

Toward the discovery of lepton creation with neutrinoless double- β decay

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Toward the discovery of lepton creation with $0\nu\beta\beta$ decay

Content:

1. Historical landscape
2. Particle physics theory and motivation → I'm not a theorist, please don't ask tough questions
3. Nuclear physics and implications → I know even less about nuclear theory! 🤔
4. Experimental aspects and methods
5. Recent and future experiments
6. Discovery probability and impact

Work developed in collaboration with:

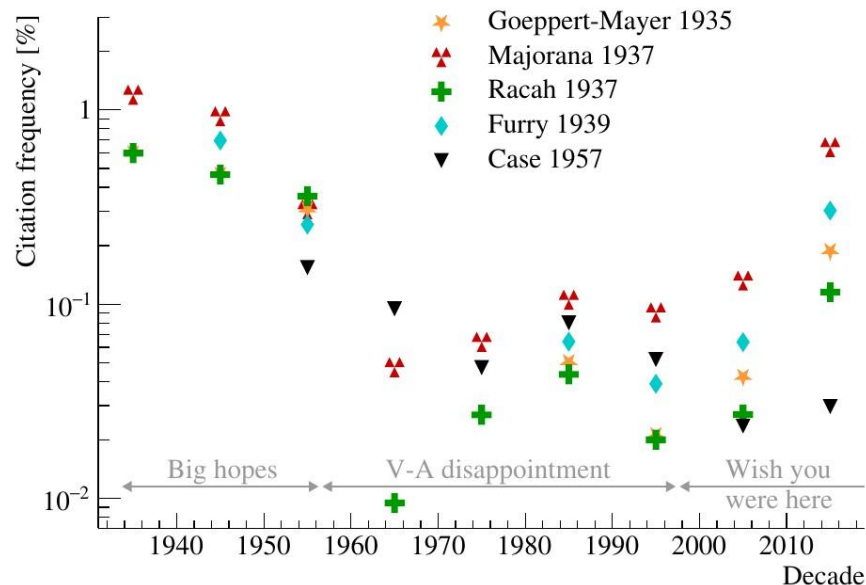
- M. Agostini, University College London
- J. Detwiler, University of Washington
- J. Menéndez, University of Barcelona
- F. Vissani, Laboratori Nazionali del Gran Sasso

Review paper foreseen to come out around mid December → Stay tuned!

Historical landscape: theoretical aspects

1920ies	Nucleus made of A protons and $(A-Z)$ <i>inner electrons</i> \rightarrow Monochromatic β spectrum
1930	Pauli proposes existence of neutrino <i>inside the nucleus</i> \rightarrow Non relativistic model
1934	Fermi theory of β decay \rightarrow Creation and destruction of matter particles
1935	Goeppert-Mayer applies Fermi theory to $\beta\beta$ decay
1937	Majorana theory: neutral elementary particles can be their own antiparticles
1938	Furry introduces $0\nu\beta\beta$ decay through emission and re-absorption of virtual Majorana ν
1950ies	(V-A) theory: $0\nu\beta\beta$ rate depends also on ν mass, not only on its nature \rightarrow Loss of interest
1960ies	Models for flavor transformation of massive neutrinos
~1980	Seesaw mechanisms
~2000	Definitive evidence of neutrino oscillations \rightarrow Mass mechanism as “standard” for $0\nu\beta\beta$ decay

Historical landscape: experimental aspects



- ~1950 Geochemical and direct experiments
- ~1970 First searches with Ge(Li) and scintillators
- ~1990 HPGe, TPCs, tracking detectors
- ~2000 O(10) kg HPGe sensitivity; KK claim 😲
- ~2010 O(100) kg HPGe, TPCs, liquid scintillators
- now moving towards the ton-scale!

Which are the fundamental symmetries and conserved quantities?

- B , L_e , L_μ , L_τ and $L=L_e+L_\mu+L_\tau$ are accidental symmetries that emerge without being *a priori* required
- Actually, these are anomalous symmetries spoiled by the full quantum theory
- Exact *global* symmetries of the standard model:

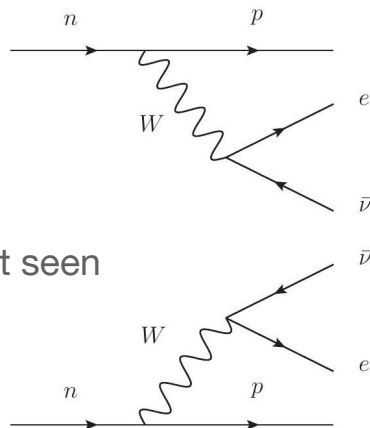
$L_e-L_\mu \rightarrow$ Violation seen in appearance mode by T2K

$L_\mu-L_\tau \rightarrow$ Violation seen in appearance mode by OPERA

$B-L \rightarrow$ We don't know yet

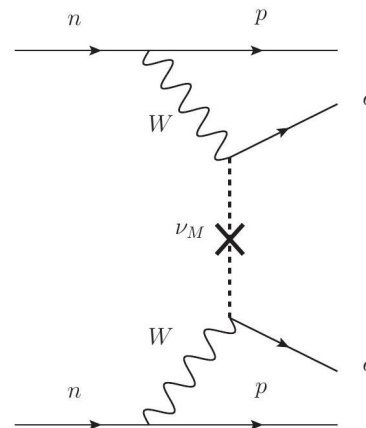
$2\nu\beta\beta$ DECAY

- \rightarrow SM process
- \rightarrow Very rare but seen



$0\nu\beta\beta$ DECAY

- \rightarrow BSM process
- \rightarrow Violates $B-L$
- \rightarrow Creates matter
- \rightarrow Yet unseen!



What generated the matter-antimatter asymmetry in the Universe?

- Non-perturbative effect violate $B+L$, but can't explain *quantitatively* the asymmetry
→ We need a *dynamical* explanation yielding the asymmetry at some point in history: *baryogenesis*
- Sakharov conditions for baryogenesis:
 - a. Violation of B (or $B-L$)
 - b. Violation of C and CP
 - c. Interactions out of thermal equilibrium
- Possible solution: baryogenesis through *leptogenesis*
→ Observing the violation of $B-L$ in the lepton sector would give a strong qualitative indication for the correctness of the baryogenesis hypothesis

What are the origin and nature of neutrino mass?

Why are neutrino masses so small?

Maybe, are they Majorana masses?

- Neutrino mass terms generated by odd-dimensional operators with $D \geq 5$
- $D=5$ case *only* implies an additional right-handed Majorana neutrino
 - Light neutrino masses generated through seesaw mechanism from the heavy Majorana one:

$$M_\nu = -M_D \cdot M_M^{-1} \cdot M_D$$

where:

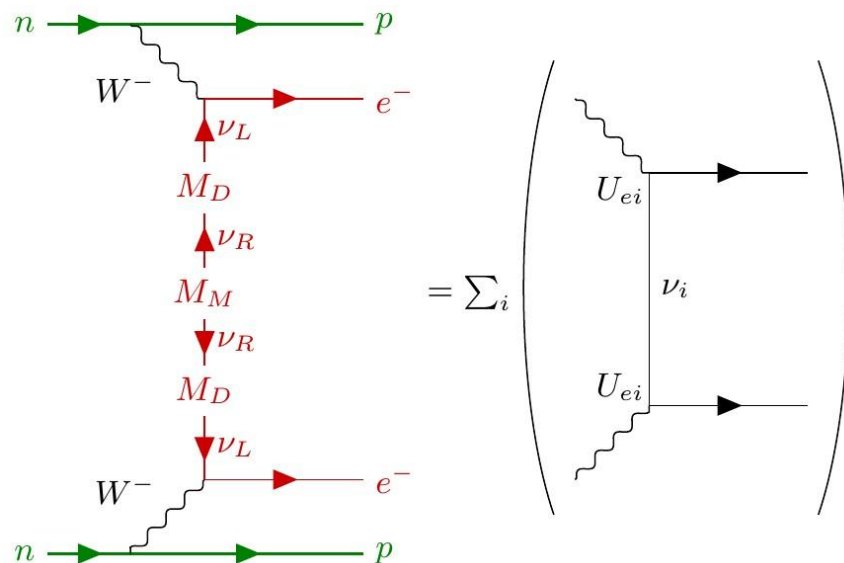
M_ν = Majorana mass for light neutrino

M_D = Dirac mass

M_M = Heavy Majorana neutrino mass

- In terms of neutrino oscillations:

$$m_{\beta\beta} = \left| \sum_{i=1}^3 |U_{ei}^2| e^{i\varphi_i} m_i \right|$$

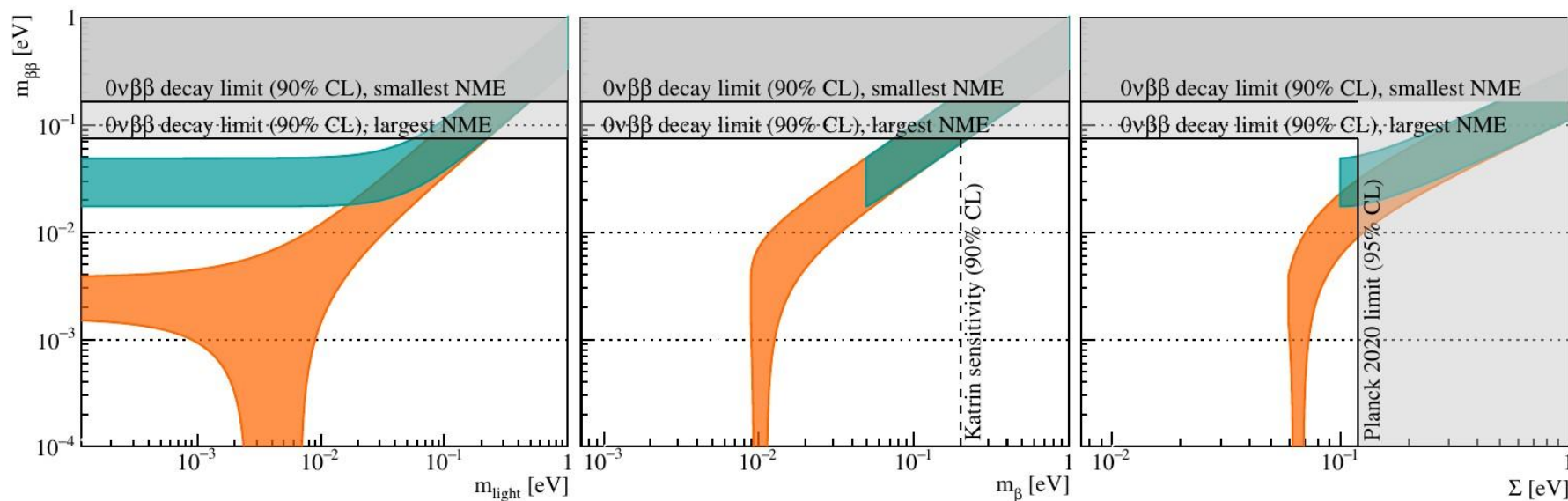


From $0\nu\beta\beta$ decay to the neutrino mass

- Connection between $0\nu\beta\beta$ decay rate and new physics terms:

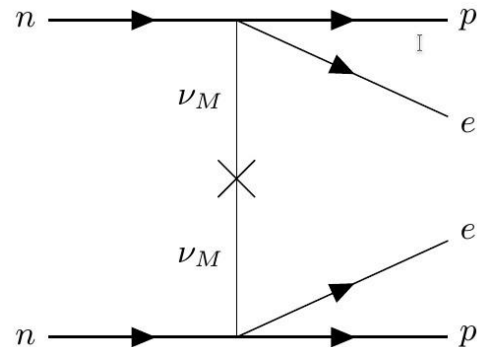
$$\frac{\Gamma_{0\nu}}{\ln 2} = \frac{1}{T_{1/2}^{0\nu}} = \sum_i G_i g_i^4 M_i^2 f_i(\Lambda) + \text{interference terms},$$

- In the case of D=5 (light neutrino exchange): $\frac{1}{T_{1/2}^{0\nu}} = G_{01} g_A^2 (M_{light}^{0\nu})^2 \frac{m_{\beta\beta}^2}{m_e^2}$



Nuclear physics theory and implications

- $0\nu\beta\beta$ decay triggered by BSM physics, but takes place in a highly correlated many-body system
- Nuclear Matrix Elements computed in two approximations:
 - Light: $m_\nu \ll p \sim 200 \text{ MeV}$
 - Heavy: $m_\nu \gg p \sim 200 \text{ MeV}$
- Different nuclear models yield NMEs differing by up to a factor 3
- New approaches worth remarking:
 - Ab-initio calculations
 - Good explanation of β decay
 - Not so good (yet) at predicting energy levels of heavy nuclei
 - Effective Field Theory
 - Compute contribution due to *any* odd-dimensional operator
 - Short-range exchange of high-energy light neutrinos might not be negligible

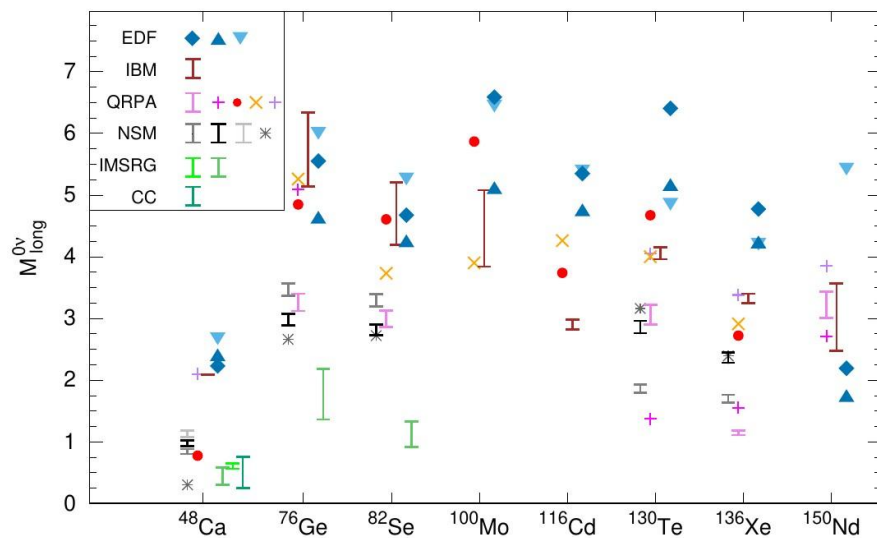


$$\begin{aligned}
 T_{1/2}^{-1} = & G_{01} g_A^4 (M_{\text{light}}^{0\nu})^2 \frac{m_{\beta\beta}^2}{m_e^2} \\
 & + \frac{m_N^2}{m_e^2} \tilde{G} \tilde{g}^4 \tilde{M}^2 \left(\frac{v}{\tilde{\Lambda}} \right)^6 \\
 & + \frac{m_N^4}{m_e^2 v^2} \tilde{G}' \tilde{g}'^4 \tilde{M}'^2 \left(\frac{v}{\tilde{\Lambda}'} \right)^{10} + \dots
 \end{aligned}$$

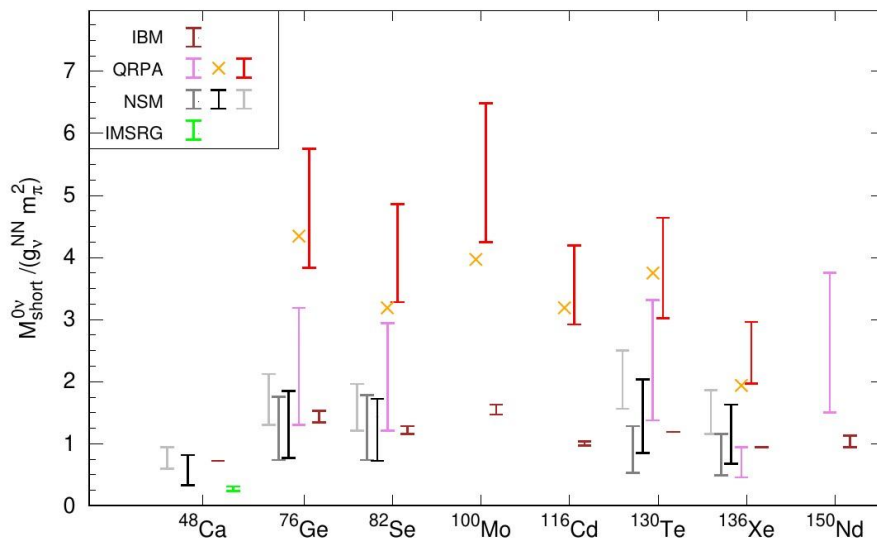
$$T_{1/2}^{-1} = G_{01} g_A^4 (M_{\text{long}}^{0\nu} + M_{\text{short}}^{0\nu})^2 \frac{m_{\beta\beta}^2}{m_e^2}$$

Nuclear physics theory and implications

Long-range, light neutrino exchange

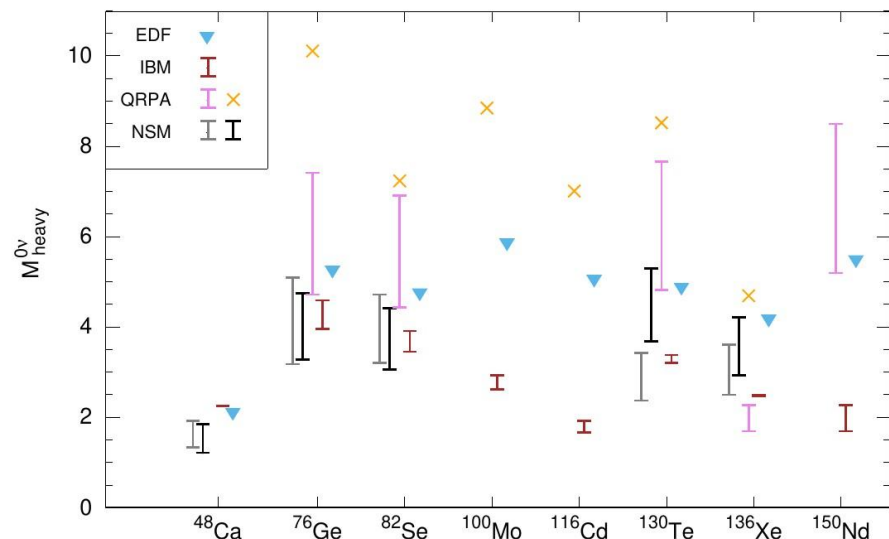


Short range, light neutrino exchange



Nuclear physics theory and implications

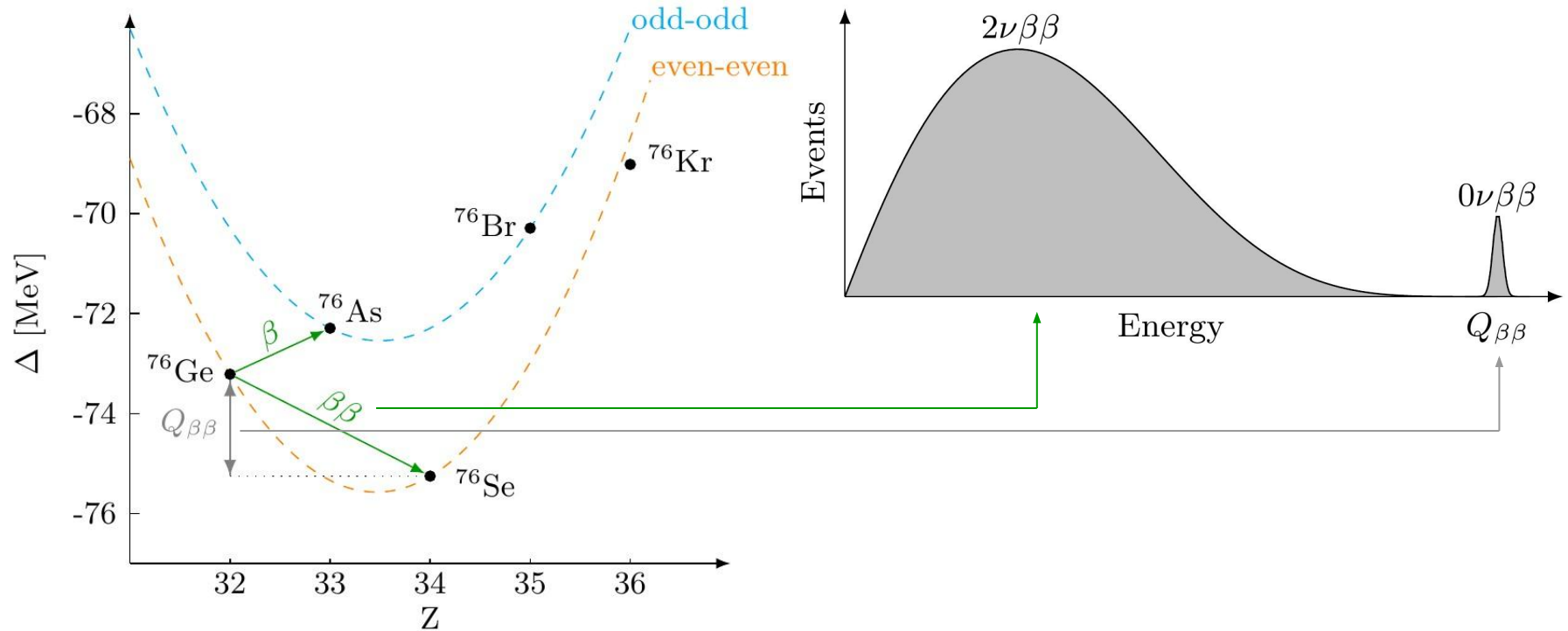
Heavy neutrino exchange



Possible experimental validations

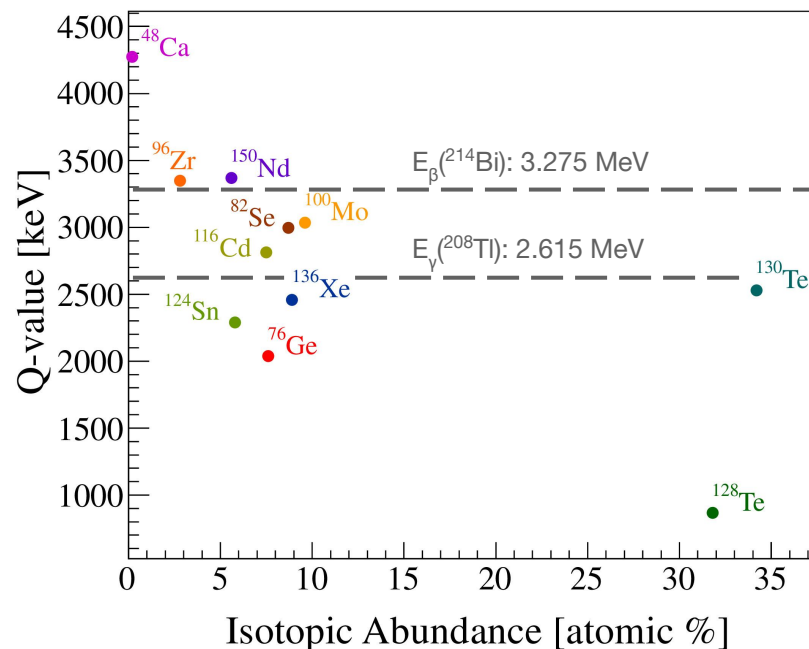
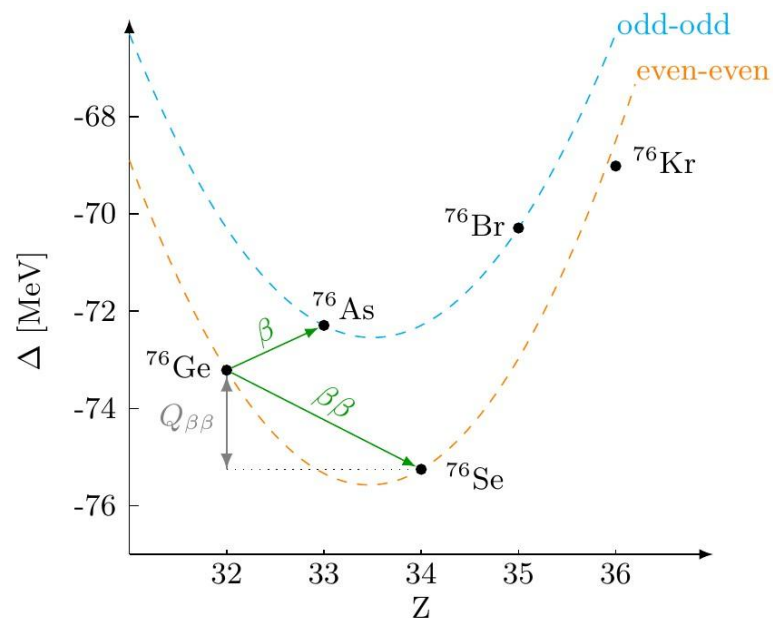
- Validation of NMEs for single β decay and $2\nu\beta\beta$ decay not enough, because of different exchanged momentum
- Ordinary muon capture: $\mu^- + p \rightarrow n + \nu_\mu$
- Inelastic neutrino-nucleon scattering
- Two-nucleon transfer reactions, e.g. $^{136}\text{Ba} (p,t) ^{138}\text{Ba}$
→ However this is a strong interaction
- Double charge-exchange reaction, e.g. $^{40}\text{Ca}(^{18}\text{O}, ^{18}\text{Ne})^{40}\text{Ar}$
→ However this is a strong interaction
- NMEs for Double Gamow-Teller transitions and $\gamma\gamma$ transitions possibly correlated to $0\nu\beta\beta$ decay NME

Experimental aspects: $0\nu\beta\beta$ decay signature



- $0\nu\beta\beta$ decay can be searched only with isotopes for which the regular β decay is forbidden
- $0\nu\beta\beta$ decay signature is an excess at the Q -value ($Q_{\beta\beta}$) in the sum electron spectrum
- $2\nu\beta\beta$ decay yields a continuum up to $Q_{\beta\beta}$

Isotope selection



- High Q-value ($Q_{\beta\beta}$) is highly preferable:
 - Larger phase space
 - Lower background from natural radioactivity
- High isotopic abundance preferable → Easier enrichment

Isotope selection

Isotope	Daughter	$Q_{\beta\beta}^a$ [keV]	f_{nat}^b [%]	f_{enr}^c [%]	$T_{1/2}^{2\nu\beta\beta d}$ [yr]	$T_{1/2}^{0\nu\beta\beta e}$ [yr]
^{48}Ca	^{48}Ti	4 267.98(32)	0.187(21)	16	$(6.4_{-0.6}^{+0.7}(\text{stat})_{-0.9}^{+1.2}(\text{syst})) \cdot 10^{19}$	$> 5.8 \cdot 10^{22}$
^{76}Ge	^{76}Se	2 039.061(7)	7.75(12)	92	$(1.926 \pm 94) \cdot 10^{21}$	$> 1.8 \cdot 10^{26}$
^{82}Se	^{82}Kr	2 997.9(3)	8.82(15)	96.3	$(8.60 \pm 0.03(\text{stat})_{-0.13}^{+0.19}(\text{syst})) \cdot 10^{19}$	$> 3.5 \cdot 10^{24}$
^{96}Zr	^{96}Mo	3 356.097(86)	2.80(2)	86	$(2.35 \pm 0.14(\text{stat}) \pm 0.16(\text{syst})) \cdot 10^{19}$	$> 9.2 \cdot 10^{21}$
^{100}Mo	^{100}Ru	3 034.40(17)	9.744(65)	99.5	$(7.12_{-0.14}^{+0.18}(\text{stat}) \pm 0.10(\text{syst})) \cdot 10^{18}$	$> 1.5 \cdot 10^{24}$
^{116}Cd	^{116}Sn	2 813.50(13)	7.512(54)	82	$2.63_{-0.12}^{+0.11} \cdot 10^{19}$	$> 2.2 \cdot 10^{23}$
^{130}Te	^{130}Xe	2 527.518(13)	34.08(62)	92	$(7.71_{-0.06}^{+0.08}(\text{stat})_{0.15}^{+0.12}(\text{syst})) \cdot 10^{20}$	$> 3.2 \cdot 10^{25}$
^{136}Xe	^{136}Ba	2 457.83(37)	8.857(72)	90	$(2.165 \pm 0.016(\text{stat}) \pm 0.059(\text{syst})) \cdot 10^{21}$	$> 1.1 \cdot 10^{26}$
^{150}Nd	^{150}Sm	3 371.38(20)	5.638(28)	91	$(9.34 \pm 0.22(\text{stat})_{-0.60}^{+0.62}(\text{syst})) \cdot 10^{18}$	$> 2.0 \cdot 10^{22}$

- Highest sensitivities achieved with isotopes that:
 - are easier to enrich
 - can be used in a well-established detector technology
- Detector technology plays primary role through several parameters:
 - energy resolution
 - detector material purity
 - background suppression

Detector technologies

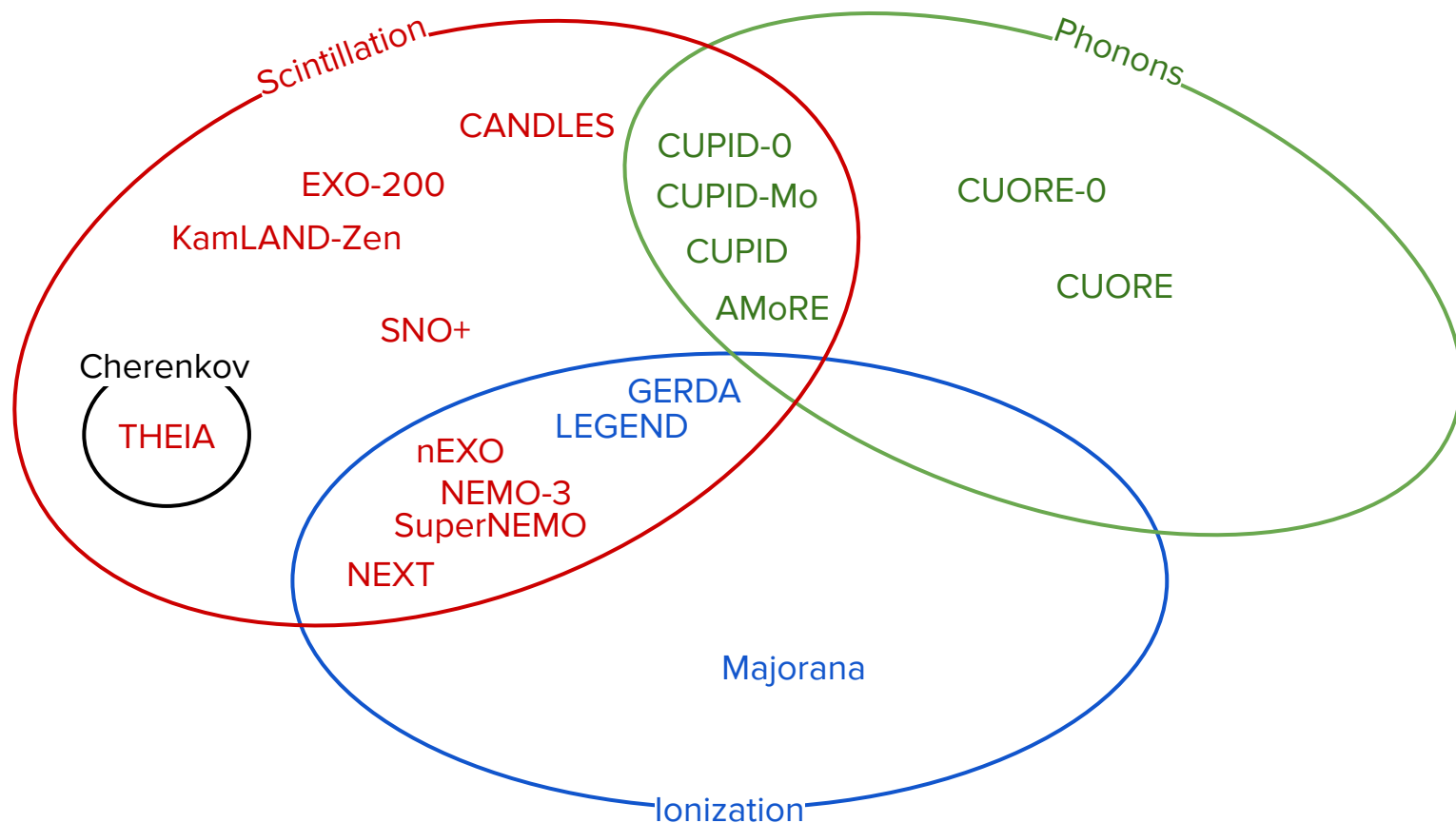
Detector type	Description	Typical isotope mass	Containment efficiency	Pros and cons
Solid state granular	Crystals containing $\beta\beta$ isotope	0.1-1 kg (single crystal)	70-95%	+ Scalable + High resolution - Many crystals required
Monolithic liquid or gaseous	Material = isotope or isotope dissolved in liquid	100-1000 kg	100%	+ Single detector + Self shielding - Poor resolution
Composite	Foil with isotope in low-pressure gas tracker	10 kg	~50%	+ Ultra-low background + Multiple isotopes - Very hard to scale

Detector technologies

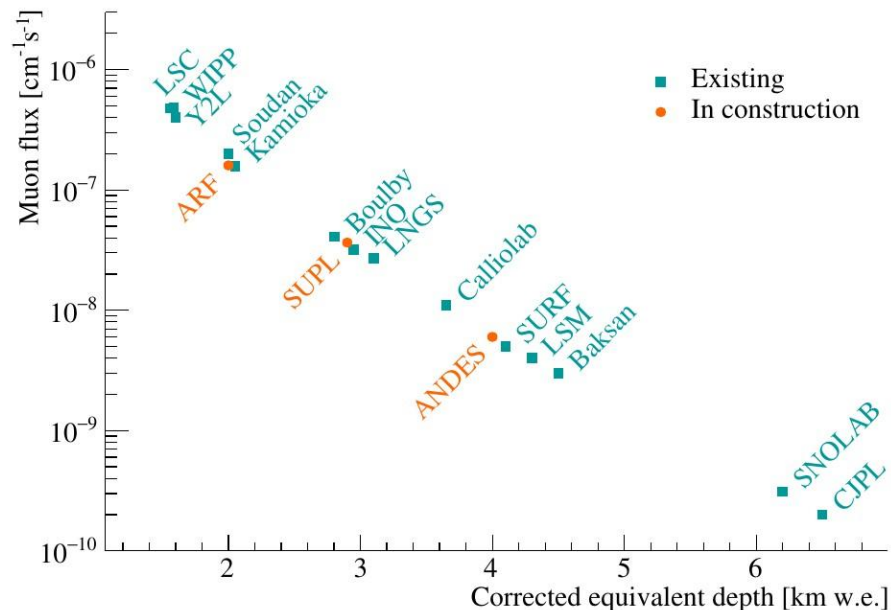
Readout channel	Energy resolution	Particle identification	Sensitivity to position	Applicable to multiple isotopes
Ionization	0.1-1%	Only in gas	Yes	Not really
Phonons	~0.2%	Nope	Nope	Yes
Scintillation	Few %	α vs β	In liquids and gases	Yes
Cherenkov	Forget it!	Visible only for β 's	Maybe	Yes

- Take-away messages:
 - The best detector technology doesn't exist
 - Combined readout channels are VERY helpful
→ good resolution with one, background suppression with the other

$0\nu\beta\beta$ decay experimental fauna



Cosmic ray background

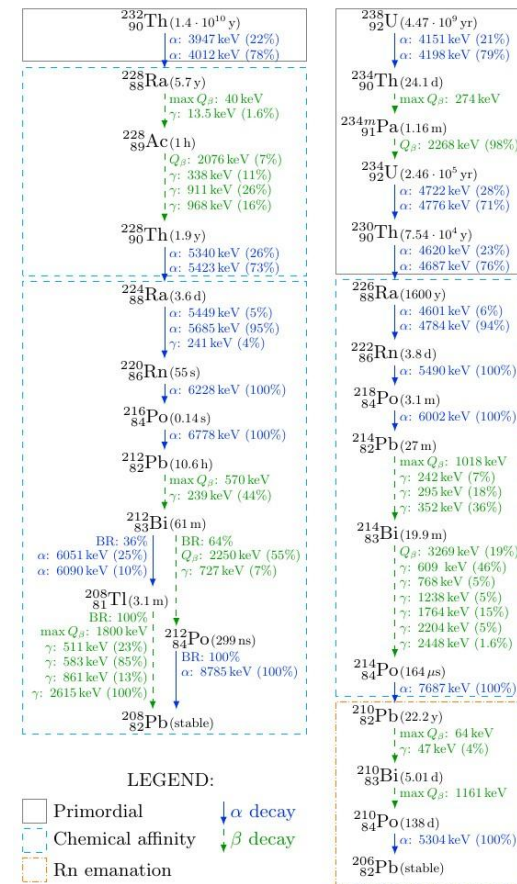


Residual backgrounds from cosmic in underground experiments:

- High energy muons (up to TeV)
 - Reconstruct muon tracks (monolithic)
 - Muon veto around granular detector
- Spallation products
 - Activate isotopes in the detector material prior to the installation underground
 - Activation in situ
 - Relevant for large volume scintillators
 - If the activated isotope decays quickly, search for delayed coincidences
 - High energy spallation neutrons
 - Well, this is a problem!

Background from actinides contamination

- Uranium and thorium contamination present in **many** materials
- Several particle types involved:
 - α between 4 and 9 MeV
 - β up to 3.3 MeV
 - γ up to 2.6 MeV, but summation possible
 - neutrons from (α ,n) reactions
- Decay chains not always in equilibrium
 - Material exposure to air
→ ^{222}Rn deposition followed by ^{210}Pb accumulation
 - Surface cleaning
→ Pb removed, ^{210}Po remains
 - Mechanical or chemical processes in material bulk or surface
→ Accumulation of “chemically active” radium
- Possible suppression techniques:
 - Use cleaner materials and surfaces
 - Minimize exposure to radon
 - Particle identification to reject α
 - Event topology
 - Delayed coincidences



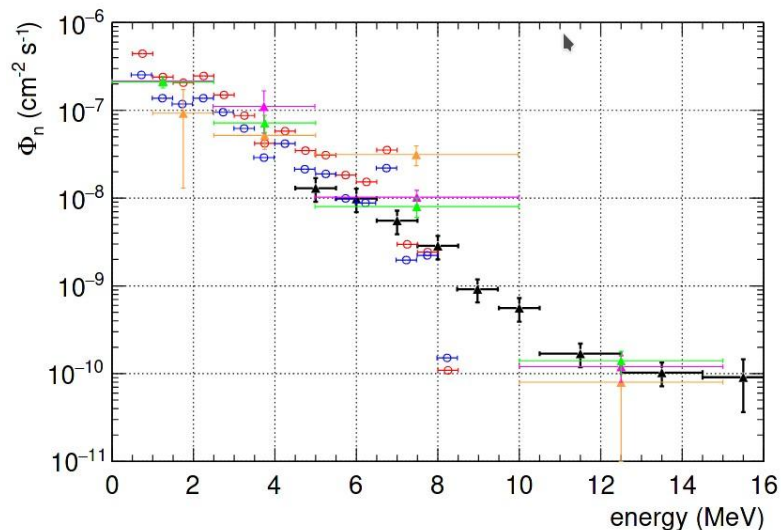
Man-made isotopes

- Several man-made isotopes visible in ultra-low background experiments
- Background only from:
 - Isotopes decaying β with $Q_\beta > Q_{\beta\beta}$
 - Isotopes or decay chains with dominant half-life comparable to experiment lifetime
- So far, only ^{110m}Ag has been found

Isotope	Half life	Q_β [keV]	Detected	Notes
^{88}Y	107 d	3008	No	Several γ lines
^{90}Sr	28.8 y	546	No	
^{90}Y	64 h	2279	No	Pure β emitter
^{110m}Ag	250 d	3008	Yes	Several γ lines
^{134}Cs	2 y	2059	No	Several γ lines
^{144}Ce	285 d	319	No	
^{144}Pr	17.3 m	2997	No	Pure β emitter

Neutrons

Source	Location	Energy
^{238}U fission	Concrete or internal	<10 MeV
(α ,n) reactions	Concrete or internal	<10 MeV
Spallation	Rock, concrete, ...	up to GeV



Neutron suppression

- Passive shielding (polyethylene, boron, water) → good for neutrons <10 MeV
- Active shielding in liquid scintillator outer layer
- Add element with high neutron cross section (e.g. ^6Li) in active volume

Neutron backgrounds

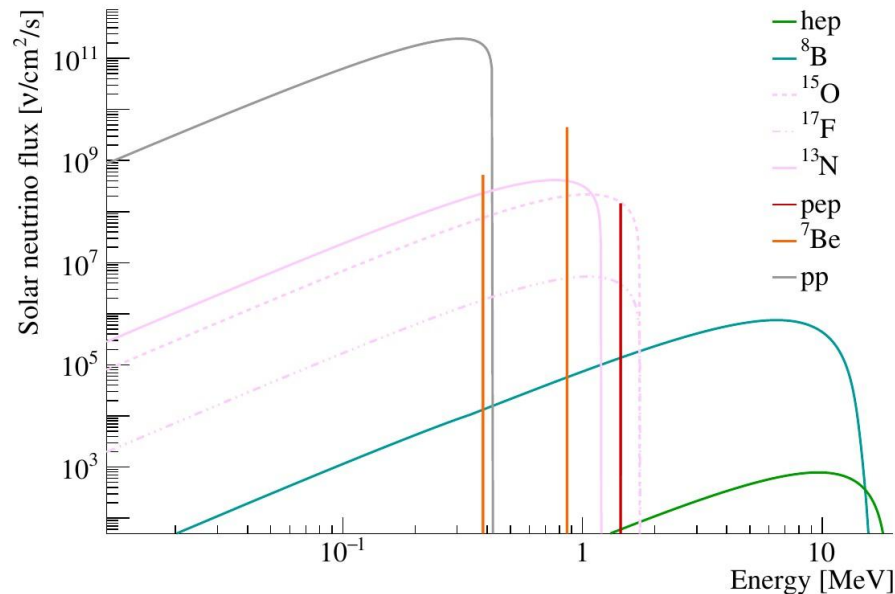
- In-situ isotope activation
- γ 's from inelastic scattering or captures

Solar neutrinos

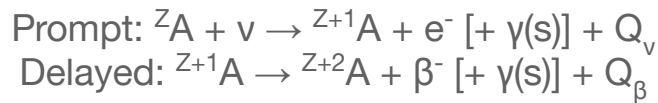
Elastic scattering

$$\nu + e^- \rightarrow \nu + e^-$$

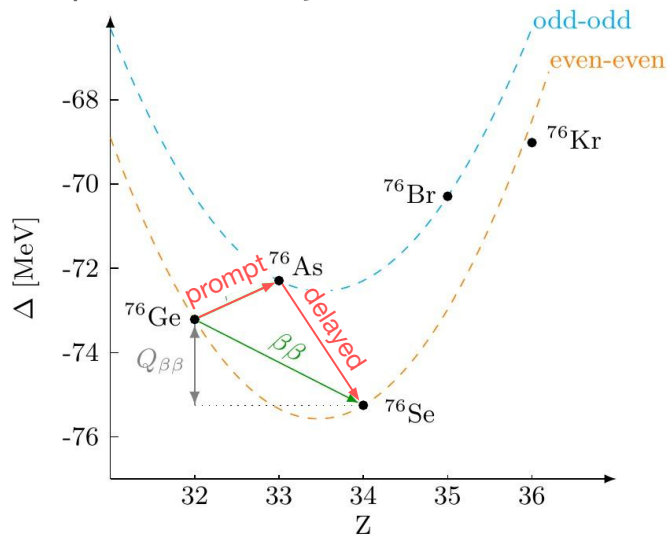
- Only relevant for large scintillators
- Can be suppressed through signal directionality



Charge Current

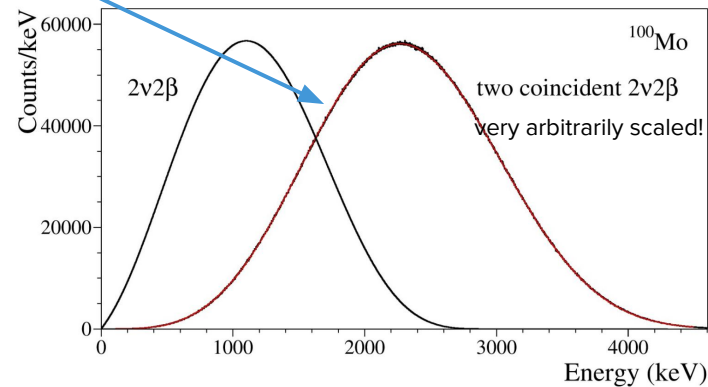
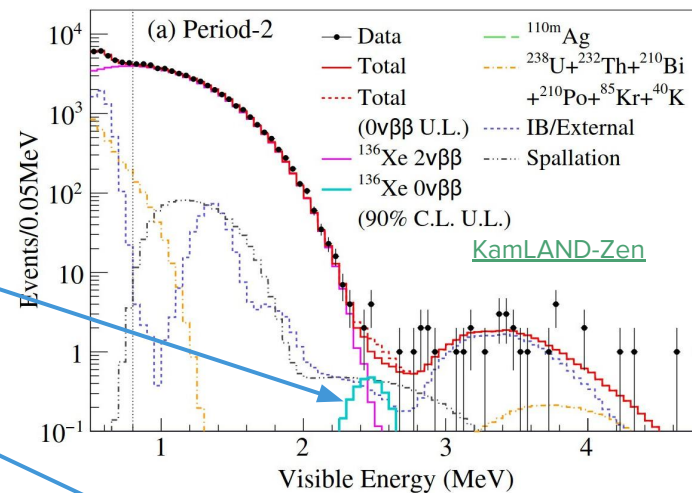


- Can be suppressed via topology and delayed coincidences, depending on isotope
- Not quite relevant, yet



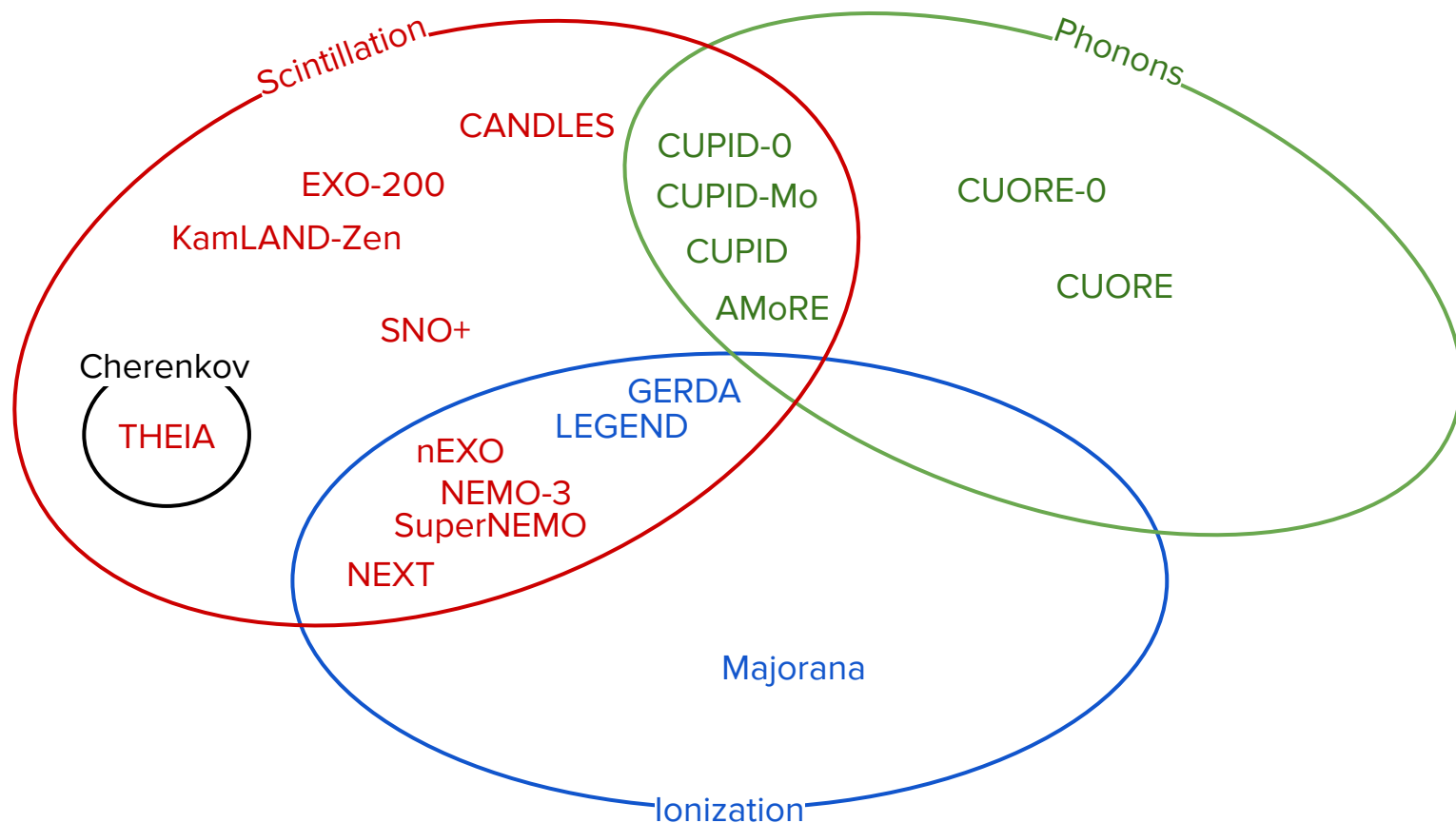
$2\nu\beta\beta$ decay

- Irreducible $2\nu\beta\beta$ background
 - Tail of $2\nu\beta\beta$ spectrum
⇒ Energy resolution
 - Pile-up of $2\nu\beta\beta$ events
⇒ Time resolution

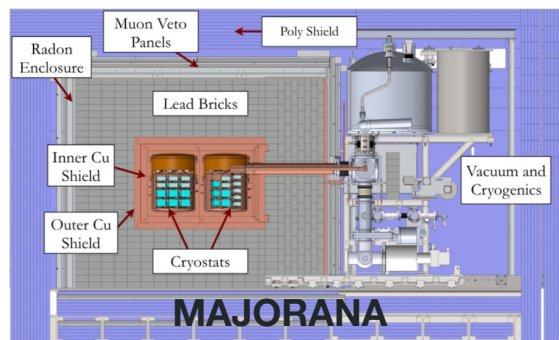


D. M. Cherniak et al., EPJ C72 (2012) 1989

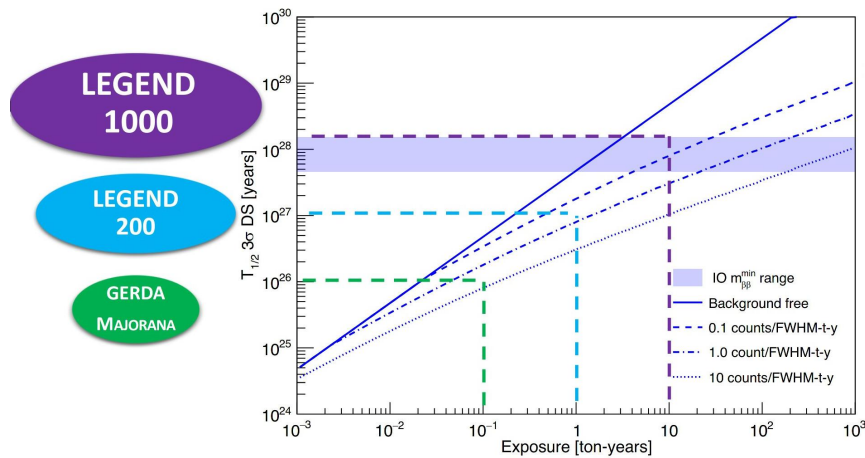
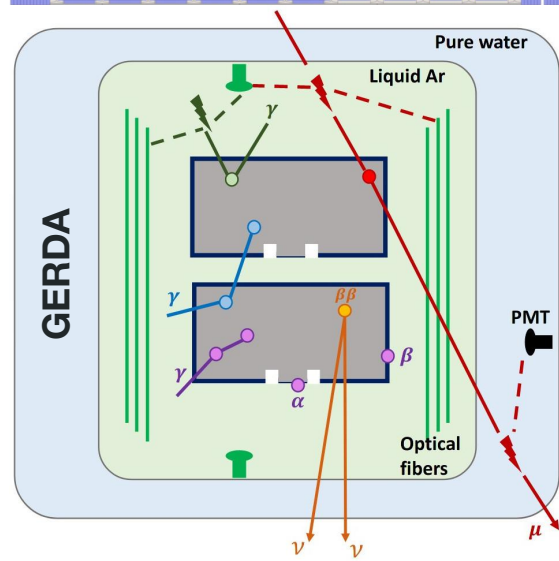
$0\nu\beta\beta$ decay experimental fauna



Germanium experiments

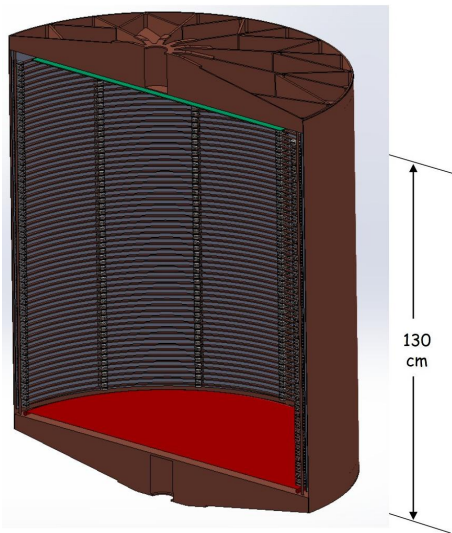
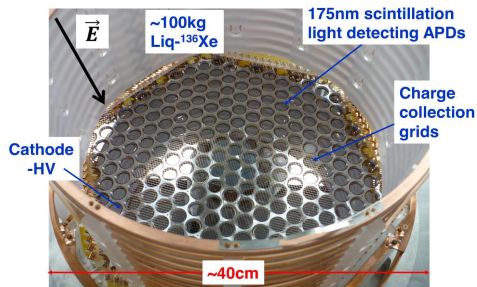


- Low Q-value: 2039 keV
- Highest energy resolution: $\sim 0.1\%$
- Extremely low bkg: $\sim 5 \cdot 10^{-4}$ counts/keV/kg/yr
→ Operating next to linear sensitivity regime
- MAJORANA + GERDA joining for next generation experiment: LEGEND

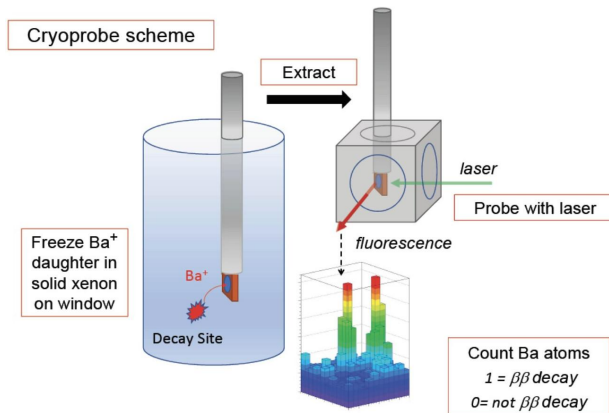


Xenon TPCs

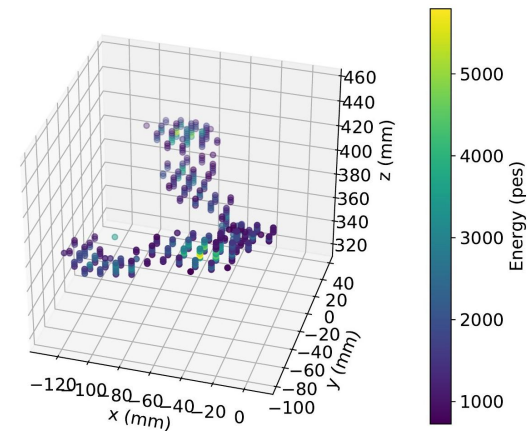
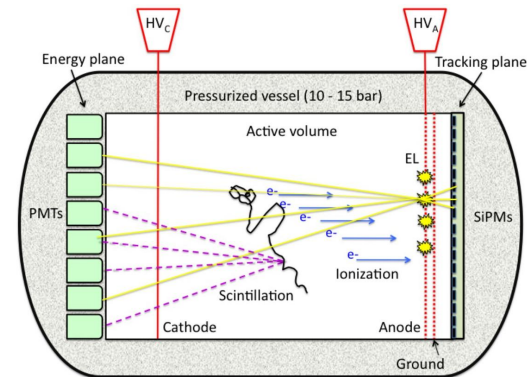
EXO-200 / nEXO



- Liquid TPC
 - Self shielding, easy to scale up
- Gas TPC
 - Energy resolution $\sim 1\%$
 - Particle tracking
- Double readout: ionization and scintillation
- Daughter tagging possible!



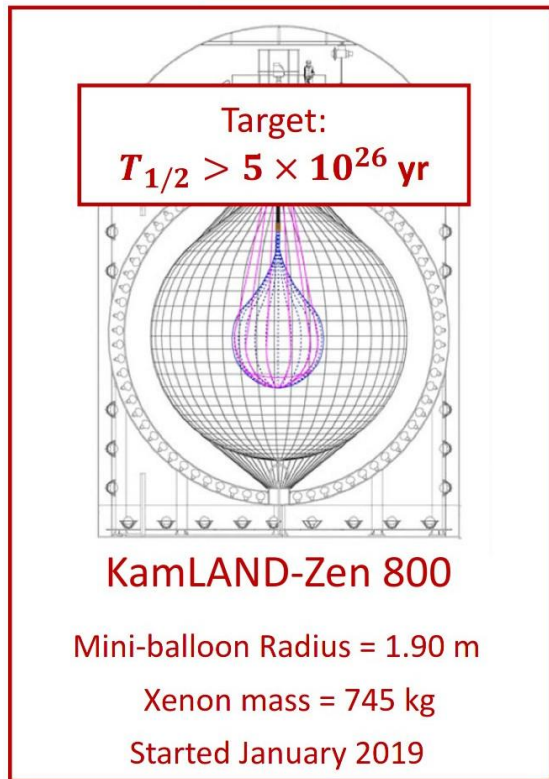
NEXT



Liquid scintillator experiments

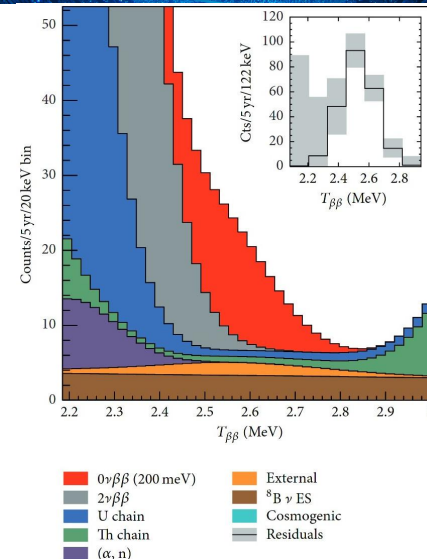
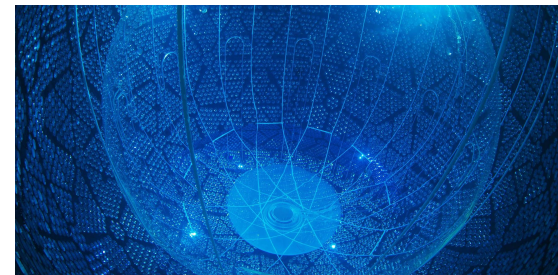
KamLAND-Zen

Current

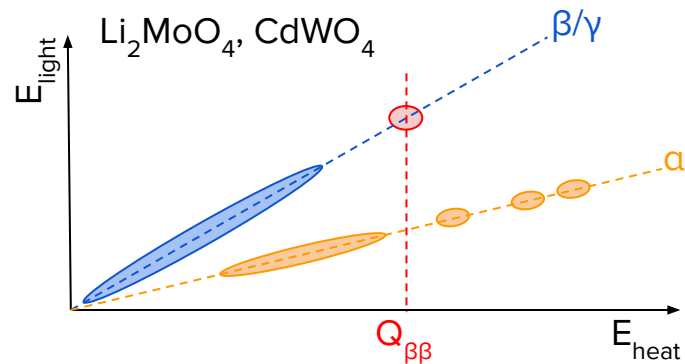
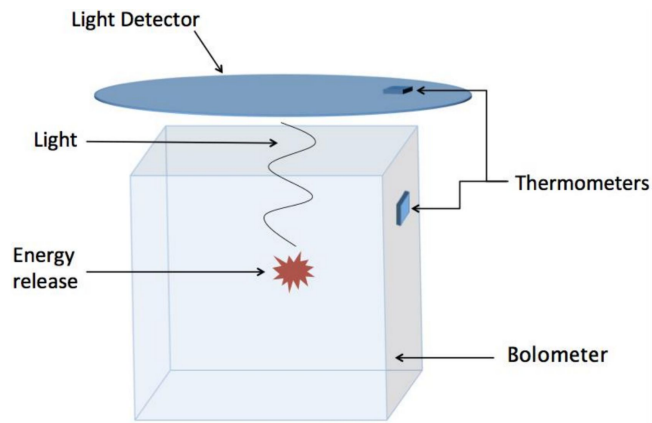


- Readout of scintillation only
→ Energy resolution of few %
→ Particle identification possible
- Very large volume
→ Isotope in central part
→ Highly effective self shielding
- Isotope (^{130}Te or ^{136}Xe) dissolved in liquid scintillator
→ Easily scalable
→ Enrichment not strictly required
- Readout of Cherenkov light possible in future experiments

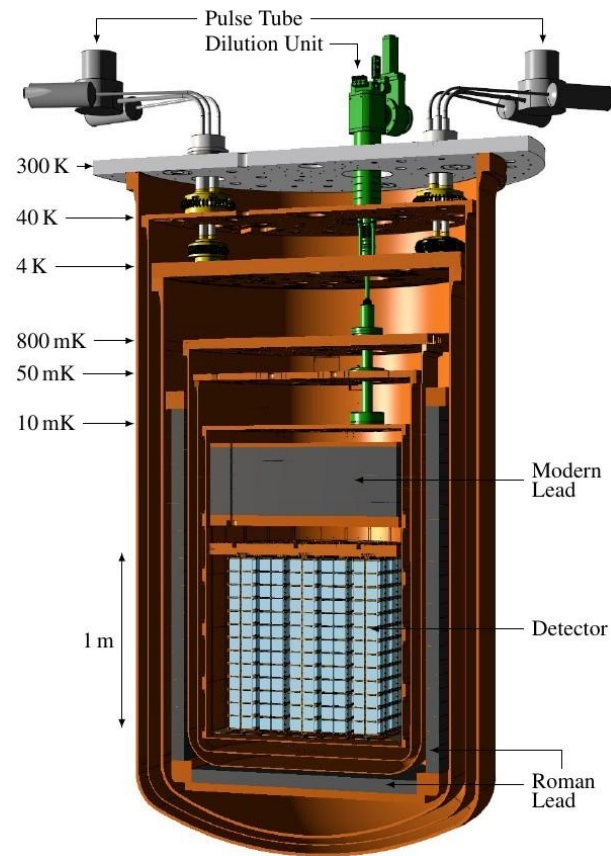
SNO+



Cryogenic calorimeters

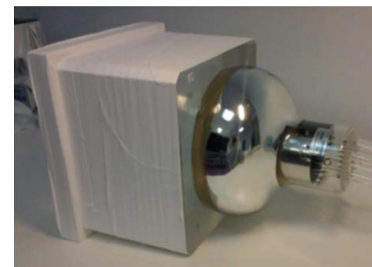
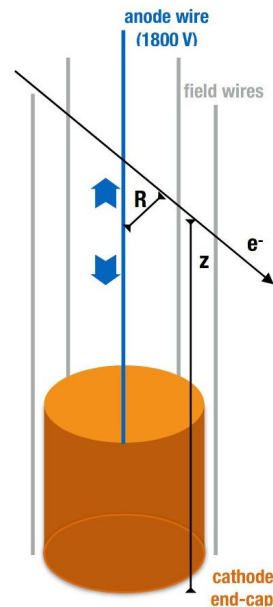
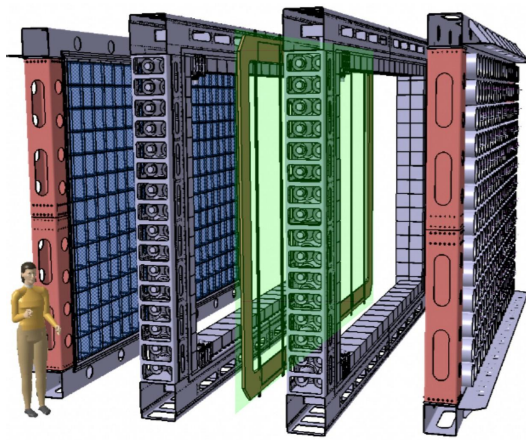
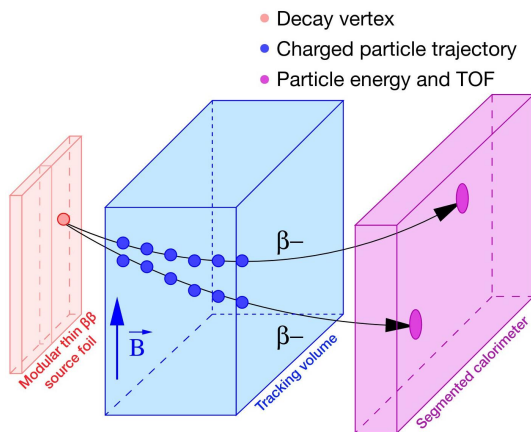


- Array of crystals operated at ~ 10 mK
- Measure temperature increase following phonon recombination
- Resolution: 5-10 keV
- Main background: α 's from support materials
- Scintillating crystal allow particle discrimination!
- CUORE: 200 kg of ^{130}Te
→ Taking data since 2017
→ $\text{BI} \sim 10^{-2}$ cts/keV/kg/yr
- CUPID: upgrade of CUORE in preparation
→ 250 kg of ^{100}Mo
→ $\text{BI} \sim 10^{-4}$ cts/keV/kg/yr thanks to light readout!



Tracking calorimeters: SuperNEMO

- Measure both energy and momentum
 - Background suppression
 - Single electrons resolved
 - Possible to study $0\nu\beta\beta$ decay mechanism
- Source \neq detector
 - Limited isotope mass
 - Any isotope is usable
- Perfect technology for precision measurement of $0\nu\beta\beta$ and $2\nu\beta\beta$ decay

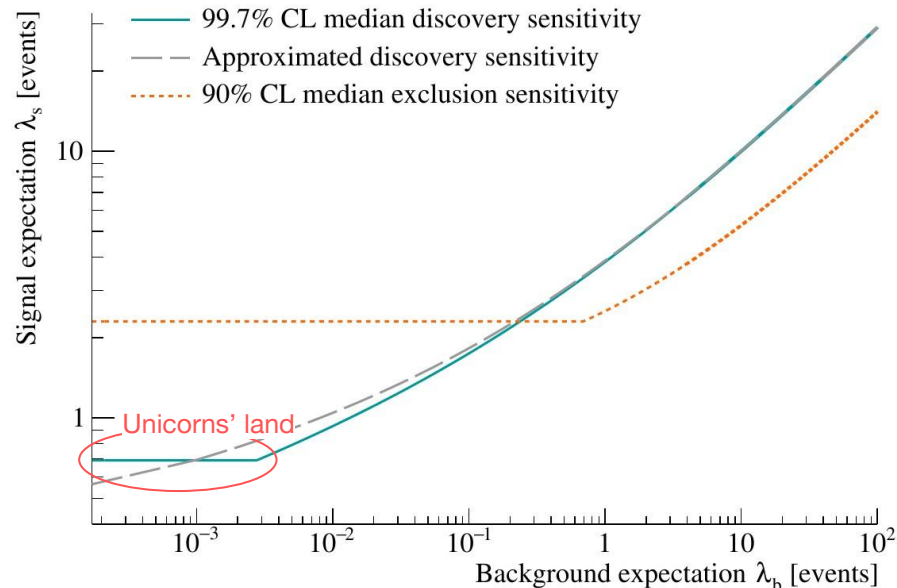


Other technologies

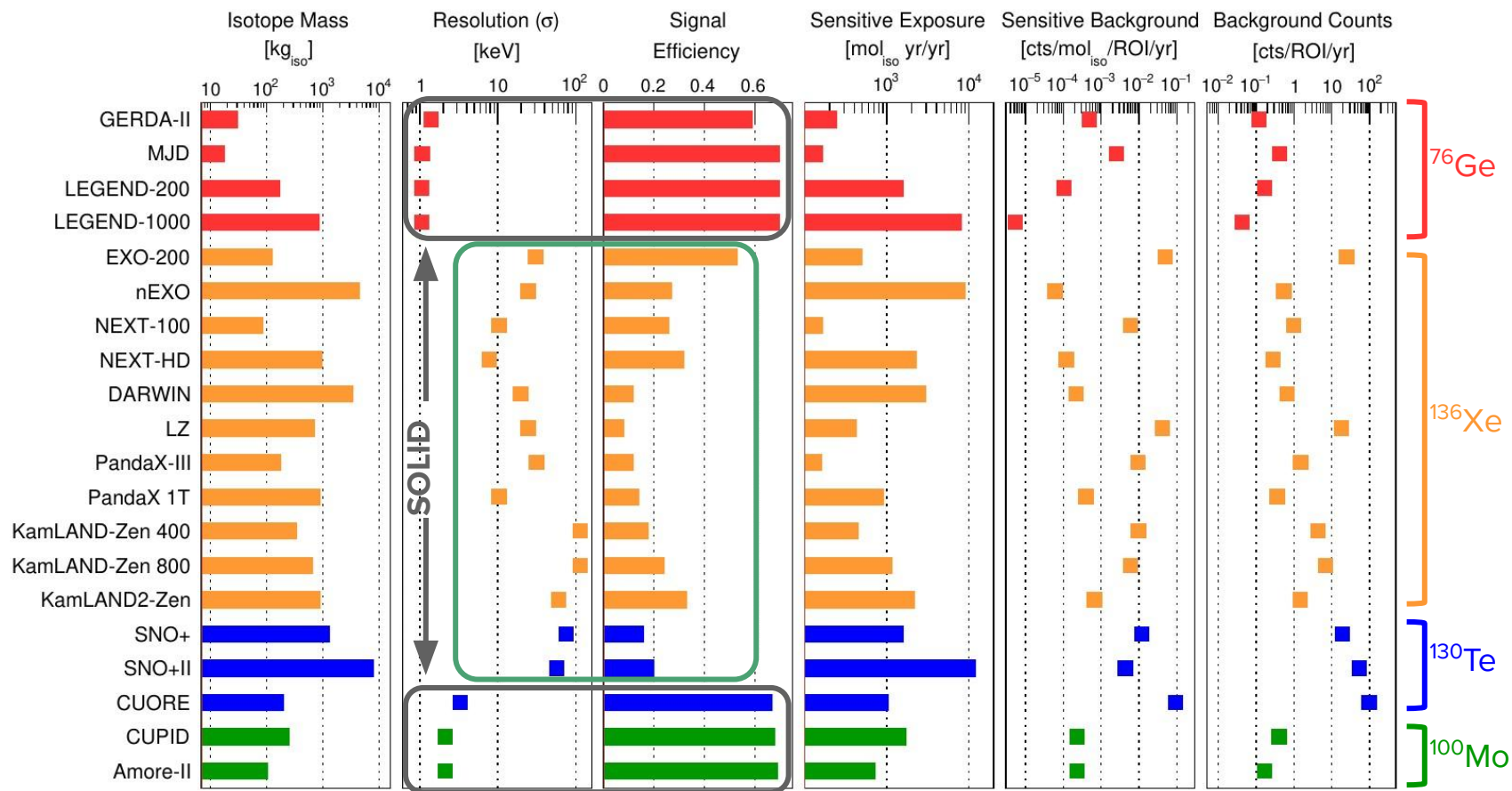
Project	Isotope(s)	Description and main features
CANDLES ^{†a}	^{48}Ca	Scintillator crystal. Possible operation as cryogenic calorimeter.
COBRA ^{†b}	^{70}Zn , $^{114,116}\text{Cd}$, $^{128,130}\text{Te}$	CdZnTe semiconductor detectors. Room temperature; multi-isotope.
Selena ^c	^{82}Se	Amorphous ^{82}Se on high resolution, high-granularity CMOS array with integrated readout. Particle discrimination via space resolution; room temperature; no self-shielding needed.
N ν DEx ^d	^{82}Se	High-pressure gaseous $^{82}\text{SeF}_6$ ion-imaging TPC. $\lesssim 1\%$ energy resolution; precise signal topology; possible multi-isotope.
R2D2 ^e	^{136}Xe	Spherical TPC. Single readout channel; inexpensive infrastructure.
AXEL ^f	^{136}Xe	High-pressure TPC operated in proportional scintillation mode. High energy resolution; possible positive ion detection.
JUNO ^g	n.d.y.	Isotope loaded liquid scintillator. Multi-isotope.
NuDot ^h	n.d.y.	Liquid scintillator loaded with quantum dots or perovskites as wavelength shifter for Cherenkov light. Discriminate directional backgrounds; multi-isotope.
ZICOS ⁱ	^{96}Zr	Zr-loaded liquid scintillator. Topology and particle discrimination via Cherenkov light readout.
THEIA ^j	n.d.y.	Water-based loaded liquid scintillator with Cherenkov light readout. Topology and particle discrimination; multi-isotope; multi-purpose.
LiquidO ^k	n.d.y.	Opaque isotope-loaded liquid scintillator with wavelength shifting fibers for event topology. Room temperature; multi-isotope; multi-purpose.

Discovery and exclusion sensitivity

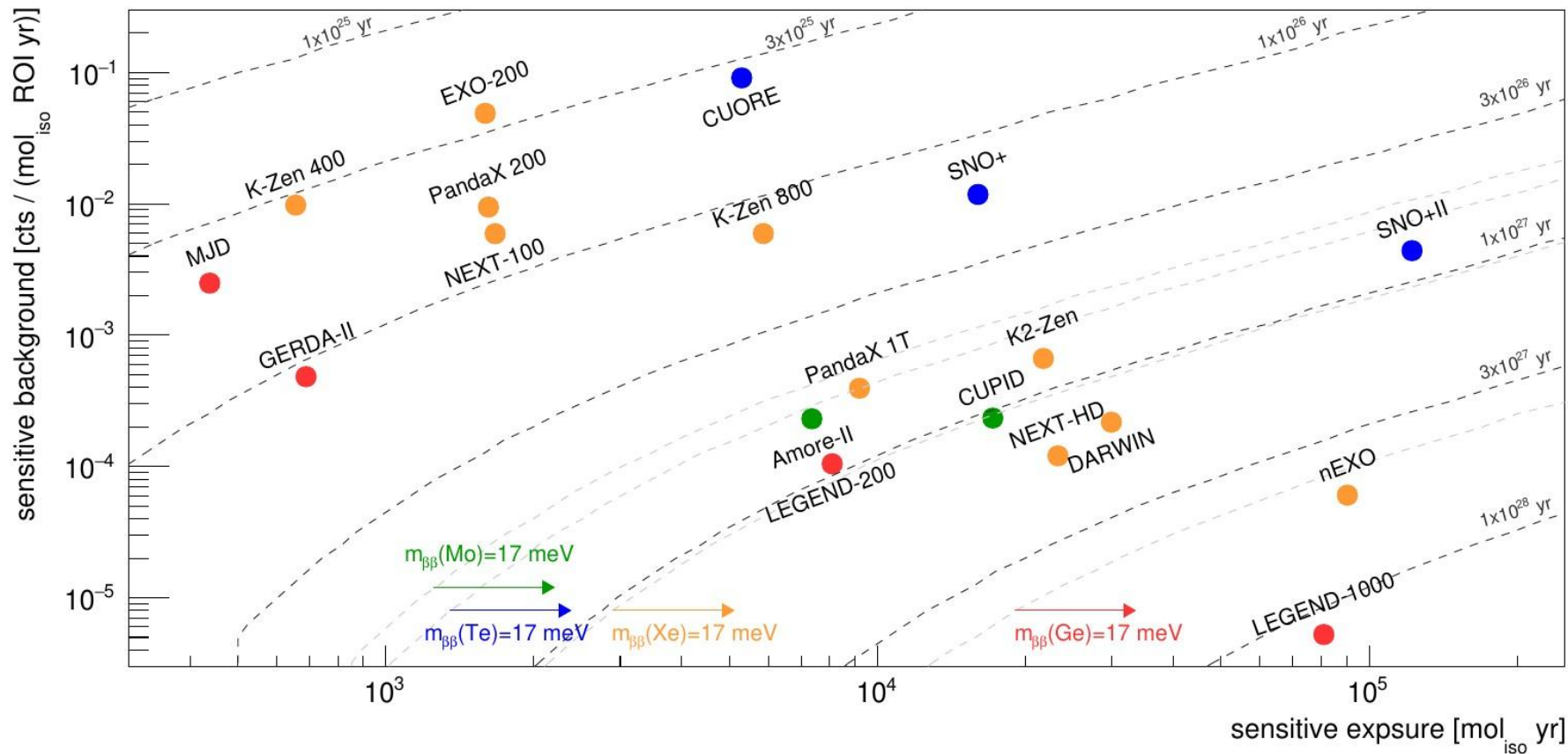
- Events measured in multi-dimensional space: energy, position, topology, particle type, ...
- Energy is the only necessary and sufficient quantity for claiming a discovery
- Option 1: multi-dimensional fit
- Option 2: cut on variables with good signal/background separation + energy fit
- Signal is restricted to small region, background dominant everywhere else
 - Counting analysis in ROI possible
 - Signal and background follow Poisson
- Discovery sensitivity:
$$\begin{cases} P(X \leq x | \lambda_b) \geq 99.73\% \\ P(X \geq x | \lambda_b + \lambda_s) \geq 50\% \end{cases}$$
- Exclusion sensitivity:
$$\begin{cases} P(X \leq x | \lambda_b) \geq 50\% \\ P(X \geq x | \lambda_b + \lambda_s) \geq 90\% \end{cases}$$



Summary of relevant experiments' parameters

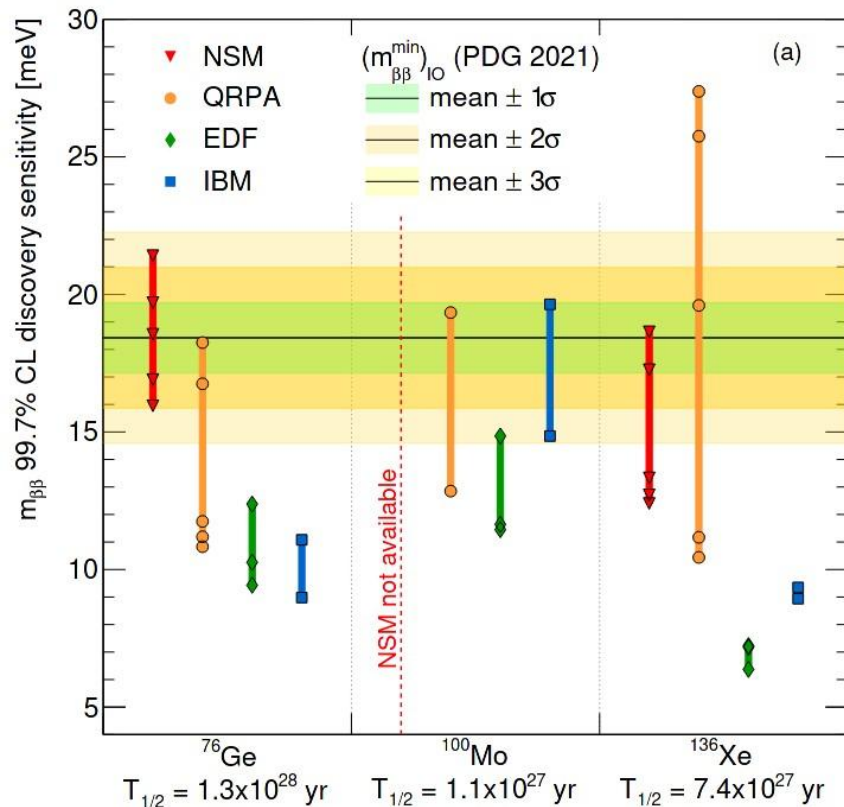


Exposure + background \rightarrow sensitivity

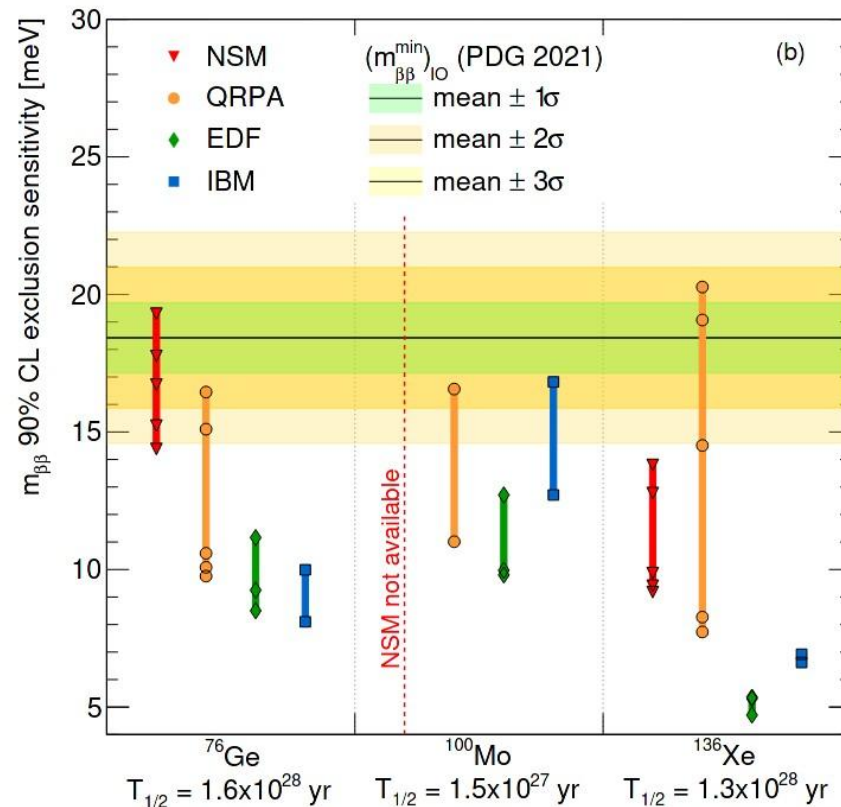


Discovery probability and impact

Discovery sensitivity



Exclusion sensitivity



THANK YOU!