

<https://arxiv.org/abs/2112.07274>

<https://arxiv.org/abs/2204.04204>

Global analysis of electroweak data in the Standard Model and beyond

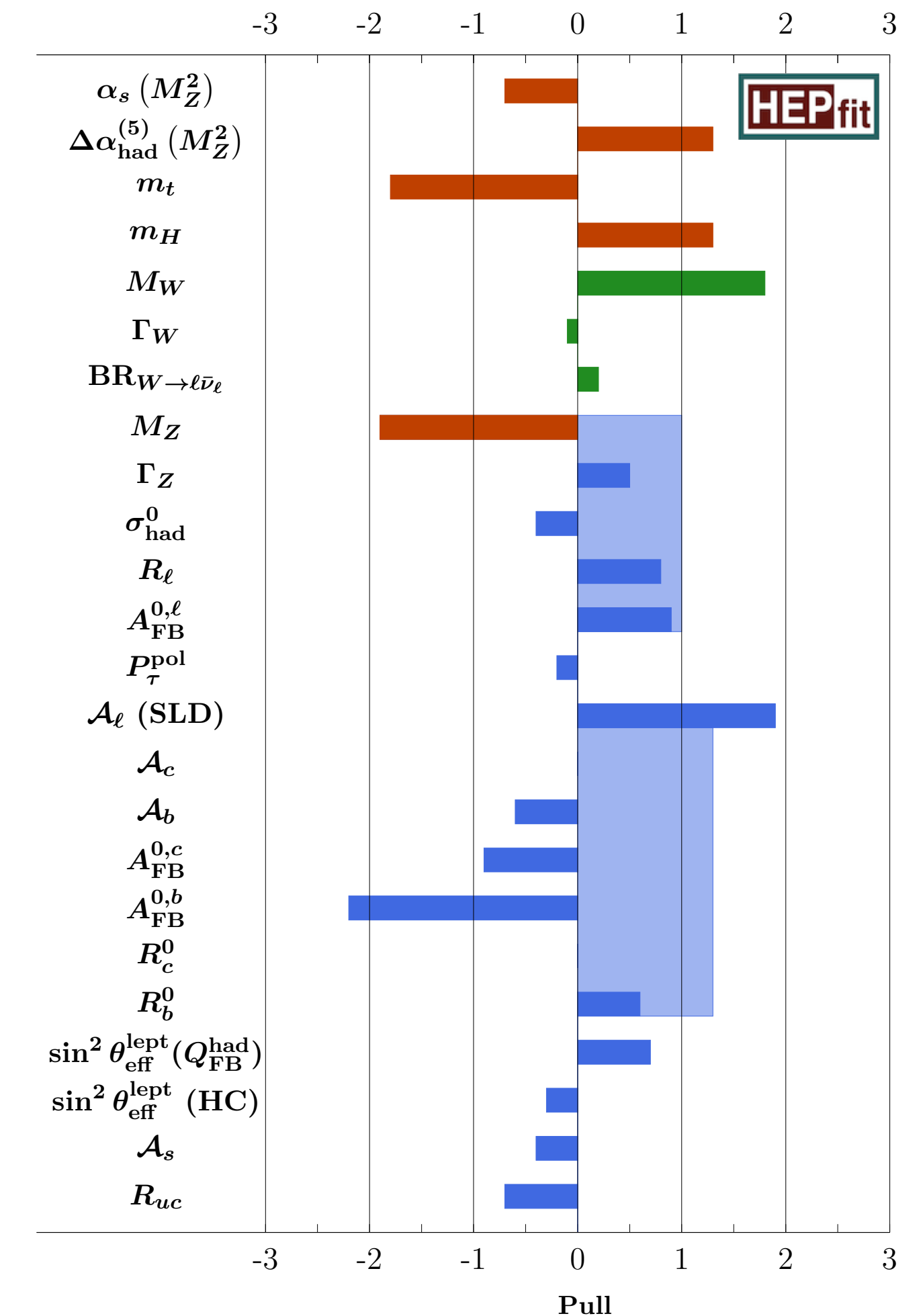
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(CERN)

in collaboration with J. de Blas, M. Ciuchini, E. Franco, A.
Goncalves, S. Mishima, L. Reina, L. Silvestrini
on behalf of the HEPfit collaboration



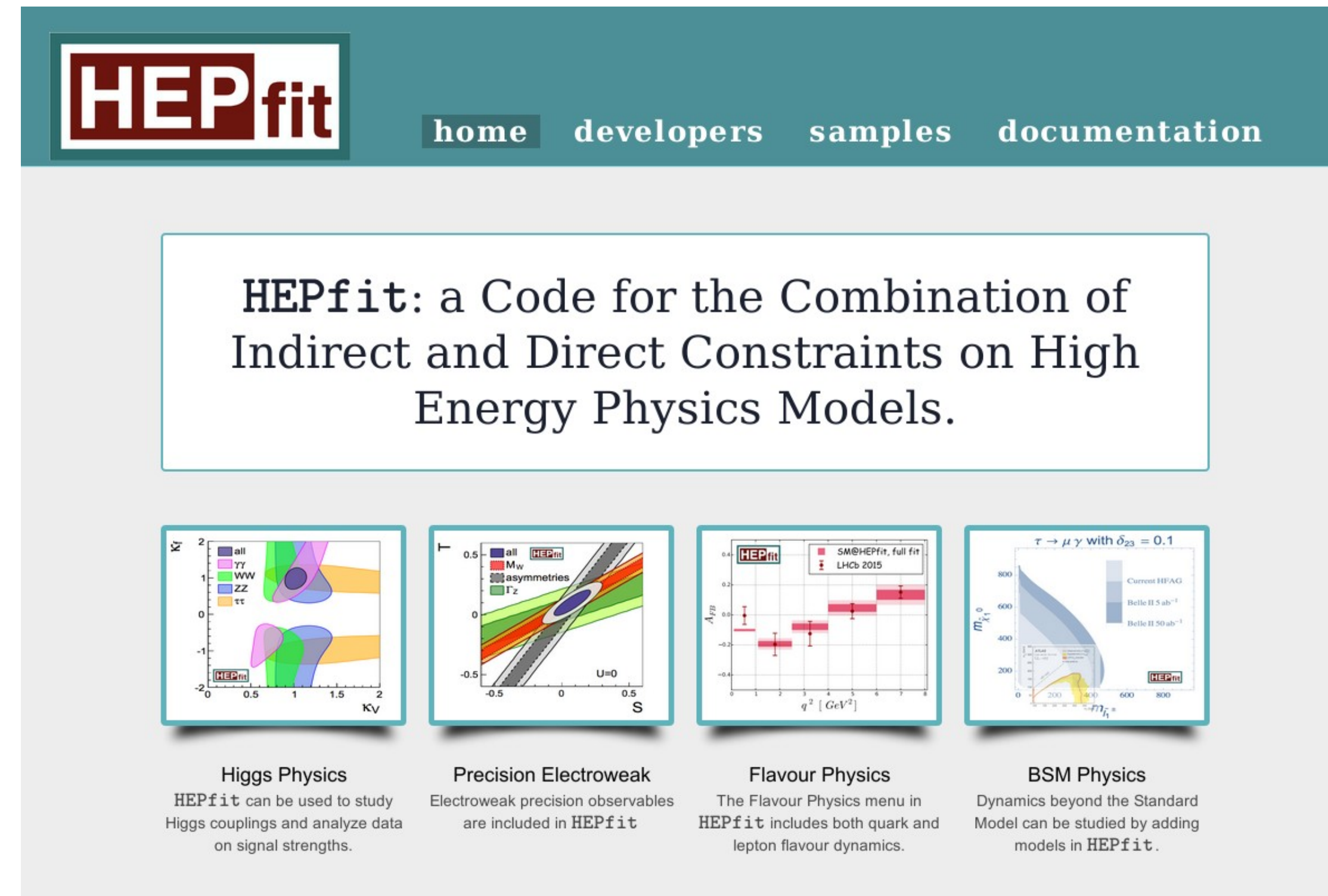
The Electroweak fit

- Exploit the over-constrained EW sector (dictated by rigid symmetry structure) to perform consistency tests of the SM with EW precision observables
- Set of input parameters (α scheme): G_F , α , m_Z , m_H , m_t , $\alpha_s(m_Z)$, $\Delta\alpha_{\text{had}}^{(5)}$
- Compute EW precision observables as functions of these quantities
 - Z-pole observables
 - W observables
- Compare computations to experimental data to learn the values of the input quantities
- Extend relations to include BSM effects and determine bounds on New Physics
 - Oblique parameters: S, T, U, ...
 - Effective interactions: SMEFT
 - ...



The HEPfit library

- Open source library to perform combined fits of HEP observables (including EWfit) in various scenarios
- Computes EWPO in SM, SMEFT, several kinds of 2HDM, some SUSY (mostly LFV), etc.
- Allows for Bayesian analysis exploiting MCMC via the [Bayesian Analysis Toolkit](#)
- Why a Bayesian fit if we have several others based on likelihood-ratios?
- Answer 1: Why not?
- Answer 2: when conclusions depend on the statistical approach, there are no solid conclusions
- Answer 3: because HEPfit is more than an EWfit tool



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Special Article - Tools for Experiment and Theory

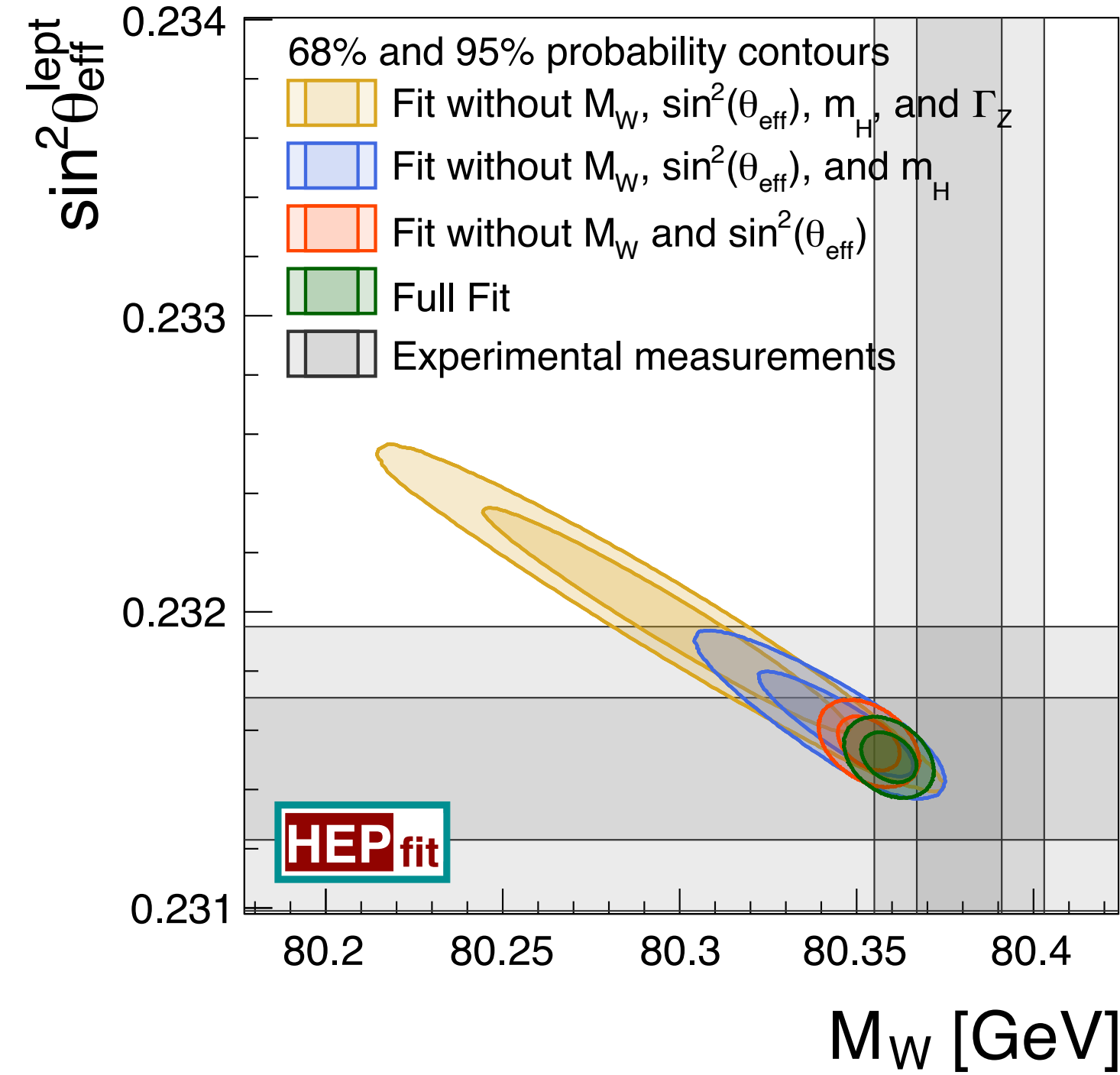
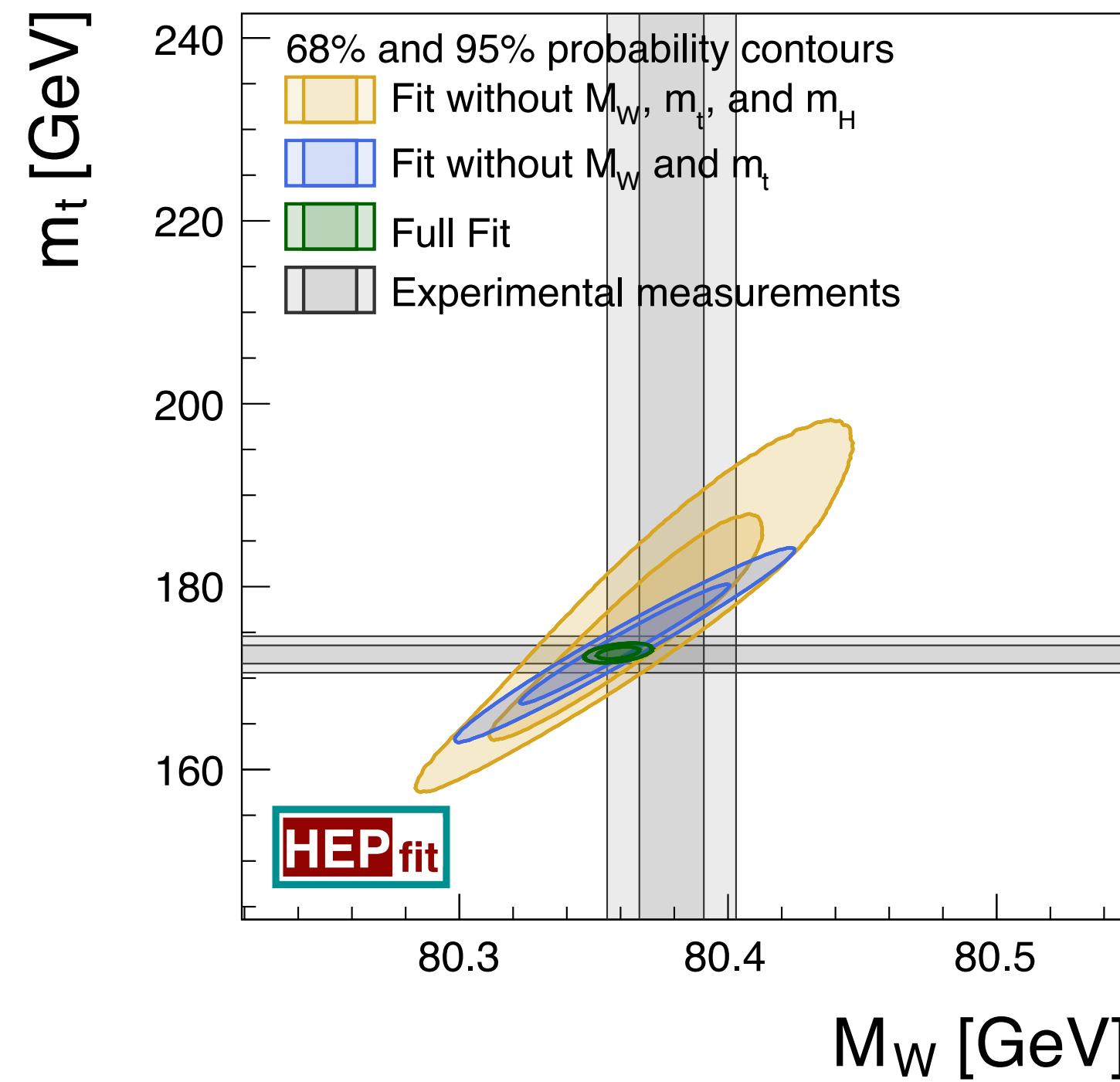
HEPfit: a code for the combination of indirect and direct constraints on high energy physics models

J. de Blas^{1,2}, D. Chowdhury^{3,4}, M. Ciuchini⁵, A. M. Coutinho⁶, O. Eberhardt⁷, M. Fedele⁸, E. Franco⁹, G. Grilli di Cortona¹⁰, V. Miralles⁷, S. Mishima¹¹, A. Paul^{12,13,a}, A. Peñuelas⁷, M. Pierini¹⁴, L. Reina¹⁵, L. Silvestrini^{9,16}, M. Valli¹⁷, R. Watanabe⁵, N. Yokozaki¹⁸

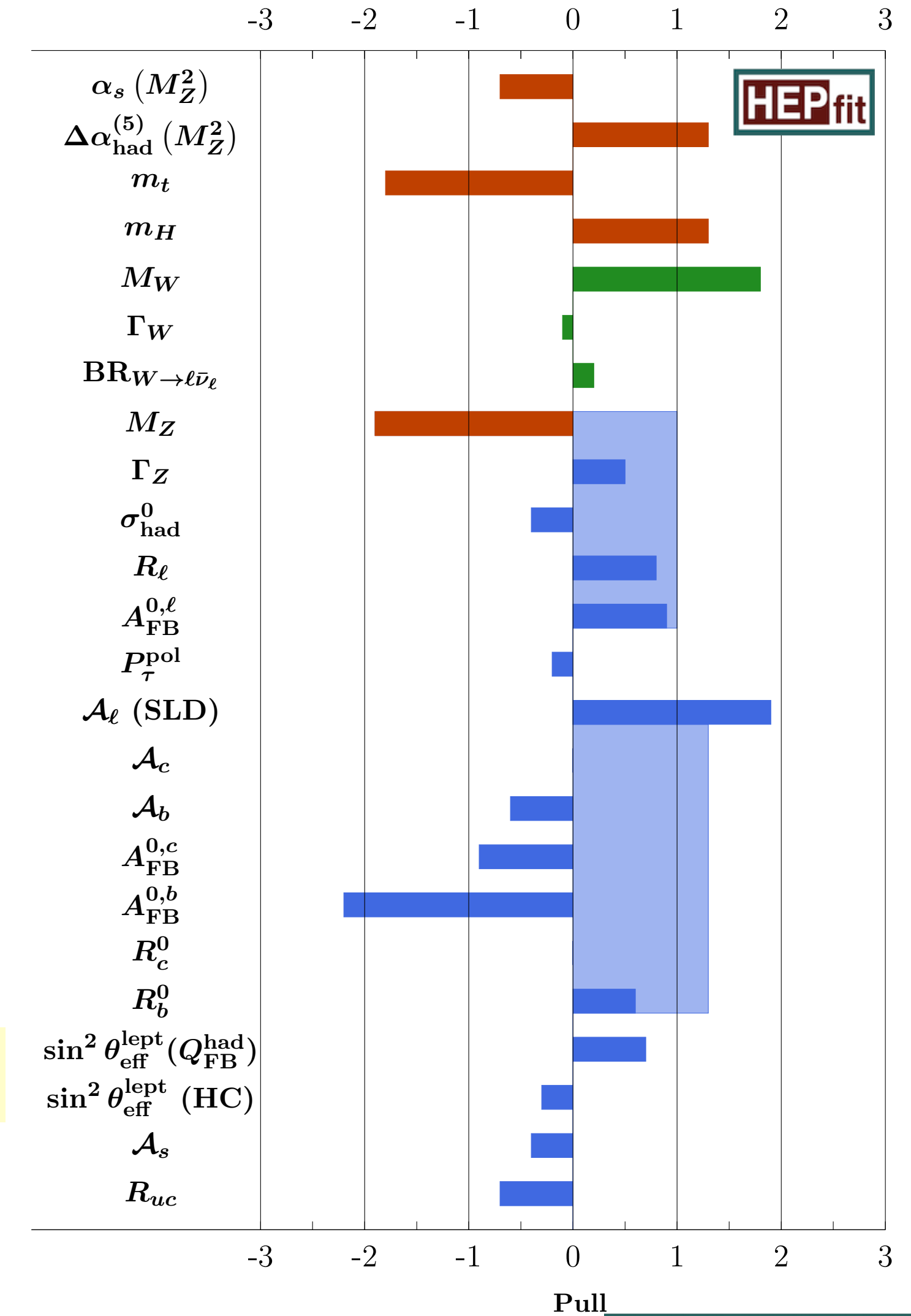
<https://arxiv.org/abs/1910.14012>



Fit result @Fall 2021



In conclusion, EWPD appear to be fully compatible with the SM, with no more tensions than expected from statistical fluctuations. In the *standard scenario* SM fit, the largest pull neglecting correlations is 2.2σ on 24 observables, while taking correlations into account it is 1.8σ on 14 observables. In both the *full indirect* and *full prediction* determinations, the largest pull neglecting correlations is 2.1σ on 24 observables. To quantify further the agreement of the SM, we generated 600 toy experiments centered on the *full prediction* with the current experimental uncertainty and computed the fraction of toys in which the largest pull was larger than the largest one observed in real data. This fraction is an estimate of the *global p-value*. Neglecting correlations, we obtain $p = 0.53$, corresponding to 0.6σ for a Gaussian distribution, while taking into account the correlations (fixed to the values observed in current data) we get $p = 0.45$, corresponding to 0.8σ .



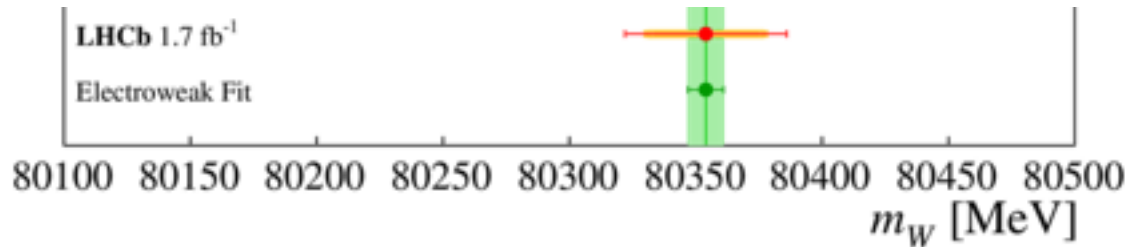
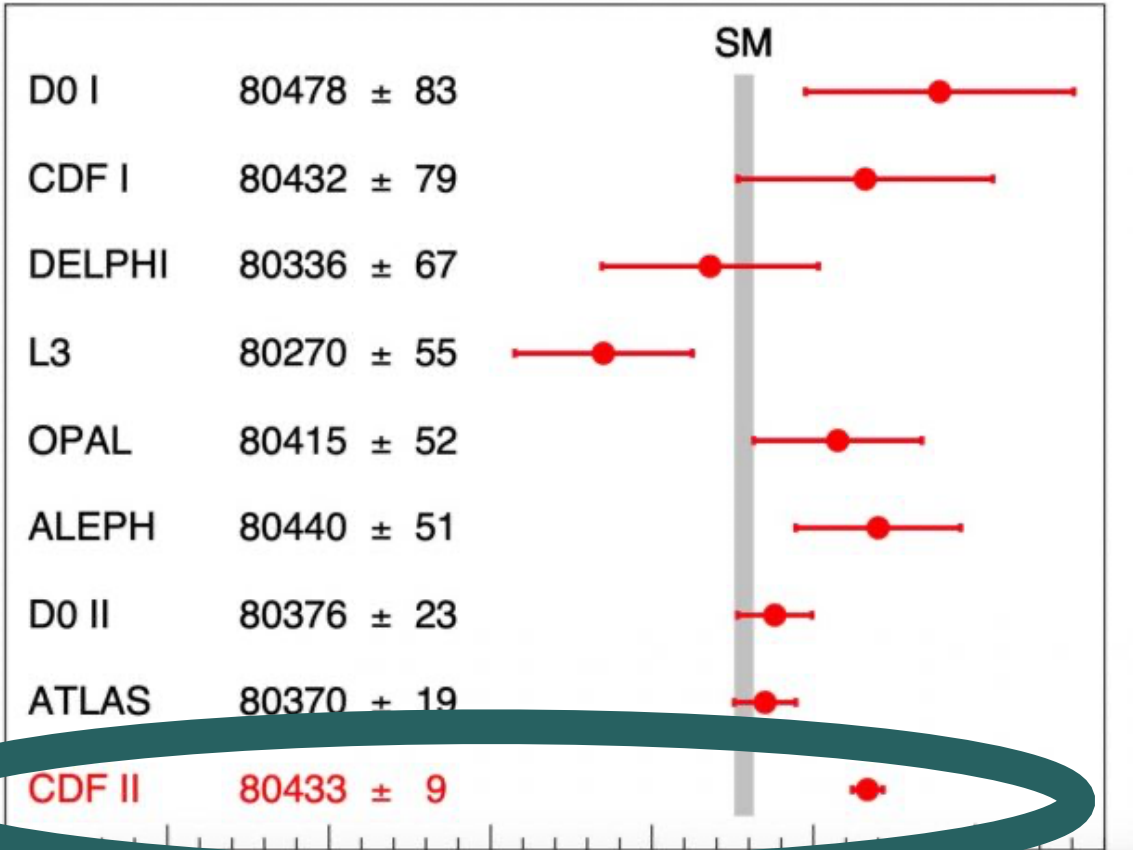
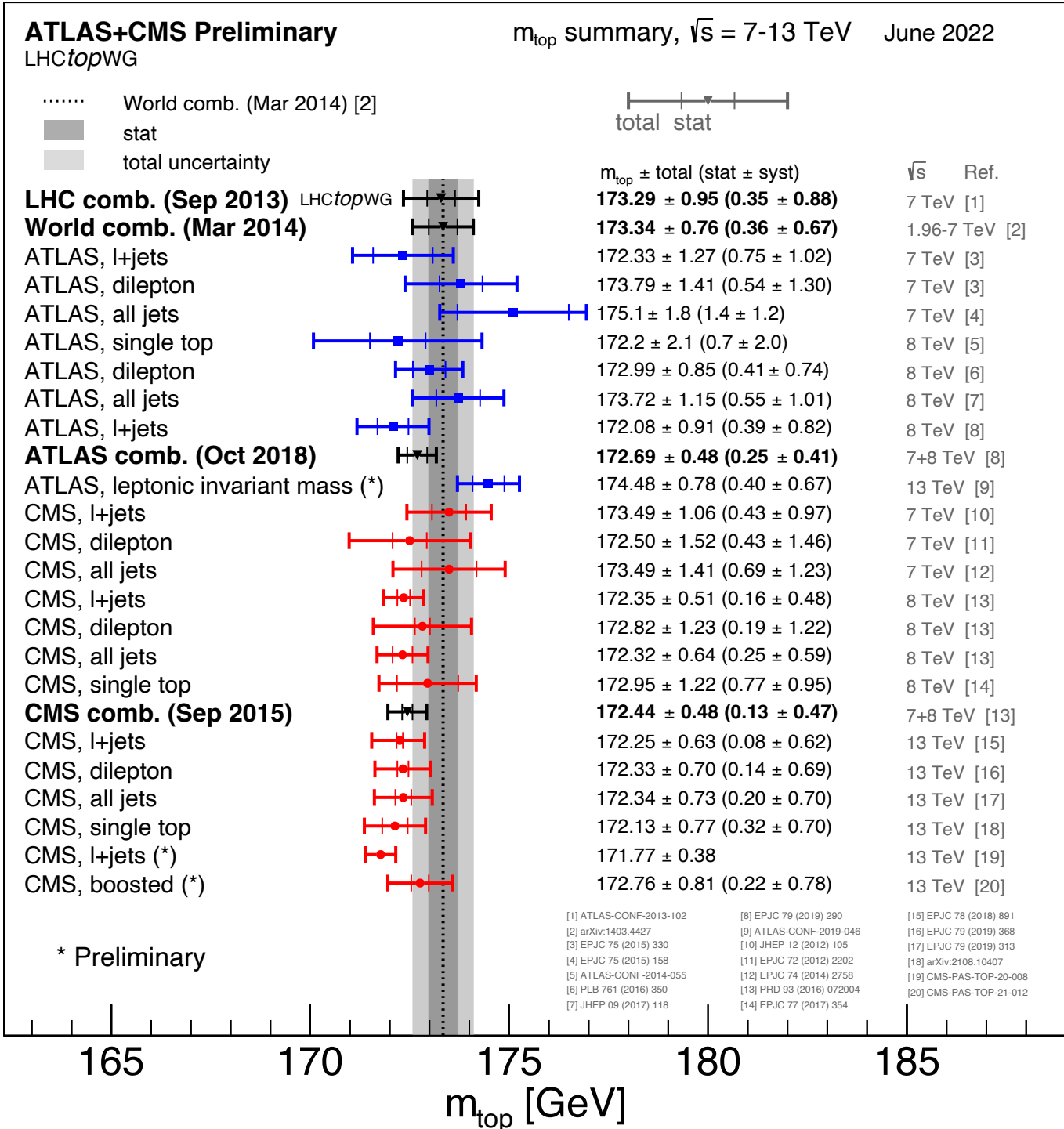
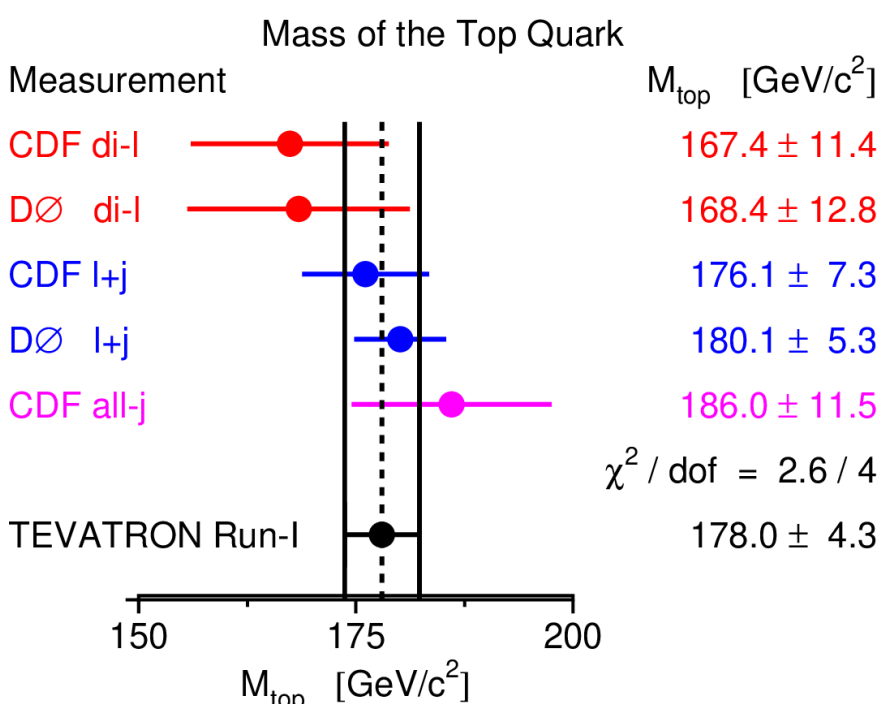
New precision reached, with some question mark

● To which extent are these new measurements compatible to previous measurements?

- ➡ Not addressed here
- ➡ This is up to the experimental collaborations to establish (see talk by S. Amoroso)

● Assuming that they are, can we learn something about New Physics

- ➡ Oblique correction scheme (S,T,U, etc.)
- ➡ SMEFT



Fit setup adding new measurements

● Input parameters $\underbrace{\alpha, G_F, \alpha_s(M_Z)}_{\text{fixed}}, M_Z, M_H, m_t, \Delta\alpha_{\text{had}}^{(5)}$

● Observables from LEP, Tevatron, and LHC

● Recent changes

● new top mass

- 2016 Tevatron combination
- ATLAS Run 1 and Run2 results
- CMS Run 1 and Run 2 results
- Recent CMS l+j measurement [$m_t = (171.77 \pm 0.38)$ GeV]

Theory intrinsic uncertainties on input parameters

$$\delta_{\text{th}} M_W = 4 \text{ MeV}, \delta_{\text{th}} \sin^2 \theta_W = 5 \times 10^{-5}$$

$$\delta_{\text{th}} \Gamma_Z = 0.4 \text{ MeV}, \delta_{\text{th}} \sigma_{\text{had}}^0 = 6 \text{ pb}$$

$$\delta_{\text{th}} R_l^0 = 0.006, \delta_{\text{th}} R_c^0 = 0.00005$$

$$\delta_{\text{th}} R_b^0 = 0.0001$$

previous average
 $m_t = 172.58 \pm 0.45 \text{ GeV}$



new average
 $m_t = 171.79 \pm 0.38 \text{ GeV}$
"standard"

new average
 $m_t = 171.79 \pm 1.00 \text{ GeV}$
"conservative"

details on backup



Waiting for official Tevatron+LHC averages to be released

Fit setup adding new measurements

● Input parameters $\underbrace{\alpha, G_F}_{\text{fixed}}, \alpha_s(M_Z), M_Z, M_H, m_t, \Delta\alpha_{\text{had}}^{(5)}$

● Observables from LEP, Tevatron, and LHC

● Recent changes

● new W mass

- All LEP 2 measurements *DO measurement*
- ~~Previous Tevatron average~~
- ATLAS and LHCb measurements
- Recent CDF measurement [$M_W = (80.4335 \pm 0.0094)$ GeV]

before
after

Theory intrinsic uncertainties on input parameters

$$\delta_{\text{th}} M_W = 4 \text{ MeV}, \delta_{\text{th}} \sin^2 \theta_W = 5 \times 10^{-5}$$

$$\delta_{\text{th}} \Gamma_Z = 0.4 \text{ MeV}, \delta_{\text{th}} \sigma_{\text{had}}^0 = 6 \text{ pb}$$

$$\delta_{\text{th}} R_l^0 = 0.006, \delta_{\text{th}} R_c^0 = 0.00005$$

$$\delta_{\text{th}} R_b^0 = 0.0001$$

Using the PDG rescaling prescription

previous average

$$M_W = 80.379 \pm 0.012 \text{ GeV}$$

new average

$$M_W = 80.4133 \pm 0.0088 \text{ GeV}$$

“standard”

new average

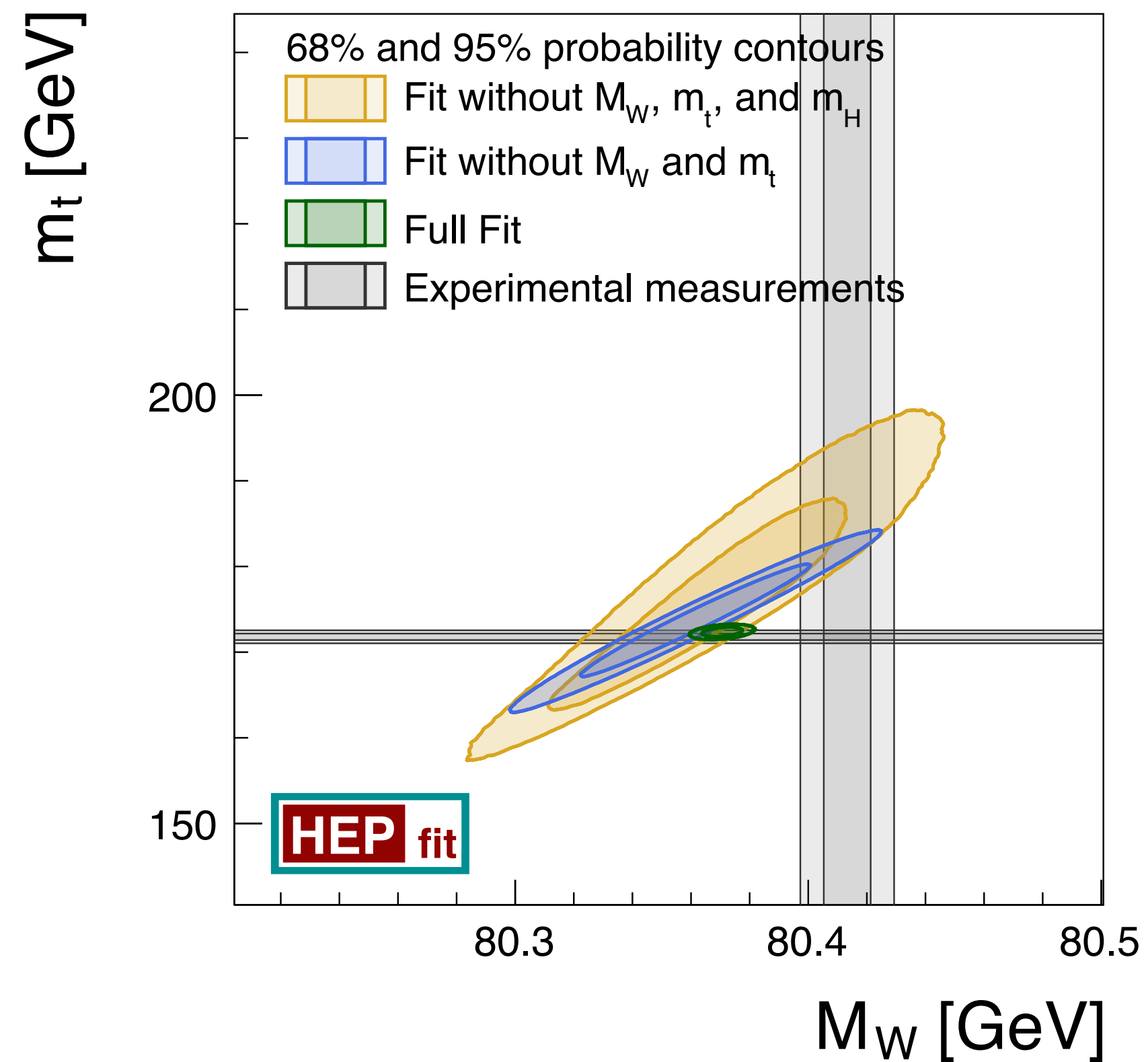
$$M_W = 80.4133 \pm 0.015 \text{ GeV}$$

“conservative”

Waiting for official Tevatron+LHC averages to be released

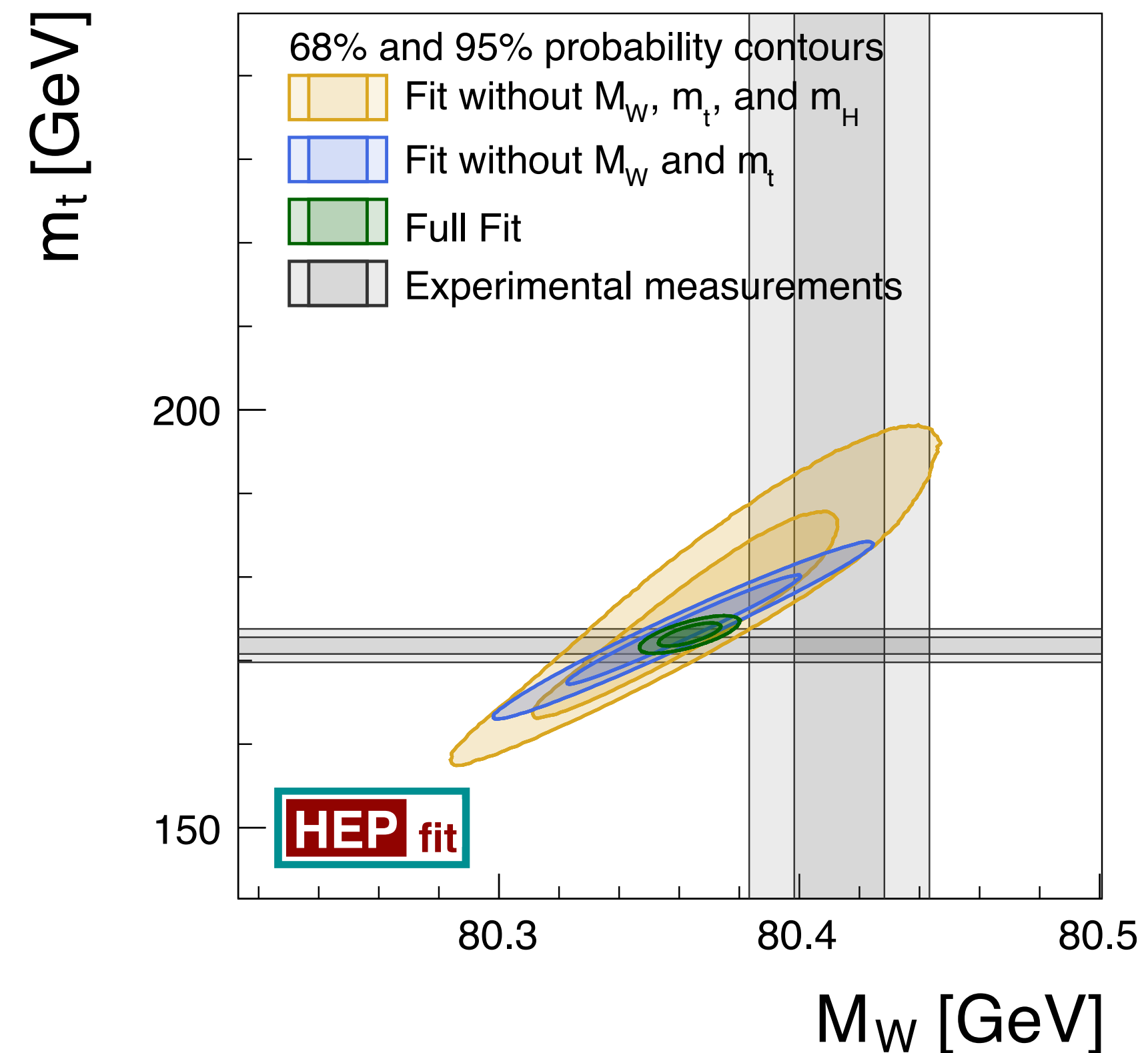
Impact of the new m_W and m_t measurements

Model	Pred. M_W [GeV] <i>standard average</i>	Pull	Pred. M_W [GeV] <i>conservative average</i>	Pull
SM	80.3499 ± 0.0056	6.5σ	80.3505 ± 0.0077	3.7σ



"standard" scenario

8

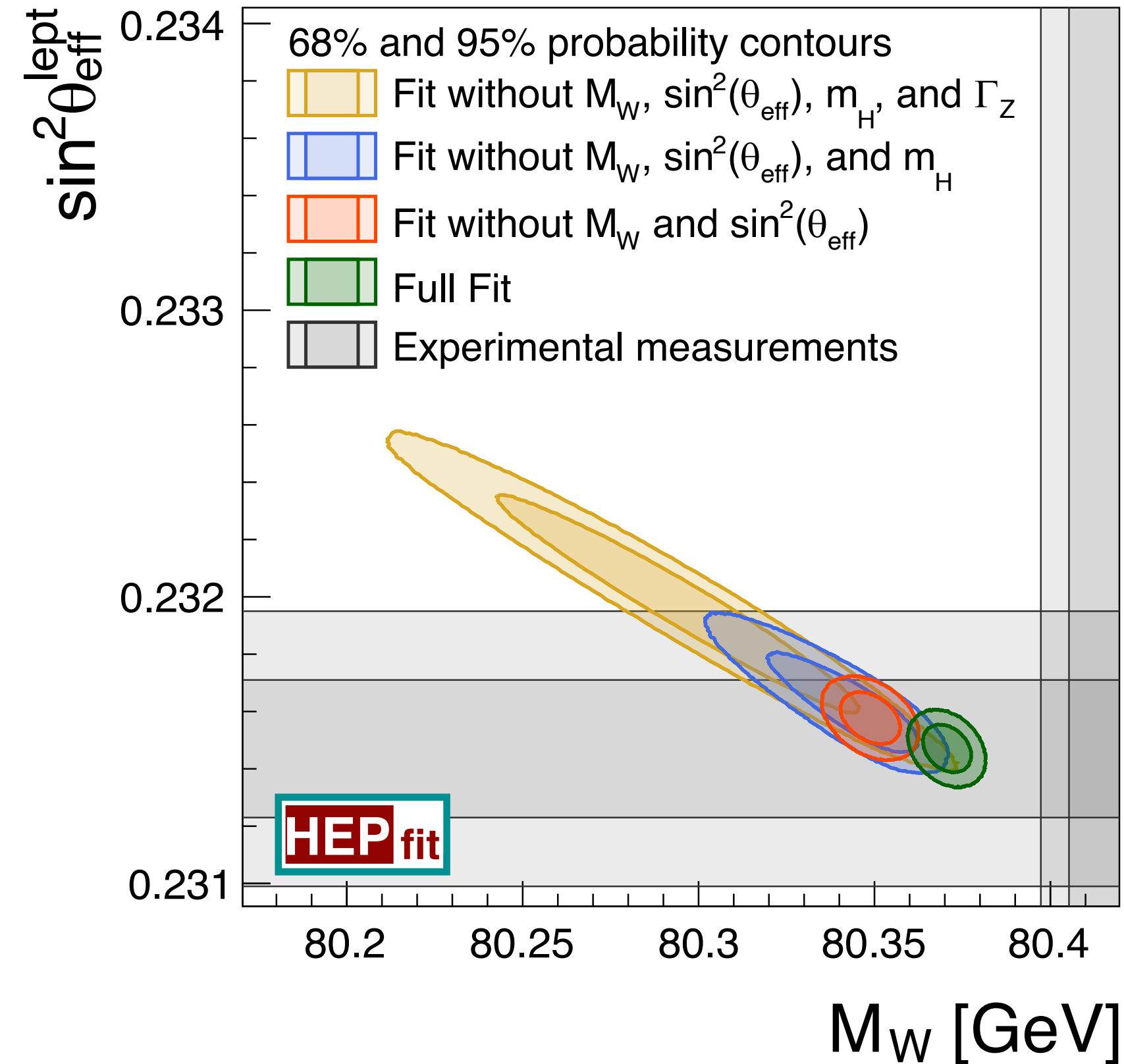


"conservative" scenario

HEPfit

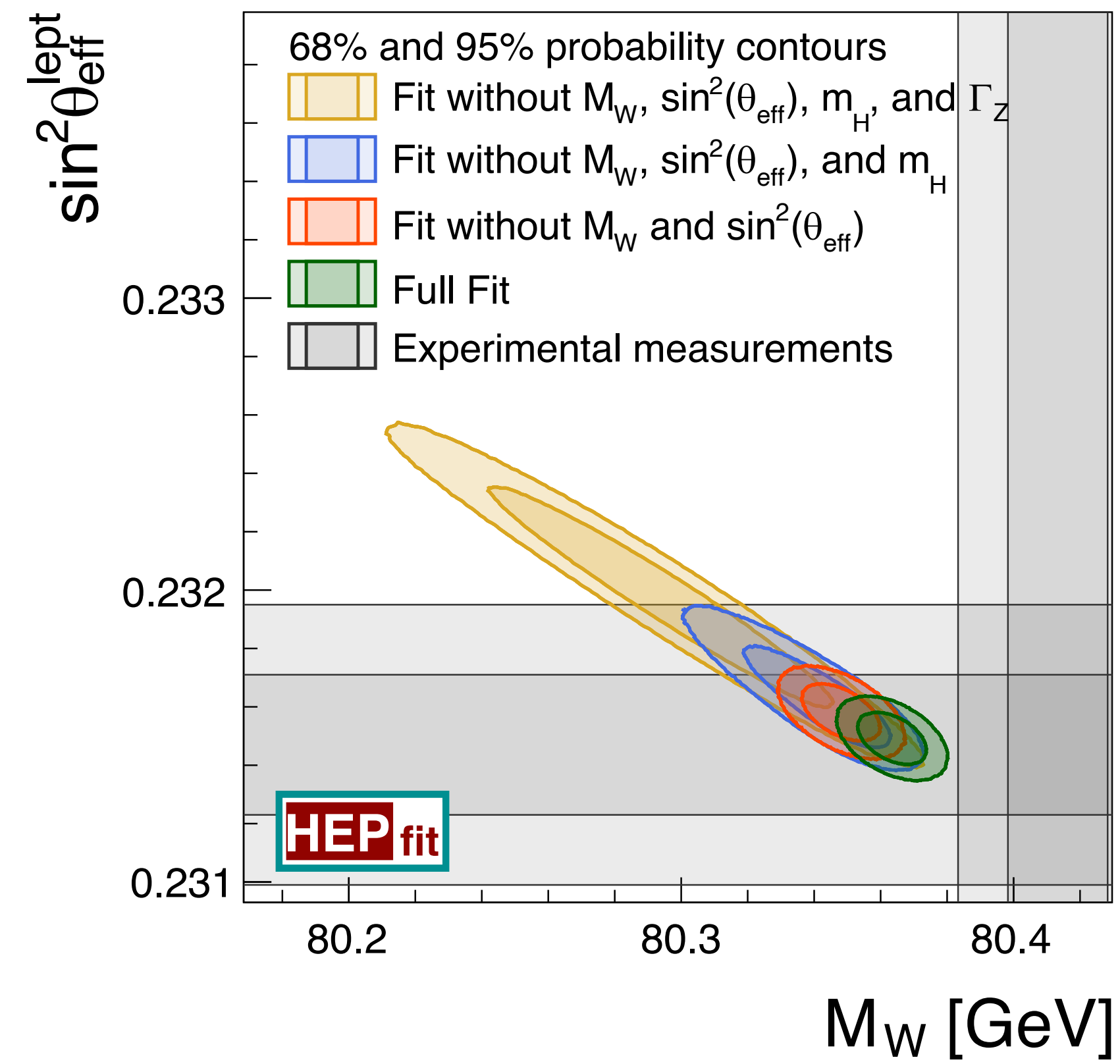
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SM	80.3499 ± 0.0056	6.5σ	80.3505 ± 0.0077	3.7σ



“standard” scenario

9



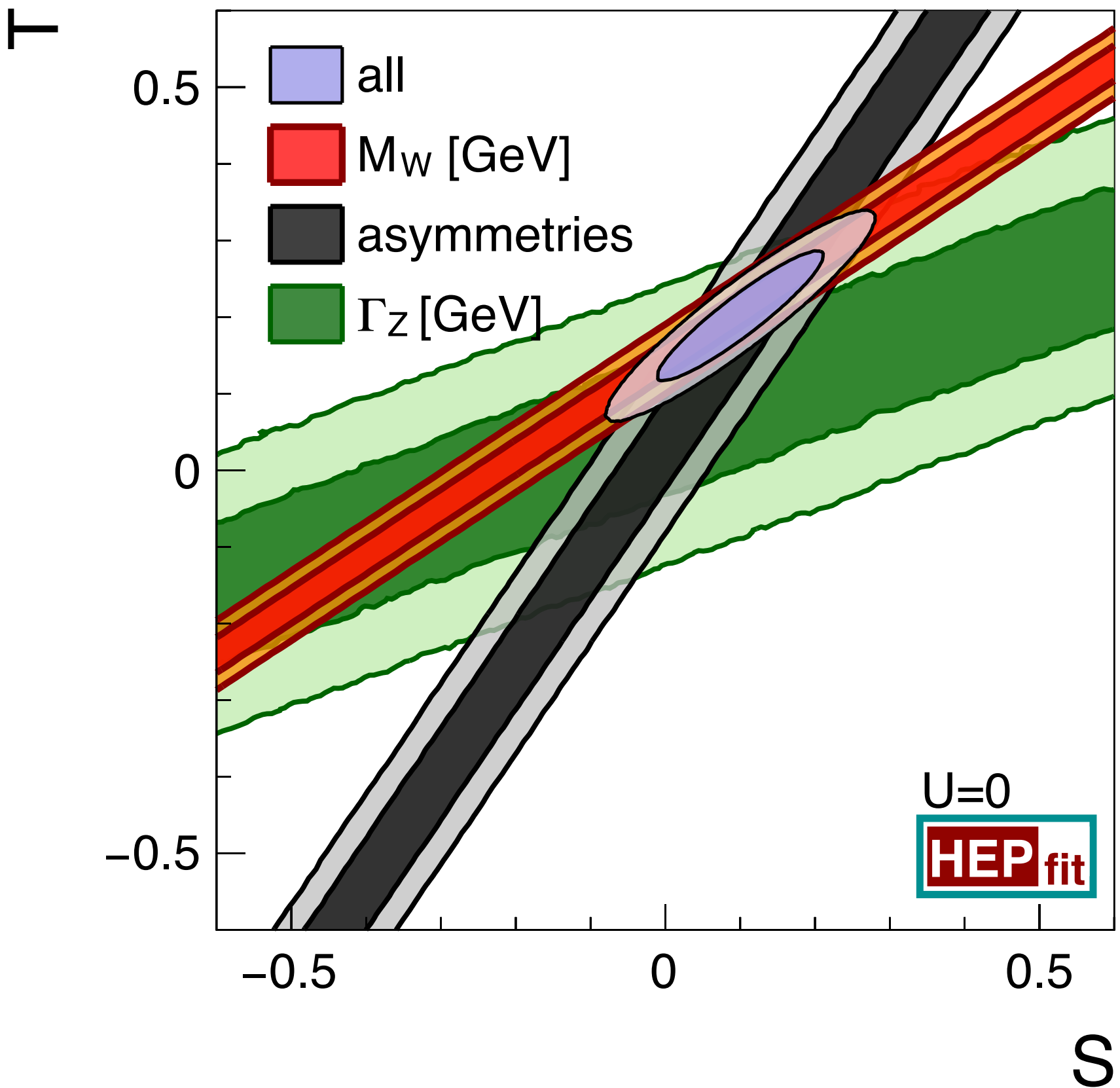
“conservative” scenario

Fitting Oblique Corrections: S, T (with U=0)

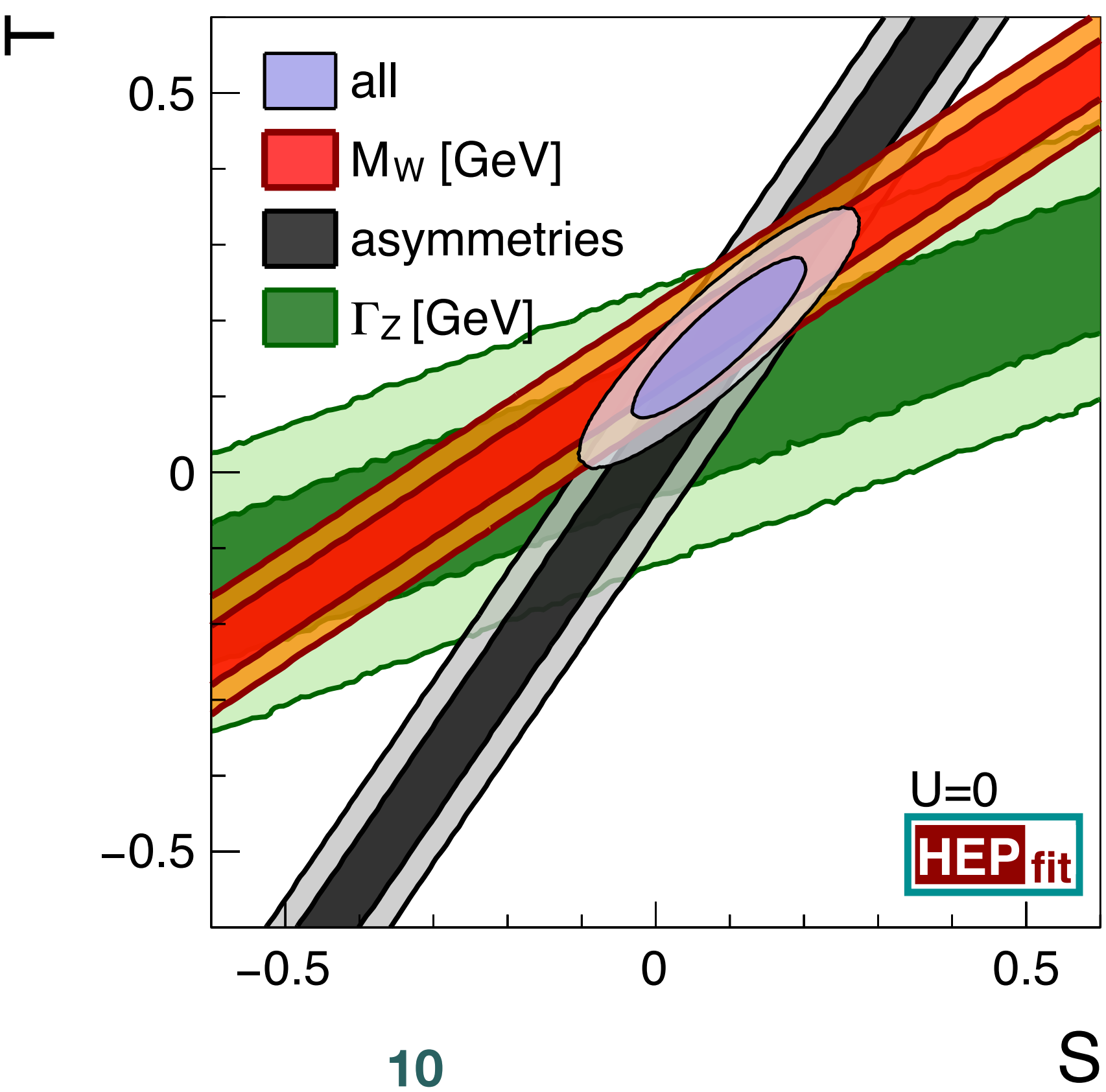
A large T value can
compensate for the W mass,
reducing the fit pull close to 1σ

Model	Pred. M_W [GeV]	Pull	Pred. M_W [GeV]	Pull
	<i>standard average</i>		<i>conservative average</i>	
SM	80.3499 ± 0.0056	6.5σ	80.3505 ± 0.0077	3.7σ
ST	80.366 ± 0.029	1.6σ	80.367 ± 0.029	1.4σ

“standard” scenario



“conservative” scenario

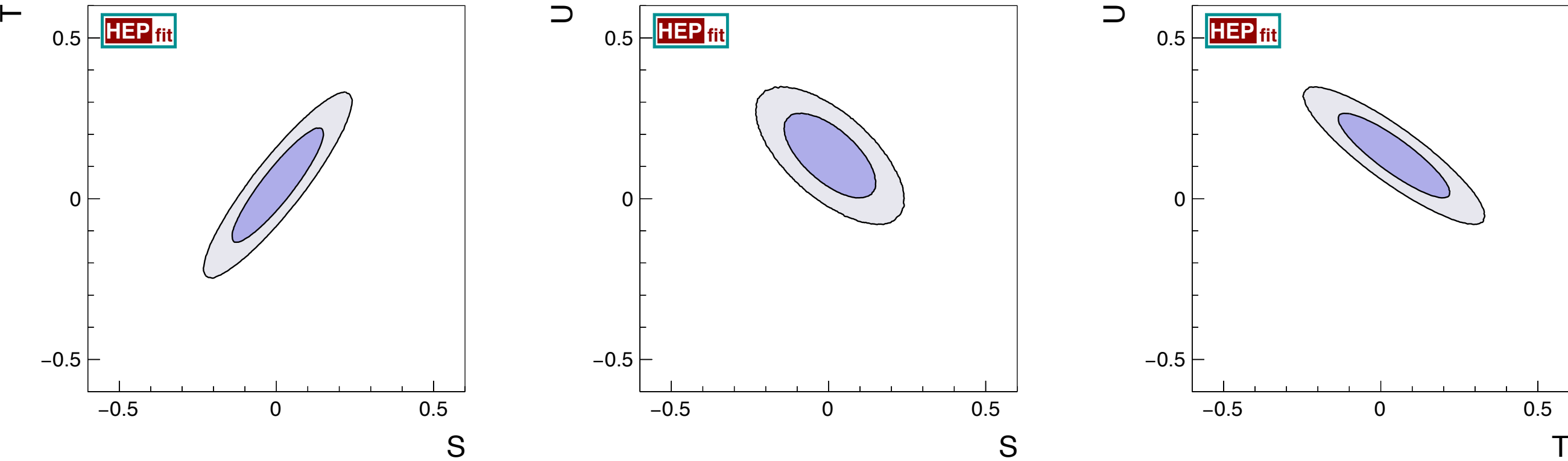


	Result	Correlation	
	(IC _{ST} /IC _{SM} = 25.0/80.2)		
S	0.100 ± 0.073	1.00	
T	0.202 ± 0.056	0.93	1.00
U	—	—	—

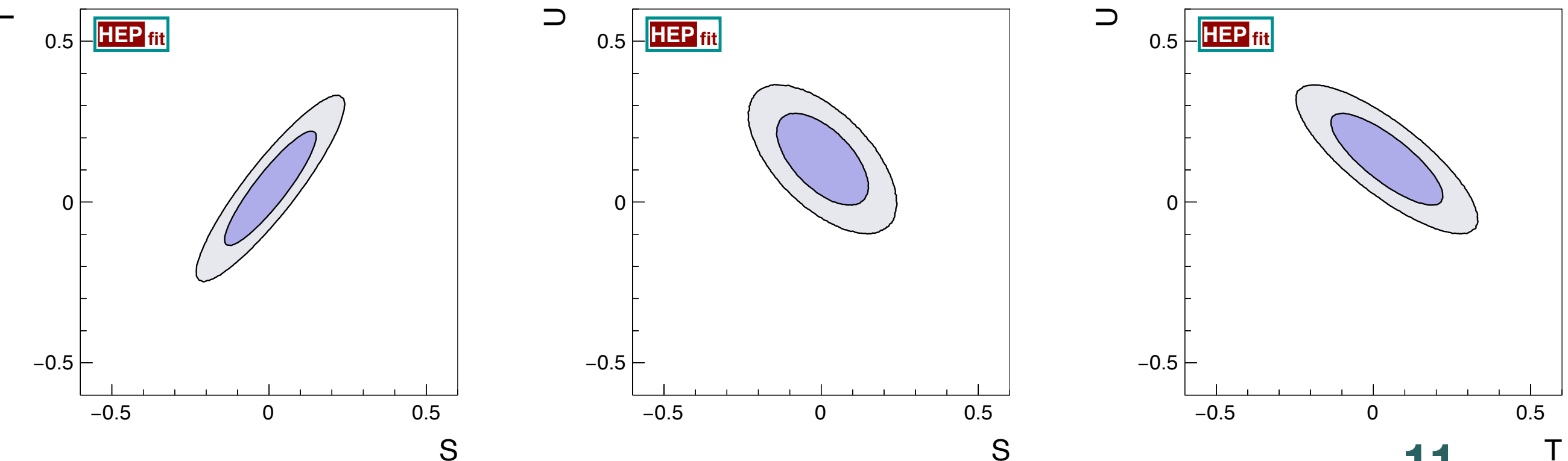
Fitting Oblique Corrections: S, T, U

A large U value can compensate for the W mass, without pulling S and T

“standard” scenario



“conservative” scenario



Model	Pred. M_W [GeV]	Pull	Pred. M_W [GeV]	Pull
	<i>standard average</i>		<i>conservative average</i>	
SM	80.3499 ± 0.0056	6.5σ	80.3505 ± 0.0077	3.7σ
ST	80.366 ± 0.029	1.6σ	80.367 ± 0.029	1.4σ
STU	80.32 ± 0.54	0.2σ	80.32 ± 0.54	0.2σ

Result	Correlation		
(IC _{STU} /IC _{SM} = 25.3/80.2)			
0.005 ± 0.096	1.00		
0.040 ± 0.120	0.91	1.00	
0.134 ± 0.087	−0.65	−0.88	1.00

Beyond Oblique Corrections: SMEFT analysis

$$\mathcal{O}_{\phi l}^{(1)} = (\phi^\dagger i \overleftrightarrow{D}_\mu \phi) (\bar{l}_L \gamma^\mu l_L) ,$$

$$\mathcal{O}_{\phi l}^{(3)} = (\phi^\dagger i \overleftrightarrow{D}_\mu^i \phi) (\bar{l}_L \sigma_i \gamma^\mu l_L) ,$$

$$\mathcal{O}_{\phi e} = (\phi^\dagger i \overleftrightarrow{D}_\mu \phi) (\bar{e}_R \gamma^\mu e_R) ,$$

$$\mathcal{O}_{\phi q}^{(1)} = (\phi^\dagger i \overleftrightarrow{D}_\mu \phi) (\bar{q}_L \gamma^\mu q_L) ,$$

$$\mathcal{O}_{\phi q}^{(3)} = (\phi^\dagger i \overleftrightarrow{D}_\mu^i \phi) (\bar{q}_L \sigma_i \gamma^\mu q_L) ,$$

$$\mathcal{O}_{\phi u} = (\phi^\dagger i \overleftrightarrow{D}_\mu \phi) (\bar{u}_R \gamma^\mu u_R) ,$$

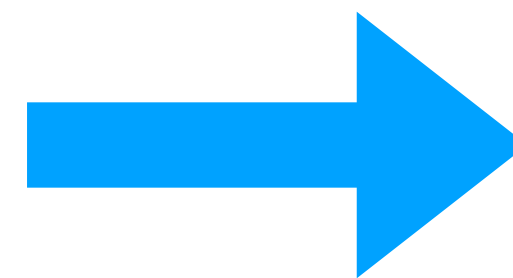
$$\mathcal{O}_{\phi d} = (\phi^\dagger i \overleftrightarrow{D}_\mu \phi) (\bar{d}_R \gamma^\mu d_R) ,$$

$$\mathcal{O}_{\phi WB} = (\phi^\dagger \sigma_i \phi) W_{\mu\nu}^i B^{\mu\nu} ,$$

$$\mathcal{O}_{\phi D} = (\phi^\dagger D^\mu \phi)^* (\phi^\dagger D_\mu \phi) ,$$

$$\mathcal{O}_{ll} = (\bar{l}_L \gamma^\mu l_L) (\bar{l}_L \gamma^\mu l_L)$$

zff/Wff
vertex
corrections



Only 8 independent combinations enter EWPO

$$\begin{aligned} \hat{C}_{\varphi f}^{(1)} &= C_{\varphi f}^{(1)} - \frac{Y_f}{2} C_{\varphi D}, \quad f = l, q, e, u, d, \\ \hat{C}_{\varphi f}^{(3)} &= C_{\varphi f}^{(3)} + \frac{c_w^2}{4s_w^2} C_{\varphi D} + \frac{c_w}{s_w} C_{\varphi WB}, \quad f = l, q, \\ \hat{C}_{ll} &= \frac{1}{2} ((C_{ll})_{1221} + (C_{ll})_{2112}) = (C_{ll})_{1221}, \end{aligned}$$

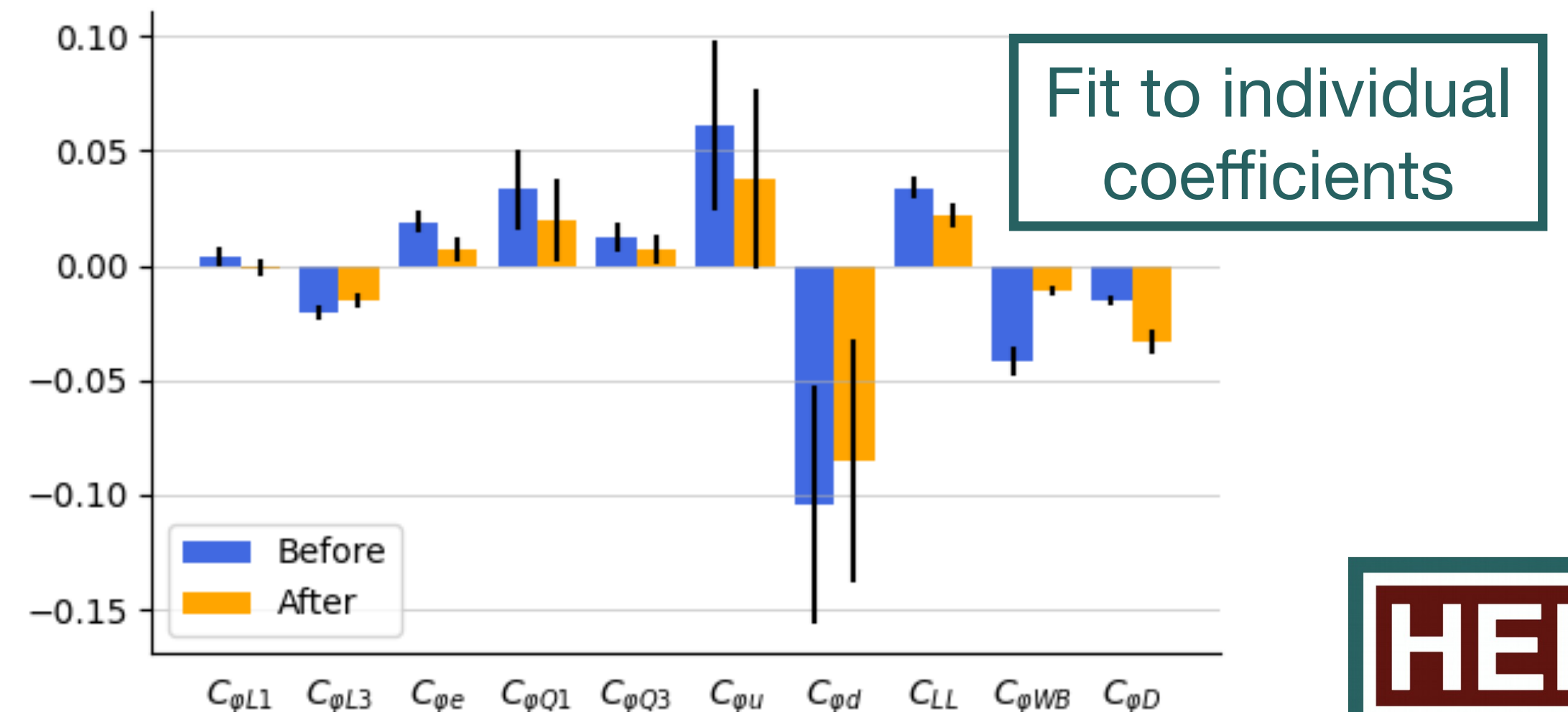
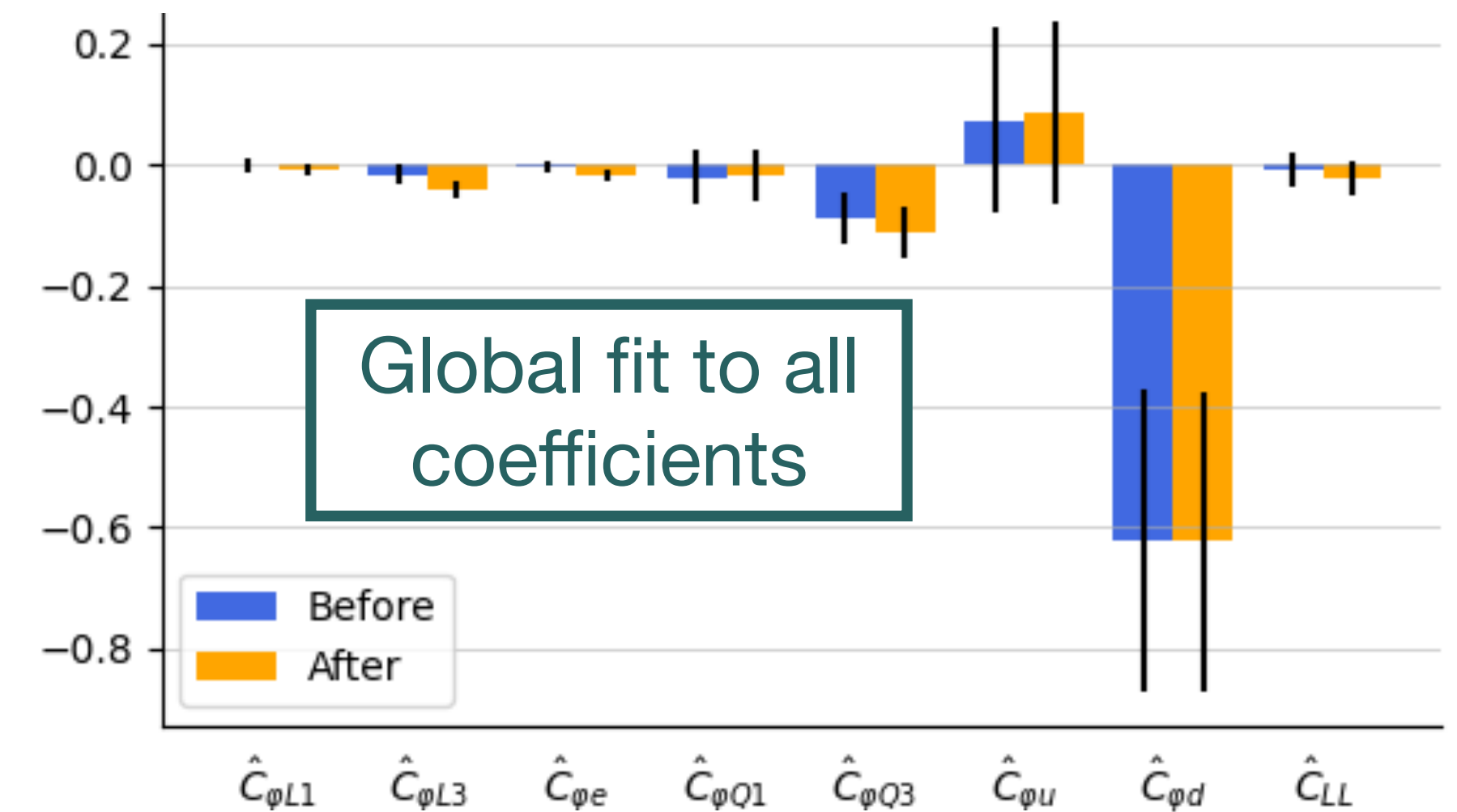
W/Z
propagators \longleftrightarrow S,T

G_F

A unique combination of coefficients enters

the W mass: $\left(\hat{C}_{\phi \ell}^{(3)} - \frac{\hat{C}_{\ell \ell}}{2} \right)$

Model	Pred. M_W [GeV] <i>standard average</i>	Pull	Pred. M_W [GeV] <i>conservative average</i>	Pull
SMEFT	80.66 ± 1.68	-0.1σ	80.66 ± 1.68	-0.1σ



Conclusions

- Since decades, the EW fit has been one of our most effective tools to test the SM and probe the presence of new physics at higher energy scale
- The recent update of m_W and m_t measurements represent a further step forward in precision
- But the new m_W measurement by CDF challenges the consistency of the fit
- Certainly, new physics effects can compensate for that
 - Invoking large T (if $U=0$) or U oblique correction
 - Adjusting the value of the $\hat{C}_{\phi\ell}^{(3)} - \frac{\hat{C}_{\ell\ell}}{2}$ combination of SMEFT coefficients
- On the other hand, this doesn't address the issue of compatibility between experimental measurements
 - Something that we tried to mitigate inflating the uncertainty on the average (PDG style)
 - Something that requires (ongoing) scrutiny by the experiments

A close-up shot of Bruce Banner, played by Mark Ruffalo, in a kitchen setting. He is wearing his signature black-rimmed glasses, a grey V-neck shirt, and a dark blue cardigan. He has a slightly open mouth as if speaking. In the background, a kitchen with stainless steel shelves, a computer monitor, and a person in a white chef's coat and blue hairnet is visible.

These are confusing times

Backup Slides

Top Mass combination

- To combine m_t measurements, we need some correlation model
 - we assume a linear correlation between systematic uncertainties from different measurements $\rho_{ij}^{sys} = \min(\sigma_i, \sigma_j) / \max(\sigma_i, \sigma_j)$ which results in the “standard average”
 $m_t = 171.79 \pm 0.38$
- Doing so, we observe a tension between some set of measurements (ATLAS and CMS l+jets)
 - By applying the PDG procedure, the uncertainty explodes to 1.7 GeV. We don’t think this reflects what we know about the W mass, so we discarded this value
 - Instead, we inflated the error up to 1 GeV (“conservative” scenario) which in any case has little impact on the fit (parametric uncertainties are subleading with respect to experimental errors on EWPO)