

Searching for Dark Matter in top quark production with the CMS experiment



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 $\widetilde{\chi}_1^0$

Motivation

 Cosmology suggests Dark Matter (DM) may couple to the Standard Model (SM) around the weak scale

- Search for spin-0 mediator with Yukawa coupling => large coupling to top quarks
- Interactions described by



Results

• Limits set on scalar (left) and pseudoscalar (right) mediator hypotheses for $g_a = q_y = 1$

• Limits independent of m_v for $m_v < m_{\phi/a}/2$



simplified Lagrangians [1] (e.g. for scalar mediator):

$$\mathcal{L}_{\phi} \supset - \frac{g_{\chi}}{2} \phi \bar{\chi} \chi - \frac{\phi}{2} \sum_{q=u,d,s,c,b,t} \frac{g_q y_q \bar{q} q}{2} - \frac{1}{2} \frac{m_{\phi}^2}{2} |\phi|^2 - m_{\chi} |\chi|^2$$

- Very similar final state to stop production in MSSM [2]
- Can also probe SM Higgs invisible decays



Analysis strategy

• Split into channels based on decay of top quarks: dileptonic (DL) [3], semileptonic (SL) [4] and all hadronic (AH) [5]

Future Plans

Single top + DM

• DM can also be produced in association with a single top quark

- Cross-section generally lower than for tt+DM, but becomes comparable for high mediator masses [6]
- Target with new categories requiring one fewer top quark and possible forward jets





Neural Networks

- DL channel usually gives weakest limits due to low branching ratio and high SM tt background
- However "clean" 2 lepton final state gives access to more sensitive 35.9 fb⁻¹ (13 TeV)

-Spin correlation observables (especially sensitive for pseudoscalar)



-Single top observables help distinguish t+DM from SM tt -p,^{miss} assigned to DM in top reconstruction

• Train a neural network on these variables + $M_{\tau_2}(II)$ and $p_{\tau_2}^{miss}$

> [1] J. Abdallah et al., doi:10.1016/j.dark.2015.08.001, arXiv:1506.03116 [2] CMS Collaboration, doi:10.1140/epjc/s10052-021-09721-5, arXiv:2107.10892 [3] CMS Collaboration, doi:10.1140/epjc/s10052-020-08701-5, arXiv:2008.05936 [4] CMS Collaboration, doi:10.1007/JHEP05(2020)032, arXiv:1912.08887 [5] CMS Collaboration, doi:10.1103/PhysRevD.104.052001, arXiv:2103.01290 [6] D. Pinna et al., doi:10.1103/PhysRevD.96.035031,573 arXiv:1701.05195