

Flavor Violating Axions in the Early Universe

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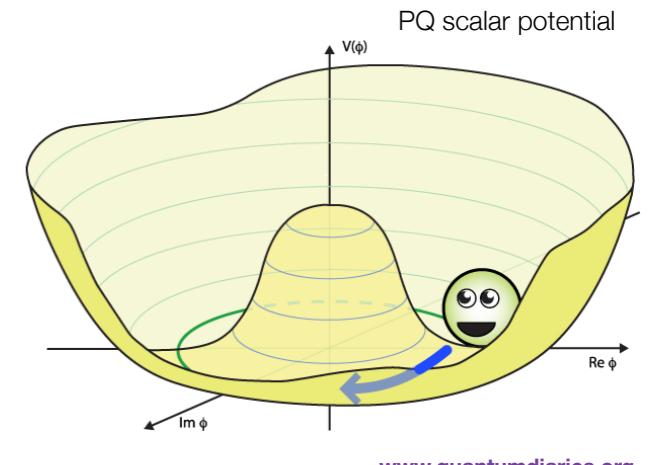
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Outline

- 01 Introduction** • Overview of axion physics
- 02 Flavor violating axion** • Flavor violating interactions
- 03 Hot flavor violating axion relic** • Production rate
• Cosmological constraints
- 04 Conclusion**

Axion

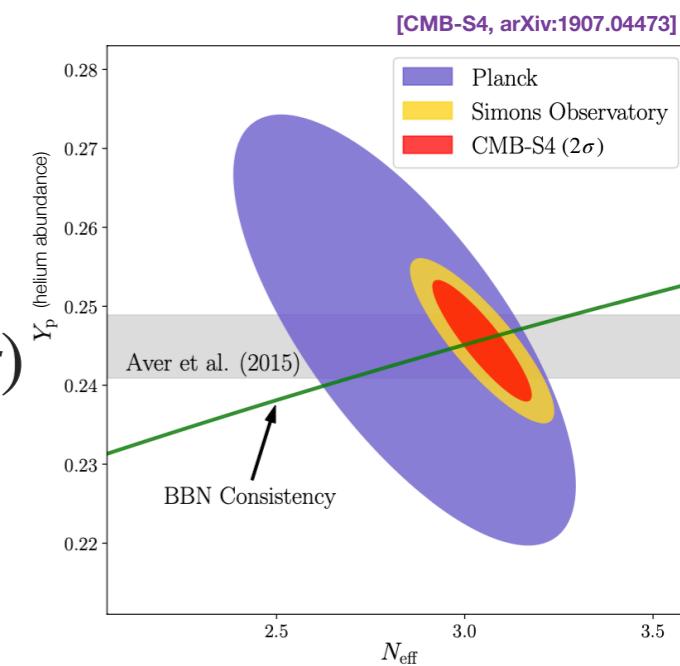
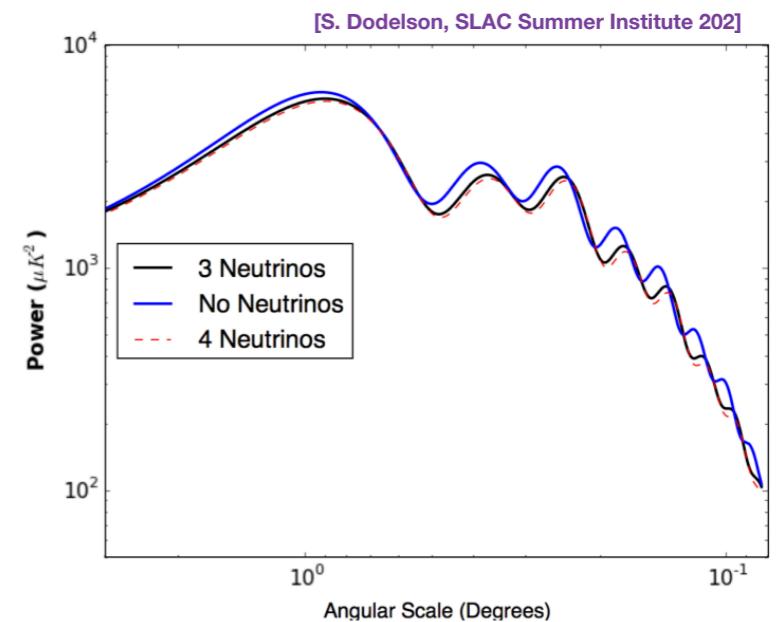
- Solution for the strong CP problem $\bar{\theta}_S G_{\mu\nu} \tilde{G}^{\mu\nu} \lesssim 10^{-10}$



- Global Peccei-Quinn symmetry - Pseudo Nambu-Goldstone boson after spontaneous breaking at $\sim f_a$
- Mass and coupling proportional to f_a^{-1} ; if only coupling, Axion-like particle (ALP)

Axionic dark radiation?

- Parametrized by $\Delta N_{\text{eff}} \equiv N_{\text{eff}} - N_{\text{eff}}^{\text{SM}}$
- Effect on CMB spectrum
(e.g. damping tails, changing the scale of matter-radiation equality, location of the first acoustic peak, etc)
- Current bound from Planck $N_{\text{eff}} = 2.99 \pm 0.17$
- Future sensitivity from CMB-S4 $\Delta N_{\text{eff}} = 0.03 \text{ (} 0.06 \text{)} \text{ for } 1\sigma \text{ (} 2\sigma \text{)}$



Boltzmann equation

- Quantitative tool to track the abundance

d=4 axion production rate

$$\frac{dn_a}{dt} + 3Hn_a = \gamma_a \left(1 - \frac{n_a}{n_a^{\text{eq}}} \right)$$

$$Y_a = \frac{n_a}{s}, \quad x = \frac{M}{T}$$

$$\rightarrow \frac{dY_a}{d \log x} = \left(1 - \frac{1}{3} \frac{d \log g_{*s}}{d \log x} \right) \frac{\gamma_a}{sH} \left(1 - \frac{Y_a}{Y_a^{\text{eq}}} \right)$$

$$\Rightarrow \Delta N_{\text{eff}} = \frac{4}{7} \left(\frac{11}{4} \right)^{4/3} \left[\frac{\frac{2\pi^4}{45\xi(3)} g_{*s}^{\text{SM}} Y_a^\infty}{1 - \frac{2\pi^4}{45\xi(3)} Y_a^\infty} \right]$$

Axion interactions

[H. Gerogi et al, 86]

- Lagrangian with the non-linearly realized PQ symmetry (i.e, $a \rightarrow a + \text{const}$)

$$\mathcal{L}_a \supset \frac{\partial_\mu a}{2f_a} \sum_{f_i f_j} \bar{f}_i \gamma^\mu \left(c_{f_i f_j}^V + c_{f_i f_j}^A \gamma^5 \right) f_j$$

\Rightarrow PQ current

- Field redefinition by axionic rotation \rightleftharpoons linear realized PQ symmetric (*non-der.*) basis
relying on div. of PQ currents $\partial_\mu j_{\text{PQ}}^\mu$
- In general, flavor dependent & violating structure

$$F_{f_i f_j}^\alpha \equiv \frac{2f_a}{c_{f_i f_j}^\alpha} \quad \begin{cases} \{\alpha = V, A\} \\ \{i \neq j\} \end{cases}$$

Flavor-violating couplings $F_{f_i \neq f_j}^\alpha$

♦ Explicit charge assignments

e.g., through the Froggatt-Nielsen mechanism

“Flaxion” [Y. Ema et al, 2016]

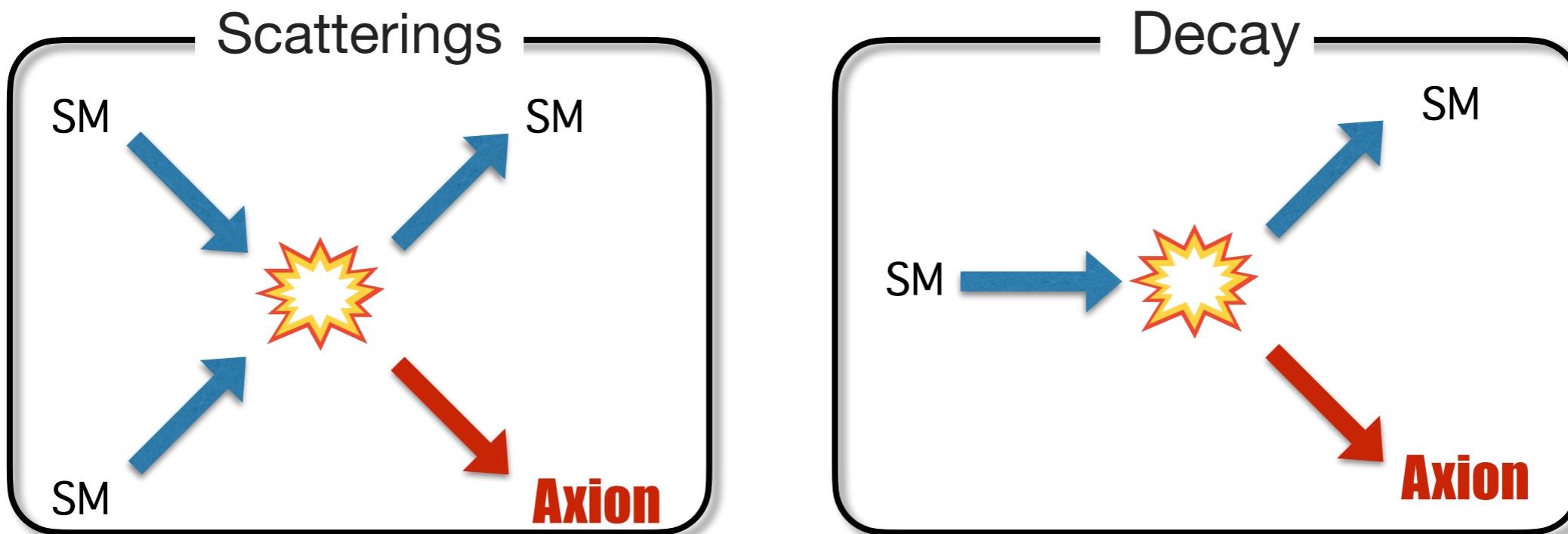
“Axiflaviton” [L. Calibbi et al, 2016]

♦ Radiatively induced

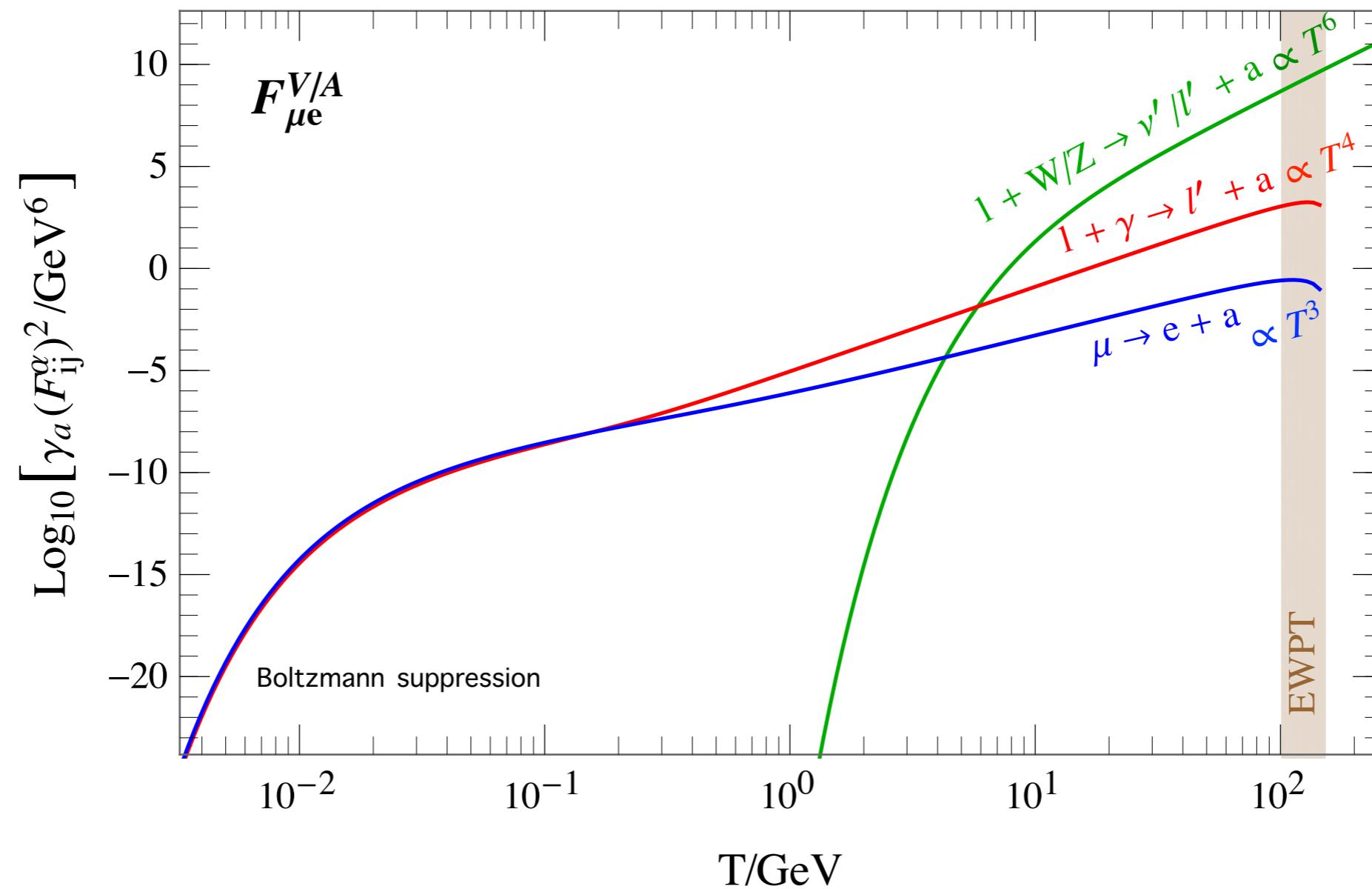
[K. Choi, S. Im, C. Park, SY, 2017]

[M. Chala, 2020]

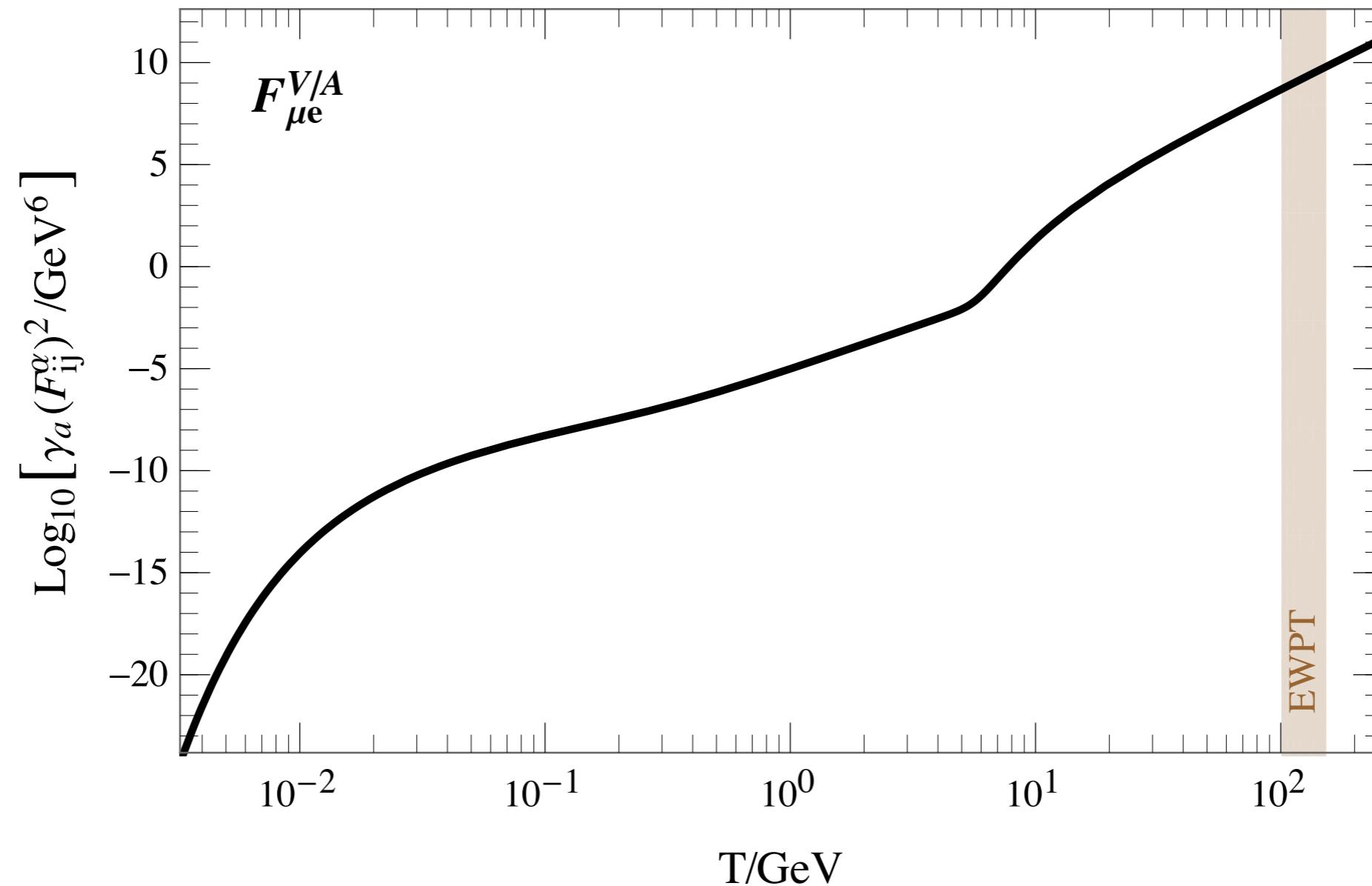
[K. Choi et al, 2021]



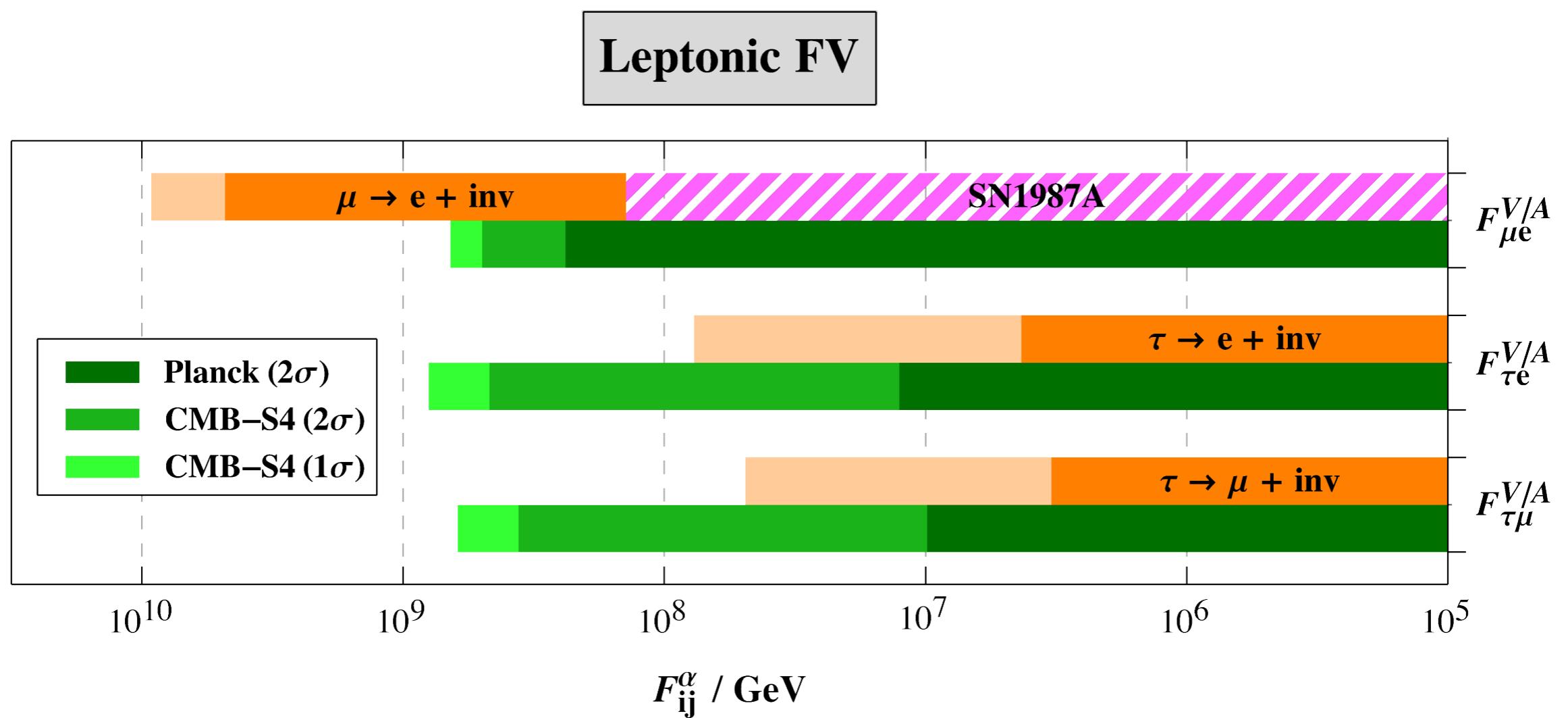
Leptonic flavor violation



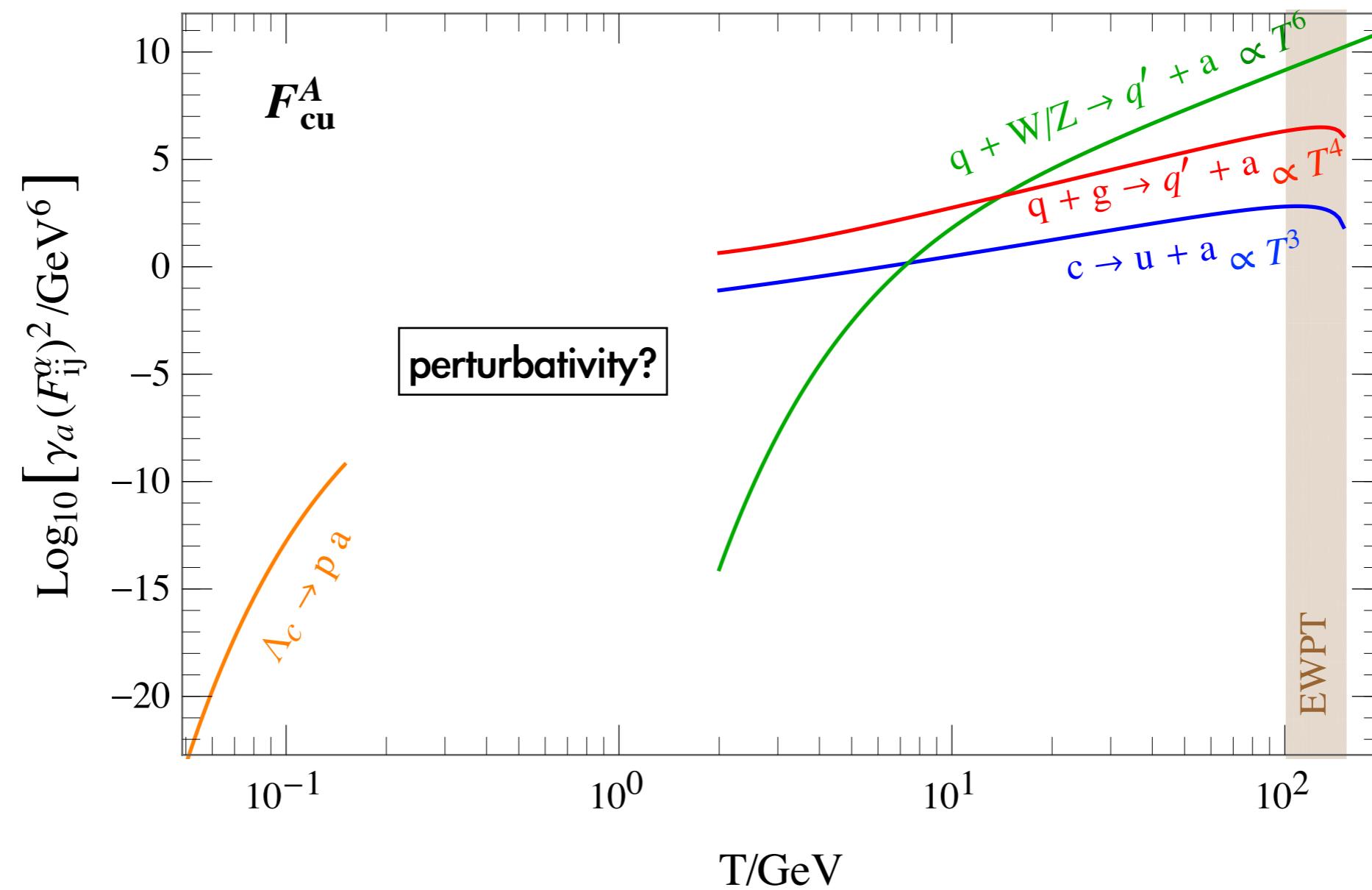
Leptonic flavor violation



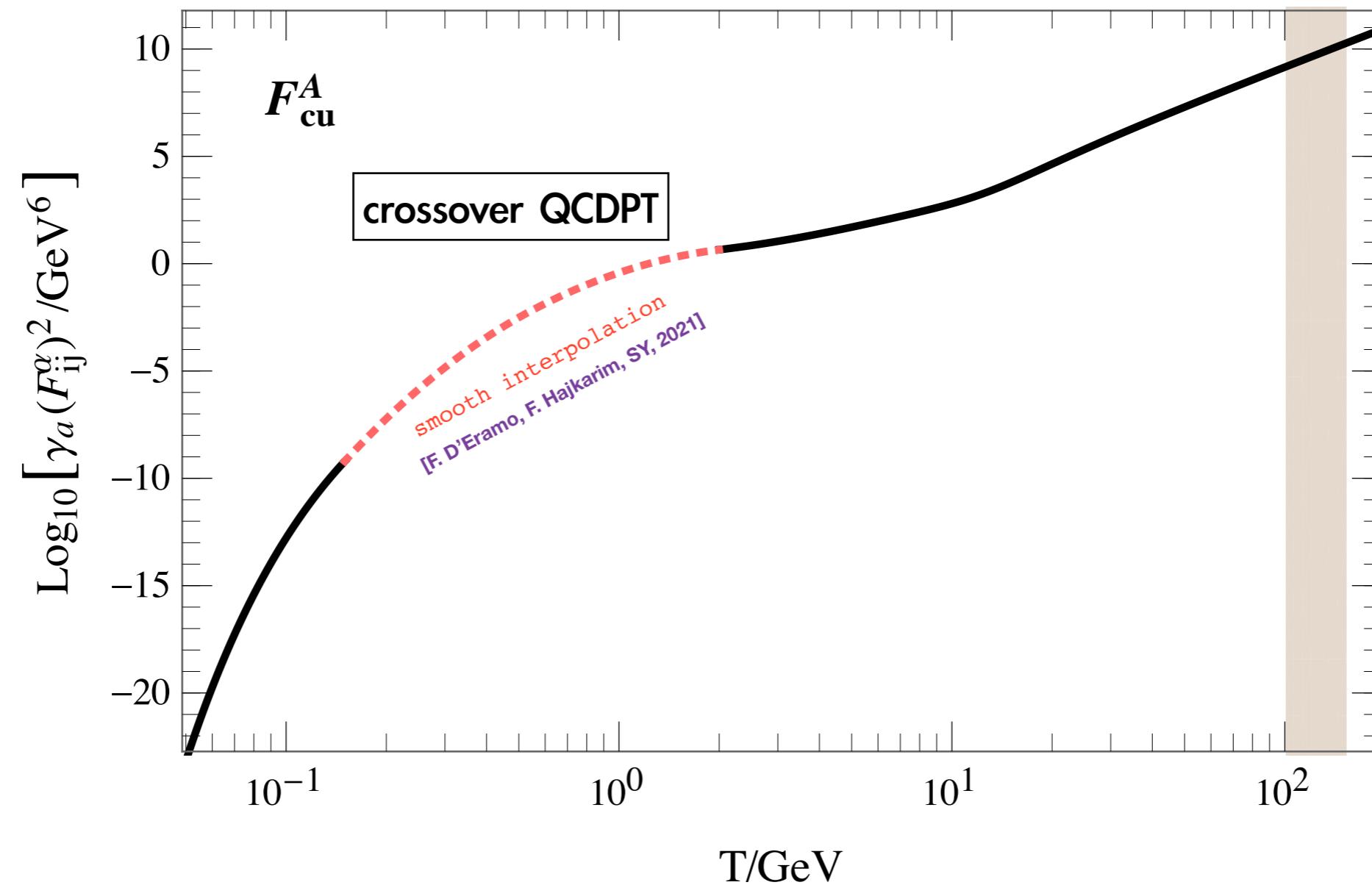
Leptonic flavor violation



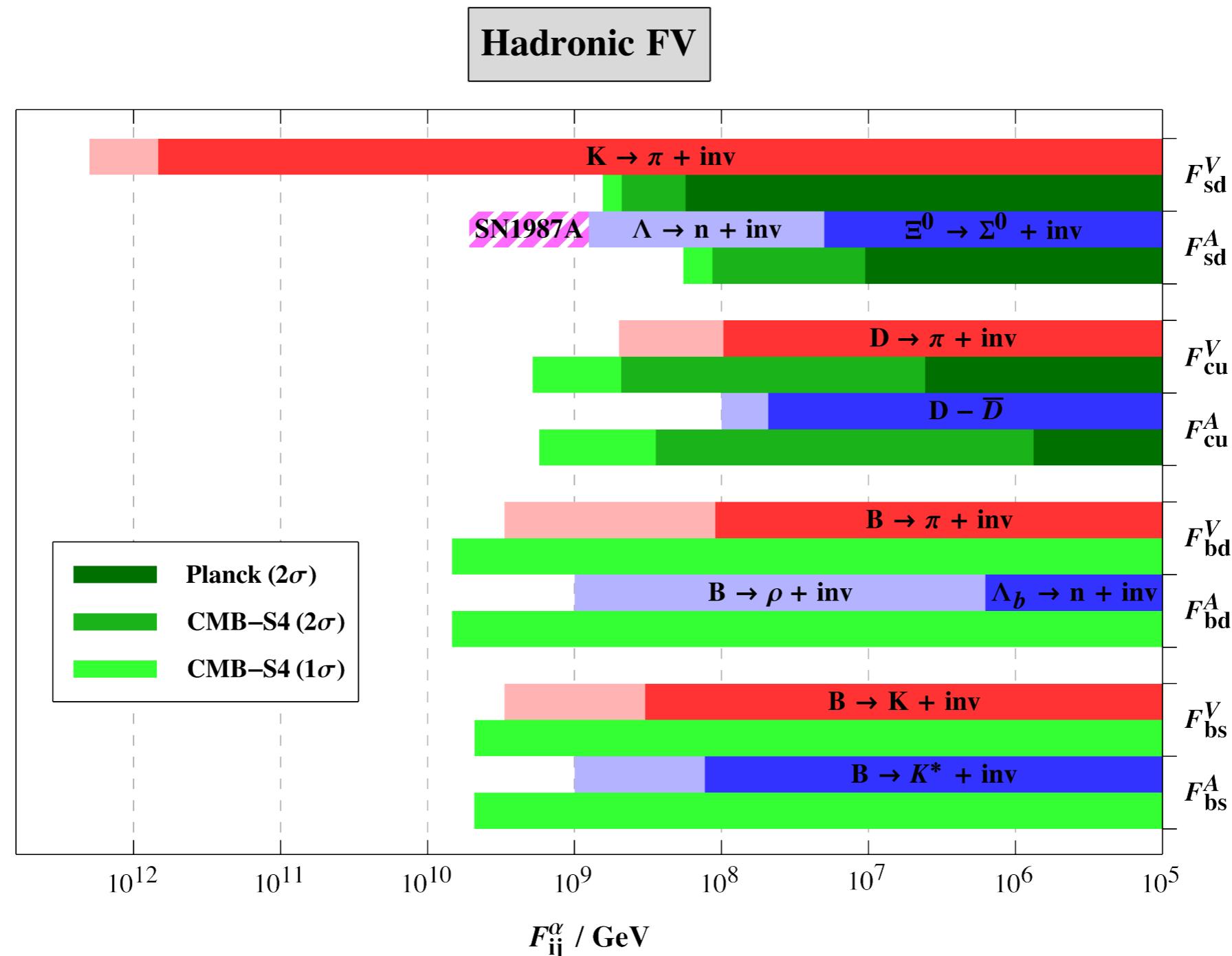
Hadronic flavor violation



Hadronic flavor violation



Hadronic flavor violation

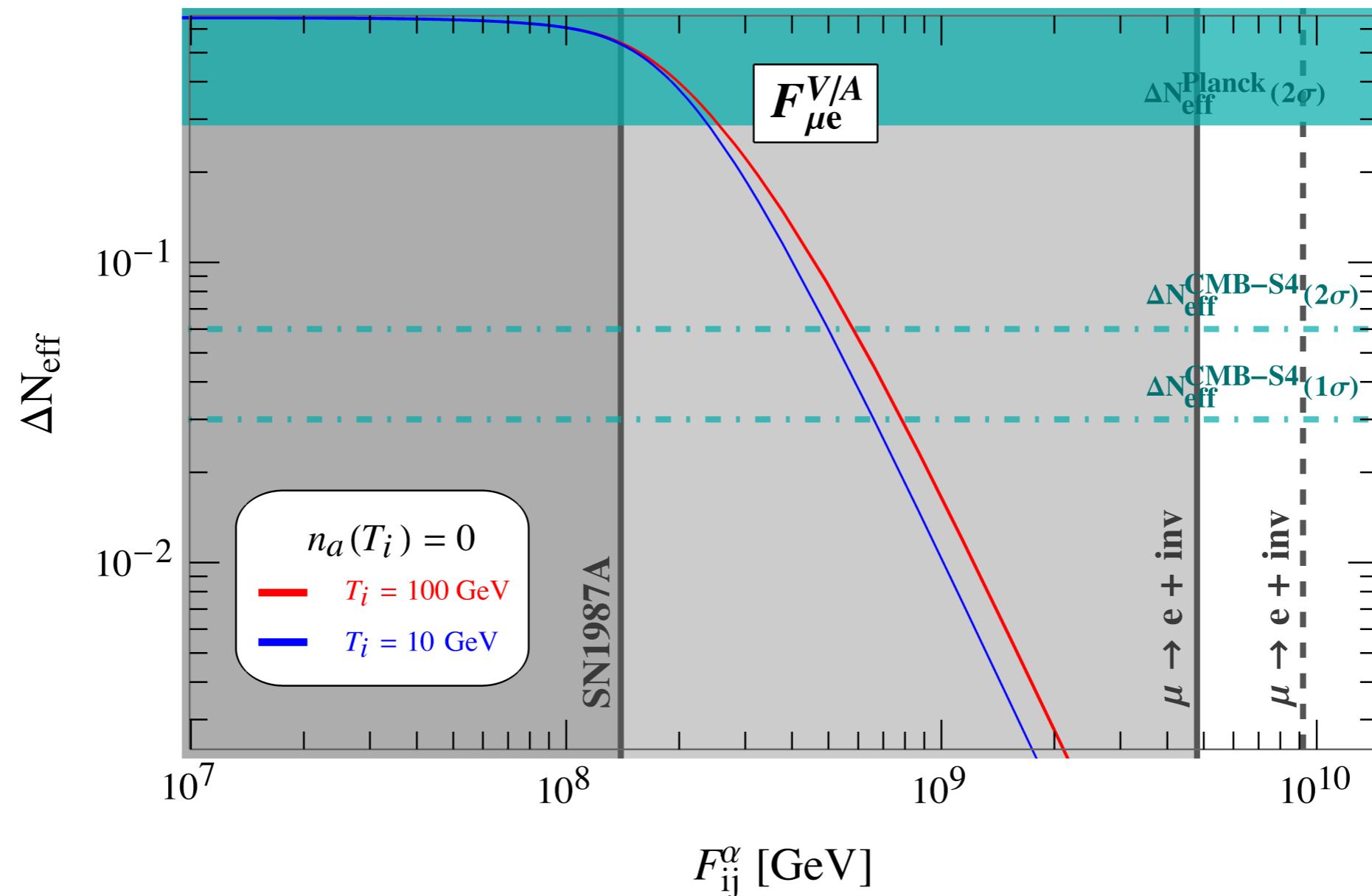


Conclusion

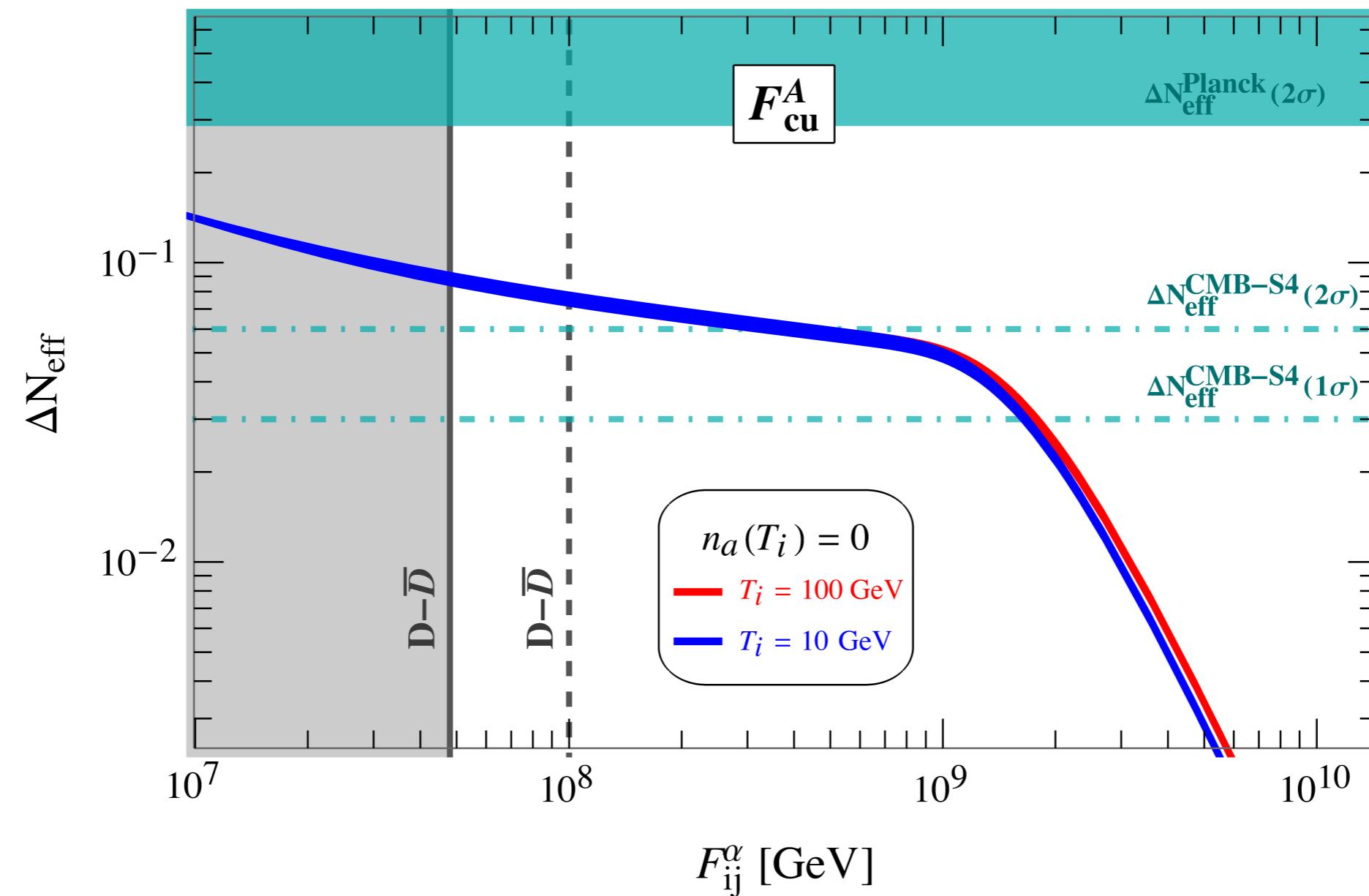
- Flavor violating axion production rate in the full temperature range
- In comparison with terrestrial experiments, future CMB surveys would give comparable or even better sensitivities to probe flavor violating axion couplings
- Possibility to detect flavor violating axion signal for $F_{f_i f_j}^\alpha \sim \mathcal{O}(10^{9-10}) \text{ GeV}$

Back up

Leptonic flavor violation



Hadronic flavor violation



Flaxion

