

# A novel Vlasov approach for modeling electron cloud instabilities

ECLOUD 22

Sofia Johannesson, Giovanni Iadarola, Beams Department CERN



## Introduction

## Simulation Model

Linear model of electron clouds

The Vlasov Equation

## Results with zero chromaticity

## Results with chromaticity



# Outline

Introduction

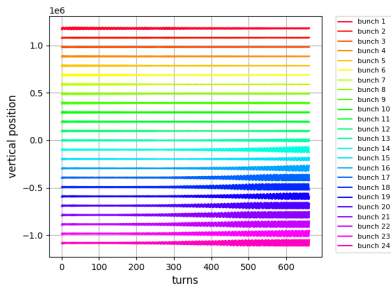
Simulation Model

Results with zero chromaticity

Results with chromaticity

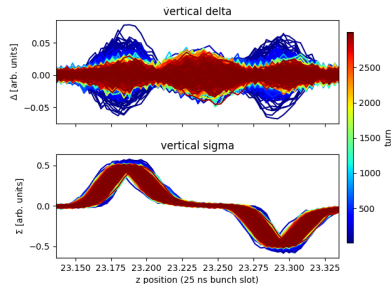
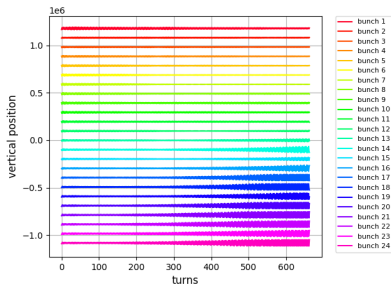


# Instabilities driven by e-cloud.



- Electron clouds can drive transverse instabilities

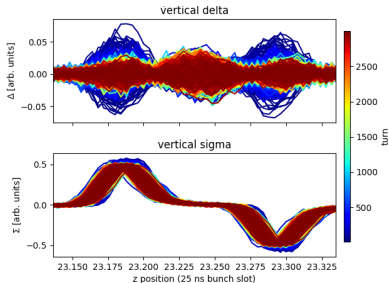
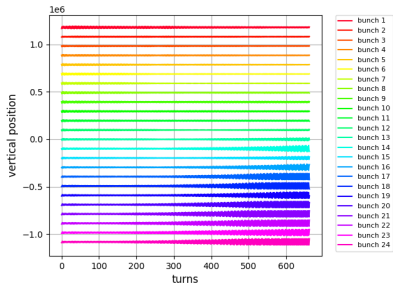
# Instabilities driven by e-cloud.



LHC measurements, June 2022

- Electron clouds can drive transverse instabilities, which cannot be mitigated by transverse feed-back system due to strong intra-bunch motion. [6] F. Zimmermann, 2004, *Review of Single bunch instabilities driven by electron cloud*

# Instabilities driven by e-cloud.



LHC measurements, June 2022

- Electron clouds can drive transverse instabilities, which cannot be mitigated by transverse feedback system due to strong intra-bunch motion. [6] F. Zimmermann, 2004, *Review of Single bunch instabilities driven by electron cloud*
- Conventional simulations using macroparticle tracking together with the PIC method for e-cloud beam interaction, are very computationally heavy. [7] G. Iadarola, et al., 2017, *Evolution of Python Tools for the Simulation of Electron Cloud Effects*

# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*

# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- The calculation time is independent of the instability growth rate.



# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- The calculation time is independent of the instability growth rate.  
⇒ Possible to study slow instabilities!

# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- The calculation time is independent of the instability growth rate.  
⇒ **Possible to study slow instabilities!**

Shifting from impedance forces to e-cloud forces in the Vlasov model requires:

# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- The calculation time is independent of the instability growth rate.  
⇒ **Possible to study slow instabilities!**

Shifting from impedance forces to e-cloud forces in the Vlasov model requires:

- A more general description of forces, since e-cloud dipolar forces cannot be modelled by conventional wakefields and impedances. [9] G. Rumolo and F. Zimmermann, 2002, *Electron Cloud simulations: beam instabilities and wakefields*

# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- The calculation time is independent of the instability growth rate.  
⇒ Possible to study slow instabilities!

Shifting from impedance forces to e-cloud forces in the Vlasov model requires:

- A more general description of forces, since e-cloud dipolar forces cannot be modelled by conventional wakefields and impedances. [9] G. Rumolo and F. Zimmermann, 2002, *Electron Cloud simulations: beam instabilities and wakefields*
- a betatron tune modulation along the longitudinal coordinate of the bunch as a result of e-cloud forces. [6] F. Zimmermann, 2004, *Review of Single bunch instabilities driven by electron cloud*

# Instabilities driven by electron cloud

- Instabilities driven by impedance effects have been modelled using the linearised Vlasov Equation, which identifies the Instability growth rate and betatron frequency shift for each instability mode. [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- The calculation time is independent of the instability growth rate.  
⇒ Possible to study slow instabilities!

Shifting from impedance forces to e-cloud forces in the Vlasov model requires:

- A more general description of forces, since e-cloud dipolar forces cannot be modelled by conventional wakefields and impedances. [9] G. Rumolo and F. Zimmermann, 2002, *Electron Cloud simulations: beam instabilities and wakefields*
- a betatron tune modulation along the longitudinal coordinate of the bunch as a result of e-cloud forces. [6] F. Zimmermann, 2004, *Review of Single bunch instabilities driven by electron cloud*

Previous attempts of using the Vlasov method to model e-cloud driven instabilities have not included these points together. [10] K. Ohmi et al, 2001, *Wake-Field and Fast Head-Tail Instability Caused by an Electron Cloud.*, [11] E. Perevedentsev, 2002, *Head-Tail Instability Caused by Electron Cloud*

# Outline

Introduction

## Simulation Model

Linear model of electron clouds

The Vlasov Equation

Results with zero chromaticity

Results with chromaticity



# General description of e-cloud forces

Begin by describing the dipolar e-cloud forces:

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*



# General description of e-cloud forces

Begin by describing the dipolar e-cloud forces:

Choose a set of **sinusoid beam distortions**,  $h_n(z)$  where  $z$  is the position along the bunch. The sinusoid test functions satisfy the **orthogonality condition**:  $\int h_n(z) h_{n'}(z) = H_n^2 \delta_{n,n'}$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*





# General description of e-cloud forces

Begin by describing the dipolar e-cloud forces:

Choose a set of **sinusoid beam distortions**,  $h_n(z)$  where  $z$  is the position along the bunch. The sinusoid test functions satisfy the **orthogonality condition**:  $\int h_n(z)h_{n'}(z) = H_n^2\delta_{n,n'}$

Each distortion,  $h_n$ , corresponds to a response function  $k_n$  calculated from the interaction with e-cloud using **single-pass PIC simulations**.

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

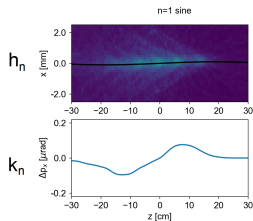


# General description of e-cloud forces

Begin by describing the dipolar e-cloud forces:

Choose a set of **sinusoid beam distortions**,  $h_n(z)$  where  $z$  is the position along the bunch. The sinusoid test functions satisfy the **orthogonality condition**:  $\int h_n(z)h_{n'}(z) = H_n^2\delta_{n,n'}$

Each distortion,  $h_n$ , corresponds to a response function  $k_n$  calculated from the interaction with e-cloud using **single-pass PIC simulations**.



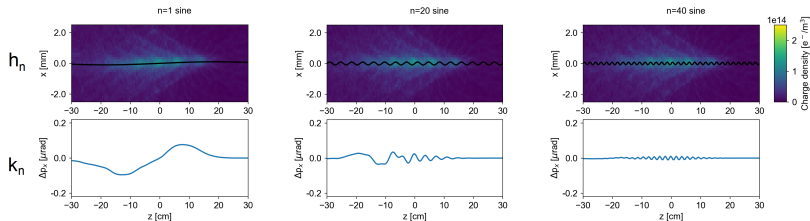
[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# General description of e-cloud forces

Begin by describing the dipolar e-cloud forces:

Choose a set of **sinusoid beam distortions**,  $h_n(z)$  where  $z$  is the position along the bunch. The sinusoid test functions satisfy the **orthogonality condition**:  $\int h_n(z)h_{n'}(z) = H_n^2\delta_{n,n'}$

Each distortion,  $h_n$ , corresponds to a response function  $k_n$  calculated from the interaction with e-cloud using **single-pass PIC simulations**.



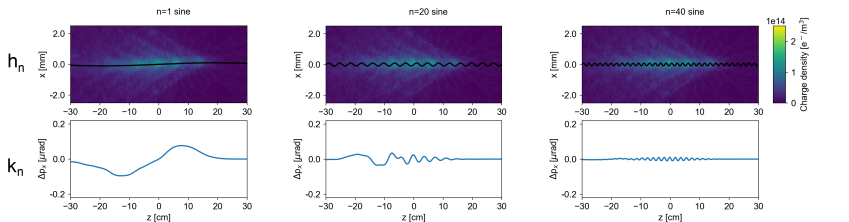
[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# General description of e-cloud forces

Begin by describing the dipolar e-cloud forces:

Choose a set of **sinusoid beam distortions**,  $h_n(z)$  where  $z$  is the position along the bunch. The sinusoid test functions satisfy the **orthogonality condition**:  $\int h_n(z)h_{n'}(z) = H_n^2\delta_{n,n'}$

Each distortion,  $h_n$ , corresponds to a response function  $k_n$  calculated from the interaction with e-cloud using **single-pass PIC simulations**.



These calculations use the **e-cloud in the superconducting quadrupoles** of the LHC for a beam energy of 450GeV.

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Linear model of e-cloud - Dipolar Forces

Describe the **transverse centroid** along the bunch,  $\bar{x}(z)$ , as a **linear combination** of test functions  $h_n$ :

$$\bar{x}(z) = \sum_{n=0}^{\infty} a_n h_n(z); \quad a_n = \frac{1}{H_n^2} \int \bar{x}(z) h_n(z) dz \quad (1)$$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

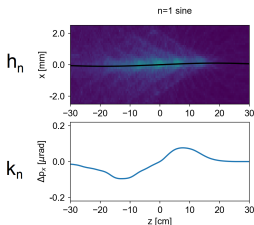


# Linear model of e-cloud - Dipolar Forces

Describe the **transverse centroid** along the bunch,  $\bar{x}(z)$ , as a **linear combination** of test functions  $h_n$ :

$$\bar{x}(z) = \sum_{n=0}^{\infty} a_n h_n(z); \quad a_n = \frac{1}{H_n^2} \int \bar{x}(z) h_n(z) dz \quad (1)$$

$k_n$  is the **resulting electron cloud kick** from a bunch distortion  $h_n$ .



[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Linear model of e-cloud - Dipolar Forces

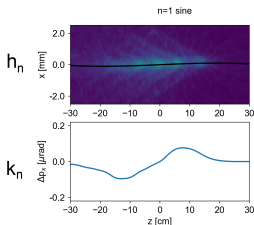
Describe the **transverse centroid** along the bunch,  $\bar{x}(z)$ , as a **linear combination** of test functions  $h_n$ :

$$\bar{x}(z) = \sum_{n=0}^{\infty} a_n h_n(z); \quad a_n = \frac{1}{H_n^2} \int \bar{x}(z) h_n(z) dz \quad (1)$$

$k_n$  is the **resulting electron cloud kick** from a bunch distortion  $h_n$ .

From simulation we verify **linear behaviour** such that the kick,  $\Delta x'$ , of arbitrary distribution  $\bar{x}(z)$  is:

$$\Delta x'(z) = \sum_{n=0}^{\infty} a_n k_n(z) \quad (2)$$



[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Linear model of e-cloud - Dipolar Forces

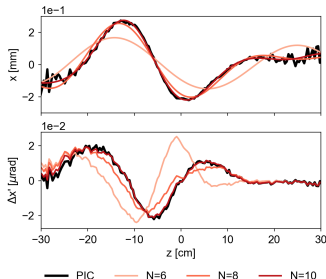
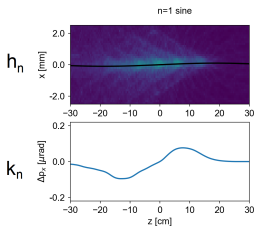
Describe the **transverse centroid** along the bunch,  $\bar{x}(z)$ , as a **linear combination** of test functions  $h_n$ :

$$\bar{x}(z) = \sum_{n=0}^{\infty} a_n h_n(z); \quad a_n = \frac{1}{H_n^2} \int \bar{x}(z) h_n(z) dz \quad (1)$$

$k_n$  is the **resulting electron cloud kick** from a bunch distortion  $h_n$ .

From simulation we verify **linear behaviour** such that the kick,  $\Delta x'$ , of arbitrary distribution  $\bar{x}(z)$  is:

$$\Delta x'(z) = \sum_{n=0}^{\infty} a_n k_n(z) \quad (2)$$



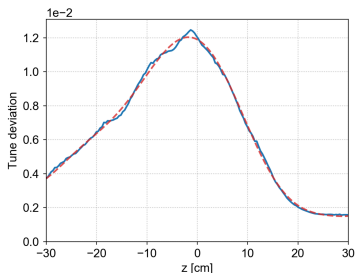
[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*



# Linear model of electron clouds - Quadrupolar forces

Model detuning using a polynomial

$$\Delta Q(z) = \sum_{n=0}^{N_p} A_n z^n \quad (3)$$



- [2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

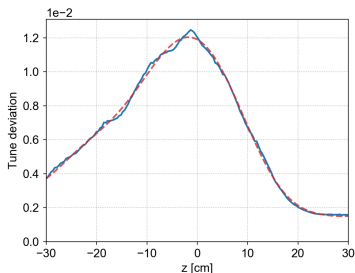
# Linear model of electron clouds - Quadrupolar forces

Model detuning using a polynomial

$$\Delta Q(z) = \sum_{n=0}^{N_p} A_n z^n \quad (3)$$

Generalize by adding chromaticity

$$\Delta Q(z, \delta) = \sum_{n=0}^{N_p} A_n z^n + B_n \delta^n \quad (4)$$



[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Linear model of electron clouds - Quadrupolar forces

Model detuning using a polynomial

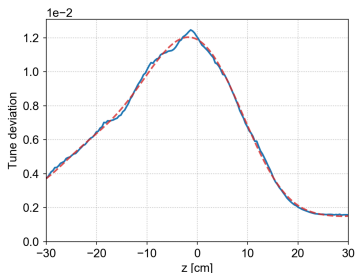
$$\Delta Q(z) = \sum_{n=0}^{N_p} A_n z^n \quad (3)$$

Generalize by adding chromaticity

$$\Delta Q(z, \delta) = \sum_{n=0}^{N_p} A_n z^n + B_n \delta^n \quad (4)$$

Including only linear chromaticity:

$$\Delta Q(z, \delta) = Q' \delta + \sum_{n=0}^{N_p} A_n z^n \quad (5)$$



[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# The Linerized Vlasov Equation - introduction

- $\psi_0$  is a distribution of particles where each individual particle obeys a Hamiltonian  $H_0$ .

[3] N. Mounet, 2018, *Direct Vlasov Solvers*



# The Linerized Vlasov Equation - introduction

- $\psi_0$  is a distribution of particles where each individual particle obeys a Hamiltonian  $H_0$ .
- The Vlasov equation describes the collective motion of the distribution  $\psi_0$

[3] N. Mounet, 2018, *Direct Vlasov Solvers*



# The Linerized Vlasov Equation - introduction

- $\psi_0$  is a distribution of particles where each individual particle obeys a Hamiltonian  $H_0$ .
- The Vlasov equation describes the collective motion of the distribution  $\psi_0$
- Introduce a perturbation,  $\Delta H$  and  $\Delta\psi$ , which means that the total Hamiltonian is  $H = H_0 + \Delta H$  and the total distribution is  $\psi = \psi_0 + \Delta\psi$

[3] N. Mounet, 2018, *Direct Vlasov Solvers*



# The Linearized Vlasov Equation - introduction

- $\psi_0$  is a distribution of particles where each individual particle obeys a Hamiltonian  $H_0$ .
- The Vlasov equation describes the collective motion of the distribution  $\psi_0$
- Introduce a perturbation,  $\Delta H$  and  $\Delta\psi$ , which means that the total Hamiltonian is  $H = H_0 + \Delta H$  and the total distribution is  $\psi = \psi_0 + \Delta\psi$
- This leads to the Linearized Vlasov Equation, which truncated to first order and expressed with Poisson brackets is:

$$\frac{\partial \Delta\psi}{\partial t} + [\Delta\psi, H_0] = -[\psi_0, \Delta H] \quad (6)$$

[3] N. Mounet, 2018, *Direct Vlasov Solvers*



# The Linearized Vlasov Equation - introduction

- $\psi_0$  is a distribution of particles where each individual particle obeys a Hamiltonian  $H_0$ .
- The Vlasov equation describes the collective motion of the distribution  $\psi_0$
- Introduce a perturbation,  $\Delta H$  and  $\Delta\psi$ , which means that the total Hamiltonian is  $H = H_0 + \Delta H$  and the total distribution is  $\psi = \psi_0 + \Delta\psi$
- This leads to the Linearized Vlasov Equation, which truncated to first order and expressed with Poisson brackets is:

$$\frac{\partial \Delta\psi}{\partial t} + [\Delta\psi, H_0] = -[\psi_0, \Delta H] \quad (6)$$

- The electron cloud forces are contained in  $\Delta H$
- The distortion  $\Delta\psi$  is the impact of the perturbation and the unknown

[3] N. Mounet, 2018, *Direct Vlasov Solvers*



# The Linearized Vlasov Equation

$$\frac{\partial \Delta \psi}{\partial t} + [\Delta \psi, H_0] = -[\psi_0, \Delta H]$$



# The Linearized Vlasov Equation

$$\frac{\partial \Delta \psi}{\partial t} + [\Delta \psi, H_0] = -[\psi_0, \Delta H]$$

$H_0$  from equations of motion



# The Linearized Vlasov Equation

$$\frac{\partial \Delta \psi}{\partial t} + [\Delta \psi, H_0] = -[\psi_0, \Delta H]$$

$H_0$  from equations of motion

$\Delta H = -\frac{x F_x^{coh}(z, t)}{m_0 \gamma v}$  from integrating Hamilton's equations

# The Linearized Vlasov Equation

$$\frac{\partial \Delta \psi}{\partial t} + [\Delta \psi, H_0] = -[\psi_0, \Delta H]$$

$H_0$  from equations of motion

switch to polar coordinates  
 $(x, x', z, \delta) \rightarrow (J_x, \theta_x, r, \phi)$

$\Delta H = -\frac{x F_x^{coh}(z, t)}{m_0 \gamma v}$  from integrating Hamilton's equations

# The Linearized Vlasov Equation

$$\frac{\partial \Delta \psi}{\partial t} + [\Delta \psi, H_0] = -[\psi_0, \Delta H]$$

$H_0$  from equations of motion

switch to polar coordinates  
 $(x, x', z, \delta) \rightarrow (J_x, \theta_x, r, \phi)$

$\Delta H = -\frac{x F_x^{coh}(z, t)}{m_0 \gamma v}$  from integrating Hamilton's equations

Factorize unperturbed bunch distribution,  $\psi_0 = \frac{\eta v}{\omega_0} f_0(J_x) g_0(r)$

# The Linearized Vlasov Equation

$$\frac{\partial \Delta \psi}{\partial t} + [\Delta \psi, H_0] = -[\psi_0, \Delta H]$$

$H_0$  from equations of motion

$\Delta H = -\frac{x F_x^{coh}(z, t)}{m_0 \gamma v}$  from integrating Hamilton's equations

switch to polar coordinates  
( $x, x', z, \delta$ )  $\rightarrow$  ( $J_x, \theta_x, r, \phi$ )

Factorize unperturbed  
bunch distribution,  $\psi_0 =$   
 $\frac{\eta v}{\omega_0} f_0(J_x) g_0(r)$

$$\frac{\partial \Delta \psi}{\partial t} - \omega_0 (Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta \psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta \psi}{\partial \phi} = -\frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$

reds are unknowns and greens come from e-cloud forces

# E-cloud in the Vlasov Equation - Dipolar Forces

The coherent force can be expressed using the responses  $k_n$  introduced earlier.



## E-cloud in the Vlasov Equation - Dipolar Forces

The coherent force can be expressed using the responses  $k_n$  introduced earlier.

Assuming the force is distributed uniformly in the accelerator:

$$F_x^{coh}(z, t) = \frac{m_0 \gamma v^2}{2\pi R} \Delta x' \quad (7)$$



# E-cloud in the Vlasov Equation - Dipolar Forces

The coherent force can be expressed using the responses  $k_n$  introduced earlier.

Assuming the force is distributed uniformly in the accelerator:

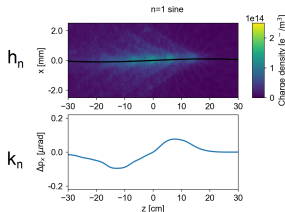
$$F_x^{coh}(z, t) = \frac{m_0 \gamma v^2}{2\pi R} \Delta x' \quad (7)$$

$\Delta x'$  is a linear combination of the responses  $k_n$  using the coefficients calculated from the projection of  $h_n$  on  $\bar{x}(x, t)$ , the average transverse position at longitudinal position  $z$ .

$$\Delta x' = \sum_{n=0}^N a_n k_n(z); \quad a_n = \frac{1}{H_n^2} \int \bar{x}(z, t) h_n(z) dz$$

(8)

Note that  $\Delta\psi$  is used to calculate  $\bar{x}(z, t)$



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$



detuning from e-cloud

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$

↑
↑

detuning from e-cloud
dipolar forces from e-cloud

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$

↑
↑

detuning from e-cloud
dipolar forces from e-cloud

First ansatz:  $\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j\Omega t} \Delta\psi(J_x, \theta_x, r, \phi)$ .

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

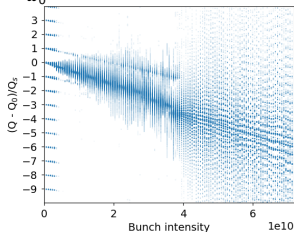
$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$

↑  
detuning from e-cloud

↑  
dipolar forces from e-cloud

First ansatz:  $\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j\Omega t} \Delta\psi(J_x, \theta_x, r, \phi)$ .

$(\frac{Re(\Omega)}{\omega_0} - Q_0)/Q_s$  is the **tune shift**



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

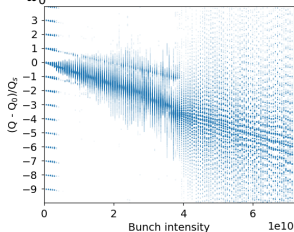
Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x \overset{\uparrow}{F_x^{coh}}(z, t)$$

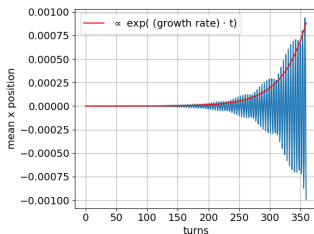
↑  
detuning from e-cloud
↑  
dipolar forces from e-cloud

First ansatz:  $\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{i\Omega t} \Delta\psi(J_x, \theta_x, r, \phi)$ .

$(\frac{Re(\Omega)}{\omega_0} - Q_0)/Q_s$  is the **tune shift**



$-Im(\Omega)$  is the **instability growth rate**



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

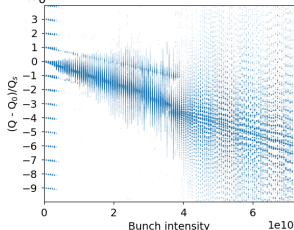
$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t)$$

↑
↑

detuning from e-cloud
dipolar forces from e-cloud

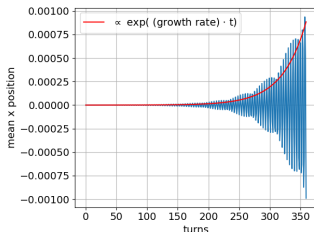
First ansatz:  $\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{i\Omega t} \Delta\psi(J_x, \theta_x, r, \phi)$ .

$(\frac{Re(\Omega)}{\omega_0} - Q_0)/Q_s$  is the **tune shift**



**Beam instabilities** are characterized by these qualities.

$-Im(\Omega)$  is the **instability growth rate**



$\omega_0$  is the angular revolution frequency,  $Q_0$  is the unperturbed tune and  $Q_s$  is the synchrotron frequency



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

Expand ansatz:

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j(\Omega t - \Delta\Phi(r, \phi))} f_1(J_x) e^{j\theta_x} \sum_{l=-\infty}^{\infty} e^{-jl\phi} W_l(r) \sum_{m=0}^{+\infty} b_{lm} f_{lm}(r)$$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

Expand ansatz:

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = \underbrace{e^{j(\Omega t - \Delta\Phi(r, \phi))}}_{\text{unknown}} f_1(J_x) e^{j\theta_x} \sum_{l=-\infty}^{\infty} e^{-jl\phi} W_l(r) \sum_{m=0}^{+\infty} \underbrace{b_{lm}}_{\text{unknown}} f_{lm}(r)$$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*



# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

Expand ansatz:

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j(\Omega t - \Delta\Phi(r, \phi))} f_1(J_x) e^{j\theta_x} \sum_{l=-\infty}^{\infty} e^{-jl\phi} W_l(r) \sum_{m=0}^{+\infty} b_{lm} f_{lm}(r)$$

↑ unknown     
 ↙ known for dipolar oscillation     
 ↑ unknown

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

Expand ansatz:

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j(\Omega t - \Delta\Phi(r, \phi))} f_1(J_x) e^{j\theta_x} \sum_{l=-\infty}^{\infty} e^{-jl\phi} W_l(r) \sum_{m=0}^{+\infty} b_{lm} f_{lm}(r)$$

↑ unknown     
 ↑ known for dipolar oscillation     
 ↑ unknown

↙ Phase shift term to be chosen

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

Expand ansatz:

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j(\Omega t - \Delta\Phi(r, \phi))} f_1(J_x) e^{j\theta_x} \sum_{l=-\infty}^{\infty} e^{-jl\phi} W_l(r) \sum_{m=0}^{+\infty} b_{lm} f_{lm}(r)$$

↑ unknown     
 ↑ known for dipolar oscillation     
 ↑ unknown

Phase shift term to be chosen

Choose  $\Delta\Phi$  to have the constraint:

$$\frac{\partial \Delta\Phi}{\partial \phi} = - \frac{\omega_0}{\omega_s} \Delta Q_\phi(r, \phi) \quad (10)$$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Solving the Vlasov Equation - Ansatz of $\Delta\psi$

Equation to solve:

$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (9)$$

Expand ansatz:

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j(\Omega t - \Delta\Phi(r, \phi))} f_1(J_x) e^{j\theta_x} \sum_{l=-\infty}^{\infty} e^{-jl\phi} W_l(r) \sum_{m=0}^{+\infty} b_{lm} f_{lm}(r)$$

↑ unknown     
 ← known for dipolar oscillation     
 ↑ unknown

Phase shift term to be chosen

Choose  $\Delta\Phi$  to have the constraint:

$$\frac{\partial \Delta\Phi}{\partial \phi} = - \frac{\omega_0}{\omega_s} \Delta Q_\Phi(r, \phi) \quad (10)$$

Detuning from e-cloud can be divided into a **detuning w longitudinal amplitude** and a **head-tail phase shift**

$$\Delta Q(z, \delta) = \Delta Q(r, \phi) = Q' \delta + \sum_{n=0}^{N_p} A_n z^n = \Delta Q_R(r) + \Delta Q_\Phi(r, \phi) \quad (11)$$

[2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*

# Solving the Vlasov Equation

The linerized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

Unkowns in red and terms including electron cloud forces in green



# Solving the Vlasov Equation

The linearized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

Unkowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packages.

# Solving the Vlasov Equation

The linearized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

Unkowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packages.

Each eigenmode of the Vlasov equation corresponds to a possible distortion  $\Delta\psi$ .

# Solving the Vlasov Equation

The linerized Vlasov Equation now becomes an **eigenvalue problem**:

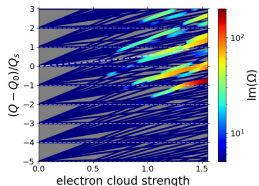
$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

Unknowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packages.

Each eigenmode of the Vlasov equation corresponds to a possible distortion  $\Delta\psi$ .

the tune shift is calculated from  $\text{Re}(\Omega)$



# Solving the Vlasov Equation

The linerized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

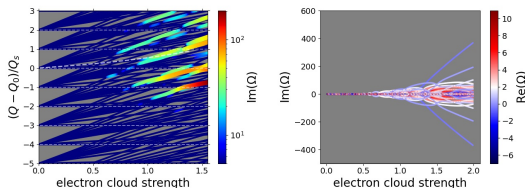
Unkowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packages.

Each eigenmode of the Vlasov equation corresponds to a possible distortion  $\Delta\psi$ .

the tune shift is calculated from  $Re(\Omega)$

$-Im(\Omega)$  is the instability growth rate



# Solving the Vlasov Equation

The linerized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

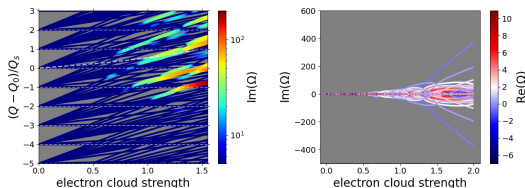
Unknowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packages.

Each eigenmode of the Vlasov equation corresponds to a possible distortion  $\Delta\psi$ .

the tune shift is calculated from  $\text{Re}(\Omega)$

$-\text{Im}(\Omega)$  is the instability growth rate



The *e-cloud strength* is a factor that multiply the e-cloud forces.

# Solving the Vlasov Equation

The linerized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (12)$$

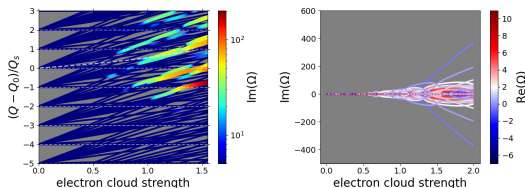
Unknowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packages.

Each eigenmode of the Vlasov equation corresponds to a possible distortion  $\Delta\psi$ .

the tune shift is calculated from  $\text{Re}(\Omega)$

$-\text{Im}(\Omega)$  is the instability growth rate



*The e-cloud strength* is a factor that multiply the e-cloud forces.

The Vlasov equation is solved for one e-cloud strength at a time yielding a set of  $\Omega$  which corresponds to a vertical line in the plots.

# Outline

Introduction

Simulation Model

Results with zero chromaticity

Results with chromaticity



# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.



# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.
- Typically  $\sim 10^6$  particles are tracked in a simulation.



# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.
- Typically  $\sim 10^6$  particles are tracked in a simulation.
- If an **LHC bunch** is simulated, each of the macro-particles represent about  $\sim 10^5$  protons.



# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.
- Typically  $\sim 10^6$  particles are tracked in a simulation.
- If an **LHC bunch** is simulated, each of the macro-particles represent about  $\sim 10^5$  protons.
- The accelerator is divided into segments, between which e-cloud forces are acting on the beam. These simulations use 8 segments.

# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.
- Typically  $\sim 10^6$  particles are tracked in a simulation.
- If an **LHC bunch** is simulated, each of the macro-particles represent about  $\sim 10^5$  protons.
- The accelerator is divided into segments, between which e-cloud forces are acting on the beam. These simulations use 8 segments.
- The conventional simulation method of this interaction is the **Particle-In-Cell** simulation method

# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.
- Typically  $\sim 10^6$  particles are tracked in a simulation.
- If an **LHC bunch** is simulated, each of the macro-particles represent about  $\sim 10^5$  protons.
- The accelerator is divided into segments, between which e-cloud forces are acting on the beam. These simulations use 8 segments.
- The conventional simulation method of this interaction is the **Particle-In-Cell** simulation method
- The linear model of forces can be used instead of PIC in macro-particle simulation to benchmark the Vlasov approach.

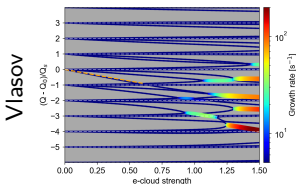
# Benchmark against macro-particle simulations

- Use **PyHEADTAIL**, a conventional macro-particle simulation code, to benchmark the Vlasov approach.
- Typically  $\sim 10^6$  particles are tracked in a simulation.
- If an **LHC bunch** is simulated, each of the macro-particles represent about  $\sim 10^5$  protons.
- The accelerator is divided into segments, between which e-cloud forces are acting on the beam. These simulations use 8 segments.
- The conventional simulation method of this interaction is the **Particle-In-Cell** simulation method
- The linear model of forces can be used instead of PIC in macro-particle simulation to benchmark the Vlasov approach.

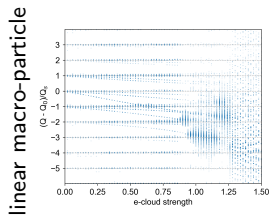
Will now compare the results from macro-particle tracking and the Vlasov approach, both using the linear model of e-cloud forces.

# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\Phi = 0$$

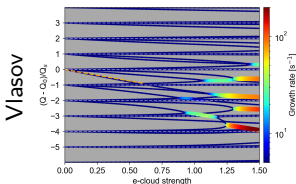


- No detuning is included, meaning  $\Delta Q_R = 0$ , and  $\Delta Q_\Phi = 0$

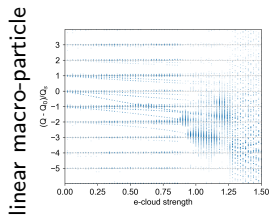


# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\Phi = 0$$



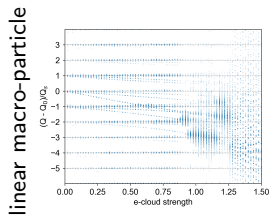
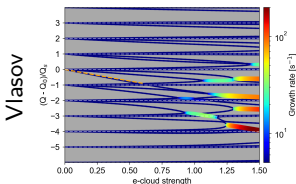
- No detuning is included, meaning  $\Delta Q_R = 0$ , and  $\Delta Q_\Phi = 0$
- The tune shift from macro-particle simulation results is calculated using the **SUSSIX** algorithm





# Results with zero chromaticity - Phase Shift

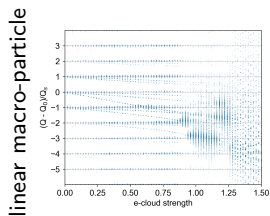
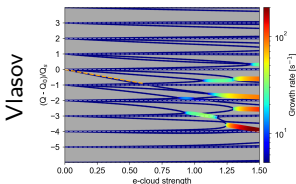
$$\Delta Q_R = 0, \Delta Q_\Phi = 0$$



- No detuning is included, meaning  $\Delta Q_R = 0$ , and  $\Delta Q_\Phi = 0$
- The tune shift from macro-particle simulation results is calculated using the SUSSIX algorithm
- The Vlasov modes gives information about the instability growth rate at each possible mode, this information is not available from the macro-particle simulations.

# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\Phi = 0$$



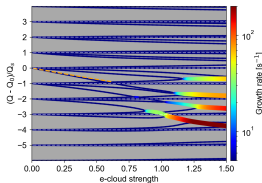
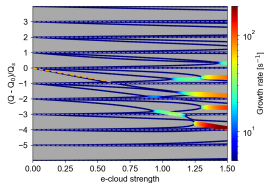
- No detuning is included, meaning  $\Delta Q_R = 0$ , and  $\Delta Q_\Phi = 0$
- The tune shift from macro-particle simulation results is calculated using the **SUSSIX** algorithm
- The **Vlasov modes** gives information about the **instability growth rate at each possible mode**, this information is **not available from the macro-particle simulations**.
- The **same modes are visible** in the results from the Vlasov approach and the macro-particle tracking.

# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\phi = 0$$

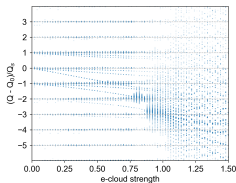
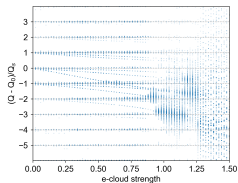
$$\Delta Q_R = 0, \Delta Q_\phi \neq 0$$

Vlasov



Now the head-tail phase shift  $\Delta Q_\phi$ , is included.

linear macro-particle

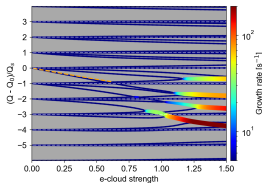
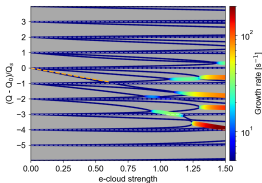


# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\phi = 0$$

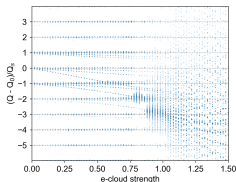
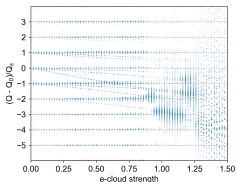
$$\Delta Q_R = 0, \Delta Q_\phi \neq 0$$

Vlasov



Now the head-tail phase shift  $\Delta Q_\phi$ , is included.

linear macro-particle



The same modes are visible in both plots!

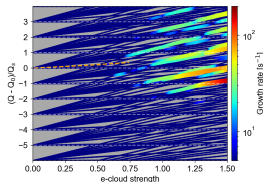
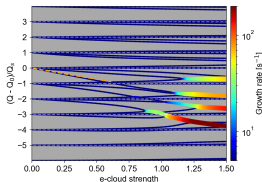
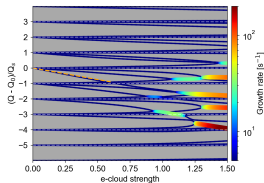
# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\Phi = 0$$

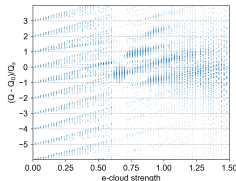
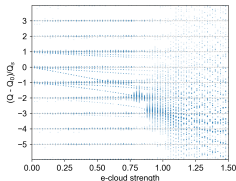
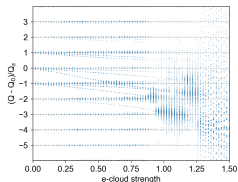
$$\Delta Q_R = 0, \Delta Q_\Phi \neq 0$$

$$\Delta Q_R \neq 0, \Delta Q_\Phi \neq 0$$

Vlasov



linear macro-particle



Include also **detuning with longitudinal amplitude.**

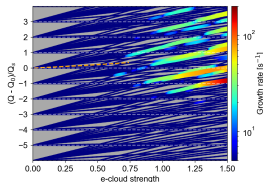
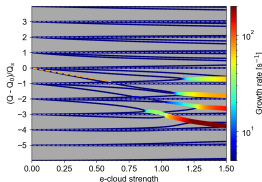
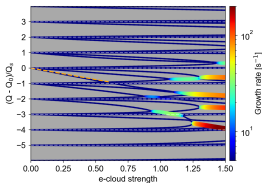
# Results with zero chromaticity - Phase Shift

$$\Delta Q_R = 0, \Delta Q_\phi = 0$$

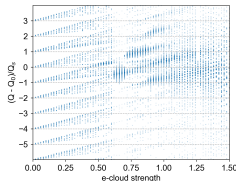
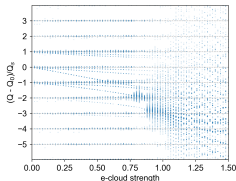
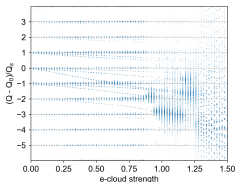
$$\Delta Q_R = 0, \Delta Q_\phi \neq 0$$

$$\Delta Q_R \neq 0, \Delta Q_\phi \neq 0$$

Vlasov



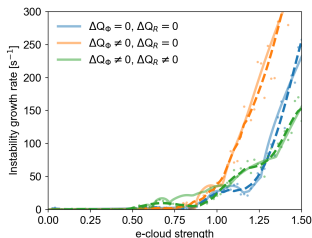
linear macro-particle



Include also detuning with longitudinal amplitude.

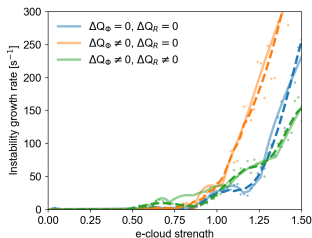
Modes agree when full detuning from e-cloud are included!

# Results with zero chromaticity - Growth Rate



- The instability growth rate of macro-particle simulations is reached by doing an exponential fit on the mean  $\times$  position (dashed lines)

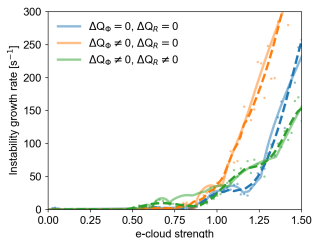
# Results with zero chromaticity - Growth Rate



- The instability growth rate of macro-particle simulations is reached by doing an exponential fit on the mean  $x$  position (dashed lines)
- The Vlasov mode with the highest instability growth rate, the worst mode, is plotted as the whole lines.



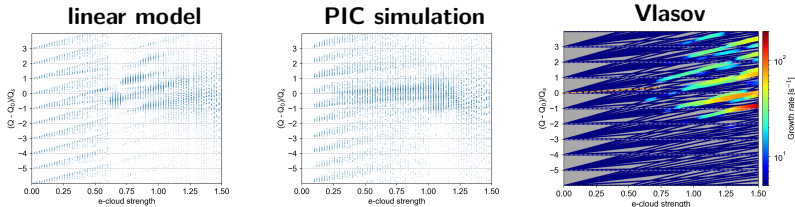
# Results with zero chromaticity - Growth Rate



- The instability growth rate of macro-particle simulations is reached by doing an exponential fit on the mean  $x$  position (dashed lines)
- The Vlasov mode with the highest instability growth rate, the worst mode, is plotted as the whole lines.

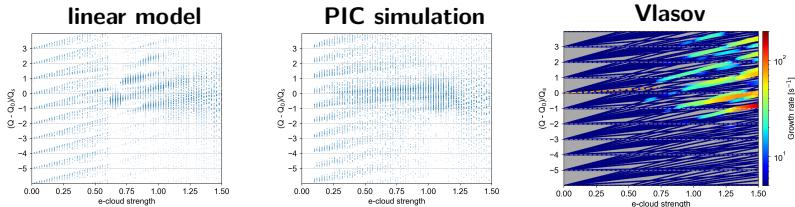
The instability growth rates agree well!

# Benchmark linear model with PIC simulations



The two left plots are made with **macro-particle simulations**.

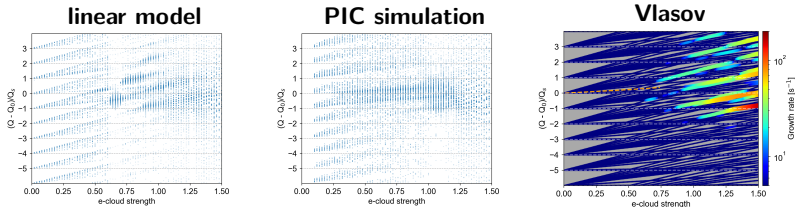
# Benchmark linear model with PIC simulations



The two left plots are made with **macro-particle simulations**.

All simulations methods have **fans of modes** for low e-cloud strengths.

# Benchmark linear model with PIC simulations

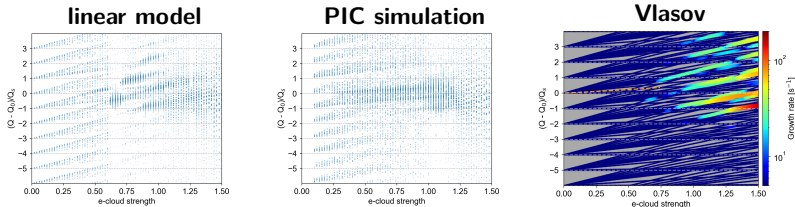


The two left plots are made with **macro-particle simulations**.

All simulations methods have **fans of modes** for low e-cloud strengths.  
**Non-linear effects** from e-cloud are present in middle plot which have a **stabilizing effect**.

Modes with a tune shift  $\Delta Q/Q_s > 0$  are **not visible in the PIC simulations**.

# Benchmark linear model with PIC simulations



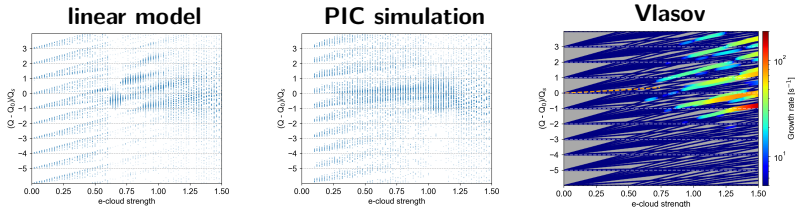
The two left plots are made with **macro-particle simulations**.

All simulations methods have **fans of modes** for low e-cloud strengths. **Non-linear effects** from e-cloud are present in middle plot which have a **stabilizing effect**.

Modes with a tune shift  $\Delta Q/Q_s > 0$  are **not visible in the PIC simulations**.

For **high e-cloud strength** ( $>1.25$ ), both macro-particle simulation methods give **similar results**

# Benchmark linear model with PIC simulations



The two left plots are made with **macro-particle simulations**.

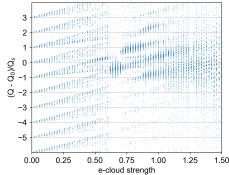
All simulations methods have **fans of modes** for low e-cloud strengths. **Non-linear effects** from e-cloud are present in middle plot which have a **stabilizing effect**.

Modes with a tune shift  $\Delta Q/Q_s > 0$  are **not visible in the PIC simulations**.

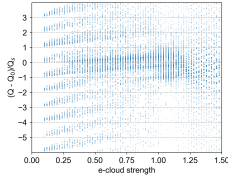
For **high e-cloud strength** ( $>1.25$ ), both macro-particle simulation methods give **similar results** and the **worst Vlasov mode** also have a tune shift between  $-1 < (Q - Q_0)/Q_s < 0$ .

# Benchmark linear model with PIC simulations

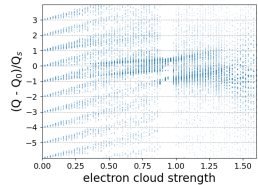
**linear model**



**PIC simulation**



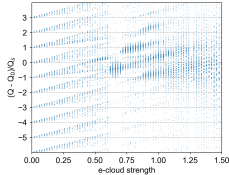
**linear model +  
static non-linear map**



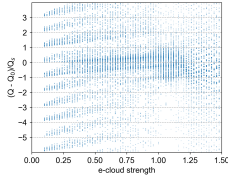
A **static non-linear map**, independent on  $z$ , can be made by removing the linear forces from the field map of the electron pinch and averaging along  $z$ .

# Benchmark linear model with PIC simulations

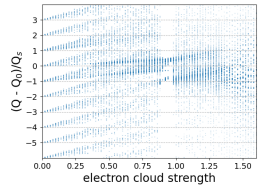
**linear model**



**PIC simulation**



**linear model +  
static non-linear map**



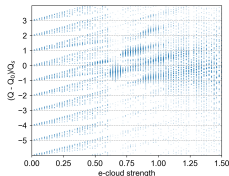
A **static non-linear map**, independent on  $z$ , can be made by removing the linear forces from the field map of the electron pinch and averaging along  $z$ .

This map is then **added to the linear e-cloud model**.

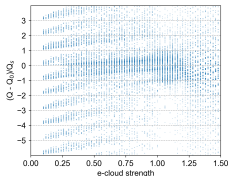


# Benchmark linear model with PIC simulations

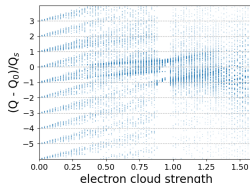
linear model



PIC simulation



linear model +  
static non-linear map

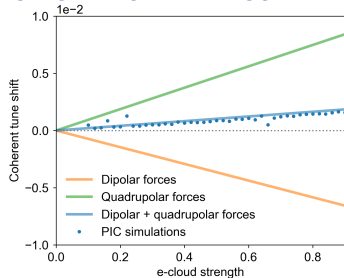


A **static non-linear map**, independent on  $z$ , can be made by removing the linear forces from the field map of the electron pinch and averaging along  $z$ .

This map is then **added to the linear e-cloud model**.

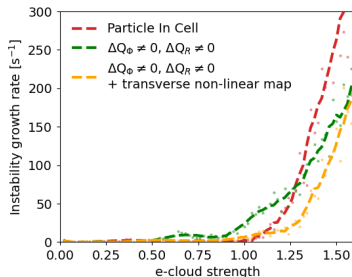
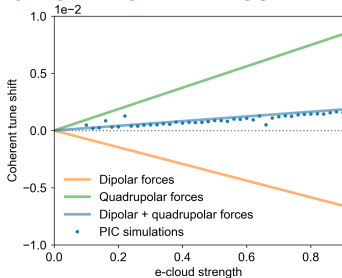
The results using the linear model + a static non-linear map are **similar to the results using PIC**: modes with  $\Delta Q/Q_s > 0$  are stabilized.

# Benchmark linear model with PIC simulations



An excellent agreement of tune shift between the Vlasov approach using **both** dipolar and quadrupolar forces and the **PIC** method.

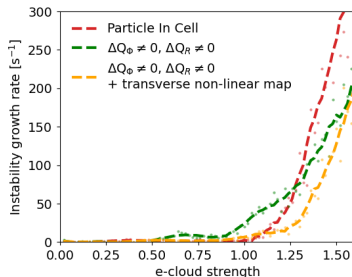
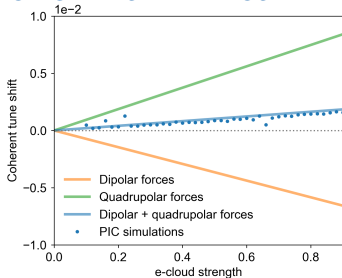
# Benchmark linear model with PIC simulations



An excellent agreement of tune shift between the Vlasov approach using **both** dipolar and quadrupolar forces and the **PIC** method.

The instability growth rate from **PIC** has **similar behaviour** to the instability growth rate from Vlasov and macro-particle simulations using the **linear model** of e-cloud.

# Benchmark linear model with PIC simulations



An excellent agreement of tune shift between the Vlasov approach using **both** dipolar and quadrupolar forces and the PIC method.

The instability growth rate from PIC has **similar behaviour** to the instability growth rate from Vlasov and macro-particle simulations using the **linear model** of e-cloud.

The static non-linear map has a **stabilizing effect** on the instability growth rate for low e-cloud strengths and the behaviour is **slightly more similar to the PIC** instability growth rate.

# Outline

Introduction

Simulation Model

Results with zero chromaticity

Results with chromaticity



# Chromaticity in Vlasov

Chromaticity is included in Vlasov together with the detuning caused by e-cloud forces:

$$\Delta Q(z, \delta) = Q' \delta + \sum_{n=0}^{N_p} A_n z^n \quad (13)$$

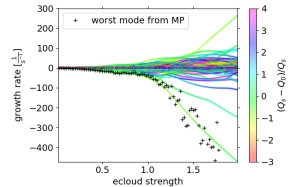
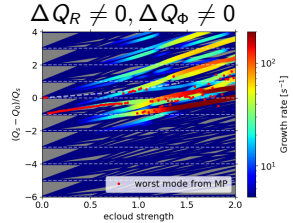
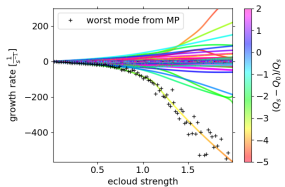
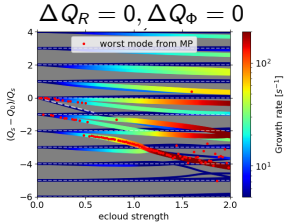
which is included in the phase shift term  $\Delta\Phi$  used in the ansatz of the distortion  $\Delta\psi$ :

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j\Omega t} \sum_{p=-\infty}^{\infty} f^p(J_x) e^{jp[\theta_x - \Delta\Phi(r, \phi)]} \sum_{l=-\infty}^{\infty} R_l^p(r) e^{-jl\phi} \quad (14)$$

Equation to solve:

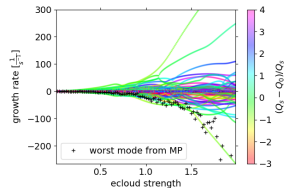
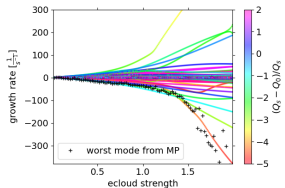
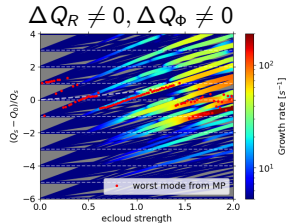
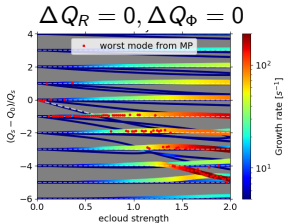
$$\frac{\partial \Delta\psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta\psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta\psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (15)$$

# Chromaticity = -5



The macro-particle simulation results follow the behaviour of the worst mode from Vlasov.

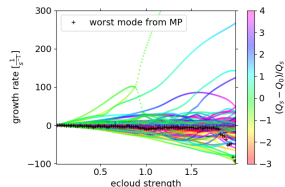
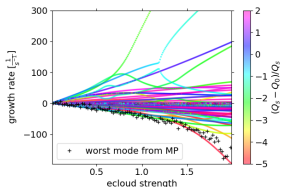
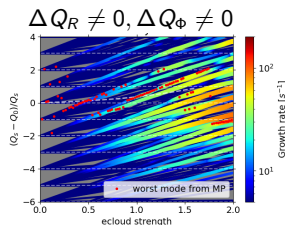
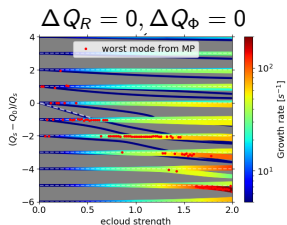
# Chromaticity = 5



Good agreement also for low and positive chromaticity.  
The instability growth rates are lower compared to chromaticity -5



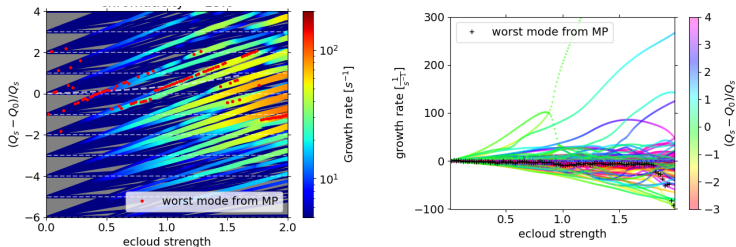
# Chromaticity = 15



Good agreement when **only dipolar forces** are included

# Chromaticity = 15

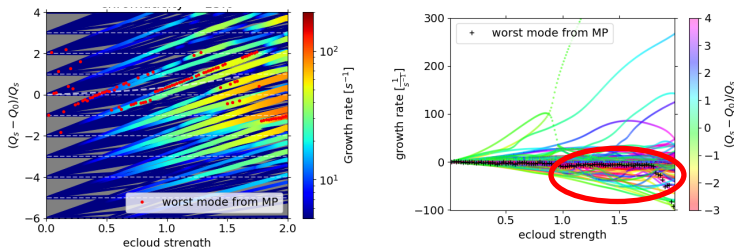
$$\Delta Q_R \neq 0, \Delta Q_\Phi \neq 0$$



The tune shift from macro-particle simulations does **not** follow the worst mode, **but always follow an existing Vlasov mode.**

# Chromaticity = 15

$$\Delta Q_R \neq 0, \Delta Q_\Phi \neq 0$$

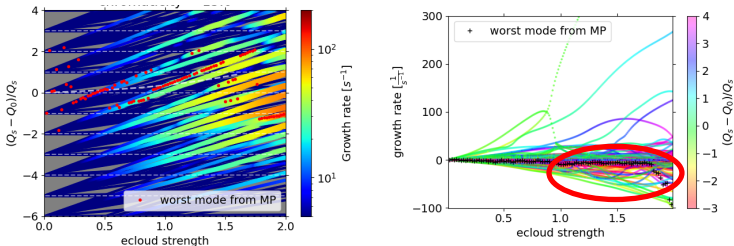


The tune shift from macro-particle simulations does **not** follow the worst mode, **but always follow an existing Vlasov mode**.

Instabilities in the **macro-particle simulations** for e-cloud strenghts <1.8 are **damped**.

# Chromaticity = 15

$$\Delta Q_R \neq 0, \Delta Q_\Phi \neq 0$$



The tune shift from macro-particle simulations does **not** follow the worst mode, **but always follow an existing Vlasov mode**.

Instabilities in the **macro-particle simulations** for e-cloud strengths < 1.8 are **damped**.

This is **not true** for the **Vlasov** modes.

# Summary and Conclusions

- A linear model of e-cloud forces including both dipolar and quadrupolar forces have been developed.



# Summary and Conclusions

- A linear model of e-cloud forces including both dipolar and quadrupolar forces have been developed.
- The linear model can be put into the Vlasov equation, which can be solved for instability modes.



# Summary and Conclusions

- A linear model of e-cloud forces including both dipolar and quadrupolar forces have been developed.
- The linear model can be put into the Vlasov equation, which can be solved for instability modes.
- The linear model can also be put into a conventional macro-particle tracking code for benchmarking.



# Summary and Conclusions

- A linear model of e-cloud forces including both dipolar and quadrupolar forces have been developed.
- The linear model can be put into the Vlasov equation, which can be solved for instability modes.
- The linear model can also be put into a conventional macro-particle tracking code for benchmarking.
- The instability modes found by Vlasov agree well with macro-particle simulation results using the same linear model for negative and low chromaticity.





# Summary and Conclusions

- A linear model of e-cloud forces including both dipolar and quadrupolar forces have been developed.
- The linear model can be put into the Vlasov equation, which can be solved for instability modes.
- The linear model can also be put into a conventional macro-particle tracking code for benchmarking.
- The instability modes found by Vlasov agree well with macro-particle simulation results using the same linear model for negative and low chromaticity.
- For high chromaticity, the mode visible in the macro-particle simulations are among the unstable Vlasov modes, but is not the worst.

# Summary and Conclusions

- A linear model of e-cloud forces including both dipolar and quadrupolar forces have been developed.
- The linear model can be put into the Vlasov equation, which can be solved for instability modes.
- The linear model can also be put into a conventional macro-particle tracking code for benchmarking.
- The instability modes found by Vlasov agree well with macro-particle simulation results using the same linear model for negative and low chromaticity.
- For high chromaticity, the mode visible in the macro-particle simulations are among the unstable Vlasov modes, but is not the worst.
- Currently checks with impedance driven instabilities are being done to further study this behaviour.

# References

- [1] G. Iadarola et al, 2018, *Electron Cloud Effects*
- [2] G. Iadarola, et. al. 2020, *Linearized method for the study of transverse instabilities driven by electron clouds*
- [3] N. Mounet, 2018, *Direct Vlasov Solvers*
- [4] L. Sabato, et. al. 2020, *Numerical simulation studies on single-bunch instabilities driven by electron clouds at the LHC*
- [6] F. Zimmermann, 2004, *Review of Single bunch instabilities driven by electron cloud*
- [7] G. Iadarola. et al., 2017, *Evolution of Python Tools for the Simulation of Electron Cloud Effects*
- [8] N. Mounet, 2017, *Vlasov Solvers and Macroparticle Simulations*
- [9] G. Rumolo and F. Zimmermann, 2002, *Electron Cloud Simulations: Beam Instabilities and Wakefields*
- [10] K. Ohmi et al, 2001, *Wake-Field and Fast Head-Tail Instability Caused by an Electron Cloud.*
- [11] E. Perevedentsev, 2002, *Head-Tail Instability Caused by Electron Cloud*



Thank you for your attention!



# e-cloud in Vlasov Equation - quadrupolar forces

Ansatz of  $\Delta\psi$ :

- Assume time dependence contained in  $e^{j\Omega t}$
- transverse Fourier decomposition (angle  $\theta_x$ )
- extract **phase shift term**  $e^{-jp\Delta\Phi(r,\phi)}$
- longitudinal Fourier decomposition (angle  $\phi$ )

$$\Delta\psi(J_x, \theta_x, r, \phi, t) = e^{j\Omega t} \sum_{p=-\infty}^{\infty} f^p(J_x) e^{jp[\theta_x - \Delta\Phi(r,\phi)]} \sum_{l=-\infty}^{\infty} R_l^p(r) e^{-jl\phi} \quad (16)$$

where the **unknown** terms are  $\Omega$  and  $f^p(J_x)$  and  $R_l^p(r)$ .

Choose  $\Delta\Phi$  in the **phase shift term** so that: 
$$\frac{\partial \Delta\Phi}{\partial \phi} = -\frac{\omega_0}{\omega_s} \Delta Q_\Phi(r, \phi) \quad (17)$$

The total detuning from e-cloud forces and chromaticity can be divided into **transverse detuning with longitudinal amplitude**,  $\Delta Q_R$ , and **head-tail phase shift**,  $\Delta Q_\Phi$ :

$$\Delta Q(z, \delta) = \Delta Q(r, \phi) = Q' \delta + \sum_{n=0}^{N_p} A_n z^n = \Delta Q_R(r) + \Delta Q_\Phi(r, \phi)$$

# Solving the linerized Vlasov Equation

$$\frac{\partial \Delta \Psi}{\partial t} - \omega_0(Q_{x0} + \Delta Q(r, \phi)) \frac{\partial \Delta \Psi}{\partial \theta_x} + \omega_s \frac{\partial \Delta \Psi}{\partial \phi} = - \frac{\eta g_0(r)}{\omega_s m_0 \gamma} \frac{df_0}{dJ_x} \sqrt{\frac{2J_x R}{Q_{x0}}} \sin \theta_x F_x^{coh}(z, t) \quad (18)$$

Using the ansatz of the distortion:

$$\Delta \psi(J_x, \theta_x, r, \phi, t) = e^{j\Omega t} \sum_{p=-\infty}^{\infty} f^p(J_x) e^{jp[\theta_x - \Delta \Phi(r, \phi)]} \sum_{l=-\infty}^{\infty} R_l^p(r) e^{-jl\phi} \quad (19)$$

after an additional decomposition of the radial functions

$R_l(r) = W_l(r) \sum_{m=0}^{\infty} b_{lm} f_{lm}$  where  $f_{lm}(r)$  are orthogonal and  $W_l(r)$  is arbitrary.

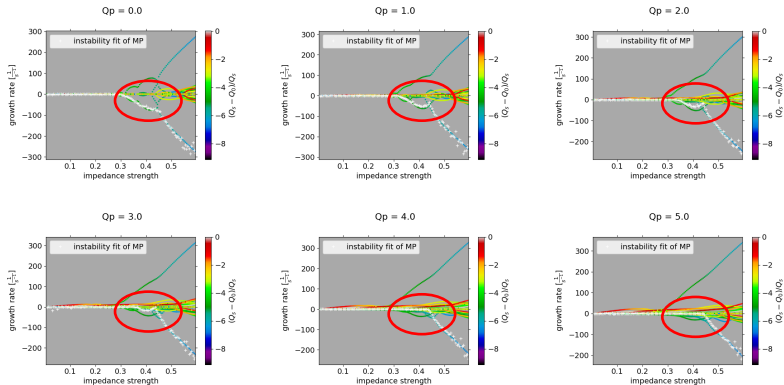
The linerized Vlasov Equation now becomes an **eigenvalue problem**:

$$b_{lm}(\Omega - Q_{x0}\omega_0 - l\omega_s) = \sum_{l'm'} (\mathbf{M}_{lm,l'm'} + \tilde{\mathbf{M}}_{lm,l'm'}) b_{l'm'} \quad (20)$$

Unkowns in red and terms including electron cloud forces in green

Solved by computing the matrices  $\mathbf{M}_{lm,l'm'}$  and  $\tilde{\mathbf{M}}_{lm,l'm'}$  and solve for "eigenvalue"  $\Omega$  and mode  $b_{lm}$  using standard linear algebra packets

# Discrepancy between simulations results



As chromaticity increases the vlasov mode for medium impedance strength is no longer visible in the macroparticle simulations results.