FIRST WIMP SEARCH RESULTS FROM THE LUX-ZEPLIN EXPERIMENT Amy Cottle, University of Oxford



COLLABORATION







Science and Technology Facilities Council







<u>@Izdarkmatter</u> https://lz.lbl.gov/



INTRODUCTION TO LZ



<u>NIM A, 163047 (2019)</u>



- Based at the Sanford Underground
- Dual-phase xenon time projection •

TPC DETECTION PRINCIPLE

- Interactions in the xenon create
 - Light prompt scintillation S1
 - Charge electrons drifted and extracted into gas -> proportional scintillation - S2
- Excellent 3D position reconstruction (~mm)
- Distinguish between single scatter (SS) and multiple scatter (MS) interactions
- S2:S1 ratio discriminate electronic recoils (ERs) from potential WIMP nuclear recoils (NRs)

Incoming Particle





VETO DETECTOR ANTI-COINCIDENCE

- 17 tonnes Gd-loaded scintillator in OD



TPC & SKIN ASSEMBLY







OD CONSTRUCTION & UNDERGROUND INSTALLATION

Water Tank Panoramic

Cryostat Insertion

Instrumented OD



FIRST SCIENCE RUN (SR1)

[sm

- 116 calendar days -> 89 live days
- Stable detector conditions •
 - Temperature of 174.1 K
 - Gas pressure of 1.791 bar
 - Drift field of 193 V/cm
 - Extraction field of 7.3 kV/cm (in gas)
- Continuous purification at 3.3 t/day through hot getter system
- Demonstration run, not blinded



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TPC CALIBRATIONS

- Backgrounds predominantly ERs; WIMPs produce NRs
- Tritiated methane (CH₃T) injection to calibrate ER band
 - Spatially homogenous β source
- DD neutron generator (NR band) •
 - Monoenergetic 2.45 MeV neutrons
- 99.9% discrimination of beta backgrounds under NR band median achieved

4.50 4.25 4.00[[phd]] 3.75 S2(3.50 \log_{10} 3.25 3.00 2.75 2.50



RADON BACKGROUNDS

- Naked Pb β decays main WIMP background
 - Produced from emanated Rn in TPC xenon
- Rn-chain alpha tagging measurements performed with S1s at energies >> WIMP ROI

Rn222 (µBq/kg)	Po218 (µBq/kg)	Po214 (µBq/
4.37 ± 0.31	4.51 ± 0.32	2.56 ± 0.2

- Rn222 activity within assay expectations
- Po218/Po214 rates bound Pb214 rate





INTERNAL BACKGROUNDS

- Kr85 constrained by sampling at 144 +/- 22 ppq g/g nat Kr/Xe



Activated xenon & contaminant rates informed via energy spectra fits & analyses >40 keVee - Pb214 (3.26 µBq/kg) & Xe127 (36.9 µBq/kg SR1-averaged) constraints for WIMP analysis



ACCIDENTAL COINCIDENCE BACKGROUNDS

- Lone S1s & S2s can accidentally combine to form WIMP ROI events
- Data-driven estimation of distribution
- Event classification drift time agnostic
 - Drift time >1 ms = definite accidentals
 - Population used to inform rate
- Lone S1 & S2 waveforms extracted and stitched together to form fake events
 - Can produce statistics to assess shape
- Analysis cuts developed to combat observed pulse/event pathologies



counts/tonne/year

12

WIMP ANALYSIS - ROI & FV

- Region of Interest definition •
 - 3 < S1c < 80 photons detected (phd); three-fold PMT coincidence
 - Uncorrected S2 > 600 phd; log10 (S2c) < 5
- Fiducial volume (FV) definition •
 - 86 us < drift time < 936.5 us cut to avoid higher background rates at TPC edges
 - Radial cut chosen to ensure <0.01 wall ____ background counts in the FV
- Calculated fiducial mass of 5.5 ± 0.2 t •



WIMP ANALYSIS - CUTS & DATA QUALITY

1.0

- Event selection criteria
 - FV, ROI, single scatter cuts
 - Veto detector anti-coincidence
 - S1/S2 shape cuts
- Cuts developed on non-WIMP ROI background & calibration data
- Rejection of live time with detector instabilities, high TPC pulse rates
 - 60.3 ± 1.2 live days



WIMP ANALYSIS - DATA & STATISTICAL INFERENCE

• 335 events after all cuts

- PDFs created with energy deposit + detector response simulations*
- Profile likelihood ratio analysis

Key

- 1 & 2-Sigma Contours
- Post-fit total background distribution
- Ar37
- B8
- 30 GeV/c² WIMP

NR band from DD

* j.astropartphys.2020.102480

WIMP ANALYSIS - BACKGROUNDS & STATISTICAL INFERENCE

Component	Expected Events	Best Fit Events
β decays & detector γs	218 ± 36	222 ± 16
Ar37	[0, 291]	$52.1^{+9.6}_{-8.9}$
Xe127	9.2 ± 0.8	9.3 ± 0.8
Xe124	5.0 ± 1.4	5.2 ± 1.4
Xe136	15.2 ± 2.4	15.3 ± 2.4
Solar v ERs	27.3 ± 1.6	27.3 ± 1.6
B8 CEvNS	0.15 ± 0.01	0.15 ± 0.01
Det. Neutrons	$0.00^{+0.02}$	$0.00^{+0.02}$
Accidentals	1.2 ± 0.3	1.2 ± 0.3
Total w/o ³⁷ Ar	276 ± 36	281 ± 16
Total w/ 37Ar		333 ± 17

WIMP ANALYSIS - SR1 LIMIT

- Two-sided PLR search as per recommended conventions*
- Minimum cross-section of σ_{SI} = 5.9 × 10⁻⁴⁸ cm² for WIMP mass of 30 GeV/c²
- No evidence for WIMPs

<u>Key</u>

- Observed limit
- --- Median expected sensitivity
- •--• Median 3-sigma evidence

CONCLUSIONS & FUTURE PLANS

- All LZ systems are performing well
- Background validations proceeding
- World-leading spin-independent WIMP search limit achieved
- Preparing for a year-long run & ultimately 1000 live days of data
 - Detector optimisation & calibrations campaign in motion
- Additional papers in preparation
- <u>XLZD consortium</u> formed, looking towards a next-generation project

First Dark Matter Search Results from the LUX-ZEPLIN (LZ) Experiment

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Paper Link

BACKUP SLIDES

TPC ENERGY RESPONSE

- S1s & S2s position-corrected using ^{131m}Xe background, ^{83m}Kr calibration
- Doke plot constructed with monoenergetic electron recoil peaks

VETO DETECTOR RESPONSE

- Skin & OD response and inter-detector timings calibrated
 - OD optical calibration system
 - External γ-ray & neutron sources (e.g. ²²Na; DD, AmLi, ²⁵²Cf)
- ¹²⁷Xe Skin tagging efficiency of 78 ± 5% based on K-shell analysis
- OD tagging efficiency of TPC-interacting neutrons of $89 \pm 1\%$ (AmLi calibrations)
 - TPC-OD coincidence window: 1200 µs; threshold equivalent to ~200 keV

LIVE TIME VETOES

- thus contributing to accidental coincidence backgrounds
- Removal of periods • after S2s (e-/ph trains) excludes ~30% of our live time
- Working on optimising this live time veto for future runs

LIMIT SHAPE

Downward fluctuation in limit caused by deficiency of events under Ar37 contour

Calibrations and Xe127 M-shell counts as expected under signal acceptance model -> background under-fluctuation

AR37 ESTIMATE

- Ar37 a significant background in early LZ data/SR1 WIMP search
 - K-shell e⁻ capture -> 2.8 keV
 - $\tau_{1/2} = 35$ days
- Can be produced via cosmic spallation on xenon
 - Calculated using the ACTIVIA package & estimated exposure of the xenon during transport*
 - Large uncertainties in spallation cross-section

* PRD 105, 082004 (2022)

DETECTOR NEUTRON CONSTRAINT

- Perform fit to events passing all WIMP search cuts except OD anticoincidence to constrain neutrons
 - Expect ~8 times more neutrons than in the WIMP search due to OD veto efficiency
 - In contrast, only 5% of nonneutron backgrounds should be OD-tagged
- Result of fit -> number of neutrons in SR1 WIMP search is <0.2 events (two-sided constraint)

5.0

4.5 ([phd]) 4.0 3.5 3.5

3.0

