

# Neutrino oscillation from reactor experiments, neutrino reactor anomaly and sterile neutrinos

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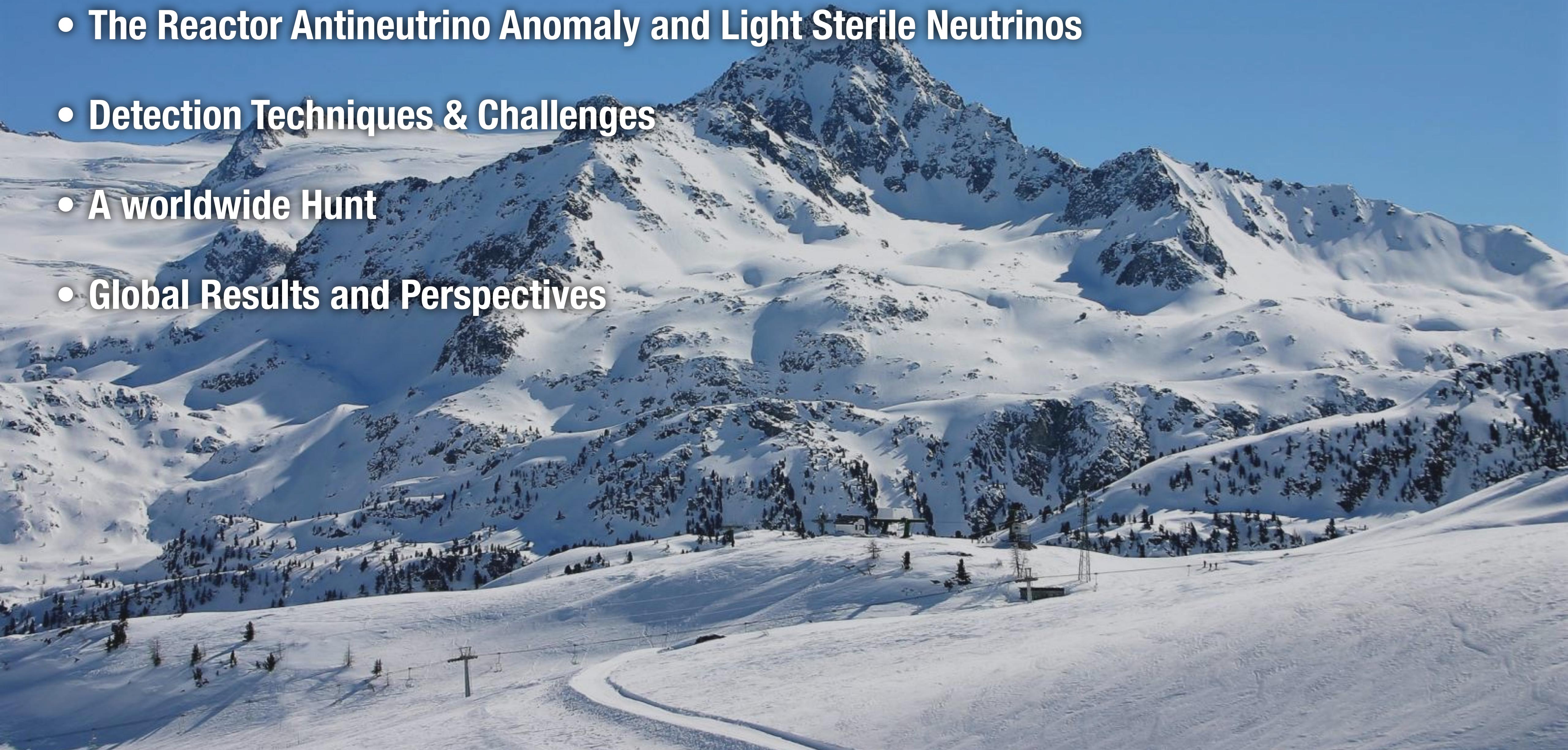
[alessandro.minotti@unimib.it](mailto:alessandro.minotti@unimib.it)

La Thuile 2022 - Les Rencontres de Physique de la Vallée d'Aoste

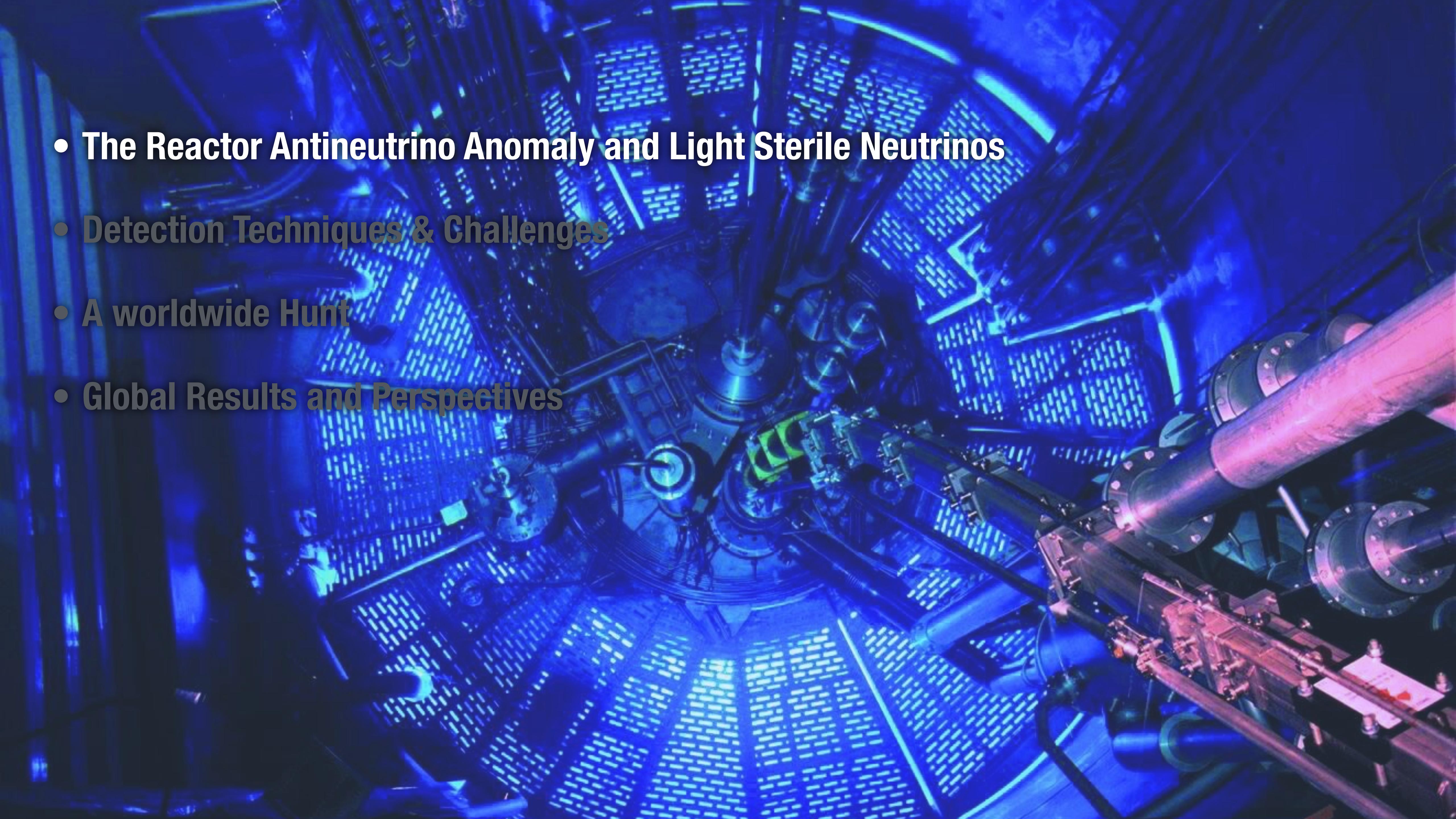
08/03/2022

# Outline

- The Reactor Antineutrino Anomaly and Light Sterile Neutrinos
- Detection Techniques & Challenges
- A worldwide Hunt
- Global Results and Perspectives



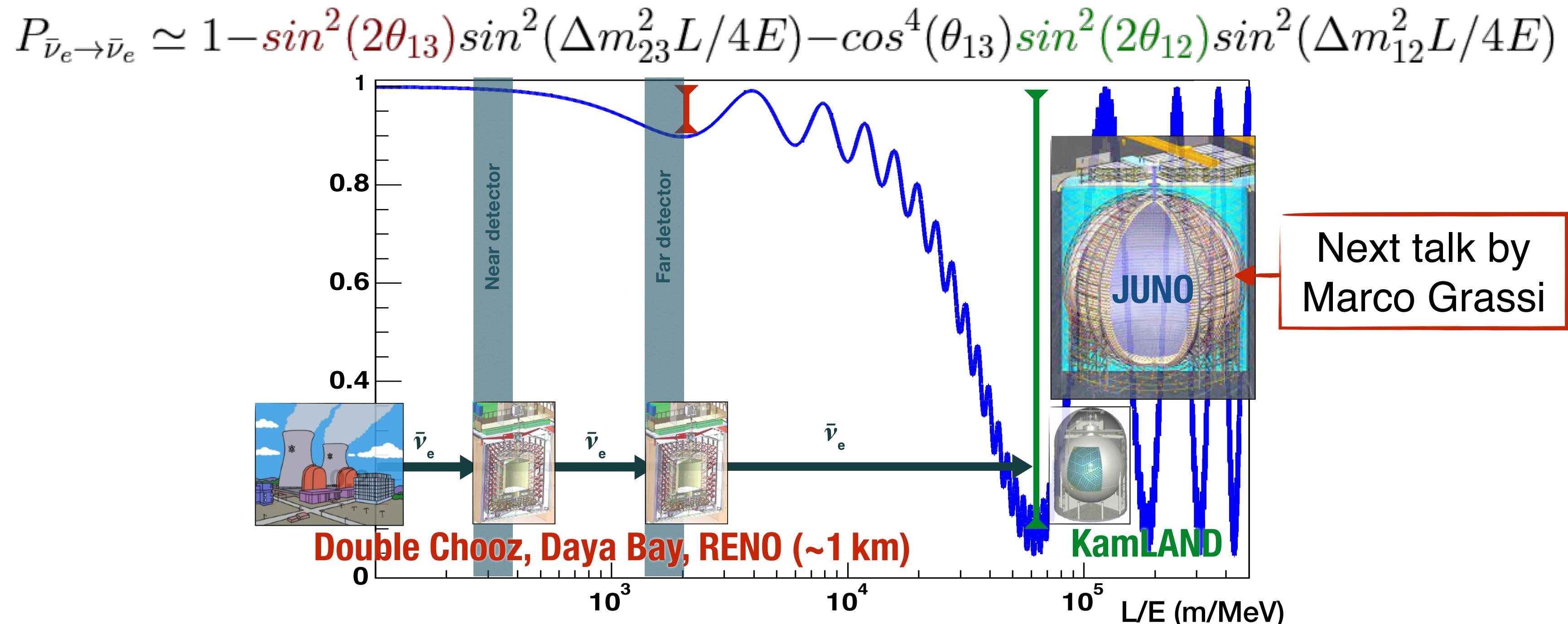
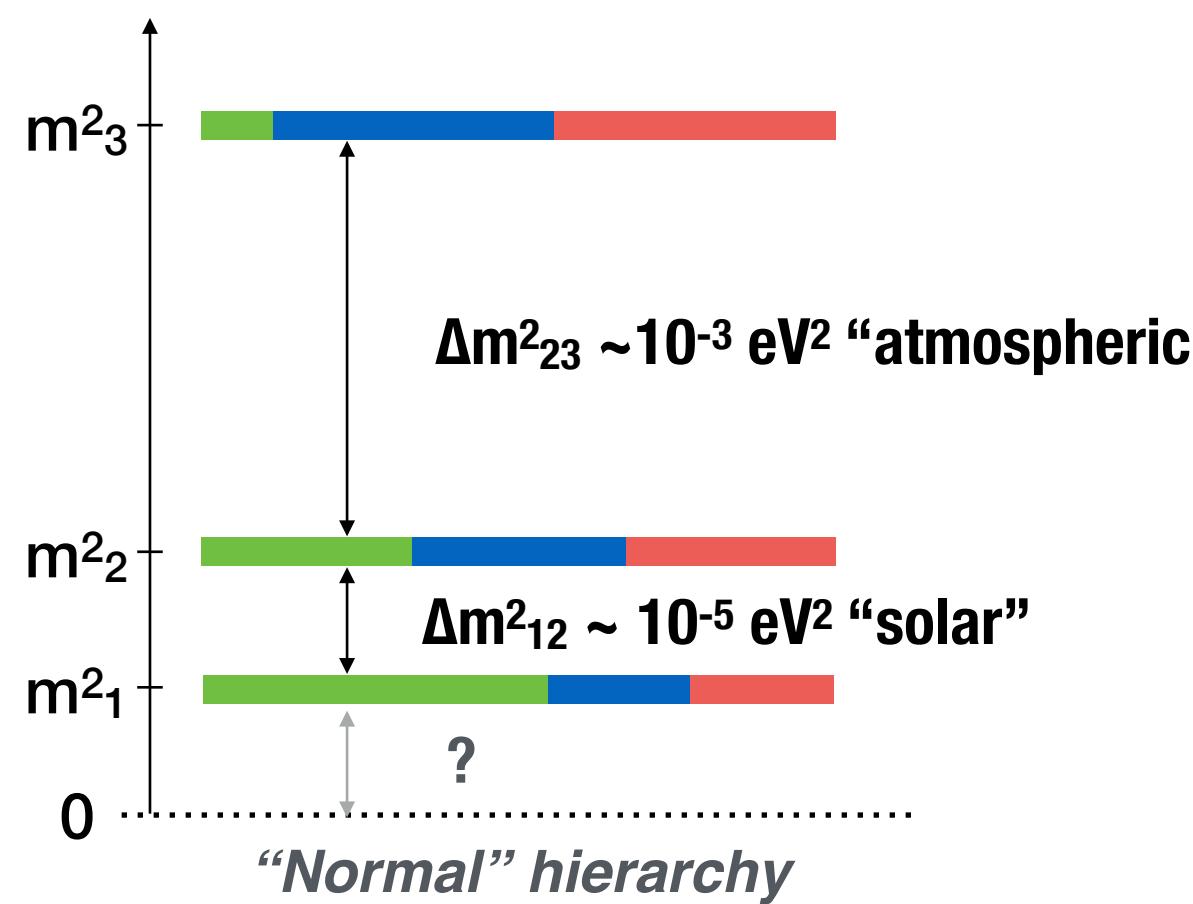
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# Reactor Neutrino Oscillation Experiments

- Reactor antineutrino oscillation is sensitive to **two mixing angles** ( $\theta_{13}$ ,  $\theta_{12}$ ) of the  $U_{PMNS}$  matrix, and **both squared-mass splittings**, including, if energy resolution is good enough, the **sign of  $\Delta m^2_{23}$**

$$U_{PMNS} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}$$



- In the 2000s we built **three experiments to measure** the yet unknown  $\theta_{13}$ , which is the nowadays most precisely measured

D. Adey, F. An, A. Balantekin, H. Band, M. Bishai, S. Blyth et al., *Measurement of the electron antineutrino oscillation with 1958 days of operation at daya bay*, *Physical review letters* **121** (2018) 241805.  
 H. de Kerret and T.D.C. Collaboration, *Double chooz  $\theta_{13}$  measurement via total neutron capture detection*, *Nature Physics* **16** (2020) 558.  
 C. Shin, Z. Atif, G. Bak, J. Choi, H. Jang, J. Jang et al., *Observation of reactor antineutrino disappearance using delayed neutron capture on hydrogen at reno*, *Journal of High Energy Physics* **2020** (2020) 1.

- Basic idea: compare spectra in near-far detector(s)



# Reactor Antineutrino Anomaly (RAA)

- Mueller ( $^{238}\text{U}$ )-Huber ( $^{235}\text{U}$ , Pu) calculated new precise IBD spectra, that could be used before a near detector was operational

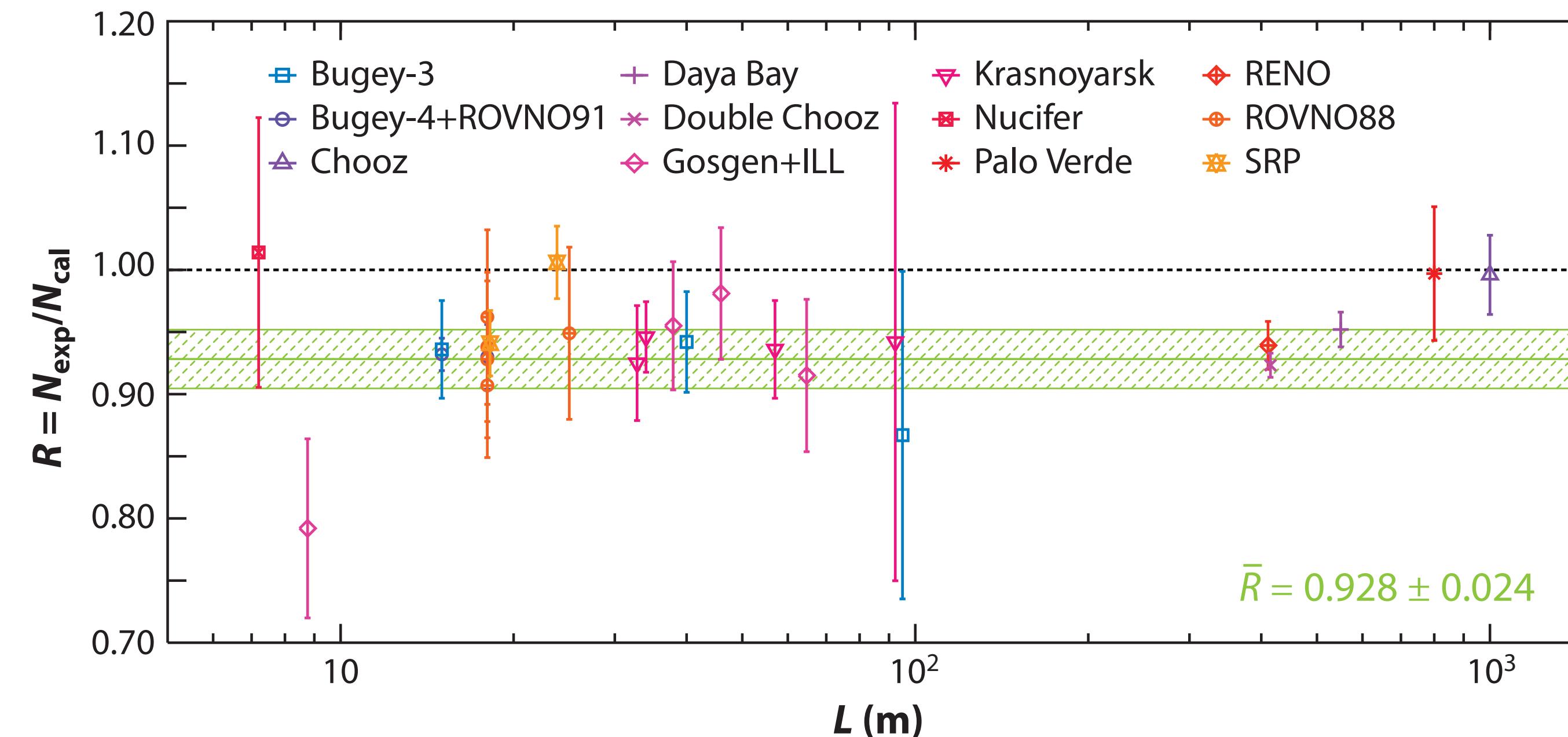
[4] T.A. Mueller, D. Lhuillier, M. Fallot, A. Letourneau, S. Cormon, M. Fechner et al., *Improved predictions of reactor antineutrino spectra*, *Physical Review C* **83** (2011) 054615.

[5] P. Huber, *Determination of antineutrino spectra from nuclear reactors*, *Physical Review C* **84** (2011) 024617.

- Rate excess of ~6% in the model** compared to previous short baseline measures

[6] G. Mention, M. Fechner, T. Lasserre, T.A. Mueller, D. Lhuillier, M. Cribier et al., *Reactor antineutrino anomaly*, *Physical Review D* **83** (2011) 073006.

- Discrepancy confirmed by Double Chooz, Daya Bay and RENO near detectors



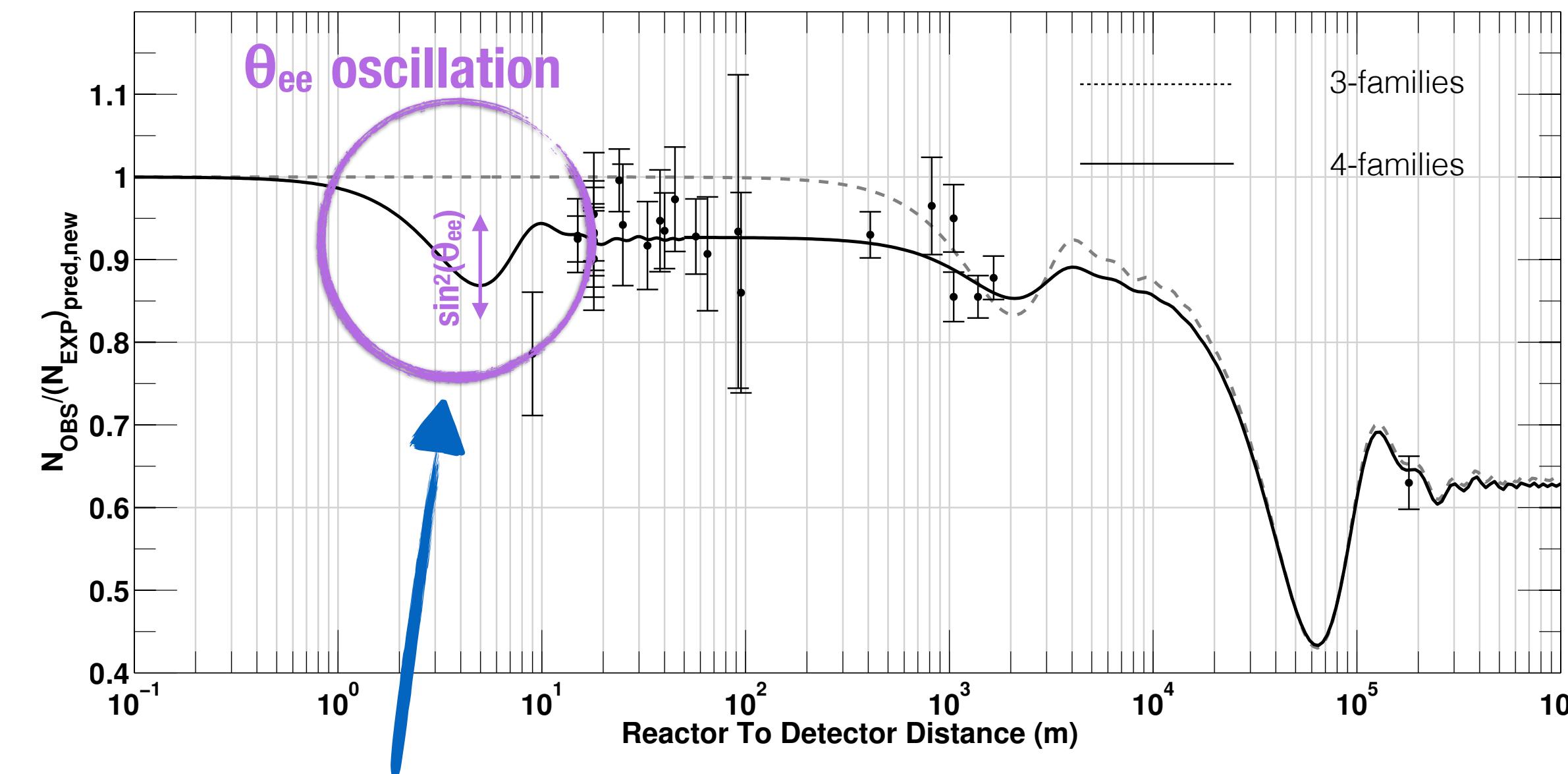
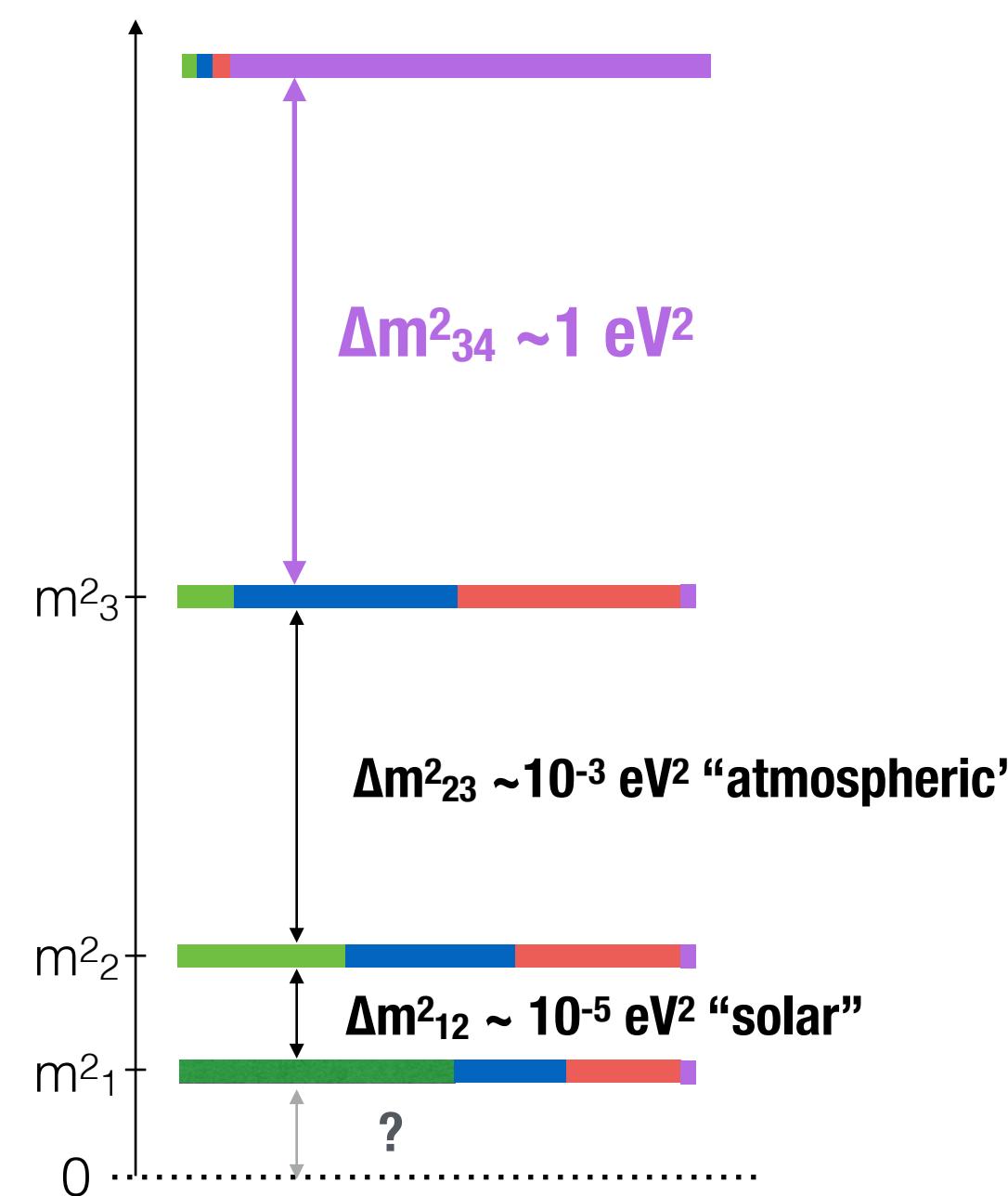
# The Light Sterile Neutrino

- Adding a **new neutrino** (0.1-1 eV mass) consisting almost exclusively of an **extra sterile flavour** can account for the observed deficit

$$U_{PMNS} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

- Sterile neutrinos** do not interact weakly but **mix with standard neutrinos**, originating the disappearance at short baseline

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L \lesssim 10m) \simeq 1 - \sin^2(\theta_{ee}) \sim^2 (1.27 \Delta m_{14}^2 L / E)$$

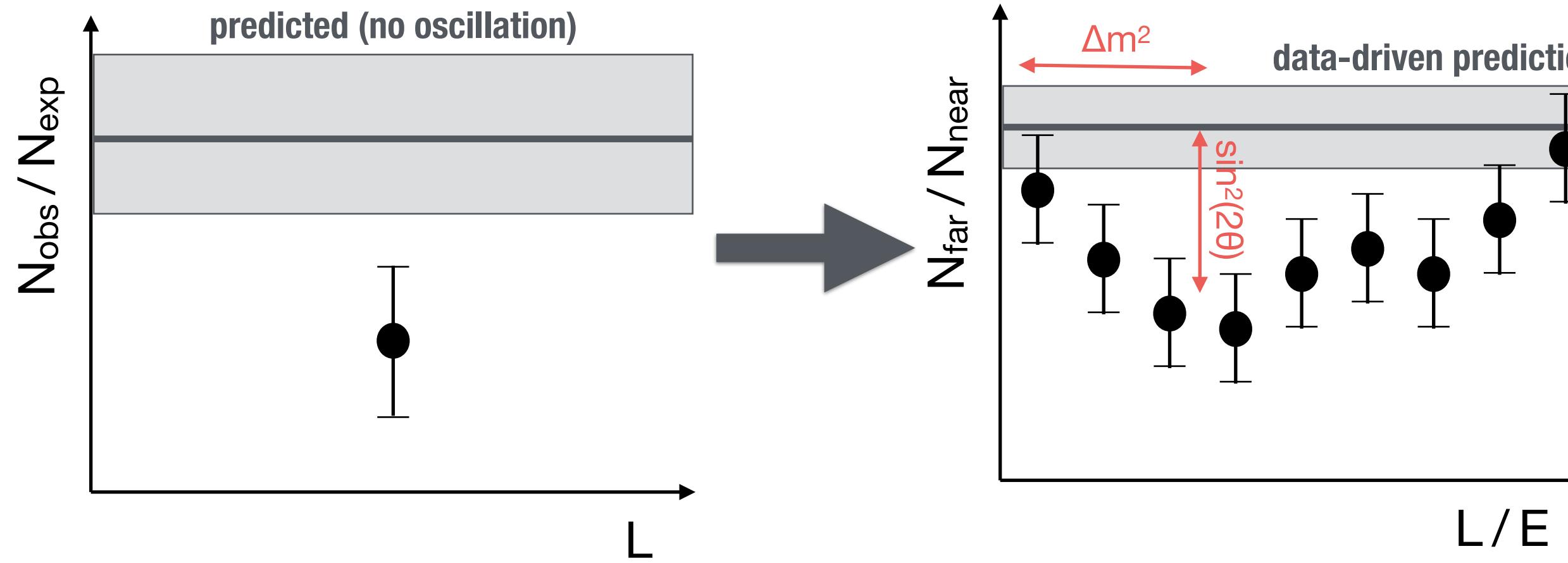


- To see this oscillation we need to **go look at very short ( $\sim 10m$ ) baselines**

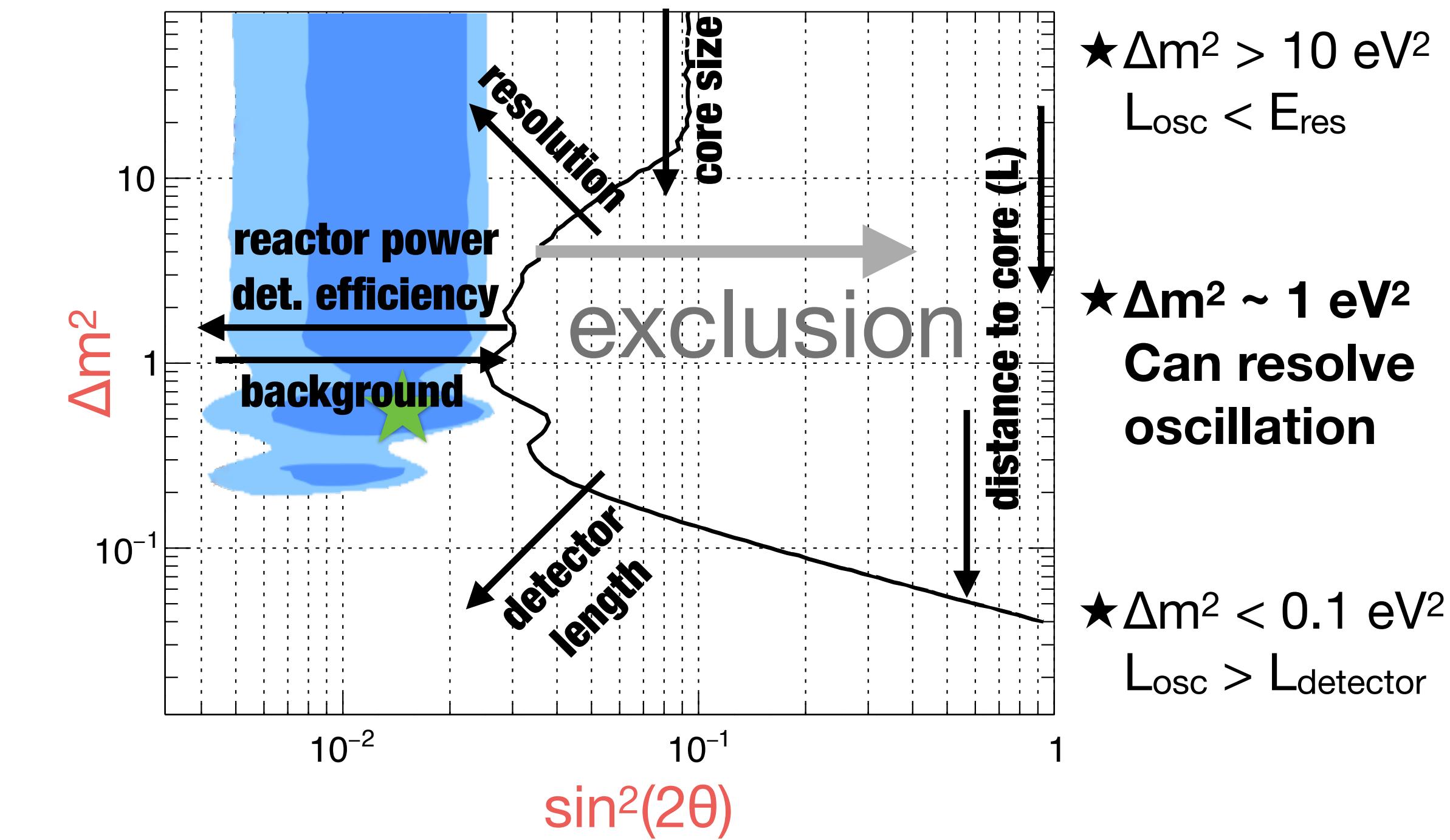
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# Search for the Light Sterile Neutrino

- Difficulty in predicting the neutrino rate limits the sensitivity of past measurements → need to **disentangle the oscillating signature from the absolute rate**

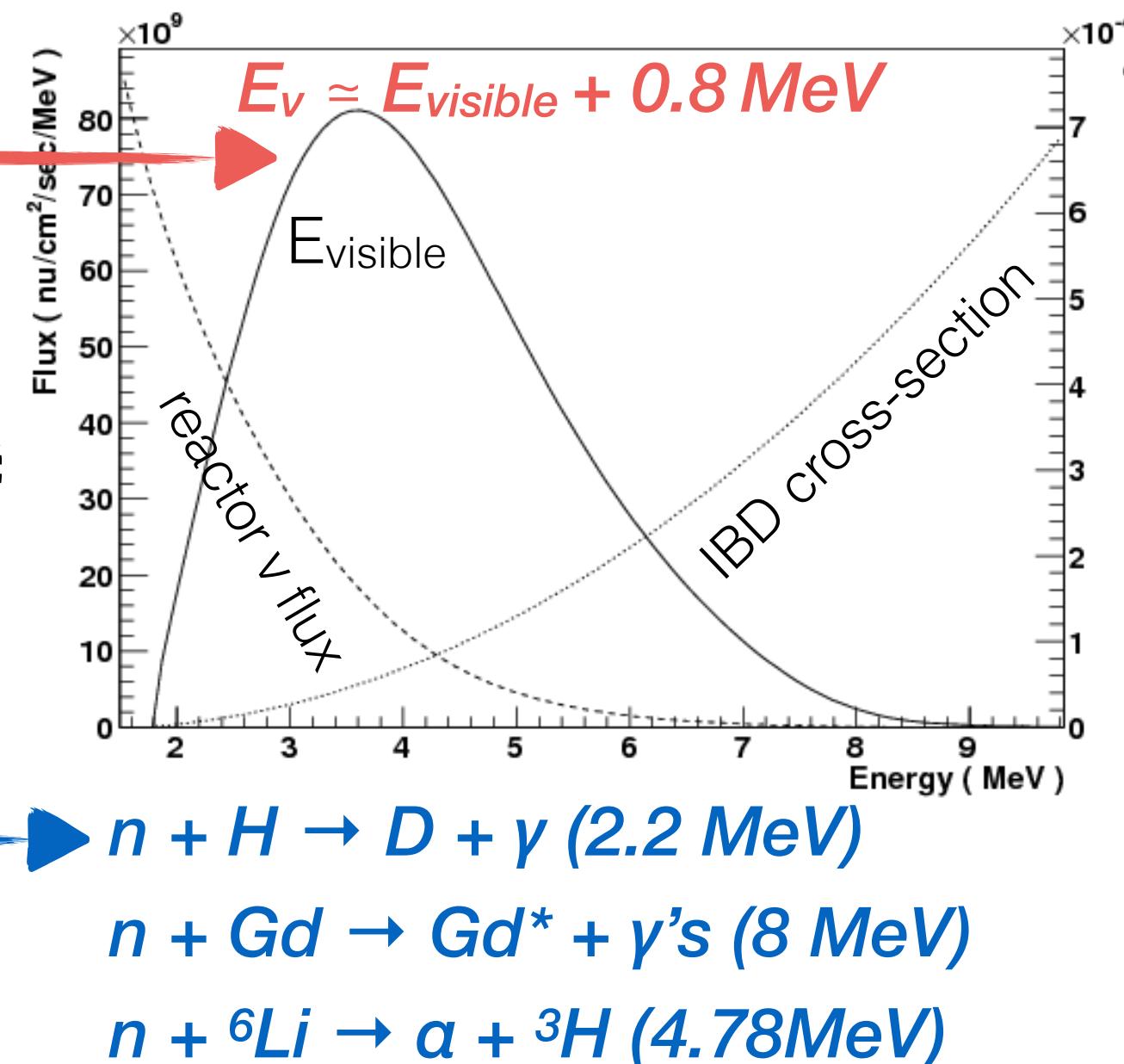
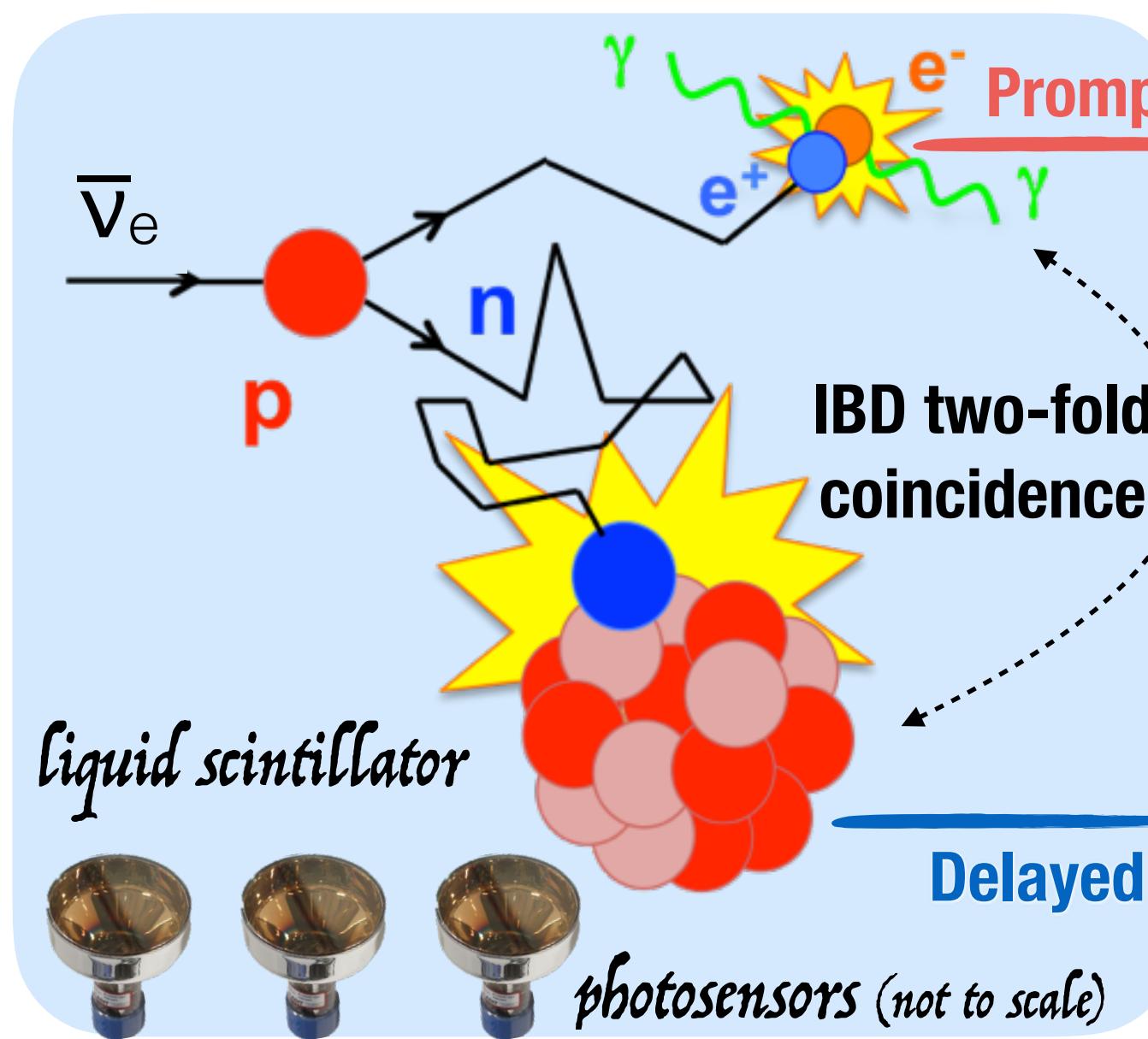


- Oscillation parameters ( $\Delta m^2$ ,  $\theta$ ) are tested against data
  - Oscillation hypothesis ⇒ **contour plot (CL)** + **best fit**
  - Null hypothesis ⇒ **exclusion plot**
- The better statistics (reactor power, detection efficiency), the larger the exclusion, but the core size washes out the oscillation, and distance from core send us far from the RAA zone



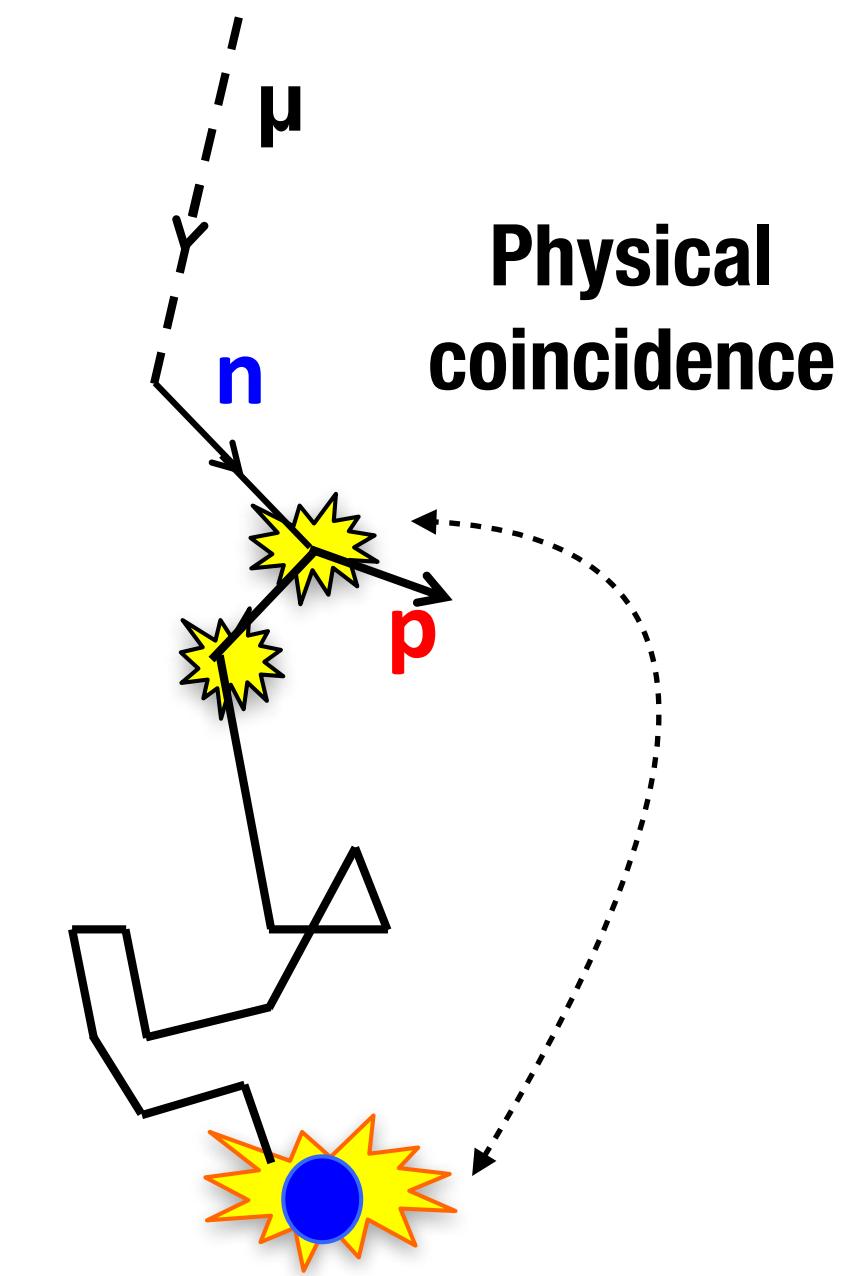
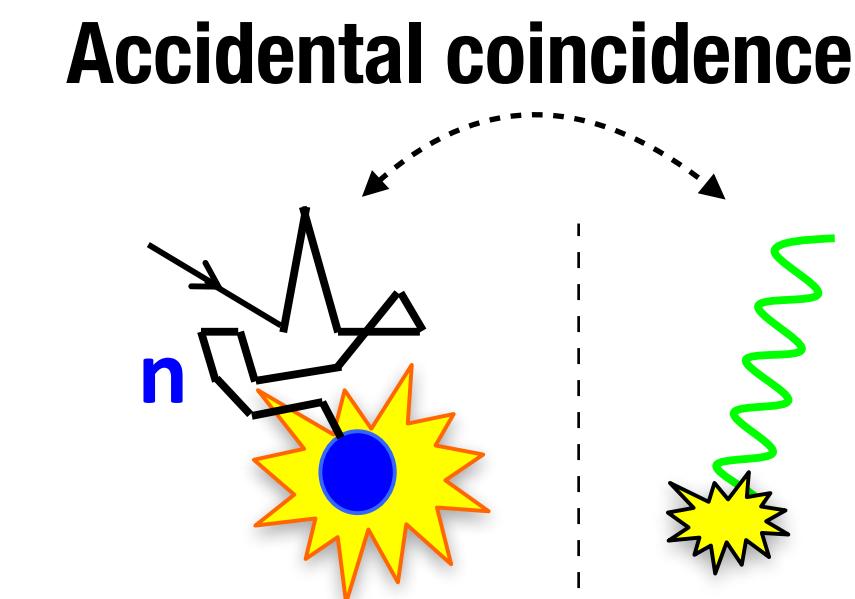
# Reactor Antineutrino Detection

- **Signal:** Inverse Beta Decay in scintillator target

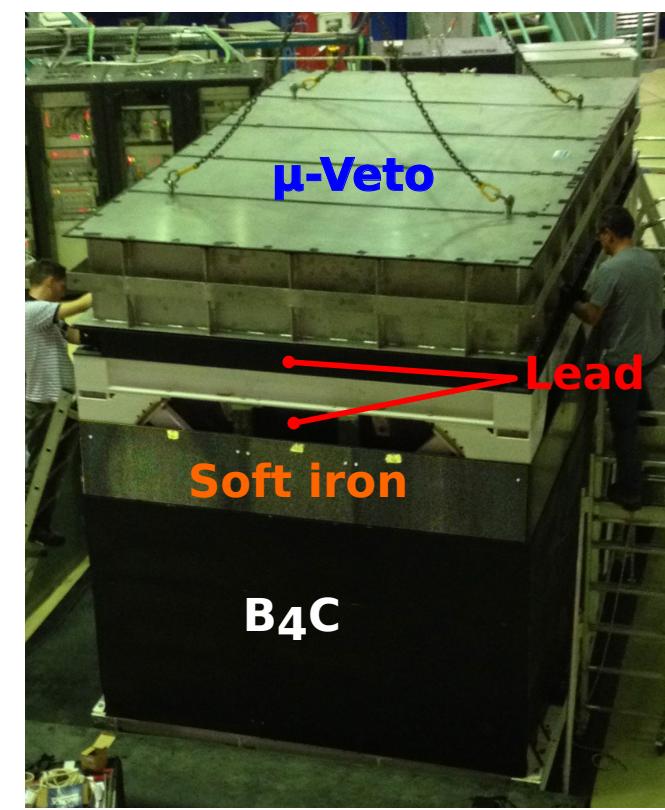
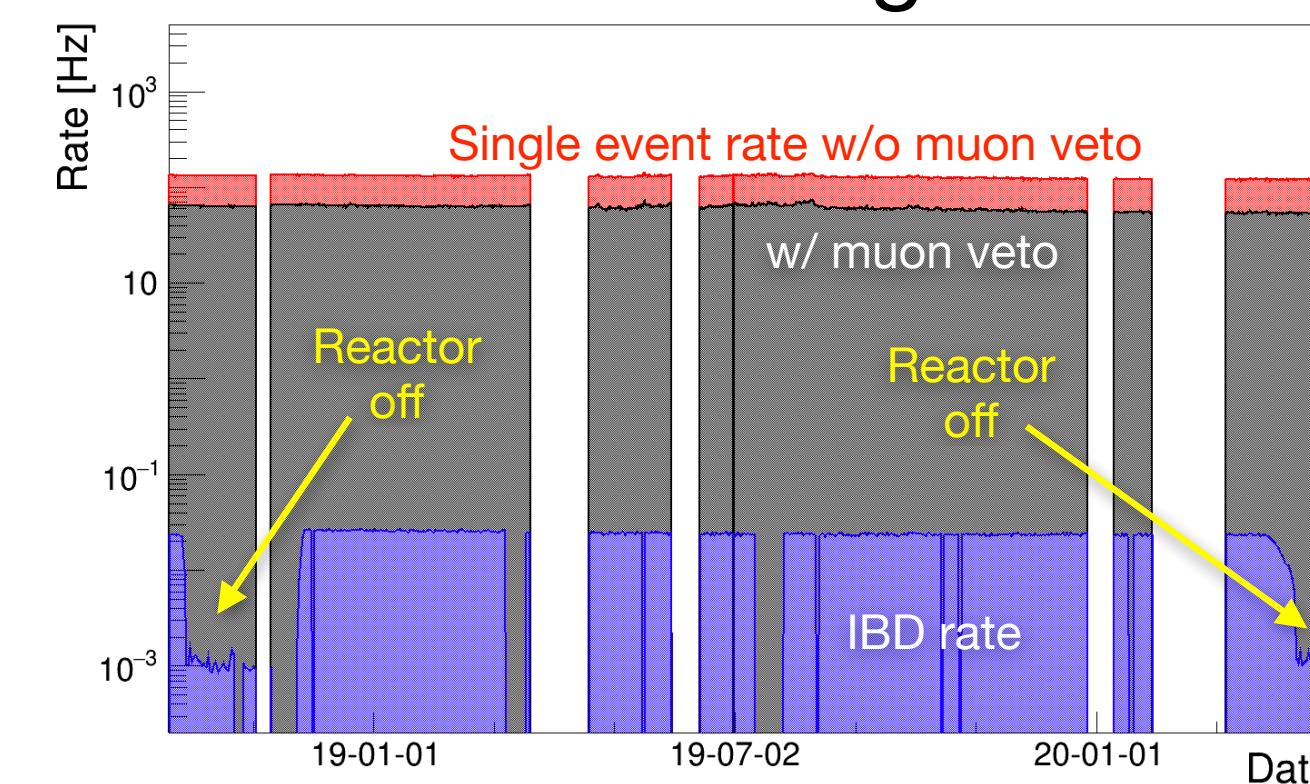


- **Background**

- Cosmic induced ( $\mu$ 's,  $n-\gamma$  from  $\mu$  spallation)
- Reactor induced ( $n$ 's,  $\gamma$ 's)



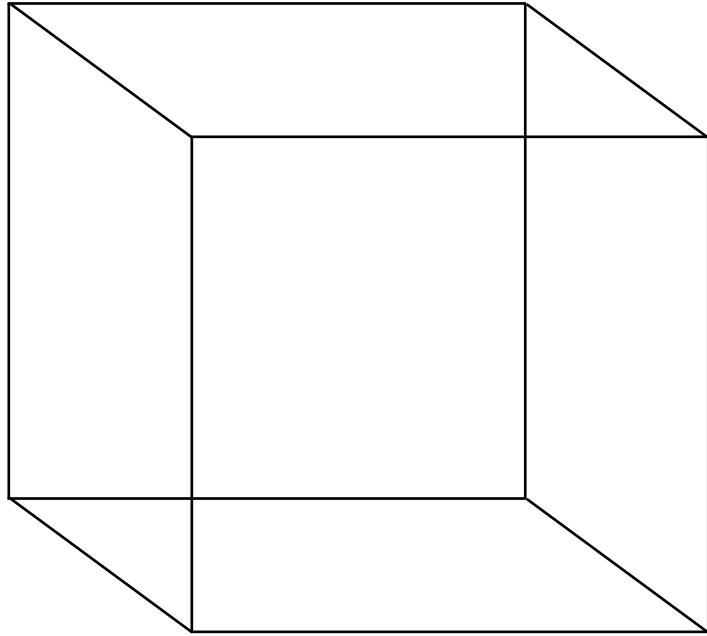
- Event topology ( $E_{\text{prompt}}$ ,  $E_{\text{delayed}}$ ,  $\Delta t$ ,  $\Delta \bar{x}$ ) allows to isolate IBD signal from the sea of single-events
- Strategies to deal with residual background
  - Passive shielding (PE, B, Fe,  $\text{H}_2\text{O}$ ) & active vetoes
  - Pulse shape discrimination (PSD)
  - Statistical subtraction of **accidental coincidences** & **cosmogenic background** (reactor OFF)



# Different Reactors and Technologies

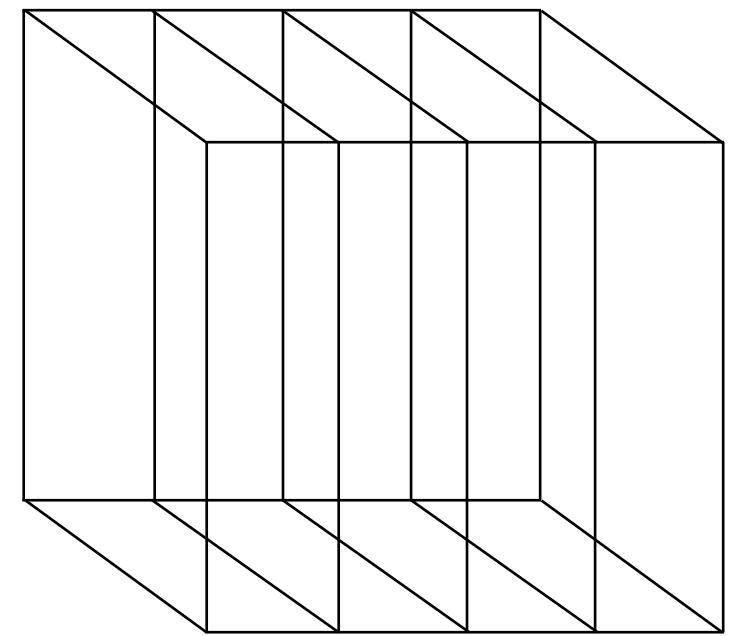
## Detector segmentation

no-segmentation



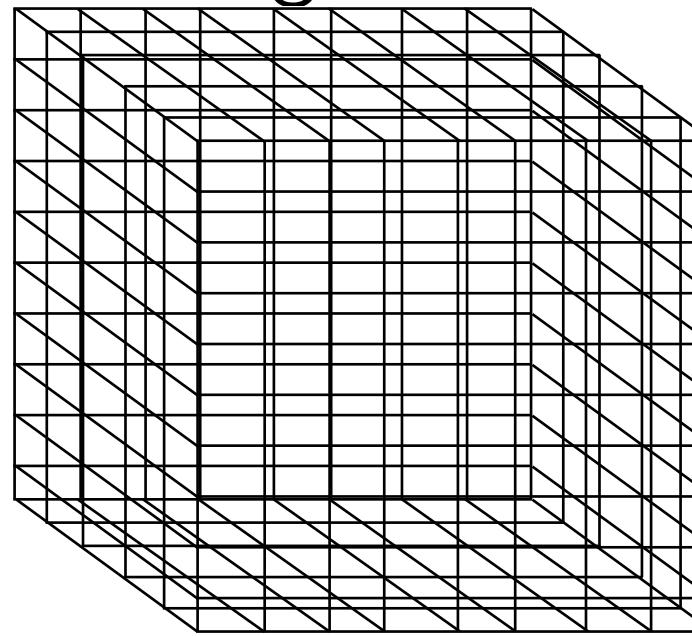
compare  $\bar{\nu}$  spectrum  
with predictions

coarse segmentation



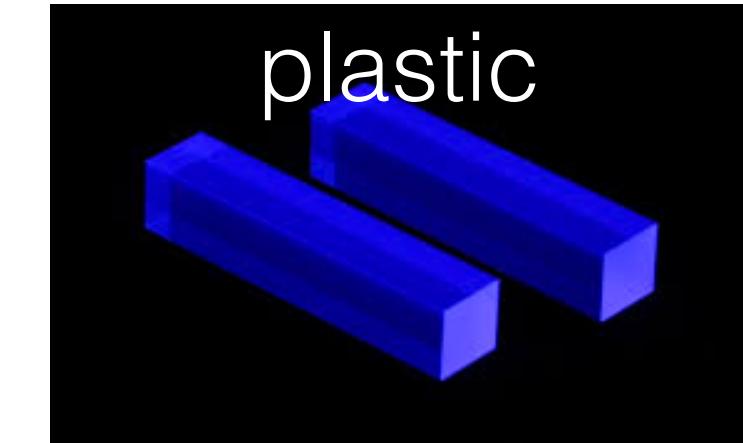
compare  $\bar{\nu}$  spectra  
in different segments  
(model free)

fine segmentation



compare  $\bar{\nu}$  spectra in  
sections + background  
rejection w/ topology

## Scintillator



better for  
segmentation  
& detection  
efficiency



Easier to have  
large volumes

## Reactor

research reactor (HEU)



Short baseline & compact  
core, no fuel evolution

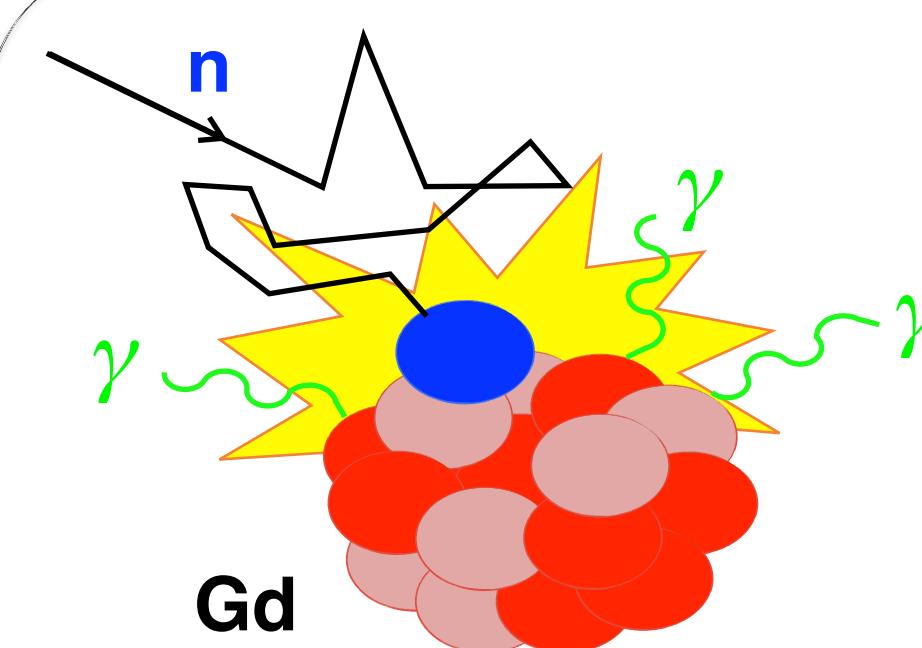
$\mathcal{O}10^2 \text{ MW}_{\text{th}}$ , limited space,  
background from facility

power reactor (LEU)

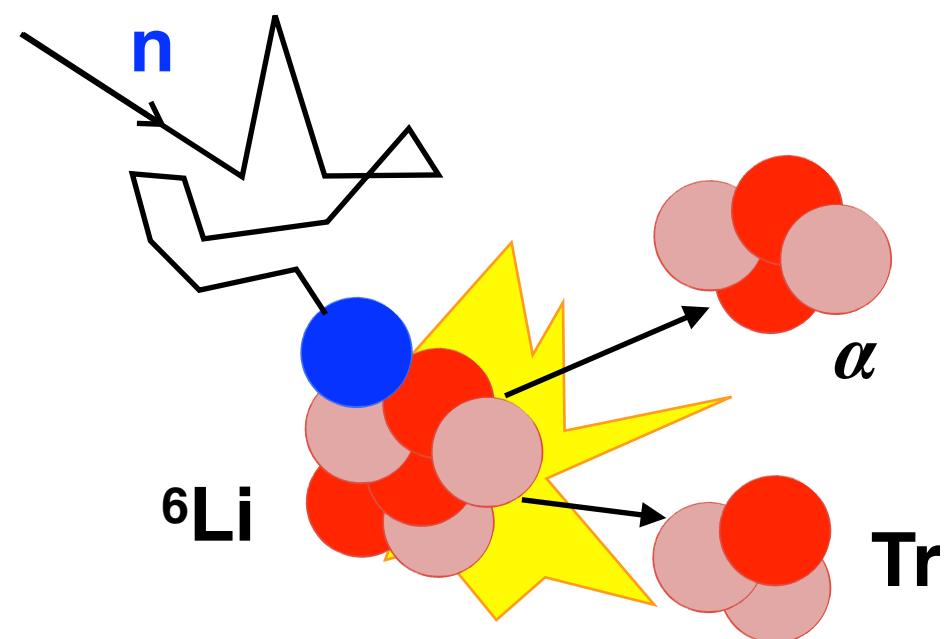


$\mathcal{O}\text{GW}_{\text{th}}$ , some  
overburden possible  
Lower sensitivity at low  
energy, fuel burnup

## Neutron-capturing isotope

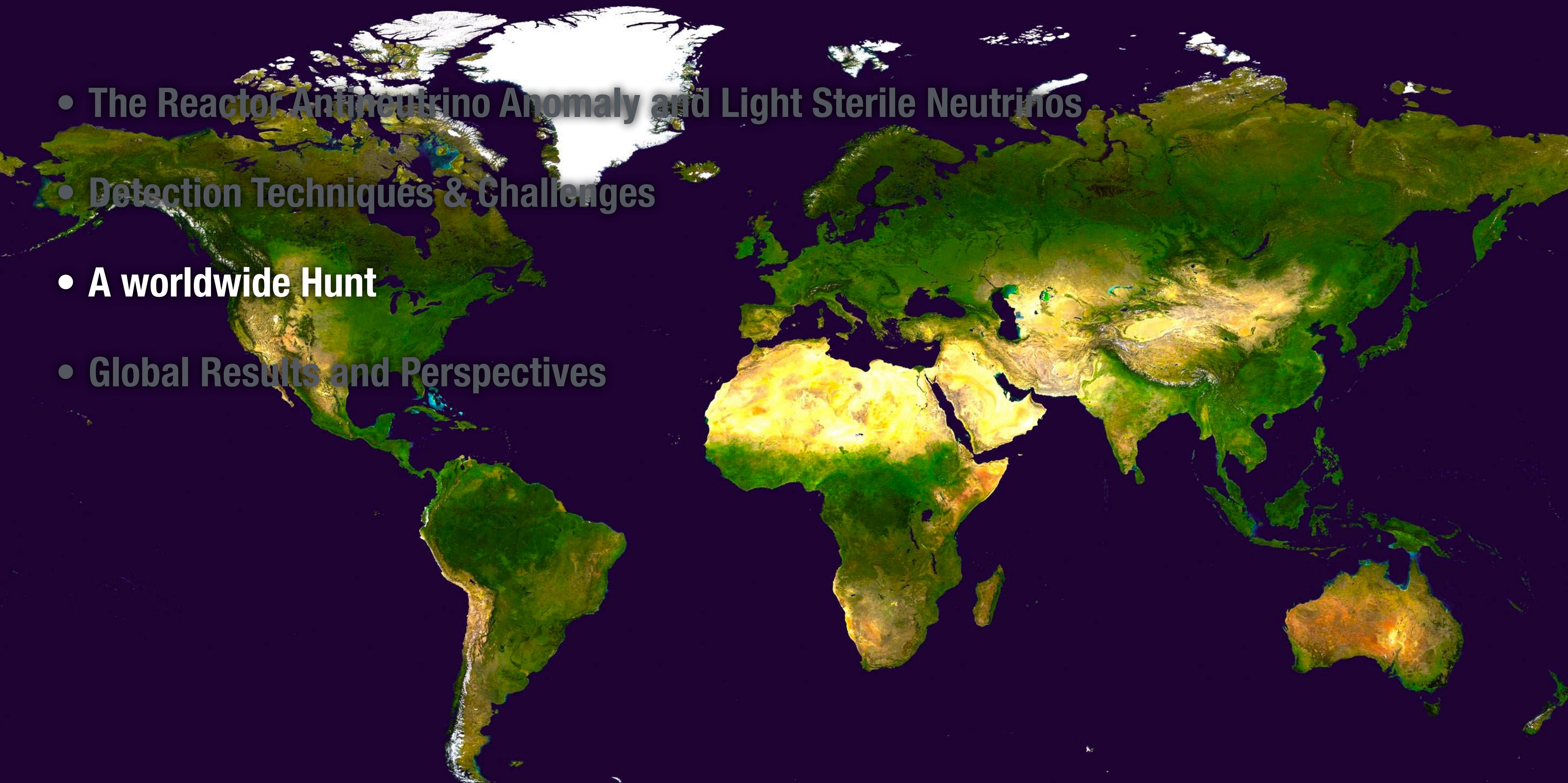


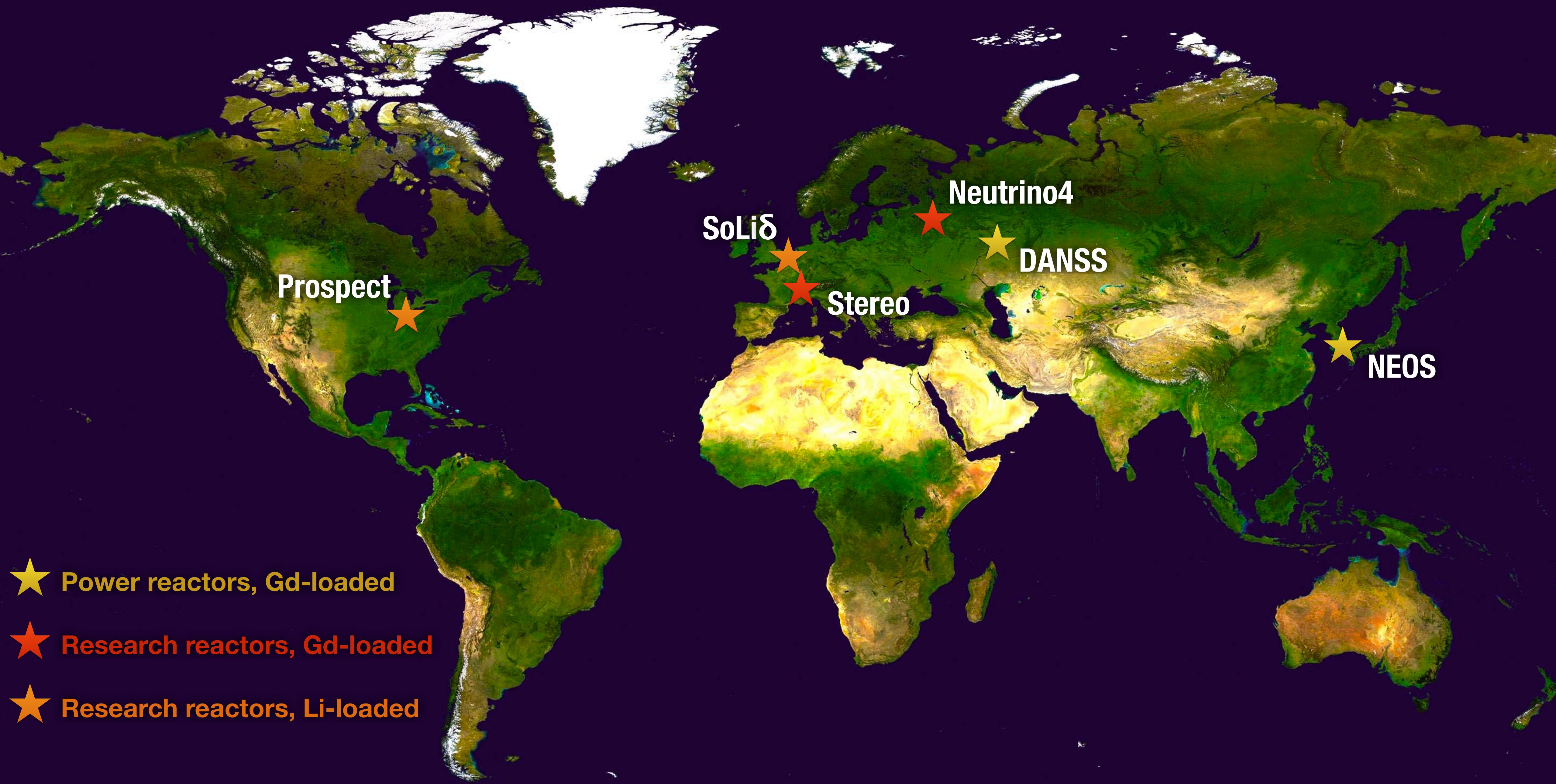
Well-established,  
high  $E_{\text{dep}}$  &  $\sigma_{\text{capture}}$



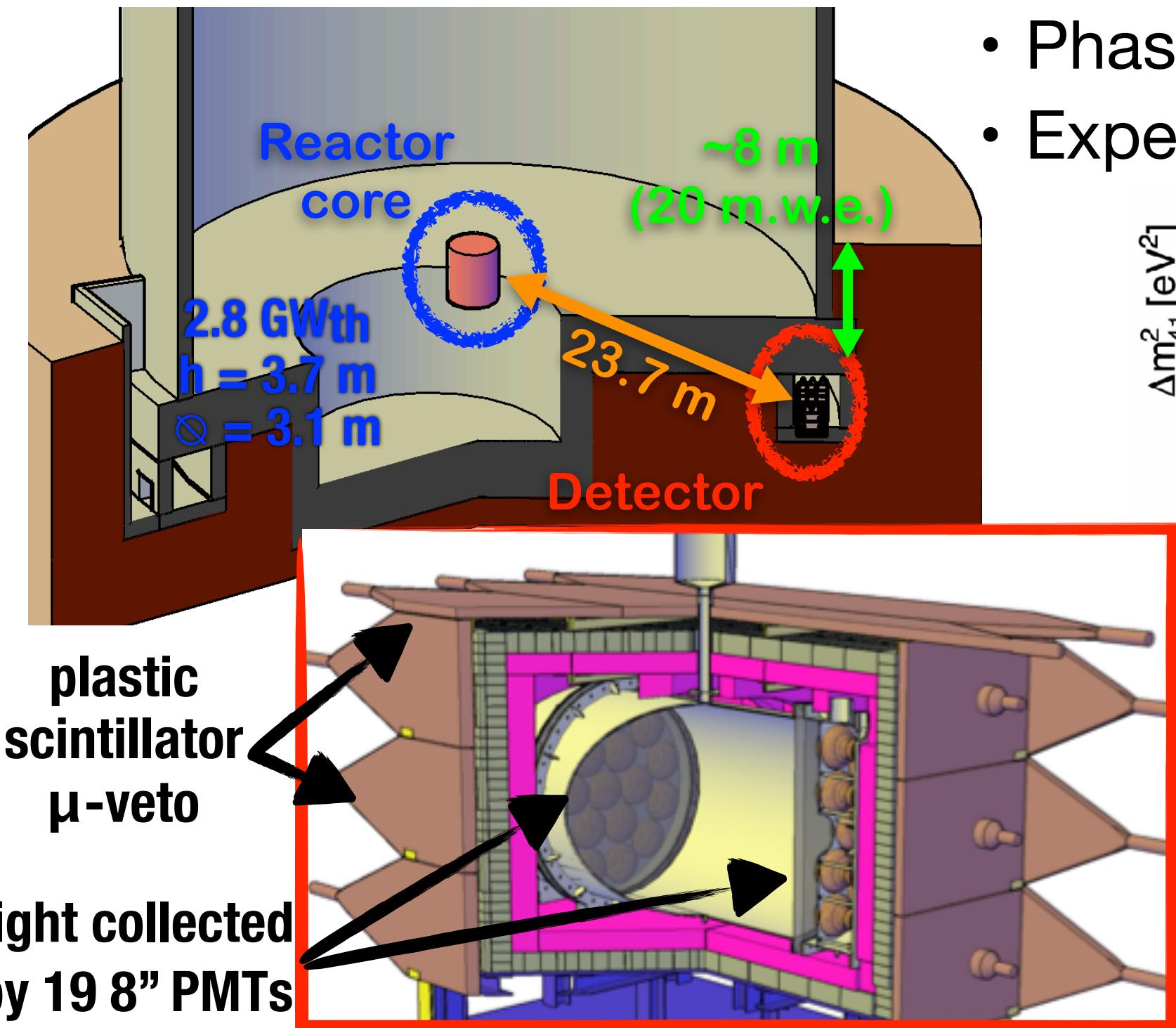
Localised  $E_{\text{dep}}$ : quenched  
but can select via PSD

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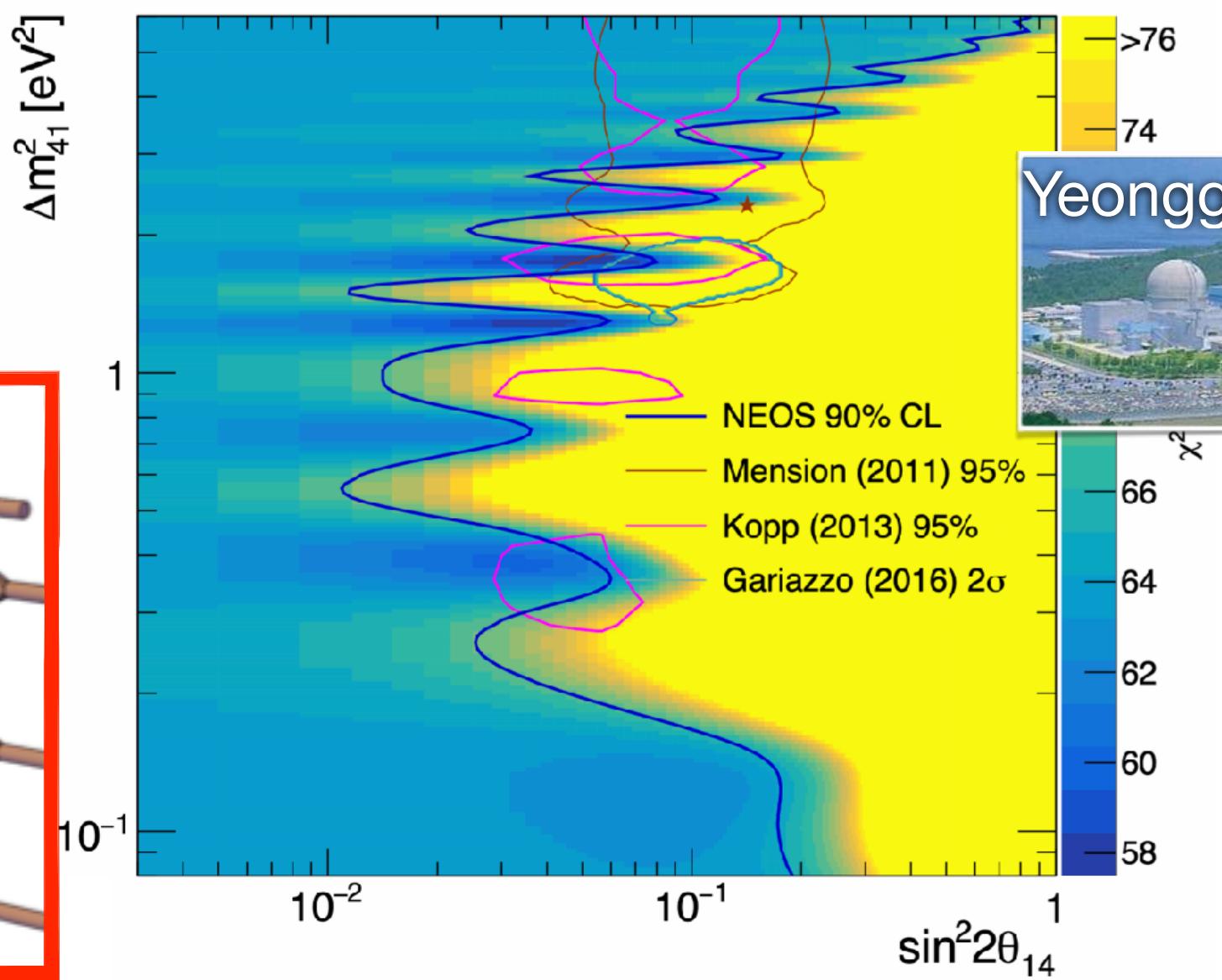




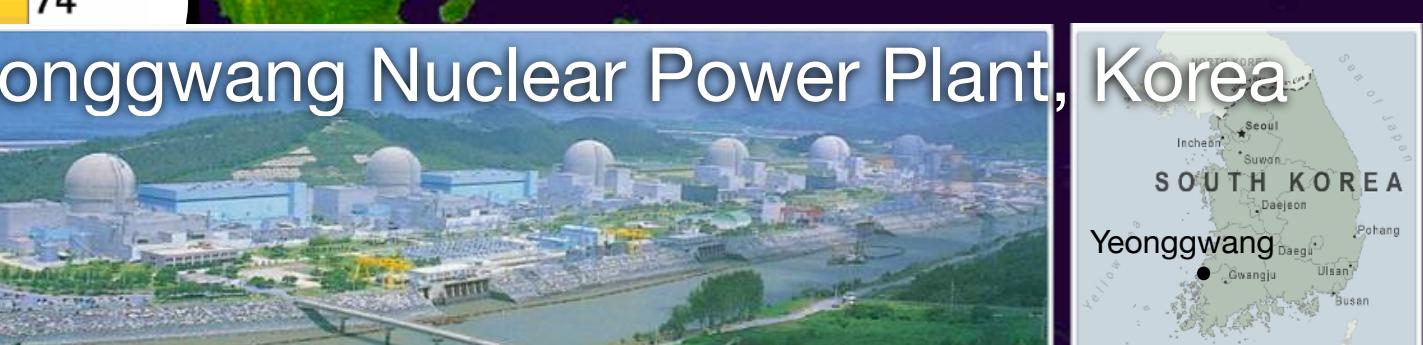
- **Simple design:** 1008 L Gd-loaded (0.48%) liquid scintillator tank spectrum compared with Data Bay
- **Very high statistics** (~2000 ν/day) thanks to the powerful (2.8 GW) commercial reactor
- **Degradation of light yield in time**



- Phase I (46 day OFF + 180 ON)
- Expected X2 sensitivity w/ phase I+II



# NEOS



Yeonggwang Nuclear Power Plant, Korea

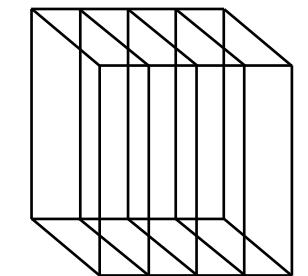
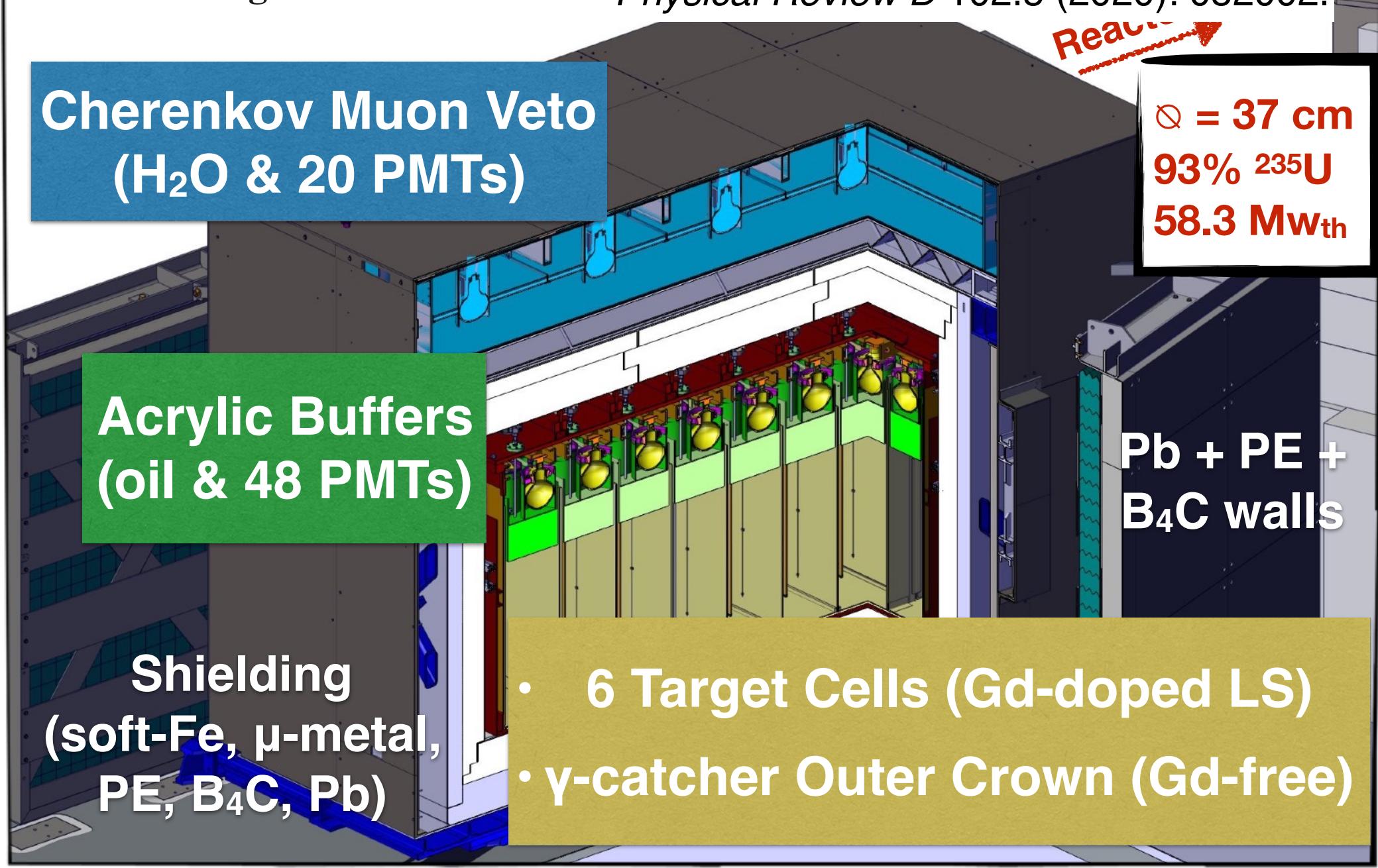
SOUTH KOREA



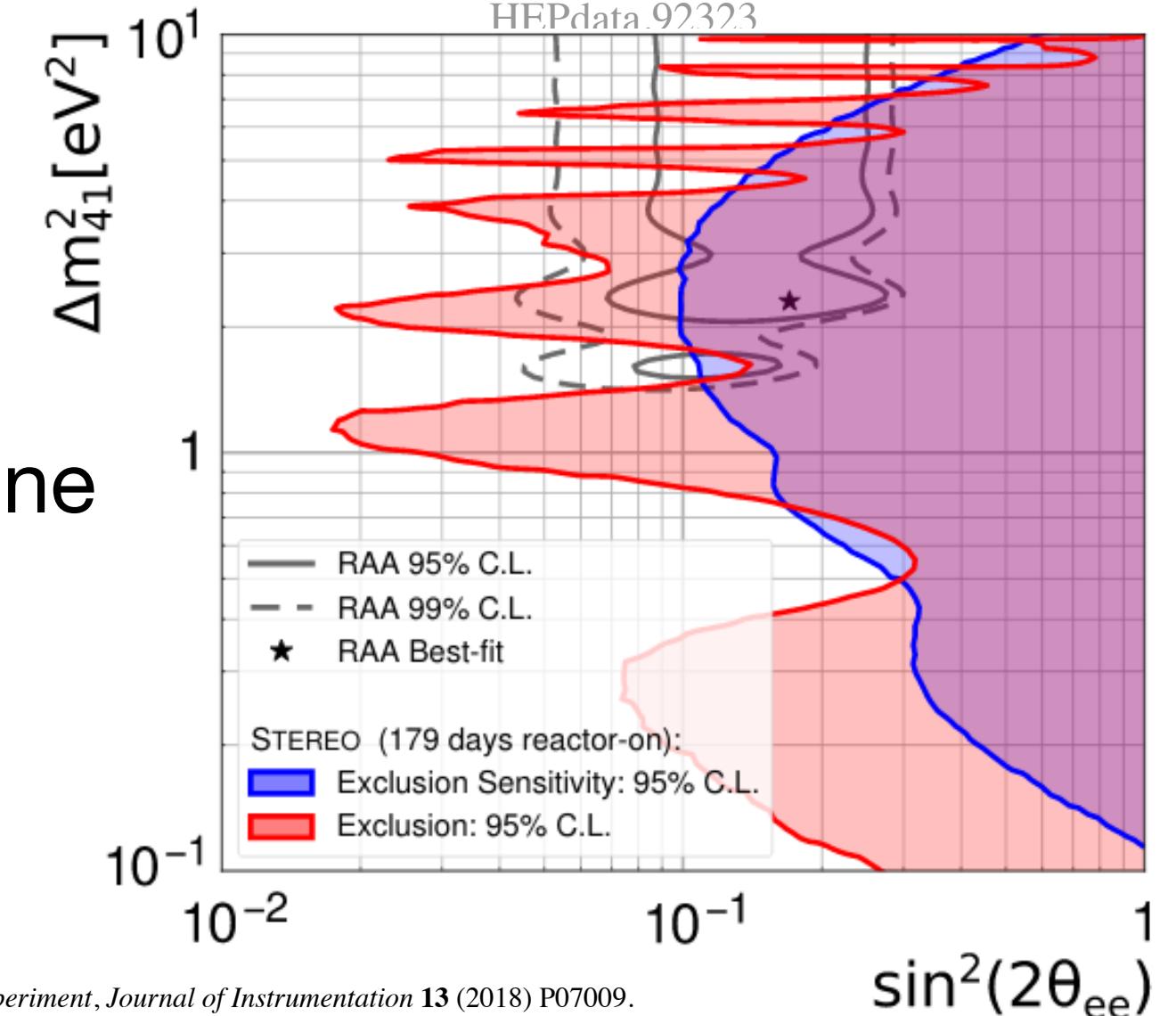
Neutrino target:  $2.2 \times 1.5 \times 1 \text{ m}^3$

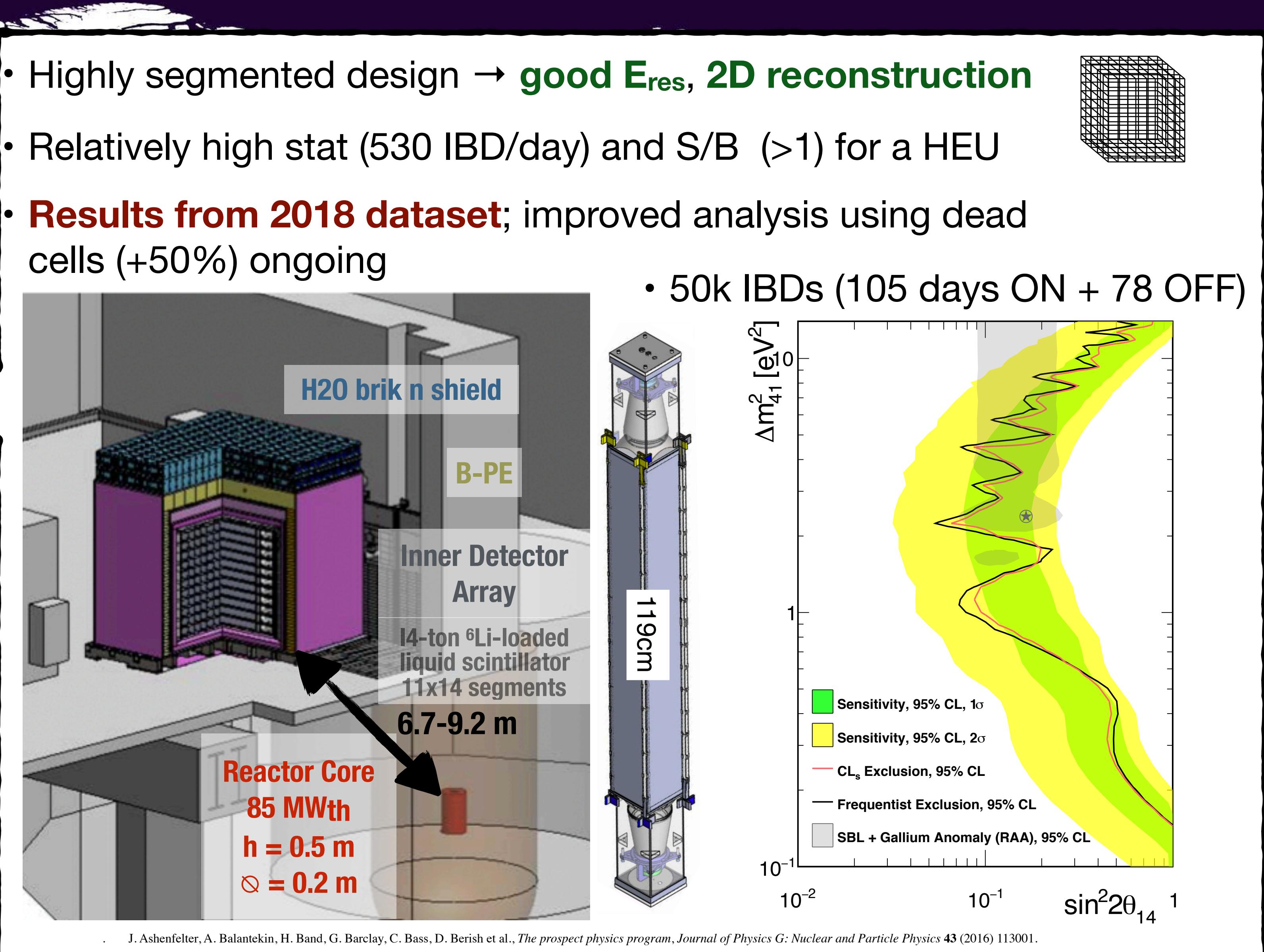
*Physical Review D* 102.5 (2020): 052002.

**Cherenkov Muon Veto**  
( $\text{H}_2\text{O}$  & 20 PMTs)



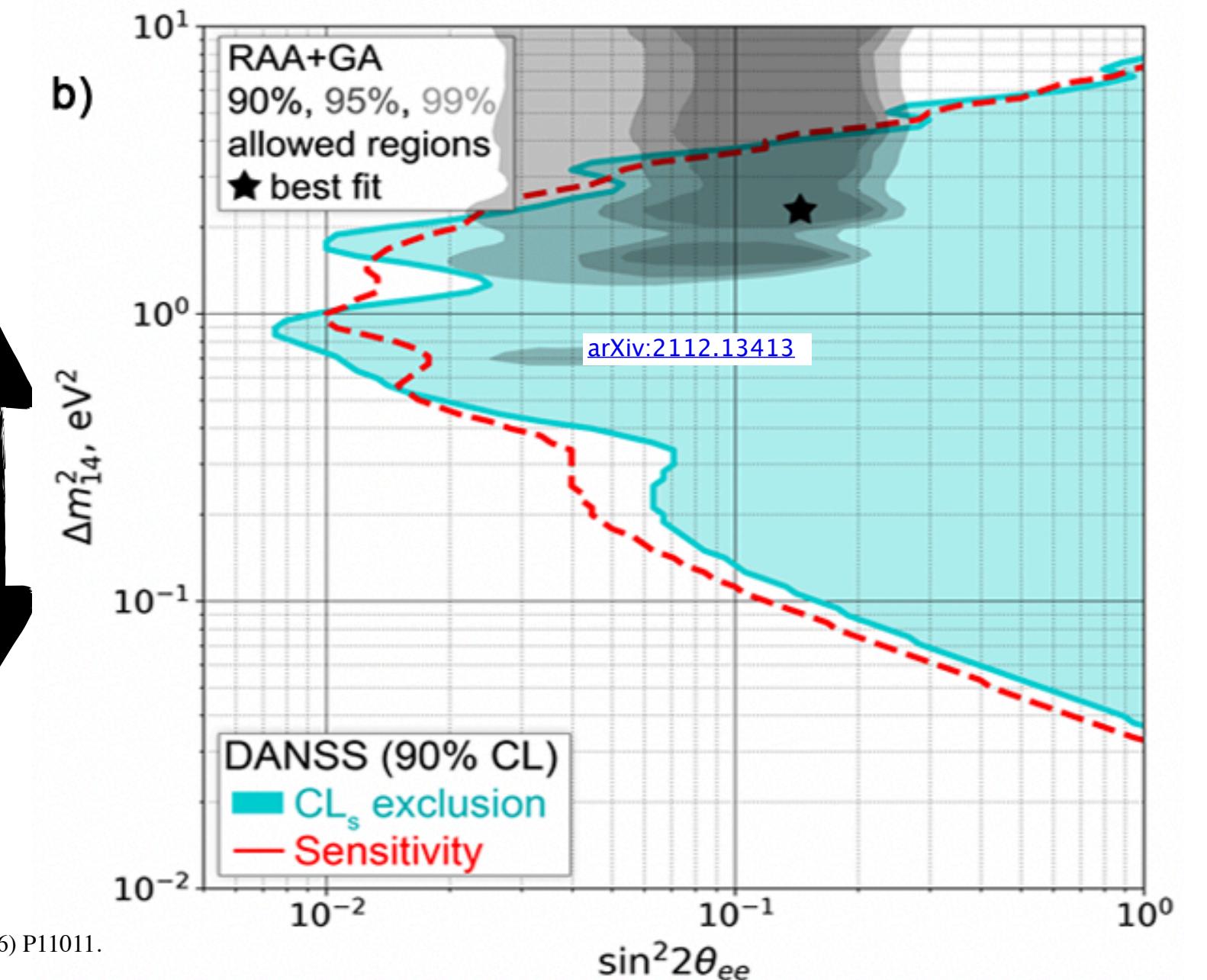
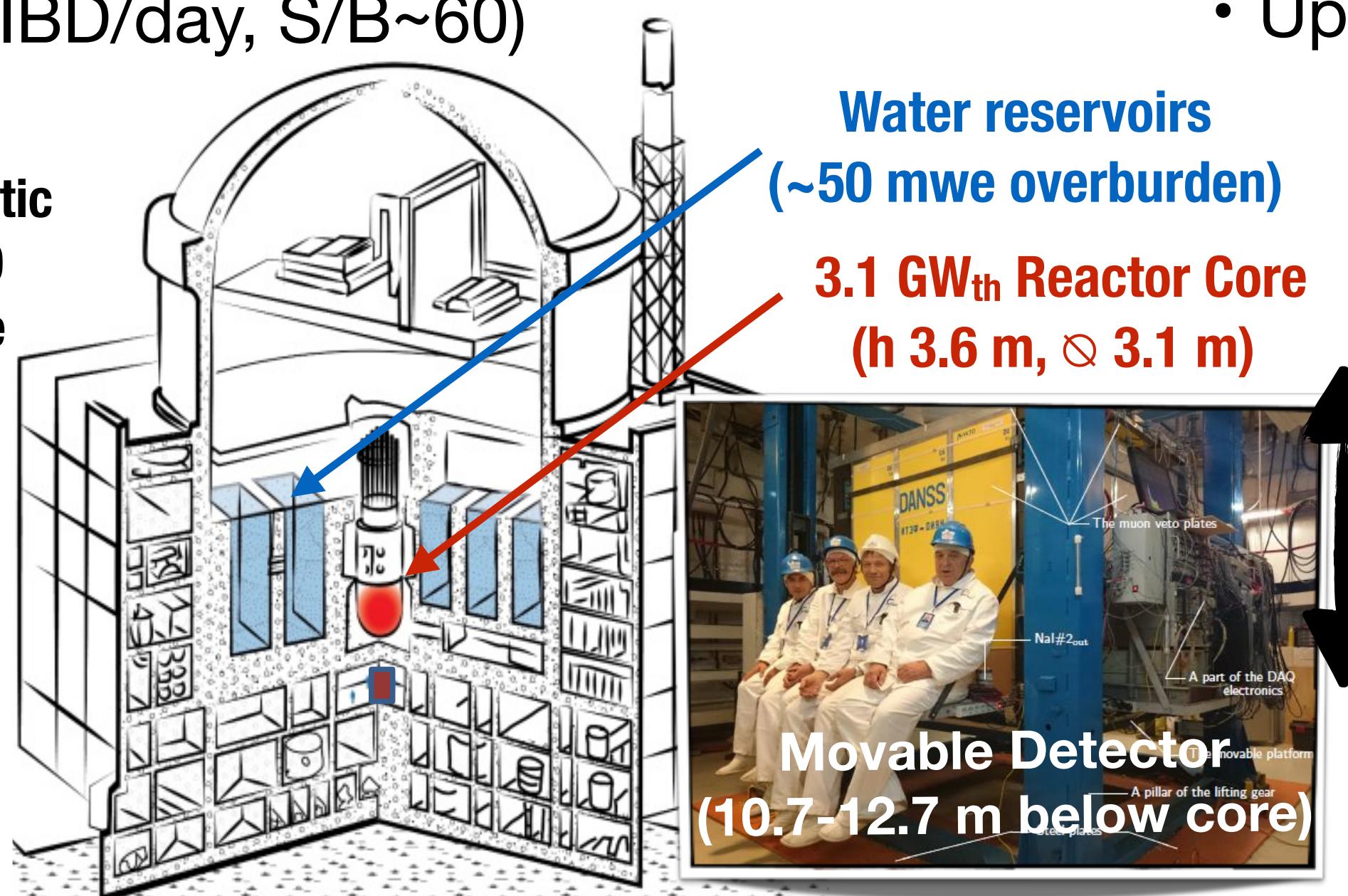
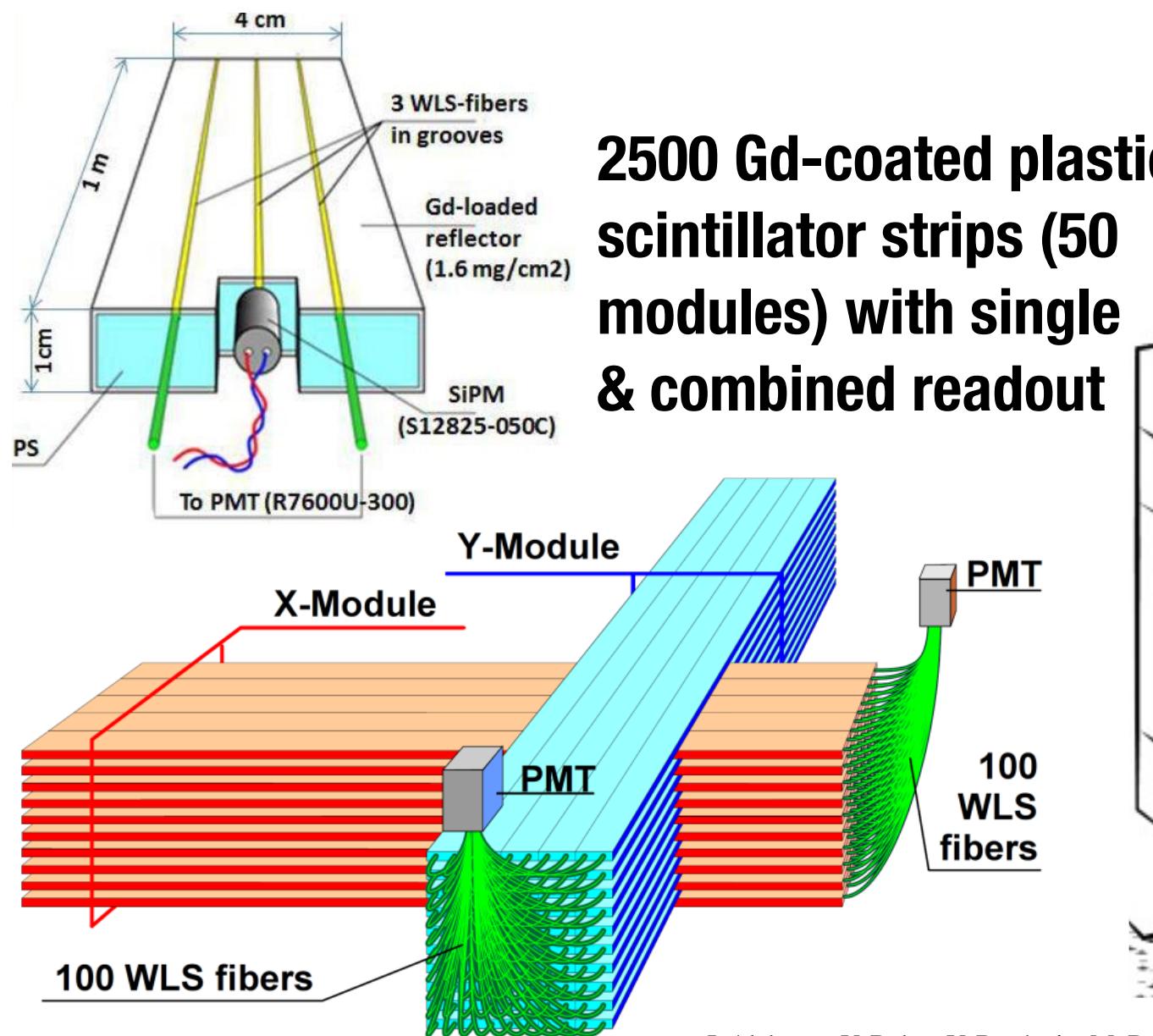
- Phase I+II (65k IBDs, 179 day ON + 235 OFF)
- Expected X2 sensitivity with full dataset







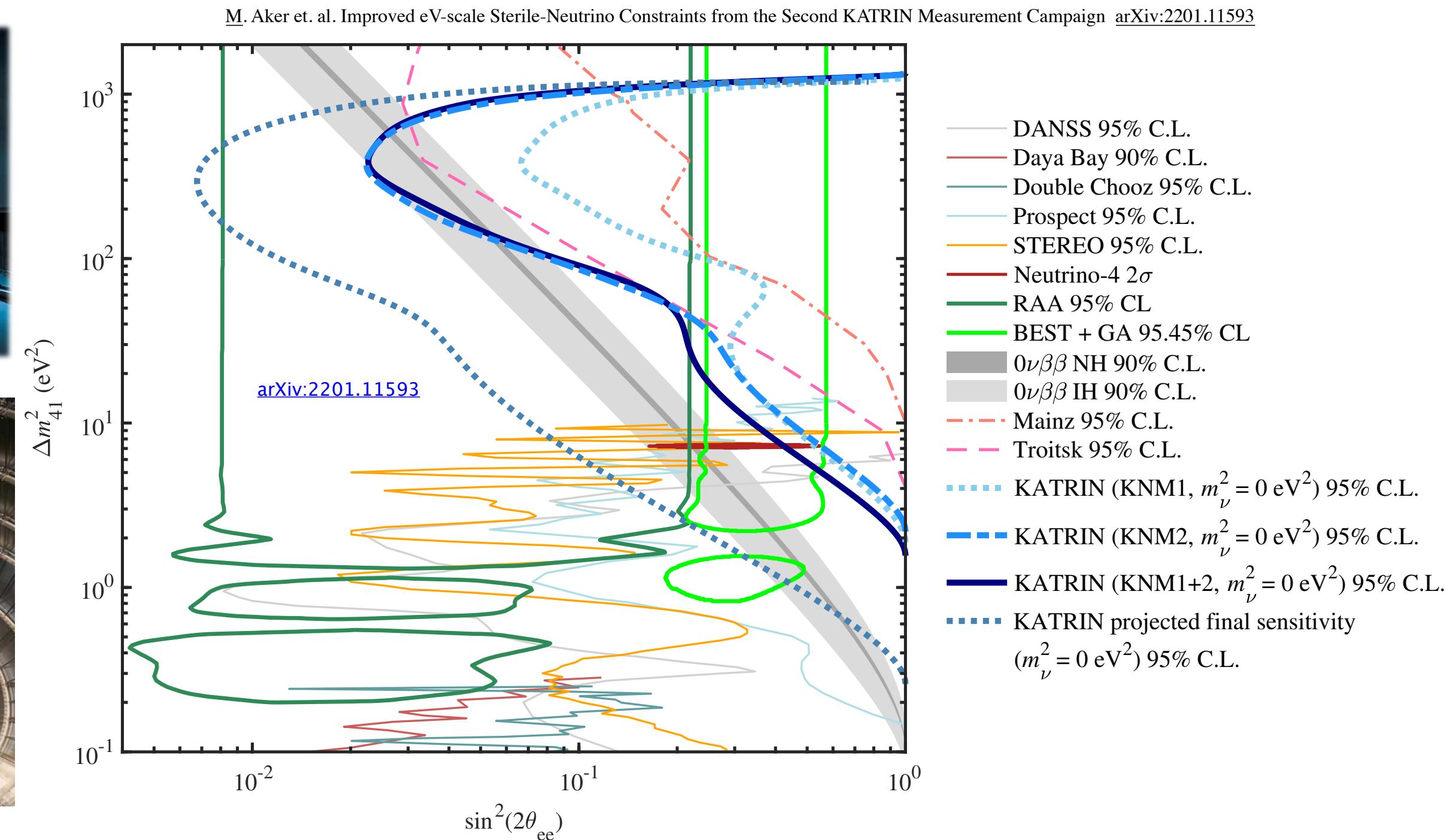
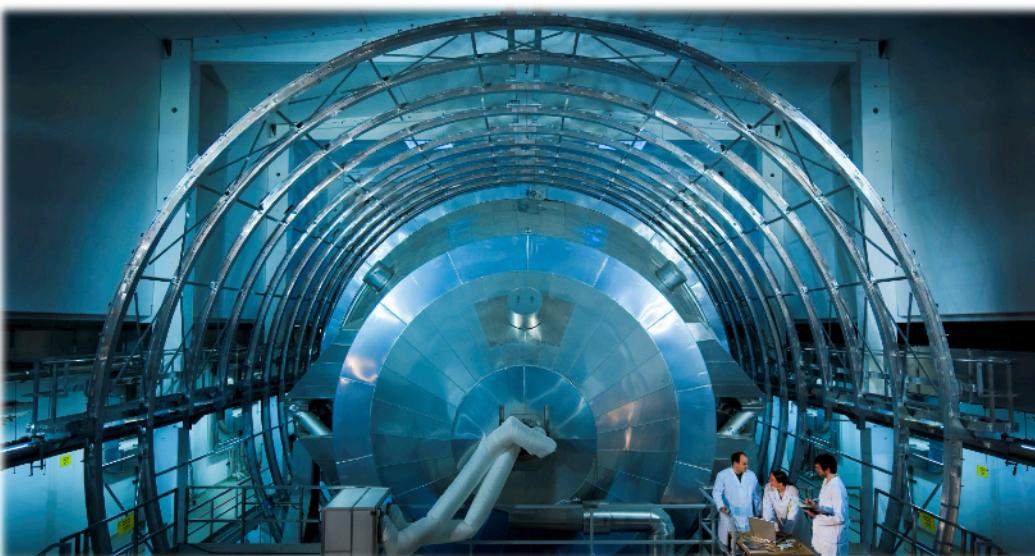
- Highly-segmented  $\nu$  spectrometer → **quasi-3D reconstruction**
- **Excellent statistics** (~5000 IBD/day, S/B~60)
- Phase-I (5.5 M IBD in 5 years)
- Upgrade ongoing → halve  $E_{\text{res}}$  increase  $V_{\text{det}}$



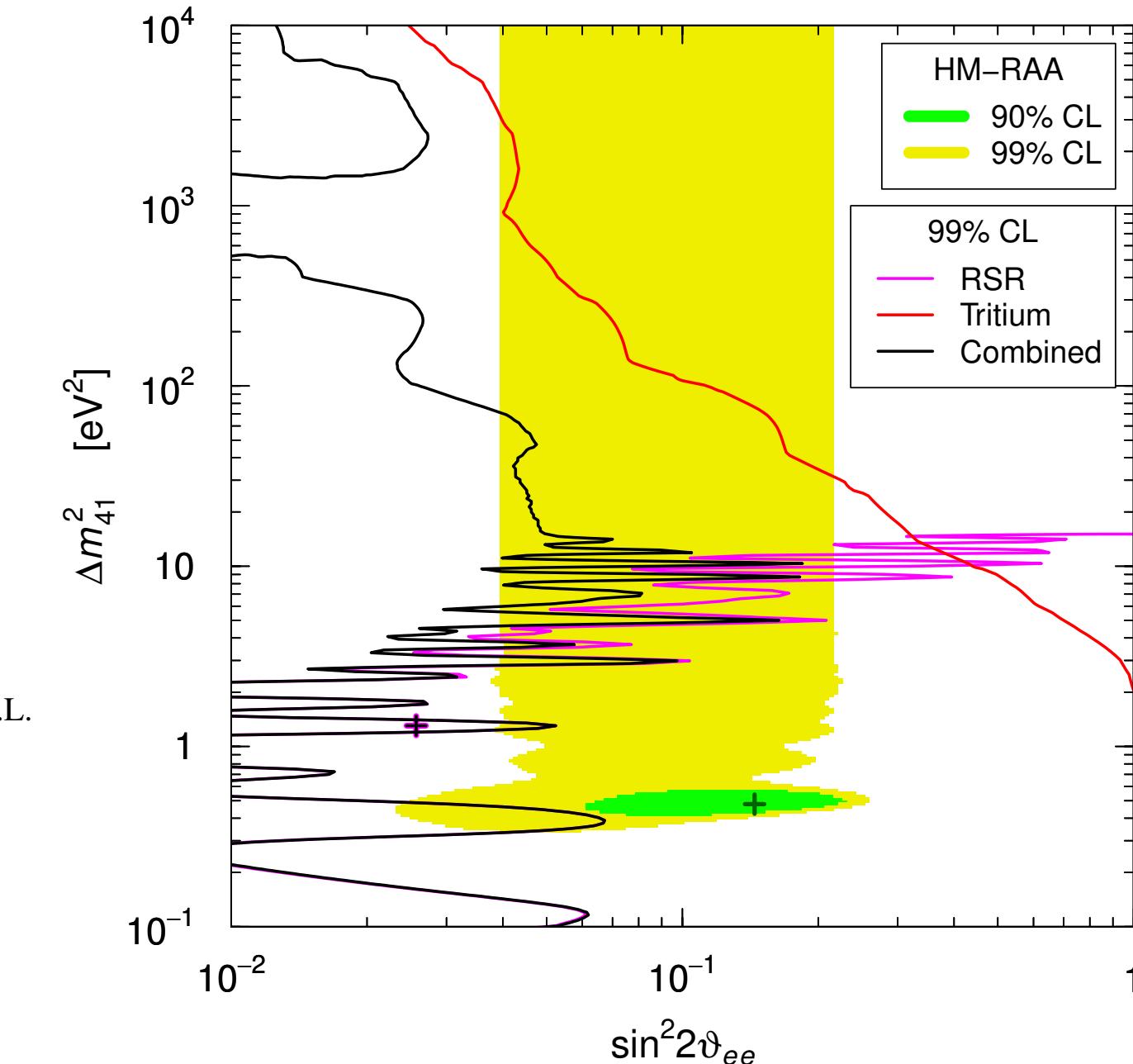
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# Reactor Anomaly and Recent Results: the Global Picture

- Each of the 4 experiments: **DANSS, NEOS, STEREO, PROSPECT** excluded large portions of the RAA region & the best fit value ( $> 90\% \text{ CL}$ ), while **Neutrino-4** claims a  **$2.8\sigma$  CL observation** of a  $\Delta m^2 \sim 7.3 \text{ eV}^2$  oscillation
- STEREO, PROSPECT, DANSS are releasing their  $\chi^2$  tables and data to help combined fits
- **KATRIN**, a 200t spectrometer for the measurement of the  $\nu_e$  mass, has also published results on sterile  
→ **exclusion of RAA region with strong synergy with the short-baseline reactor searches**



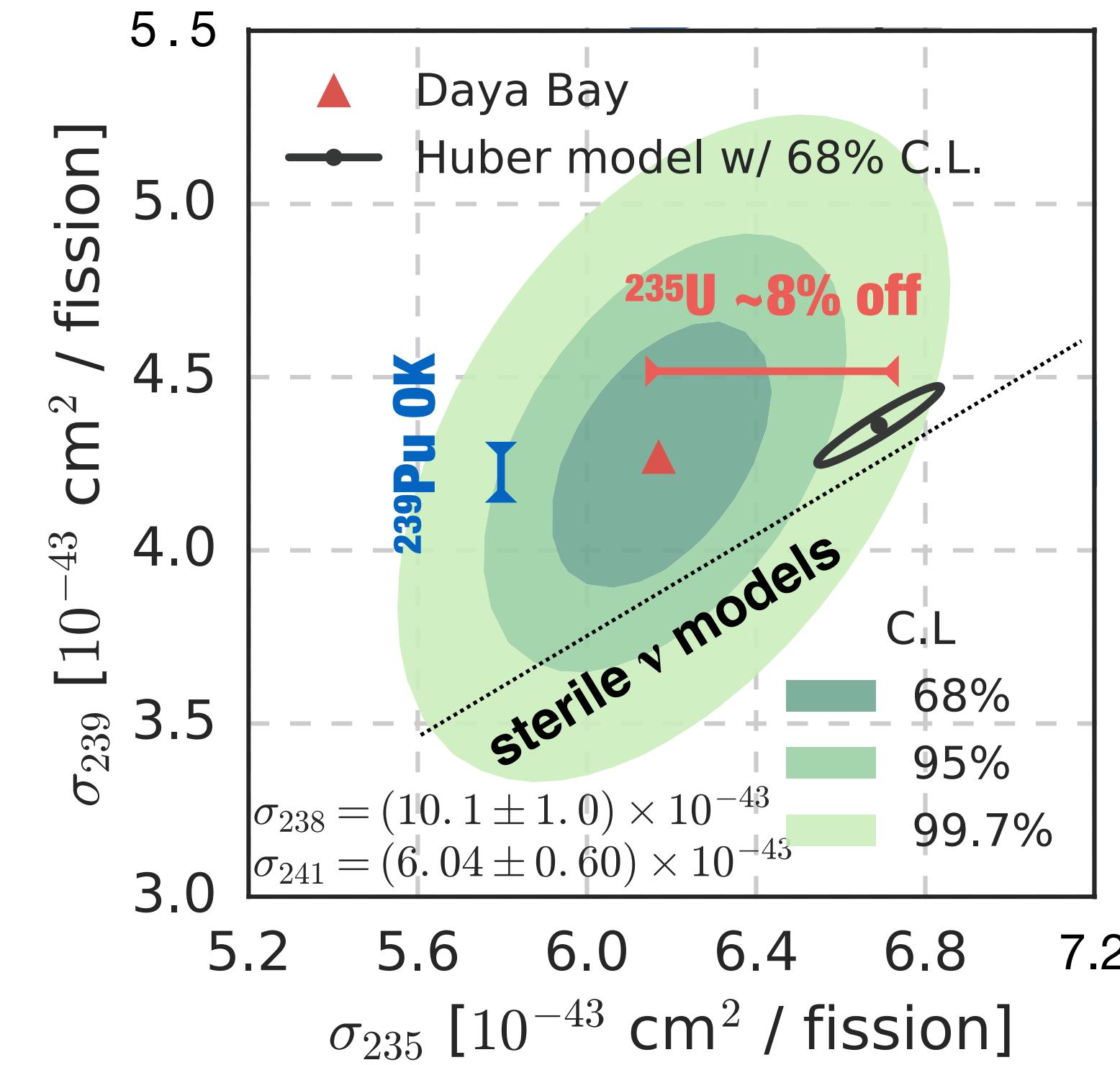
Giunti, C., Li, Y. & Zhang, Y. KATRIN bound on 3+1 active-sterile neutrino mixing and the reactor antineutrino anomaly. *J. High Energ. Phys.* **2020**, 61 (2020).



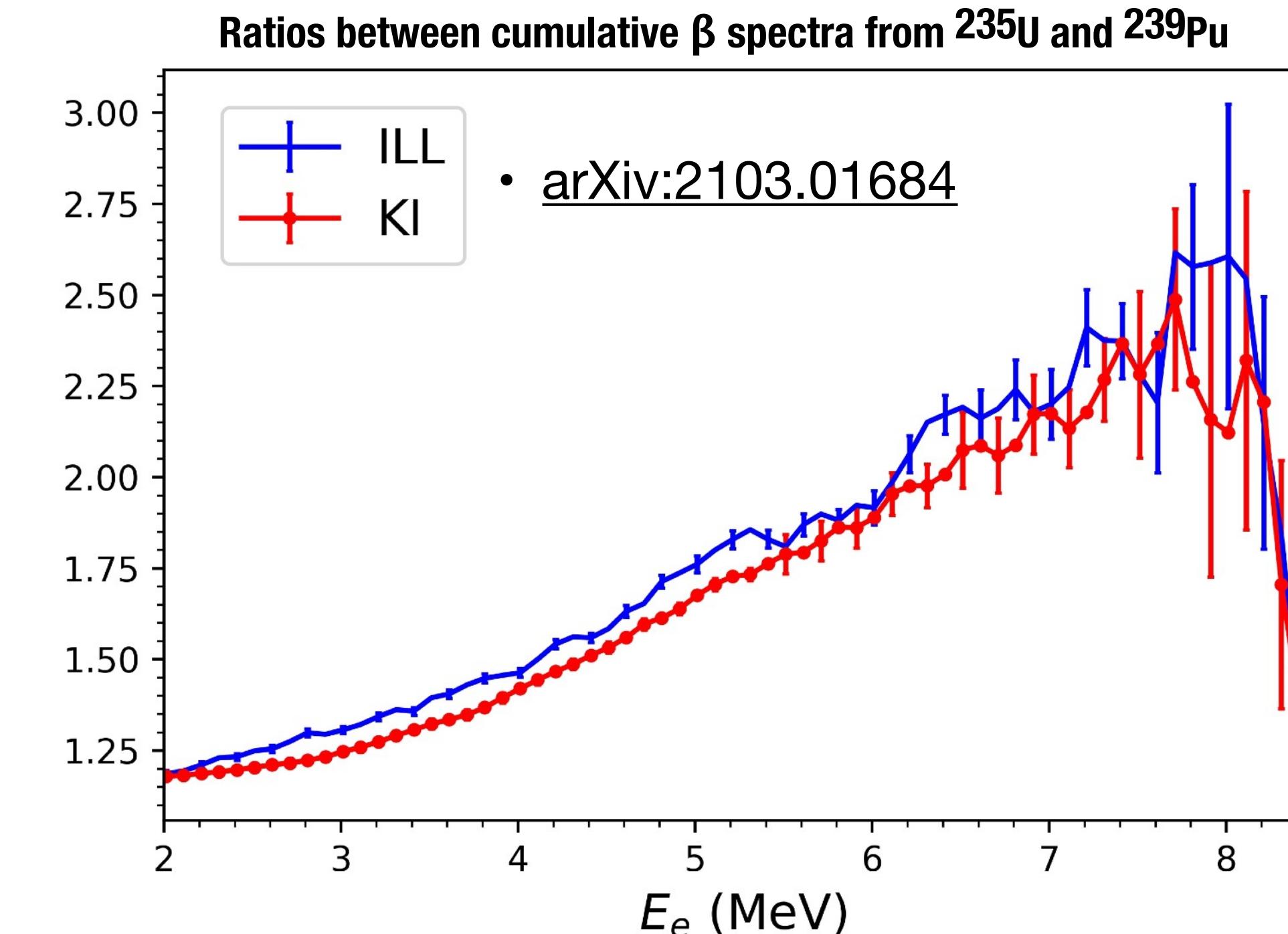
# On Isotopes and Antineutrino Flux & Spectrum

- Daya Bay and RENO can separate  $^{235}\text{U}$  and  $^{239}\text{Pu}$  contribution to  $\bar{\nu}_e$  flux → **RAA rate deficit mainly from  $^{235}\text{U}$**
- Estimation of antineutrino spectra based on global  $\beta$  spectra measured at ILL; recent re-evaluation in Kurchatov Institute suggest a ~5% excess in  $^{235}\text{U}$  to  $^{239}\text{Pu}$  ratio for ILL (compatible with the RAA excess)

F. An, A. Balantekin, H. Band, M. Bishai, S. Blyth, D. Cao et al., *Evolution of the reactor antineutrino flux and spectrum at daya bay*, *Physical review letters* **118** (2017) 251801.



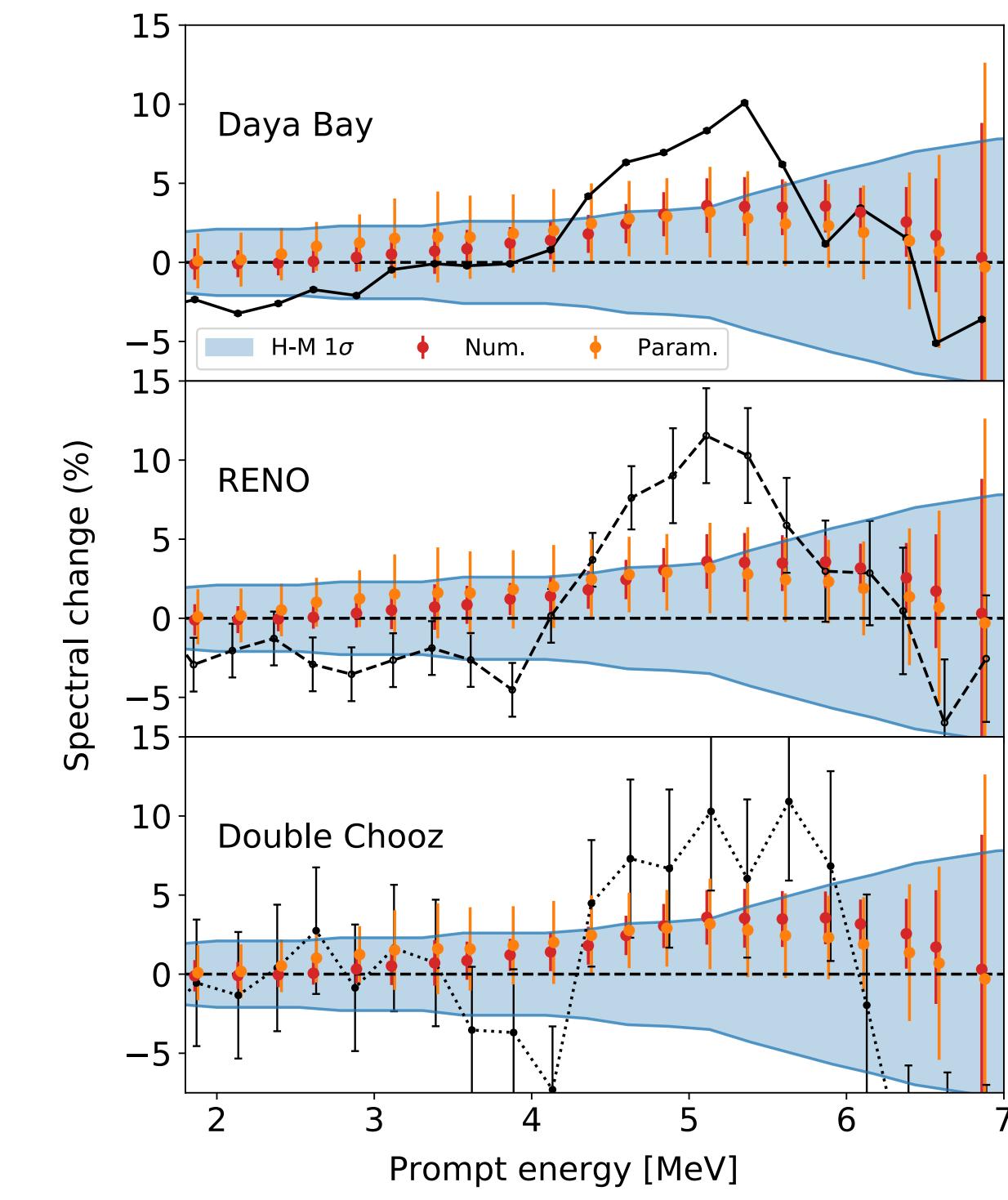
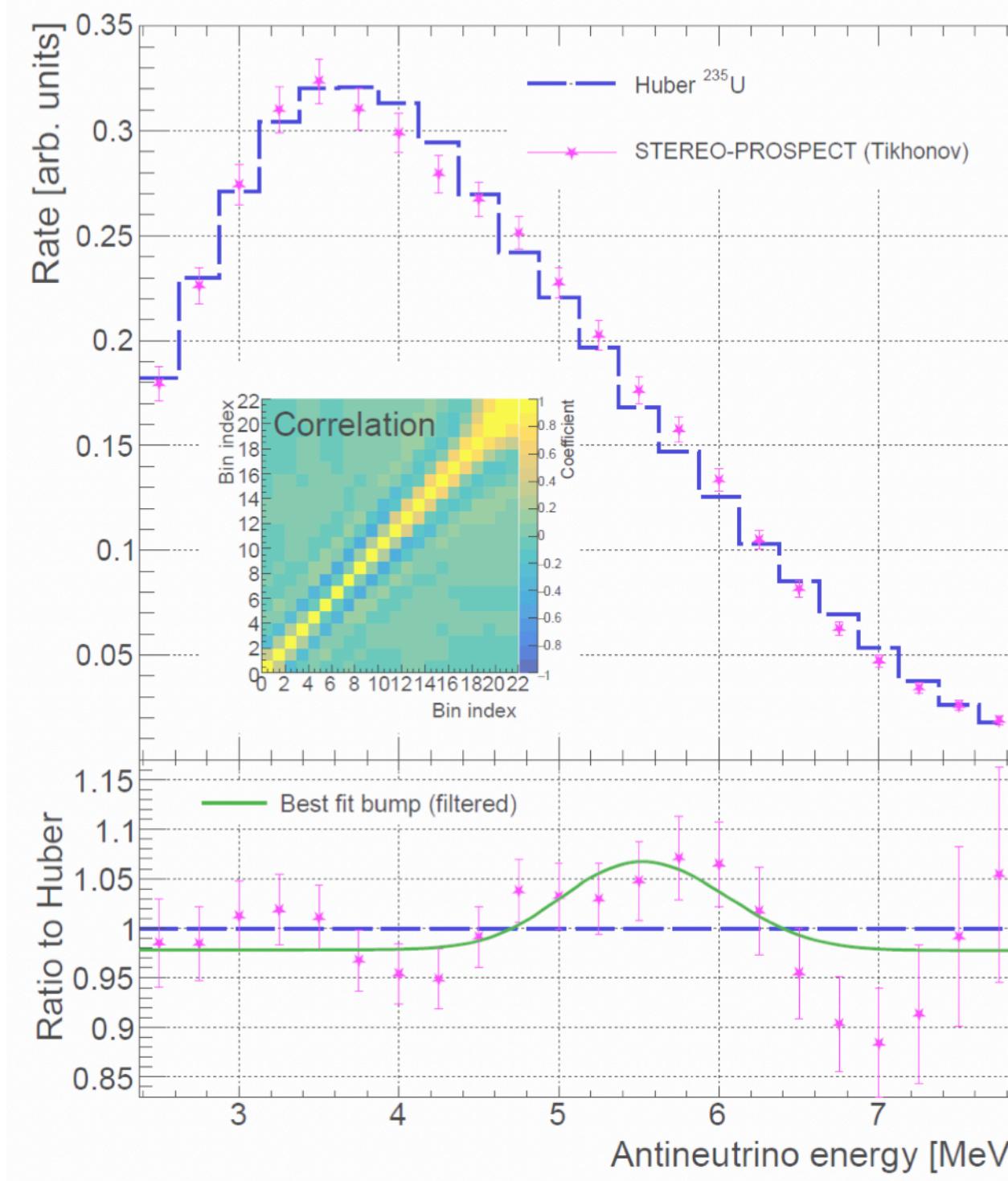
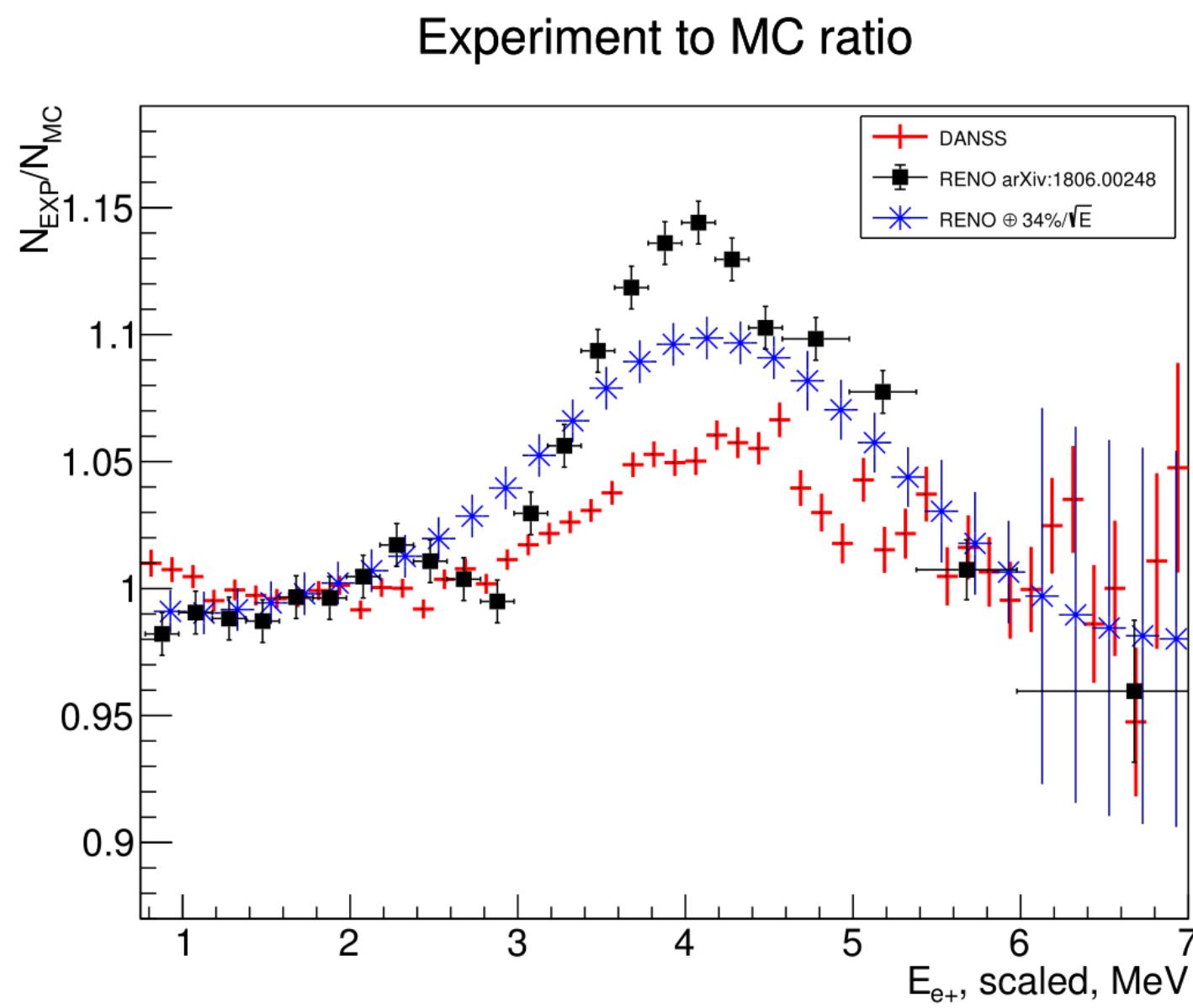
V. Kopeikin, M. Skorokhvatov, O. Titov Reevaluating reactor antineutrino spectra with new measurements of the ratio between  $^{235}\text{U}$  and  $^{239}\text{Pu}$   $\beta$  spectra [arXiv:2103.01684](https://arxiv.org/abs/2103.01684)



# On Isotopes and Antineutrino Flux & Spectrum

- A spectral distortion @  $E_\nu \sim 6$  MeV was observed in  $\theta_{13}$ -aimed neutrino experiments in 2014
- **STEREO & PROSPECT** released a **combine spectral analysis** confirming the distortion w/  **$2.4\sigma$  significance** and  **$A = 9.9 \pm 3.3\%$  for pure  $^{235}\text{U} \rightarrow$  distortion independent of other isotopes**
- H. Almazán, M. Andriamirado, A. Balantekin, H. Band, C. Bass, D. Bergeron et al., *Joint measurement of the  $^{235}\text{U}$  antineutrino spectrum by prospect and stereo*, arXiv preprint arXiv:2107.03371 (2021).
- Meanwhile, limits of current spectrum models are emerging, and the **treatment of forbidden decays could change both normalisation and spectral shape**

. A. Hayes, J. Friar, G. Garvey, D. Ibeling, G. Jungman, T. Kawano et al., *Possible origins and implications of the shoulder in reactor neutrino spectra*, Physical Review D **92** (2015) 033015.  
. L. Hayen, J. Kostensalo, N. Severijns and J. Suhonen, *First-forbidden transitions in reactor antineutrino spectra*, Physical Review C **99** (2019) 031301.

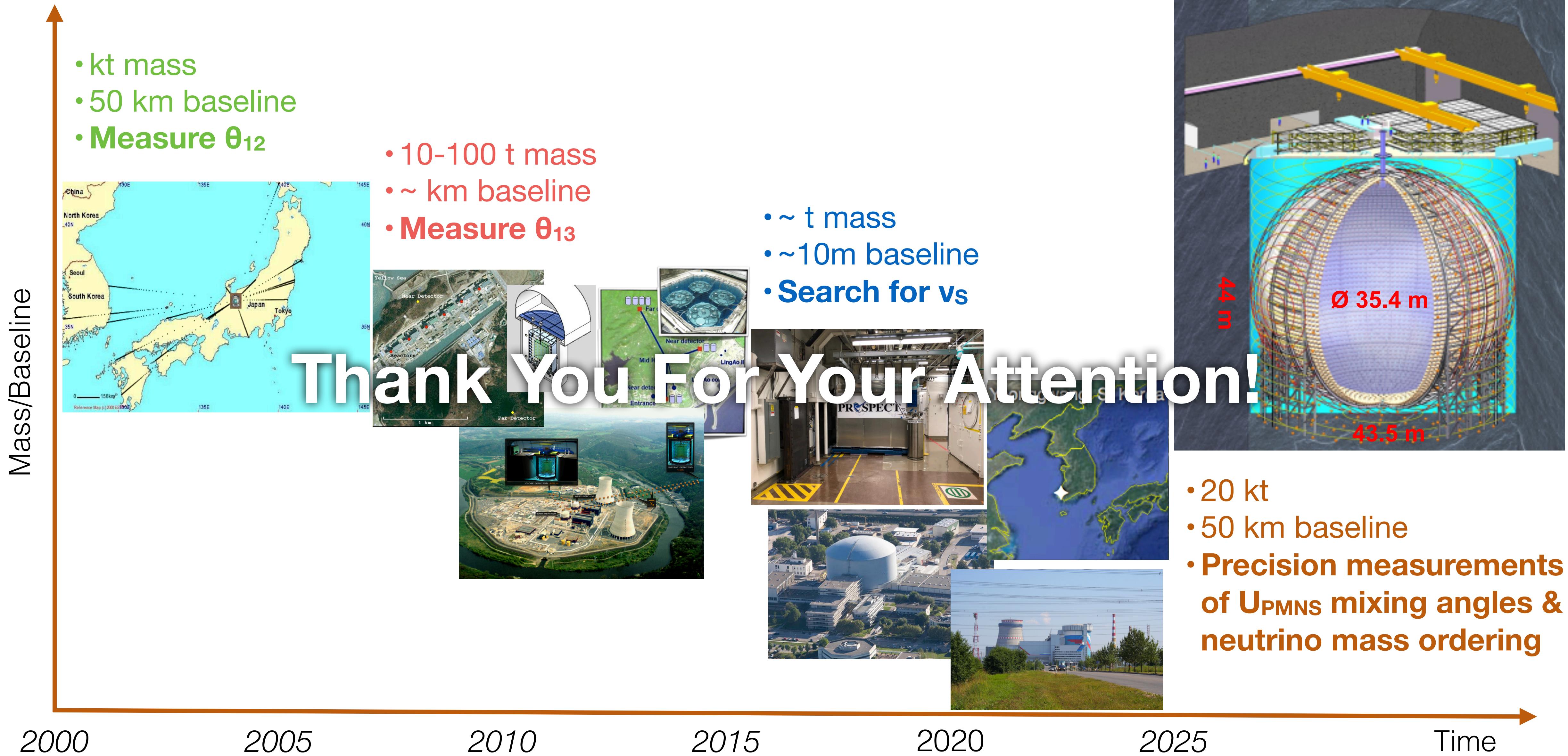


# Conclusions

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- The quest for  $\theta_{13}$  in the 2010's prompted **new models for reactor  $\bar{\nu}_e$  spectra** & a **~6% discrepancy with measured  $\bar{\nu}_e$  rates** was found (reactor antineutrino anomaly)
- **Several projects worldwide** were launched to study the anomaly and test the sterile neutrino hypothesis by **looking for neutrino oscillations at very short baseline**, and produced compelling results
- Overall, the **sterile neutrino hypothesis** as a solution of the RAA is **under increasing pressure** by experimental results and advancements in theoretical models
- Thanks to the these different contributions, we are starting to **better understand our antineutrino rates & spectra**, as well as our detectors; an important effort in view of the future of reactor neutrino physics

# ...I'll Leave You With This





Extra Slides

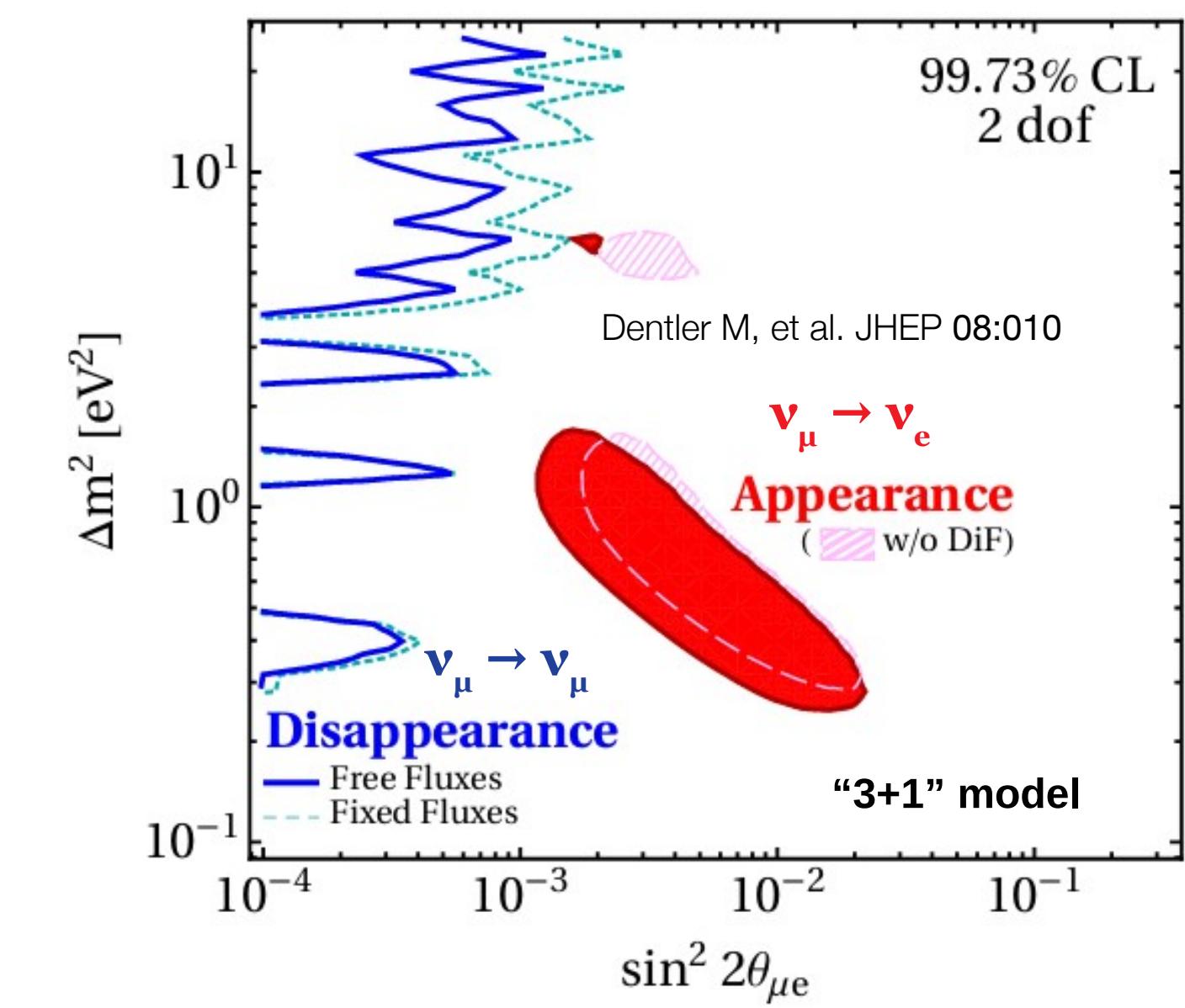
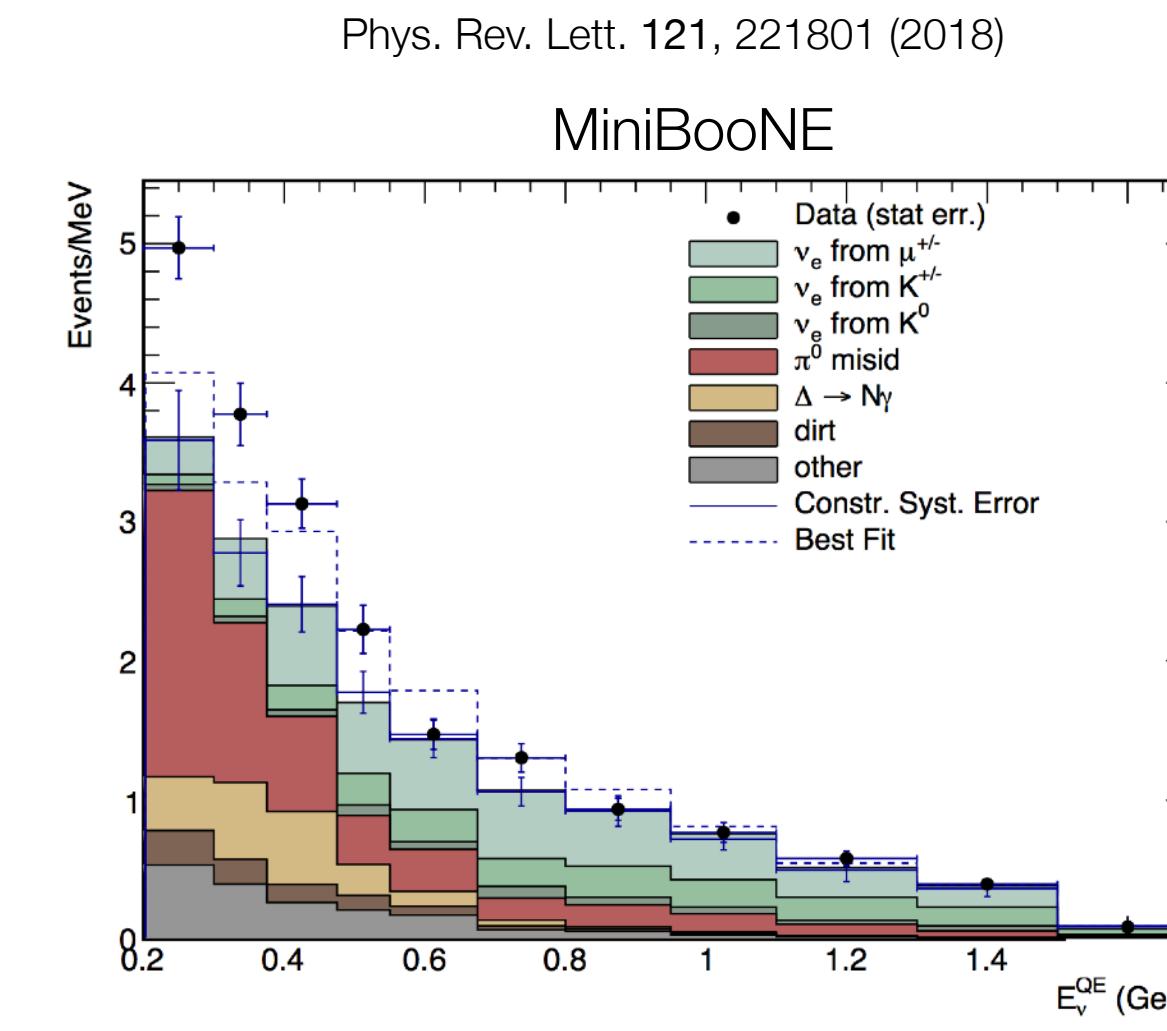
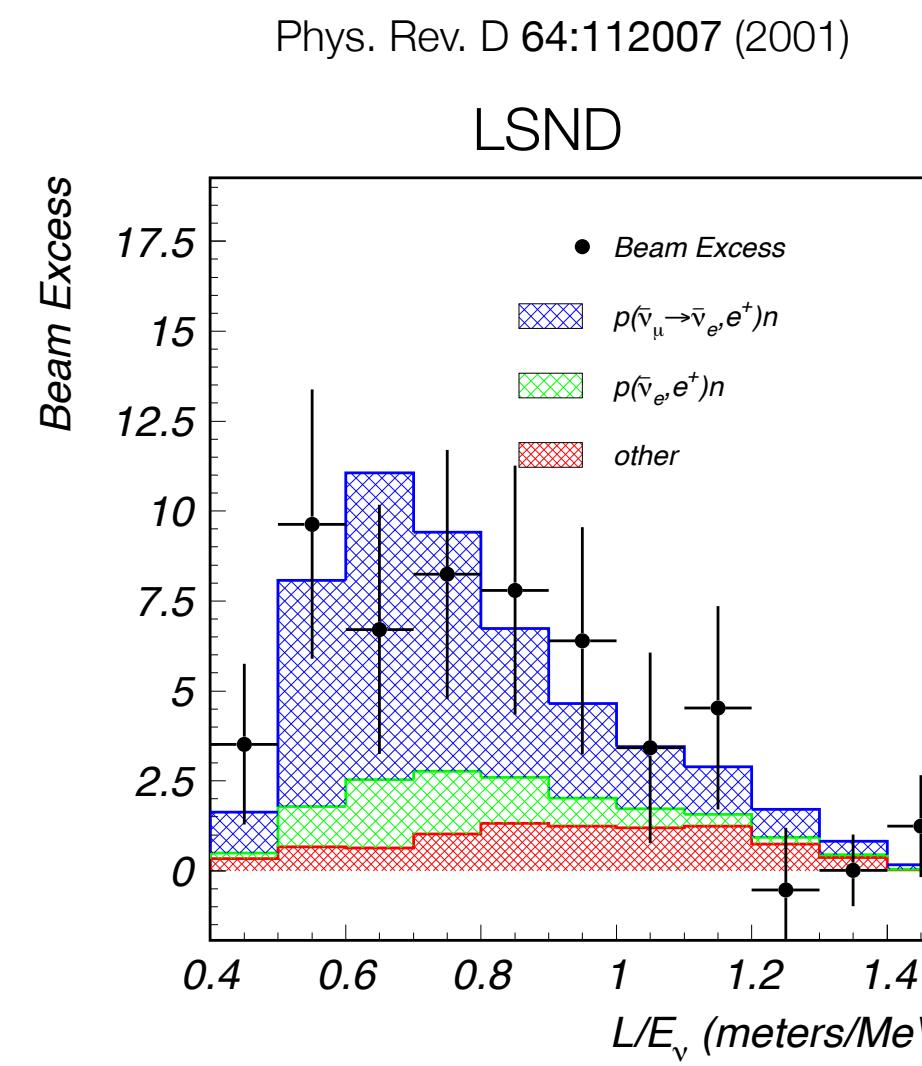
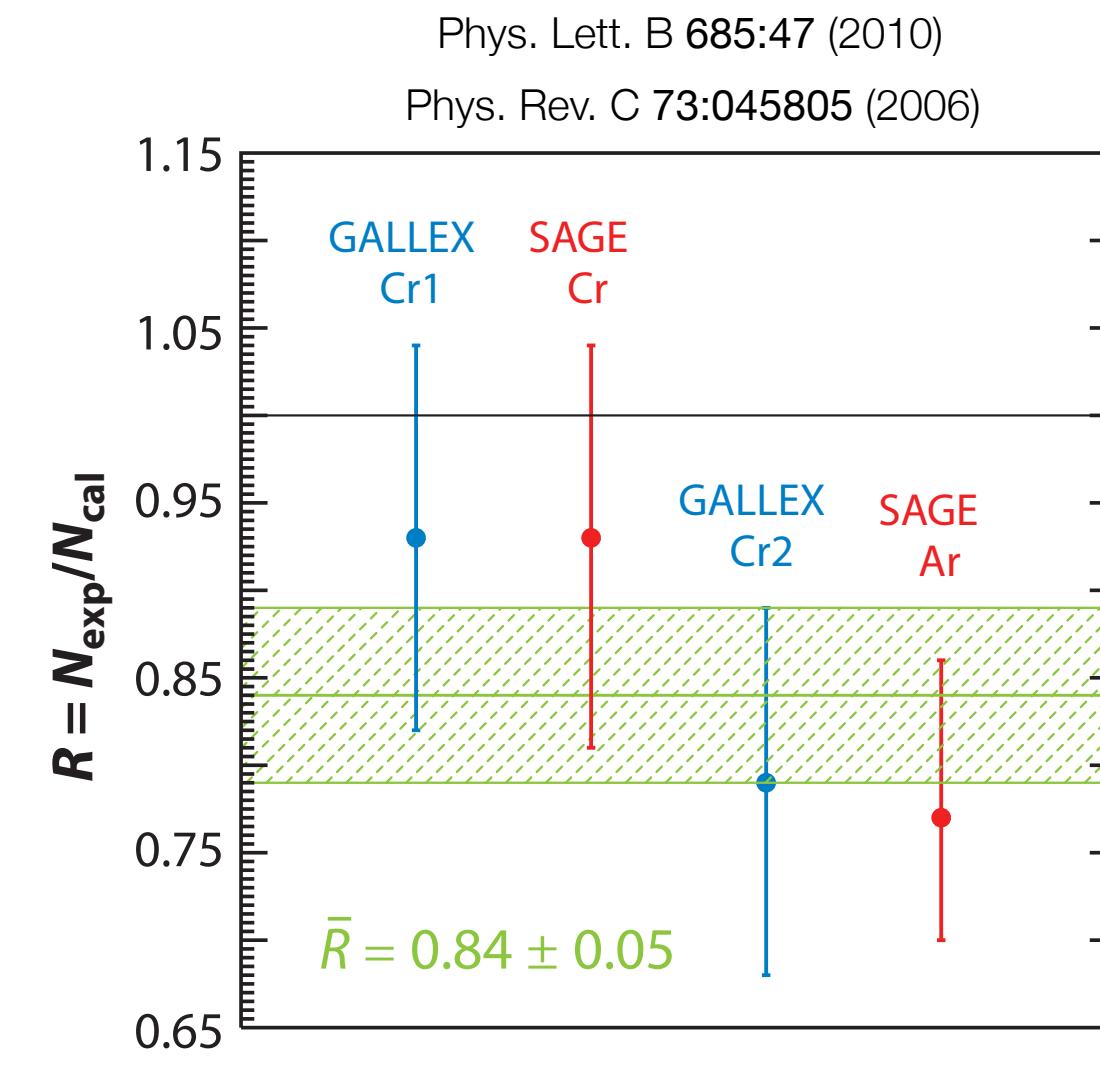
# Not the Only Anomaly

- **Gallium anomaly - disappearance of  $\nu_e$**  measured with radioactive sources in the solar neutrinos gallium experiments GALLEX and SAGE (rate only)
- **LSND/MiniBooNE anomaly** - energy-dependent event **excess in  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  channel** measured in LSND, consistent with an active-sterile oscillation with  $\Delta m^2 \gtrsim 0.1 \text{ eV}^2$ ; a similar excess was later seen by MiniBooNE
- All these anomalies can be explained by the existence of a light sterile neutrino, but a global simple solution combining them all is not possible
- **LSND/MiniBooNE anomaly** ( $\nu_\mu \rightarrow \nu_e$ ) is **highly disfavoured** by disappearance ( $\nu_\mu \rightarrow \nu_\mu$ ) results, while the **Reactor/Gallium anomalies** remain yet **untested**

$$P_{\nu_e \rightarrow \nu_e} \simeq 1 - 2|U_{e4}|^2(1 - |U_{e4}|^2)$$

$$P_{\nu_\mu \rightarrow \nu_\mu} \simeq 1 - 2|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2)$$

$$P_{\nu_\mu \rightarrow \nu_e} \simeq 2|U_{e4}|^2|U_{\mu 4}|^2$$



# Antineutrino Spectrum Estimation

- In low-enriched-uranium (LEU) facilities four isotopes contribute to neutrino spectrum ( $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{238}\text{U}$ ,  $^{241}\text{Pu}$ ), their fraction  $\alpha_k$  evolves with time (burnup)

$$N_{IBD}(E_{\bar{\nu}_e}, t) = \frac{N_p \epsilon}{4\pi L^2} \times \frac{P_{th}(t)}{\langle E_f \rangle(t)} \times \langle \sigma_f \rangle(E_{\bar{\nu}_e}, t)$$

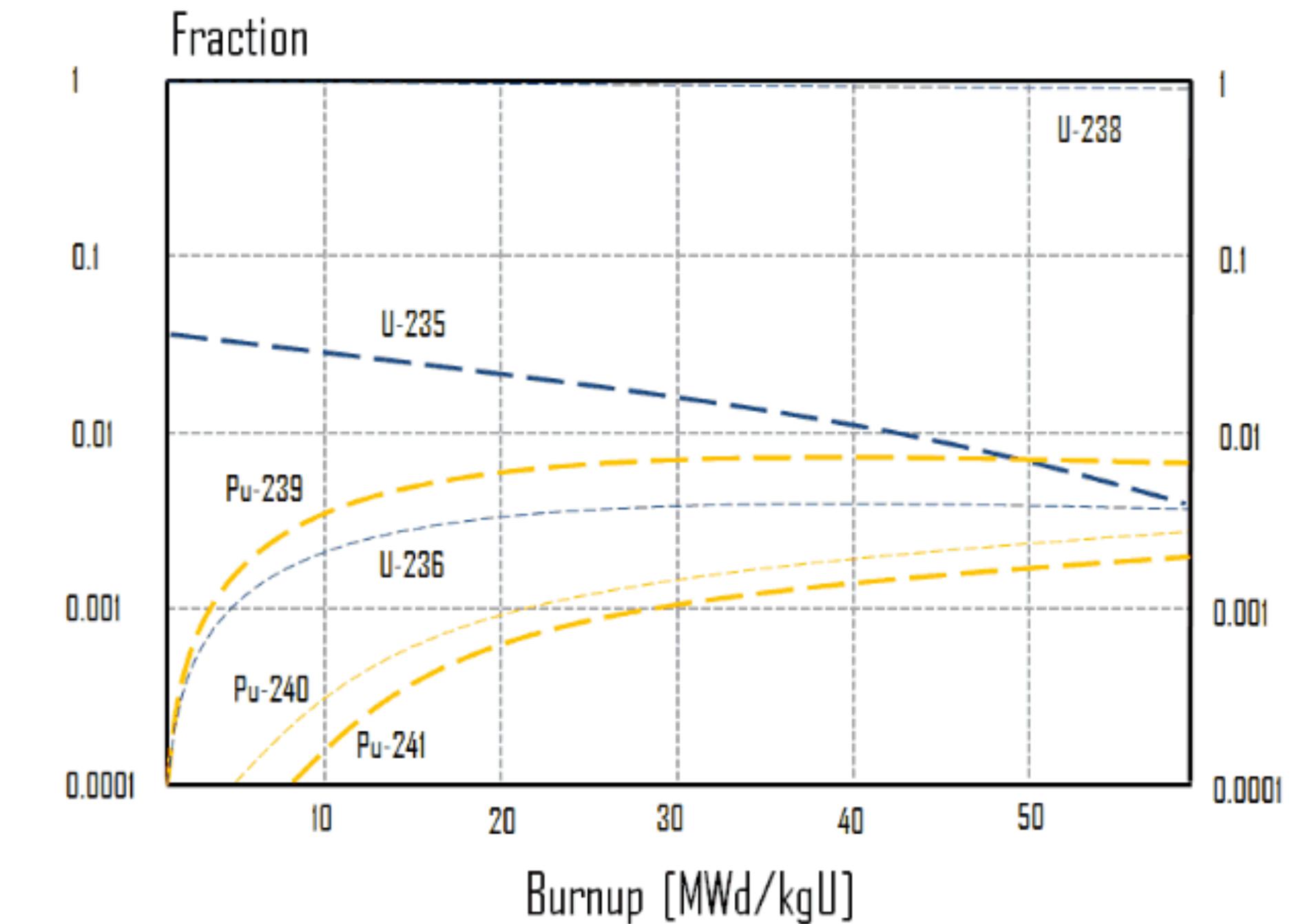
reactor thermal power  
average energy released per fission

$$\langle E_f \rangle = \sum_k \alpha_k(t) \langle E_f \rangle_k$$

average IBD cross-section per fission

$$\langle \sigma_f \rangle_k = \int S_k(E) \sigma_{IBD}(E) dE$$

- IBD cross-section from theoretical calculations
- Single  $\bar{\nu}$  spectra  $S_k(E)$  unavailable, obtained from global  $\beta$  spectrum ( $\mathcal{O}10^3$  branches)
  - Start with known branches from nuclear data tables...
  - ... and complement with effective decay branches



# Limits of Current Neutrino Spectrum Models

- Converted spectra method (used for the  $^{235}\text{U}$  and Pu contribution)
- Large uncertainty for the weak magnetism term
- Underestimated impact on uncertainties of the selection of average effective Z distributions used in the fit of the ILL spectra (up to 5%)
- Treatment of forbidden decays could change both normalisation and spectral shape measurement of the shape factors for the most important forbidden decays is crucial

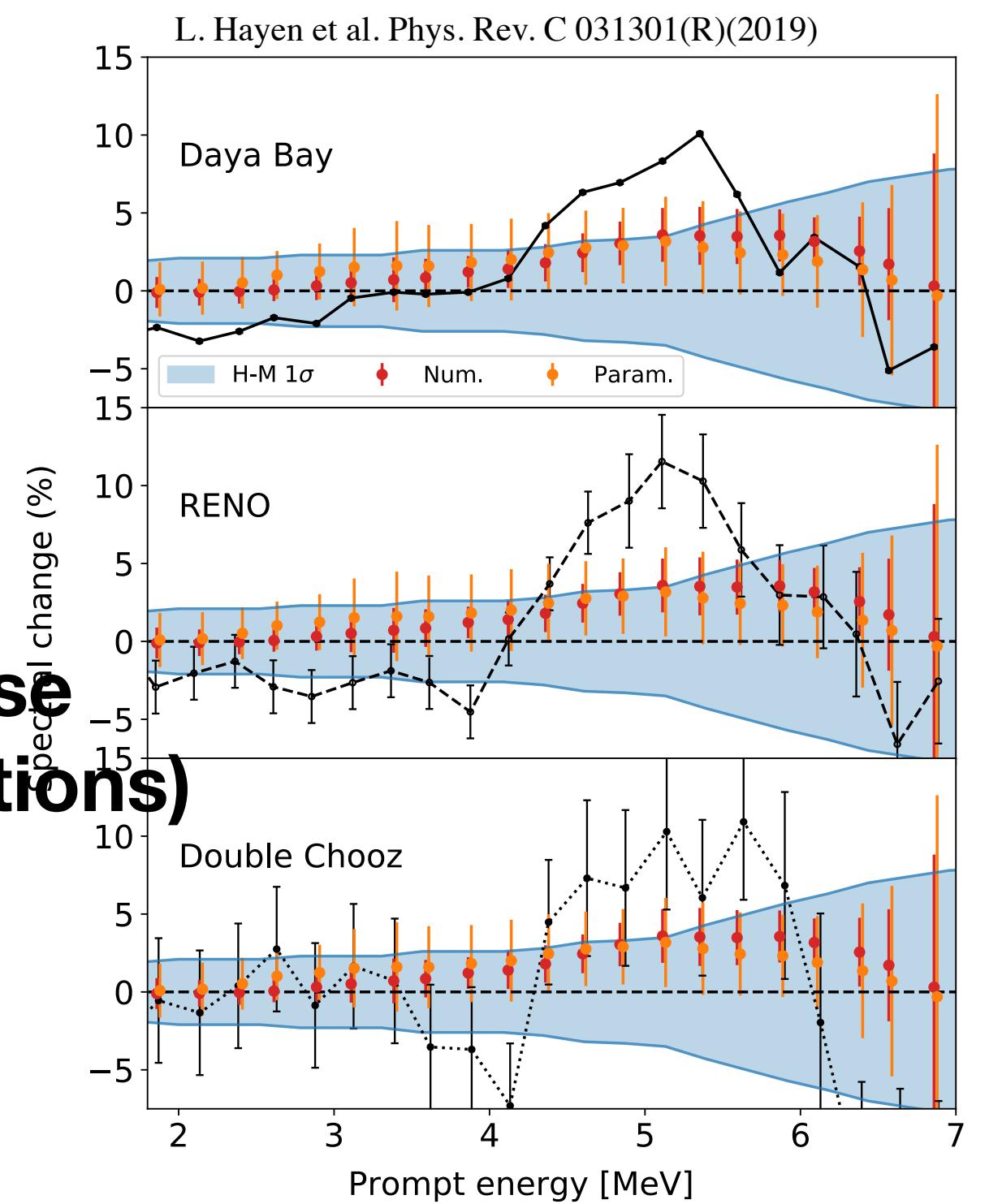
P. Huber PRC84,024617(2011)  
D.-L. Fang and B. A. Brown, Phys. Rev. C 91, 025503 (2015)

- Summation method (used for  $^{238}\text{U}$ )
- Incomplete or biased nuclear decay schemes
- Pandemonium effect, which can be solved by total absorption  $\gamma$  spectroscopy measurements (data-model discrepancy reduced to < 2%)

J.C.Hardy et al., Phys. Lett. B, 71, 307 (1977)

- To solve the RAA, we must tackle the problem from both experimental (increase statistics, detector upgrades) and theoretical side (new models, better corrections)**

A . Hayes et al. Phys. Rev. Lett. 112, 202501 (2014)  
D.-L. Fang and B. A. Brown, Phys. Rev. C 91, 025503 (2015)  
X.B. Wang, J. L. Friar and A. C. Hayes Phys. Rev. C 95 (2017) 064313  
L. Hayen et al. Phys. Rev. C 031301(R)(2019) and PRC.100.054323

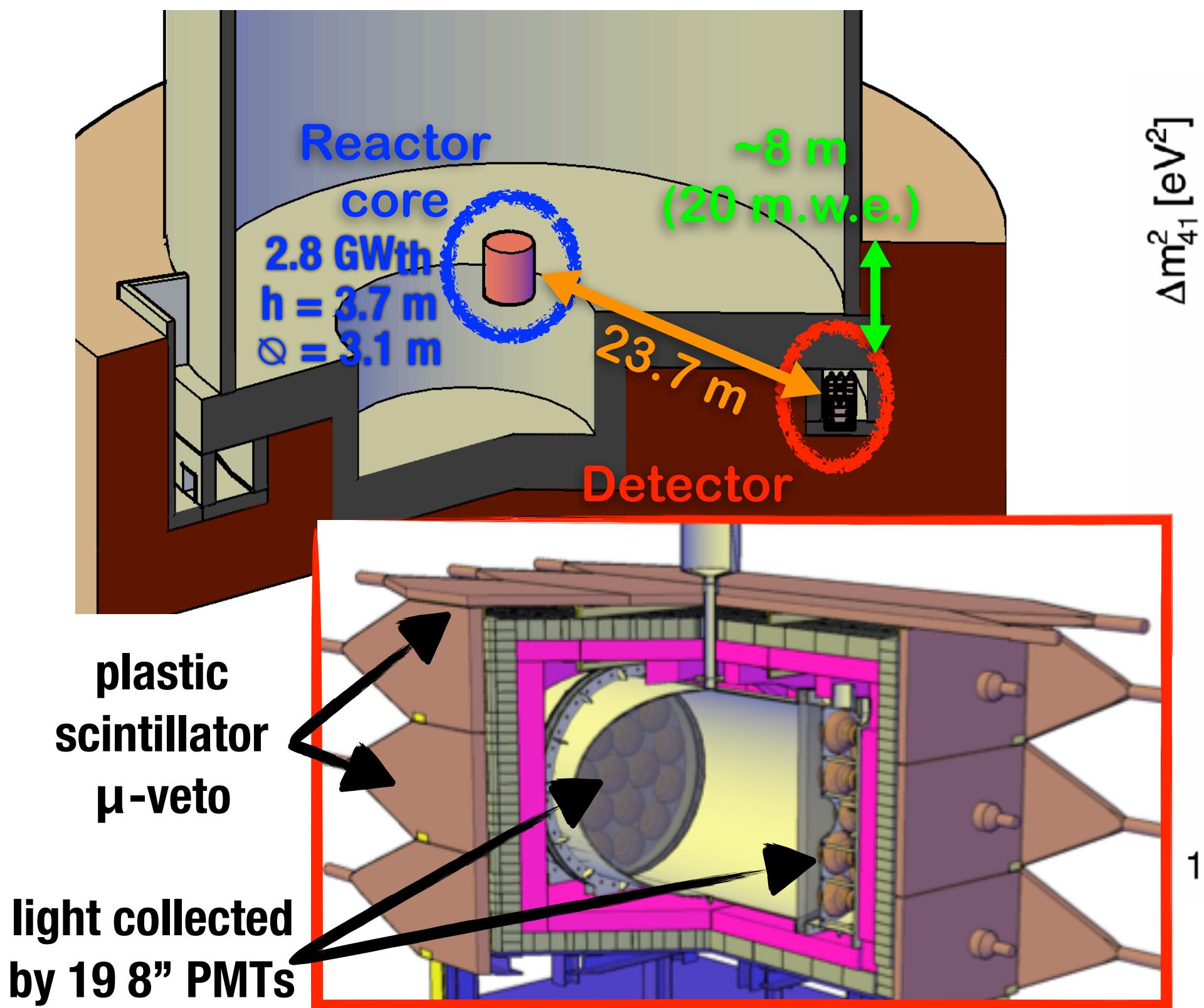


# A World-Wide Hunt

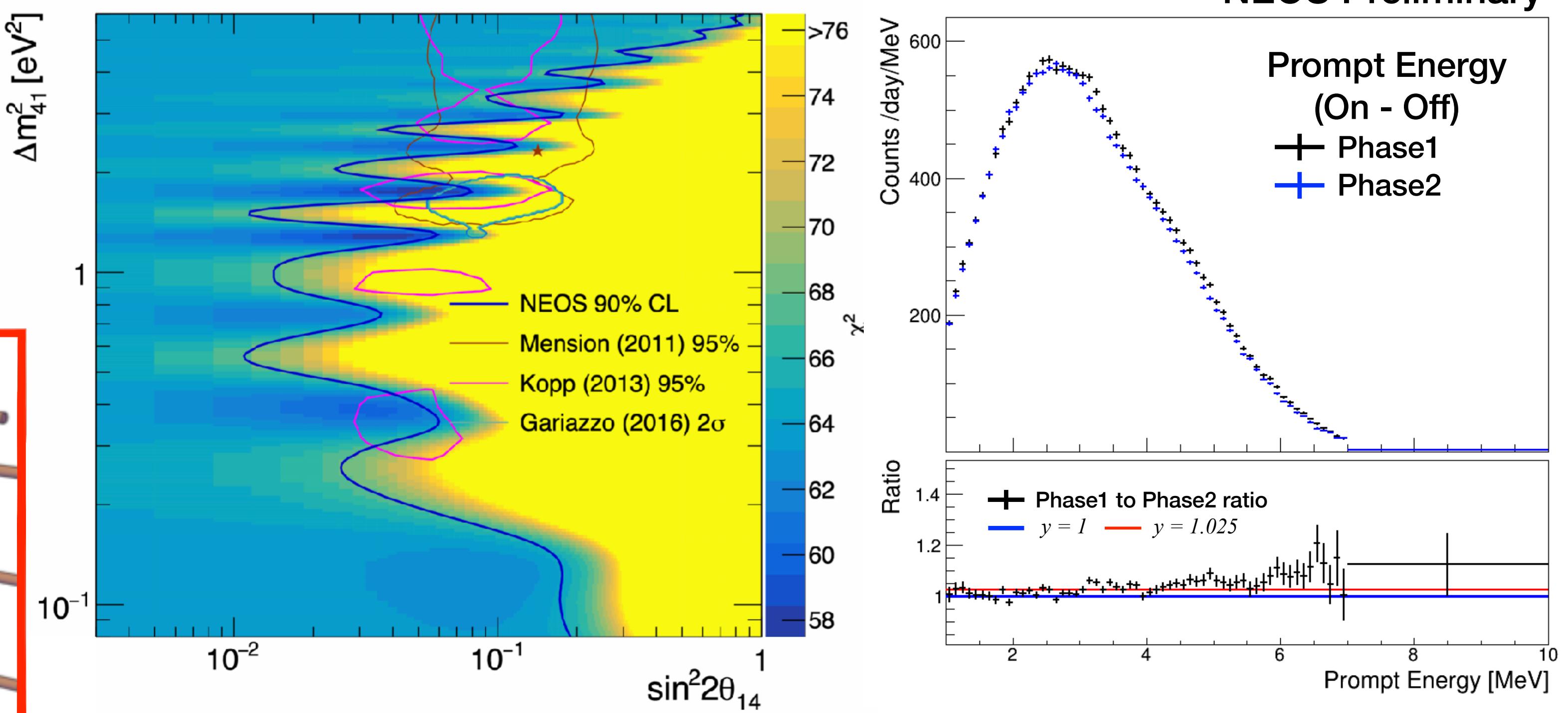
	<b>Core P<sub>Th</sub></b>	<b>Core Size</b>	<b>Overburden</b>	<b>Segmentation</b>	<b>Baseline</b>	<b>Material</b>
<b>Chandler</b>	72 MW ( <sup>235</sup> U)	$Q = 50 \text{ cm}$	$\sim 10 \text{ mwe}$	6.2 cm (3D)	5.5 m	PS + Li layer
<b>DANSS</b>	3 GW (LEU)	$h = 3.6 \text{ m}$ $Q = 3.1 \text{ m}$	$\sim 50 \text{ mwe}$	5 cm (2D)	10.7-12.7 m	Gd-doped PS
<b>NEOS</b>	2.8 GW (LEU)	$h = 3.7 \text{ m}$ $Q = 3.1 \text{ m}$	$\sim 20 \text{ mwe}$	-	23.7 m	Gd-doped LS
<b>Neutrino4</b>	90 MW ( <sup>235</sup> U)	$35 \times 42 \times 42 \text{ cm}^3$	few mwe	22.5 cm (2D)	6-12 m	Gd-doped LS
<b>NuLat</b>	40/1790 MW ( <sup>235</sup> U/LEU)		few mwe	6.35 cm (3D)	4.7/24 m	Li-doped PS
<b>Prospect</b>	85 MW ( <sup>235</sup> U)	$h = 0.5 \text{ m}$ $Q = 0.2 \text{ m}$	few mwe	15 cm (2D)	7 m	Li-doped LS
<b>SoLi<math>\delta</math></b>	72 MW ( <sup>235</sup> U)	$Q = 0.5 \text{ m}$	$\sim 10 \text{ mwe}$	5 cm (3D)	5.5 m	PS + Li layer
<b>Stereo</b>	58 MW ( <sup>235</sup> U)	$Q = 37 \text{ cm}$	$\sim 15 \text{ mwe}$	25 cm (1D)	8.8-11.2 m	Gd-doped LS

# The NEOS Experiment

- **Simple design:** 1008 L Gd-loaded (0.48%) liquid scintillator tank, spectrum compared with Data Bay
- **Very high statistics** (~2000 ν/day) thanks to the 2.8 GW commercial reactor
- **Degradation of light yield in time**

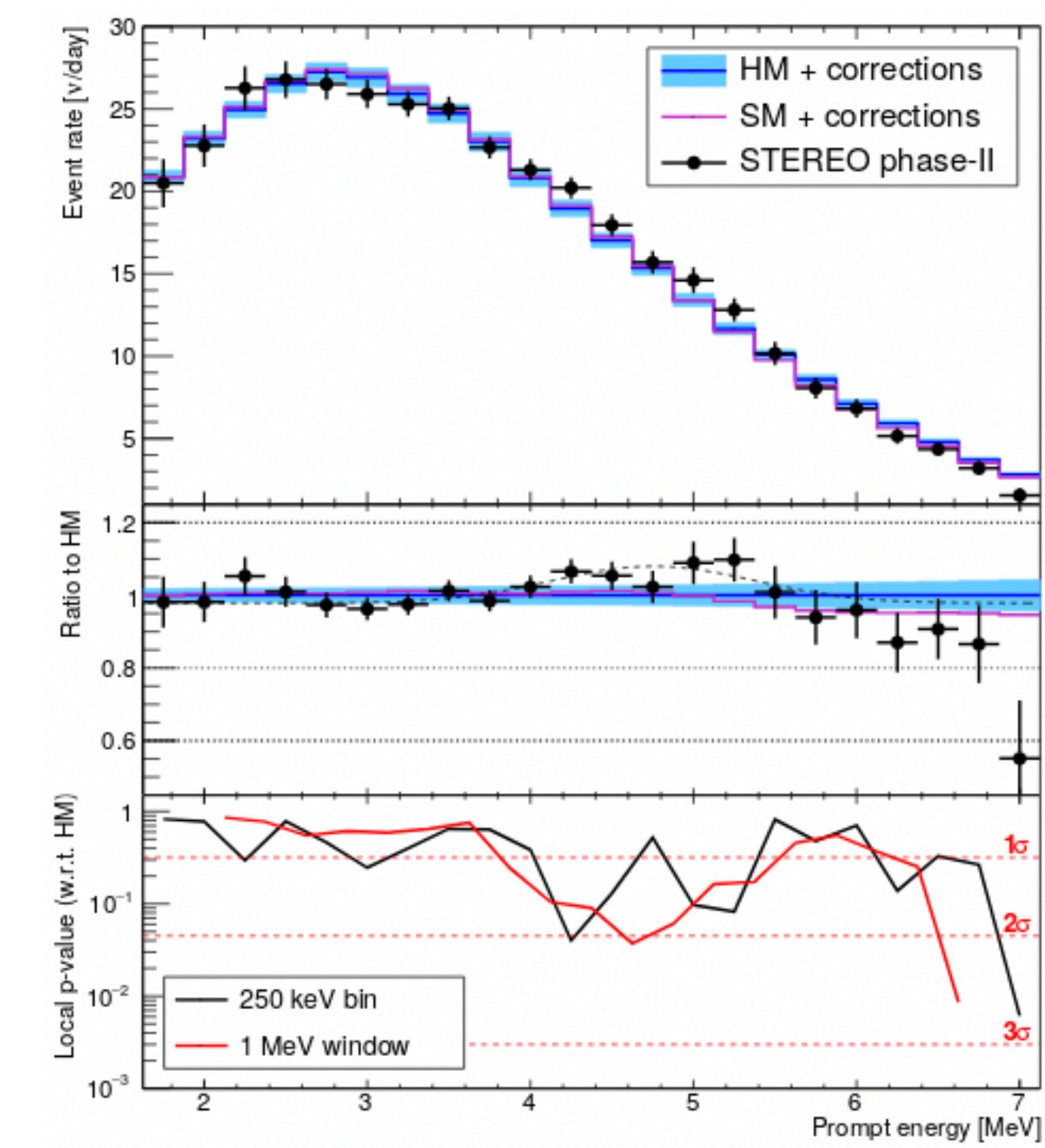
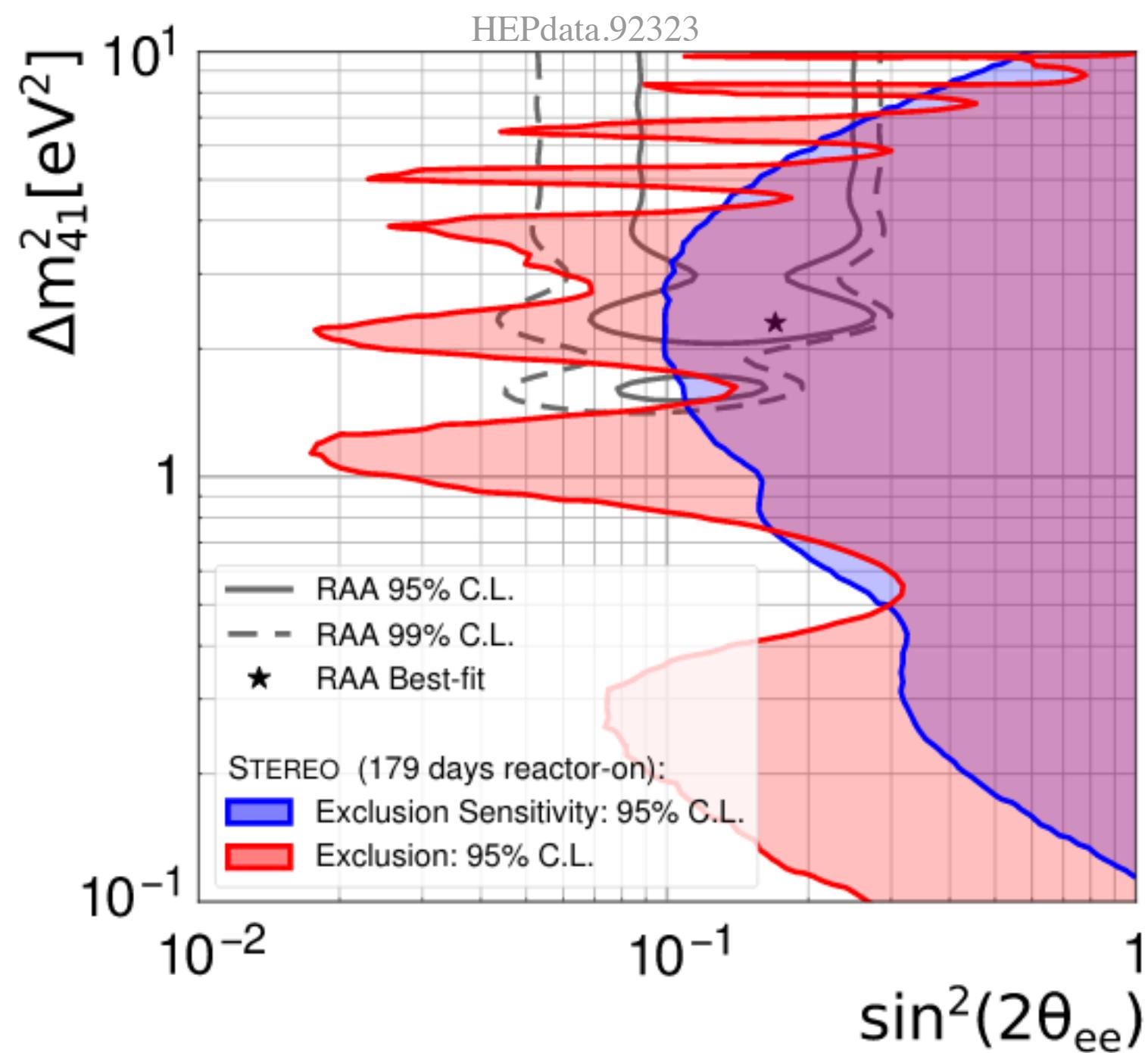
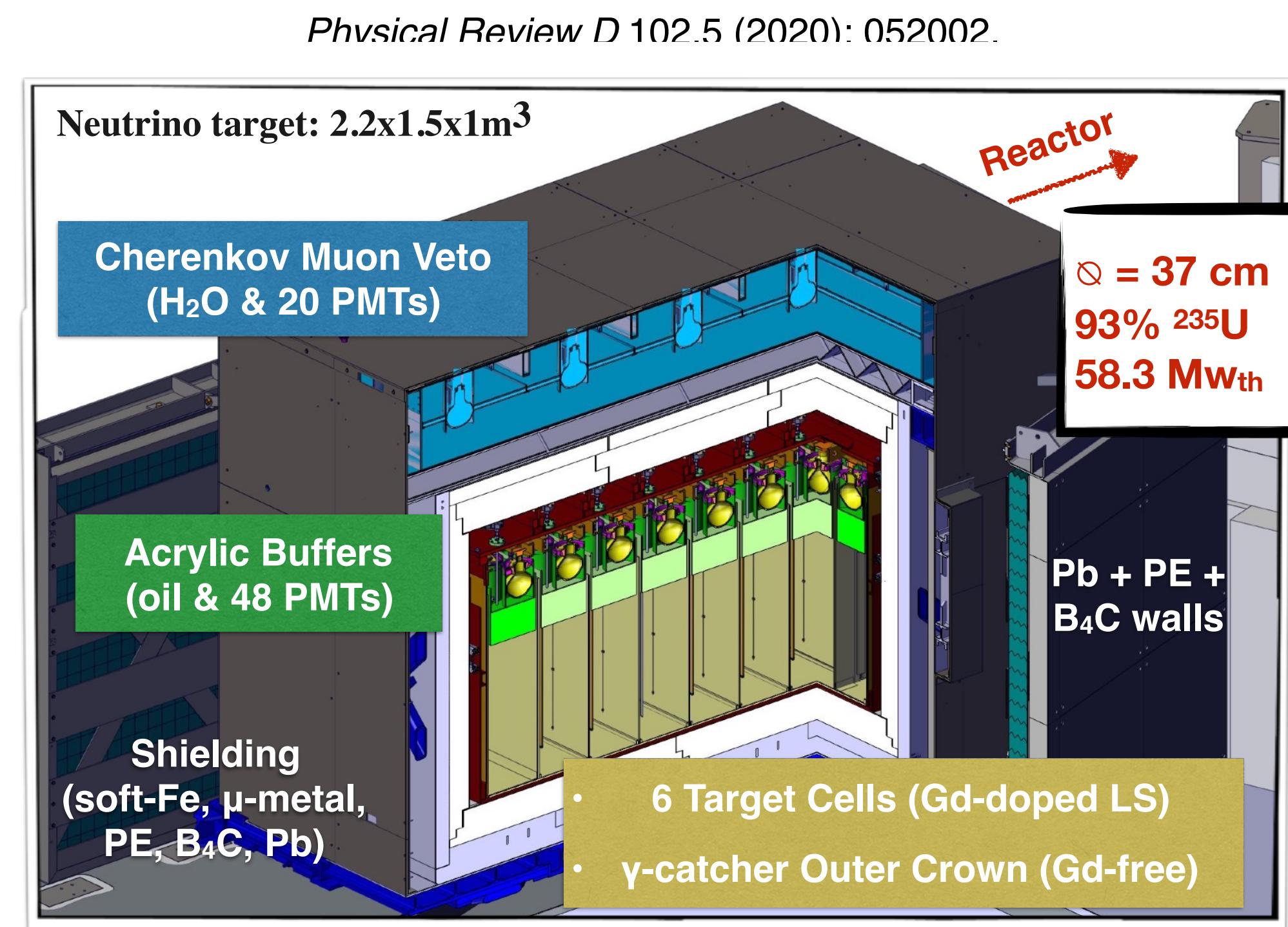


- Phase I (46 days OFF + 180 days ON)
- Oscillation analysis with RAA best fit excluded @90%CL
- Phase I+II
- Energy spectrum released
- Oscillation analysis ongoing (expected X2 sensitivity)



# The STEREO Experiment

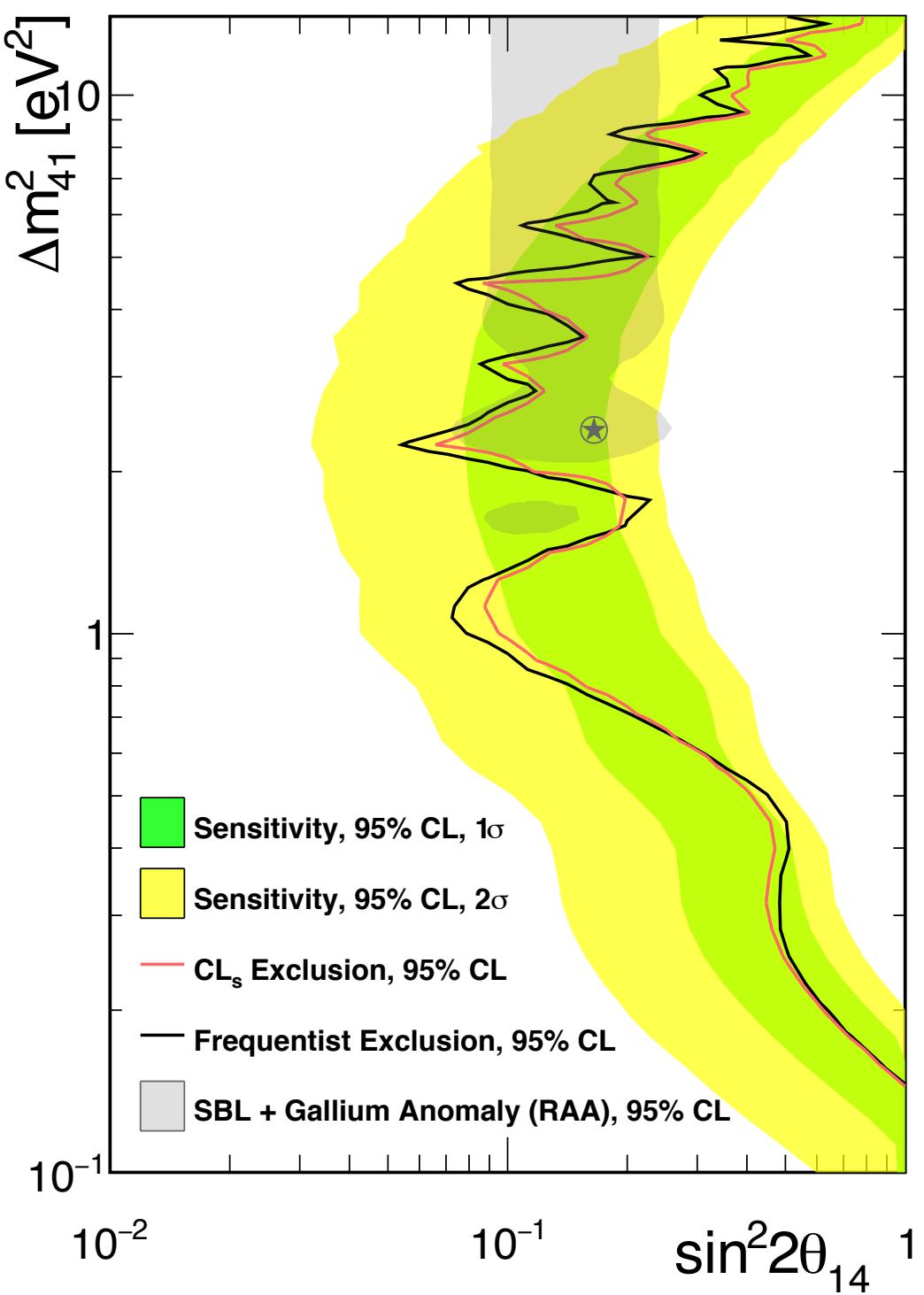
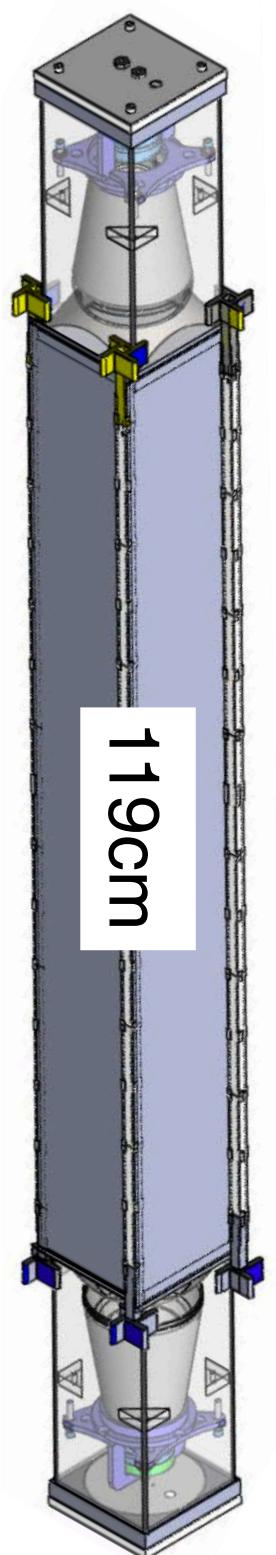
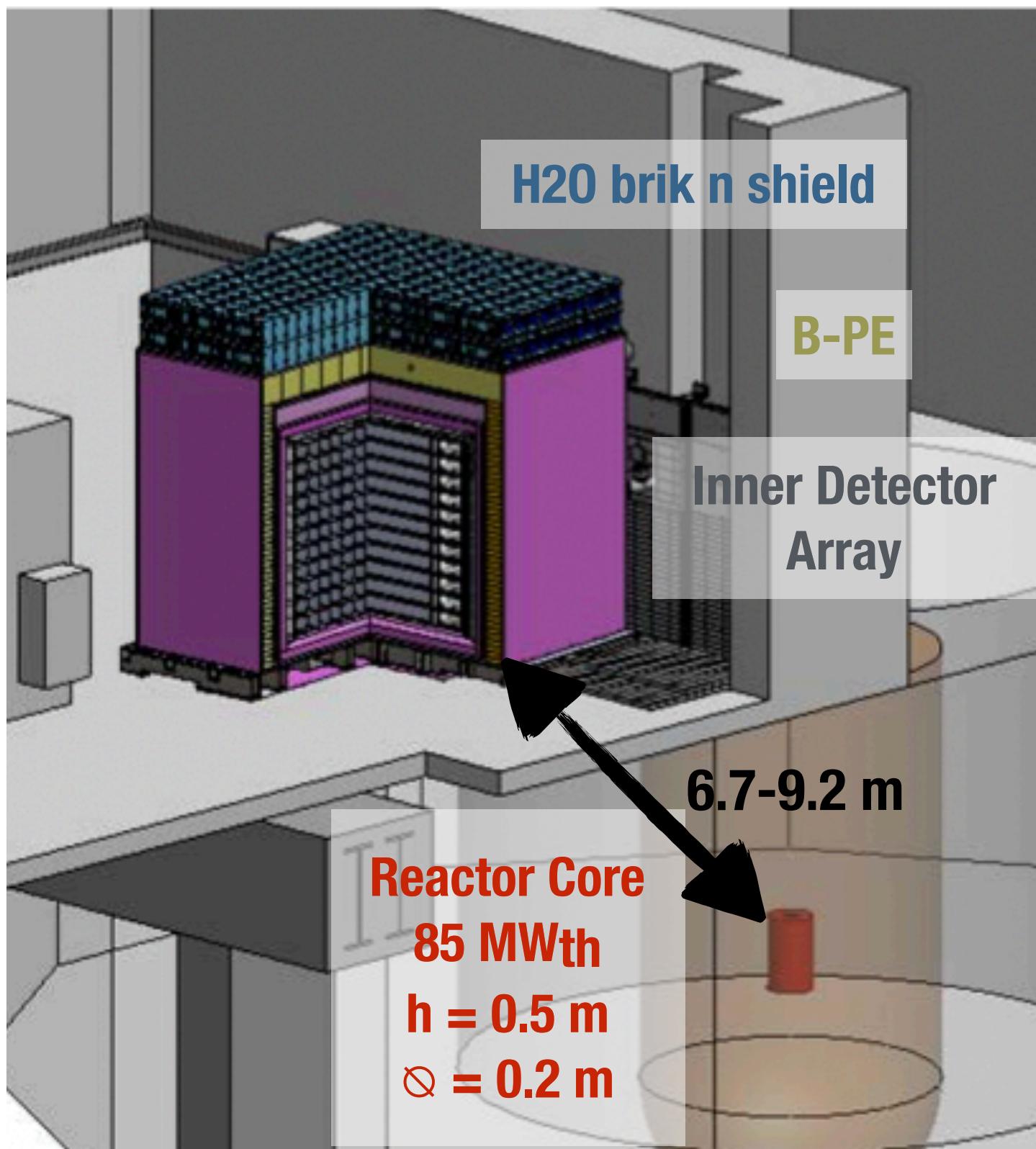
- Segmented design: 6 cells filled with Gd-loaded liquid scintillator → cell-to-cell relative oscillation analysis
- Compact HEU (58 MW) reactor core & short baseline (9-11 m from core) → **little damping of oscillation**
- **Little overburden and noise from reactor facility**
- Phase-I & -II combined data (65k IBDs, 179 days ON + 235 OFF) with S/B ~1 → **RAA best-fit rejected at > 99% CL**
- Expected X2 increase in sensitivity with full dataset
- **Absolute  $^{235}\text{U}$  rate and spectral shape** released using phase-II data (consistent with models)



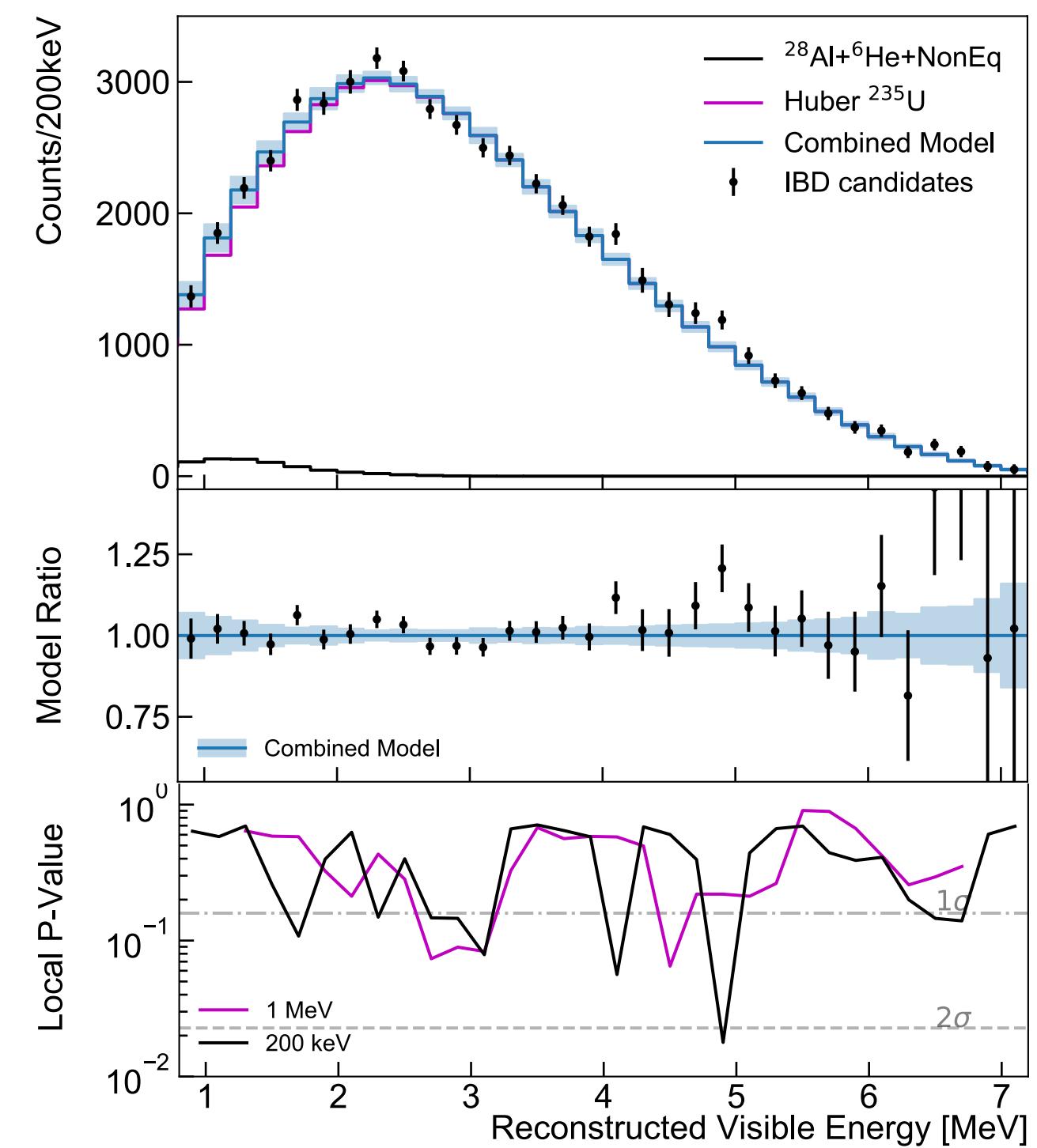
# The PROSPECT Experiment

- Highly segmented design: 4-ton  ${}^6\text{Li}$ -loaded liquid scintillator in 11x14 optically separated segments  
→ **good  $E_{\text{res}}$ , and 3D reconstruction**
- Relatively **high statistics** (530 IBD/day) and S/B ( $>1$ ) **for a HEU experiment**

PHYSICAL REVIEW D 103, 032001 (2021)

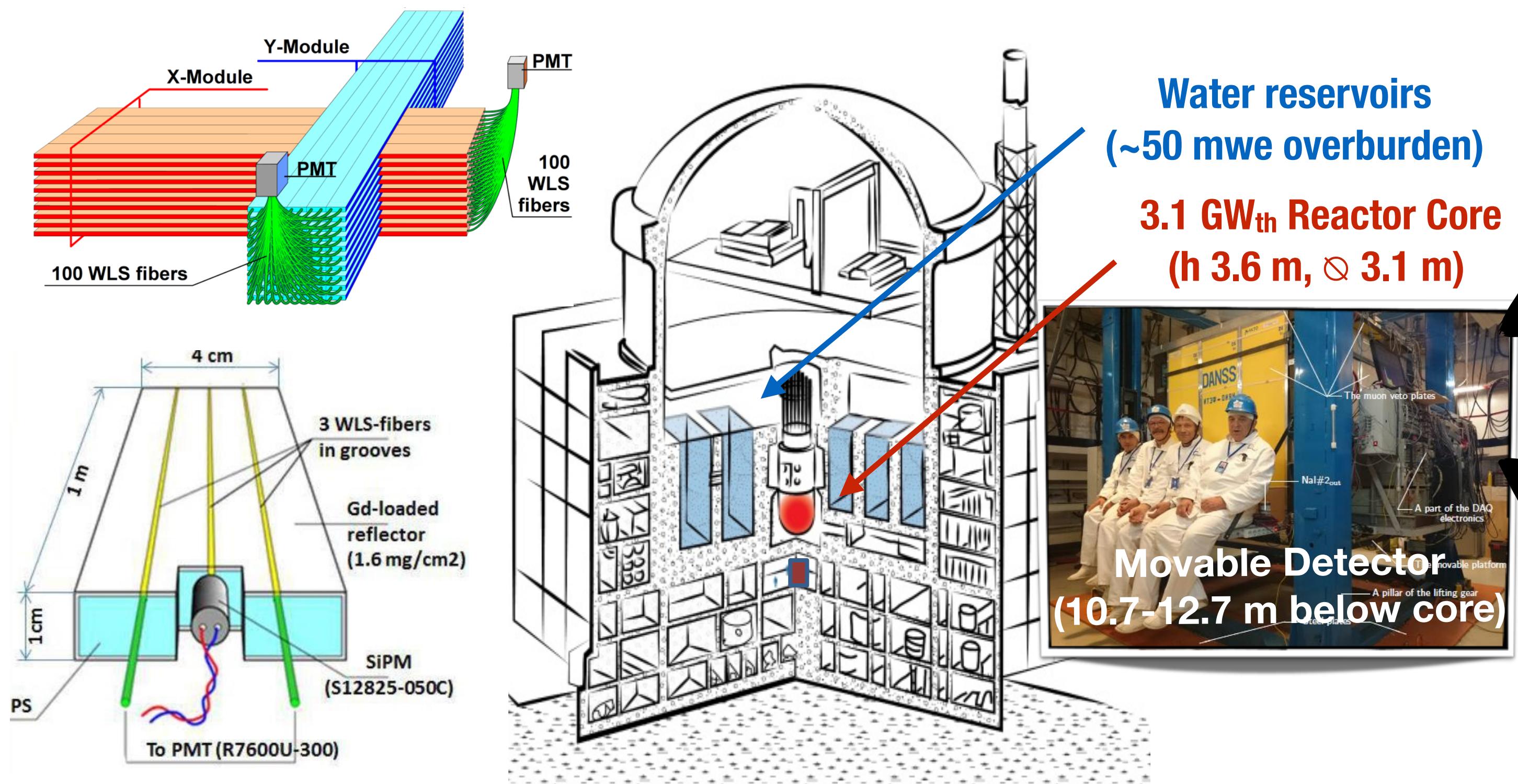


- Published results with 50k IBDs (105 days ON + 78 days OFF) → **RAA best-fit rejected at 98.5% CL**
- Pure  ${}^{235}\text{U}$  spectrum measured (consistent with models) and combined analyses with Data Bay and STEREO
- Results based on dataset from 2018**; improved analysis using dead cells (+50%) ongoing

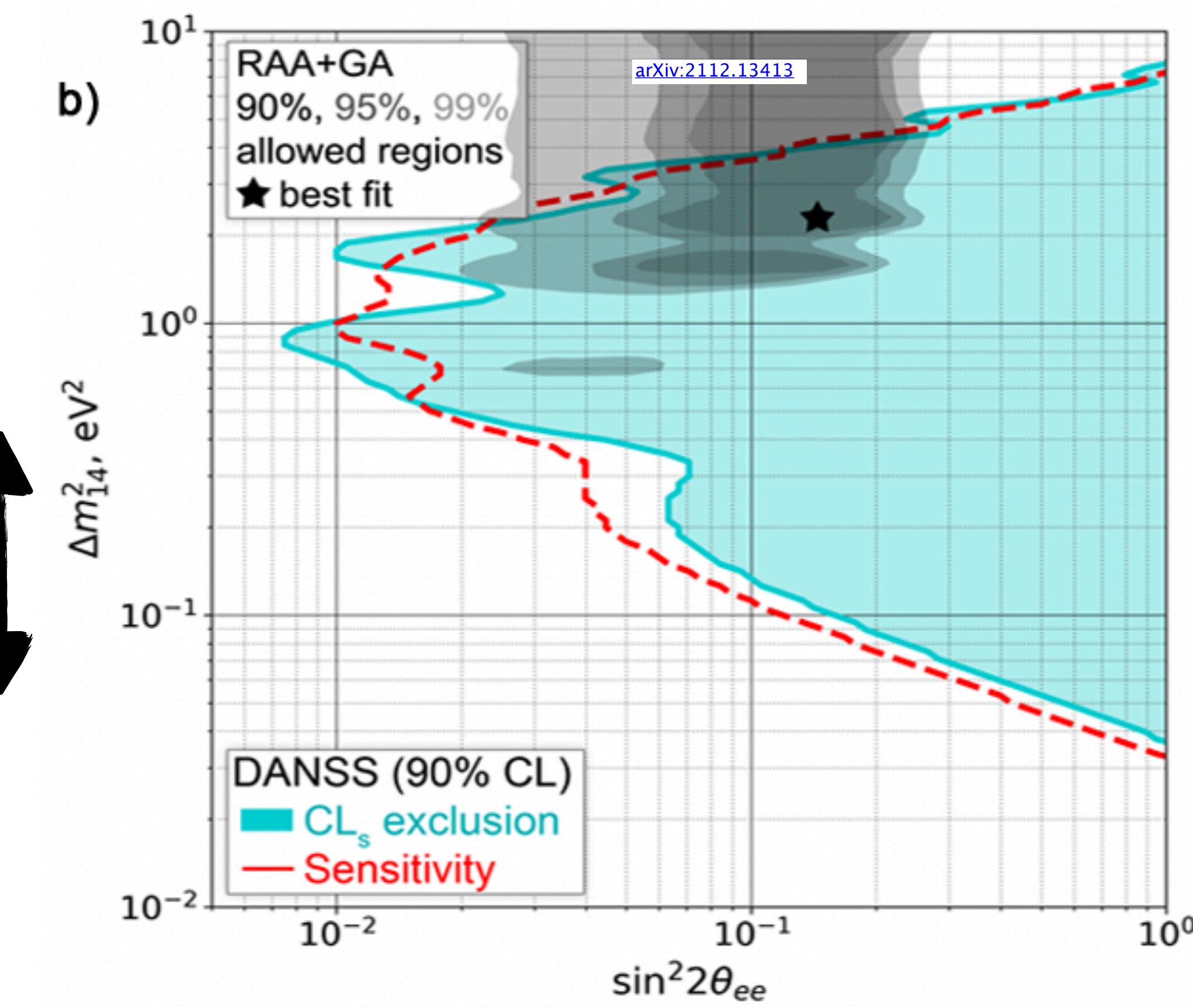


# The DANSS Experiment

- Highly-segmented  $\nu$  spectrometer: 2500 Gd-coated plastic scintillator strips in 50 modules  
→ **quasi-3D reconstruction**
- Excellent statistics** (~5000 IBD/day, S/B~60)
- Movable detector under the reactor (suppression of systematic + overburden)

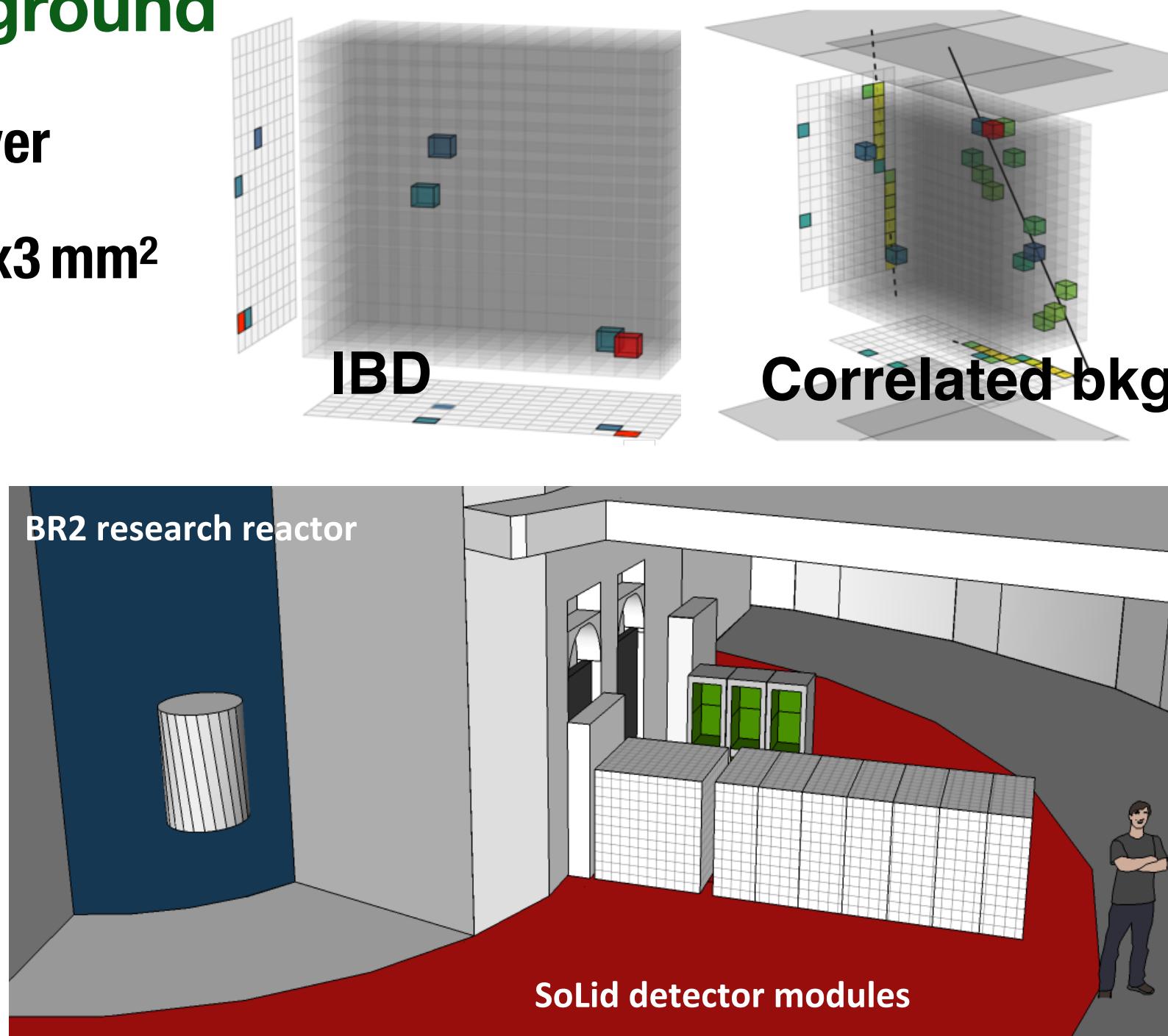
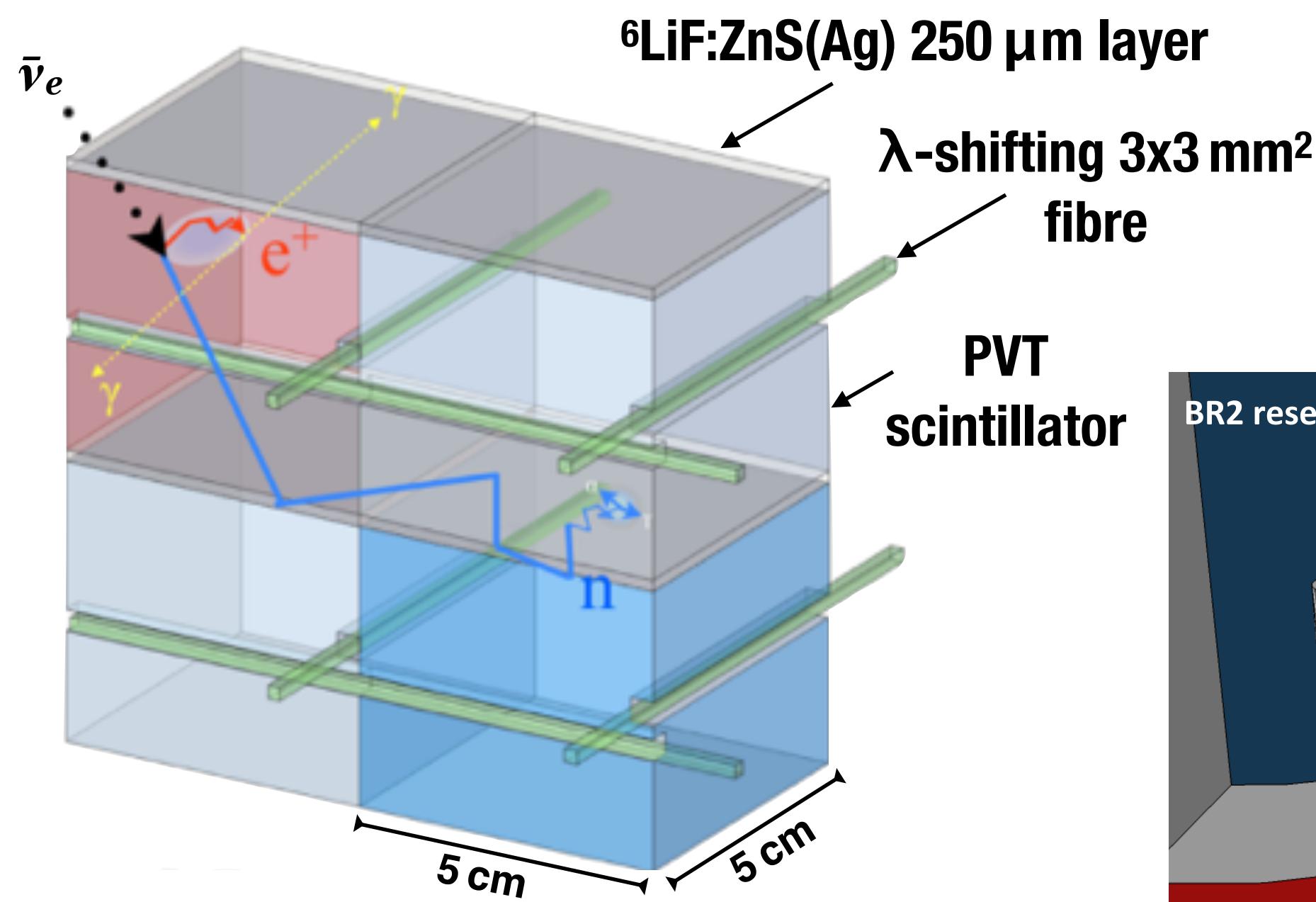


- Phase-I results (5.5 M IBD in 5 years) **exclude a large portion of the RAA** allowed region and its best fit
- Phase-II upgrade ongoing → goal is to halve energy resolution and increase detector volume (new strips)



# The SoLiD Experiment

- **Highly-segmented 3D detector design:** 12800  $5 \times 5 \times 5 \text{ cm}^3$  optically separated PVT cubes, with a  ${}^6\text{LiF:ZnS(Ag)}$  layer for neutron identification
- Relatively powerful HEU compact core & very short baseline (6-9 m from the core)
- Very little overburden, but **can use topology to separate IBDs from cosmic background**



- **Challenging background from BiPo coincidences** due to the internal  ${}^{238}\text{U}/{}^{230}\text{Th}$  series isotopes (mainly in  ${}^6\text{LiF:ZnS(Ag)}$ )
  - Analysis of phase-I data (326 days ON + 87 days OFF) ongoing
  - Detector upgrade for Phase-II with new SiPMs

