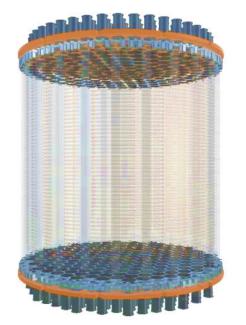
Recent Progress and Plan of PandaX Experiment

Qing Lin

University of Science and Technology of China

On behalf of the PandaX Collaboration

ICHEP, 2022/07/06 - 2022/07/13

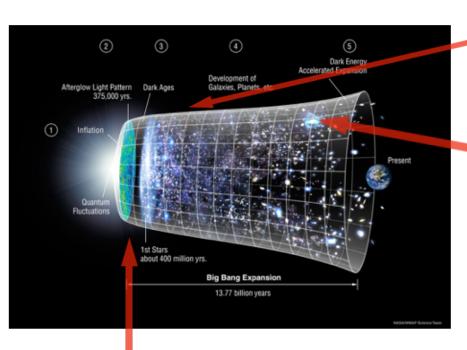






Dark Matter and its Gravitational evidence



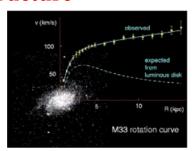


Large Structure

Small Structure

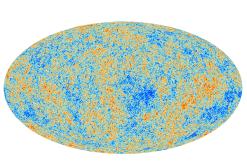


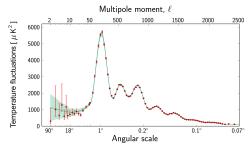
Bullet cluster collision

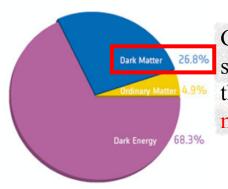


Galaxy rotation curve

Primordial Universe







Gravitational evidences suggest dark matter is the dominant form of matter in Universe!

PandaX Collaboration





























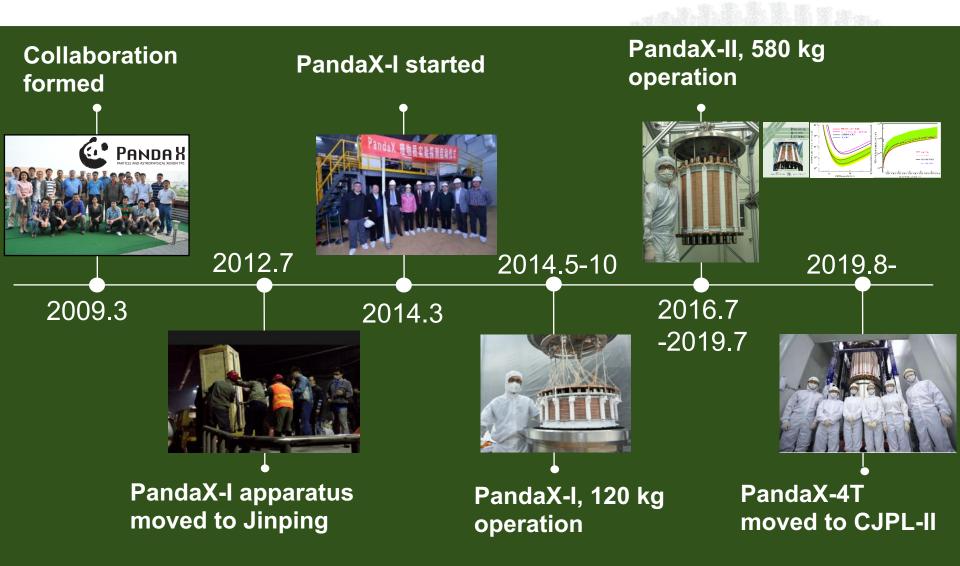




PandaX Experiments



Particle and Astrophysical Xenon Experiments

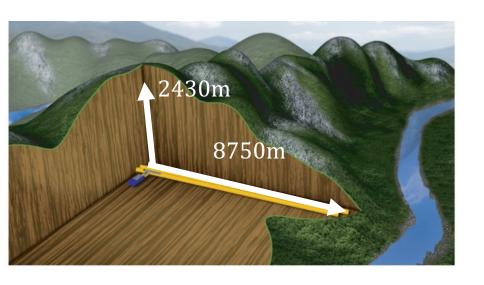


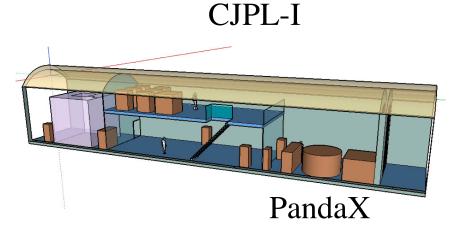
CJPL: Deepest Underground Lab





- Deepest (6800 m.w.e): < 0.2muons/m²/day
- Horizontal access with ~9 km long tunnel: large truck can drive in.
- National key science research facility for dark matter searches, neutrino physics, and astroparticle physics, etc.





CJPL-II: Much Larger



300k m³ with eight main halls of 14x14x64 m (4k m³ of CJPL-I)



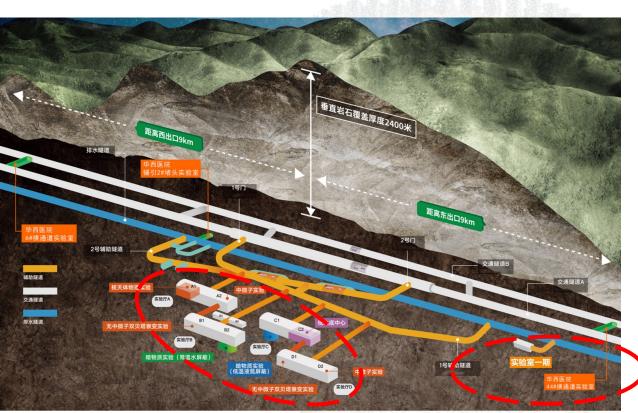
PHYSICS

China supersizes its underground physics lab Planned expansion could pave way for "ultimate dark matter experiment"

he world's deepest physics laboratory is about to become one of its largest. Early next year, workers will start carving four cavernous experiment halls along a tunnel through Jinping Mountain in China's Sichuan province. Once the science at the China Jinping Underground Laboratory (CJPL) is scaled up as well, "it will be a milestone for Chinese physics," says Nigel Smith, director of the

underground SNOLAR in Sudbury Canada

vsics Science, Nov. 30, 2014

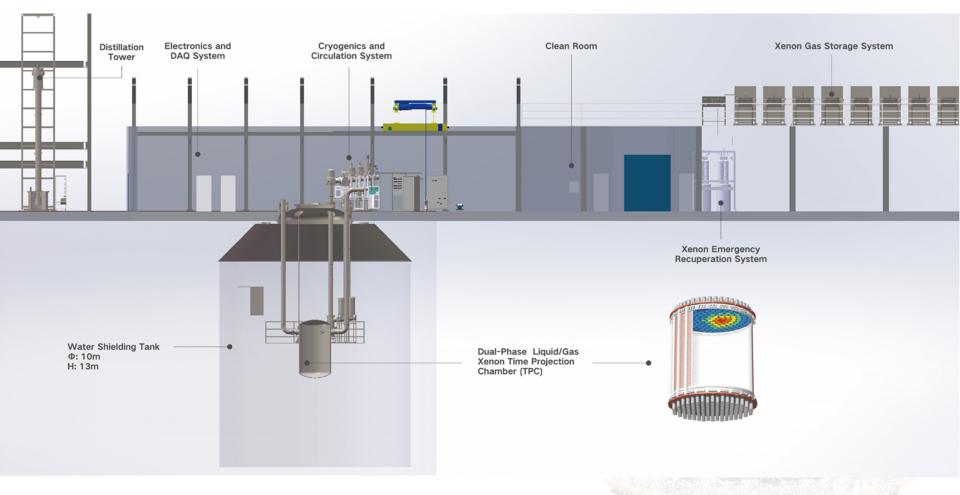


CJPL-II

CJPL-I

PandaX-4T overview @ CJPL-II



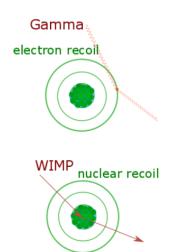


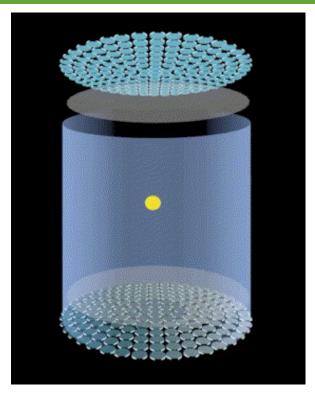
- ☐ Ultrapure water shield: 13 m (H) x 10 m (D) ~ 900 m³
- ☐ TPC: 1.2 m (H) x 1.2 m (D)
- ☐ 3-in PMTs: 169 top/199 bottom

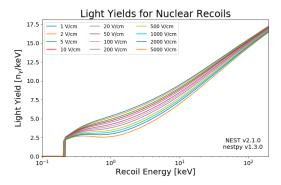
• Sensitive volume: 3.7-tonne LXe

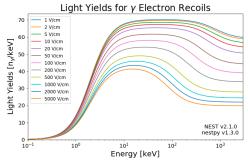
Dual phase xenon TPC



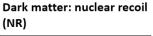




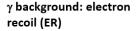


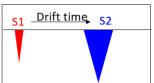


https://nest.physics.ucdavis.edu/benchmark-plots



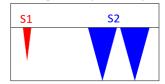






 $(S2/S1)_{NR} < < (S2/S1)_{ER}$

Multi-site scattering background (ER or NR)



Detector capability:

- ☐ Large monolithic target
- □ 3D reconstruction and fiducialization
- ☐ Good ER/NR rejection
- Calorimeter capable of seeing a couple of photons/electrons

PandaX-4T Experiment



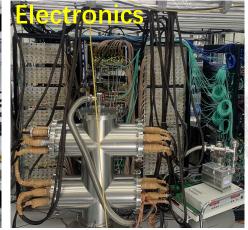


PandaX-4T Subsystems











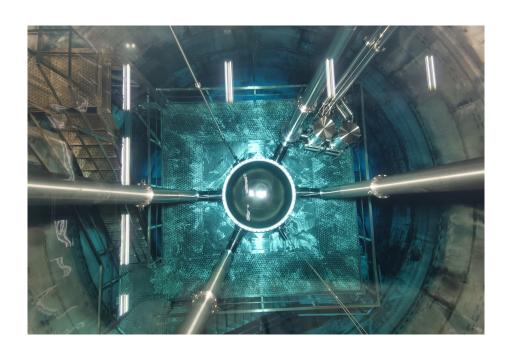


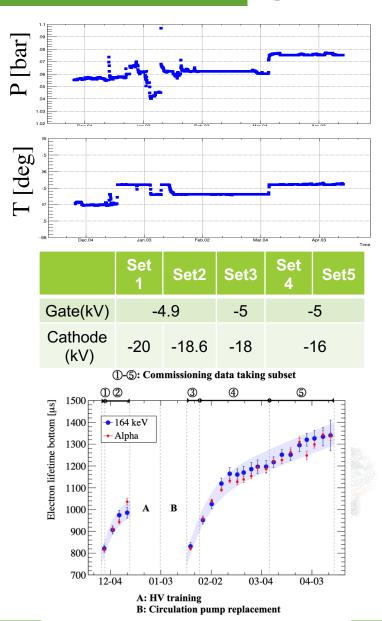


PandaX-4T Commissioning



- ☐ Electron lifetime: *in situ* S2 vertical uniformity calibration
- \square Ref: the maximum drift time ~ 840 μ s (field dependent)
- ☐ Two gas loops for purification
- ☐ Stable data running period: 95.0 calendar days (86 days after selection)



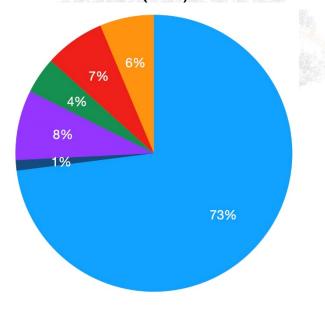


Background composition



Component	Nominal (evts)
³ T (from fit to data)	532 (32)
Flat ER* (18-30keV side band)	492 (31)
Rn	347 (190)
Kr	53 (34)
Material	33 (4)
Xe127	8 (1)
Neutron	0.9 (0.5)
Neutron-X	0.2 (0.1)
Surface	0.5 (0.1)
Accidental	2.4 (0.5)
B8	0.6 (0.3)
Sum	1037 (45)

Expected below-NR-median events: 9.8 (0.6) evts

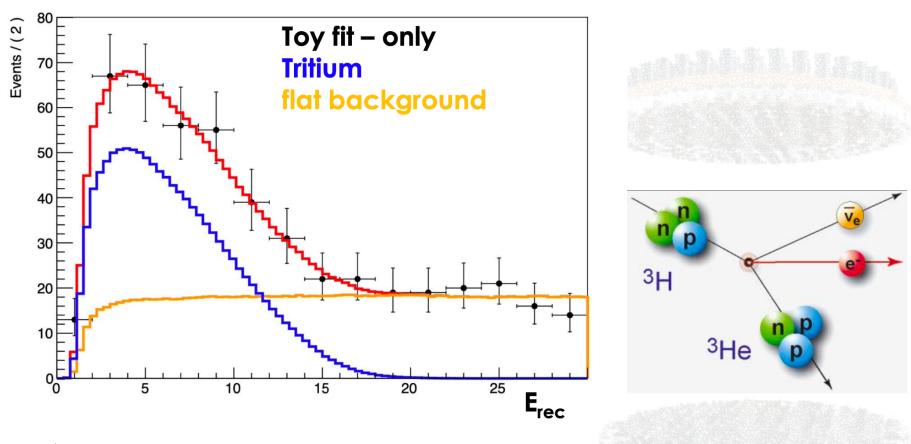


- ☐ Flat ER (Rn+Kr+Material) is determined from side band in DM data
- Background per unit target is improved from PandaX-II by 4 times (<10 keV)

- ER (flat ER + tritium)
- Xe127
 - Accidental
- Surface
- Neutron
- B8

Tritium background

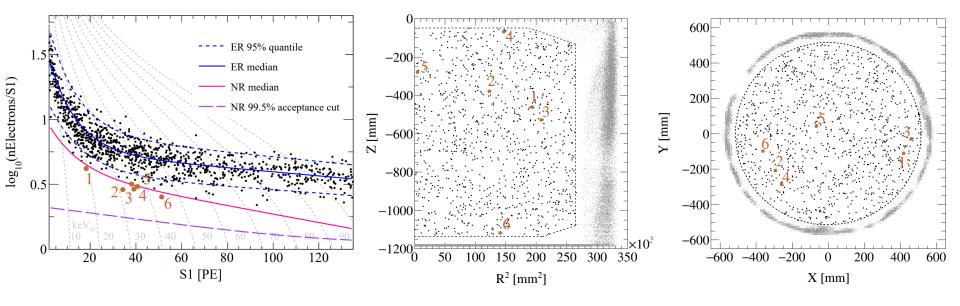




- > Tritium spectrum identified in data;
- Likely originated from CH3T calibration in the end of PandaX-II;
- \triangleright Rate floated in final statistical fit which obtained 5(0.3)x10⁻²⁴(mol/mol)

DM candidates & position distribution

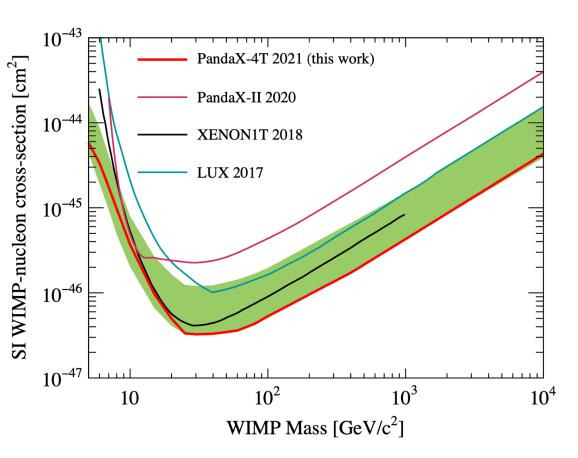




- \square S1 = (2, 135) PE, S2_{raw} > 80 PE, S2 < 20000
- \square FV = 2.67 tonne
- □ 1058 candidates (expected 1054±39), 6 below NR median curve (expected 9.8 ± 0.6)
- ☐ Events uniformly distributed in the FV, as expected if dominated by tritium and radon

WIMP-nucleon SI exclusion limits





- □Exposure: 0.63 tonne•year
- □Sensitivity improved from PandaX-II final analysis by 2.9 times (30 GeV/c²)
- □Our limit is ~1.24 times stronger than XENON1T around 30 GeV/c²

Post-WIMP Analyses



Dark Matter Energy

<keV

B8 CEvNS;
Low-mass WIMP;
WIMP-electron
scatter;
Self-interacting
DM;

 $\sim \text{keV} - 30 \text{ keV}$

Spin-dependent DM;
Self-interacting DM;
DM Absorption;
Migdal effect
Solar axion & axionlike particles;
EFT;
Superheavy DM;

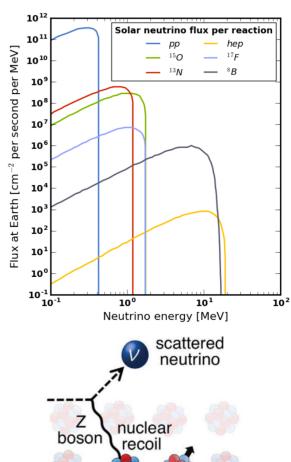
>30 keV

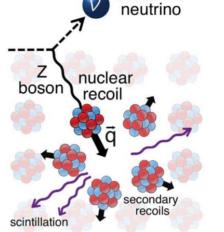
(Neutrinoless) Double
beta decay;
Xe124 double electron
capture;
Solar pp neutrino
electron scatter;

We are taking data with lowered background for more exposure.

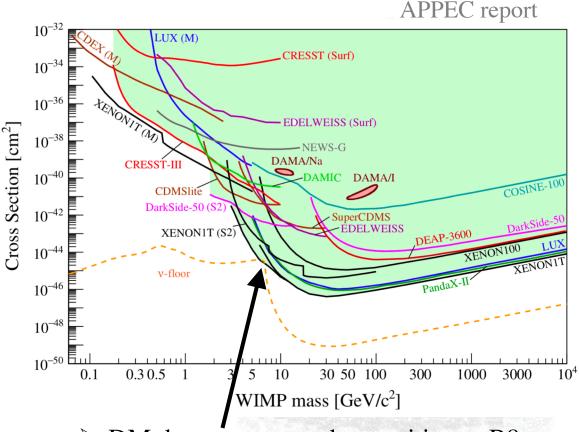
Solar B8 neutrino







10.1126/science.aao0990

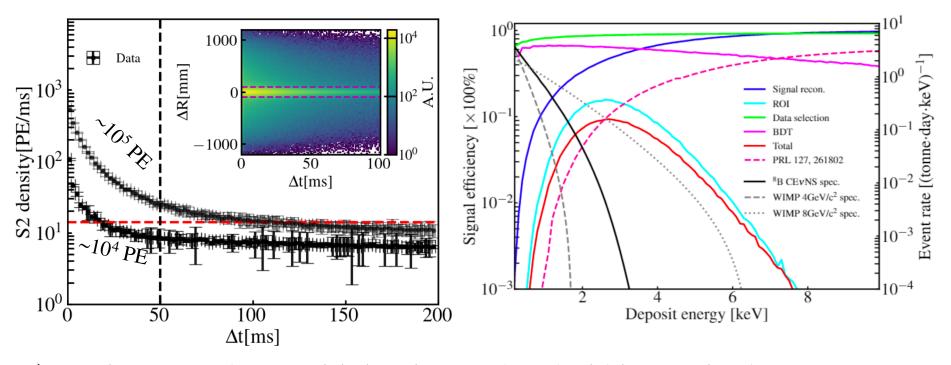


➤ DM detector start to be sensitive to B8 neutrino CEvNS;

> XENON1T (PRL 126, 091301) expects 2, sees 0.

PandaX-4T search on B8 CEvNS





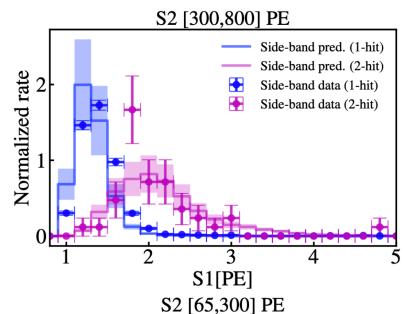
- To increase the sensitivity, lower threshold is required.
- > Key difficulty: Accidental pileup emerges.
- ➤ Various techniques used: waveform simulation & boosted decision tree.
- ➤ Blind analysis is performed with 0.48 tonne-year data, which has a software veto excluding time and position with high rates.

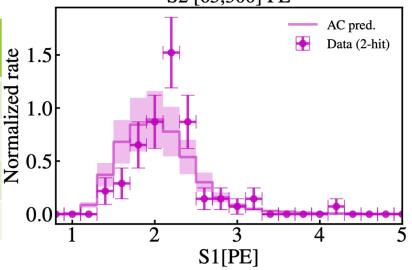
AC background



- ➤ AC samples are "scrambled" S1s & S2s.
- ➤ Boosted Decision Tree (BDT) is trained to maximize the B8-to-AC ratio.
- ➤ Trained with simulation + "scrambed" AC sample;
- Tested on neutron calibration and alternative AC samples, gaining ~25% and ~20% uncertainty for signal and background.
- \triangleright Downward probability of ~20%.

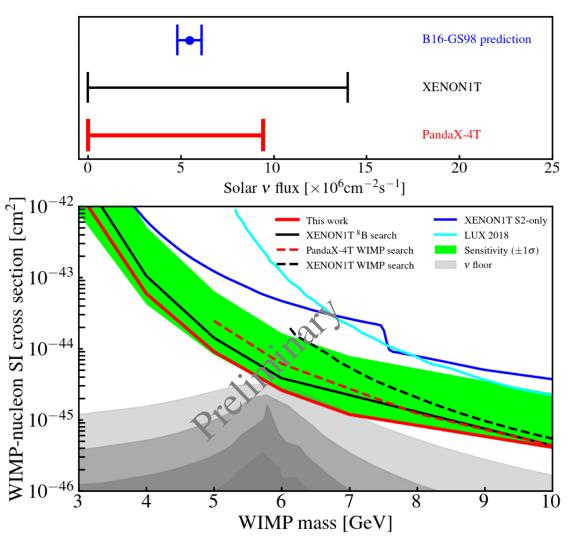
Nhit	BDT	Expected bkg	Expected B8	Observed	
2	off	62.6	2.3	59	
2	on	1.5	1.4	1	-
3	off	0.8	0.4	2	1
3	on	0.04	0.3	0	





B8 search preliminary results

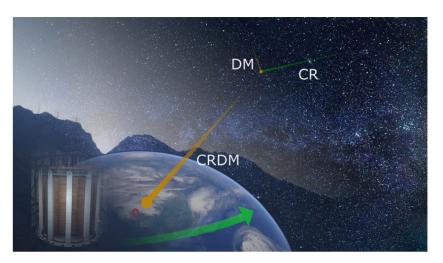




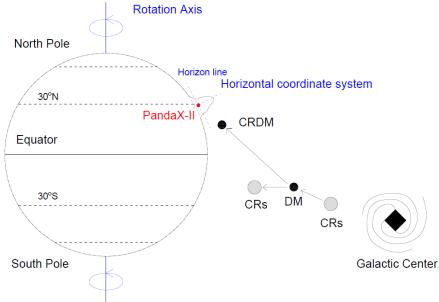
- Leading constraint on B8 neutrino flux among CEvNS detection results.
- ➤ Into sensitivity of the "neutrino floor". Can cast new insight on neutrinonucleus interactions.
- Assuming B8 as bkg, it also gives the most stringint upperlimit for WIMP-nucleon interactions with mass from 3-10GeV.

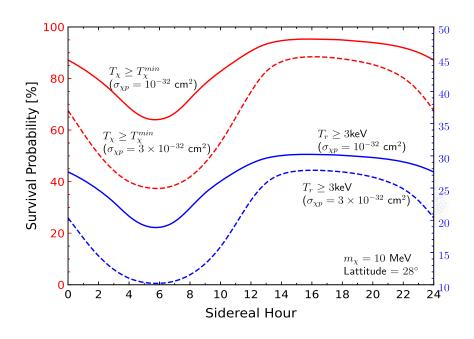
Cosmic ray boosted Dark Matter





- Energetic cosmic ray in Galactic center scatters off low-mass dark matter, boosting it to close to higher speed.
- For low-mass DM, the attenuation of Earth may not be negligible.

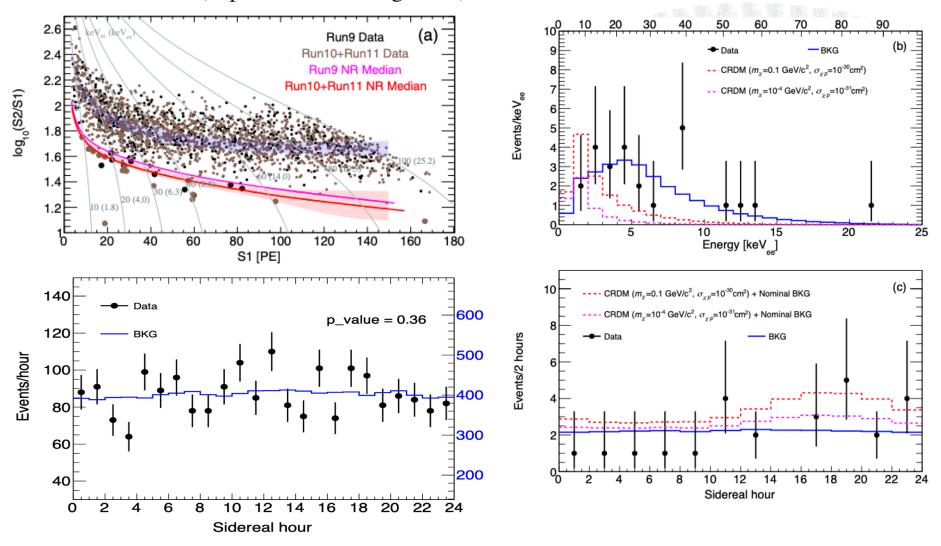




Diurnal modulation in PandaX-II



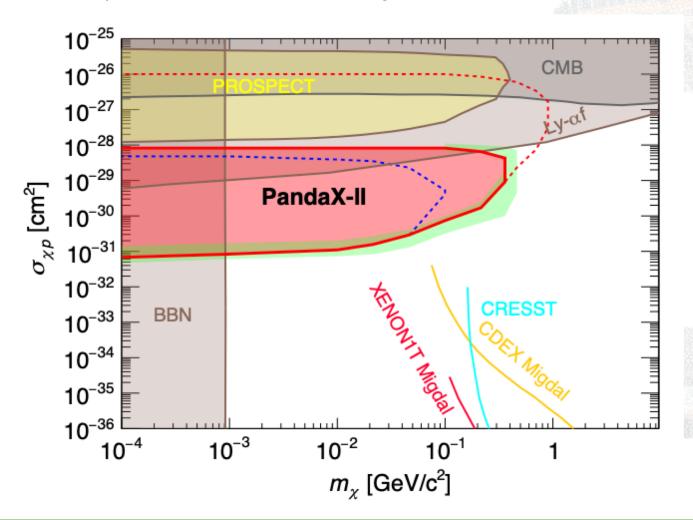
Using events below NR median 25 events (expected 26.6 background)



Constraint on sub-GeV DM



- Expand to the region beyond the astrophysical and cosmological probes;
- The analysis will be extended using PandaX-4T data;



PRL 128, 171801 (2022) Editors' Suggestion

Summary



- ☐ PandaX-4T has completed its commissioning run
- ☐ With a 0.63 tonne·year exposure, PandaX-4T produced the strongest WIMP-nucleon interaction constraint at 2021.
- ☐ An offline tritium removal campaign has been performed, new physics run is on going
- ☐ PandaX-4T analysis on other physics topics, including B8 CEvNS, double beta decay, and etc.
- ☐ In parallel, the collaboration is developing the plan for the next generation experiment at CJPL, we welcome collaborators!





Thank you for listening!







Backups



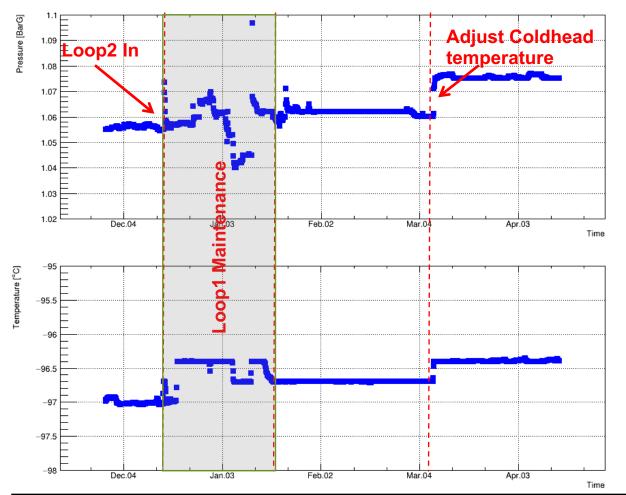






Cryogenics stability



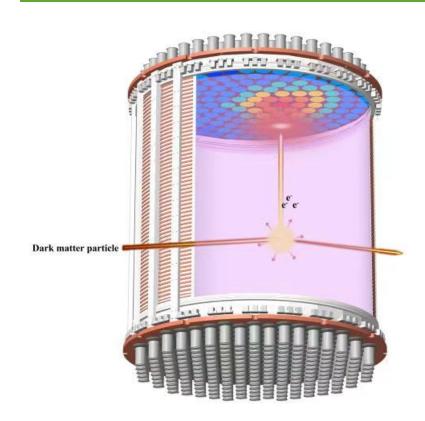


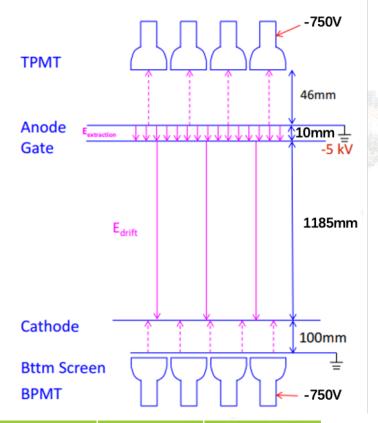
☐ In the stablerunning period, theP and T are stablewithin 0.5% and0.1K, separately.

Parameters	Heating load (No purification)	Maximum Cooling Power	Purification flow rate	Outer Vacuum
Value	~50W	~580W	~110 SLPM (40 kg/h)	<2E-4Pa

TPC operation conditions





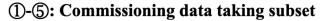


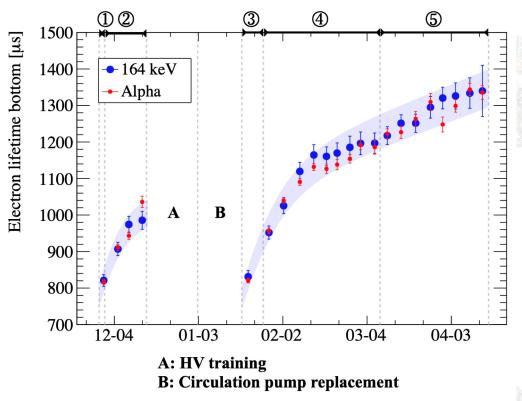
	Set1	Set2	Set3	Set4	Set5
Gate(kV)	-4.9		-5	-5	
Cathode (kV)	-20	-18.6	-18	-	16

During the run, HV set at a few different values to avoid excessive discharges.

Data taking history



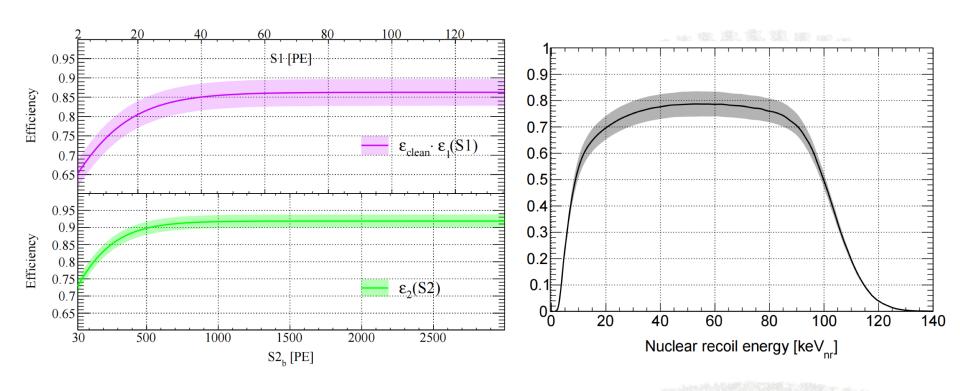




- ☐ Electron lifetime: *in situ* S2 vertical uniformity calibration
- □ Ref: the maximum drift time ~ 840 µs (field dependent)
- ☐ Two gas loops for purification
- ☐ Stable data running period: 95.0 calendar days (86 days after selection)

Efficiencies obtained from calibration data

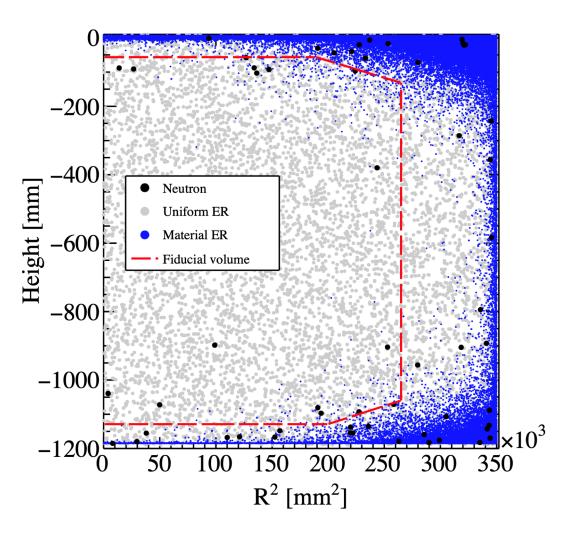




- ☐ Same S1 and S2 efficiency obtained from the ER and NR data
- ☐ Plateaued efficiency at 40 keV_{nr} ~78%.

FV determination

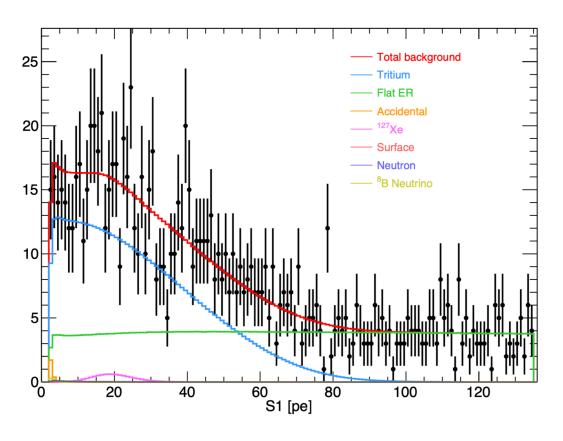




- □ Based on background simulation (10 t-year)
- ☐ Uniform ER (including tritium) normalization come from data
- \Box Define FoM = sqrt(B)/M
- ☐ Best FV = 2.67 tonne

Spectral comparison





- ☐ Fit data with unbinned

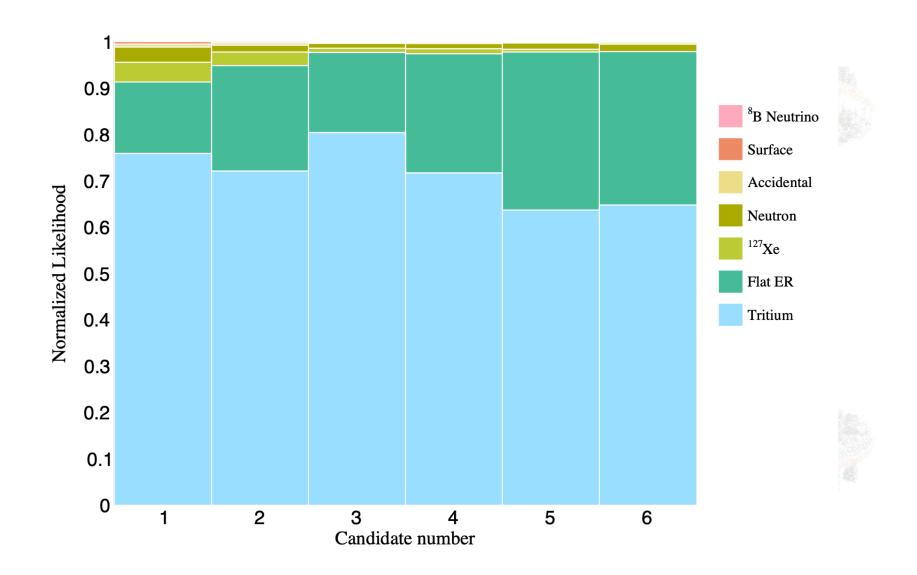
 likelihood with all

 signal/background PDFs in

 (S1, S2_b)
- □ No excess found,background-only p-value0.58
- Spectrum agrees with expected background

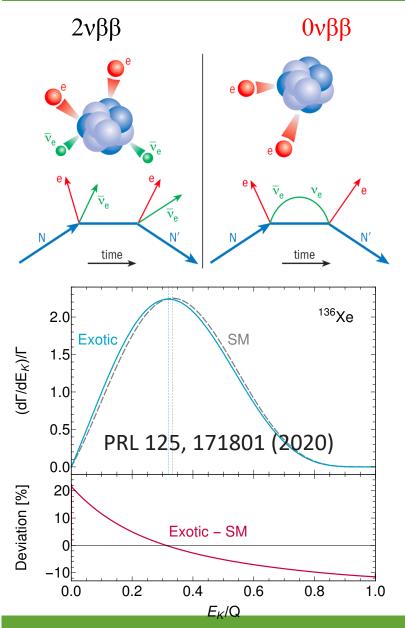
Likelihoods of the six below-NR events

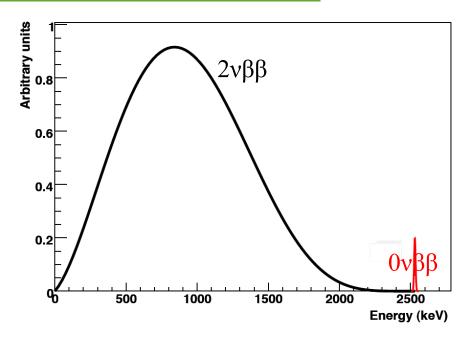




Double beta decay (DBD) measurement





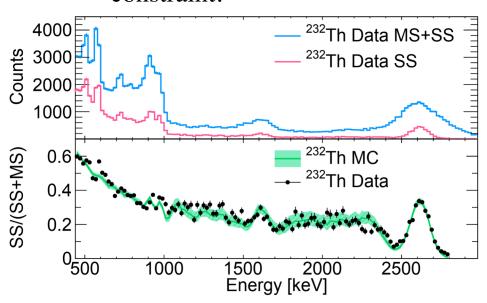


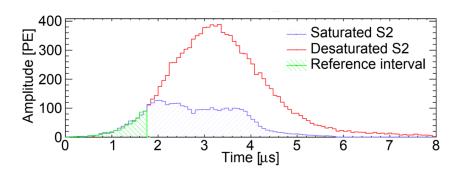
- Precise DBD measurement is essential for neutrinoless DBD, to better understand background;
- > Searching for possible shape distortion, exploring beyond-standard-model physics.

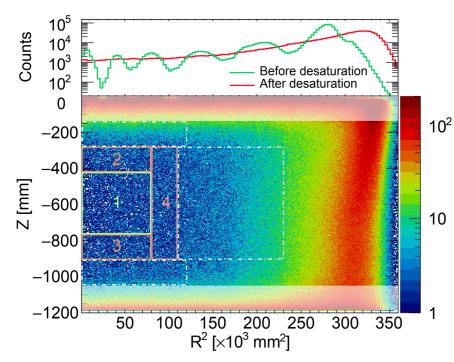
Saturation correction & MC-data comparsion PANDAR



- ➤ Single-site vs. Multi-site identification
 - ➤ Agreement in 2.7% in ROI
- > PMT desaturation
 - > Great improvement in pos. rec. of high-energy events.
- > FV divided into 4 regions
 - ➤ For different bkg component constraint.





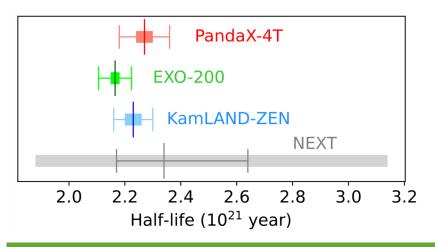


Final fit and Xe136 DBD half-life

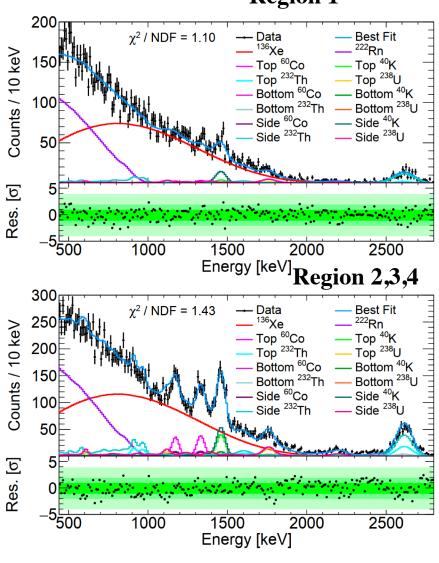


- Simultaneous fit in 4 regions of FV
- 649.7±6.5 kg natural xenon in FV and 94.9 days of data,
 - the total fitted number of Xe-136 DBD events is 17468±243 in ROI of 440 to 2800 keV
- \times Xe-136 DBD half-life is measured as: $2.27\pm0.03(\text{stat.})\pm0.09(\text{syst.})\times10^{21} \text{ yr}$

Source	Percentage	Source	Percentage
Quality cut	0.39%	SS cut	1.7%
FV cut	1.0%	Bin size	0.05%
Fit range	1.2%	Energy resolution	0.58%
Energy scale	0.26%	Regional weight	1.6%
²¹⁴ Pb spectrum	2.0%	LXe density	0.13%
¹³⁶ Xe abundance	1.9%	Total	4.1%







2205.12809