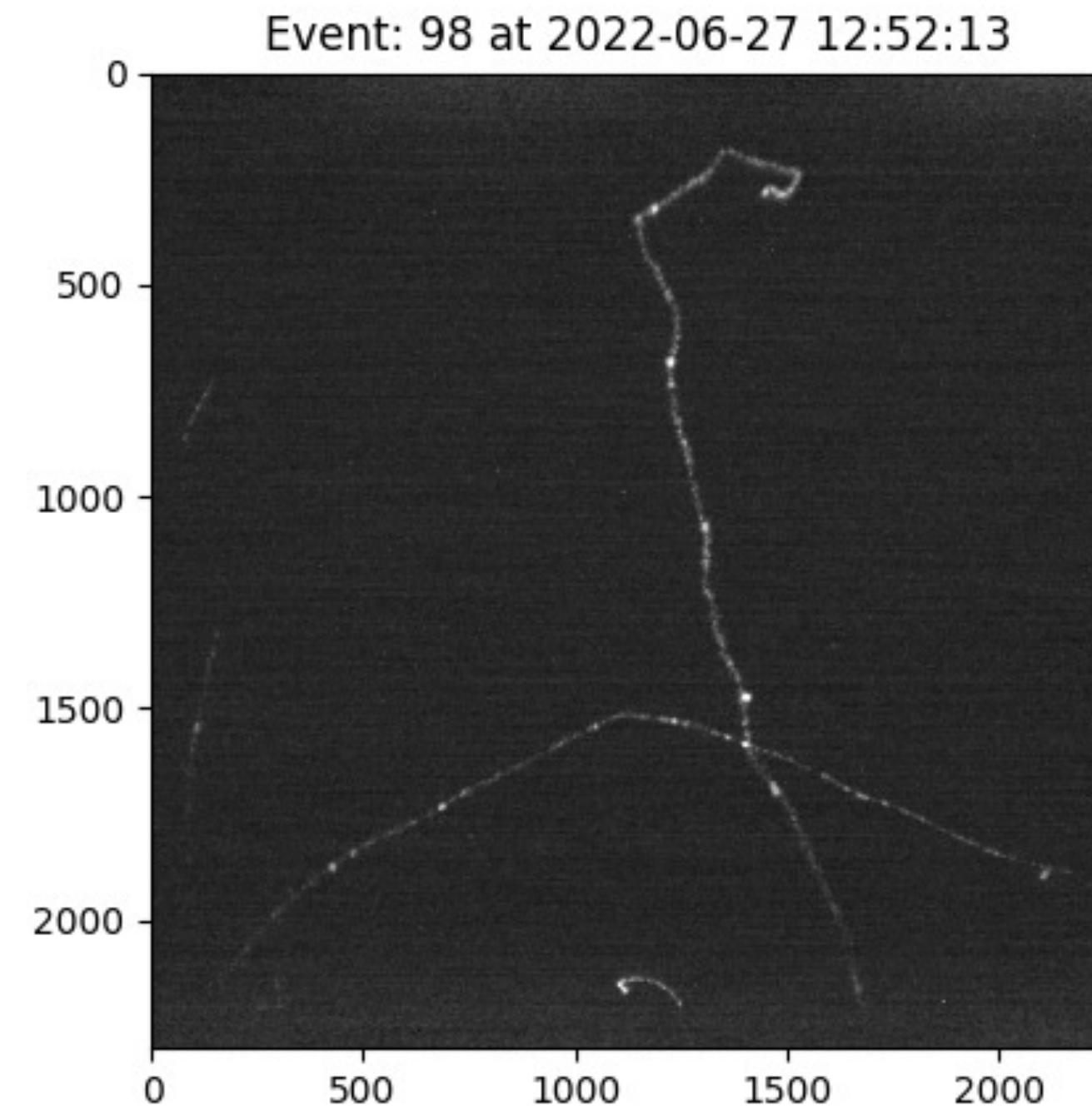
HV and DAQ crate



First images underground



- **LIME is currently collecting its first data** underground @ LNGS
- We are **testing the stability** of the chamber and the **gas system**
- In parallel, a clone of LIME, called **LIME2**, has been installed at **LNF** for the assessment and the setup of an **upgraded trigger and DAQ system**.



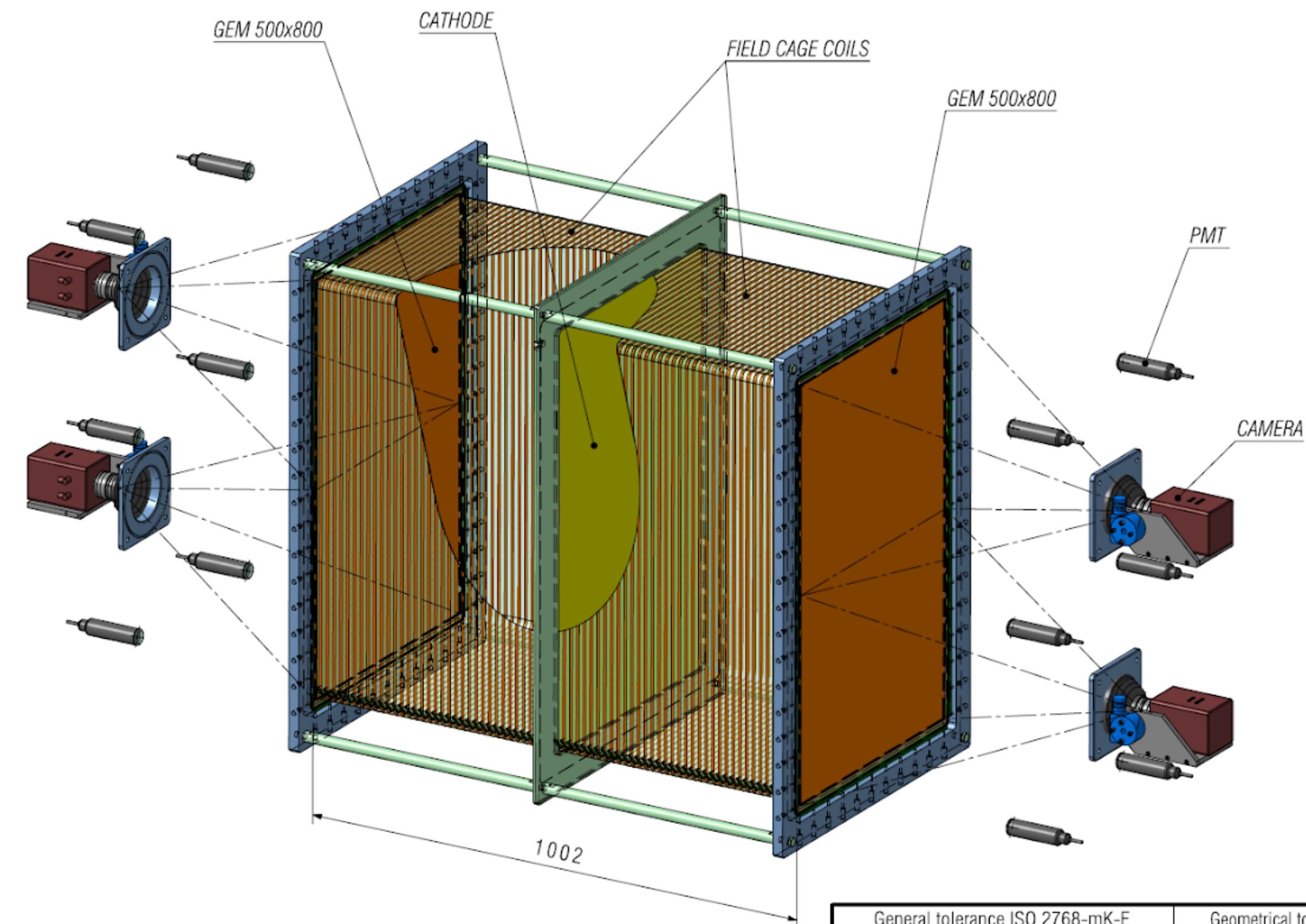
CXGNO PHASE 1: CYGNO_04



- **Preliminary design:**

- TPC made of **2 chambers** with a **common cathode**.
- Closed by 2 sets of **50 cm x 80 cm triple GEMs**
- **Readout** of each GEM side: 2 cameras with rectangular sensors (ORCA Quest) + 6 PMTs
- **Vessel**: low radioactivity PMMA
- **Shielding**: 10 cm copper + 100 cm water with a polyethylene base

Designed at LNF and to be installed at LNGS



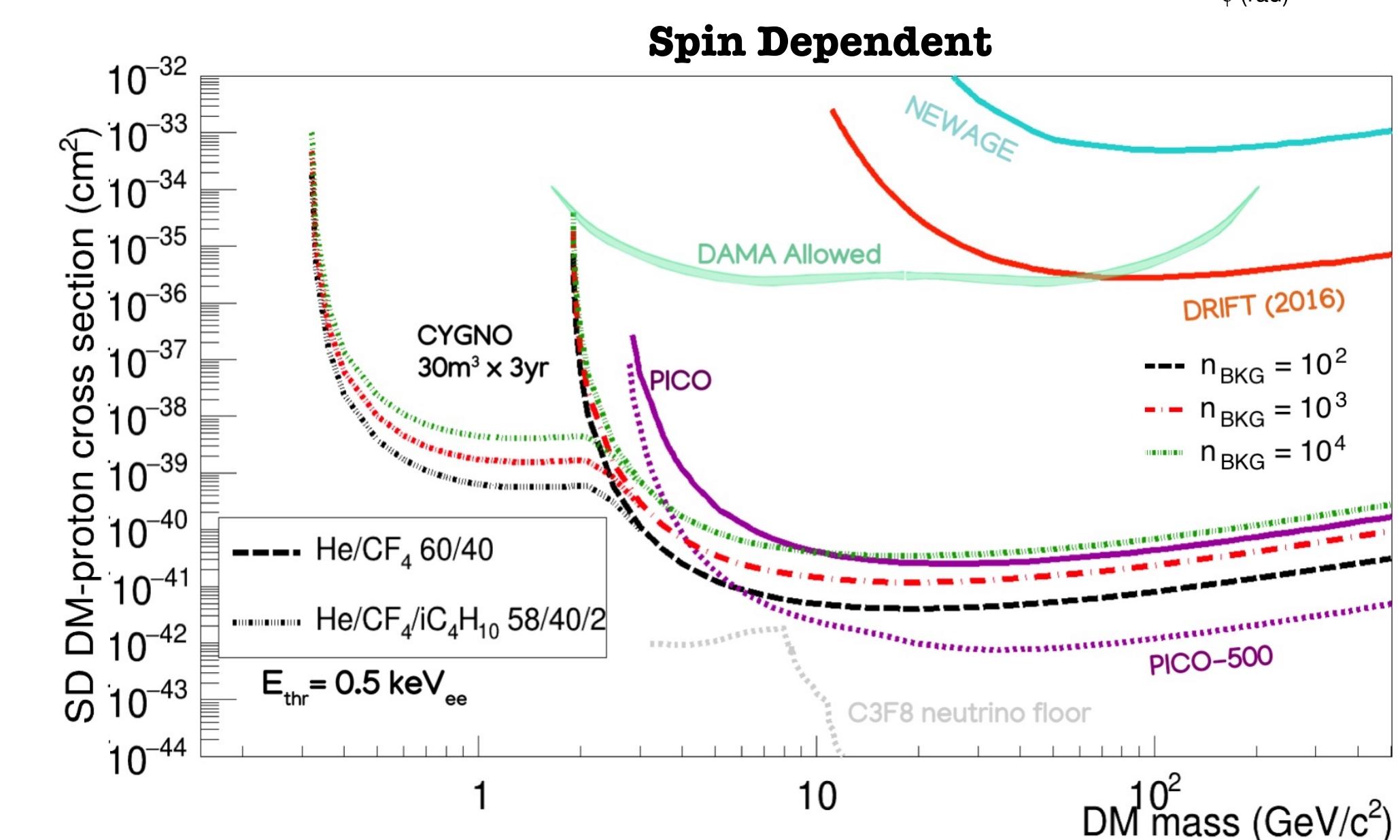
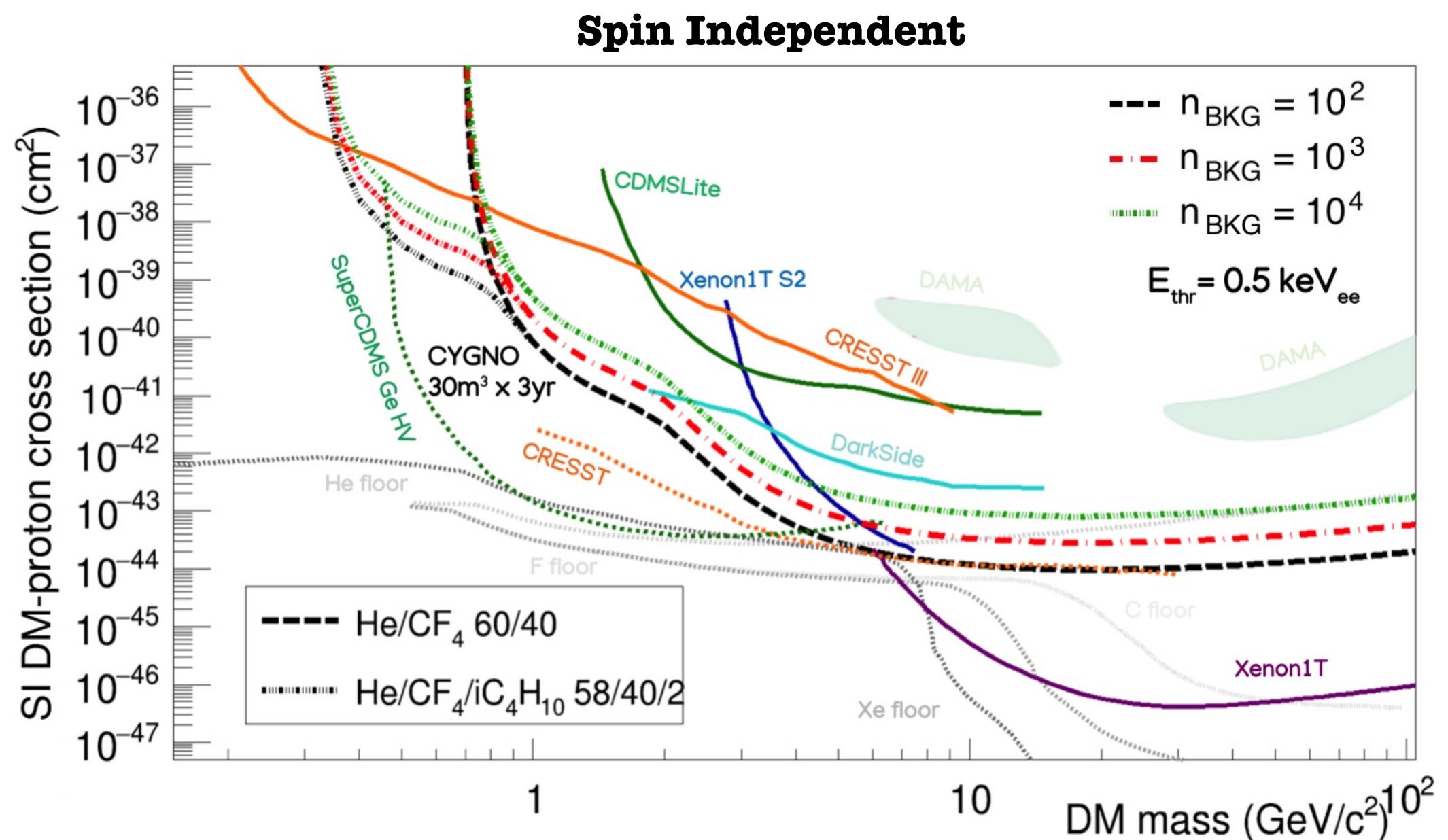
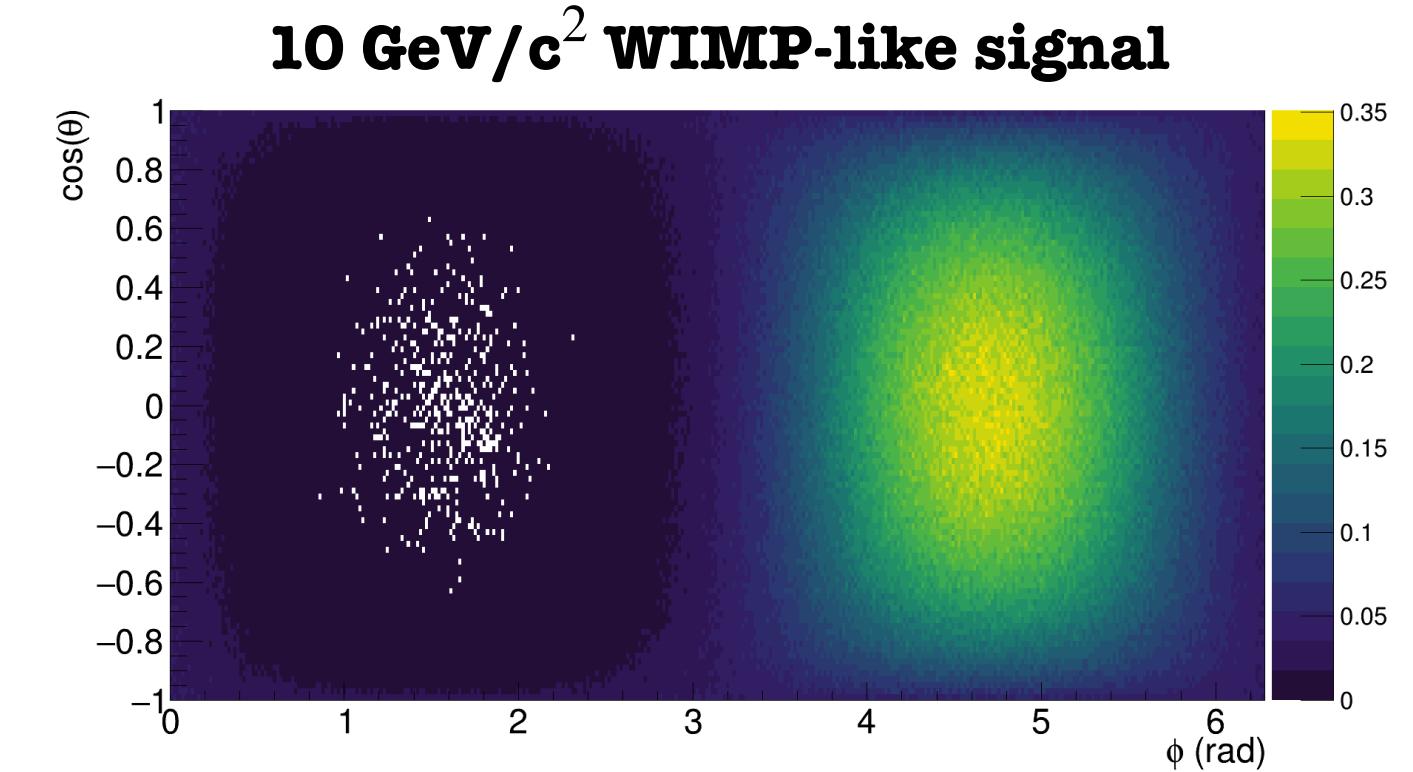
CYGNO 0.4 TPC
500x800x1000=0.4 m³

General tolerance ISO 2768-mK-E		Geometrical tolerance ISO 8015-E		Roughness ISO 1302	
REVISION	SIZE A3	DATE:	NAME:	DATE:	NAME:
	NATIONAL INSTITUTE FOR NUCLEAR PHYSICS				
	Frascati National Lab				
	RESEARCH DIVISION - SEM				
	STUDIO MECCANICO				
PROJECTION					
TOTAL WEIGHT (kg)	11/06/2022	DATE	DRAWN		
SCALE		DATE	CHECKED		
1:8					
SHEET		DATE	APPROVED		
1/3					
CYGNO EXPERIMENT					
CYGNO 0.4 DETECTOR					
TPC COMPONENTS SCHEME					
CY4-01-P					

CYGNOPHASE 2: the 30 m^3 experiment for low mass DM searches



- $\mathcal{O}(30\text{-}100\text{ m}^3)$ detector for directional dark matter (**DM**) search in the $\sim\text{GeV}/c^2$ mass region.
- Preliminary sensitivity projections. **Assumptions:**
 - Low energy **threshold**: 1 keV (0.5 keV)
 - **Quenching Factor**: SRIM simulation
 - **Observable**: angular distribution
 - **Angular resolution**: 30°
 - **Background**: different scenarios (isotropic background)



Conclusions



- The **CXGNO** collaboration is developing a **gaseous TPC with optical readout**
- **LIME**, a 50 L prototype, has recently been moved @ LNGS and is now taking data.
- This is just a step towards larger detectors for **directional dark matter search**
- **Other R&D activities:**
 - **Internal background reduction:** custom sCMOS sensor, custom lens
 - **Studies of the gas mixture:** addition of hydrogen-rich gas (CH_4), study of the light yield from hydrocarbons, possible eco-friendly gas mixture to substitute CF_4
 - **Negative ion drift:** negative ions with the addition of SF_6 for better fiducialization, 1 atm tests with encouraging results

Acknowledgements

This project has received fundings under the European Union's Horizon 2020 research and innovation programme from the Marie Skłodowska-Curie grant agreement No 657751 and from the European Research Council (ERC) grant agreement No 818744

CYGNO Project is funded by INFN.



Thanks for the attention!

The **CYGNO** collaboration:

F.. D. Amaro, E. Baracchini, L. Benussi, S. Bianco, C. Capoccia, M. Caponero, D. S. Cardoso, G. Cavoto, A. Cortez, R. J. de Cruz Roque, I. A. Costa, E. Dané, E. Di Marco, G. Grilli di Cortona, G. D'Imperio, G. Dho, F. Di Giambattista, R. R. M. Gregorio, F. Iacoangeli, H. P. Lima Júnior, G. Maccarrone, R. D. P. Mano, M. Marafini, G. Mazzitelli, A. G. Mc Lean, A. Messina, M. L. Migliorini, C.M.B. Monteiro, R. A. Nóbrega, A. Orlandi, I. F. Pains, E. Paoletti, L. Passamonti, F. Petrucci, S. Pelosi, S. Piacentini, D. Piccolo, D. Pierluigi, D. Pinci, A. Prajapati, F. Renga, F. Rosatelli, A. Russo, J.M.F. dos Santos, G. Saviano, A. da Silva Lopes Júnior, N. Spooner, R. Tesauro, S. Tomassini, S. Torelli



Istituto Nazionale di Fisica Nucleare

Backup

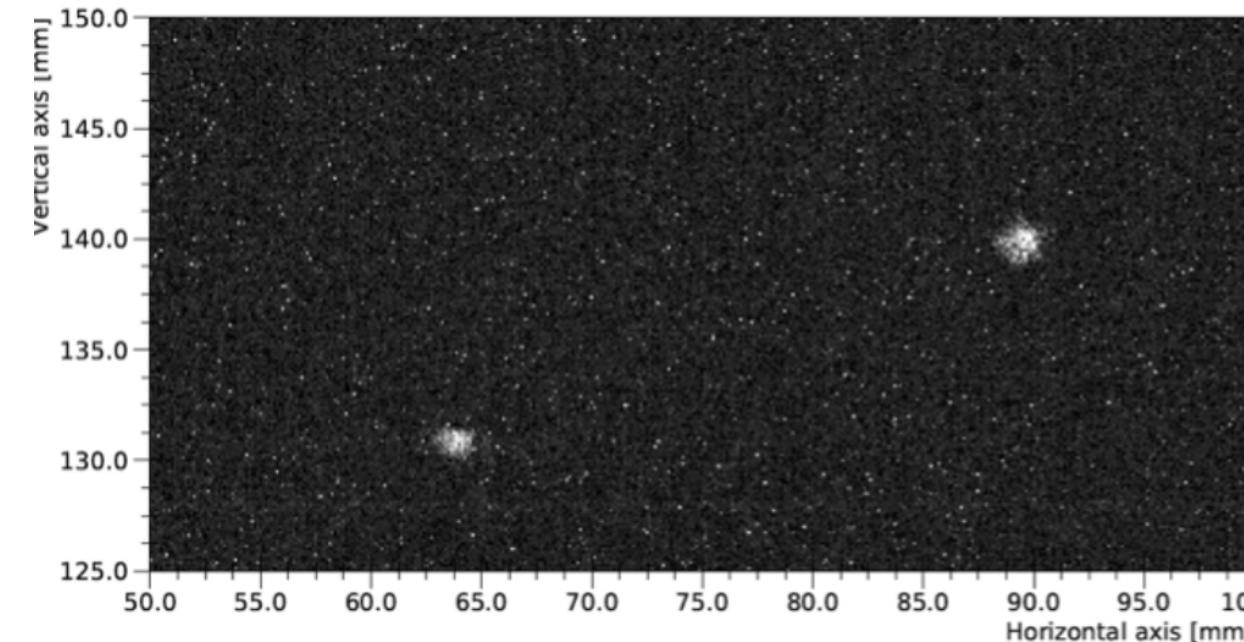
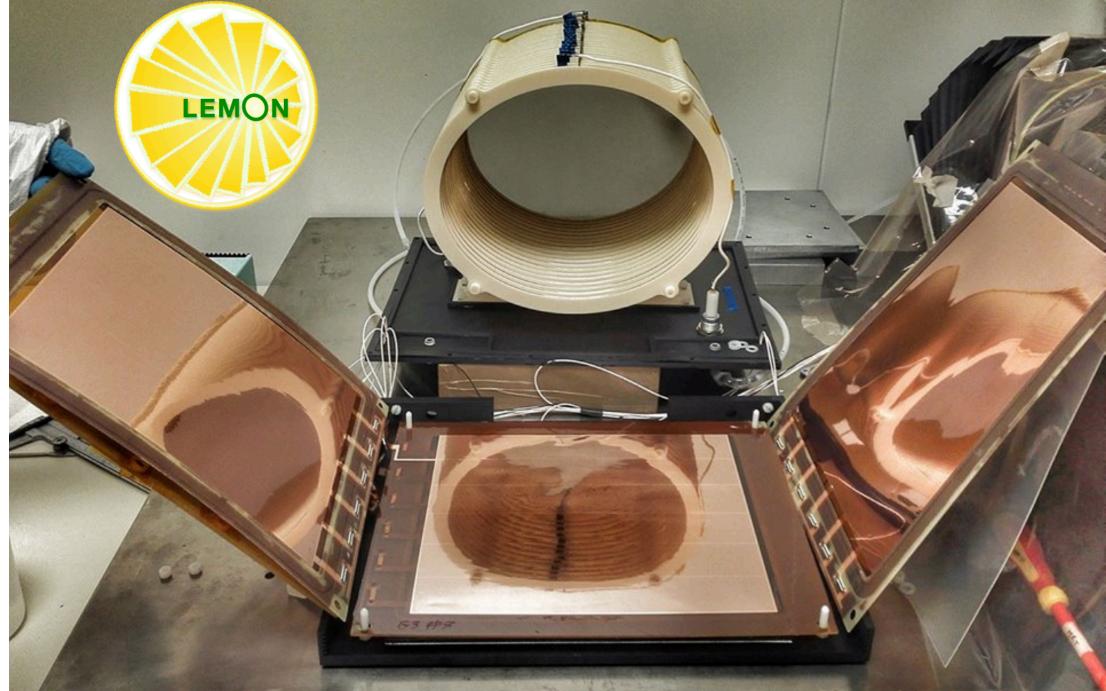
Detector performances: LEMOn



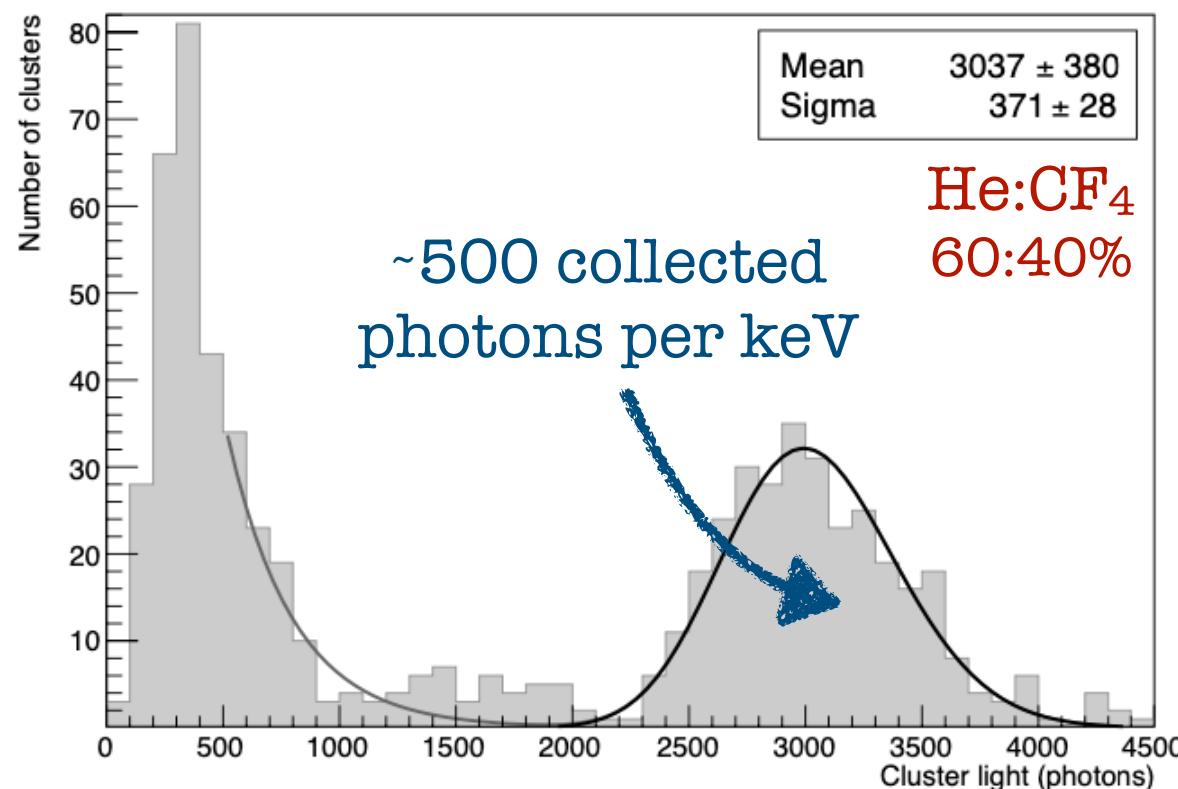
Energy resolution



^{55}Fe X-ray signal

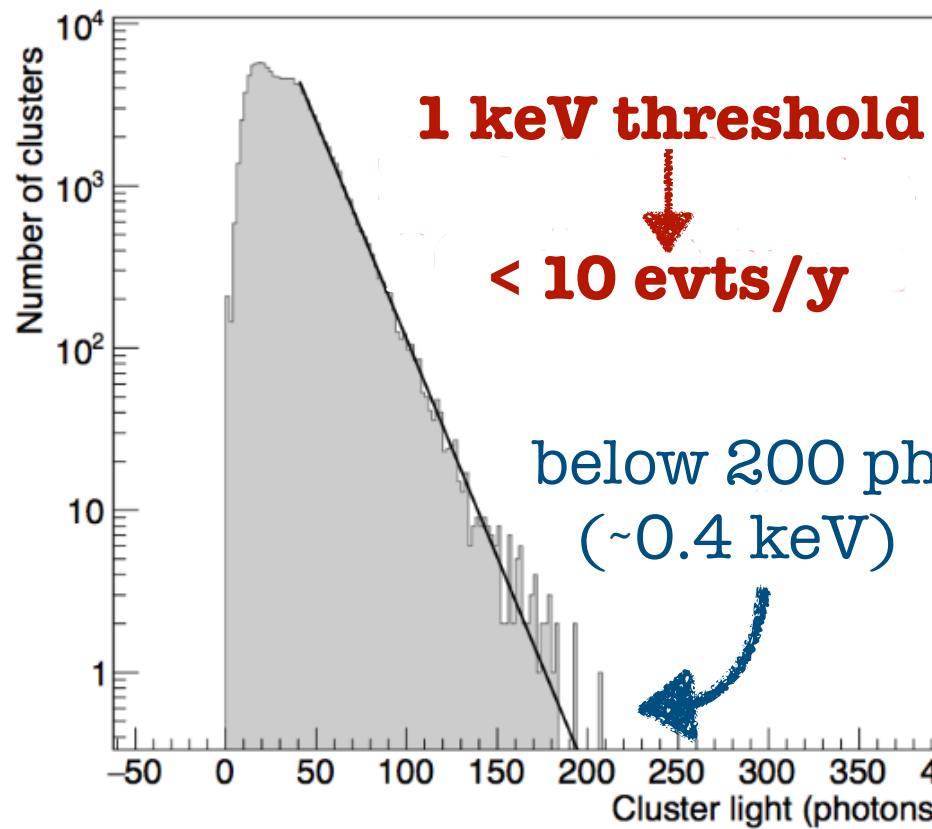


Energy resolution



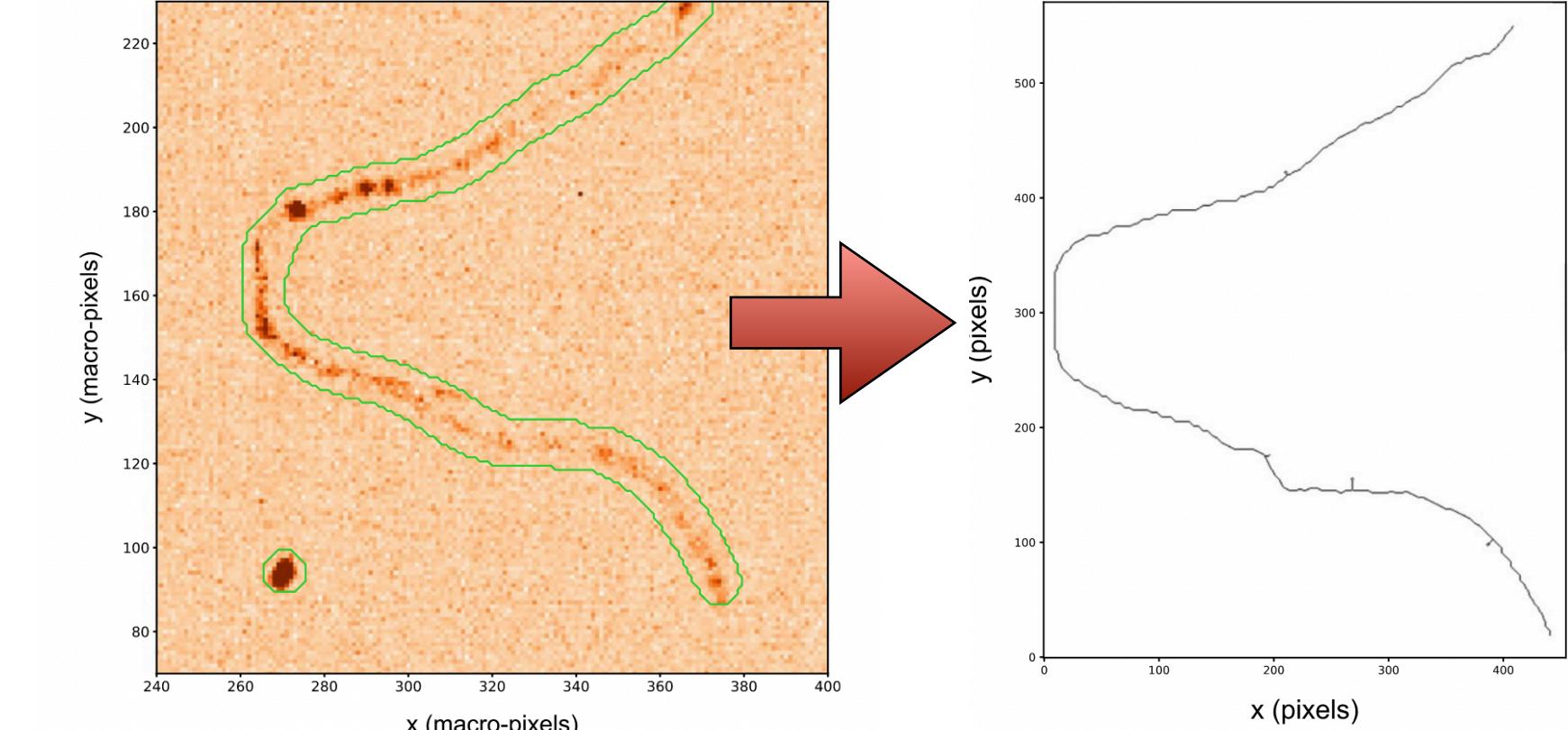
Energy resolution of **15% at 5.9 keV_{ee}** with both sCMOS and PMT

sCMOS sensor noise

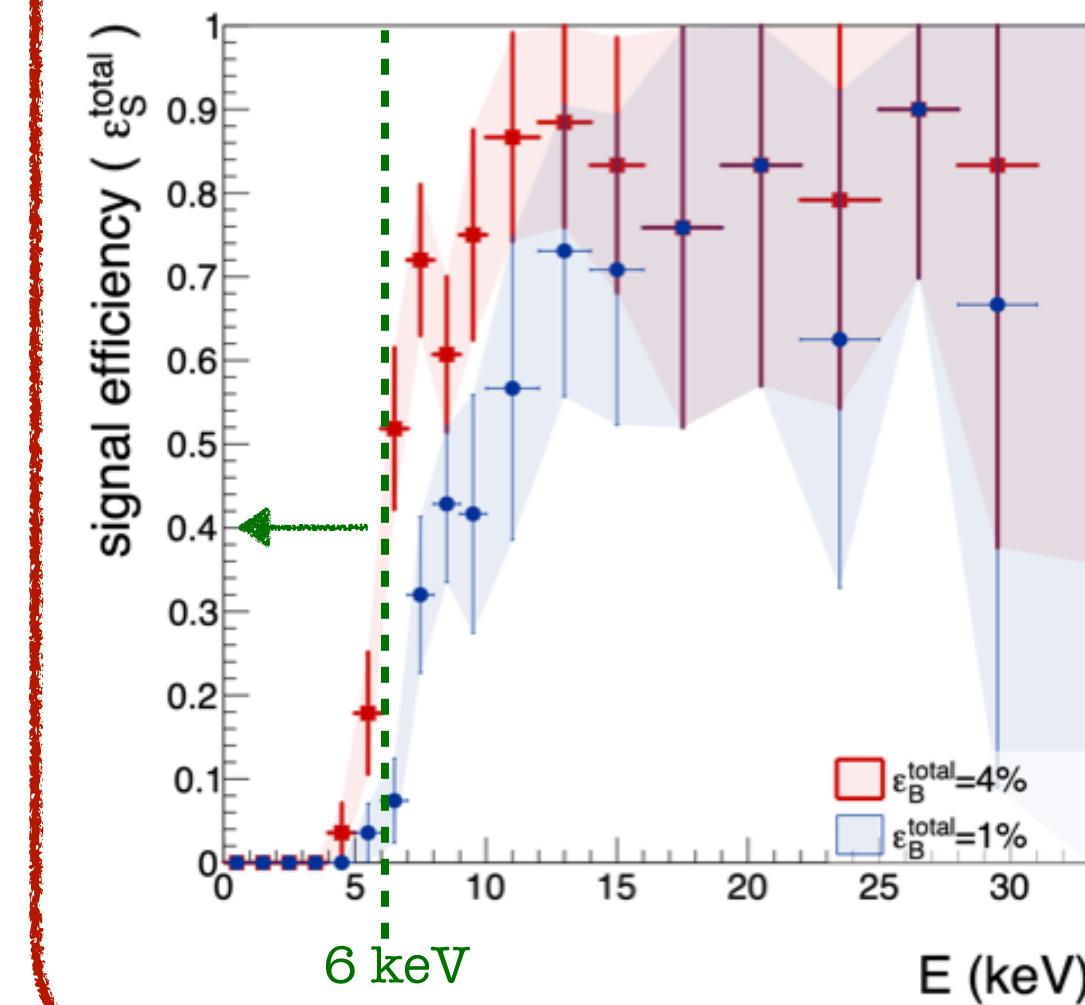


Reconstruction efficiency

Multiple DBSCAN + GAC* pattern recognition algorithm



ER/NR discrimination power



40% NR efficiency at 6 keV_{ee}
obtained with 96% rejection against ^{55}Fe

Detector performances: LEMOn

z position reconstruction



$$\eta = \frac{\sigma}{A}$$

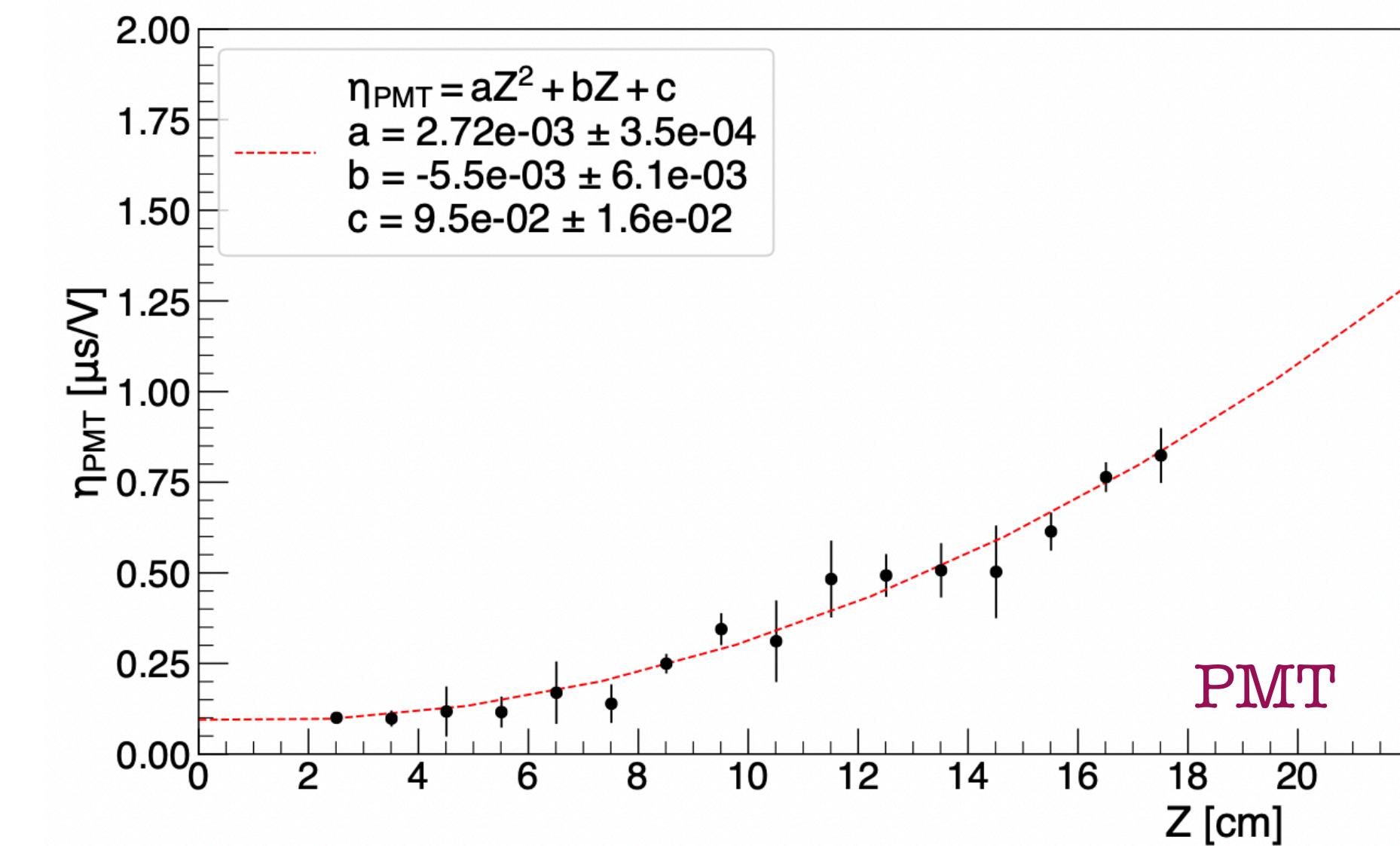
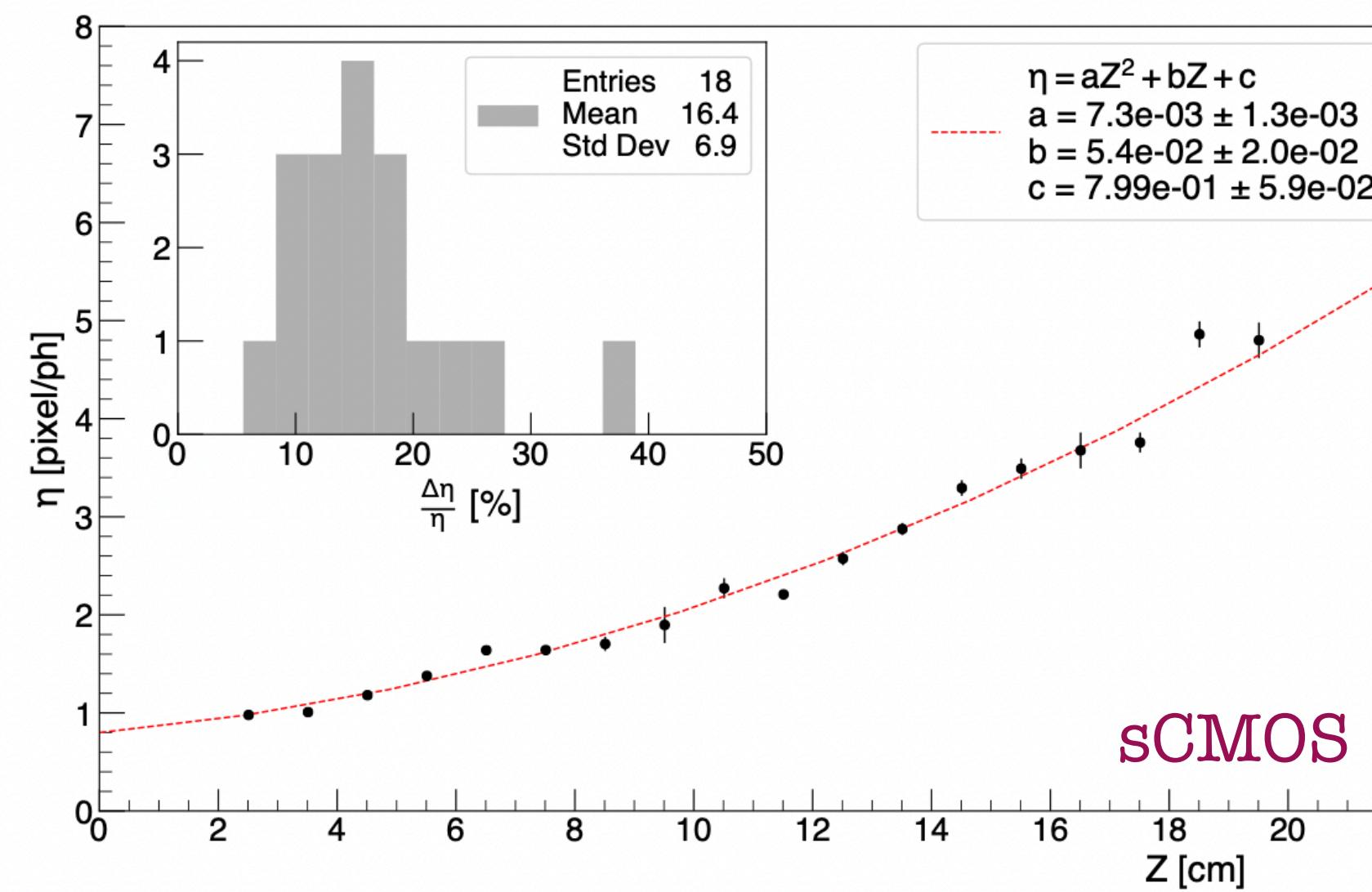
Width of the transverse light profile

Amplitude of the transverse light profile

$$\eta_{\text{PMT}} = \frac{\sigma_{\text{PMT}}}{A_{\text{PMT}}}$$

Time width of the PMT signal

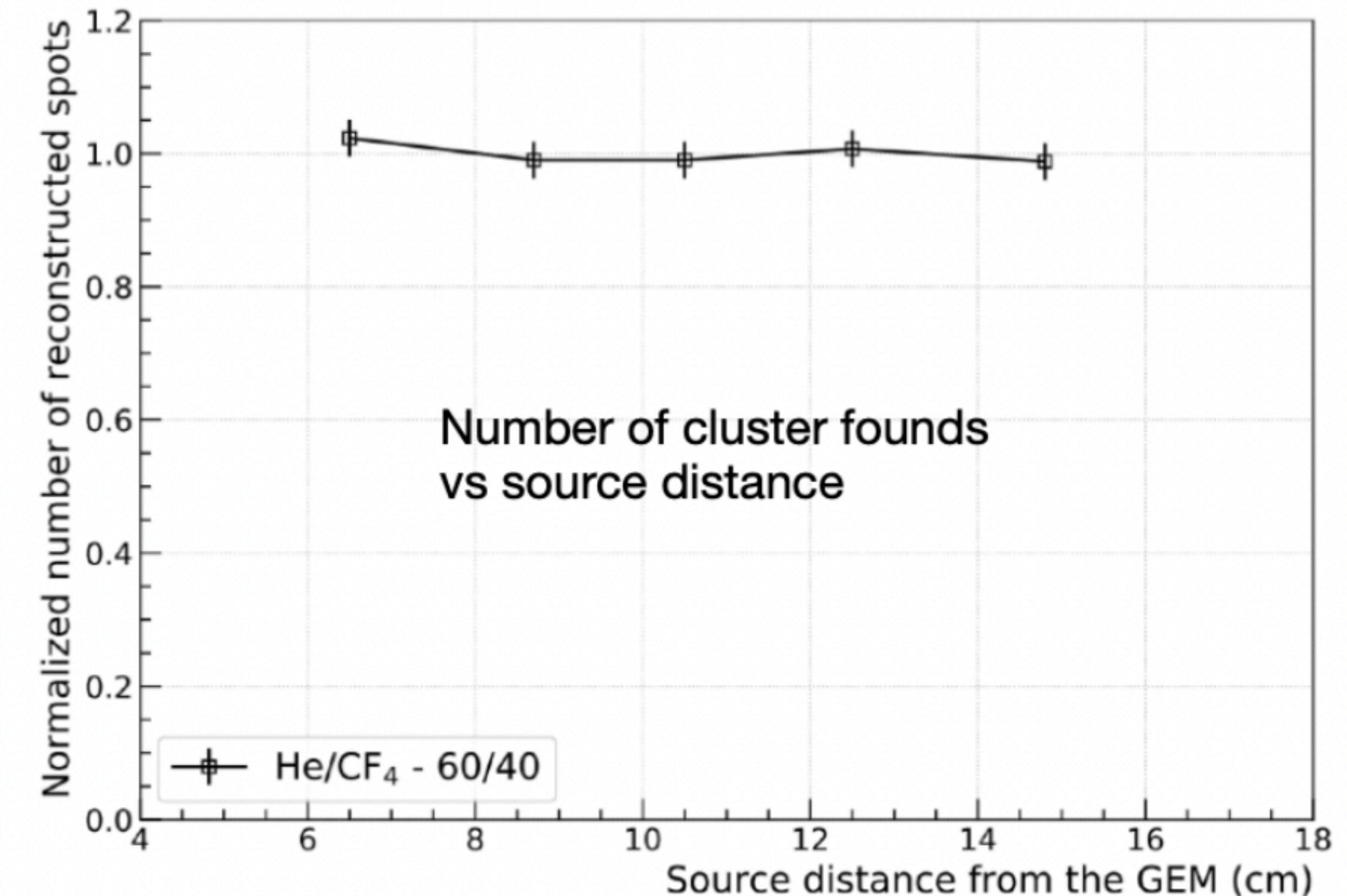
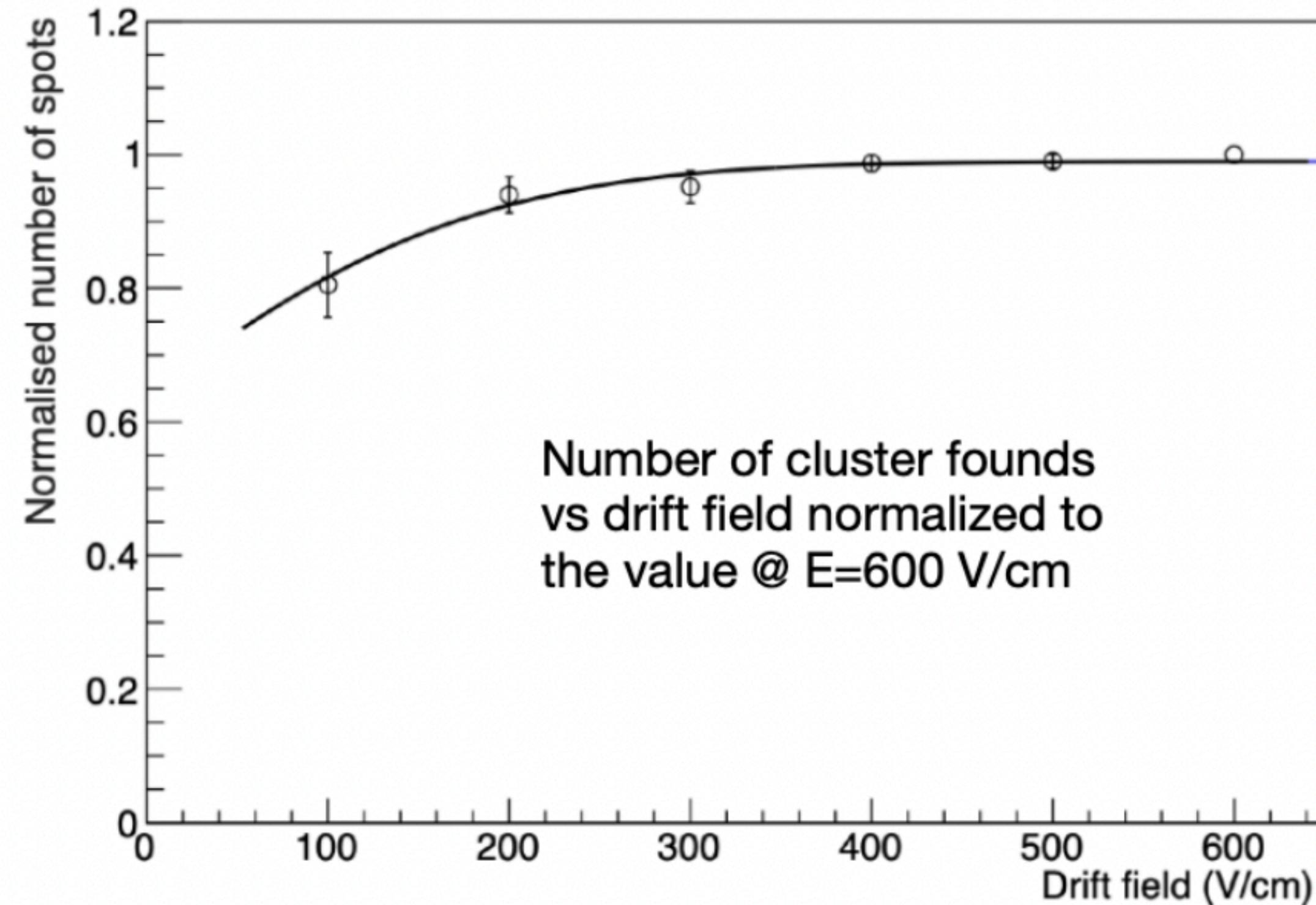
Amplitude of the PMT signal



Using both sCMOS and PMT: 15% z resolution

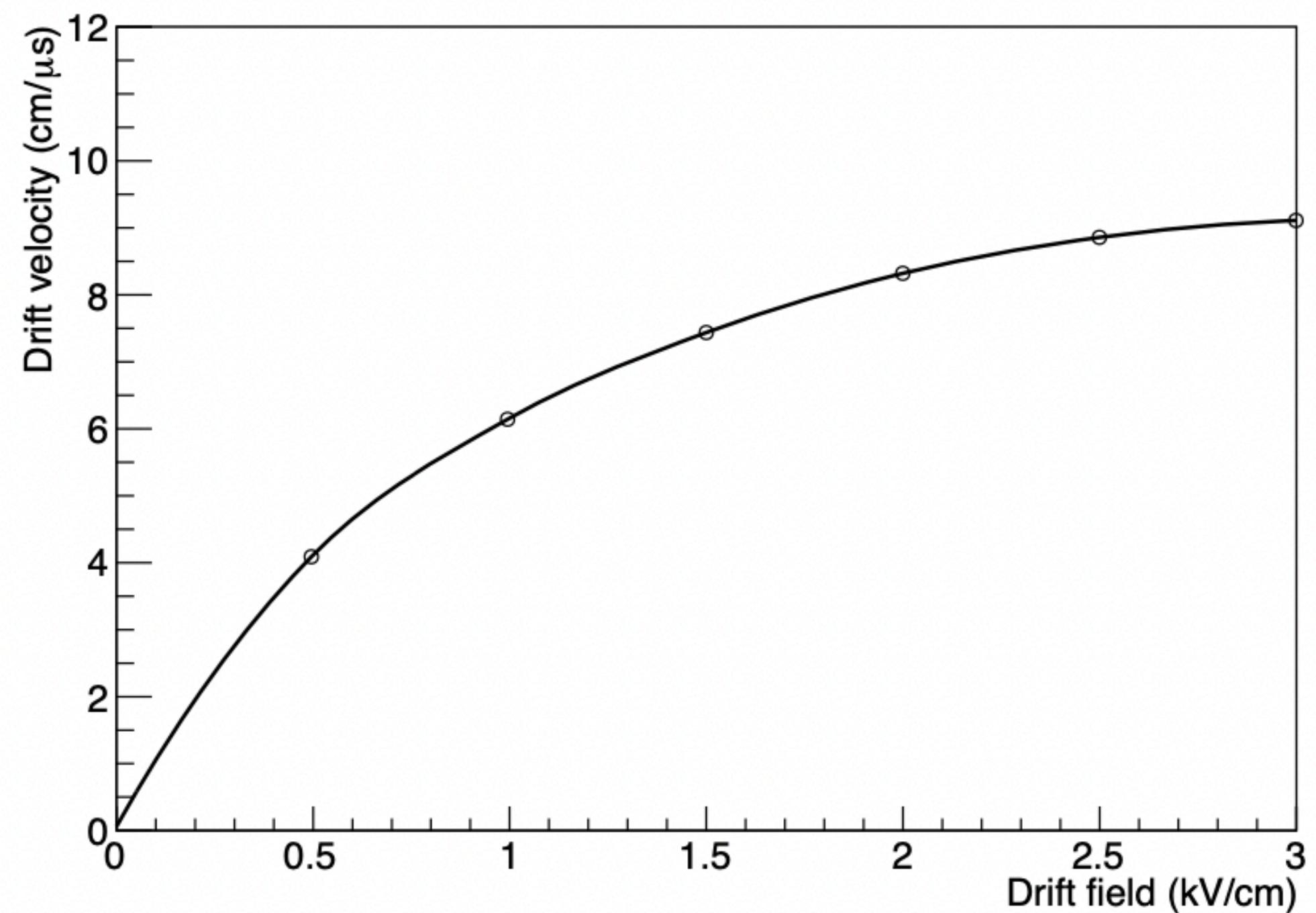
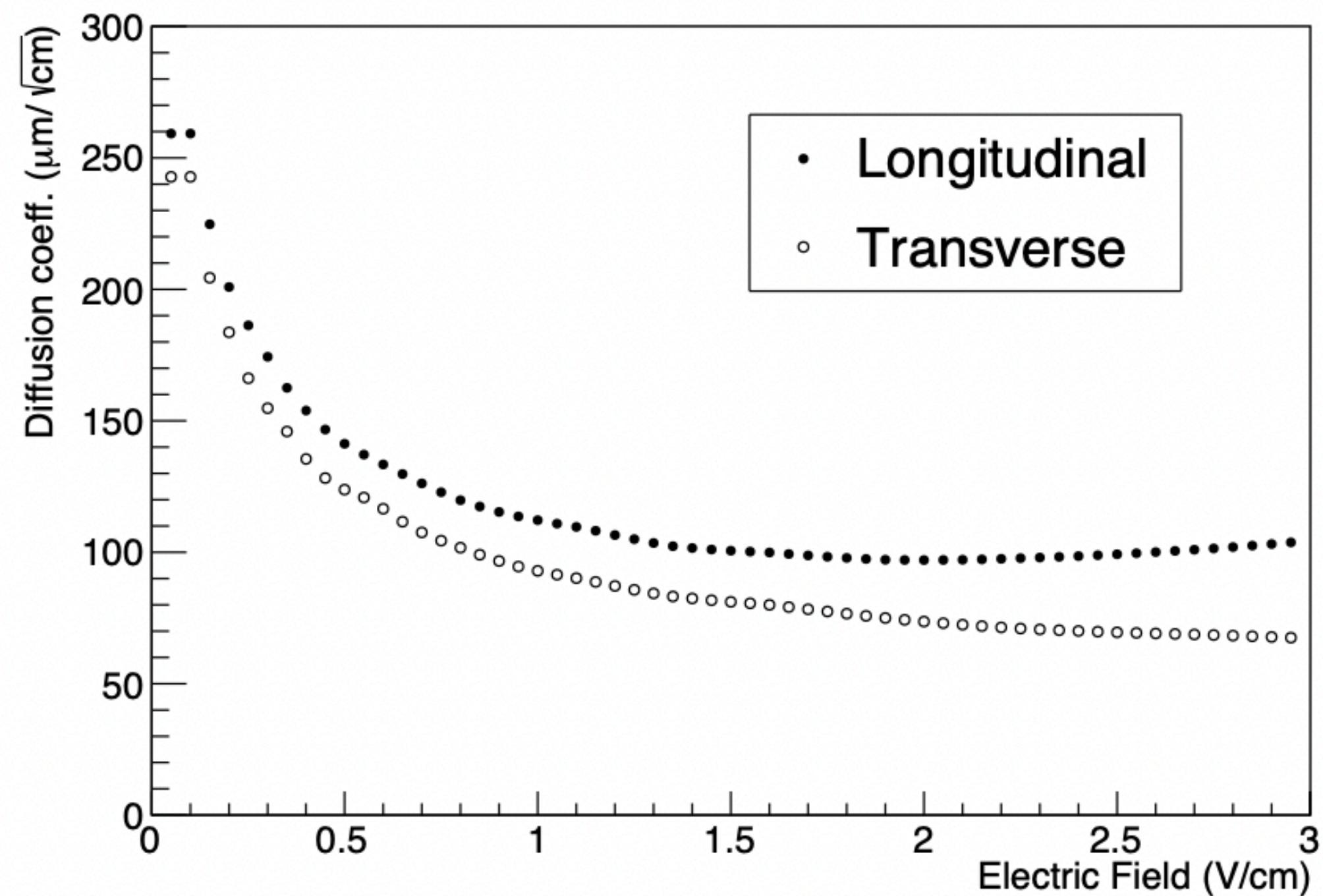
Detector performances: LEMOn

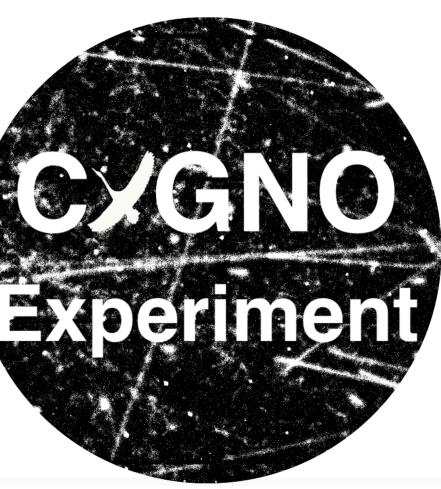
Detector efficiency



Detector performances: LEMOn

Diffusion

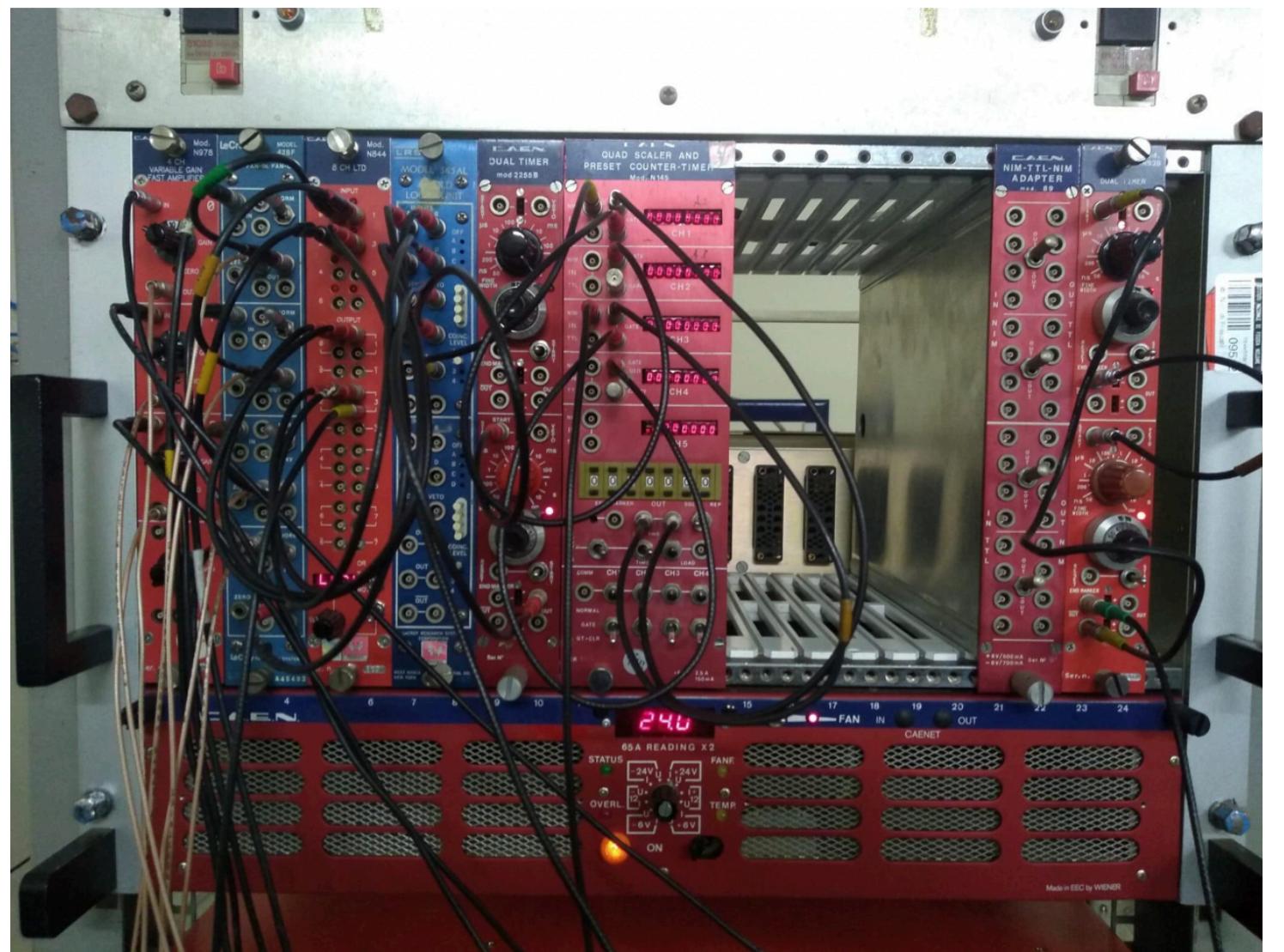




DAQ system for LIME underground

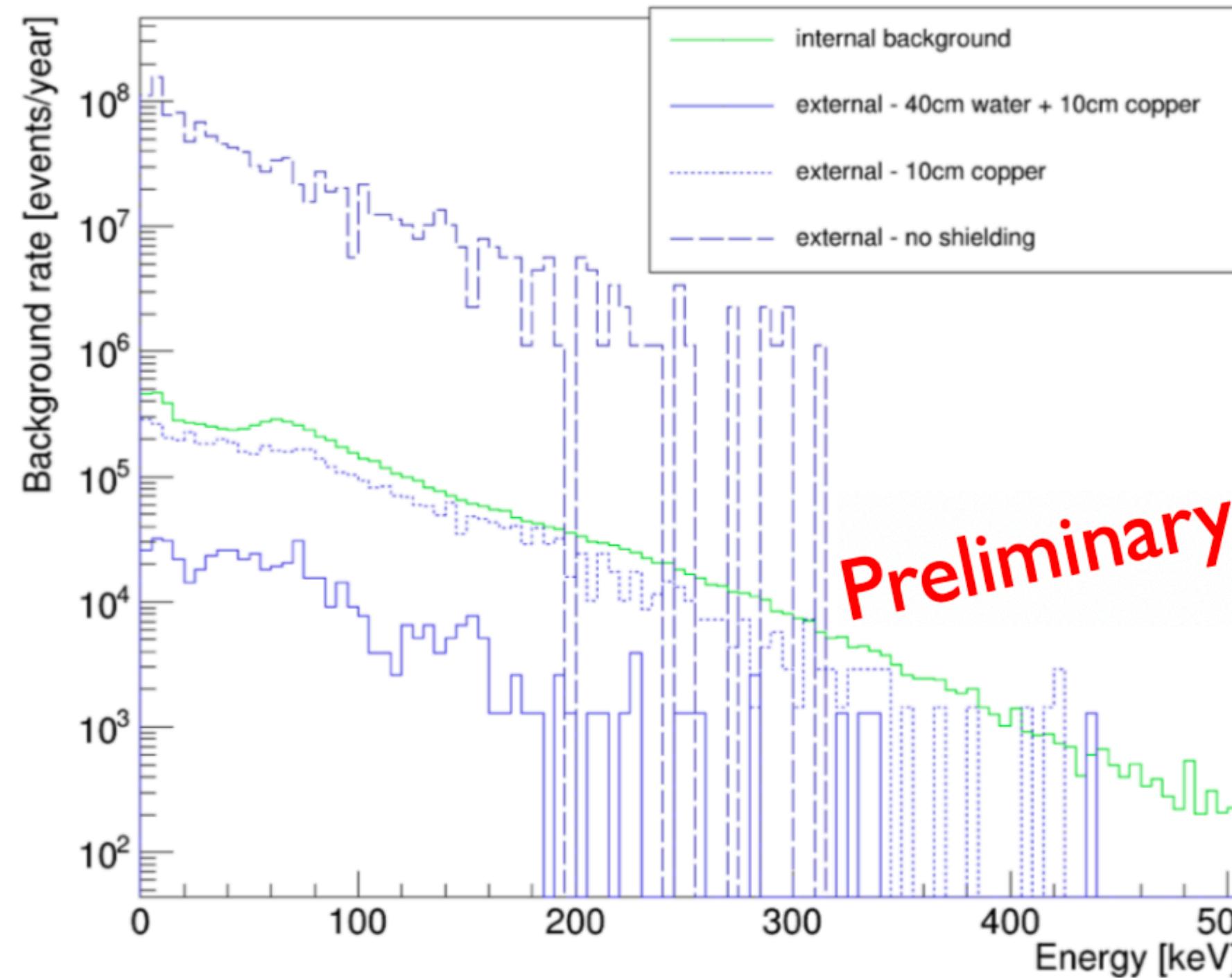
DAQ system based on USB for the camera + NIM and VME modules for analog signals:

- **NIM**: Splitting (N625), Discriminating (N840)
- **VME**: Fast digitization for PMT (V1742), slow digitization for GEMs (V1718, V3718), and logic (V976 or custom trigger module)
- **Two servers Xeon** 16 core 4216, 32 GB RAM+ 2x8 TB HD, NVIDIA RTX5000 24 GB

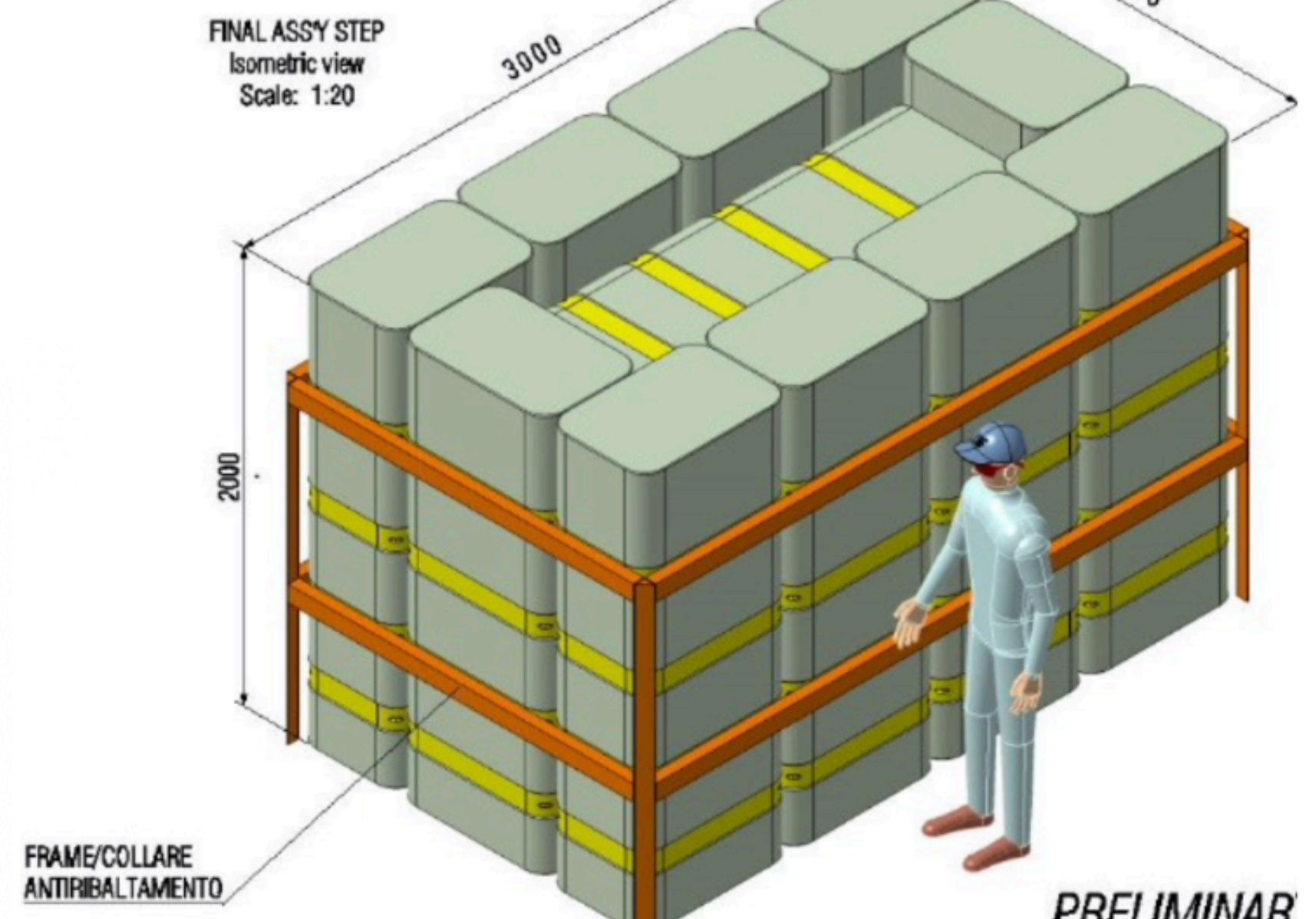
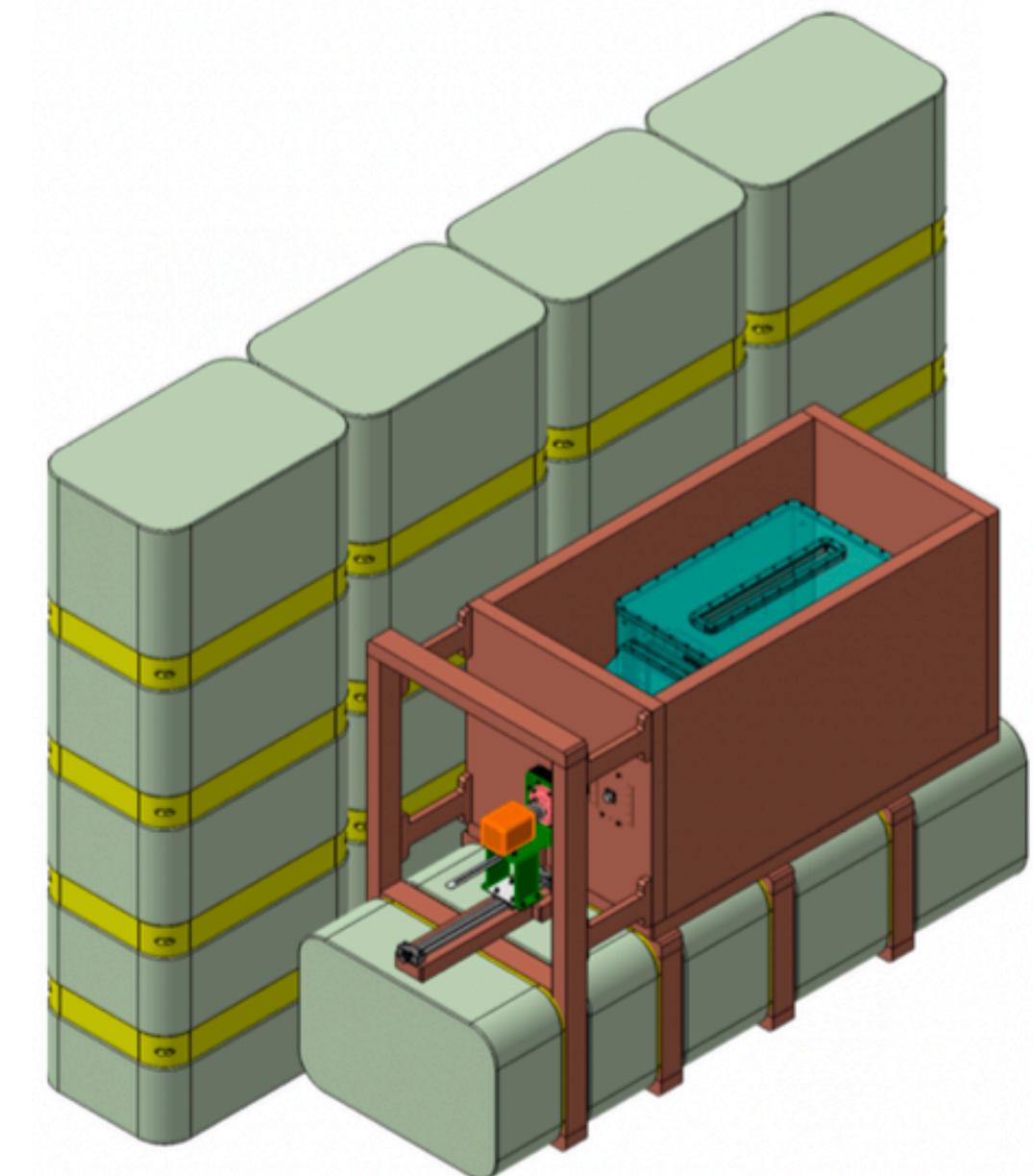


LIME: the underground campaign

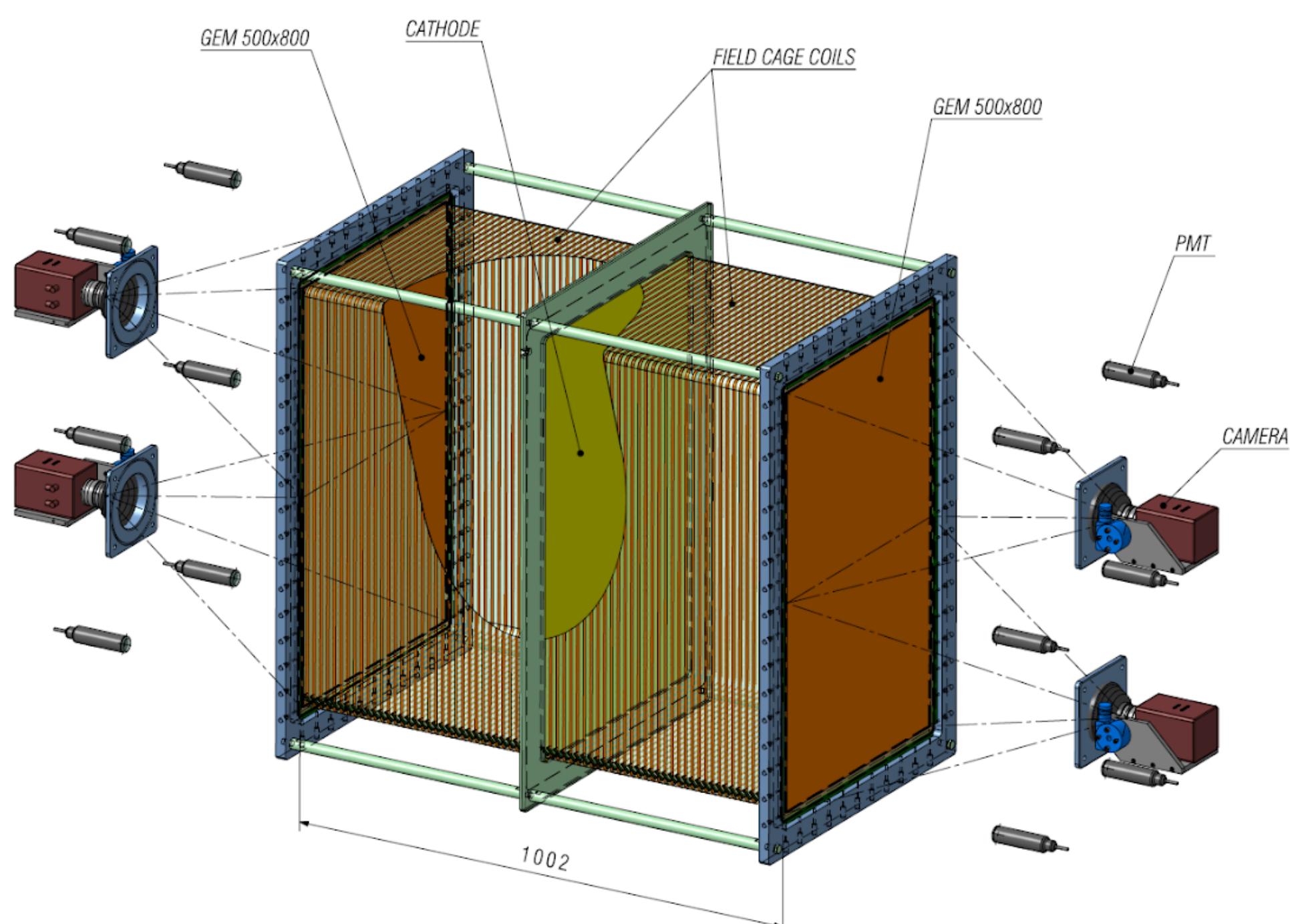
Shielding	Internal [ev/yr] (1-20 keV)	External* [ev/yr] (1-20 keV)
No shield	$1.5344(7) \times 10^6$	$4.061(8) \times 10^8$
5cm copper	$1.5344(7) \times 10^6$	$1.90(2) \times 10^7$
10cm copper	$1.5344(7) \times 10^6$	$1.024(2) \times 10^6$
40cm water + 10cm copper	$1.5344(7) \times 10^6$	$2.46(1) \times 10^5$



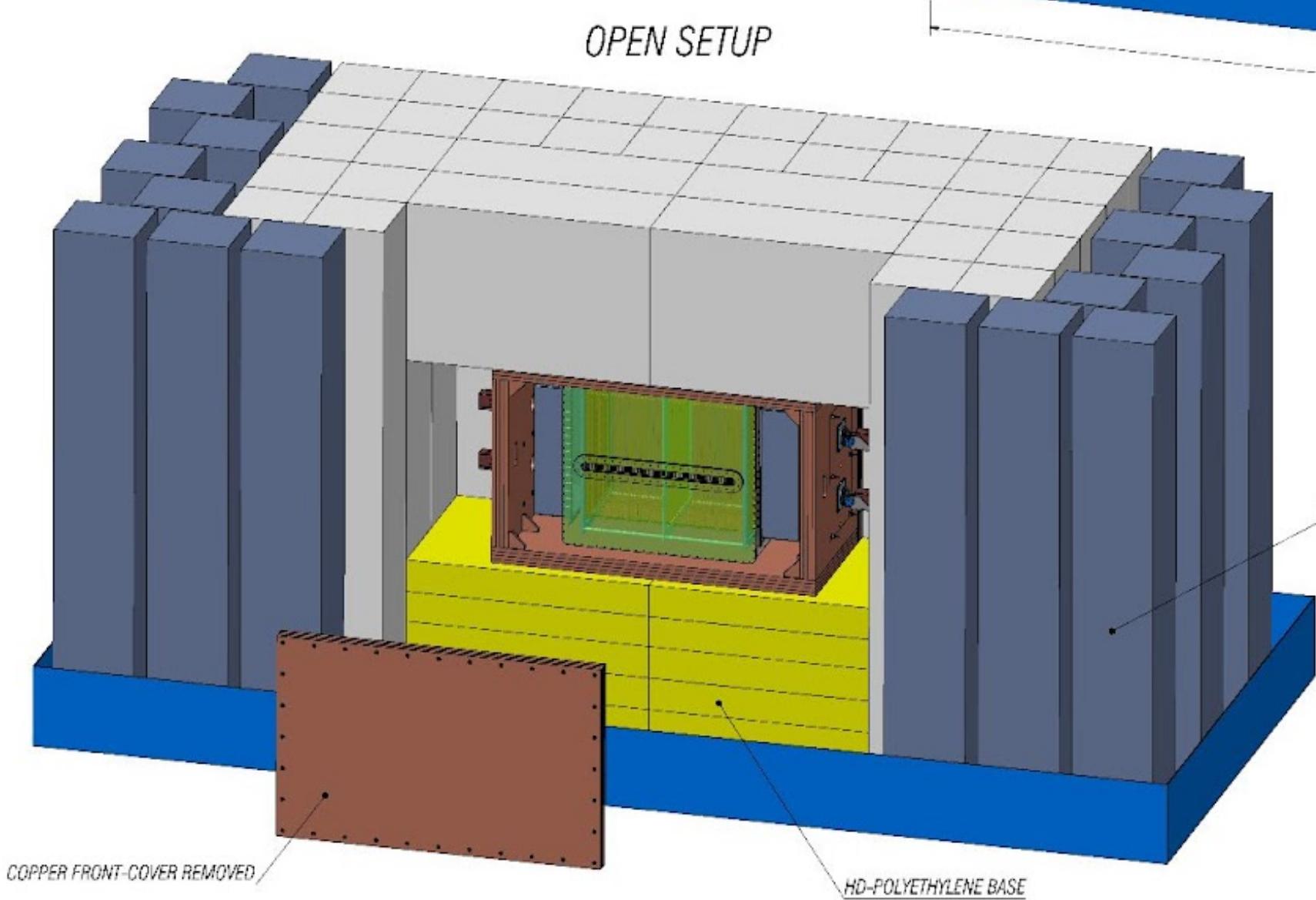
- Shielding: 10 cm of copper + 40 cm of water
- Validate detector MC simulation
- Measure the neutron flux (expected 200 NRs from neutrons in 4 months)



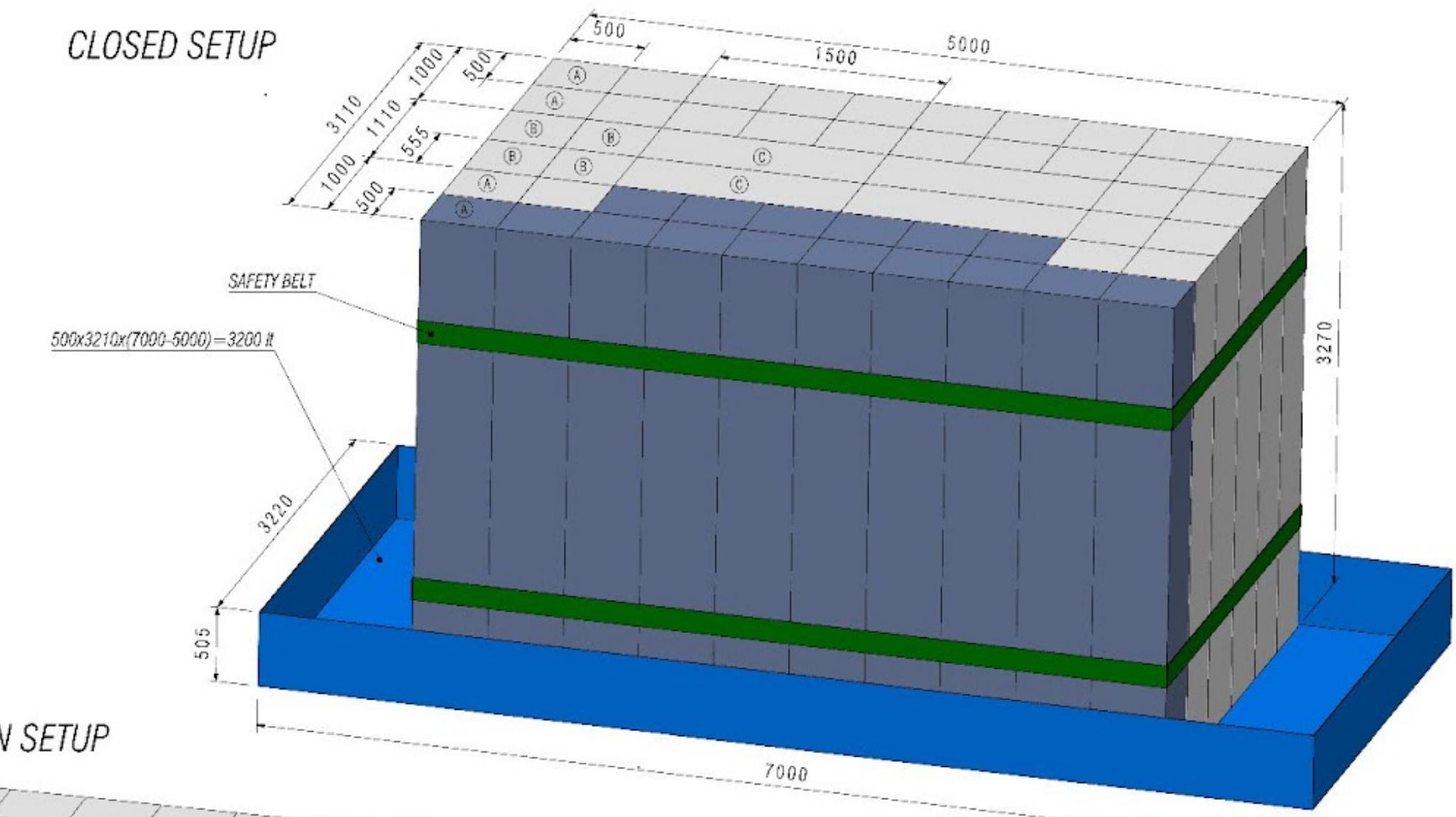
CYGNO PHASE 1: CYGNO_04



PRELIMINARY NEUTRON
SHIELDING SETUP
(1 METER THICKNESS)



CLOSED SETUP



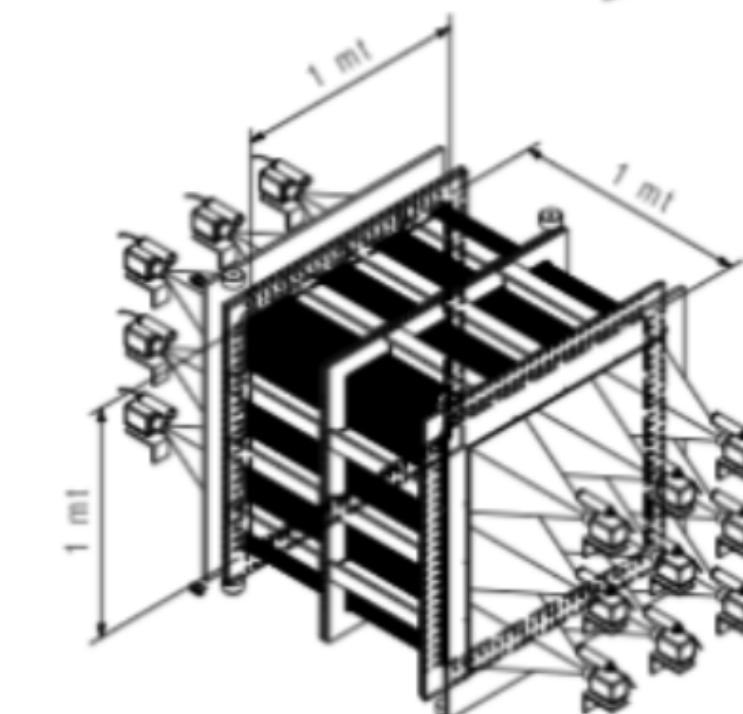
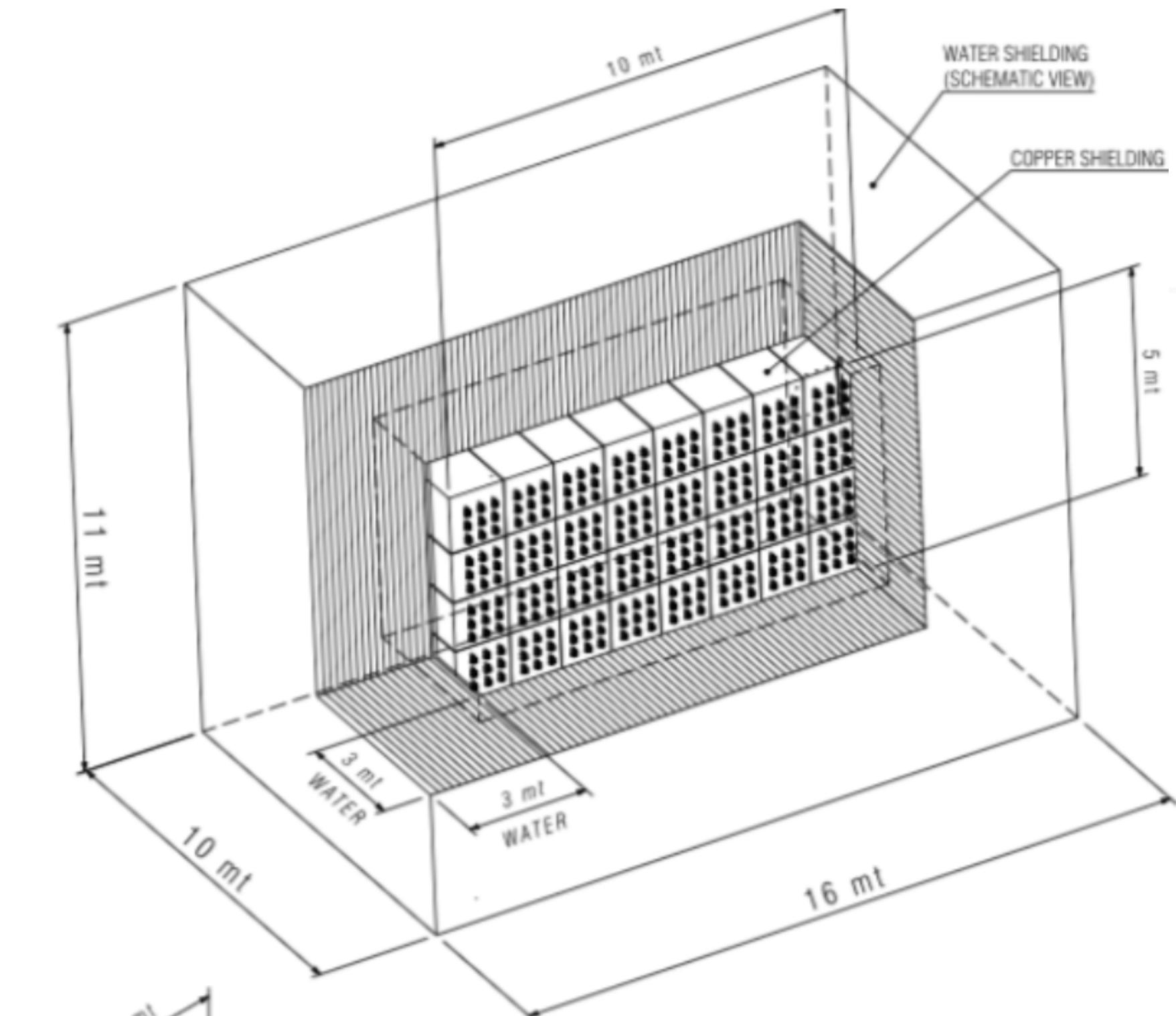
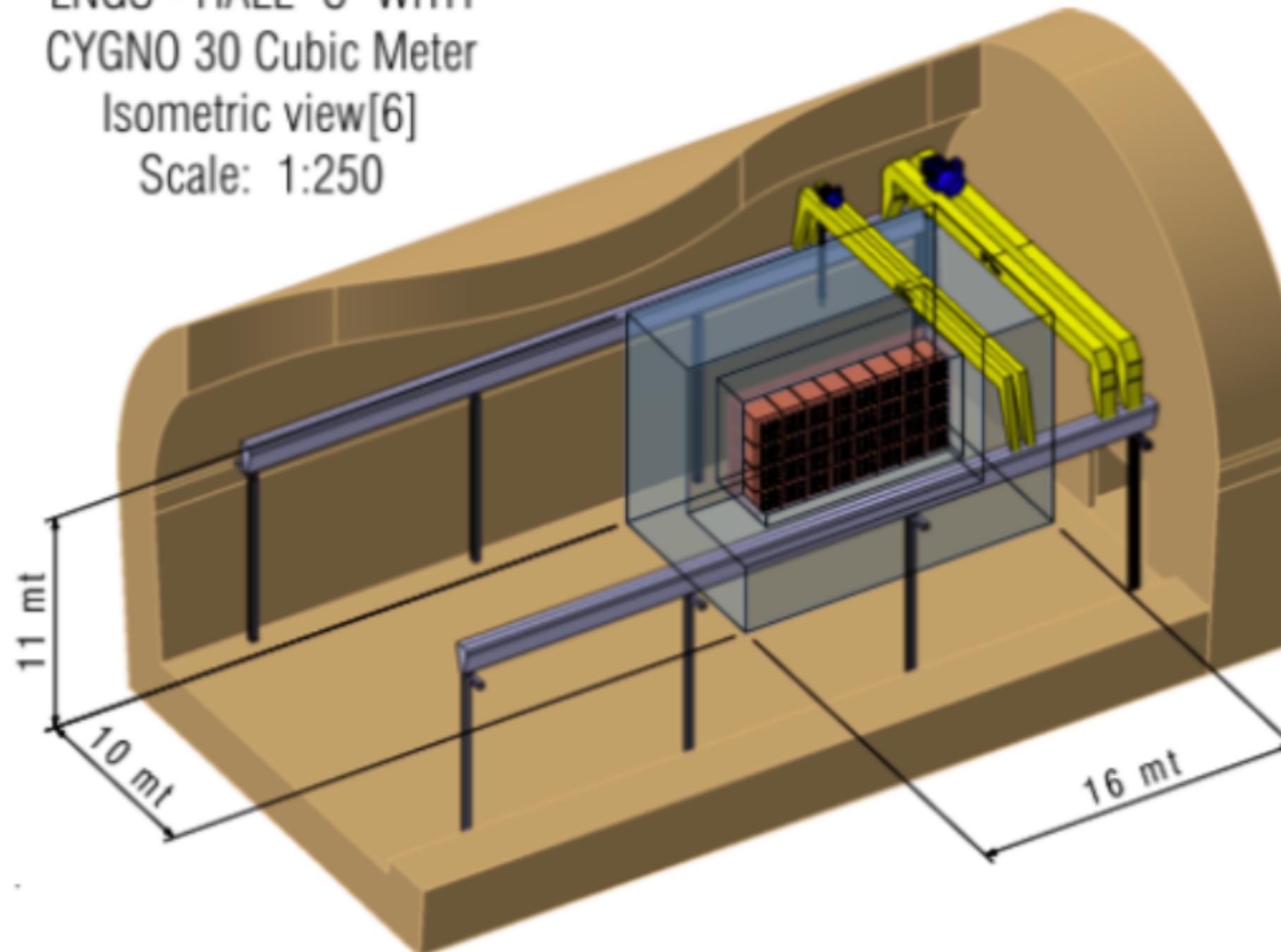
NEUTRON SHIELDING SETUP:

- 40 mc WATER
- 3.4 mc POLYETHYLENE
- N. 40 TANKs (A) -> 500x500x3270 -> 817 lt x 40 = 33 mc (WATER)
- N. 04 TANKs (B) -> 555x500x3270 -> 908 lt x 4 = 3.6 mc (WATER)
- N. 04 TANKs (C) -> 555x1500x1000 -> 833 lt x 4 = 3.4 mc (WATER)
- N. 01 BASE IN POLYETHYLENE -> 1110x3000x1000 = 3330 Kg

MAX WEIGHT OF EMPTY TANK IS 72 Kg

CYGNO PHASE 2: the 30 m^3 experiment

LNGS - HALL "C" WITH
CYGNO 30 Cubic Meter
Isometric view[6]
Scale: 1:250



CYGNO 1 Cubic Meter

Cygnos of 30 cubic meters made by an array of $4 \times 8 = 32$ Cygnos modulus of 1 cubic meter. The setup is completed by a copper (100 mm thick) and water (3m thick) shielding

General tolerance ISO 2768-mK-E		Geometrical tolerance ISO 8015-E		Roughness ISO 1302
SIZE A2				
INRAN	NATIONAL INSTITUTE FOR NUCLEAR PHYSICS			
PROJECTION	Frascati National Lab			
	RESEARCH DIVISION - SEM			
	TOTAL WEIGHT kg	DATE	DRAWN	
		23/11/2020	C. Capoccia	
	SCALE	DATE	CHECKED	
	1:1			
	SHEET	DATE	APPROVED	
	1/1			
CYGNO EXPERIMENT				
CYGNO 30 Mc @ LNGS-U SETUP				
FIRST EVALUATION OF SETUP-ENVELOPPE				
CYGNO 30 Mc				