Nine ways of evaluating a 3-loop diagram
David Broadhurst, Open University, UK, 10 December 2021, virtually at Inspired by Precision, held in Bologna in honour of Ettore Remiddi

1. I begin with a 3-loop result from Ettore and Juan Alberto Mignaco in October 1968, when I began my graduate work.
2. I describe a meeting in France in January 1992, where I was kidnapped by Ettore. This resulted in a stimulating week in Bologna, where we discussed integration by parts for massive 3 -loop diagrams in dimensional regularization.
3. I recall the monumental 3-loop result of Ettore and Stefano Laporta in February 1996 for the electron's magnetic moment.
4. The Greek for Ettore is $\mathcal{E} \kappa \tau \omega \rho$ : "he who holds all things together". Accordingly, I exhibit 9 ways of evaluating a 3 -loop massive diagram. Some of the tools were ably forged by Ettore.

Juan Alberto Mignaco and Ettore Remiddi, Fourth-order vacuum polarization contribution to the sixth-order electron magnetic moment, CERN TH.953, 30 October 1968, Il Nuovo Cimento A 60 (1969) 519-529.
Abstract: Il contributo della polarizzazione del vuoto al quarto ordine alla correzione radiativa al sesto ordine al momento magnetico dell'elettrone è calcolata analiticamente e risulta essere $0.055(\alpha / \pi)^{3}$.

Their exact coefficient of $(\alpha / \pi)^{3}$ for this 3 -loop contribution was

$$
\frac{269}{81}-\frac{434 \zeta_{2}}{135}+\frac{61 \zeta_{3}-8 \pi^{2} \log 2}{18}+\frac{\zeta_{4}-16 U_{3,1}}{3}
$$

$$
=0.05542917741228434613027275265283443562314047694171107625457697 \ldots
$$

with an alternating double sum of weight 4 :

$$
\begin{aligned}
U_{3,1} & =\sum_{m>n>0} \frac{(-1)^{m+n}}{m^{3} n}=\frac{1}{2} \int_{0}^{1} \frac{\log ^{2}(x) \log (1-x) \mathrm{d} x}{1+x} \\
& =\frac{\zeta_{4}}{2}+\frac{\left(\pi^{2}-\log ^{2} 2\right) \log ^{2} 2}{12}-2 \sum_{n>0} \frac{1}{2^{n} n^{4}}
\end{aligned}
$$

$=-0.1178759996505093268410139508341376187152175131759750633222452 \ldots$

## 2: Visiting Bologna in 1992

In January 1992, Denis Perret-Gallix (1949-2018) organized a workshop Artificial Intelligence in High Energy and Nuclear Physics near Toulon, at La-Londe-Les-Maures. I recall rather vividly Ettore's greeting, said with a friendly smile: "you are the person who has been giving seminars saying that I have wasted 20 years of my life by not using dimensional regularization for massive diagrams." Of course, I had said no such thing.

I explained that integration by parts in $D=4-2 \varepsilon$ dimensions had been useful for me, working with massive two-loop two-point integrals and massive three-loop vacuum diagrams. It gave recurrence relations that related many terms to a few master integrals, which I had obtained as hypergeometric series with parameters of the form half integers plus multiples of $\varepsilon$. Then the Laurent expansion in $\varepsilon$ gave alternating sums as well as zeta values. Ettore seemed interested and invited me to join him for a week in Bologna, after the workshop. I explained that my family was expecting me to return home. "You must tell them that you have been kidnapped by someone called Hector!"

It was a very enjoyable week in Bologna. I grouped diagrams into families that might have the same number content. I guessed that the only numbers up to weight 4 would be those found by Ettore in 1968. At weight 5, I was agnostic.

## 3: Inspired by precision in 1996

The magnetic moment of the electron, in Bohr magnetons, has electrodynamic contributions $\sum_{L \geq 0} a_{L}(\alpha / \pi)^{L}$ given up to $L=3$ loops by

$$
\begin{aligned}
a_{0} & =1 \quad[\text { Dirac, 1928] } \\
a_{1} & =\frac{1}{2} \quad[\text { Schwinger, 1947] } \\
a_{2} & =\frac{197}{144}+\frac{\zeta_{2}}{2}+\frac{3 \zeta_{3}-2 \pi^{2} \log 2}{4} \quad[\text { Petermann, Sommerfield, 1957] } \\
a_{3} & =\frac{28259}{5184}+\frac{17101 \zeta_{2}}{135}+\frac{139 \zeta_{3}-596 \pi^{2} \log 2}{18}-\frac{39 \zeta_{4}+400 U_{3,1}}{24} \\
& -\frac{215 \zeta_{5}-166 \zeta_{3} \zeta_{2}}{24} \quad[\text { Laporta \& Remiddi, 1996] }
\end{aligned}
$$

with 3 -loops terms up to weight 4 containing the constants encountered by Ettore in 1968, as expected, and terms of weight 5 that contain only zeta values, which is a pleasing simplicity for the sum of so many difficult diagrams.

Stefano's amazing semi-analytical 4-loop result, in February 2017, gives 4800 digits of $a_{4}$ and contains many new constants, two of which are Bessel moments.

## 4: Three-loop sunrise, 2007-2018

As a showcase, I choose the 3-loop equal-mass sunrise diagram in $D=2$ space-time dimensions. In momentum space it gives

$$
J\left(p^{2}\right) \propto\left(\prod_{n=1}^{4} \frac{\mathrm{~d}^{2} k_{n}}{k_{n}^{2}-m^{2}+\mathrm{i} \epsilon}\right) \delta^{(2)}\left(p-\sum_{n=1}^{4} k_{n}\right)
$$

with 4 massive particles in the intermediate state. For the magnetic moment of the electron, no diagram has an intermediate state with an even number of massive particles, since fermion number is conserved. Nevertheless, Stefano Laporta and I identified this integral, at $p^{2}=m^{2}$, as an instructive stepping stone, en route to the 5 -fermion intermediate states at 4 loops. The number theory at $D=2$ is the same as for minimal subtraction at $D=4$, where there are UV divergences.
4.1 Schwinger parameters: Let us set $m=1$ and normalize the integral as

$$
J(t)=\int_{0}^{\infty} \int_{0}^{\infty} \int_{0}^{\infty} \frac{\mathrm{d} x \mathrm{~d} y \mathrm{~d} z}{x y z((1+x+y+z)(1+1 / x+1 / y+1 / z)-t-\mathrm{i} \epsilon)}
$$

in the cut $t$-plane with a branchpoint at $t=16$. Since the denominator of the integrand is quadratic in all three parameters, we encounter an obstruction to polylogarithmic reduction at the first integration.
4.2 Co-ordinate space: Here we obtain the Bessel moment

$$
\begin{equation*}
J(t)=8 \int_{0}^{\infty} I_{0}(\sqrt{t} x) K_{0}^{4}(x) x \mathrm{~d} x \tag{1}
\end{equation*}
$$

for real $t<16$. The Bessel function $I_{0}(z)=\sum_{n \geq 0} 2^{2 n} /\left(2^{n} n!\right)^{2}$ grows exponentially at large $z$, while $K_{0}(z)$ has a logarithmic singularity at $z=0$ and falls off exponentially at large $z$. This one-dimensional integral is rather convenient for $16>t>0$. In the space-like region, with $t<0$, it becomes less convenient, since $J_{0}(z)=I_{0}(\mathrm{i} z)$ is oscillatory. On the cut, with $t>16$, it makes no sense at all.
4.3 Dispersion relation: The discontinuity across the cut is given by squares of complete elliptic integrals of the first kind and hence by reciprocals of arithmetic-geometric means that may be calculated at lightning speed:

$$
\begin{gather*}
J(t)=\int_{16}^{\infty} \frac{f_{-}^{2}(s)-3 f_{+}^{2}(s)}{s-t-\mathrm{i} \epsilon} \mathrm{~d} s, \quad f_{ \pm}(s)=\frac{2 \pi}{\operatorname{agm}(\sqrt{2 \mu}, \sqrt{\mu \pm \sqrt{\nu}})}  \tag{2}\\
\mu=\frac{2 \alpha+\beta}{2}, \quad \nu=\mu^{2}-48 \frac{\alpha-\beta}{\alpha+\beta}, \quad \alpha=\sqrt{s^{2}-4 s}, \quad \beta=\sqrt{s^{2}-16 s} .
\end{gather*}
$$

I obtained this using a remarkable discovery, made by Geoff Joyce in 1973.
4.4 Differential equation: $J(t)$ satisfies a third-order inhomogeneous differential equation, with singularities at $t=0,4,16, \infty$. For the two-loop sunrise integral $I\left(w^{2}\right)=4 \int_{0}^{\infty} I_{0}(w x) K_{0}^{3}(x) x \mathrm{~d} x$, the second-order differential equation has singularities at $w=0,1,3, \infty$. Joyce's discovery (in condensed matter physics) was that the third-order differential operator is a symmetric square of the second-order one, after a transformation of variables. I shall use the transformation

$$
t=10-w^{2}-\frac{9}{w^{2}}
$$

to relate the three-loop and two-loop cases, after modular parametrization.
4.5 Modular parametrization at 2 loops: Here the break-through came from Spencer Bloch and Pierre Vanhove in 2013. I summarize it as follows.
With $q=\exp (2 \pi \mathrm{i} z)$ and $\Im(z)>0$, the Dedekind eta function is given by

$$
\eta(z)=q^{1 / 24} \prod_{n=1}^{\infty}\left(1-q^{n}\right)=\sum_{n=-\infty}^{\infty}(-1)^{n} q^{(6 n+1)^{2} / 24}=\frac{\eta(-1 / z)}{\sqrt{-\mathrm{i} z}}
$$

With $\eta_{n}=\eta(n z)$ and $\chi_{3}(n)=0,1,-1$, for $n=0,1,2 \bmod 3$, we may expand

$$
\begin{aligned}
\frac{w}{3} & =\frac{\eta_{2}^{2} \eta_{3}^{4}}{\eta_{1}^{4} \eta_{6}^{2}}=1+4 q+12 q^{2}+28 q^{3}+60 q^{4}+120 q^{5}+O\left(q^{6}\right), \\
\frac{w^{2}-1}{8} & =\frac{\eta_{2}^{9} \eta_{3}^{3}}{\eta_{1}^{9} \eta_{6}^{3}}=1+9 q+45 q^{2}+171 q^{3}+549 q^{4}+1566 q^{5}+O\left(q^{6}\right), \\
\frac{w^{2}-9}{72} & =\frac{\eta_{2} \eta_{6}^{5}}{\eta_{1}^{5} \eta_{3}}=q+5 q^{2}+19 q^{3}+61 q^{4}+174 q^{5}+O\left(q^{6}\right), \\
f & =\frac{\eta_{1}^{6} \eta_{6}}{\eta_{2}^{3} \eta_{3}^{2}}=1-6 \sum_{n=1}^{\infty} \frac{\chi_{3}(n) q^{n}}{1+q^{n}}, \text { at weight } 1, \text { for an elliptic integral, } \\
g & =\frac{\eta_{2}^{5} \eta_{3}^{4} \eta_{6}}{\eta_{1}^{4}}=\frac{\eta_{3}^{9}}{\eta_{1}^{3}}+\frac{\eta_{6}^{9}}{\eta_{2}^{3}}=\sum_{n=1}^{\infty} \frac{n^{2}\left(q^{n}-q^{5 n}\right)}{1-q^{6 n}}, \text { at weight 3, for } 2 \text { loops, } \\
h & =\frac{\eta_{2}^{16}}{\eta_{1}^{8}}-9 \frac{\eta_{6}^{6}}{\eta_{3}^{8}}=\sum_{n=1}^{\infty} \frac{n^{3}\left(q^{n}-8 q^{3 n}+q^{5 n}\right)}{1-q^{6 n}}, \text { at weight 4, for } 3 \text { loops. }
\end{aligned}
$$

In 1993, Broadhurst, Fleischer and Tarasov gave the differential equation for the $D$-dimensional equal-mass 2-loop sunrise integral. At $D=2$, this yields

$$
-\left(q \frac{\mathrm{~d}}{\mathrm{~d} q}\right)^{2} \frac{I\left(w^{2}\right)}{6 f}=g, \quad I\left(w^{2}\right)=4 \int_{0}^{\infty} I_{0}(w x) K_{0}^{3}(x) x \mathrm{~d} x,
$$

with $w^{2}$ the norm of the external momentum. The elliptic periods

$$
\begin{aligned}
f & =\frac{4 \sqrt{3}}{\operatorname{agm}\left(\sqrt{(w+3)(w-1)^{3}}, \sqrt{16 w}\right)}=\frac{\sqrt{3} \Im I\left(w^{2}+\mathrm{i} 0\right)}{\pi^{2}} \text { for } w>3, \\
2 z f & =\frac{4 \sqrt{-3}}{\operatorname{agm}\left(\sqrt{(w+3)(w-1)^{3}}, \sqrt{(w-3)(w+1)^{3}}\right)},
\end{aligned}
$$

yield the nome $q=\exp (2 \pi \mathrm{i} z)$. Their Wronskian determines

$$
g=\frac{w^{2}\left(w^{2}-1\right)\left(w^{2}-9\right) f^{3}}{2^{6} 3^{4}}=\frac{\eta_{2}^{5} \eta_{3}^{4} \eta_{6}}{\eta_{1}^{4}}=\frac{\eta_{3}^{9}}{\eta_{1}^{3}}+\frac{\eta_{6}^{9}}{\eta_{2}^{3}}=\sum_{n=1}^{\infty} \frac{n^{2}\left(q^{n}-q^{5 n}\right)}{1-q^{6 n}} .
$$

With $\chi_{6}(n)=\chi_{2}(n) \chi_{3}(n)$ and $\chi_{2}(n)=\left(1-(-1)^{n}\right) / 2$, the solution

$$
\frac{I\left(w^{2}\right)}{f}=\frac{\pi \log (-1 / q)}{\sqrt{3}}-3 \sum_{n=1}^{\infty} \frac{\chi_{6}(n)}{n^{2}} \frac{1+q^{n}}{1-q^{n}}
$$

was obtained by Bloch and Vanhove, using the finiteness of $I(1)=\pi^{2} / 4$.

### 4.6 Atkin-Lehner transformations and optimal nomes:

$$
\begin{gathered}
z_{2}=\frac{2 z-1}{6 z-2}, \quad z_{3}=\frac{3 z-2}{6 z-3}, \quad z_{6}=\frac{-1}{6 z}, \quad q_{k}=\exp \left(2 \pi \mathrm{i} z_{k}\right), \\
f_{2}(z)=\frac{\eta_{2}^{6} \eta_{3}}{\eta_{1}^{3} \eta_{6}^{2}}, \quad f_{3}(z)=\frac{\eta_{2} \eta_{3}^{6}}{\eta_{1}^{2} \eta_{6}^{3}}, \quad f_{6}(z)=\frac{\eta_{1} \eta_{6}^{6}}{\eta_{2}^{2} \eta_{3}^{3}}, \\
-\left(q_{k} \frac{\mathrm{~d}}{\mathrm{~d} q_{k}}\right)^{2} \frac{I\left(w^{2}\right)}{6 f_{k}\left(z_{k}\right)}=g_{k}\left(z_{k}\right), \\
g_{2}(z)=\frac{\eta_{1}^{5} \eta_{3} \eta_{6}^{4}}{\eta_{2}^{4}}, \quad g_{3}(z)=\frac{\eta_{1}^{4} \eta_{2} \eta_{6}^{5}}{\eta_{3}^{4}}, \quad g_{6}(z)=\frac{\eta_{1} \eta_{2}^{4} \eta_{3}^{5}}{\eta_{6}^{4}},
\end{gathered}
$$

from which I obtain alternative expansions

$$
\begin{aligned}
& \frac{I\left(w^{2}\right)}{f_{2}\left(z_{2}\right)}=I(0)-\sum_{n=1}^{\infty} \frac{6 \chi_{3}(n)}{n^{2}} \frac{q_{2}^{n}}{1+q_{2}^{2 n}}, \\
& \frac{I\left(w^{2}\right)}{f_{3}\left(z_{3}\right)}=I(1)-\sum_{n=1}^{\infty} \frac{6 \chi_{2}(n)}{n^{2}} \frac{q_{3}^{n}}{1+q_{3}^{n}+q_{3}^{2 n}} \\
& \frac{I\left(w^{2}\right)}{f_{6}\left(z_{6}\right)}=-3 \log ^{2}\left(-q_{6}\right)+\sum_{n=1}^{\infty} \frac{6}{n^{2}} \frac{q_{6}^{n}}{1-q_{6}^{n}+q_{6}^{2 n}} .
\end{aligned}
$$

For a given $w^{2}$ we choose the smallest of the 4 nomes, for best convergence.
4.7 Three-loop sunrise: In 2008, Bailey, Borwein, Broadhurst and Glasser used the third-order differential equation to develop the momentum expansion

$$
\begin{equation*}
J(t)=8 \int_{0}^{\infty} I_{0}(\sqrt{t} x) K_{0}^{4}(x) x \mathrm{~d} x=7 \zeta_{3}+\left(7 \zeta_{3}-6\right) \frac{t}{16}+\left(49 \zeta_{3}-54\right) \frac{t^{2}}{1024}+\ldots \tag{3}
\end{equation*}
$$

about $t=0$. A neat $q$-expansion comes from the transformation

$$
\begin{align*}
t=10-w^{2}-\frac{9}{w^{2}} & =-64\left(\frac{\eta_{2} \eta_{6}}{\eta_{1} \eta_{3}}\right)^{6}=-64 q+O\left(q^{2}\right) \\
\left(q \frac{\mathrm{~d}}{\mathrm{~d} q}\right)^{3} \frac{J(t)}{(w f / 3)^{2}} & =24 h=24 \sum_{n=1}^{\infty} \frac{n^{3}\left(q^{n}-8 q^{3 n}+q^{5 n}\right)}{1-q^{6 n}} \\
\frac{J(t)}{(w f / 3)^{2}} & =7 \zeta_{3}+24 \sum_{n=1}^{\infty} \frac{\phi(n)}{n^{3}} \frac{q^{n}}{1-q^{n}} \tag{4}
\end{align*}
$$

with $\phi(n)=0, \mathbf{1}, 0,-8,0, \mathbf{1}$, for $n=0,1,2,3,4,5 \bmod 6$. This expansion works well for $|t|<8$, where $|q|<\exp (-\pi \sqrt{2} / 3)<0.22742$. On-shell, at $t=1$, we have good convergence with $-q=\exp (-\pi \sqrt{5 / 3})<0.017322$.
4.8 Atkin-Lehner transformation at three loops: For $|t|>8$, it is better to use the more ornate Fourier expansion of

$$
h_{6}(z)=\frac{-64 h}{t}=1+2 h-30 \sum_{n=1}^{\infty} \frac{n^{3}\left(q^{2 n}+q^{4 n}-8 q^{6 n}\right)}{1-q^{6 n}} .
$$

Using the alternative nome $q_{6}=\exp \left(2 \pi \mathrm{i} z_{6}\right)$, with $z_{6}=-1 /(6 z)$, we obtain

$$
\begin{gather*}
\left(q_{6} \frac{\mathrm{~d}}{\mathrm{~d} q_{6}}\right)^{3} \frac{J(t)}{\left(w f_{6}\left(z_{6}\right)\right)^{2}}=-24 h_{6}\left(z_{6}\right) \\
\frac{J(t)}{\left(w f_{6}\left(z_{6}\right)\right)^{2}}=-4 \log ^{3}\left(q_{6}\right)+24 \sum_{n=1}^{\infty} \frac{15 \phi(n+3)-\phi(n)}{n^{3}} \frac{1+q_{6}^{n}}{1-q_{6}^{n}} . \tag{5}
\end{gather*}
$$

Bloch, Kerr and Vanhove obtained this alternative solution, neglecting issues of analytic continuation or efficient convergence, I resolved both issue at Les Houches in June 2014. This simple Pari-GP procedure

$$
\begin{aligned}
& \mathrm{Z}(\mathrm{t})=\{\operatorname{local}(\mathrm{x}=2 /(\operatorname{sqrt}(4-\mathrm{t})+\operatorname{sqrt}(16-\mathrm{t})), \mathrm{a}=\operatorname{sqrt}((1-\mathrm{x}) \wedge 3 *(1+3 * \mathrm{x}))) ; \\
& \mathrm{I} / 2 * \operatorname{agm}(\mathrm{a}, 4 * \mathrm{x} * \operatorname{sqrt}(\mathrm{x})) / \operatorname{agm}(\operatorname{a}, \operatorname{sqrt}((1+\mathrm{x}) \wedge 3 *(1-3 * x))) ;\}
\end{aligned}
$$

determines $z$ and hence the nomes $q=\exp (2 \pi \mathrm{i} z)$ and $q_{6}=\exp (-\pi \mathrm{i} /(3 z))$.
4.9 On-shell evaluation at the 15 th singular value: As Laporta and I discovered at Bielefeld in July 2007, the on-shell integral

$$
\begin{equation*}
J(1)=\frac{\Gamma\left(\frac{1}{15}\right) \Gamma\left(\frac{2}{15}\right) \Gamma\left(\frac{4}{15}\right) \Gamma\left(\frac{8}{15}\right)}{30 \sqrt{5}} \tag{6}
\end{equation*}
$$

has a stunningly simple evaluation. Its proof by Bloch, Kerr and Vanhove, elucidated by Detchat Samart in 2016, was complicated by expansion in complex $q_{6}$. By contrast, expansion in real $q$ reduces the burden of proof to showing that

$$
\sum_{n=1}^{\infty} \frac{\phi(n)}{n^{3}} \frac{1-\exp (-\pi \sqrt{5 / 3} n)}{1+\exp (-\pi \sqrt{5 / 3} n)}=\frac{\pi^{3}}{12 \sqrt{15}}
$$

which results from an Eichler integral that is rational:

$$
\int_{0}^{\sqrt{1 / 15}} h\left(\frac{1+\mathrm{i} y}{2}\right) y \mathrm{~d} y=\frac{1}{120}, \quad h(z)=\frac{\eta_{2}^{16}}{\eta_{1}^{8}}-9 \frac{\eta_{6}^{16}}{\eta_{3}^{8}},
$$

somewhat to the surprise of Bruce Berndt.
4.10 Theta series at the 15 th singular value: With $b>a \geq 0$ and $c \geq 0, \mathrm{I}$ define Bessel moments

$$
M(a, b, c)=\int_{0}^{\infty} I_{0}^{a}(x) K_{0}^{b}(x) x^{c} \mathrm{~d} x
$$

Then the 5-Bessel matrix

$$
\left[\begin{array}{ll}
M(1,4,1) & M(1,4,3) \\
M(2,3,1) & M(2,3,3)
\end{array}\right]=\left[\begin{array}{cc}
\pi^{2} C & \pi^{2}\left(\frac{2}{15}\right)^{2}\left(13 C-\frac{1}{10 C}\right) \\
\frac{\sqrt{15 \pi} C}{2} & \frac{\sqrt{15 \pi}}{2}\left(\frac{2}{15}\right)^{2}\left(13 C+\frac{1}{10 C}\right)
\end{array}\right]
$$

has a determinant $2 \pi^{3} / \sqrt{3^{3} 5^{5}}$ that is free of the 3-loop constant $C$ in

$$
\begin{align*}
J(1)= & 8 \pi^{2} C=\frac{64 \pi^{3}(1-1 / \sqrt{5})}{\operatorname{agm}^{2}(8 \sqrt{2}, \sqrt{15}+3 \sqrt{5}+\sqrt{3}-1)}  \tag{7}\\
& =\frac{\pi^{3}}{2}\left(1-\frac{1}{\sqrt{5}}\right)\left(1+2 \sum_{n=1}^{\infty} q_{\text {tiny }}^{n^{2}}\right)^{4}, \tag{8}
\end{align*}
$$

$=8.570280443374461268149695824246537286105438217336449736197549 \ldots$
where $q_{\text {tiny }}=\exp (-\pi \sqrt{\mathbf{1 5}})<0.0000051975$ is the tiny nome of Bologna.
4.11 Critical $L$-value: The L-series for 5 Bessel functions comes from a modular form of weight 3 and level 15:

$$
\begin{gather*}
f_{3,15}(z)=\left(\eta_{3} \eta_{5}\right)^{3}+\left(\eta_{1} \eta_{15}\right)^{3}=\sum_{n>0} A(n) q^{n}, \\
L(s)=\sum_{n>0} \frac{A(n)}{n^{s}}, \quad \text { for } s>2, \\
J(1)=\frac{8 \pi^{2}}{5} L(1)=\frac{8 \pi^{2}}{5} \sum_{n>0} \frac{A(n)}{n}\left(2+\frac{\sqrt{\mathbf{1 5}}}{2 \pi n}\right) \exp \left(-\frac{2 \pi n}{\sqrt{\mathbf{1 5}}}\right) \tag{9}
\end{gather*}
$$

as Francis Brown and I discovered at Les Houches in June 2010.
4.12 Finite part in 4 dimensions: With the measure $d^{D} k /\left(\pi^{D / 2} \Gamma(1+\varepsilon)\right)$ for a loop momentum $k$, in $D=4-2 \varepsilon$ dimensions, the Laurent expansion of the on-shell 3-loop sunrise diagram is

$$
\frac{2}{\varepsilon^{3}}+\frac{22}{3 \varepsilon^{2}}+\frac{577}{36 \varepsilon}+\frac{6191}{216}-\frac{\pi^{2}}{3}\left(4 C+\frac{7}{40 C}\right)+O(\varepsilon)
$$

as I discovered using PSLQ in 2007.

Summary: I have given 9 ways of evaluating a 3-loop diagram, using

1. a Bessel moment, in co-ordinate space,
2. a dispersion integral over squares of elliptic integrals,
3. a momentum expansion from an inhomogeneous differential equation,
4. an expansion in powers of a nome, using eta quotients,
5. an alternative expansion obtained by an Atkin-Lehner transformation,
6. a product of Gamma values at the 15 th singular value,
7. the inverse square of an arithmetic-geometric mean,
8. the fourth power of a theta series, for the fastest method,
9. a critical value of an $L$-series of a modular form of weight 3 .

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