

Space-like method for hadronic vacuum polarization contributions to muon $g-2$

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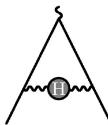
STRONG 2020 Virtual Workshop on “Spacelike and Timelike determination of the Hadronic Leading Order contribution
to the Muon $g-2$ ”

Padova

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- Very brief summary of time-like method for LO hadronic vacuum polarization contribution to muon $g-2$
- Space-like method for LO hadronic vacuum polarization contribution to muon $g-2$
- NLO hadronic vacuum polarization contributions

Reminder: Time-like method



leading order (LO) hadronic vacuum polarization contribution to muon $g-2$.

$$a_\mu^{\text{HVP}}(\text{LO}) = \frac{\alpha}{\pi^2} \int_{s_0=m_\pi^2}^{\infty} \frac{ds}{s} K^{(2)}(s/m^2) \text{Im}\Pi(s) = 6931(40) \times 10^{-11} \text{ (WP20)}$$

$$\text{optical theorem} \rightarrow \text{Im}\Pi(s) = \frac{\alpha}{3} R(s) \quad R(s) = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{4\pi\alpha^2/(3s)}$$

- $R(s)$ fluctuating at low energy due to resonance and particle production threshold effects
- $K^{(2)}(s/m^2)$: 1-loop QED $g-2$ contribution with a massive photon of mass \sqrt{s}

$$K^{(2)}(s/m^2) = \int_0^1 dx \frac{x^2(1-x)}{x^2 + (1-x)(s/m^2)}$$

$$K^{(2)}(z) = \frac{1}{2} - z + \left(\frac{z^2}{2} - z \right) \ln z + \frac{\ln y(z)}{\sqrt{z(z-4)}} \left(z - 2z^2 + \frac{z^3}{2} \right)$$

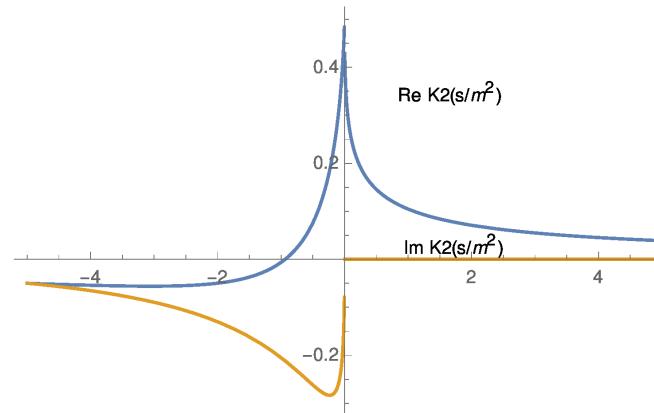
$$z = \frac{s}{m^2} \quad \rightarrow \quad (\text{conformal}) \text{ rationalizing variable} \quad y(z) = \frac{\sqrt{z} - \sqrt{z-4}}{\sqrt{z} + \sqrt{z-4}}$$

$$K^{(2)}(z) = \frac{1}{\pi} \int_{-\infty}^0 dz' \frac{\text{Im} K^{(2)}(z')}{z' - z}, \quad z > 0 \quad \frac{1}{\pi} \int_{s_0}^{\infty} \frac{ds}{s} \frac{\text{Im} \Pi(s)}{s - q^2} = \frac{\Pi(q^2)}{q^2}, \quad q^2 < 0$$

$$a_{\mu}^{\text{HVP}}(\text{LO}) = -\frac{\alpha}{\pi^2} \int_{-\infty}^0 \frac{dt}{t} \Pi(t) \text{Im} K^{(2)}(t/m^2)$$

we can evaluate the correct expression of the imaginary part for $z < 0$ from the exact $K^{(2)}(z)$ with the replacement $y(z) \rightarrow 1/y(z)$

$$\text{Im} K^{(2)}(z + i\epsilon) = \pi \theta(-z) \left[\frac{z^2}{2} - z + \frac{z - 2z^2 + \frac{z^3}{2}}{\sqrt{z(z-4)}} \right]$$



$$K^{(2)}(0) = 1/2, \quad K^{(2)}(z) \rightarrow 1/(3z) \quad z \rightarrow \infty$$

$\text{Im}K^{(2)}$ expressed in terms of $y(z)$ is simpler

$$\text{Im}K^{(2)}(z + i\epsilon) = \pi\theta(-z)F^{(2)}(y(z)) , \quad F^{(2)}(y) = \frac{y+1}{y-1}y^2$$

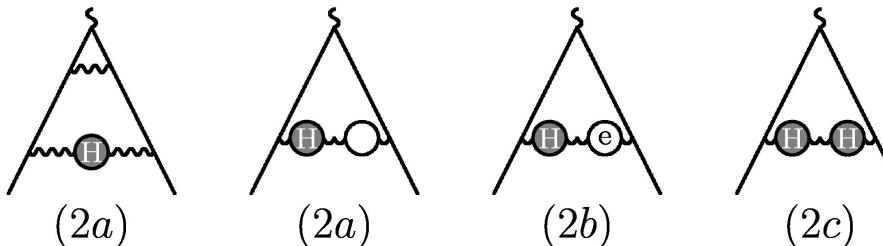
changing again variable in the dispersive integral $t \rightarrow y \rightarrow x$ ($t < 0 \rightarrow -1 < y < 0 \rightarrow 0 < x < 1$)

$$t(x) = m^2 \frac{x^2}{1-x} , \quad x = 1+y$$

$$a_\mu^{\text{HVP}}(\text{LO}) = \frac{\alpha}{\pi} \int_0^1 dx \kappa^{(2)}(x) \Delta\alpha_{\text{had}}(t(x)) \quad \text{Lautrup Peterman de Rafael 1975}$$

- $\kappa^{(2)}(x) = 1 - x$ simple space-like kernel
- $\Delta\alpha_{\text{had}}(t) = -\Pi(t)$ hadronic contribution to the running of the effective fine-structure constant in the space-like region

The above expression was proposed for the first time (Carloni Calame Passera Trentadue Venanzoni 2015) to determine a_μ^{HVP} measuring the electromagnetic effective coupling in the space-like region through scattering data.



- Class a: 1 HVP insertion in one photon line of 2-loop QED vertex diagrams
- Class b: 1 HVP insertion in the photon line of 2-loop QED vertex with one electron vacuum polarization
- Class c: 2 HVP insertion in the 1-loop QED vertex diagram

$$a_\mu^{\text{HVP}}(\text{NLO}; a) = -209.0 \times 10^{-11}$$

$$a_\mu^{\text{HVP}}(\text{NLO}; b) = +106.8 \times 10^{-11}$$

$$a_\mu^{\text{HVP}}(\text{NLO}; c) = +3.5 \times 10^{-11}$$

$$a_\mu^{\text{HVP}}(\text{NLO}; \text{total}) = -98.7(9) \times 10^{-11}$$

(Hagiwara Martin Nomura Toebner 2004, Hagiwara Liao Martin Nomura Toebner 2011, Kurz Liu Marquard Steinhauser 2014)

We write the time-like expression

$$a_\mu^{\text{HVP}}(\text{NLO}; a) = \frac{\alpha}{\pi^2} \int_{m_\pi^2}^{\infty} \frac{ds}{s} \, \mathbf{2}K^{(4)}(s/m^2) \, \text{Im}\Pi(s)$$

$2K^{(4)}$ is the anomaly from all 2-loop QED diagrams with 1 photon massless and 1 photon of mass \sqrt{s}

(factor 2 due to normalization chosen, as in the limit $z \rightarrow 0$ the anomaly is the $2 \times$ the 2-loop QED g-2 anomaly)

The analytical expression is lengthy

$$\begin{aligned}
 K^{(4)}(z) = & \left(\frac{z^2}{2} - \frac{7z}{6} + \frac{1}{2} \right) \left[-3\text{Li}_3(-y(z)) - 6\text{Li}_3(y(z)) + 2(\text{Li}_2(-y(z)) + 2\text{Li}_2(y(z))) \ln(y(z)) \right. \\
 & \left. + \ln(1-y(z)) \ln^2(y(z)) + \frac{1}{2} (\ln^2(y(z)) + \pi^2) \ln(y(z) + 1) \right] \\
 & + \frac{\left(-\frac{z^3}{6} + \frac{z^2}{4} - \frac{7z}{6} - \frac{4}{z-4} + \frac{13}{3} \right) \left(\text{Li}_2(-y(z)) + \frac{\ln^2(y(z))}{4} + \frac{\pi^2}{12} \right)}{\sqrt{(z-4)z}} \\
 & + \frac{\left(-\frac{7z^3}{12} + \frac{17z^2}{6} - 2z \right) \left(\text{Li}_2(y(z)) - \frac{1}{4} \ln^2(y(z)) + \ln(1-y(z)) \ln(y(z)) - \frac{\pi^2}{6} \right)}{\sqrt{(z-4)z}} \\
 & + \left(-\frac{29z^2}{96} + \frac{53z}{48} + \frac{2}{z-4} - \frac{1}{3z} + \frac{19}{24} \right) \ln^2(y(z)) + \frac{\left(\frac{23z^3}{144} - \frac{115z^2}{72} + \frac{127z}{36} - \frac{4}{3} \right) \ln(y(z))}{\sqrt{(z-4)z}} \\
 & + \frac{\left(-\frac{7z^3}{48} + \frac{17z^2}{24} - \frac{z}{2} \right) \ln(y(z)) \ln(z)}{\sqrt{(z-4)z}} + \frac{1}{6} \pi^2 \left(-\frac{z^2}{2} + \frac{5z}{24} - \frac{2}{z} + \frac{9}{4} \right) + \frac{5}{96} z^2 \ln^2(z) \\
 & + \left(\frac{23z^2}{144} - \frac{7z}{36} + \frac{1}{z-4} + \frac{19}{12} \right) \ln(z) + \frac{115z}{72} - \frac{139}{144} \quad \text{analytic: Barbieri Remiddi 1975}
 \end{aligned}$$

As in the LO case we write the dispersive relation for $K^{(4)}(z)$

$$K^{(4)}(z) = \frac{1}{\pi} \int_{-\infty}^0 dz' \frac{\text{Im}K^{(4)}(z')}{z' - z}, z > 0$$

$$a_\mu^{\text{HVP}}(\text{NLO};\text{a}) = -\frac{\alpha}{\pi^2} \int_{-\infty}^0 \frac{dt}{t} \Pi(t) \text{Im}K^{(4)}\left(\frac{t}{m^2}\right)$$

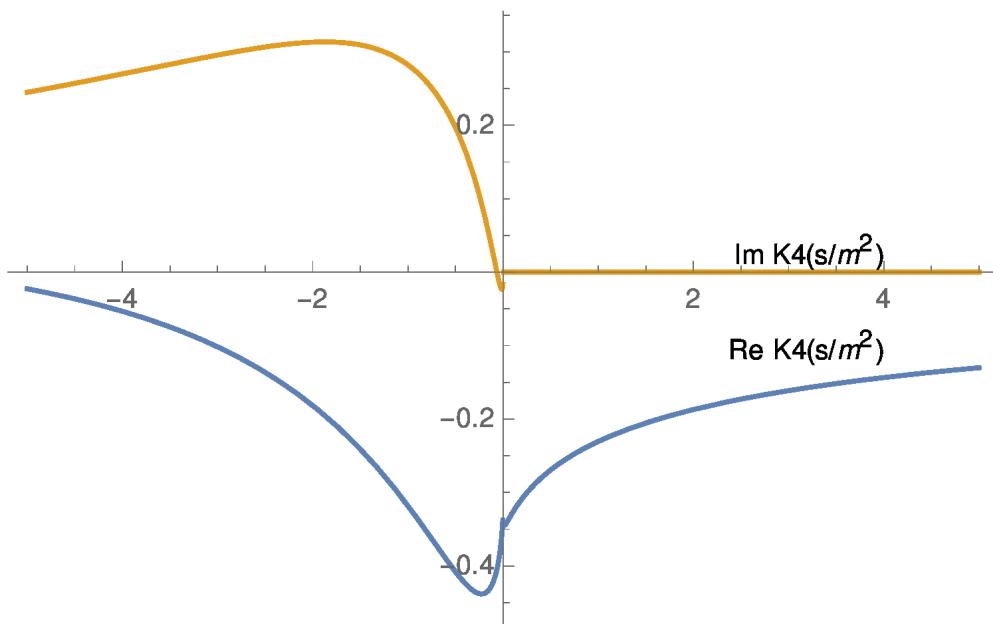
As in the LO case, the correct imaginary part is obtained from $K^{(4)}(z)$ replacing $y(z)$ with $1/y(z)$ for $z < 0$. The expression is more compact if written in terms of y instead of z :

$$\text{Im}K^{(4)}(z + i\epsilon) = \pi\theta(-z)F^{(4)}(y(z))$$

$$F^{(4)}(y) = \frac{(-3y^4 - 5y^3 - 7y^2 - 5y - 3)(2\text{Li}_2(-y) + 4\text{Li}_2(y) + \ln(-y) \ln((1-y)^2(y+1)))}{6y^2}$$

$$+ \frac{(y+1)(-y^3 + 7y^2 + 8y + 6) \ln(y+1)}{12y^2} + \frac{(-7y^4 - 8y^3 + 8y + 7) \ln(1-y)}{12y^2}$$

$$+ \frac{23y^6 - 37y^5 + 124y^4 - 86y^3 - 57y^2 + 99y + 78}{72(y-1)^2y(y+1)} + \frac{(12y^8 - 11y^7 - 78y^6 + 21y^5 + 4y^4 - 15y^3 + 13y + 6) \ln(-y)}{12(y-1)^3y(y+1)^2}, y < 0$$



$$K^{(4)}(0) = \frac{197}{144} + \frac{1}{12}\pi^2 - \frac{1}{2}\pi^2 \ln 2 + \frac{3}{4}\zeta(3) = -0.328479$$

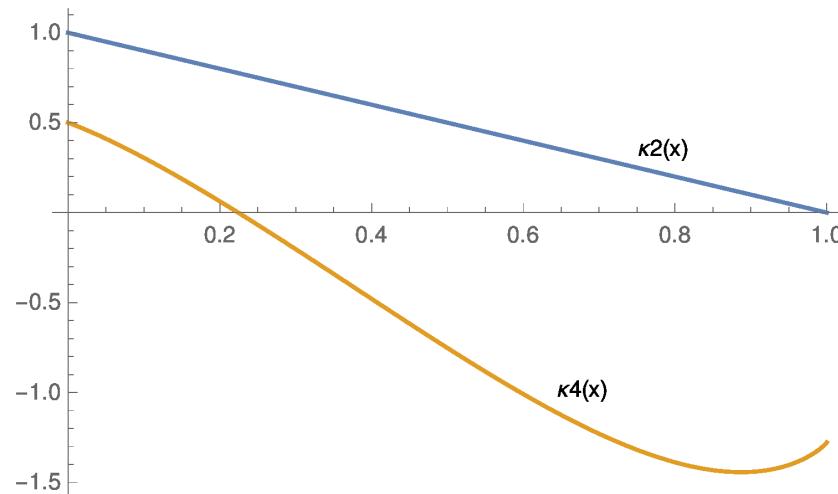
2-loop g-2

$$K^{(4)}(z) \rightarrow \frac{1}{z} \left(-\frac{23\ln(z)}{36} - \frac{\pi^2}{3} + \frac{223}{54} \right)$$

$$a_\mu^{\text{HVP}}(\text{NLO};\alpha_s) = \left(\frac{\alpha_s}{\pi}\right)^2 \int_0^1 dx \kappa^{(4)}(x) \Delta\alpha_{\text{had}}(t(x))$$

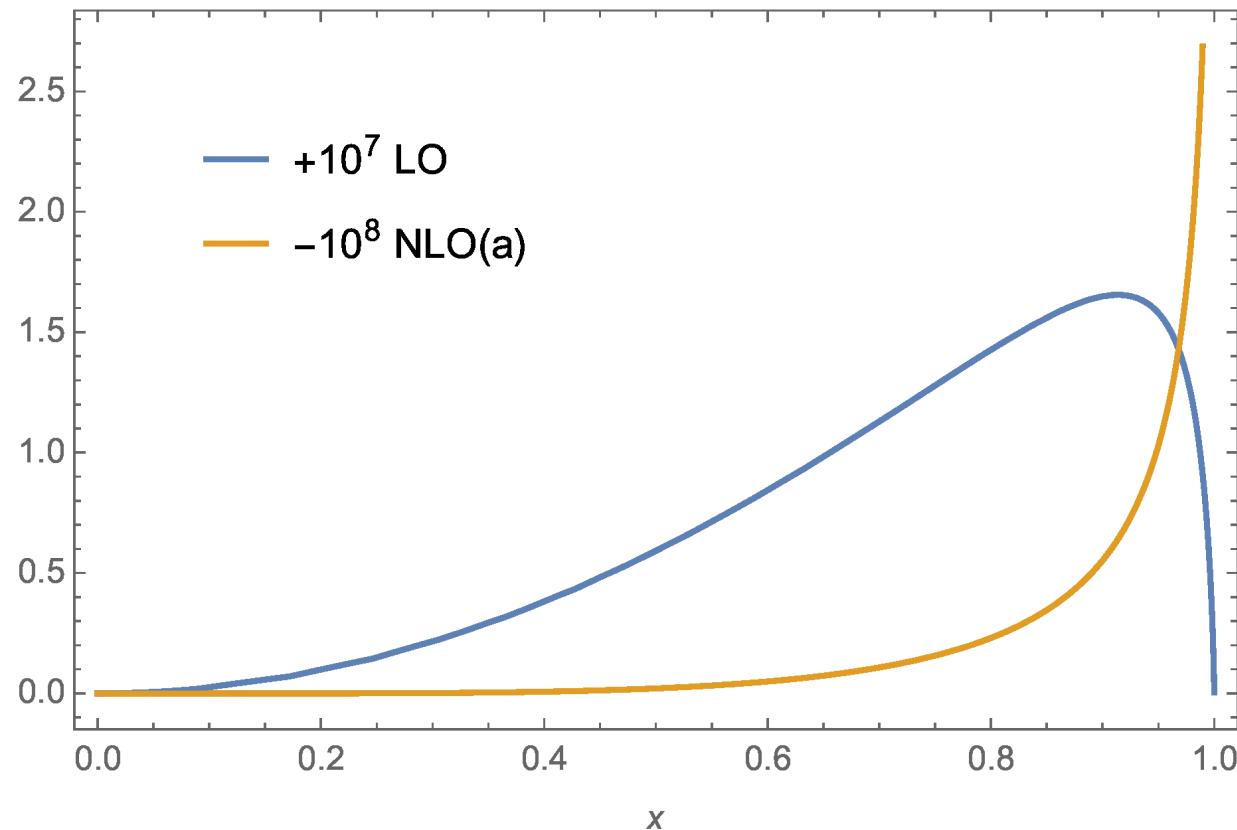
space-like kernel $\kappa^{(4)}(x)$:

$$\kappa^{(4)}(x) = \frac{2(2-x)}{x(x-1)} F^{(4)}(x-1)$$



$$\kappa^{(4)}(1) = -\frac{23}{18}$$

- $\kappa^{(4)}(x)$ provides stronger weight a large $q^2 < 0$ ($x \rightarrow 1$) than $\kappa^{(2)}(x)$



the integrands $(\alpha/\pi)\kappa^{(2)}(x)\Delta\alpha_{\text{had}}(t(x))$ (blue) $(\alpha/\pi)^2\kappa^{(4)}(x)\Delta\alpha_{\text{had}}(t(x))$ (orange)

- LO integrand has a peak at $x \approx 0.914$
- NLO has an (integrable) logarithmic singularity at $x \rightarrow 1$

Asymptotic expansion of $K^{(4)}(z)$ for large z (Lautrup 1997)

$$K^{(4)}(z) = \frac{1}{z} \left(-\frac{23 \ln(z)}{36} - \frac{\pi^2}{3} + \frac{223}{54} \right) + \frac{1}{z^2} \left(\frac{19 \ln^2(z)}{144} - \frac{367 \ln(z)}{216} - \frac{37\pi^2}{48} + \frac{8785}{1152} \right) \\ + \frac{1}{z^3} \left(\frac{141 \ln^2(z)}{80} - \frac{10079 \ln(z)}{3600} - \frac{883\pi^2}{240} + \frac{13072841}{432000} \right) + \dots$$

from these expansions it is possible to derive some approximations to the space-like kernel $\kappa^{(4)}(x)$ [Chakraborty Davies Kobonen Lepage VandeWater 2018] uses expansions for small $1/z$ up to $1/z^4$, with the ansatz from [Groote Körner Pivovarov 2002]

$$a_\mu = \left(\frac{\alpha}{\pi} \right)^2 \int_0^1 dw \left[(a_0 + a_1 w + a_2 w^2 + a_3 w^3) \Pi \left(\frac{m^2}{w} \right) + \frac{b_0 + b_1 w + b_2 w^2 + b_3 w^3}{w} \Pi(m^2 w) \right]$$

Unknown coefficients are found matching the terms of large z asymptotic expansion: $\{a_i\}$ match terms $(\ln z)/z^n$, $\{b_i\}$ match terms $1/z^n$.

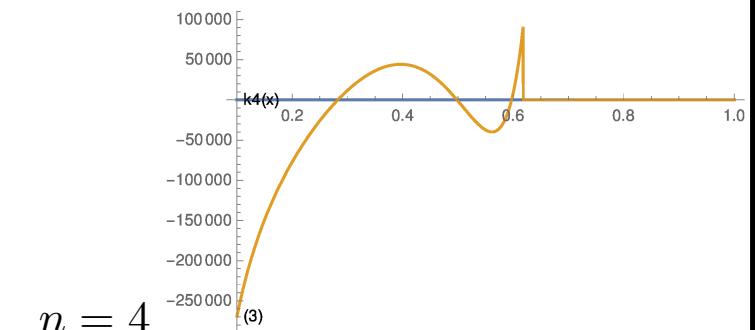
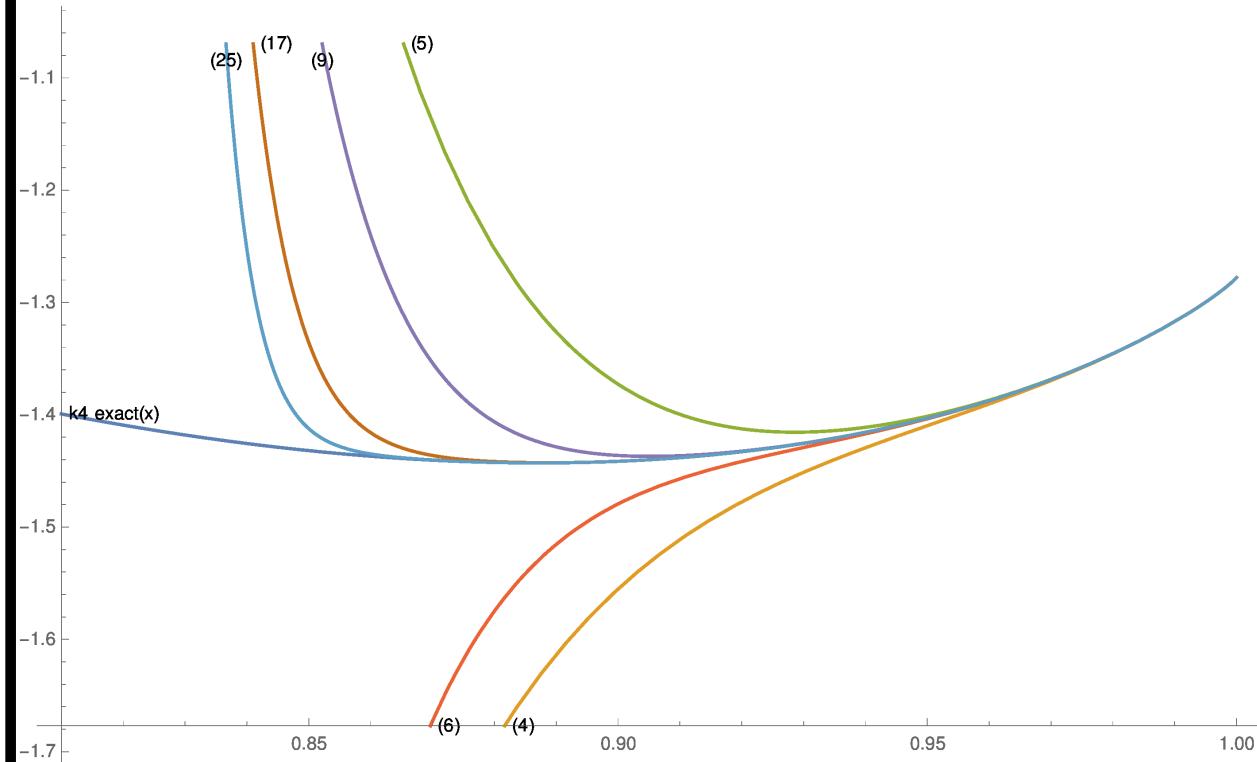
This approximation gives an error of 6% on $a_\mu^{\text{HVP}}(\text{NLO};a)$. The reason is that this ansatz does not fit terms with $\ln^2 z/z^n$.

In order to fit all terms, we added to the ansatz a logarithmic term like
 $+ \ln w (c_0 + c_1 w + c_2 w^2 + c_3 w^3) \Pi \left(\frac{m^2}{w} \right)$.

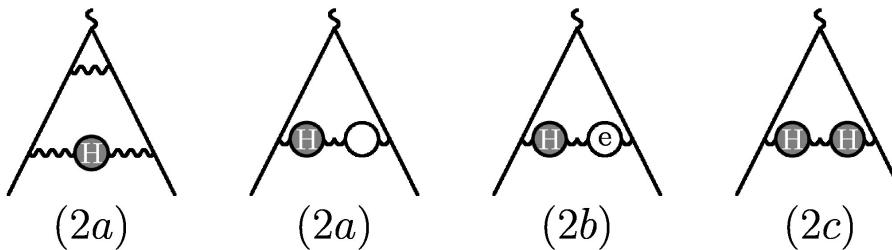
(E.Balzani, S.L. and M.Passera, in preparation)

NLO class a: approximated space-like kernels $\bar{\kappa}^{(4)}(x)$

Expansion of analytical $K_{(4)}(z) \rightarrow \bar{\kappa}_n^{(4)}(x)$ for $n_{terms} = 4$ to $n_{terms} = 25$.



- $\bar{\kappa}_n^{(4)}(x)$: good approximation for x near 1;
- Discontinuity for $x = (\sqrt{5} - 1)/2 \approx 0.618$
- Wild oscillations for small x , worse for large n .
- For $n = 25$ up to $\sim \pm 10^{30}!$ But the integral reproduces the exact result with error $10^{-20} \rightarrow$ deep numerical cancellations!
- Large n not necessary, $n = 4$ reproduces $a_\mu^{\text{HVP}}(\text{NLO};\text{a})$ with error $\lesssim 0.1\%$



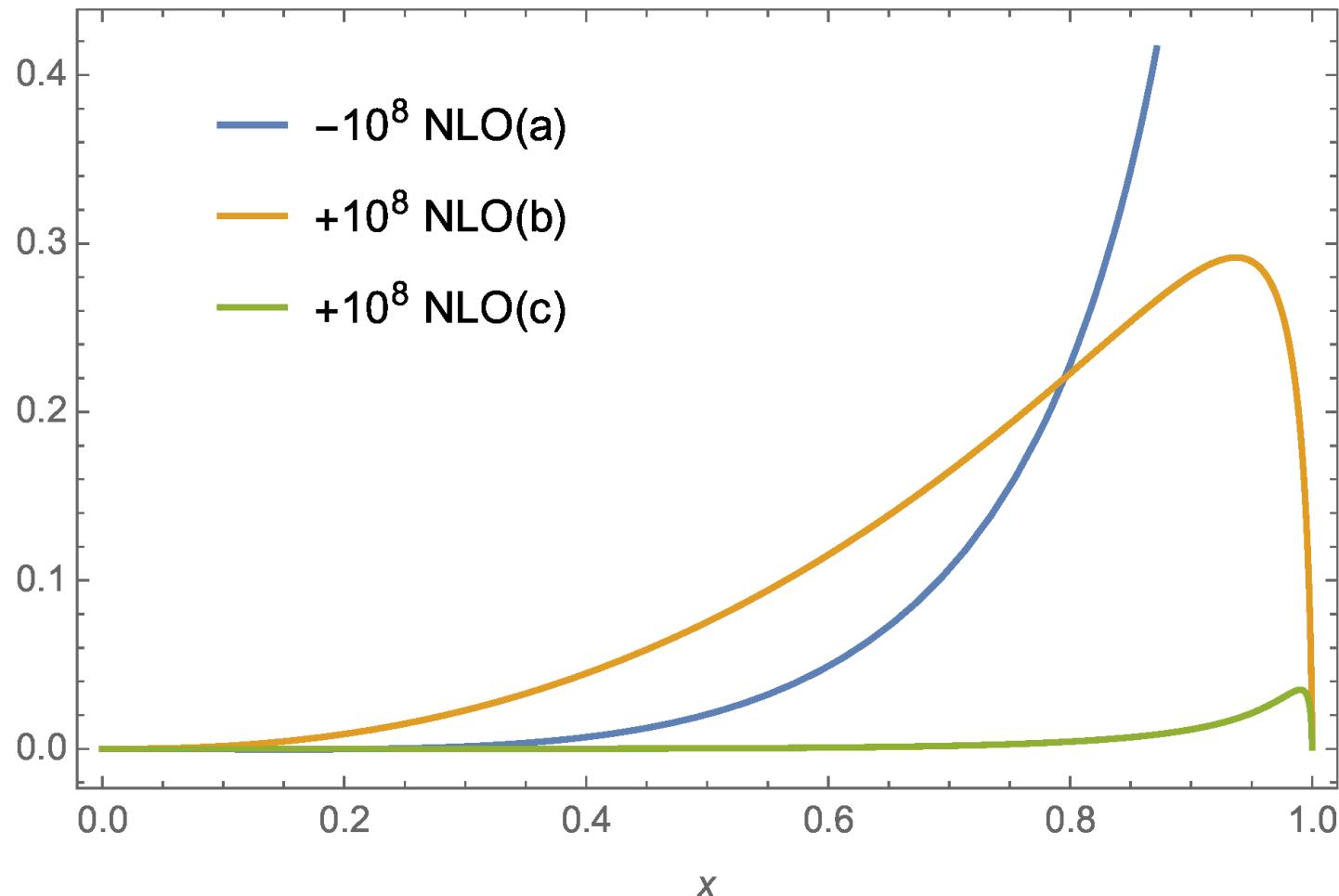
starting from timelike expressions one finds

$$a_\mu^{\text{HVP}}(\text{NLO};\text{b}) = \frac{\alpha}{\pi} \int_0^1 dx \kappa^{(2)}(x) \Delta\alpha_{\text{had}}(t(x)) 2 \left(\Delta\alpha_e^{(2)}(t(x)) + \Delta\alpha_\tau^{(2)}(t(x)) \right)$$

$$a_\mu^{\text{HVP}}(\text{NLO};\text{c}) = \frac{\alpha}{\pi} \int_0^1 dx \kappa^{(2)}(x) (\Delta\alpha_{\text{had}}(t(x)))^2$$

$\Pi_l^{(2)}(t) = -\Delta\alpha_l(t)$ renormalized QED vacuum polarization function

$$\Pi_l^{(2)}(t) = \left(\frac{\alpha}{\pi}\right) \left[\frac{8}{9} - \frac{\beta_l^2}{3} + \beta_l \left(\frac{1}{2} - \frac{\beta_l^2}{6} \right) \ln \frac{\beta_l - 1}{\beta_l + 1} \right] \quad \beta_l = \sqrt{1 - 4m_l^2/t}$$



the 3 NLO integrands

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