



#### FLArE: Forward Liquid Argon Experiment for High Energy Neutrino and Dark Matter Searches at LHC



### What are we missing ?



- Most interesting physics is believed to be at high pT, and so are we missing physics in the forward direction?
- The largest flux of high energy light particles, pions, kaons, D-mesons, and neutrinos of all flavors is in the forward direction.
- This could be true of new particles also: dark photons, axion-like particles, millicharged particles, light dark matter, etc.
- The high laboratory energies (>100 GeV), and kinematically focused nature of the particles presents a unique opportunity that should not be missed with the high-luminosity LHC.

#### Forward Physics Facility (FPF) and FLArE

- FPF: Proposal to create forward underground space for experiments during HL-LHC.
- FLArE: a liquid argon time projection chamber (LArTPC) detector for FPF to detect very high-energy neutrinos and search for dark matter at LHC@CERN
- The central goal of FPF is to extend the current LHC forward physics programs into the HL-LHC era with x10-100 exposure



The FPF will be located 620-680 m west of the ATLAS IP along the line of sight (LOS). Also shown is the location of FASER and FASERv, which are also located along the LOS, but 480 m east of the ATLAS IP

#### **Physics Goals**

Neutrino physics

- Measure the flux of tau neutrinos
- Limit the oscillations of tau neutrinos into other neutrinos over the parameter range defined by the energy spectrum and distance
- Determine the cross section of neutrino interactions in the energy range of hundreds of GeV to few TeV
- QCD physics with far forward neutrinos

Dark matter search sector

- Detection of dark photon and other dark sector particle decays in the detector volume
- Detection of light dark matter scattering in detector
- Search for milli-charged dark matter particles
- Search for dark tridents



#### **Proposed Detectors for FPF**

Experiment	Science Priority	Technology	
Faser 2	Long-live neutral particles decay	Large decay volume (super-conducting) magnetic spectrometer	
FASERnu2	SERnu2 Neutrino Interactions Tungsten/Emulsion 20 tons. Veto and interface tracker for muons		
AdvSNDNeutrino Interactions on/off axisHybrid electronic ar		Hybrid electronic and tungsten/emulsion detector with had. cal.	
FORMOSA	FORMOSA Milicharged particles Scintillation bars with photomultiplier readout.		
FLARE         DM scattering and neutrino interactions         Liquid Argon <sup>-1</sup>		Liquid Argon TPC 10-20 tons	





# **Neutrino physics**



- The current data from accelerators ends around 300 GeV. FPF would provide data that fills in the gap between accelerators and atmospheric neutrinos.
- There are three proposed detectors at 10 ton each: FASERv2 (emulsion), SND(TBD), and FLArE.
- Total rate will be  $\sim 100$ k electron neutrinos,  $\sim 1$ M muon, and  $\sim few$  thousand tau neutrino events.

## Light Dark Matter scattering

#### **Elastic scattering from electrons or nuclei**

- Mass of the  $\chi$  alters the kinematics of the outgoing electron or nucleus.
- Signal is at low energy (~1 GeV)
- Background is from neutrino interactions and muons.
- The sensitivity plot assumes reasonable cuts for background suppression
- Makes use of the huge flux of mesons for this *direct detection* technique to get to the relic density target.





## **Milicharged particles**

- These emerge in models with massless dark photons which couple weakly to dark particles.
- The idea is to see them using dE/dx in a low noise detector.
- Deep bars of scintillator coupled to PMTs: milliQan (central location) and FORMOSA (at FPF)
- The FPF sensitivity assumes high efficiency light sensitivity in 1 meter bars of plastic scintillator with coincidence of 4.
- How can we do better ? Is it possible to use a liquid argon TPC with very good single electron sensitivity (with 2phase)



Foroughi-Abari, Kling, Tsai (2020)

# **Experimental conditions (without sweeping magnet, very preliminary, work in progress)**



- This rate will be lower at 612 m.
- Both charged and neutral hadron interactions present significant background.
- Total neutrino interaction rate normalized to per ton per fb<sup>-1</sup>
- Observed nu rate: ~45/ton/fb-1 at 480 m

Minimum distance	612 m	
Total Lumi/max lumi	3000/fb;5x10 <sup>34</sup> /cm2/sec	
Lumi per day	~1 /fb assuming 10 year running	
pseudorapidity coverage	>6.4, (~5.4-6.0 for off-axis)	
Mu+/Mu- flux > 10 (100) GeV	1.5/0.93 (0.94/0.39) 10 <sup>4</sup> /cm <sup>2</sup> /fb <sup>-1</sup>	
track density (from data)	1.7x 10 <sup>4</sup> /cm²/fb <sup>-1</sup>	
max track density per sec (per crossing)	0.85/cm <sup>2</sup> /sec (2x10 <sup>-8</sup> /cm2/crossing)	
Tracks in detector/1 ms	8.5/m^2/1msec	
Neutral hadron flux > 10 GeV (10 <sup>-4</sup> of muons)	~3 /cm²/fb <sup>-1</sup>	
Total neutrino rate (all flavors)	~50/ton/fb <sup>-1</sup>	

arxiv 2105.06197

#### Sweeper Magnet: Ongoing Studies

- Preliminary design of sweeper magnet by TE-MSC
  - Based on permanent magnet to avoid power converter in radiation area
  - Consider 7m long (20x20cm<sup>2</sup> in transverse plane) magnet, 7Tm bending power
- To install such a magnet would require some modifications to cryogenic lines in relevant area
  - Possibility of modifications to be investigated with LHC cryo
  - Integration/installation aspects to be studied
- FLUKA and BDSIM studies ongoing to assess effectiveness of such a magnet in reducing the muon background in the FPF

Consider to add a sweeper magnet upstream of the FPF (e.g. where the LOS leaves the LHC beampipe) to deflect muons from the on-axis neutrino detectors.



## **Cryostat options for FLArE**

Very important for space considerations.



- Space in FPF hall currently is limited to 3.5 m X 3.5 m X 9.6 m for FLArE.
- 80 cm GTT membrane occupies 1.6 m out of 3.5 m. More space might be needed for corrugations.
- GTT is easy to install, DUNE ND-LAr design has installation from top, this would also simplify things.
- Further Engineering will be needed, but we can settle on GTT membrane for now.



#### FLARE Detector Preliminary Sketch (without HadCal and Muon Catcher)





Simulations have confirmed that these dimensions allow reasonable containment of neutrino events in LAr and total energy measurement.

They also fit within the cryostat allowed transverse space.



either 2 X 7 vertical modules with 0.45 m gap or 3 x 7 vertical modules or with 0.3 m gap

#### **Tau Neutrino event simulation in FLArE**





First conclusions from simulations:

 Need 1.8 X 1.8 to contain transverse events in fiducial of 1 m x 1 m x 7 m.

2) Need Hadronic calorimeter to contain showers that start downstream.

**3)** Even a modest resolution is sufficient.

4) excellent muon identification could result in quick selection of  $v_{\tau}$  events.

5) Studies: can we combine spatial resolution in drift dimension with kinematics to get good S/N ?

#### Identify $v_{\tau}$ CC at FLArE

TABLE I. Dominant decay modes of  $\tau^-$ . All decays involving kaons, as well as other subdominant decays, are in the "Other" category.

Dec	ay mode	Branching ratio	
Lep	tonic	35.2%	
$e^{-\overline{\nu}}$	$e \nu_{\tau}$	17.8%	
$\mu^- \bar{ u}$	$_{\mu} u_{ au}$	17.4%	
Had	ronic	64.8%	
$\pi^{-}\pi$	$^{.0}\nu_{\tau}$	25.5%	
$\pi^{-}\nu$	τ	10.8%	
$\pi^{-}\pi$	$^{0}\pi^{0}\nu_{\tau}$	9.3%	
$\pi^{-}\pi$	$\pi^{+}\nu_{\tau}$	9.0%	
$\pi^{-}\pi$	$\pi^{+}\pi^{0}\nu_{\tau}$	4.5%	
Oth	er	5.7%	

 $v_{\tau}$  CC,  $\tau \rightarrow \mu$  and  $v_{\mu}$  CC are distinct from other channels in dE/dx and energy deposit



20.0

 $v_{\tau}$  CC,  $\tau \rightarrow \mu$  have more neutrinos in the final state than  $\nu_{\mu}$ CC, thus more missing momentum in the transverse plane



A BDT shows promising results to select  $v_{\tau}$  CC,  $\tau \rightarrow \mu$ from backgrounds, working on other  $\tau$  decay modes



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## Summary

- A forward physics facility FPF is being considered at CERN for neutrino and dark matter physics.
- The physics interest is
  - Neutrino cross sections in the 1 TeV range: ~20-50 events/ton/day
  - Tau neutrino flux and associated heavy flavor physics:  $\sim 0.1-0.2$  events/ton/day
  - Light dark matter search with decays and interactions.
- Liquid Argon detector FLArE for FPF is being considered
- Detector capability, event rate and backgrounds of FLArE are preliminarily studied, showing that a LAr detector is feasible and groundbreaking

Thank you!

# Backup

## Current forward physics programs with Run 3

- 4 experiments in progress for LHCrun3 for 150fb<sup>-1</sup> 2022-24.
- FASER (March 2019), Magnetic spectrometer for neutral decays.
- FASERv (Dec 2019), Emulsion/tungsten detector (~1 ton)
- SND@LHC (Mar. 2021) Hybrid Emulsion/active target. (~1 ton)
- Also MilliQan located near CMS (not forward); scintillation bars to see milli-charged particles.
- This program will provide excellent experience for the FPF.



### FASERv pilot run

#### First collider neutrinos detected at 2.7 sigma.

- 2018 pilot emulsion detector with 11 kg was deployed for 12.2/fb
- May 2021, announced 6 candidates with 12 backgrounds.
- Same stack able to measure the muon rate at that location.
- muon and neutrino rates in rough agreement with expectations.
- <u>https://arxiv.org/abs/2105.06197</u>





FIG. 6. The BDT outputs of the observed neutral vertices, and the expected signal and background distributions (stacked) fitted to data. Higher BDT output values are associated with neutrino-like vertex features.



FIG. 1. Structure of the pilot emulsion detector. Metallic plates (1-mm-thick lead or 0.5-mm-thick tungsten) are interleaved with 0.3-mm-thick emulsion films. Only a schematic slice of the detector is depicted.

#### Neutrinos at FPF

# Uncertainties are large. 2105.08270 (Kling) is standard simulations. 2002.03012 (Bai et al.) and 2112.11605 is deep analysis of the tau neutrino flux.



New HL-LHC geometry BDSIM simulation. Where to place a sweeper magnet ? Any Placement requires movement of cryogenic lines and other services. But so far possible. L. Nevay



The MU+ and MU- have very different spatial distributions. And the rate is actually higher away from the LOS. We have to assume that this will be solved with ultimate rate of << 1 Hz/cm2 at the FPF. it will need more people !

#### Neutrino event rates (with large uncertainties)

evts/ton /fb-1	ν	$\overline{\nu}$	Total	$10^{14} \boxed{ \begin{array}{c} \\ \end{array} \begin{array}{c} SIBYLL 2.3.d \\ \end{array} \begin{array}{c} \\ \end{array} \begin{array}{c} \\ \end{array} \begin{array}{c} \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\$		
е	2.1	1.0	3.1			
ти	15	5	20			
tau	0.1	0.05	0.15	$10^{10} \downarrow flare (10 \text{ tons})$ $FPF paper$ $10^{3} 10^{3} 10^{3}$		
_		1	•	Neutrino Energy [GeV]		

# During HL-LHC fb<sup>-1</sup> is approximately per day.

Mean energy of interactions is ~500 GeV

#### QCD interest

#### Neutrino interactions neutrino-ion collisions at $\sqrt{s} \approx 50 GeV$



- Forward hadron production, instrinsic charm (large-x), ultra-small x proton structure
- Extremely well motivated by the astrophysics UHE cosmic rays.

#### Preliminary Parameters for the FLArE Detector

Cryostat outer	35mX35mX96m	Membrane
Insulation thickness	0.8 m	including corrugations
Detector dimension	1.8 m X 1.8m x 7 m	good for >90 % containment
Fiducial volume	1 m x 1m x 7 m (10 tons)	Length may be adjusted later
TPC Modules	2 X 7 or 3 X 7	Keep two options
Module opt1 dimensions	0.9 m (W) X 1.8 m (H) X 1 m (L)	Central cathode: gap: 0.45 m
Module opt2 dimensions	0.6 m (W) X 1.8 m (H) X 1 m (L)	gap: 0.3 m
Anode design fiducial region	5 mm x 5 mm for 1 m x 1 m	80000 chan/mod
Anode design containment region	10 mm x 10 mm for 0.8 m x 1 m	16000 chan/mod
photon sensor	Bare SiPM or X-ARAPUCA	~50 chan/mod
Downstream cryo wall	80 cm	Can it be thinned down
HADCAL	2 m x 2 m x (5 cm Fe + 1+1 cm scint, 15 layers) x (1.05 m)	Optimize for resolution
Murange	• 2 m x 2 m x (16 cm Fe + 1 + 1 cm scint, 2 layers) x (0.36 m)	Increase to 1 m to get clean muID

#### Muon simulation in FLArE



- •Muon flux above 1 Hz/cm2 presents a difficult problem for all detectors.
- •For Liquid argon TPC, the flux also presents a space charge problem for large gaps.
- •Showering muons will also present a trigger problem since if the incoming muon is missed the event will look like a neutrino.