

# Phase transitions in the early universe

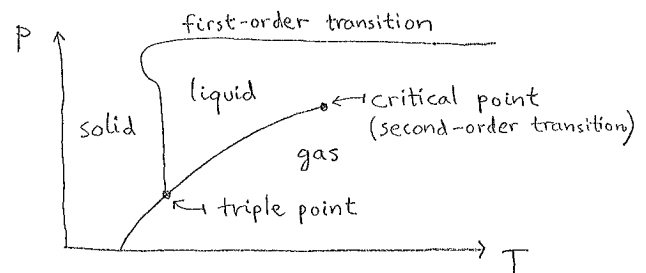
(M. Laine // GGI 2022)

1. Basics of phase transitions in field theory
2. Modern methods and applications
3. Nucleations and dynamics
4. Gravitational wave production
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## 1. Basics of phase transitions in field theory

What is a phase transition?

Phase diagram of water:



To get going, let us compute the partition function for a harmonic oscillator.

$$Z = \text{Tr}(e^{-\beta \hat{H}}), \quad \beta \equiv \frac{1}{T}, \quad \hat{H} = \text{Hamiltonian}$$

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{m\omega^2 \hat{x}^2}{2} = \underbrace{\hbar\omega}_{\equiv \varepsilon} \left( \hat{a}^\dagger \hat{a} + \frac{1}{2} \right)$$

$$\Rightarrow Z = \sum_{n=0}^{\infty} \langle n | e^{-\beta \hat{H}} | n \rangle = \sum_{n=0}^{\infty} e^{-\beta \varepsilon (n + \frac{1}{2})} = \frac{e^{-\beta \varepsilon / 2}}{1 - e^{-\beta \varepsilon}}$$

From the partition function we get the (canonical) free energy:

$$Z = e^{-\beta F} \Leftrightarrow F = -T \ln Z = \frac{\varepsilon}{2} + T \ln(1 - e^{-\beta \varepsilon})$$

We also need a "propagator" ( $\hat{x} \leftrightarrow \hat{\phi}$ ):

$$\begin{aligned} \langle \hat{x}^2 \rangle &= \frac{\text{Tr}(\hat{x}^2 e^{-\beta \hat{H}})}{\text{Tr}(e^{-\beta \hat{H}})} = (-T) \cdot \frac{2}{m} \frac{\partial}{\partial \omega^2} \ln \text{Tr}(e^{-\beta \hat{H}}) \\ &= \frac{2\hbar^2}{m} \frac{\partial}{\partial \varepsilon^2} F = \frac{\hbar^2}{m} \frac{1}{\varepsilon} \frac{\partial F}{\partial \varepsilon} \\ &= \frac{\hbar^2}{m} \frac{1}{\varepsilon} \left[ \frac{1}{2} + T \frac{\beta e^{-\beta \varepsilon}}{1 - e^{-\beta \varepsilon}} \right] \\ &= \frac{\hbar^2}{m} \frac{1}{\varepsilon} \left[ \frac{1}{2} + \frac{1}{e^{\beta \varepsilon} - 1} \right] \end{aligned}$$

"Vacuum contribution"  
(independent of T)

$\equiv n_B(\varepsilon)$   
 $\stackrel{!}{=} \text{Bose distribution}$

Scalar field theory is a collection of oscillators:  $\varepsilon \rightarrow \varepsilon_k \equiv \sqrt{k^2 + m^2}$ .  
 Moreover that system is extensive, so it is useful to consider the free energy density rather than the free energy,  $f \equiv F/V$ .

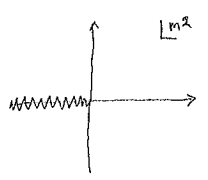
$$Z_\phi = \prod_k \exp \left\{ -\frac{1}{T} \left[ \frac{\varepsilon_k}{2} + T \ln(1 - e^{-\beta \varepsilon_k}) \right] \right\}$$

$$f_\phi = -\frac{T \ln Z_\phi}{V} = \frac{1}{V} \sum_k \left[ \frac{\varepsilon_k}{2} + T \ln(1 - e^{-\beta \varepsilon_k}) \right] \stackrel{V \rightarrow \infty}{=} \int_k \left[ \frac{\varepsilon_k}{2} + T \ln(1 - e^{-\beta \varepsilon_k}) \right]$$

The vacuum part  $\int_k \frac{\varepsilon_k}{2}$  looks divergent and requires renormalization.  
 The thermal part  $f_T(m) \equiv T \int_k \ln(1 - e^{-\beta \varepsilon_k})$  is exponentially convergent.  
 Let us look at it more precisely:

\* low-T limit,  $T \ll m$ :  $\varepsilon_k \geq m$ ,  $\beta \varepsilon_k \geq \frac{m}{T} \gg 1 \Rightarrow e^{-\beta \varepsilon_k} \ll 1$   
 $\Rightarrow f_T(m) \approx T \int_k (-e^{-\beta \varepsilon_k}) \approx -\left(\frac{mT}{2\lambda}\right)^{3/2} e^{-m/T}$

\* high-T limit,  $T \gg m$ : try a Taylor expansion in  $m^2$ ?  
 Actually, this is very non-trivial: there is a branch cut, but a generalized expansion exists:

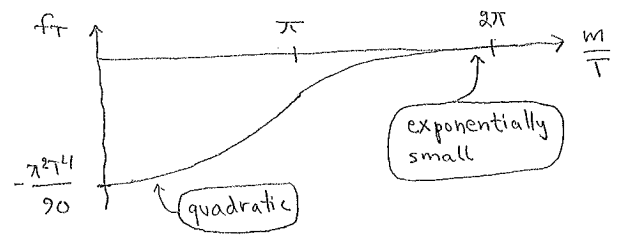


Some details on how to compute this kind of expansions are given on p.7.

$$f_T(m) = -\frac{\pi^2 T^4}{90} + \frac{m^2 T^2}{24} - \frac{m^3 T}{12\pi} + O(m^4)$$

$m := \sqrt{m^2}$

\* numerically:



We will need these results soon.

Towards Matsubara / imaginary-time formalism

The following representation can be proven\* for the "propagator":

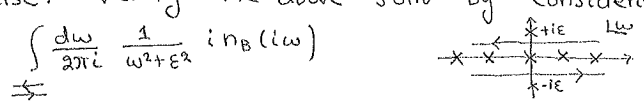
$$\frac{1}{\varepsilon_k} \left[ \frac{1}{2} + \frac{1}{e^{\varepsilon_k/T} - 1} \right] \stackrel{!}{=} T \sum_{\omega_n} \frac{1}{\omega_n^2 + \varepsilon_k^2}, \quad \omega_n \equiv 2\pi T n, \quad n \in \mathbb{Z}$$

The  $\omega_n$  are called Matsubara frequencies and  $\omega_n=0$  is "Matsubara zero mode".  
 Particularly interesting is the behaviour for  $\varepsilon_k \ll \pi T$ .

LHS:  $\frac{1}{\varepsilon_k} \left[ \frac{1}{2} + \frac{1}{e^{\varepsilon_k/T} + \varepsilon_k^2/2T^2 + \dots} \right] = \frac{1}{\varepsilon_k} \left[ \frac{1}{2} + \frac{T}{\varepsilon_k} \left( 1 - \frac{\varepsilon_k}{2T} + \dots \right) \right] = \frac{T}{\varepsilon_k^2} + O\left(\frac{1}{T}\right)$

RHS: from  $\omega_n=0$ :  $\frac{T}{\varepsilon_k^2} \Rightarrow$  much simpler to extract the limit!

\* Exercise: verify the above sum by considering the contour integral



and closing the contour in two possible ways.

Recipe of imaginary-time formalism (more on proof on p.5)

The Matsubara formalism corresponds to an "imaginary-time"/"Euclidean" path integral:

$$Z_\phi = \int_{\text{b.c.}} \mathcal{D}\phi e^{-S_E}$$

$$S_E \equiv \int_0^\beta d\tau \int_V d^3\vec{x} L_E$$

$$L_E \equiv -\mathcal{L}_M(t \rightarrow -i\tau)$$

Here boundary conditions (b.c.) are periodic for bosons (antiperiodic for fermions).

Check:  $\phi(\beta, \vec{x}) = \phi(0, \vec{x}) \Rightarrow e^{i\omega_n\beta} = 1 \Rightarrow \omega_n = 2\pi T n$  OK!

Apply this to a real scalar field with  $V(\phi) = -\frac{1}{2}m^2\phi^2 + \frac{1}{4}\lambda\phi^4$ .

$$\mathcal{L}_M = \frac{1}{2}(\partial_t\phi)^2 - \frac{1}{2}(\nabla\phi)^2 - V(\phi)$$

$$\Rightarrow L_E = \frac{1}{2}(\partial_\tau\phi)^2 + \frac{1}{2}(\nabla\phi)^2 + V(\phi)$$

Effective potential

We return to a finite volume for a moment. Let  $\bar{\phi}$  be the mode with  $\omega_n=0, \vec{k}=\vec{0}$ , and  $\phi'$  the modes with  $K=(\omega_n, \vec{k}) \neq 0$ . Note that  $\int_0^\beta d\tau \int_V d^3\vec{x} \phi' \stackrel{!}{=} 0$ .

$$\Rightarrow Z_\phi = \int_{-\infty}^{\infty} d\bar{\phi} \int \mathcal{D}\phi' e^{-S_E[\phi = \bar{\phi} + \phi']}$$

$$\equiv \int_{-\infty}^{\infty} d\bar{\phi} e^{-\frac{V}{T} V_{\text{eff}}(\bar{\phi})} \approx \int_{-\infty}^{\infty} d\bar{\phi} e^{-\frac{V}{T} [V_{\text{eff}}(\bar{\phi}_{\text{min}}) + \frac{1}{2} V_{\text{eff}}''(\bar{\phi}_{\text{min}}) (\bar{\phi} - \bar{\phi}_{\text{min}})^2 + \dots]}$$

$$\approx e^{-\frac{V}{T} V_{\text{eff}}(\bar{\phi}_{\text{min}})} \sqrt{\frac{c}{V}}$$

vanishes in thermodynamic limit

$$\Rightarrow f_\phi = V_{\text{eff}}(\bar{\phi}_{\text{min}}) + \mathcal{O}\left(\frac{\ln V}{V}\right)$$

Insert  $\phi = \bar{\phi} + \phi'$  in  $L_E$ :

$$\frac{1}{2}(\partial_\tau\phi)^2 \rightarrow \frac{1}{2}(\partial_\tau\phi')^2$$

$$-\frac{1}{2}m^2\phi^2 \rightarrow -\frac{1}{2}m^2\bar{\phi}^2 - m^2\bar{\phi}\phi' - \frac{1}{2}m^2\phi'^2$$

$$\frac{1}{4}\lambda\phi^4 \rightarrow \frac{1}{4}\lambda\bar{\phi}^4 + \lambda\bar{\phi}^3\phi' + \frac{3}{2}\lambda\bar{\phi}^2\phi'^2 + \lambda\bar{\phi}\phi'^3 + \frac{1}{4}\lambda\phi'^4$$

independent of  $\tau, \vec{x}$

vanishes because  $\int d^3\vec{x} \int d\tau \phi' = 0$

like free theory but with  $m_{\text{eff}}^2 \equiv -m^2 + 3\lambda\bar{\phi}^2$

interactions

So, in summary:

- \* tree-level potential:  $V_{\text{eff}}^{(0)}(\bar{\phi}) = -\frac{1}{2}m^2\bar{\phi}^2 + \frac{1}{4}\lambda\bar{\phi}^4$
- \* 1-loop potential:  $V_{\text{eff}}^{(1)}(\bar{\phi}) = f_\phi(m_{\text{eff}}) = \int_{\vec{k}} \left[ \frac{\epsilon_k}{2\omega_k} + T \ln(1 - e^{-\beta\epsilon_k}) \right]_{\epsilon_k = \sqrt{k^2 + m_{\text{eff}}^2}}$
- \* higher order corrections, with powers of  $\lambda$ .

What can be said about the properties of the transition?

Let us insert the high-T expansion from p.2 and see what kind of an effect  $V_{\text{eff}}^{(1)}$  has:

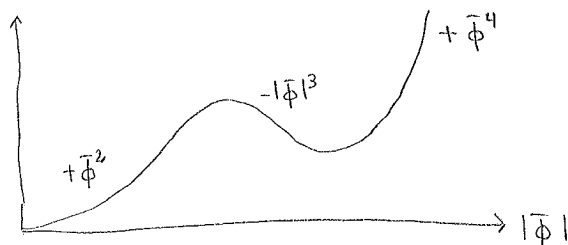
- \* leading term  $-\frac{\pi^2 T^4}{90} \Rightarrow$  independent of  $\bar{\phi} \Rightarrow$  no effect.
- \* NLO term  $\frac{m_{\text{eff}}^2 T^2}{24} = \frac{-m^2 + 3\lambda \bar{\phi}^2}{24} T^2 \Rightarrow$  thermal mass correction!

$$V_{\text{eff}}^{(0)} + V_{\text{eff}}^{(1)} \approx [\bar{\phi}\text{-indep.}] + \frac{1}{2} \left(-m^2 + \frac{\lambda T^2}{4}\right) \bar{\phi}^2 + \frac{1}{4} \lambda \bar{\phi}^4$$

Thus there is a phase transition: symmetry gets restored at  $T_c \approx \frac{2m}{\sqrt{\lambda}}$ .

- \* NNLO term  $-\frac{m_{\text{eff}}^3 T}{192\pi}$ ; let us for simplicity set  $m^2 = 0$

$$\Rightarrow V_{\text{eff}}^{(0)} + V_{\text{eff}}^{(1)} \approx [\bar{\phi}\text{-indep.}] + \frac{\lambda}{8} T^2 \bar{\phi}^2 - \frac{T}{192\pi} (3\lambda)^{3/2} |\bar{\phi}|^3 + \frac{1}{4} \lambda \bar{\phi}^4$$



This looks like a discontinuous, i.e. first-order transition.

How reliable are such predictions?

Whenever we do something perturbatively, we should estimate afterwards whether there is a small expansion parameter, i.e. whether the series converges (asymptotically).

Since the change from a continuous to a discontinuous transition is due to a 1-loop effect, it is questionable whether there is convergence in the present case.

Schematically:

$$\begin{aligned} \text{expansion parameter} &\sim \frac{T \lambda^{3/2} |\bar{\phi}|^3}{\pi \lambda \bar{\phi}^4} \sim \frac{T \lambda}{\pi \lambda^{1/2} |\bar{\phi}|} \sim \frac{T \lambda}{\pi m_{\text{eff}}}; \\ \text{no convergence if } m_{\text{eff}} &\lesssim \frac{\lambda T}{\pi} ! \end{aligned}$$

(from  $V_{\text{eff}}^{(1)}$ )

(from  $V_{\text{eff}}^{(0)}$ )

Physically, these effects originate from the Bose-enhanced  $\frac{T}{\epsilon_R}$  that we saw on p.2: there are very many quanta at small energies, which combine to give a large contribution.

Upshot (obtained with techniques discussed in 2<sup>nd</sup> lecture):

in  $O(N)$  scalar field theory, the transition is of 2<sup>nd</sup>, not 1<sup>st</sup> order.